# The Medium and High Frequency Telescopes of LiteBIRD

Team LiteBIRD Join Study Group: => 211 researchers

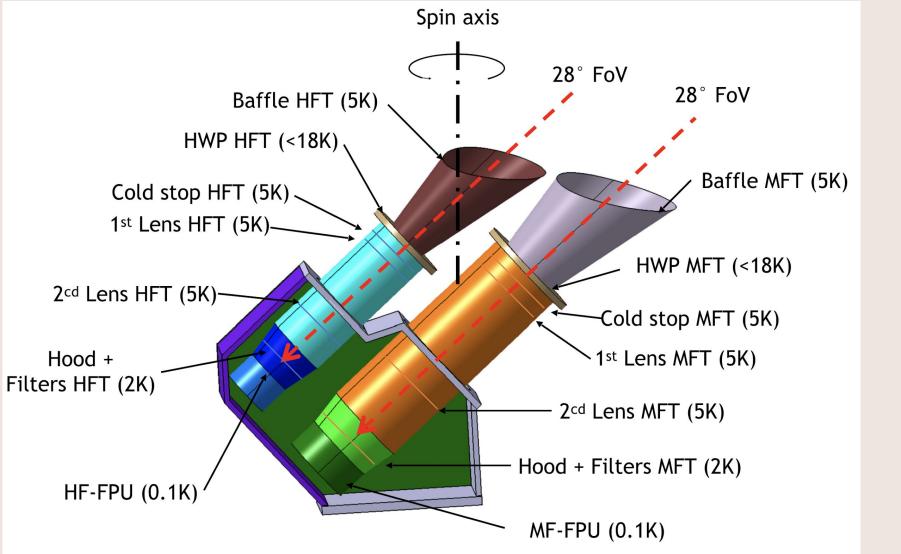
. Ade<sup>71</sup>, Y. Akiba<sup>19,30</sup>, D. Alonso<sup>71</sup>, K. Arnold<sup>8</sup>, J. Aumont<sup>44</sup>, J. Austermann<sup>5</sup>, C. Baccigalupi<sup>60,52,58</sup>, A. J R. Banerji<sup>66</sup>, R. B. Barreiro<sup>68</sup>, S. Basak<sup>49,60</sup>, J. Beall<sup>5</sup>, S. Beckman<sup>7</sup>, M. Bersanelli<sup>63,55</sup>, J. Borrill<sup>3,7</sup>, F. Boulanger<sup>42</sup>, M L. Brown<sup>75</sup>, M. Bucher<sup>38</sup>, A. Buzzelli<sup>62</sup>, E. Calabrese<sup>71</sup>, F. J. Casas<sup>68</sup>, A. Challinor<sup>74,72</sup>, V. Chan<sup>37</sup>, Y. Chinone<sup>7,18</sup>, J.-F. Cliche<sup>34</sup>, Columbro<sup>59,57</sup>, A. Cukierman<sup>7</sup>, D. Curtis<sup>7</sup>, P. Danto<sup>40</sup>, P. de Bernardis<sup>59,57</sup>, T. de Haan<sup>3</sup>, M. De Petris<sup>59,57</sup>, C. Dickinson<sup>75</sup>, M. Dobbs<sup>34</sup>, T. Dotani<sup>15</sup>, L. Duband<sup>39</sup>, A. Ducout<sup>18</sup>, S. Duff<sup>5</sup>, A. Duivenvoorden<sup>70</sup>, J.-M. Duval<sup>39</sup>, K. Ebisawa<sup>15</sup>, T. Elleflot<sup>8</sup>, H. Enokida<sup>13</sup>, H. K. Eriksen<sup>66</sup>, J. Errard<sup>38</sup>, T. Essinger-Hileman<sup>4</sup>, F. Finelli<sup>53</sup>, R. Flauger<sup>8</sup>, C. Franceschet<sup>63,55</sup>, U. Fuskeland<sup>66</sup>, K. Ganga<sup>38</sup>, J.-R. Gao<sup>65</sup>, R. Génova-Santos<sup>67,69</sup>, T. Ghigna<sup>76,18</sup>, A. Gomez<sup>40</sup>, M. L. Gradziel<sup>50</sup>, J. Grain<sup>42</sup>, F. Grupp<sup>48,47</sup>, A. Gruppuso<sup>53</sup>, J. E. Gudmundsson<sup>70</sup>, N. W. Halverson<sup>9</sup>, P. Hargrave<sup>71</sup>, T. Hasebe<sup>15</sup>, M. Hasegawa<sup>19,30</sup>, M. Hattori<sup>31</sup>, M. Hazumi<sup>19,15,18,30</sup>, S. Henrot-Versille<sup>45</sup>, D. Herranz<sup>68</sup>, C. Hill<sup>7,3</sup>, G. Hilton<sup>5</sup>, Y. Hirota<sup>13</sup>, E. Hivon<sup>41</sup>, R. Hlozek<sup>37</sup>, D.