

*ArduSiPM (Aduino+SiPM)*  
*Un rivelatore di raggi cosmici e radiazioni nucleari*  
*Utilizzo per divulgazione e didattica.*



# ArduSiPM

Low cost Cosmic ray and  
Nuclear Radiation Detector



Cosmic Ray Survey



Medical application



Dosimetry



Nuclear Waste  
monitor

Dr. Valerio Bocci  
Istituto Nazionale Fisica Nucleare  
Sezione di Roma

# Application Example 1: Intraoperative $\beta^-$ - Detecting Probe

nature.com > scientific reports > articles > article



A novel radioguided surgery technique exploiting  $\beta^-$  decays

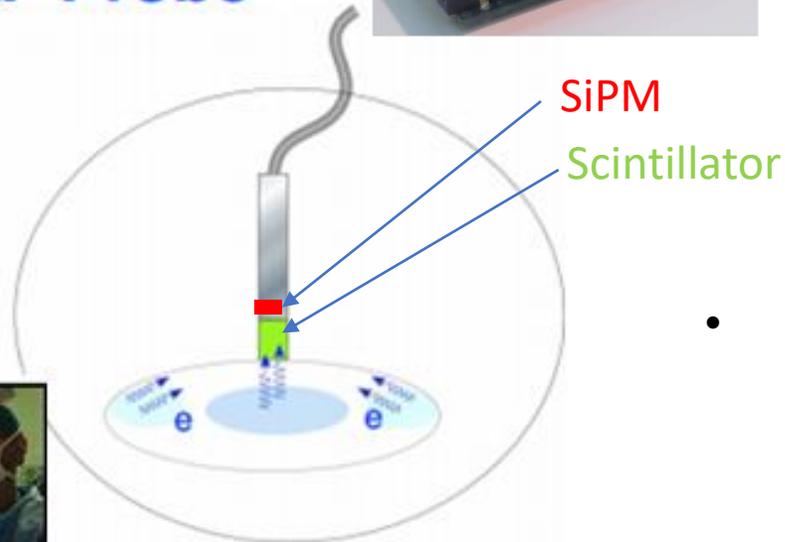
E. Solfaroli Camillocci, G. Baroni, F. Bellini, V. Bocci, F. Collamati, M. Cremonesi, E. De Lucia, P. Ferroli, S. Fiore, C. M. Grana, M. Marafini, I. Mattei, S. Morganti, G. Paganelli, V. Patera, L. Piersanti, L. Recchia, A. Russomando, M. Schiariti, A. Sarti, A. Sciubba, C. Voena & R. Faccini

**Beta- Probe**

ArduSiPM



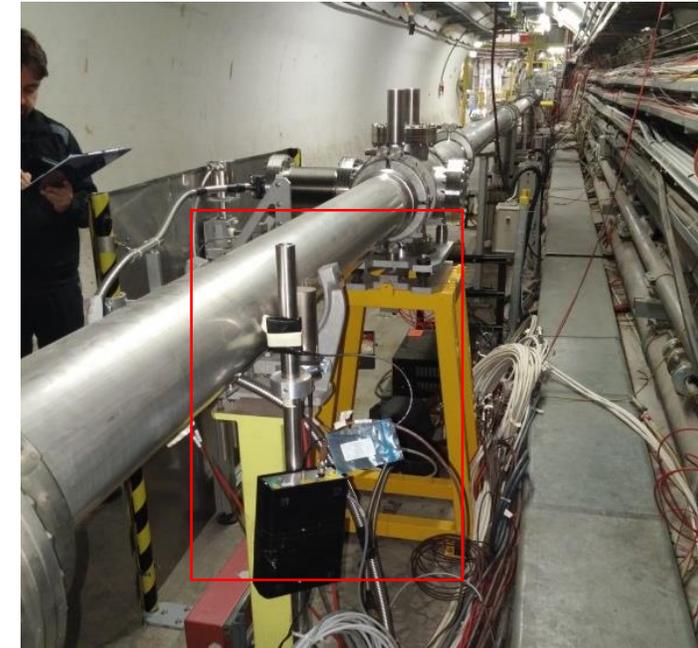
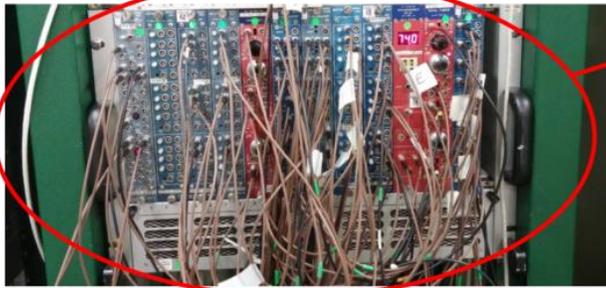
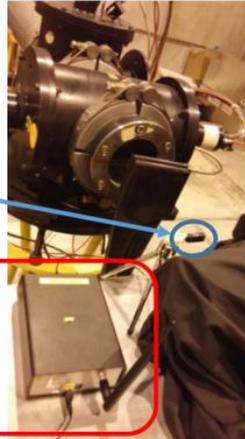
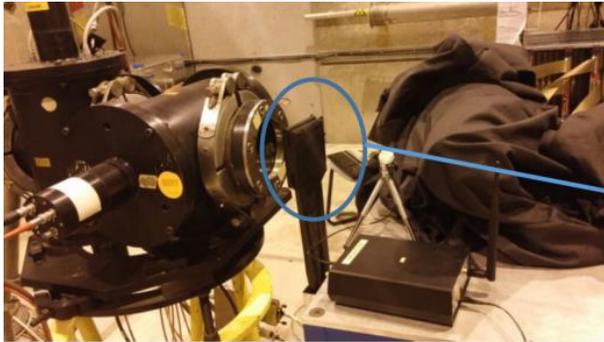
Control and readout  
Android App



- Radioguided **intraoperative beta probe**, with scintillation material coupled with SiPM detector.

# Application Example 2: Use of ArduSiPM in the CERN UA9 and CRYSBREAM activity

(substitute old Scintillator and electronics for PM)



- As beam trigger @ extracted beam line H8 (CERN)

- As beam losses counter @ SPS

# ArduSiPM è stato utilizzato per la dimostrazione della fattibilità del Brevetto n.102015000040903 “Sonda intraoperatoria”,

## Use of bremsstrahlung radiation to identify hidden weak $\beta^-$ sources: feasibility and possible use in radio-guided surgery

D. Carlotti,<sup>a,b,c</sup> F. Collamati,<sup>a</sup> R. Faccini,<sup>a,e,1</sup> P. Fresch,<sup>a</sup> F. Iacoangeli,<sup>a</sup> C. Mancini-Terracciano,<sup>a</sup> M. Marafini,<sup>a,f</sup> R. Mirabelli,<sup>a,d</sup> L. Recchia,<sup>a</sup> A. Russomando,<sup>g,h</sup> E. Soffaroli Camillocci,<sup>a,e</sup> M. Toppi,<sup>i</sup> G. Traini,<sup>a,d</sup> and V. Bocci<sup>a</sup>

<sup>a</sup>INFN Sezione di Roma, Roma, Italy  
<sup>b</sup>Scuola di Specializzazione di Fisica Medica, Sapienza Università di Roma, Roma, Italy  
<sup>c</sup>U.O.C. Fisica Sanitaria, Az. Osp. San Camillo-Forlanini, Roma, Italy  
<sup>d</sup>Dipartimento di Scienze di Base e Applicate per Ingegneria, Sapienza Università di Roma, Roma, Italy  
<sup>e</sup>Dipartimento di Fisica, Sapienza Università di Roma, Roma, Italy  
<sup>f</sup>Museo Storico della Fisica e Centro Studi e Ricerche “E. Fermi”, Roma, Italy  
<sup>g</sup>Center for Life Nano Science@Sapienza, Istituto Italiano di Tecnologia, Roma, Italy  
<sup>h</sup>Centro Científico Tecnológico de Valparaíso-CCTVal, Universidad Técnica Federico Santa María, Chile  
<sup>i</sup>Laboratori Nazionali di Frascati INFN, Frascati, Italy  
 E-mail: Riccardo.Faccini@roma1.infn.it

**ABSTRACT:** The recent interest in  $\beta^-$  radionuclides for radio-guided surgery derives from the feature of the  $\beta$  radiation to release energy in few millimeters of tissue. Such feature can be used to locate residual tumors with a probe located in its immediate vicinity, determining the resection margins with an accuracy of millimeters. The drawback of this technique is that it does not allow to identify tumors hidden in more than few mm of tissue. Conversely, the bremsstrahlung X-rays emitted by the interaction of the  $\beta^-$  radiation with the nuclei of the tissue are relatively penetrating.

To complement the  $\beta^-$  probes, we have therefore developed a detector based on cadmium telluride, an X-ray detector with a high quantum efficiency working at room temperature. We measured the secondary emission of bremsstrahlung photons in a target of Polymethylmethacrylate (PMMA) with a density similar to living tissue. The results show that this device allows to detect a 1 ml residual or lymph-node with an activity of 1 kBq hidden under a layer of 10 mm of PMMA with a 3:1 signal to noise, i.e. with a five sigma discrimination in less than 5 s.

**KEYWORDS:** Intra-operative probes; X-ray detectors

<sup>1</sup>Corresponding author.

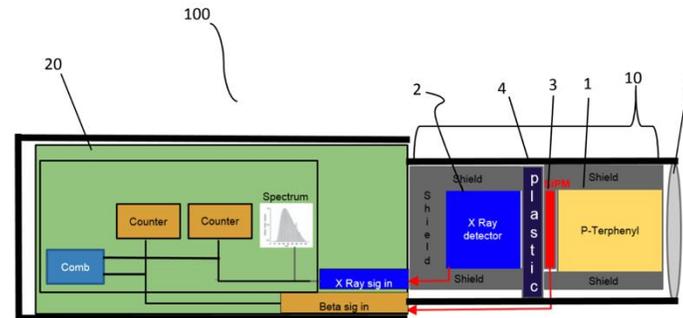


Fig. 1

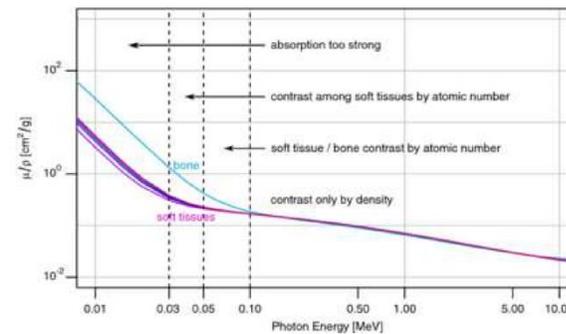


Fig. 2



### ATTESTATO DI BREVETTO PER INVENZIONE INDUSTRIALE

Il presente brevetto viene concesso per l'invenzione oggetto della domanda:

N. 102015000040903

TITOLARE/I: • ISTITUTO NAZIONALE DI FISICA NUCLEARE

Balzano Francesca  
 DOMICILIO: IP Sextant S.r.l.  
 via Antonio Salandra 18  
 00187 Roma

INVENTORE/I: • Bocci Valerio  
 • Giordano Raffaele

TITOLO: SONDA INTRAOPERATORIA

CLASSIFICA: G01T

DATA DEPOSITO: 31/07/2015

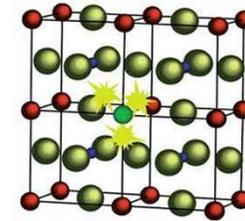
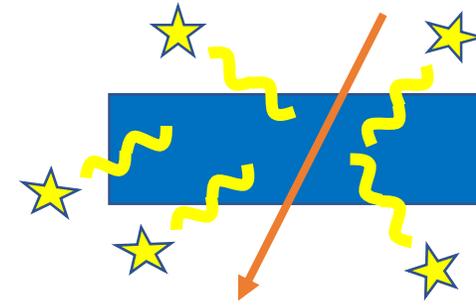
Roma, 02/01/2018

Il Dirigente della Divisione  
 Loredana Guglielmetti

# Using Scintillation materials

The use of scintillation materials is not for everyone in the past.

The only way was to use photomultiplier.

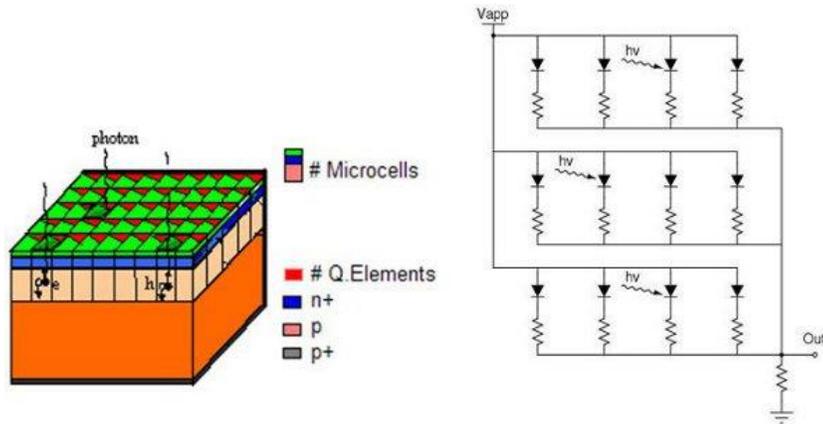
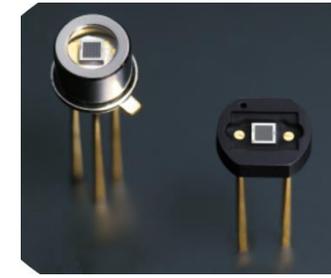


The Photomultiplier (1934).

