Final Results of CUPID-0 Phase I



Laura Cardani on behalf of the CUPID-0 collaboration **INFN-Roma**





The Lesson of CUORE



- Excellent energy resolution (0.3%)





Proved the feasibility of a tonne-scale experiment

ni (316), S. Copello (326),



From CUORE to CUPID



- Same CUORE infrastructure at LNGS
- Isotopic Enrichment (increase emitters without increasing volume)
- Background suppression by x100 wrt CUORE 10⁻² counts/keV/kg/yr
 - background of CUORE dominated by α particles —> particle ID



CUORE Upgrade with Particle IDentification



Particle ID in cryogenic calorimeters



Heat signal



- Couple each calorimeter to a light detector
- Different light yield enables particle identification
- Reject the (dominant) α background
- R&D of many years to identify the best technology
- C. Arnaboldi et al Astropart. Phys. 34 (2011) 344-353 C. Arnaboldi et al, Astropart.Phys. 34 (2010) 143-150 J.W. Beeman et al JINST 8 (2013) P05021 L. Cardani, et al., JINST 8, P10002 (2013). L. Cardani, et al., J. Phys. G 41, 075204 (2014). E. Armengaud, et al., JINST 10 (05), P05007 (2015). L. Berge, et al., JINST 9 P06004 (2014). . Bekker, et al., Astropart. Phys. 72, 38 (2016). D. R. Artusa et al., Eur. Phys. J. C 76 no.7, 364 (2016). E. Armengaud et al., Eur. Phys. J. C 77 no.11, 785 (2017). G.B Kim et al, Astropart. Phys 91, 105-112 (2017). G. Buse et al., Nucl. Instr. Meth. A 891 87 (2018) O. Azzolini et al. Eur. Phys. J. C 78 428 (2018) A. Barabash et al, Eur. Phys. J. C 76 487 (2018)
- J.W. Beeman et al. JINST 8 (2013) P07021
- L. Cardani et al. Appl.Phys.Lett. 107 (2015) 093508
- M. Biassoni et al. Eur.Phys.J. C75 (2015) no.10, 480
- L. Pattavina et al., Journal of Low Temp Phys 1-6 (2015)
- K.Schaeffner et. al, Astropart.Phys. 69 (2015) 30-36
- M. Willers et al., JINST 10 P03003 (2015)
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- F. Bellini et al. Appl.Phys.Lett. 110 (2017) 033504
- L. Cardani et al. Supercond.Sci.Technol. 31 (2018) 075002
- V. Novati et al. Nucl.Instrum.Meth. (2018)
- M.Barucci et al NIM A. 935 (2019) 150

and many others...

an many many others...



CUPID-0: absorber



Cryogenic Calorimeter for 0 $\nu\beta\beta$:

- Good bolometric performance
- Emission of scintillation light at 10 mK
- Grown from high Q-value emitter



CUPID-0: Zn⁸²Se

- Enrichment of Se at 96.3% in ⁸²Se
- Non-standard crystal (R&D to purify powder, grow large size crystals)
- Equipped with 3 x 2.8 x 1 mm³ NTD Ge thermistor









CUPID-0: light detector



Light Detector:

- ``Standard'' LD: not satisfactory at 10 mK
- Use a second calorimeter
- Light emitted by ZnSe: tens-hundreds of keV —> smaller capacitance



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- 4.4 cm x 170 μm
- Coated with SiO on a single face
- Equipped with 2 x 2.8 x 0.5 mm³ NTD Ge thermistor
- In previous test ~40 eV RMS, 1.8 ms rise-time, reproducible

CUPID-0: Ge wafer (produced by UMICORE)





The CUPID-0 Demonstrator

First medium-scale demonstrator of CUPID

- 24 ZnSe crystals enriched (95%) in ⁸²Se
- 2 natural ZnSe crystals
- 10.5 kg of ZnSe (3.8x10²⁵ nuclei)
- Hosted in the same CUORE-0 cryogenic facility (LNGS, Italy)















The CUPID-0 time-line



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Data Selection (1)

Reject non-physics events

(earthquakes, electronics noise...)