-T. Hoang<sup>1,77</sup>. Hubmayr<sup>5</sup>, K. Ichiki<sup>23</sup>, T. Iida<sup>18</sup>, H. Imada<sup>45</sup>, K. Ishimura<sup>32</sup>, H. Ishino<sup>26,19</sup>, G. C. Jaehnig<sup>9</sup>, M. Jones<sup>76</sup>, T. Kaga<sup>15</sup>, S. Kashima<sup>24</sup>, Y Kataoka<sup>26</sup>, N. Katayama<sup>18</sup>, T. Kawasaki<sup>20</sup>, R. Keskitalo<sup>3</sup>, A. Kibayashi<sup>26</sup>, T. Kikuchi<sup>10</sup>, K. Kimura<sup>27</sup>, T. Kisner<sup>3</sup>, Y. Kobayashi<sup>16</sup> N. Kogiso<sup>27</sup>, A. Kogut<sup>4</sup>, K. Kohri<sup>19</sup>, E. Komatsu<sup>46</sup>, K. Komatsu<sup>26</sup>, K. Konishi<sup>14</sup>, N. Krachmalnicoff<sup>60,52,58</sup>, C. L. Kuo<sup>6,2</sup>, N. Kurinsky<sup>6,2</sup>, A. Kushino<sup>22</sup>, M. Kuwata-Gonokami<sup>12</sup>, L. Lamagna<sup>59,57</sup>, M. Lattanzi<sup>54</sup>, A. T. Lee<sup>7,3</sup>, E. Linder<sup>3,7</sup>, B. Maffei<sup>42</sup>, D. Maino<sup>63,55</sup>, M. Maki<sup>19</sup>, A. Mangilli<sup>44</sup>, E. Martínez-González<sup>68</sup>, S. Masi<sup>59,57</sup>, R. Mathon<sup>44</sup>, T. Matsumura<sup>18</sup>, A. Mennella<sup>63,55</sup>, M. Migliaccio<sup>62,57</sup>, Y. Minami<sup>19</sup>, K. Mistuda<sup>15</sup>, D. Molinari<sup>61</sup>, L. Montier<sup>44</sup>, G. Morgante<sup>53</sup>, B. Mot<sup>44</sup>, Y. Murata<sup>15</sup>, J. A. Murphy<sup>50</sup> M. Nagai<sup>24</sup>, R. Nagata<sup>19</sup>, S. Nakamura<sup>33</sup>, T. Namikawa<sup>74</sup>, P. Natoli<sup>61,54</sup>, S. Nerval<sup>37</sup>, T. Nishibori<sup>17</sup>, H. Nishino<sup>28</sup>, Y. Nomura<sup>29</sup>, H. Noviello<sup>71</sup>, C. O'Sullivan<sup>50</sup>, H. Ochi<sup>33</sup>, H. Ogawa<sup>27</sup>, H. Ogawa<sup>15</sup>, H. Ohsaki<sup>13</sup>, I. Ohta<sup>21</sup>, N. Okada<sup>15</sup>, N. Okada<sup>27</sup>, L. Pagano<sup>61,