Based on Photoelectric effect (1921 Einstein Nobel) and electron secondary emission.

The Photomultiplier are expensive and need high voltage(1000 Volt).

# SiPM (Silicon Photo Multiplier)

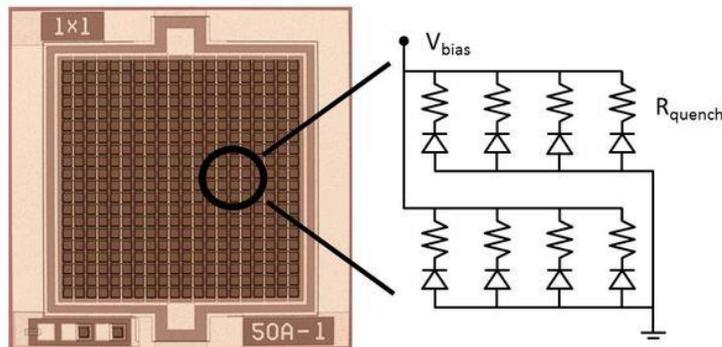


The idea behind this device is the detection of single photon events in sequentially connected SiAPDs.

The dimension of each single APD can vary from 20 to 100 micrometres, and their density can be up to 1000 per square millimeter.

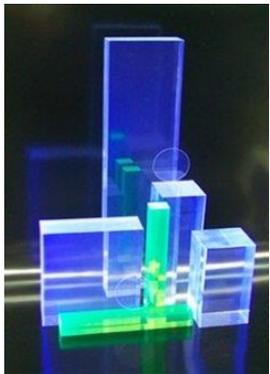
Every APD in SiPM operates in Geiger-mode and is coupled with the others by a polysilicon quenching resistor.

Although the device works in digital/switching mode, the SiPM is an analog device because all the microcells are read in parallel making it possible to generate signals within a dynamic range from a single photon to 1000 photons for just a single square millimeter area device.



The supply voltage ( $V_b$ ) depends on APD technology used, and typically varies between 20 V and 100 V, thus being from 15 to 75 times lower than the voltage required for a traditional photomultiplier tubes (PMTs) operation.

# Is it possible to build a complete particle detector and data acquisition system using Arduino microcontroller and Arduino Language ?



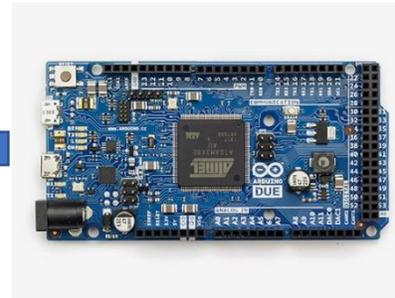
Scintillator



Photons Sensor  
(SiPM)



Custom Electronics  
(ArduSiPM Shield)



Arduino DUE



ArduSiPM Software

Particle Detector

ArduSiPM

# ArduSiPM Shield



Dr. Valerio Bocci



Giacomo Chiodi



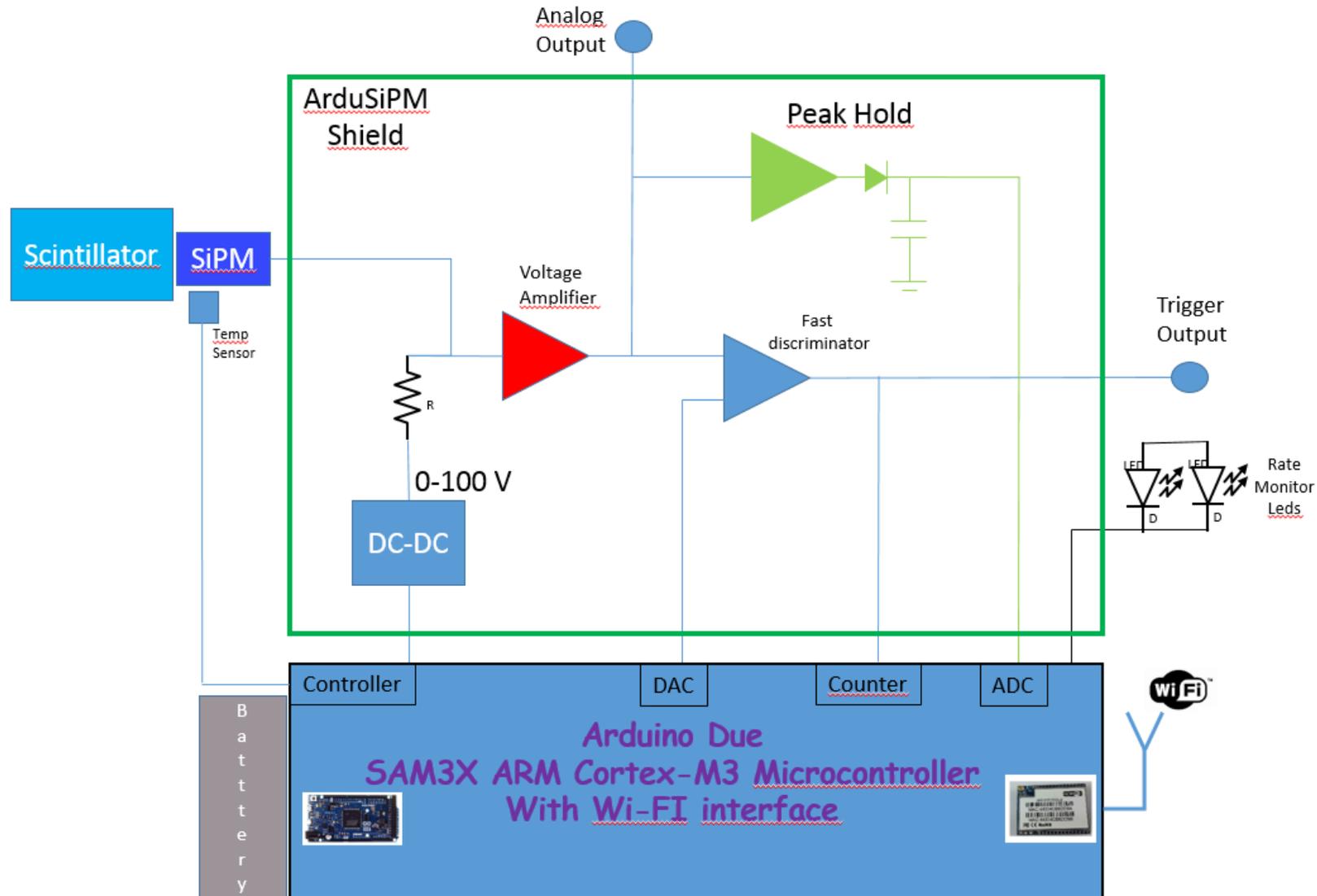
Dr. Francesco Iacoangeli



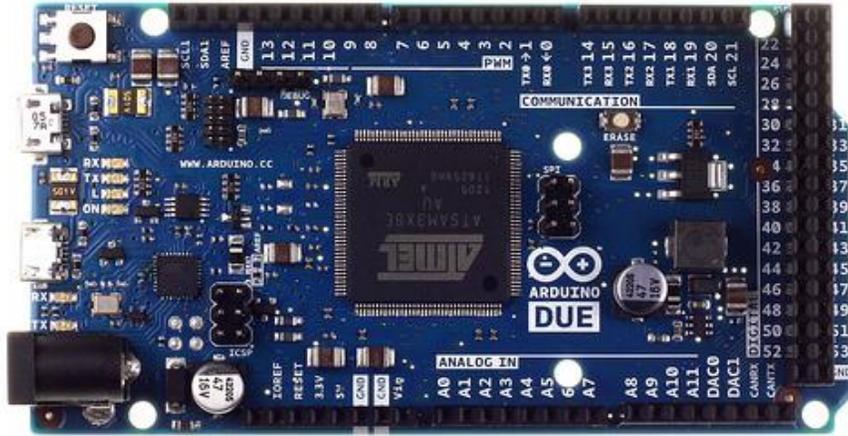
Eng. Luigi Recchia

- Shield for Arduino DUE : a custom designed piggy-back board with all the analog parts
- Same form factor of Arduino DUE
- The ArduSiPM Shield contains:
  - Fast preamplifier
  - Analog Output monitor
  - Fast discriminator with threshold and pulse output width (both digitally programmable ).
  - Trigger output
  - Peak hold circuit with programmable sample window.
  - Digitally controlled voltage bias up to 90 Volt (8 bits precision for temperature compensation)
  - Monitor of: temperature, threshold ,HV Bias

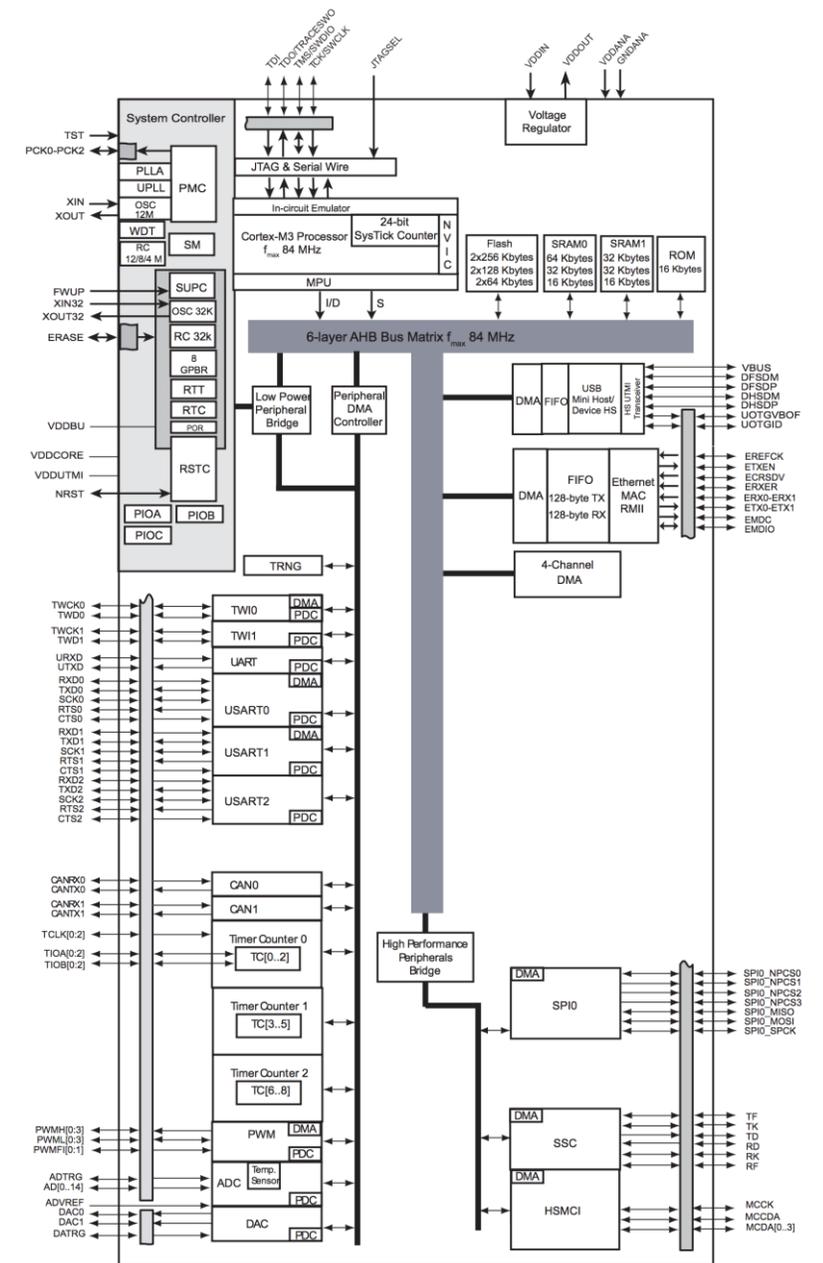
# ArduSiPM Block Diagram



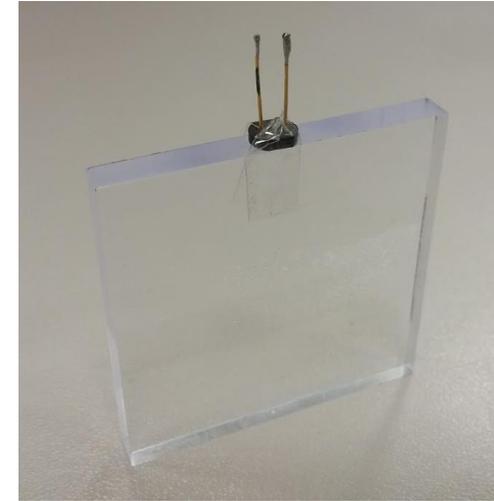
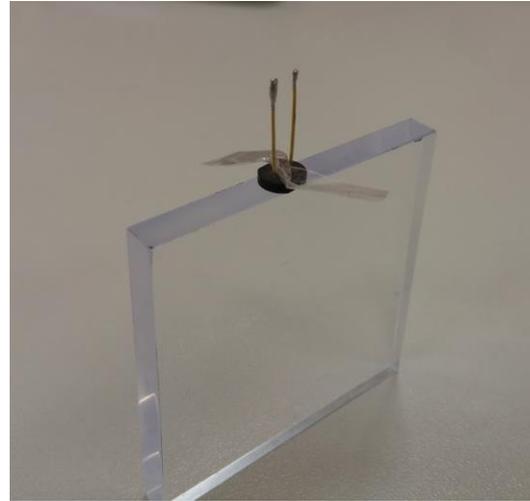
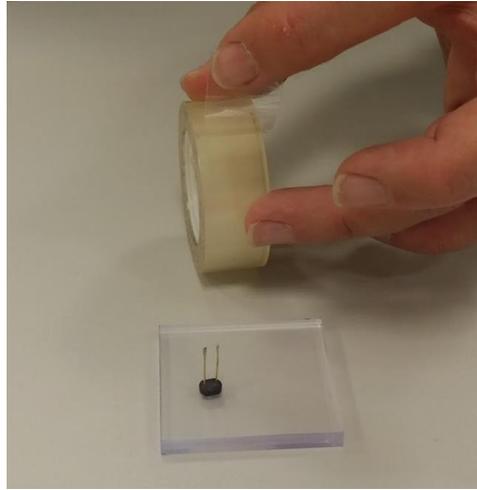
# Arduino Due