Select number of triggering crystals (1 for $0\nu\beta\beta$ searches)

Background in ROI (2800-3200 keV):

3.2 x 10⁻² counts/keV/kg/yr





Data Selection (2)

Add information of light detectors



Background in ROI (2800-3200 keV):

1.3 x 10⁻² counts/keV/kg/yr

Particle ID provides x3 suppression

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Data Selection (3)



Background in ROI (2800-3200 keV):

3.5 x 10⁻³ counts/keV/kg/yr

Lowest Background For cryogenic calorimeters

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Ovßß decay of ⁸²Se

• Total exposure:

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- 5.29 (kg x yr) of ⁸²Se
- 3.88x10²⁵ (emitters x yr) of ⁸²Se
- Final efficiency: (70 ± 1) %
 - **Probability of 0\nu\beta\beta electrons containment (81.0 ± 0.2)** %
 - **Trigger and energy reconstruction 99.5 %**
 - **Selection on heat (shape + time veto) 88%**
 - **Electrons selection: 98%**
- Energy resolution in ROI: (20.05 ± 0.34) keV

 $T_{1/2} > 3.5 \times 10^{24} \text{ yr} (90\% \text{ C.I.})$ Sensitivity on $T_{1/2} > 5.0 \times 10^{24} \text{ yr}$ (90% C.I.) Surpasses by more than x 10 previous limits

[counts/keV/kg/yr] PRL 123 032501 (2019) 10^{-4} 10^{-} 3200 3400 3800 3000 3600 2800Energy [keV] pdf 0.03Posterior 0.02 $\Gamma^{0v} < 0.20 \times 10^{-24} \text{ yr}^{-1} (90\% \text{ C.I.})$ $T_{1/2}^{0v} > 3.5 \times 10^{24} \text{ yr} (90\% \text{ C.I.})$ 0.015 0.01 0.005 0.15 0.2 0.25 0.3 0.35 0.4

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N.Casali et al





Background Model

Where does the residual (3.5x10⁻³ counts/keV/kg/yr) background come from?



+ higher multiplicity spectra to normalise cosmic rays

D. Chiesa et al Eur. Phys. J. C 79 583 (2019)

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- ~17% cryostat
- ~6% reflecting foil and holders



- ~44% muons
- ~33% contaminations ZnSe crystals
- ~17% cryostat
- ~6% reflecting foil and holders

Upgrade (January-May 2019)

- Muon-veto surrounding the cryostat (lateral +top)
- Removal of reflecting foils (coincidences)
- Addition of internal copper shield

CUPID-0 is now back in data-taking for its II scientific run



What Next







Conclusions and Perspectives



- First scientific run (June 2017 December 2018) concluded
- Most stringent limit on ⁸²Se $0\nu\beta\beta$ decay to fundamental and excited states of ⁸²Kr
- Lowest background for cryogenic calorimeters 3.5x10⁻³ counts/keV/kg/yr • Upgrade for second scientific run concluded

Reached the sensitivity to:

- Study the ⁸²Se $2\nu\beta\beta$ with unprecedented precision (HSD vs SSD)
- Improve sensitivity on $0\nu\beta\beta$ decay of ⁶⁴Zn and ⁷⁰Zn (by x100)
- Search for Lorentz violations in the $2\nu\beta\beta$ spectrum





https://cupid-0.lngs.infn.it







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Calibration of Energy Scale

Periodical (~ 4 days every month) calibration

with ²³²Th sources





Cross-check with 56Co calibration

(Q-value ~ 4.57 MeV, $T_{1/2}$ ~ 77 days)











Ονββ decay of ⁸²Se to excited states



Studied first with Ge spectroscopy of ⁸²Se powder [L. Pattavina et al. Eur.Phys.J. C75 (2015) no.12, 591]

Analysis repeated with CUPID-0 [L. Cardani et al. Eur.Phys.J. C78 (2018) no.11, 888]

$$\begin{split} &\Gamma(^{82}Se \to ^{82} Kr_{0_{1}^{+}}) < 8.55 \times 10^{-24} yr \\ &\Gamma(^{82}Se \to ^{82} Kr_{2_{1}^{+}}) < 6.25 \times 10^{-24} yr \\ &\Gamma(^{82}Se \to ^{82} Kr_{2_{2}^{+}}) < 8.25 \times 10^{-24} yr \end{split}$$

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