54</sup>, A. Paiella<sup>59,57</sup>, D. Paoletti<sup>53</sup>, G. Patanchon<sup>38</sup>, F. Piacentini<sup>59,57</sup>, G. Pisano<sup>71</sup>, G. Polenta<sup>51</sup>, D. Poletti<sup>60,52,58</sup>, T. Prouvé<sup>39</sup>, G. Puglisi<sup>6</sup> D. Rambaud<sup>44</sup>, C. Raum<sup>7</sup>, S. Realini<sup>63,55</sup>, M. Remazeilles<sup>75</sup>, G. Roudil<sup>44</sup>, J. A. Rubiño-Martín<sup>67,69</sup>, M. Russell<sup>8</sup>, H. Sakurai<sup>12</sup>, Y. Sakurai<sup>18</sup>, M. Sandri<sup>53</sup>, G. Savini<sup>73</sup>, D. Scott<sup>36</sup>, Y. Sekimoto<sup>15,19,30</sup>, B. D. Sherwin<sup>74,72</sup>, K. Shinozaki<sup>17</sup>, M. Shiraishi<sup>25</sup>, P. Shirron<sup>4</sup> G. Signorelli<sup>56</sup>, G. Smecher<sup>35</sup>, P. Spizzi<sup>40</sup>, S. L. Stever<sup>18</sup>, R. Stompor<sup>38</sup>, S. Sugiyama<sup>29</sup>, A. Suzuki<sup>3</sup>, J. Suzuki<sup>19</sup>, E. Switzer<sup>4</sup>, R. Takaku<sup>12</sup>, H. Takakura<sup>11</sup>, S. Takakura<sup>18</sup>, Y. Takeda<sup>15</sup>, A. Taylor<sup>76</sup>, E. Taylor<sup>7</sup>, Y. Terao<sup>13</sup>, K. L. Thompson<sup>6,2</sup>, B. Thorne<sup>76</sup>, M. Tomasi<sup>63,55</sup>, H. Tomida<sup>15</sup>, N. Trappe<sup>50</sup>, M. Tristram<sup>45</sup>, M. Tsuji<sup>25</sup>, M. Tsujimoto<sup>15</sup>, C. Tucker<sup>71</sup>, J. Ullom<sup>5</sup>, S. Uozumi<sup>26</sup>, S. Utsunomiya<sup>18</sup>, J. Van Lanen<sup>5</sup>, G. Vermeulen<sup>43</sup>, P. Vielva<sup>68</sup>, F. Villa<sup>53</sup>, M. Vissers<sup>5</sup>, N. Vittorio<sup>62</sup>, F. Voisin<sup>38</sup>, I. Walker<sup>71</sup>, N. Watanabe<sup>20</sup>, I. Wehus<sup>66</sup>, J. Weller<sup>48,47</sup>, B. Westbrook<sup>7</sup>, B. Winter<sup>73</sup>, E. Wollack<sup>4</sup>, R. Yamamoto<sup>10</sup>, N. Y. Yamasaki<sup>15</sup>, M. anagisawa<sup>26</sup>, T. Yoshida<sup>15</sup>, J. Yumoto<sup>12</sup>, M. Zannoni<sup>64,55</sup>, A. Zonca<sup>8</sup>