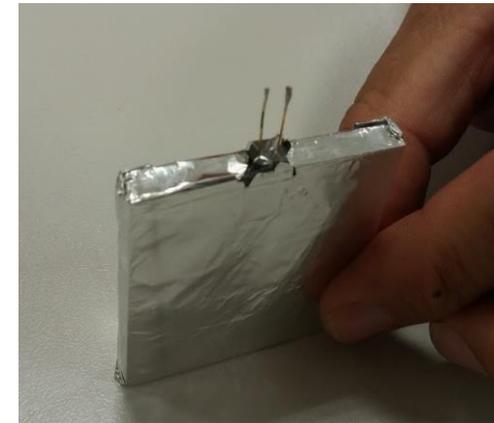
- Arduino is an open-source electronics platform based on easy-to-use hardware and software.
- Arduino Due is the first Arduino board based on SoC (System on Chip) SAM 3X8E a 32-bit ARM core microcontroller.
- Main features available on Arduino Due to build up around an acquisition system are:
  - 16 Channel Multiplexed Analog to Digital converter with 12 bit and 1 MHz sample rate
  - Multiple Input output pins
  - 9 fast Counter and pulse generator
  - 2 Digital to Analog converter with 12 bit resolution
  - Different serial interface like I2C, SPI, onewire, RS232, Ethernet MAC in SAM3X8 (not routed ☹)
  - An easy to use development software, with high level instruction for main program and interrupt handling, with the possibility to use all the complex features of the SoC SAM3X8.



# How to build a Scintillation detector with SiPM(1/2)

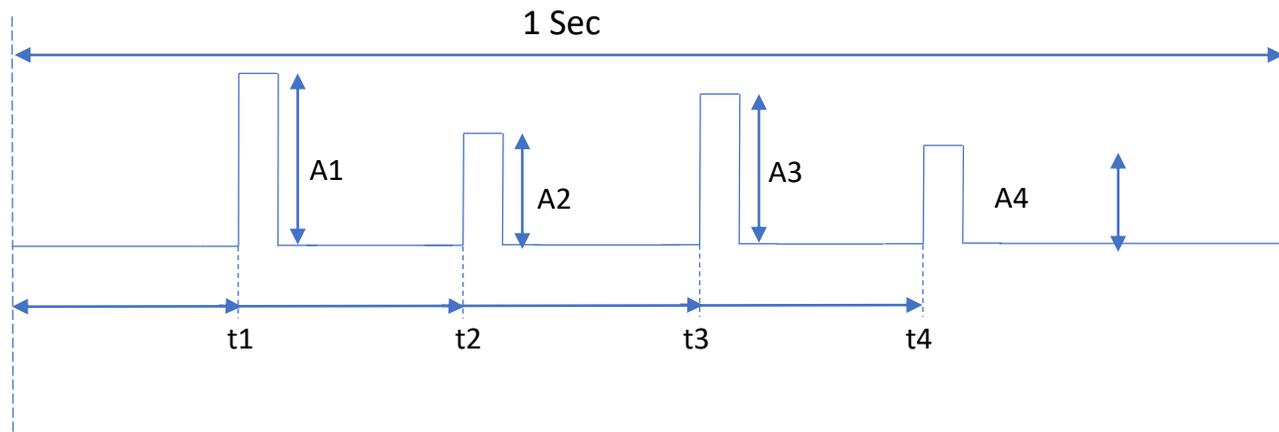


Attaching a SiPM to the scintillator with the scotch



Package with cooking aluminium foil

# ArduSiPM measurements



We split the measurements in 1 second windows, acquiring number of pulses, amplitude and time of each one.

Using a 200KBits/s serial stream

We can measure and dump (depending from amplitude and distribution of pulses):

- Only the frequency up to 40 MHz
- ADC value up to 4-6 KHz
- ADC, TDC and rate 1 -2 KHz

Using the SAM3X8 built-in ethernet it is possible to increase data acquisition performance.

## Data Stream example:

Only rate:

```
$10  
$50  
$244
```

ADC+Rate:

```
v1Fv1Dv22v27v1Dv19v20v23v20v1Cv19v1F$12  
v18v1Ev1Ev1Bv19v1Bv29v19v1Av1Dv1Bv1Dv2Av18v1B$15  
v15v20v21v21v1Dv1Fv1Av1Av1A$9  
v19v17v1Bv18v1Cv1Dv1D$7
```

TDC+ADC+RATE:

```
taedvataf0v7tv9v3$3
```

Legend:

vXXX ADC Value in HEX MSB zero suppressed

tXXXXXXXXX TDC value in HEX MSB zero suppressed

\$XXX rate in Hz

# Trasferimento Tecnologico del progetto per la produzione (non in esclusiva)

 **ROBOT DOMESTICI**  
www.robot-domestici.it

Nome utente

Utente dimenticato? Password dimenticata? Registrati

Prodotti Idea Chi siamo Contatti MEPA Supporto KIT Sviluppo progetti

Categorie [» Arduino](#) [» Arduino Kit](#)

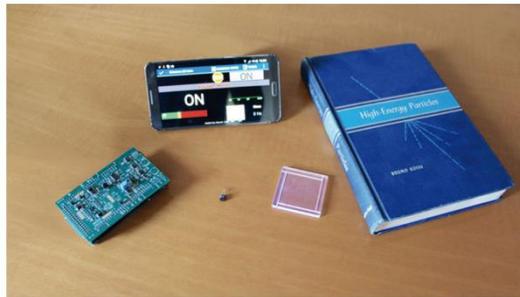
## Sistema di misura ArduSiPM Shield KIT

ArduSiPM - Un rivelatore di raggi cosmici e radiazioni nucleari

COD : RD-ArduSiPMKIT

596,83 € (excl. VAT)  
728,13 € (incl. IVA)

Disponibilità :  in arrivo in magazzino



728,13 €

Chiedi informazioni per questo prodotto (Ask a question about this product)

**PRODOTTO DISPONIBILE SUL MEPA CON IL SEGUENTE CODICE : RD-ArduSiPMKIT**

**scrivere a [commerciale@robot-domestici.it](mailto:commerciale@robot-domestici.it) per ricevere l'offerta per pacchetti multipli**

**OFFERTA PER 2 ARDUSIPM 1.149,50 € + iva**

**OFFERTA PER 4 ARDUSIPM 2.179,50 € + iva**

**OFFERTA PER 8 ARDUSIPM 3.878,50 € + iva**

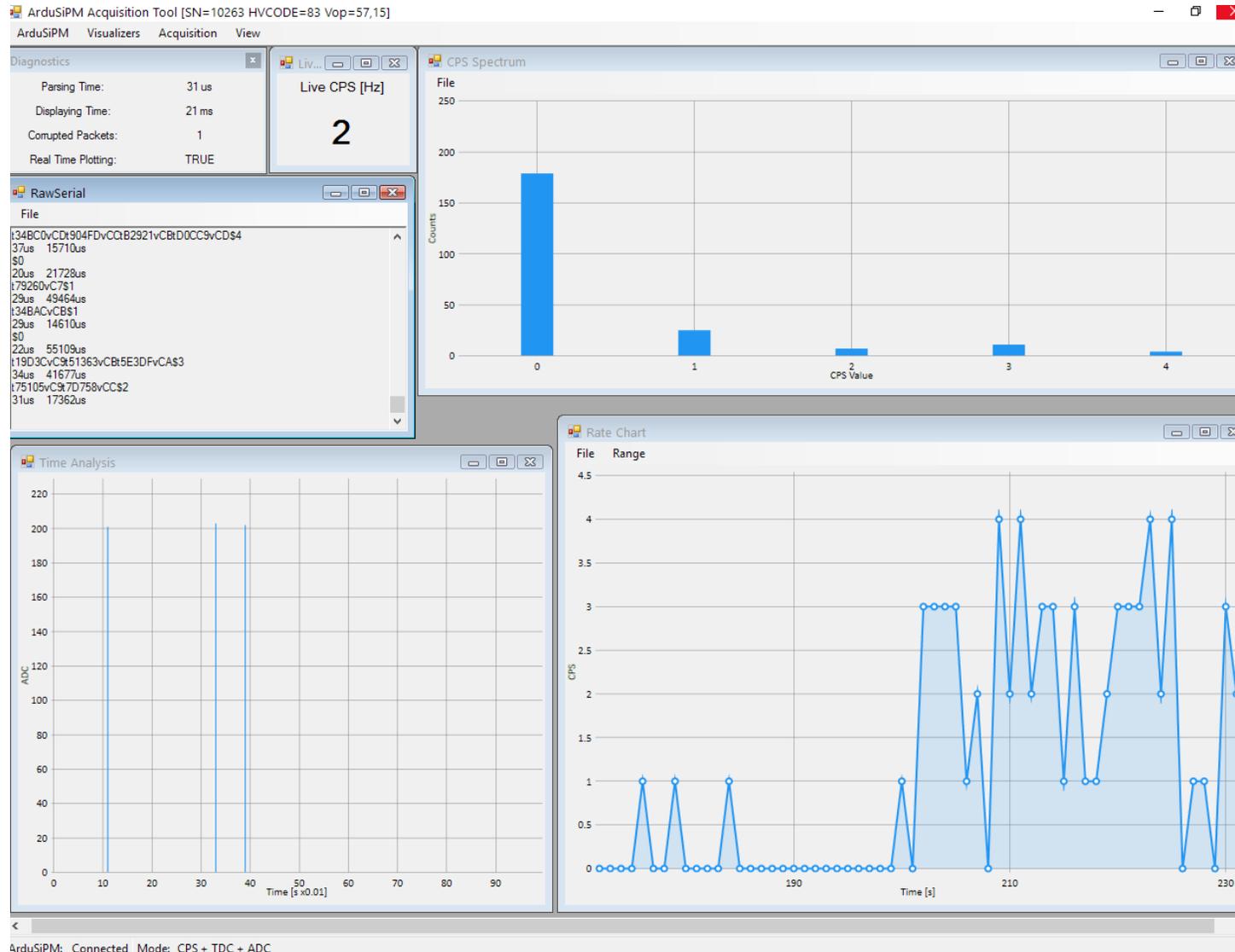
**OFFERTA PER 16 ARDUSIPM 6.786,50 € + iva**



The ArduSiPM Data format is open...Users can write custom programs for data aquisition and visualization.

Example 1. **ArduSiPM Acquisition Tool** by Filippo Curti ([Filippo.curti1@gmail.com](mailto:Filippo.curti1@gmail.com))

(written in C#, fast running also with high rate, at the moment poor in documentation freeware)



# Example 2. ArduSiPM Monitor by Ciro e Dario Chiaiese ([cirochiaiese@gmail.com](mailto:cirochiaiese@gmail.com)) Written in VisualBasic (slow good for cosmic and low rate source) well documented.

