

Towards 100,000-pixel microcalorimeter arrays using multi-absorber transition-edge sensors

Presented by Stephen Smith NASA Goddard Space Flight Center

Key contributors:

NASA GSFC: J.S. Adams, S.R. Bandler, S. Beaumont, J.A. Chervenak, A.M. Datesman, M.E. Eckart (LLNL), F.M. Finkbeiner, R. Hummatov, R.L. Kelley, C.A. Kilbourne, A.R. Miniussi, F.S. Porter, J.S. Sadleir, K. Sakai, N. Wakeham, E.J. Wassell, M.C. Witthoeft. Lincoln Labs MIT: K. Ryu.



Microcalorimeters for imaging and high resolution spectroscopy

- Microcalorimeters are ideal detectors for next gen.
 X-ray space telescopes.
- NASA's Lynx X-ray Telescope
 - Proposed launch late 2030's.
 - 5' field-of-view + 0.5-1.0" angular resolution.
 - => 100,000 pixels, 25-50 μm pitch.
 - $\Delta E_{\text{FWHM}} < 3 \text{ eV} @ 7 \text{ keV}.$



- Lynx requirements extremely challenging.
 - Not practical if 1 pixel = 1 TES.
 - Prohibitively high density of wiring on detector chip.
 - Space needed for large # of bias circuit and readout components (shunt resistors, inductors, u-mux resonators etc).
- Practical solution: Exploring the use of multi-absorber TESs + 3d buried wiring layers.



Sub-µm multi-layer wiring

- Buried wiring layers.
 - Collaborating with MIT / Lincoln Laboratory.
 - High yield, planarized, multi-layer (0.5 μm width) Nb wiring.
 - TES fabricated at NASA GSFC on top of polished SiO₂ surface.
 - Enables high density of wiring within array.

See poster **11-94**. J. Chervenak et al. And later talk **221**. T. Stevenson et al. for fabrication details on this





Multi-absorber TES 'Hydras'.

- Multiple absorbers with a different thermal conductance to a single readout sensor.
 - Different characteristic pulse shape for each element within the hydra.
 - Enables position discrimination between pixels.
- Introduction of position discrimination comes at some trade with ΔE (1) Internal thermal fluctuation noise between absorbers and TES. (2) Reduced signal bandwidth relative to the noise.

 ΔE_{FWHM} typically 1.2 -1.3 x worse than single pixel with same total heat capacity



Lynx X-ray Microcalorimeter (LXM) Array Concept





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Fabricated first prototype Lynx arrays with 25-absorbers / TES.

- Example here shows 50,000 pixel array (half final requirement).
- Includes:
 - 25-pixel hydras with 25 μm and 50 μm absorber pitch.
 - Includes 3d buried wiring.
- No back-side Au heat-sink layers yet. (susceptible to thermal noise from x-talk)



125 μ m pitch

250 μ m pitch



25-pixel hydra layout - 25 μ m pitch absorbers

Hydra before absorber deposition $20 \ \mu m$ Mo/Au TES, $T_0 = 50 mK$ 125 μm Thermal Link Vary length of link to tune G. Absorber 25x Au absorbers, 25 μ m pitch, outline Absorber-link Absorbers cantilevered above TES contact region the substrate, contact links at

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Total C = 0.1 pJ/K

3.2 μ m thick.

single point.

Au links 1-2 μm widths,

300 nm thick



25-pixel hydra measured pulse shapes at 1.5 keV (Al-K α x-rays)



- Variety of different potential metrics to characterize pulse shape.
- Different discrimination algorithms studied in detail in Poster 153.
 Beaumont et al.
- Example here is simple but reasonably effective on all pixels.
 - Uses 2 parameterizations of the rise-time.
 - X-axis slow rise-time from smoothed pulse
 - Y-axis fast rise-time from unsmoothed pulse





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- Energy dependence still needs to be explore (0.2-7 keV).
 - Broader distributions at low E (worse signal to noise).
 - Possibly broader at higher E because of non-linear.

rise-time scatter plot for 25 pixel hydra AI-K α x-rays (1.5 keV) 10-50% Rise time (µs) 10⁻⁶ 12.0 11.6 11.8 12.2 12.4 12 14 16 18 20 20-80% Rise time with 200pt smooth (us)



25-pixel hydra ΔE_{FWHM} - 25 μm pitch absorbers

- Measured $\Delta E_{FWHM} = 1.66 \pm 0.02 eV$ (all pixels coadded).
- All pixels < 2eV, $\langle \Delta E_{FWHM} \rangle = 1.68 \pm 0.13 eV$.
- Estimate effect of additional x-talk noise as 0.2 eV => \sim 1.5 eV possible (with heatsinking).



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25-pixel hydra ΔE_{FWHM} - 50 μm pitch absorbers

- $\Delta E_{FWHM} \propto \sqrt{C} = > \sim 2x$ worse than 25 μ m pixel design.
- Measured $\Delta E_{FWHM} = 3.34 \pm 0.02 eV$ (all pixels coadded).
- Estimate effect of additional thermal noise as $0.4 \text{ eV} = > \sim 3 \text{ eV}$ possible.
- 25-pixels identifiable.



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Average measured pulses



Summary and outlook

• Demonstrated 1st proof-of-concept results from 25-pixel hydras with buried wiring.

- Achieved $\Delta E_{FWHM} = 1.66 \pm 0.02 eV$ for $25 \mu m$ version.
- Achieved $\Delta E_{FWHM} = 3.34 \pm 0.02 eV$ for 50 μm version.
- Position discrimination was straightforward at 1.5 keV.
- Demonstrates feasibility of 100,000 pixel array for Lynx.

• Still to do:

- Measure broad-band response up to 7 keV (resolution and position discrimination).
- Next generation design will include back heatsinking layer for x-talk mitigation.
- Study compatibility with state-of-the art readout systems (microwave mux).
 - Testing with higher inductance to slow-rise time and reduce slew-rate to match readout.





ADDITIONAL BACK-UP MATERIAL





Effect of increasing inductance to slow rise-time



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SEM Images







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Fig. 17 (a) Illustration showing how bump-bonded connections are made between the detector chip and the side panels containing the readout. (b) Illustration showing the geometry of the flex and bumps. (c) Scanning electron microscope image of some prototype indium bumps fabricated and bump-bonded at GSFC. (d) Photograph a hexagonal detector chip connected to several silicon chips with wire-bond pads using superconducting microstrip flex and bump-bond connections. (e) Set-up for testing the bump-bonded connections between two chips through flex with superconducting microstrip.





Stephen Smith