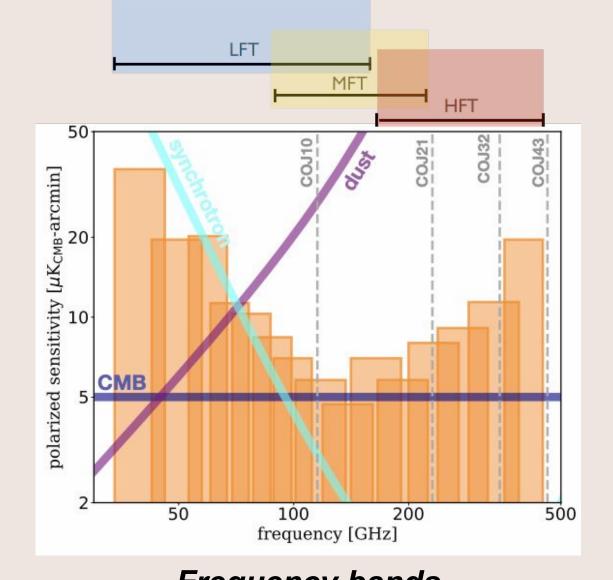
## **Mechanical Structure**

Cooled down to 5 K a thermoelastic study is performed to estimate the thermal contractions and reach the constraints on the optical alignements. Made of Aluminum, the mechanical structure of the

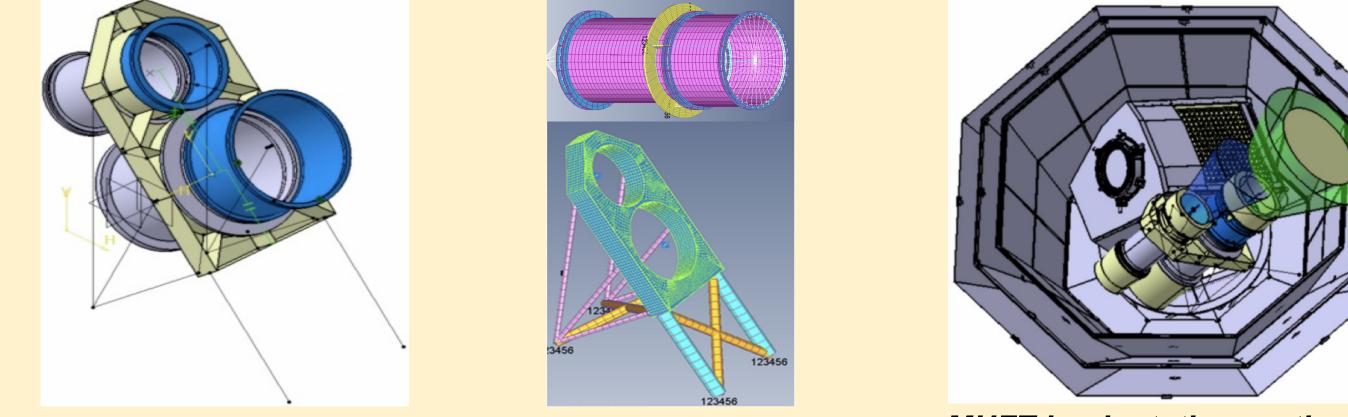
## **MHFT Instruments**

The European collaboration is responsible of the sub-K stage of the cryo-chain, and of the Mid- and High-Frequency Telescopes (MHFT), covering a large frequency range from 89 to 224 Hz and 166 to 448GHz, respectively. Both refractive telescopes are cooled down at 4.8K, with two half-wave plates rotating continuously. The two focal planes, cooled down at 100mK, consist, respectively, in 2074 and 1324 detectors delivered by the US.





telescopes ensure the thermal link between the MHFT's subsystems and the cryo-chain.