**ArduSiPM Monitor 1.2.0.2**  
realizzato da Ciro e Dario Chiaiese

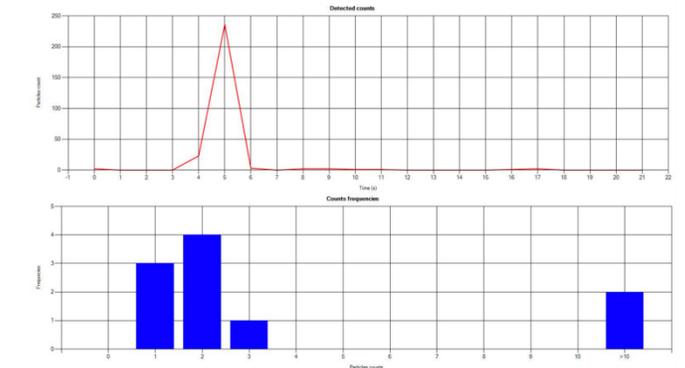
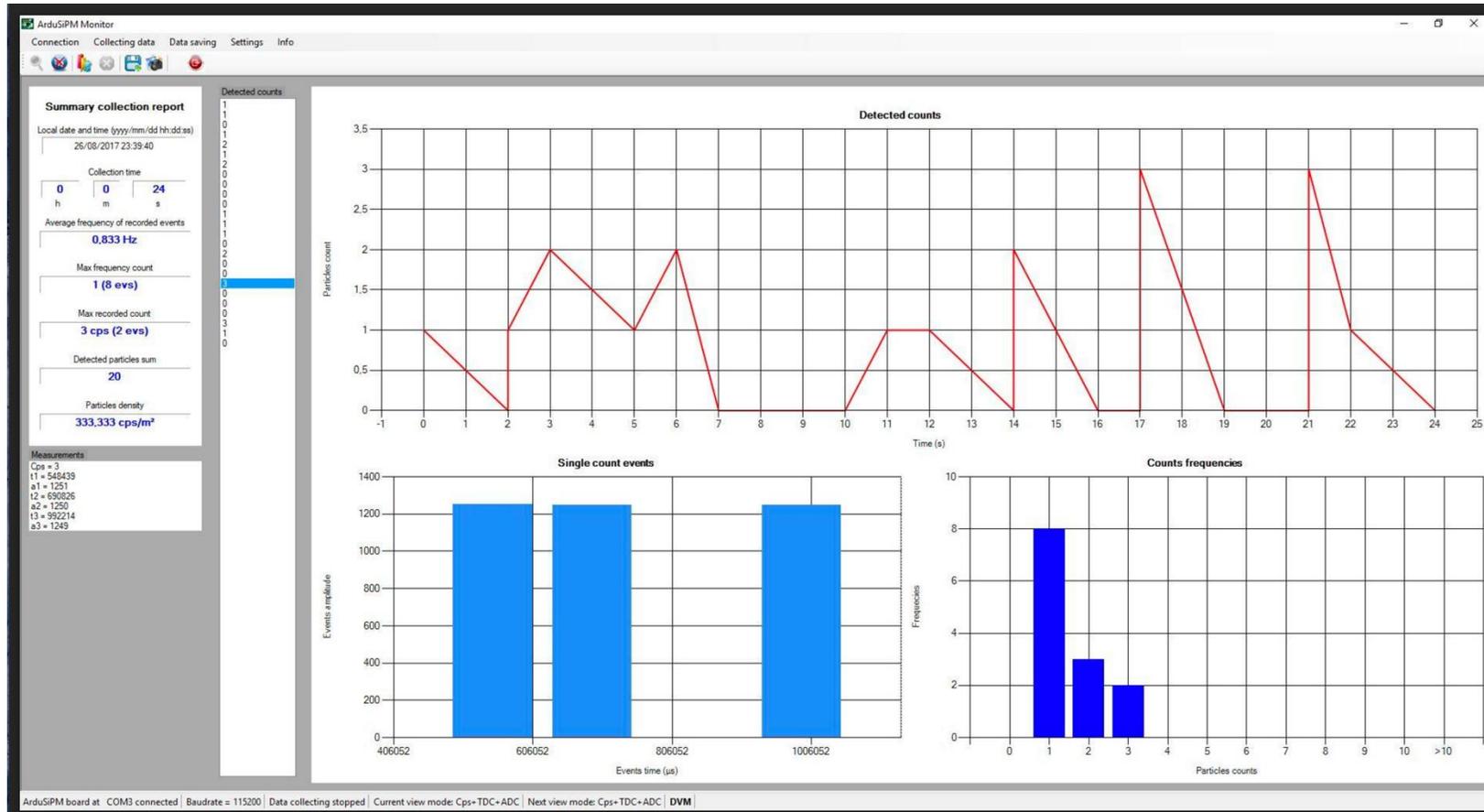
## GUIDA UTENTE

Questo software è stato realizzato per interfacciarsi con la scheda ArduSiPM (da cui il nome) equipaggiata con firmware 2.1.5. Non è stato testato sul firmware precedente e probabilmente andrebbe in errore. Si sconsiglia assolutamente di tentare di usare il software con altri tipi di schede (la qual cosa non avrebbe alcun significato, fra l'altro).

Il programma ha una funzione di riconoscimento della corretta versione di scheda (Autodetect) che è fortemente consigliata. E' comunque possibile usare il *manually detect* (🔍) per visualizzare tutti i dispositivi connessi alle USB e scegliere la porta cui è collegata ArduSiPM.

A connessione avvenuta, si può richiedere di raccogliere e graficare i dati tramite

apposito pulsante (📊) o voce di menu. La raccolta che si richiede può, precedentemente, essere stabilita in 4 diverse modalità, in seguito descritte, dal menu *Settings/Measurements*. Le modalità più ricche d'informazioni hanno una raccolta dati con frequenza più bassa (perché richiedono più tempo). La durata della raccolta può essere prima stabilita con un timer (in s) o fermata manualmente. Appena inizia la raccolta vengono visualizzati i grafici dei conteggi delle particelle rilevate per ogni secondo (Cps) e, sotto, è riportato il grafico delle frequenze dei conteggi registrati (i conteggi superiori a 10 sono accomunati in un'unica classe di frequenze ">10").



# Misura dei raggi cosmici su un aereo di linea

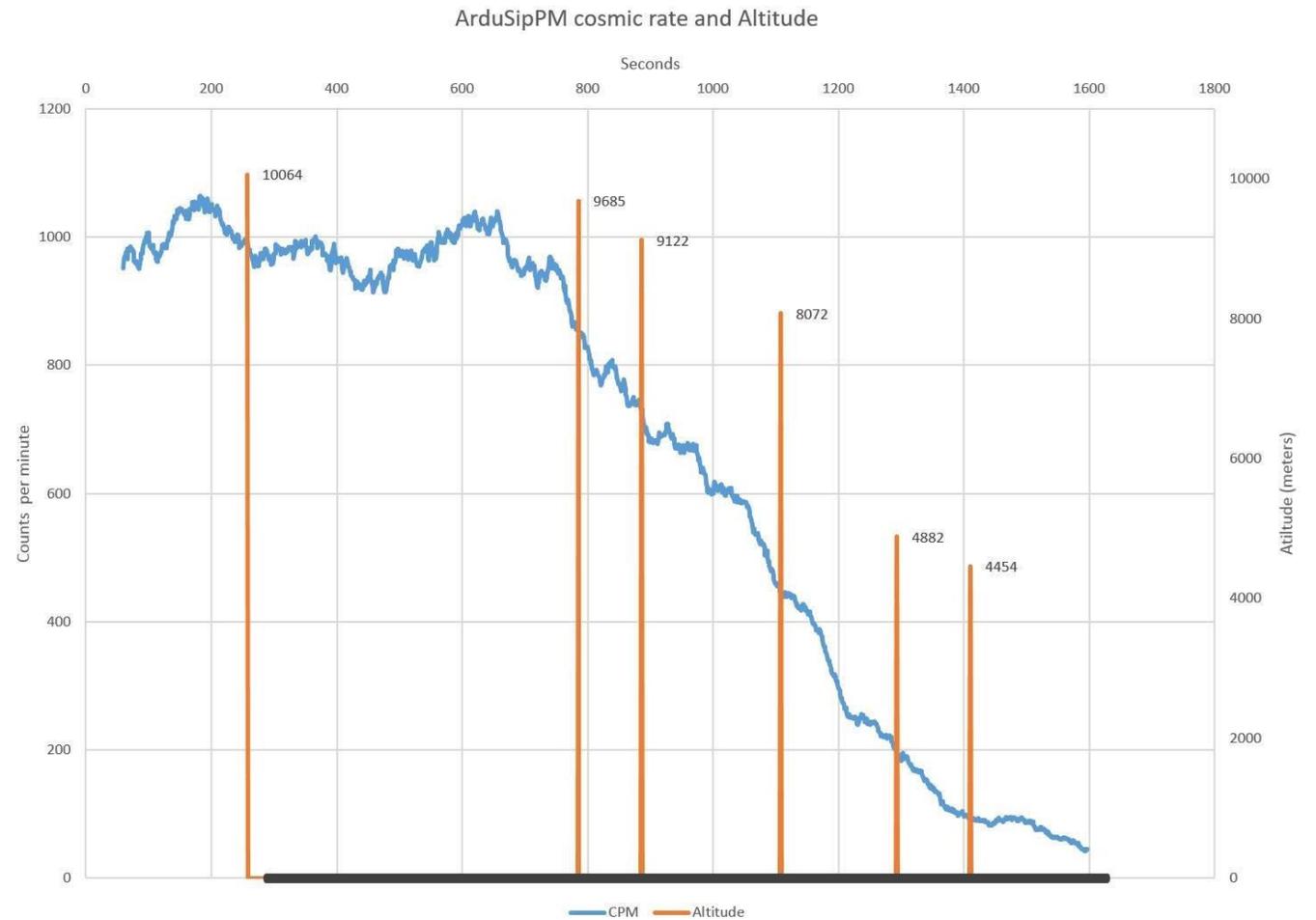
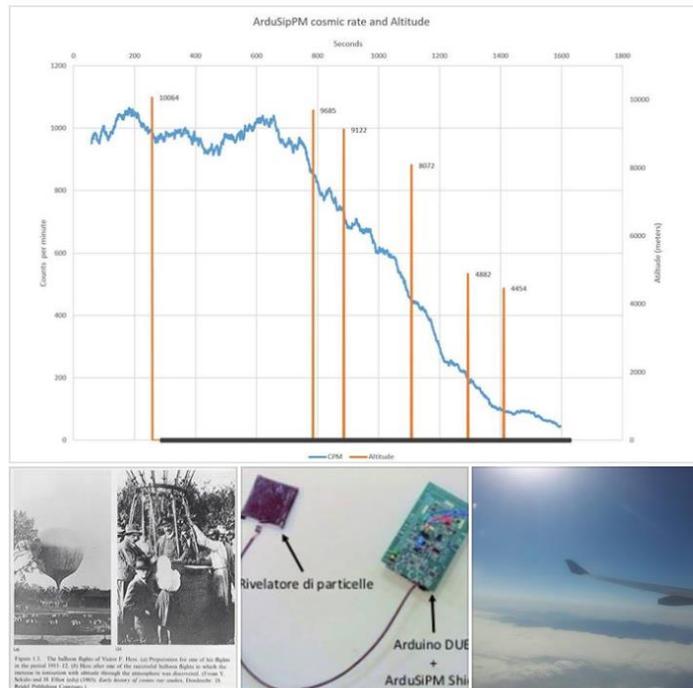


Valerio Bocci

July 6 at 12:57am

Il buon Victor Hess, nel 1911 misurò il flusso di raggi cosmici fino ad una quota di 5300 metri a bordo di un pallone aerostatico usando degli elettroscopi e una buona dose di avventurismo. Nel 1936 le sue misure gli valsero il Nobel per la fisica.

