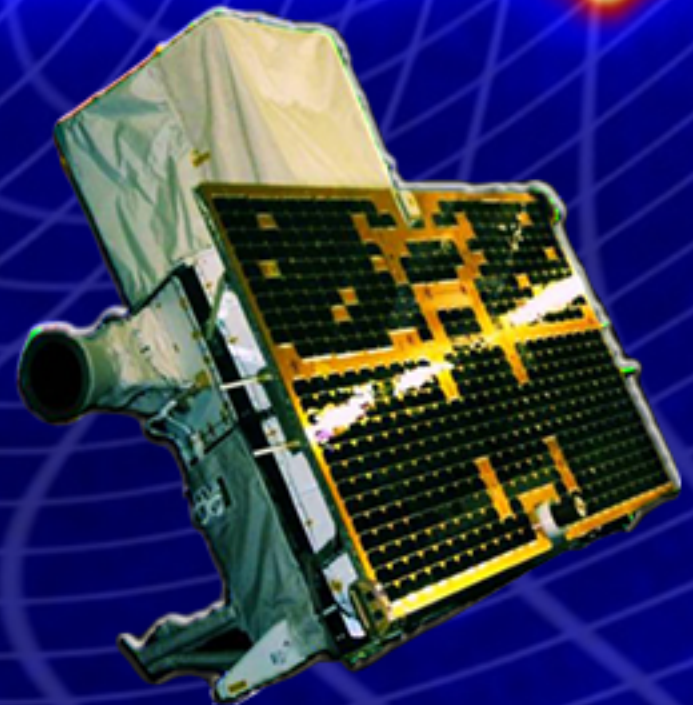


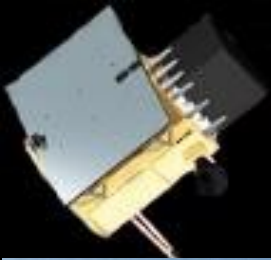
# Multi messenger astronomy with the $\gamma$ -ray satellite AGILE: gravitational wave events and ultra high energy astrophysical neutrinos