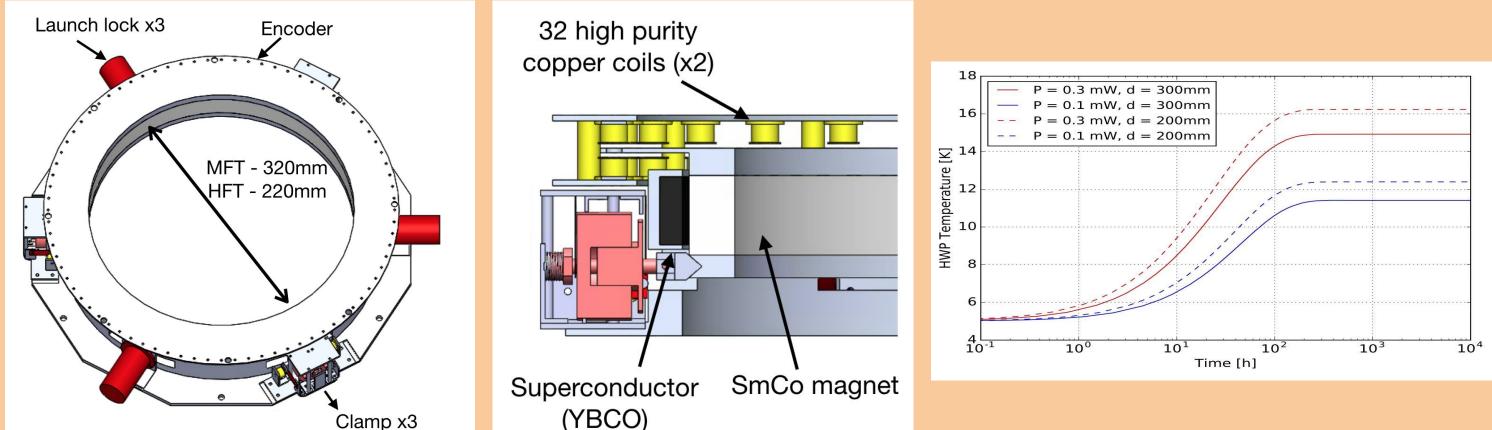


Mechanical structure

#### MHFT implantation on the PLM Finite Element Modeling

### **HWP** rotator

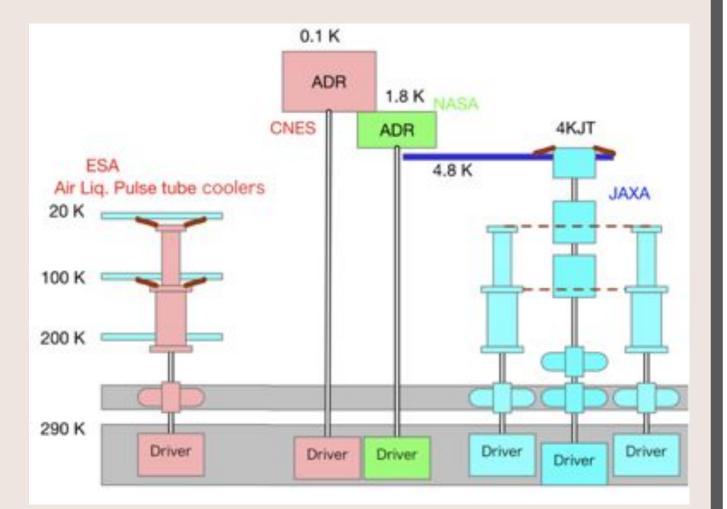
The HWP diameters of MFT and HFT are 320mm and 220mm, respectively. The Sapienza proposal is to have two scaled polarization modulators with the HWP mounted on a nearly frictionless bearing. This is obtained via the magnetic levitation of a permanent magnet rotor ring (cream grey) above a stator ring hosting an array of super-conducting bulks (black). The magnet ring is sandwiched between an Aluminum grooved ring, coupled with 3 actuators (and 3 launch locks), and an encoder ring.



#### **Ovwerview of the MFT and HFT telescopes**

The cryocahin design is based on passive cooling and mechanical coolers. A 4K stage is cooled with a <sup>4</sup>He Joule Thomson cooler from JAXA, precooled by a Stirling coolers. The current baseline of the sub-5K stage is based on a Multi-stage ADR (Adiabatic Demagnetization Refrigerator) which will be optimized to provide continuous cooling at 1.8 K, 300 mK and 100 mK. 7 ADR stages provided by NASA and CEA/CNES are required.