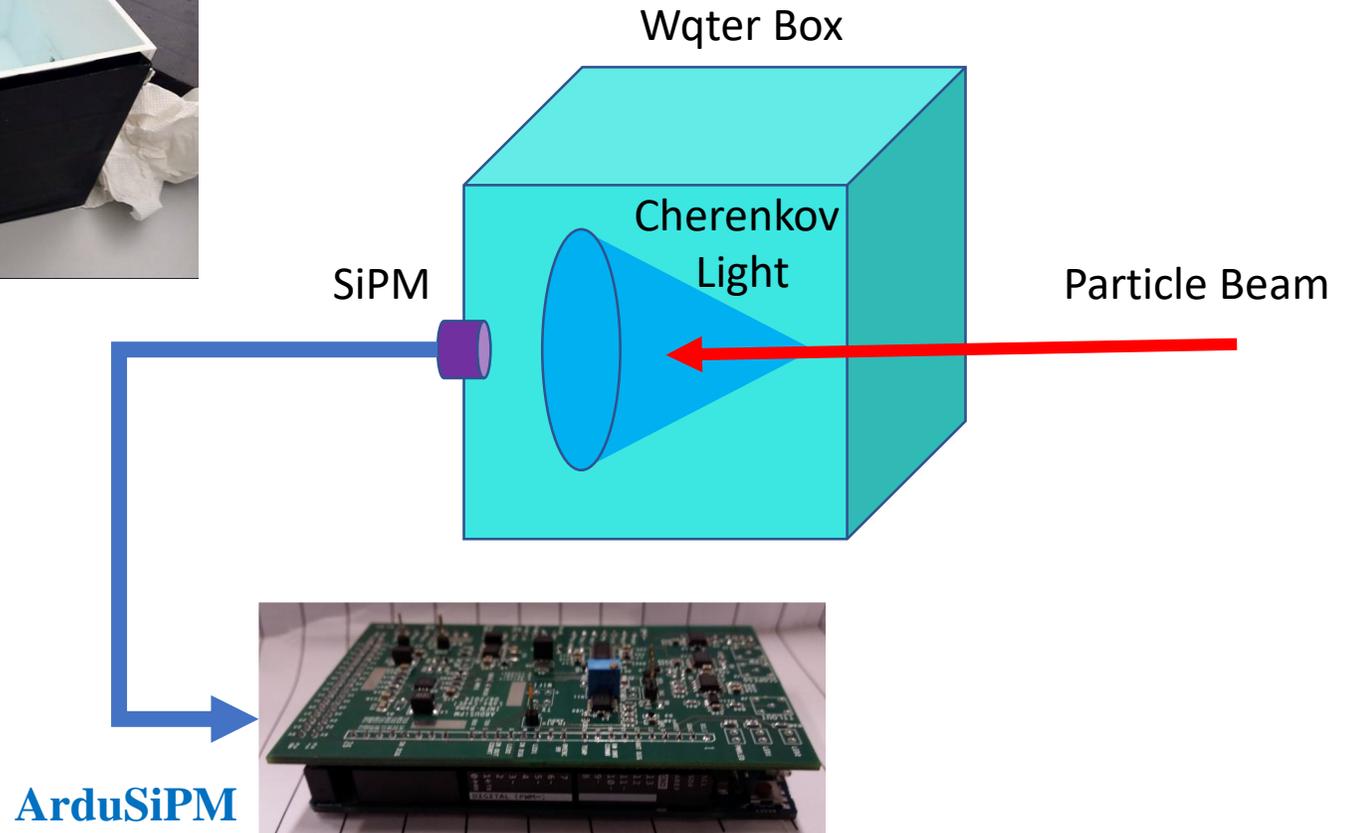
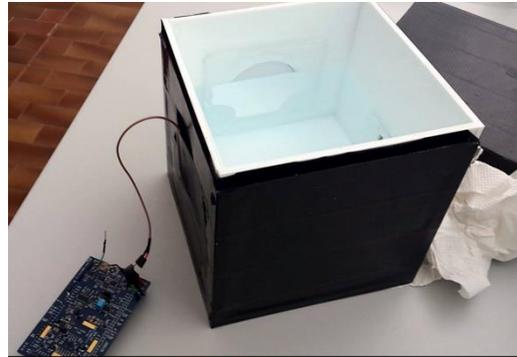
Più di cento anni dopo grazie ad un tranquillo viaggio in Aereo una misura con ArduSiPM passando da 10000 Metri e scendendo fino a circa 4000 metri giusto per il gusto di replicare un famoso esperimento.



# A School made Cherenkov light detector

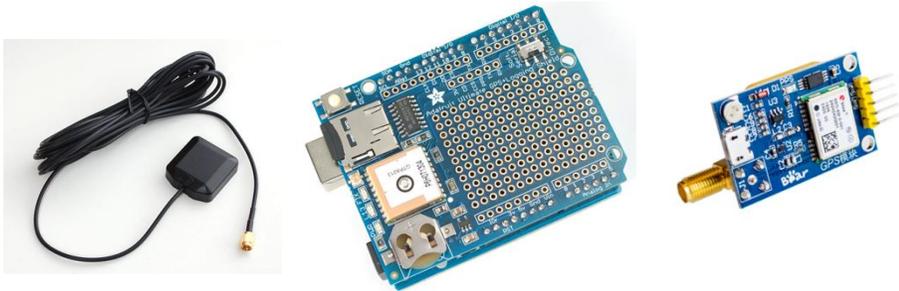
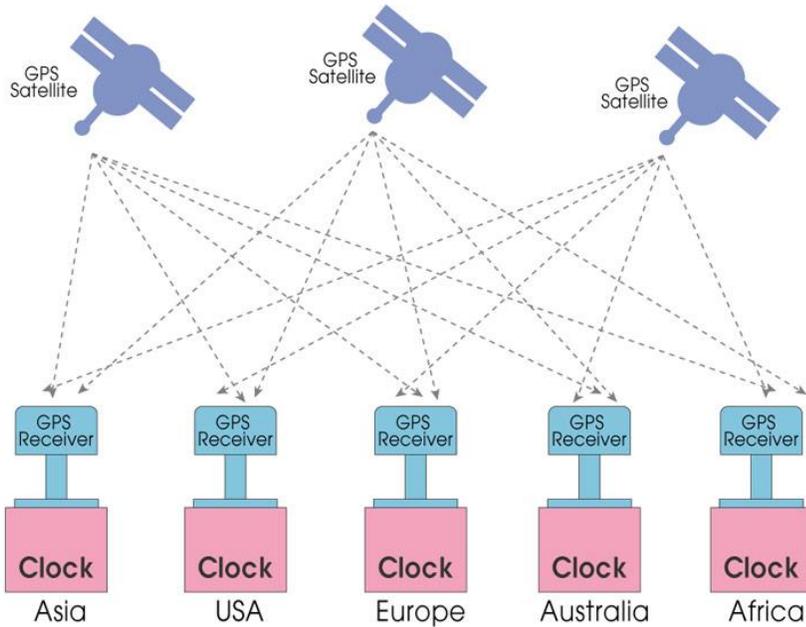
(Winner of CERN "A beamline for schools" 2017)

LICEO SCIENTIFICO STATALE T. C. ONESTI (prof Maria Rita Felici)



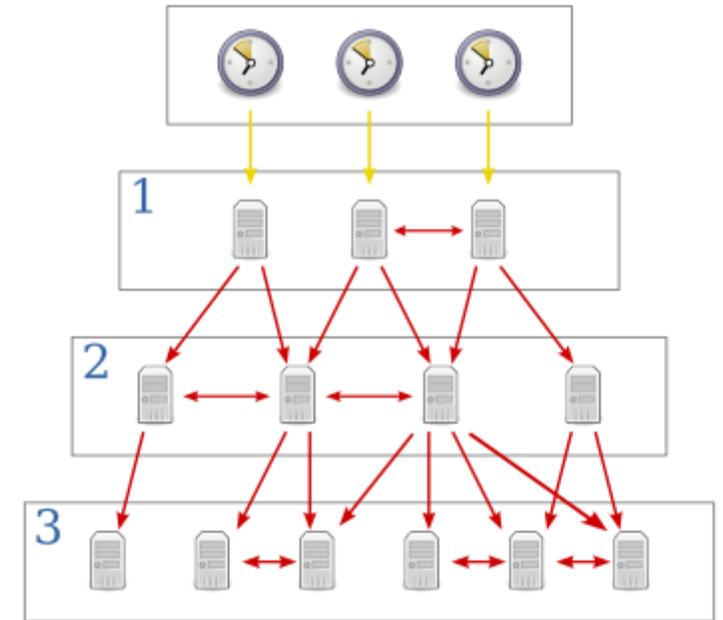
# Precise time distribution methods

GPS time (from Global Position Systems satellite)



Low cost GPS module (<30 Euro) 25 ns precision

Network Time Protocol (from internet)



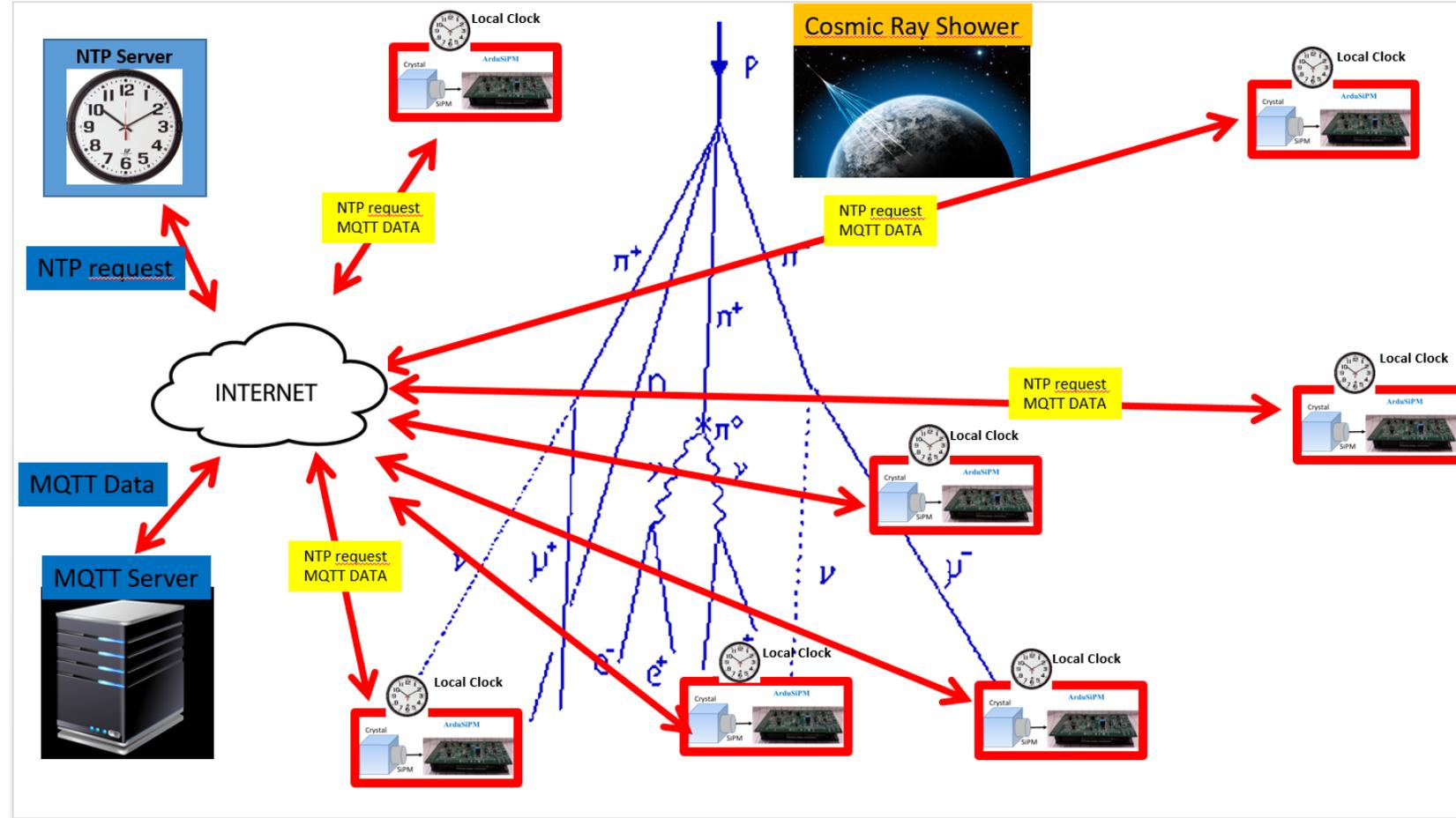
Low cost wi-fi internet processor precision <10 ms

# Search of cosmic Airshower in a wide area using ArduSiPM

Multiple ArduSiPM can be used for the research extended AirShower

The advent of microcontrollers with enough CPU power and with analog and digital peripherals give the possibility to design a complete acquisition system in one chip. The existence of an world wide data infrastructure as internet allows to think at distributed network of detectors capable to elaborate and send data or respond to settings commands.

The internet infrastructure allow us to do things unthinkable a few years ago, like to distribute the absolute time with tens of milliseconds precision to simple devices far apart from a few meters to thousands of kilometers and to create a Crowdsourcing experiment platform using simple detectors.

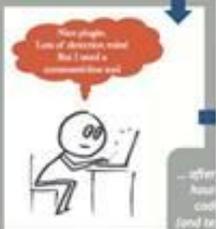




# Web Application Detection (WAD)

for asset inventory and vulnerability management

Sebastian Lopeficki / CERN Computer Security Team



# CERN Document Server

Towards the Next Generation CERN Institutional Repository

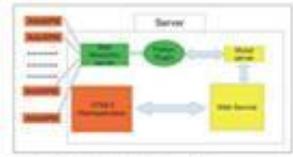
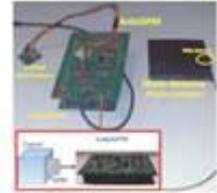
# INFN

## An educational distributed Cosmic Ray detector network based on ArduSiPM using microcontrollers as data acquisition node NTP protocol as time distribution and IoT technology for data aggregation.

Valerio Bocci, Giacomo Chiodi, Paolo Fresch, Francesco Iacangelo, Luigi Recchia

INFN Roma, Piazzale A.Moro, 2 - 00185 Roma  
valerio.bocci@roma1.infn.it

The advent of microcontrollers with enough CPU power and with analog and digital peripherals give the possibility to design a complete acquisition system in one chip. The existence of a world wide data infrastructure as internet allows to think at distributed network of detectors capable to elaborate and send data or request to settings commands. The internet infrastructure allow us to do things unthinkable a few years ago, like to distribute the absolute time with tens of milliseconds precision to single devices far apart from a few meters to thousands of kilometers and to create a Distributed experiment platform using simple detectors.