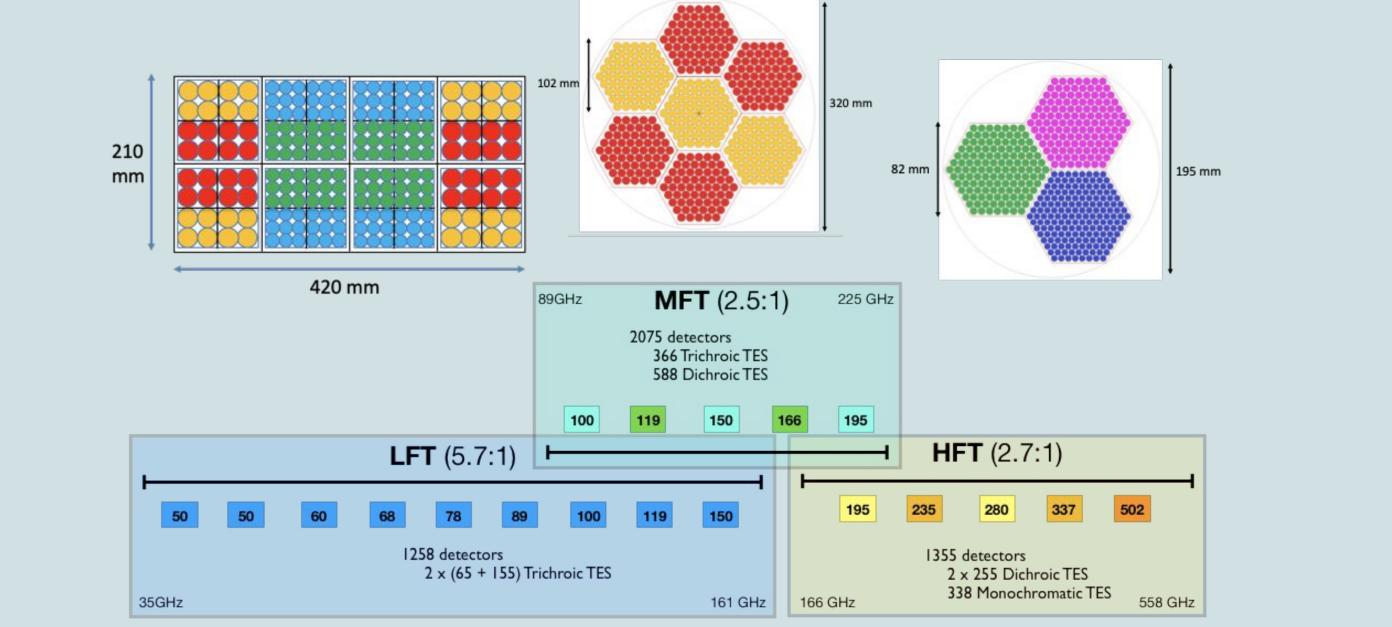
#### Frequency bands

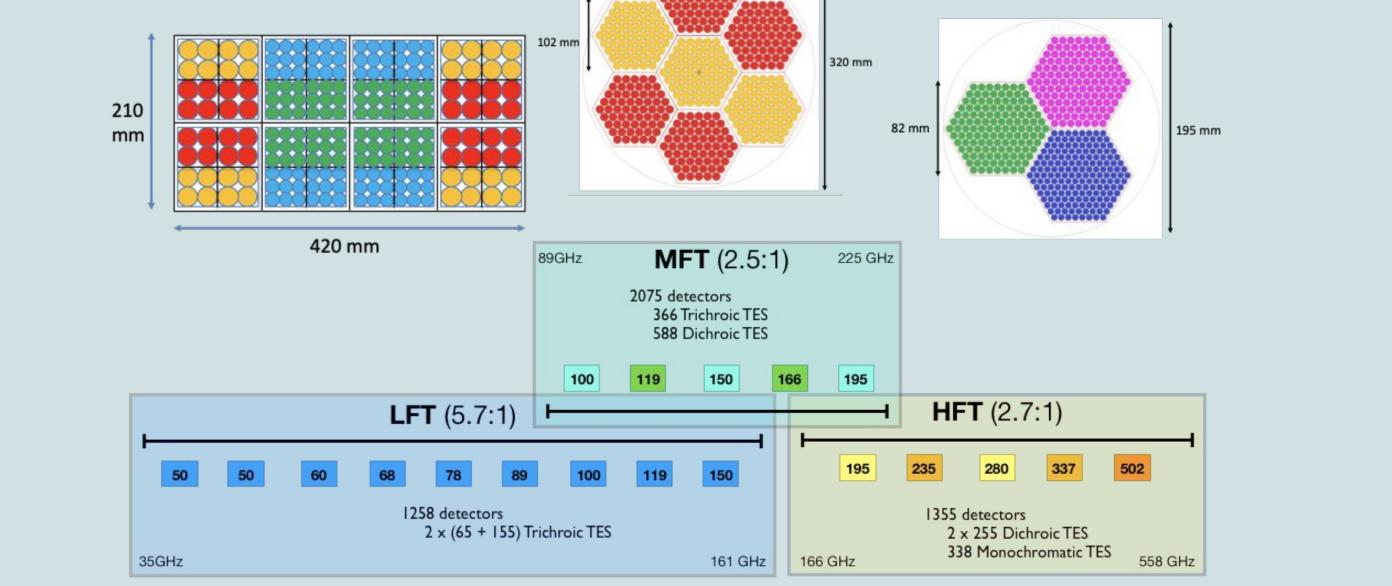


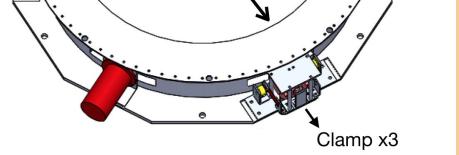
**Ovwerview of the cryo-chain** 

## **Focal Plane Units**

More than 4600 detectors cooled down to 100mK (more than 3400 for both MHFT & HFT ). Focal planes are made of TES bolometers coupled with a silicon lenslet & a sinuous antenna for MFT and coupled with silicon platelet for HFT. Each focal plane is composed of Monochromatic, Dichroic and Trichoic pixels sensitive to the polarisation.







The rotation is performed thanks to 32 coupled coils on the stator, paired with 8 small SmCo magnets on the rotor. Eddy currents produced on the rotor will heat up the HWP. For its thermal model we assumed:

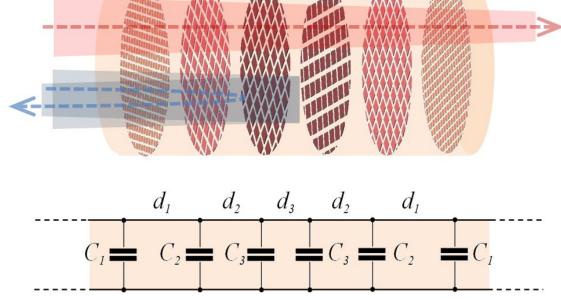
- motor composed by 32 coupled copper coils (RRR > 100)
- ~10 mA current needed in the motor
- ~0.1 mW dissipated during operation on the rotor
- Emissivities: HWP = 0.03 and AI = 0.4.

After a drift of a few hours, the expected temperatures are both below 20K A similar concept is being used for the LSPE-SWIPE mission.