The ArduSiPM (Si) is an easy hand-held battery operated data acquisition system based with an Arduino board, which is used to detect cosmic ray and neutron radiation. The ArduSiPM uses an Arduino DUE (an open source/hardware board based on an ARM Cortex-M3 microcontroller) as processor board and a signalpath custom designed board (SiBoard), these are controlled by custom developed software and interface. The SiBoard contains different electronics features both to monitor, to set and to acquire the SiPM signal using the microcontroller board. The SiPM photon counting detector can be coupled to a cheap plastic scintillator to realize a cosmic ray detector (mainly muon particles). An ArduSiPM channel give informations about rate of events, arrival time and number of photons produced by muons. It contains all the feature from controls to data acquisition, typical of high energy Physics channel at a cost affordable for single user or school.

The terms of 'all (Internet of Things) define a set of data communication protocols and the capability of single embedded electronics objects to communicate using the internet.

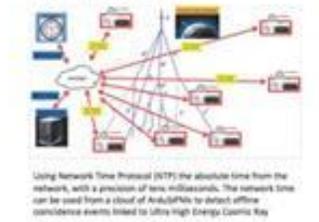
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### ESP8266 MQTT, NTP and WiFi processor



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### ArduSiPM web configuration pages.



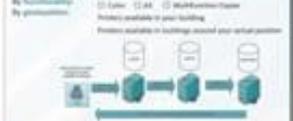
Using Network Time Protocol (NTP) the absolute time from the network, with a precision of few milliseconds. The network time can be used from a cloud of ArduSiPMs to detect offline coincidence events linked to Ultra-high Energy Cosmic Ray

# Web Services

for visitors, conference participants, anyone

worked print devices in 230 buildings a robust and user-focused service

Improving user experience



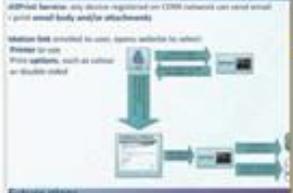
# Easy printer installation

Printer status page gives basic information including map



# Simple print solution for visitors

Without printer, any device registered on CERN network can send email



# Future plans

Reduce printer footprint, find solution for areas with multiple devices

# Users

ICHEP Chicago



# Operations



# Big Data Repositories at the Fingertips



# Shared File Spaces for All Platforms



# Data hub for analysis, shared file spaces and more

CERNBox

CONTACT: [sebastian.lopeficki@cern.ch](mailto:sebastian.lopeficki@cern.ch)

# Cloud data analysis storage backend



# Education & Research Community

Cloud Services for Synchronisation and Sharing (CS3)

Next applications, cloud storage technology, collaborations

SURFpaysa Amsterdam  
30 Jan - 1 Feb 2017

PROGRAMME COMMITTEE

# Integration Architecture



GEANT

Secure, open, frictionless. No sharing, everywhere

Federation

Multi-vendor APIs to connect on-premise clouds and sync/share solutions.



# ArduSiPM Social Media



<http://www.roma1.infn.it/conference/ArduSiPM/>



<https://www.facebook.com/groups/ardusipm/>



@ardusipm



@ArduinoSiPM



<https://groups.google.com/forum/#!forum/ardusipm>

# Attività più importanti

Attività di didattiche all'interno dell'INFN:

- Laboratorio didattico per insegnanti nell'ambito degli Incontri di Fisica IDF (LNF)
- Laboratorio didattico per studenti dell' International School on modern PhYsics and Research (LNF)

Riconoscimenti ed inviti

- Tra i 110 migliori progetti del Maker Faire 2015 (primo tra gli scientifici)
- Invito Makers Town di Bruxelles (tra i migliori 50 progetti Makers in europa) (2016)
- Invito al EU-Institutional Maker Faire at European Parliament (tra i migliori 10 progetti in europa) (2016)

**Il progetto grazie all'idea di utilizzare il canale del trasferimento tecnologico (10 % Royalty INFN o sconto 15 % per acquisto INFN) ha diffuso il rivelatore a costo zero.**

Diverse scuole hanno iniziato ad usarlo con risultati anche eccelsi :

- vincitori CERN Beam Line for school 2017 (Liceo Onesti Fermo (FM) ),
- Articolo su rivista Associazione Insegnanti Fisica "La Fisica nella Scuola" (Liceo Scientifico Cariati (CS) )
- Sperimentazione didattica (20 classi tra elementari e medie ,oltre 500 alunni)  
(Istituto Comprensivo Largo Cocconi Roma)

Le scuole di loro iniziativa e atonomamente hanno utilizzato ArduSiPM in eventi di divulgazione scientifica

- International Cosmic Ray (Liceo Scientifico Casiraghi (MI))
- A caccia di raggi cosmici sul grattacielo Pirelli (Liceo Scientifico Casiraghi Casiraghi (MI) )
- Stand al Global Junior Challenge (Liceo Scientifico Manicini Avellino)
- Utrecht Science Festival. (Amadeus Lyceum di Utrecht )

Seminari in cui si è parlato di ArduSiPM:

- 3rd Hamamatsu Photonics "Technology Days 2016 (Prof. Massimo Caccia)

## Quale futuro per ArduSiPM ?

- ArduSiPM è una novità nel panorama mondiale (2014), ma anche altri si stanno attrezzando con dispositivi simili. Esempi: Progetto MIT uscito su Symmetry magazine (2016) che ci onora citandoci ma che sempre più assomiglia ad ArduSiPM ,  
Cosmic PI al CERN(2016) prodotto e regalato dal CERN in occasioni specifiche.
- Il mercato italiano è limitato per pensare a grandi produzioni che diminuirebbero sensibilmente il costo del Kit.
- Trovare altri laboratori esteri che lo utilizzino o partecipino a progetti di ricerca comuni per nuove versioni possibili (Contatti con SUPSI per una versione spazializzata).
- Realizzazione di laboratori didattici in vari paesi (Ottimo anche per i paesi in via di sviluppo).
- Creazione di una rete di sedi INFN in grado di dare il supporto tecnico /scientifico ma anche di pensare nuovi esperimenti ed applicazioni ( contatti con INFN Pavia, Roma 2, Lecce, Milano, LNF, Roma 3, Milano Bicocca).
- Consolidare o trovare nuovi partner industriali disposti non solo a produrre ma anche ad investire e rischiare risorse su scala globale.
- Necessità di internazionalizzazione e divulgazione della sua esistenza (presentazione a conferenze scientifiche ma anche a conferenze internazionali di insegnanti)

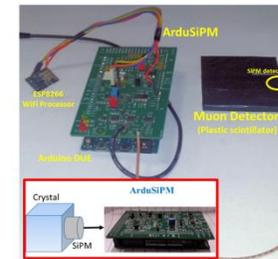
# Backup Slides

# An educational distributed Cosmic Ray detector network based on ArduSiPM using microcontrollers as data acquisition node NTP protocol as time distribution and IoT technology for data aggregation.

Valerio Bocci, Giacomo Chiodi, Paolo Fresch, Francesco Iacoangeli, Luigi Recchia

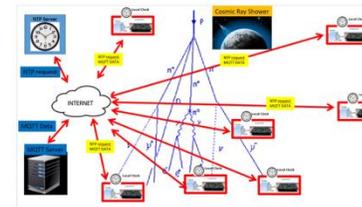
INFN Roma, Piazzale A.Moro, 2 - 00185 Roma  
valerio.bocci@roma1.infn.it

The advent of microcontrollers with enough CPU power and with analog and digital peripherals give the possibility to design a complete acquisition system in one chip. The existence of a world wide data infrastructure as internet allows to think at distributed network of detectors capable to elaborate and send data or respond to settings commands. The internet infrastructure allow us to do things unthinkable a few years ago, like to distribute the absolute time with tens of milliseconds precision to simple devices far apart from a few meters to thousands of kilometers and to create a Crowdsourcing experiment platform using simple detectors.

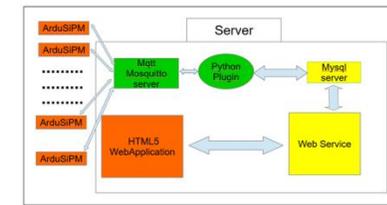


The ArduSiPM (1) is an easy hand-held battery operated data acquisition system based with an Arduino board, which is used to detect cosmic rays and nuclear radiation. The ArduSiPM uses an Arduino DUE (an open Software/Hardware board based on an ARM Cortex-M3 microcontroller) as processor board and a piggyback custom designed board (Shield), these are controlled by custom developed software and interface. The Shield contains different electronics features both to monitor, to set and to acquire the SiPM signal using the microcontroller board. The SiPM photon counting detector can be coupled to a cheap plastic scintillator to realize a cosmic ray detector (mainly muon particles). An ArduSiPM channel give informations about rate of events, arrival time and number of photons produced by muons, it contains all the feature from controls to data acquisition typical of High Energy Physics channel at a cost affordable for single user or school.

(1) The ArduSiPM: a compact transportable Software/Hardware Data Acquisition system for SiPM detector V. Bocci et al. IEEE NSSMIC 2014 Page 1 - 5 DOI: 10.1109/NSSMIC.2014.7431252 arXiv:1411.7814

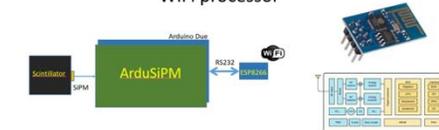


Using Network Time Protocol (NTP) the absolute time from the network, with a precision of tens milliseconds. The network time can be used from a cloud of ArduSiPMs to detect offline coincidence events linked to Ultra High Energy Cosmic Ray



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## ArduSiPM web configuration pages.



# CHEP 2016

## 22<sup>nd</sup> International Conference on Computing in High Energy and Nuclear Physics

October 10-14, 2016  
Marriott Marquis, San Francisco, CA, USA

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**http://chep2016.org**

CHEP 2016 Chairs:  
Richard Mount, SLAC and Craig Tull, LBNL

Contact:  
Program and Sponsorship: chairs@chep2016.org  
Other Issues: support@chep2016.org

The CHEP conferences address challenges in computing, networking and software for the world's leading data-intensive science experiments that currently analyze hundreds of petabytes of data using highly distributed worldwide computing resources.

CHEP 2016 will explore the increasing importance of the computing-related connections between high-energy and nuclear physics and data-intensive astronomy, X-ray science and computational science.

The CHEP conferences are held at 18-month intervals. Recent conferences took place in Okinawa, Amsterdam and New York. CHEP 2016 will take place on October 10-14, 2016 at the Marriott Marquis, San Francisco.

A WLCG Workshop will take place October 9-9 at the Marriott.

The conference is jointly organized by SLAC National Accelerator Laboratory and Lawrence Berkeley National Laboratory with logistical support from Stanford University Conference Services.

SLAC and LBNL offer research facilities at the forefront of data-intensive experimental and computational science.

Participation by students, young scientists and minorities is strongly encouraged. The organizers expect to be able to offer partial financial support to a limited number of such participants.

# The ArduSiPM in the World



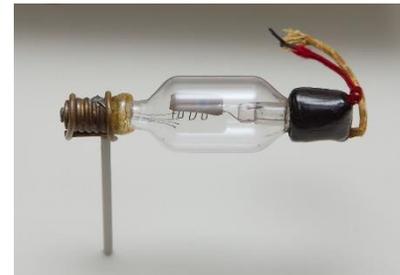
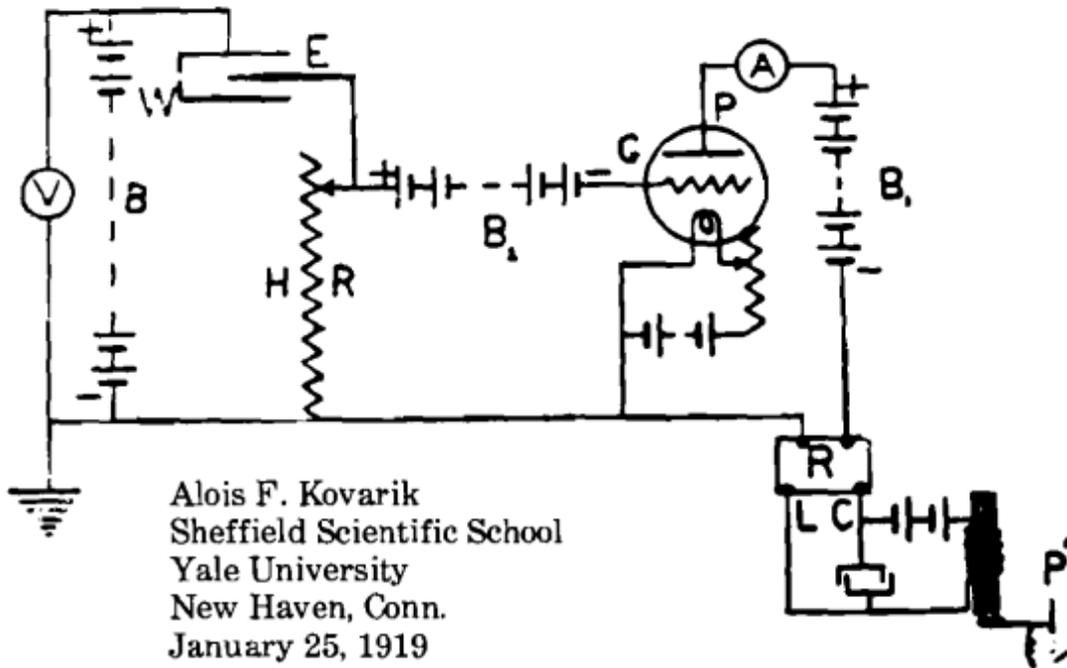
# How to build a Scintillation detector with SiPM(2/2)



Using a black tape to avoid external lighth.

# First Electronic particle detector 1919

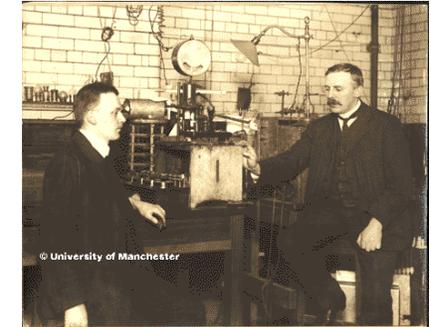
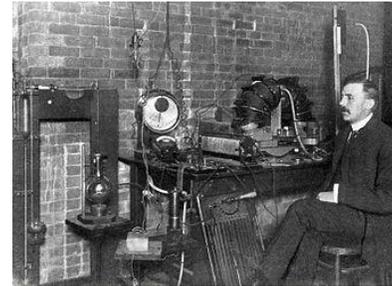
## ON THE AUTOMATIC REGISTRATION OF $\alpha$ -PARTICLES, $\beta$ -PARTICLES AND $\gamma$ -RAY AND X-RAY PULSES



Lee De Forest Audion tube from 1908, the first triode. its ability to amplify was recognized around 1912.

# The Geiger-Muller: A '900 detector.

- Robust Technology 100 years old
- Economical
- Easy to find
- There are some Makers project
- The detector is preassembled from the factory

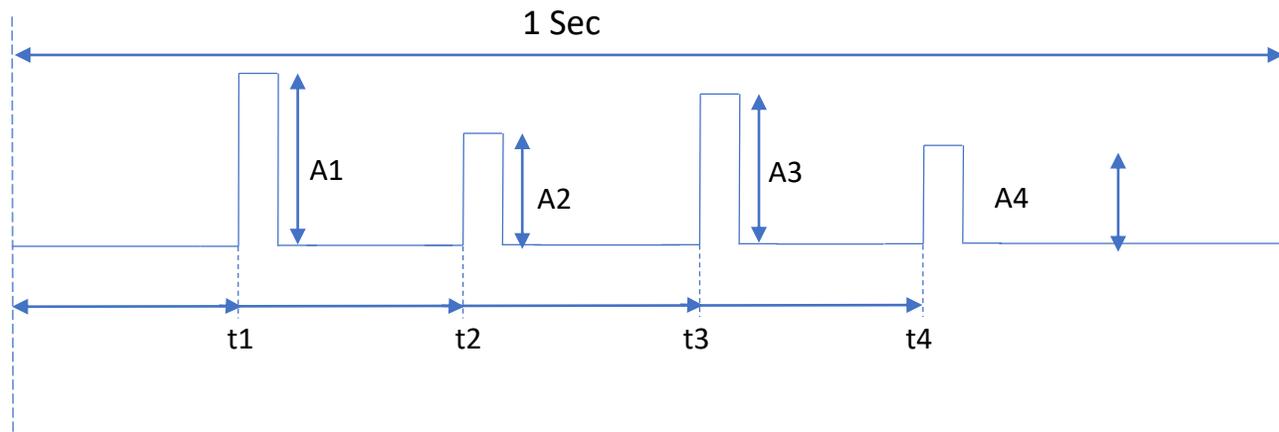


The discovery of atomic nuclei. Rutherford Hans Geiger and Ernest Marsden

- High voltage discharge (need robust electronics)
- Low efficiency.
- Yes or No detector
- Fragile.



# ArduSiPM measurements



We split the measurements in 1 second windows, acquiring number of pulses, amplitude and time of each one.

Using a 200KBits/s serial stream

We can measure and dump (depending from amplitude and distribution of pulses):

- Only the frequency up to 40 MHz
- ADC value up to 4-6 KHz
- ADC, TDC and rate 1 -2 KHz

Using the SAM3X8 built-in ethernet it is possible to increase data acquisition performance.

## Data Stream example:

### Only rate:

```
$10  
$50  
$244
```

### ADC+Rate:

```
v1Fv1Dv22v27v1Dv19v20v23v20v1Cv19v1F$12  
v18v1Ev1Ev1Bv19v1Bv29v19v1Av1Dv1Bv1Dv2Av18v1B$15  
v15v20v21v21v1Dv1Fv1Av1Av1A$9  
v19v17v1Bv18v1Cv1Dv1D$7
```

### TDC+ADC+RATE:

```
taedvataf0v7tv9v3$3
```

### Legend:

vXXX ADC Value in HEX MSB zero suppressed

tXXXXXXXXX TDC value in HEX MSB zero suppressed

\$XXX rate in Hz

# What time is it ?



$$\text{UTC}(2017)=\text{TAI}+37 \text{ sec}$$

$$\text{GPS}=\text{TAI}+19 \text{ sec}$$

## International Atomic Time

From Wikipedia, the free encyclopedia

**International Atomic Time** (**TAI**, from the French name *Temps Atomique International*<sup>[1]</sup>) is a high-precision atomic [coordinate time standard](#) based on the notional passage of [proper time](#) on [Earth's geoid](#).<sup>[2]</sup> It is the basis for [Coordinated Universal Time](#) (UTC), which is used for civil timekeeping all over the Earth's surface, and for [Terrestrial Time](#), which is used for astronomical calculations. As of 31 December 2016 when another [leap second](#) was added,<sup>[3]</sup> TAI is exactly 37 seconds ahead of UTC. The 37 seconds results from the initial difference of 10 seconds at the start of 1972, plus 27 leap seconds in UTC since 1972.

$$\text{UTC}(1958)=\text{TAI} \text{ Time}$$

## Coordinated Universal Time

From Wikipedia, the free encyclopedia

**Coordinated Universal Time** (*French*: *Temps universel coordonné*), abbreviated to **UTC**, is the primary [time standard](#) by which the world regulates clocks and time. It is within about 1 second of [mean solar time](#) at 0° longitude;<sup>[1]</sup> it does not observe [daylight saving time](#). It is one of several closely related successors to [Greenwich Mean Time](#) (GMT). For most purposes, UTC is considered interchangeable with GMT, but GMT is no longer precisely defined by the scientific community.

The first Coordinated Universal Time was informally adopted on 1 January 1960.<sup>[2]</sup>

The system was adjusted several times, including a brief period where time coordination radio signals broadcast both UTC and "Stepped Atomic Time (SAT)" until a new UTC was adopted in 1970 and implemented in 1972.<sup>[2]</sup> This change also adopted [leap seconds](#) to simplify future adjustments.

## Global Positioning System

From Wikipedia, the free encyclopedia

[Universal Time](#) (UTC), the atomic clocks on the satellites are set to GPS time (GPST; see the page of [United States Naval Observatory](#)). The difference is that GPS time is not corrected to match the rotation of the Earth, so it does not contain [leap seconds](#) or other corrections that are periodically added to UTC. GPS time was set to match UTC in 1980, but has since diverged. The lack of corrections means that GPS time remains at a constant offset with [International Atomic Time](#) (TAI) (TAI – GPS = 19 seconds). Periodic corrections are performed to the on-board clocks to keep them synchronized with ground clocks.<sup>[126]</sup>

The GPS navigation message includes the difference between GPS time and UTC. As of January 2017, GPS time is 18 seconds ahead of UTC because of the leap second added to UTC on December 31, 2016.<sup>[127]</sup> Receivers subtract this offset from GPS time to calculate UTC and specific timezone values. New GPS units may not show the correct UTC time until after receiving the UTC offset message. The GPS-UTC offset field can accommodate 255 leap seconds (eight bits).

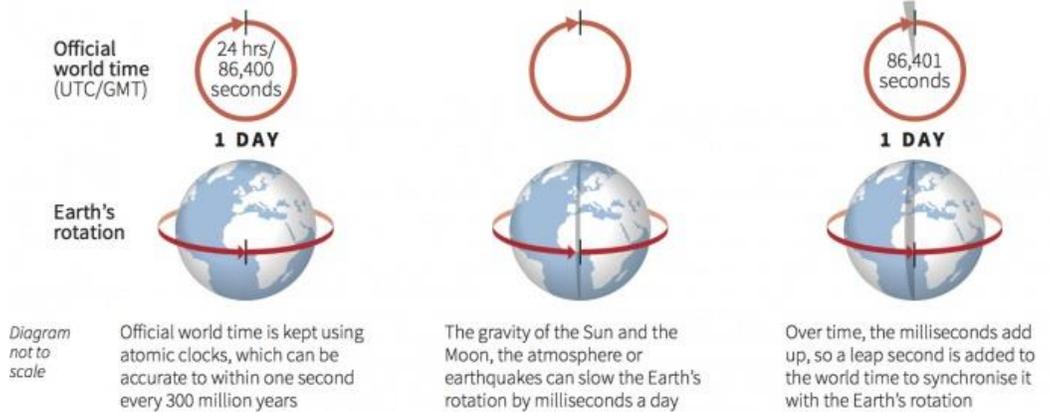
$$\text{UTC}(1980)=\text{GPS Time}$$

# Leap Second



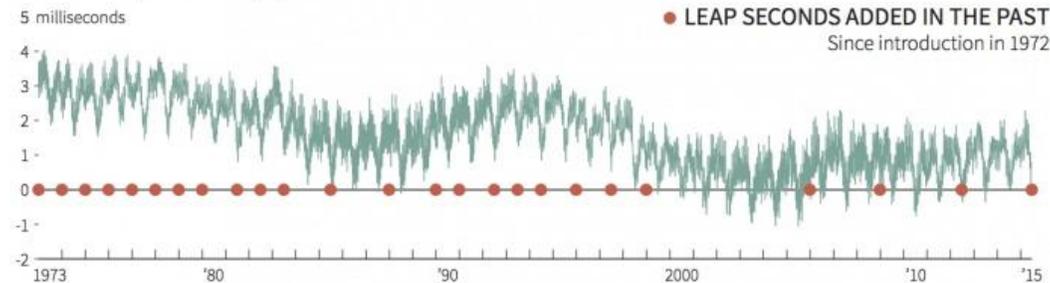
## Just a second

A leap second will be added on June 30, 2015, to allow the Earth's rotation to synchronise with atomic clocks.



## VARIATIONS IN THE LENGTH OF A DAY

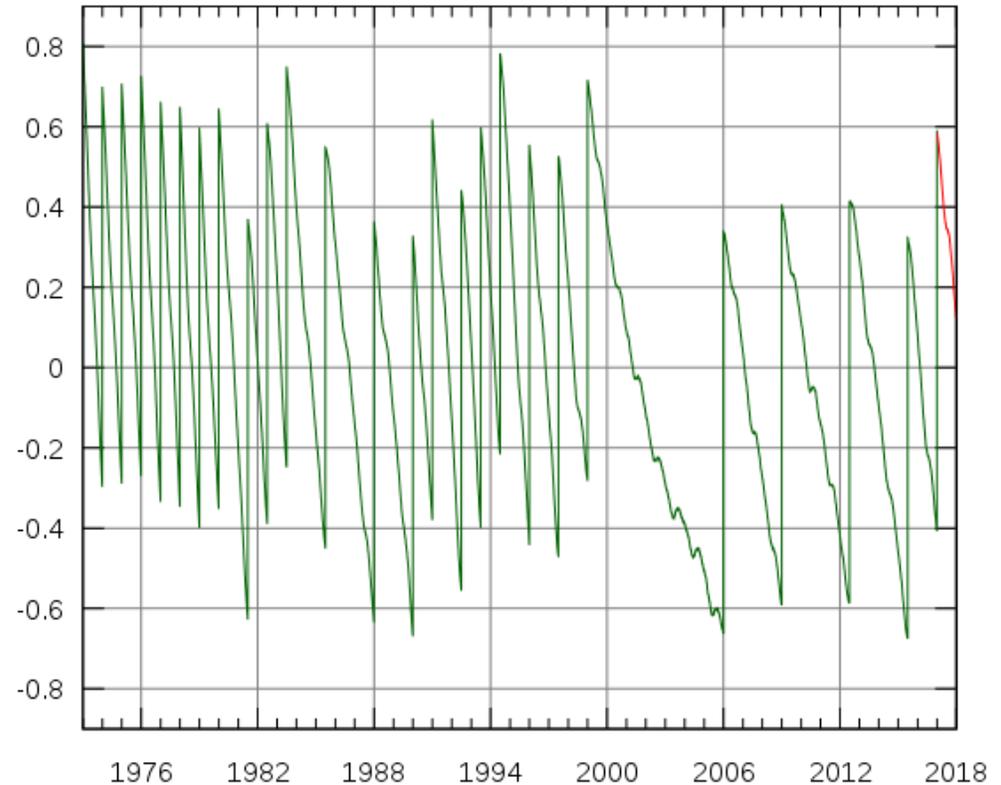
The speed of the Earth's rotation is not constant, which in turn affects the length of each day by a few milliseconds.