### **Quasi-optical components**

#### **Filters**

### **Embedded Mesh-Filter**

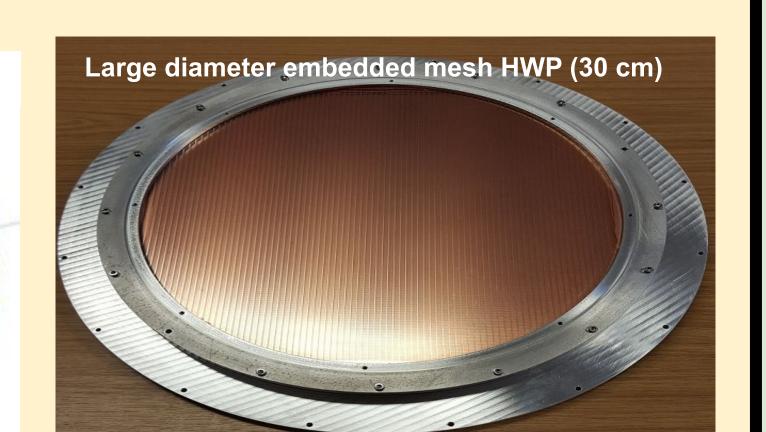


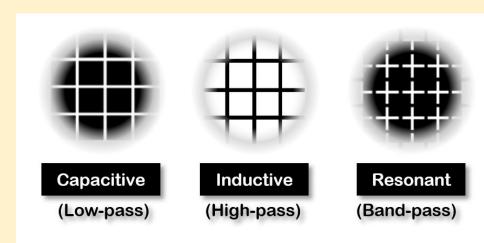


#### **Mesh-HWPs**

Pol 2

# **Embedded Mesh-HWP** Pol 1





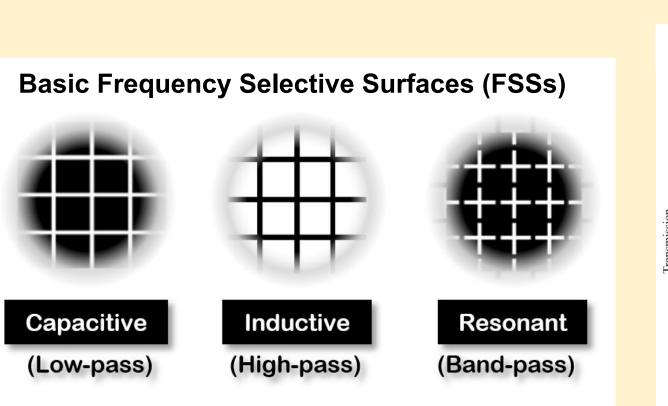
- Polypropylene embedded mesh-filter technology.
- Thermal shielding and spectral bands definition.
- Space qualified technology employed in many other ground-based, balloon-borne and satellite instruments.
- Filter diameters up to 500mm

#### Lenses

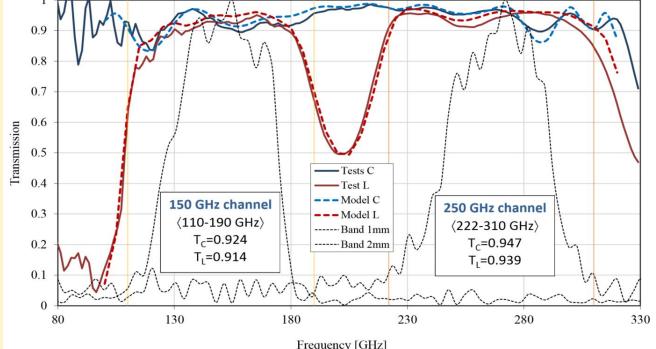
- Millimetre-wave Anti-Reflection Coated dielectric lenses
- Coating made with <sup>1</sup>/<sub>4</sub> wave Porous PTFE
- Process developed and refined under ESA contract
- Deployed on several ground & balloon-based experiments
- MFT and HFT telescopes based on two A/R coated polypropylene lenses



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#### **Mesh-HWP C- & L-axis transmissions example**



- Solution based on polypropylene embedded mesh-filter technology
- Anisotropic grids providing phase-shifts in opposite directions
- Employed in ground-based and balloon-borne instruments
- Further performance developments funded by ESA contract

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