Sources: Space Geodesy Project, NASA; National Institute of Standards and Technology (NIST), U.S. Dept. of Commerce; International Earth Rotation and Reference Systems Service

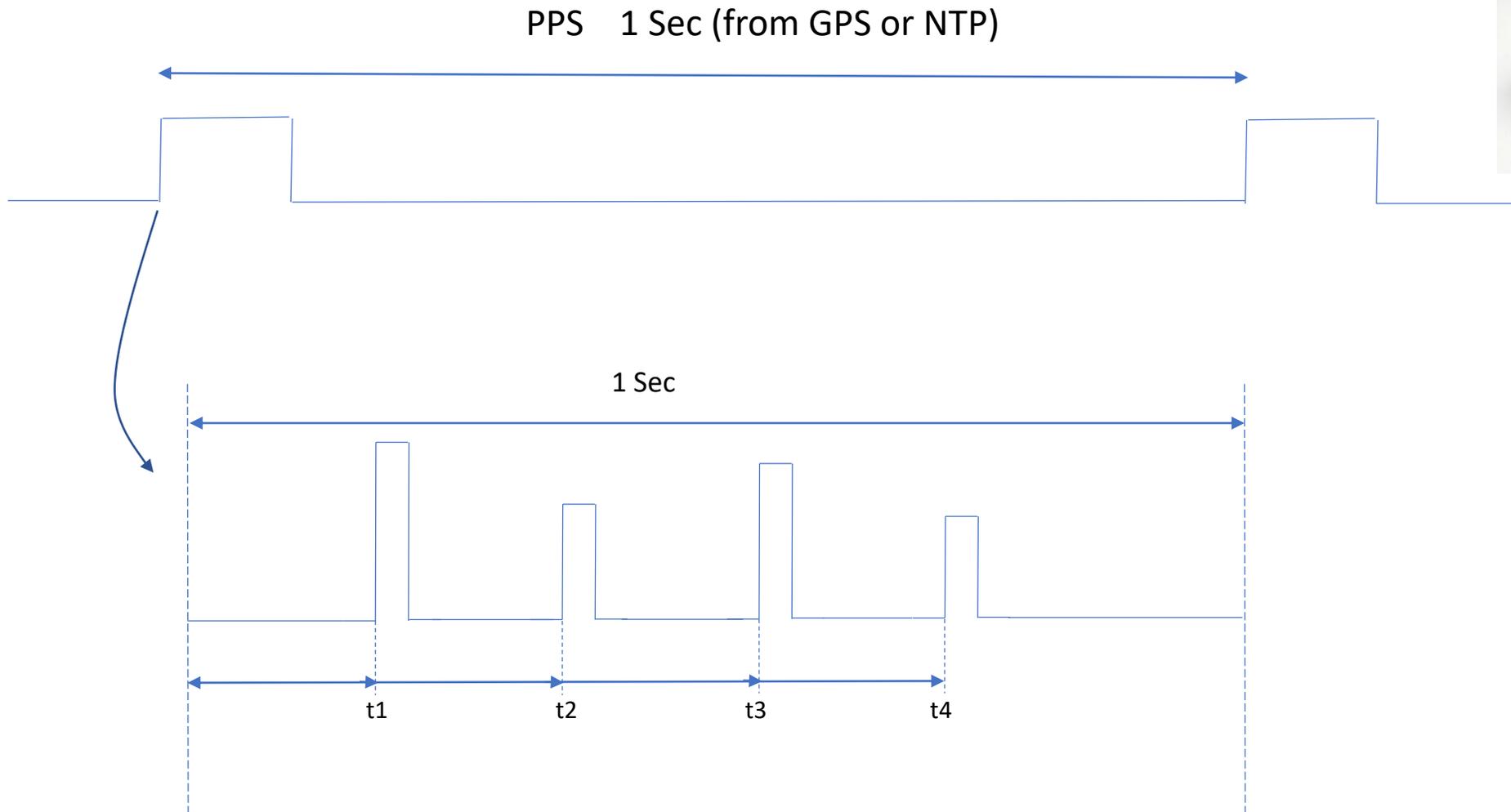
1 millisecond = 0.001 second

REUTERS



Graph showing the difference between UT1 and UTC. Vertical segments correspond to leap seconds.

# Synchronization with ArduSiPM

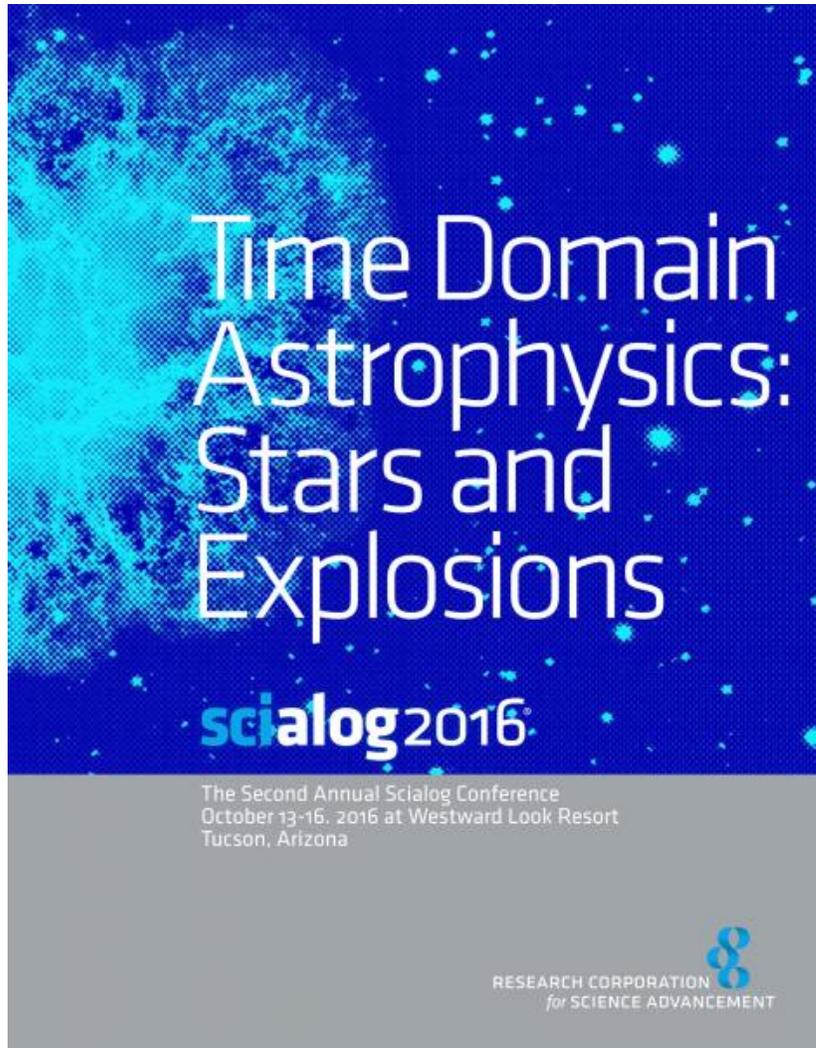


Nuove idee di utilizzo di ArduSiPM

Pronti per raccogliere nuove idee e progetti...

# Is it possible to use ArduSIPM in Time domain Astrophysics ?

Photons counting capability with timing resolution in the order of hundreds of nanoseconds

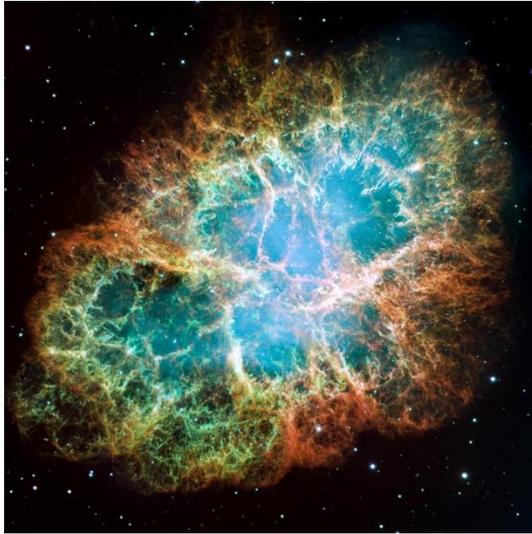


## A few thoughts about time-domain astronomy

- flares (minutes – hours)
- flickering (msec – minutes)
- outbursts (days – months)
- eruptions (months – years)
- explosions (months – years)
- eclipses / transits (minutes – hours)
- motion (minutes – years)

- stellar activity
- stellar structure & evolution
- binary evolution
- exo-planets
- accretion discs/processes
- cosmology

# Optical pulsar Real time photon detection



Fast astronomical photometers based on new technology detectors are presently used by a limited number of research groups: for instance the OPTIMA team of the Max Planck Institute (Kanbak 2003), the AQUEYE/IQUEYE team of the Padova University (Barbieri et al. 2009, Naletto et al. 2009), the S-Cam device (Oosterbroek et al. 2008), the Ultracam device (Dhillon 2008), the GASP device (Collins et al. 2008) .

ACTA POLYTECHNICA VOL. 51 No. 6/2011

## A New Fast Silicon Photomultiplier Photometer

F. Meddi, F. Ambrosino, C. Rossi, R. Nesci, S. Sclavi,  
A. Ruggieri, S. Sestito, I. Bruni, R. Gualandi

### Abstract

The Crab pulsar is one of the most intensively studied X-ray/optical objects, but up to now only a small number of research groups have based their photometers on SiPM technology. In early February 2011, the Crab pulsar signal was observed with our photometer prototype. With low-cost instrumentation, the results of the analysis are very significant: the processed data acquired on the Crab pulsar gave both a good light curve and a good power spectrum, in comparison with the data analysis results of other more expensive photometer instrumentation.

**Keywords:** Silicon PhotoMultiplier detector (SiPM), photometer, fast variability, Pulsar.

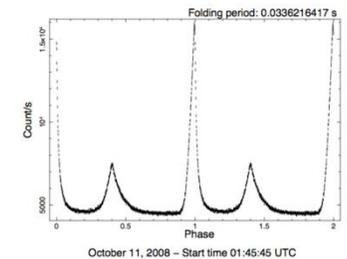
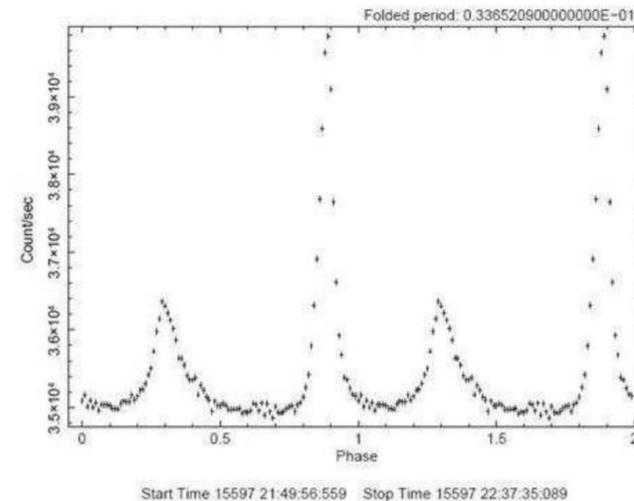
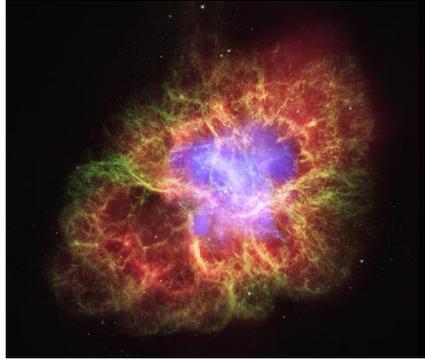


Figure 3.14: Folded light curve over the average period of the Crab pulsar for the Aquaro observation 4 in Table 2.6. Two rotations of the neutron star are shown (Germani et al. 2011).

# Raw design of ArduSiPM as fast photometer

Magnitudo	Telescope diameter cm	telescope Area cm2		
17	50	7853.981634		
	photons/cm2/s	Photons telescope/s	every us	every 5 ns
	1584.391685	12443783.2	12.4437832	0.06221891598



Optical Pulsar  
Crab Pulsar  
(CM Tauri, PSR B0531+21)  
M17