**Paolo W. Cattaneo, INFN Pavia on behalf of the AGILE Team**

CRIS 2018, Porto Palo di Capo Passero, June 18, 2018



Prepared with help from C.Pittori, F.Verrecchia and F.Lucarelli



# AGILE mission

*Science Data Center*



**Launched from Sriharikota  
(India) on 23th April 2007.**

**Orbit at ~500 km**

**Period ~90 minutes**

**Inclined 2.5° on Equator**

**80% of sky every 7 minutes**

**100-150 passes each region**

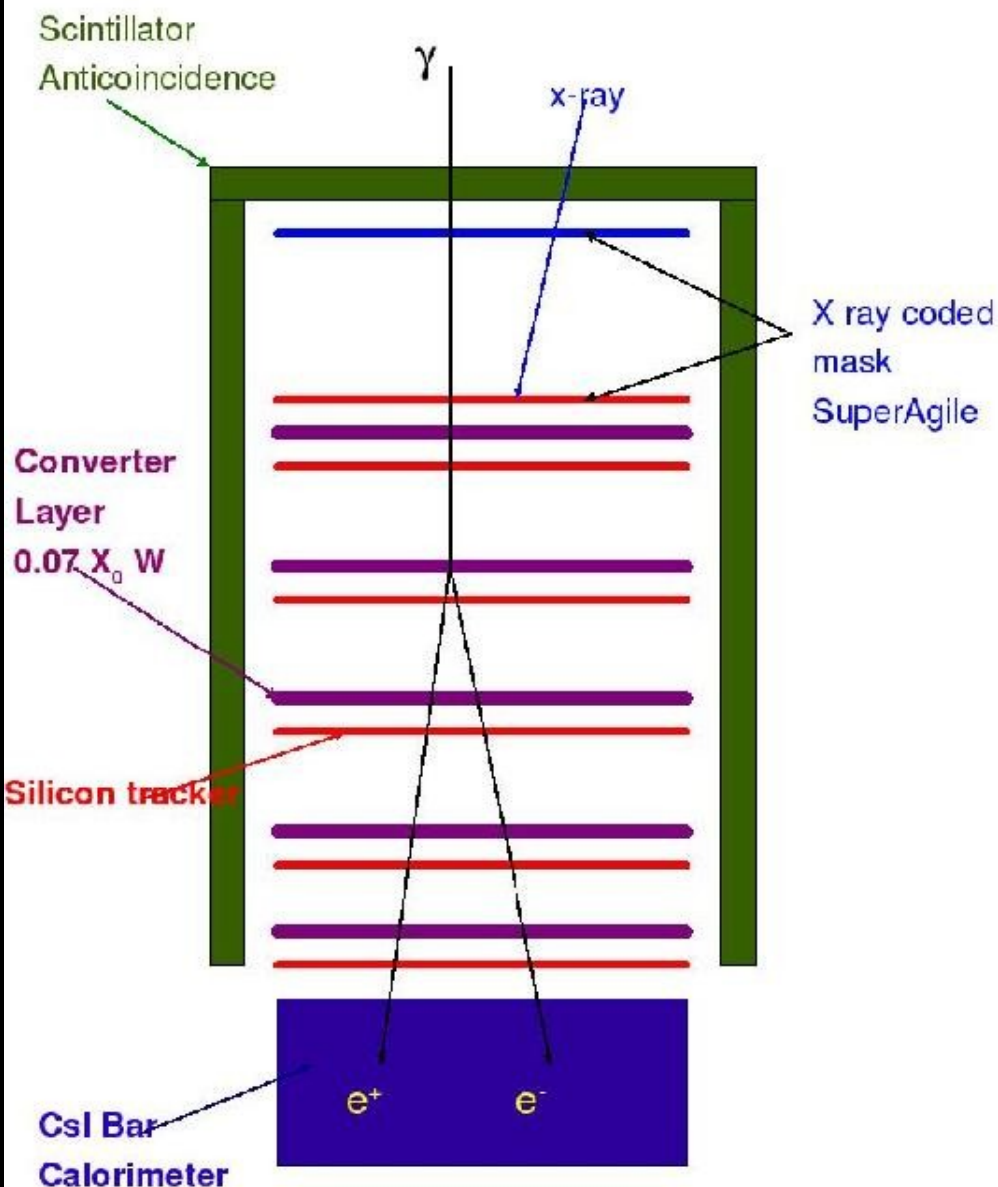


# The AGILE Payload: the most compact instrument for HE astrophysics

only ~100 kg ( $\sim 60 \times 60$  cm)

- GRID gamma-ray imager (30 MeV - 30 GeV)
- SuperAGILE hard X-ray imager (18 - 60 keV)
- MCAL Minicalorimeter (0.3 - 100 MeV)

Italian Space Agency (ASI) Mission with INFN, INAF participation



**$\gamma$ -rays detected through pair conversion ( $e^+e^-$ ) then measured in the Silicon Tracker**

**Hard X-ray detected with Coded mask followed by silicon plane**

**MiniCALorimeter of CsI Bar ( $1.5 X_0$ ). Acting standalone as GRB, TGF detector.**



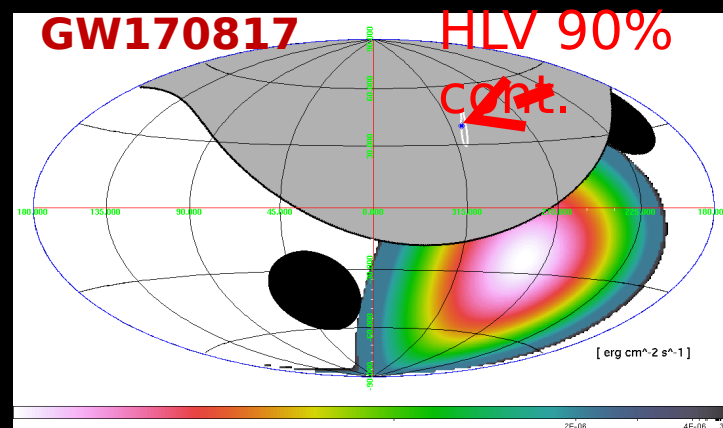
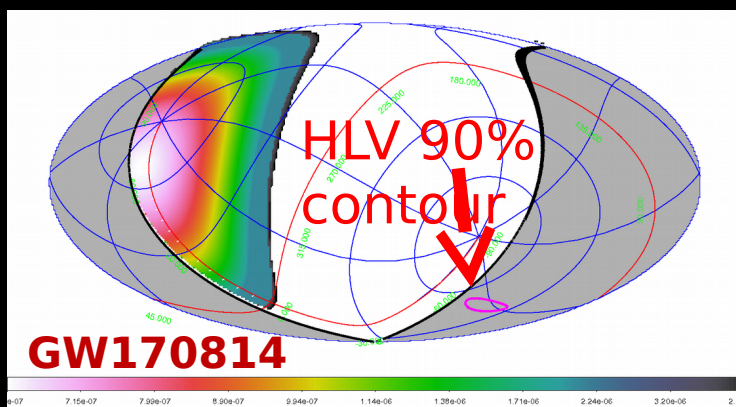
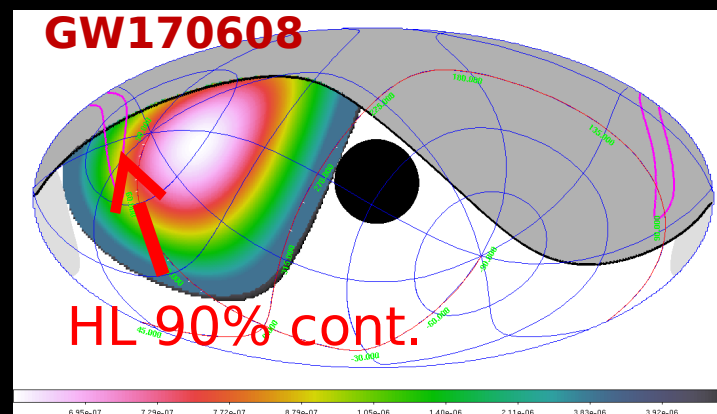
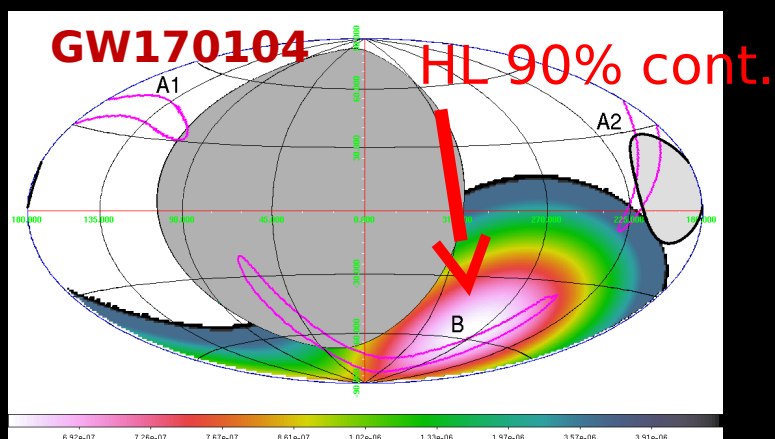
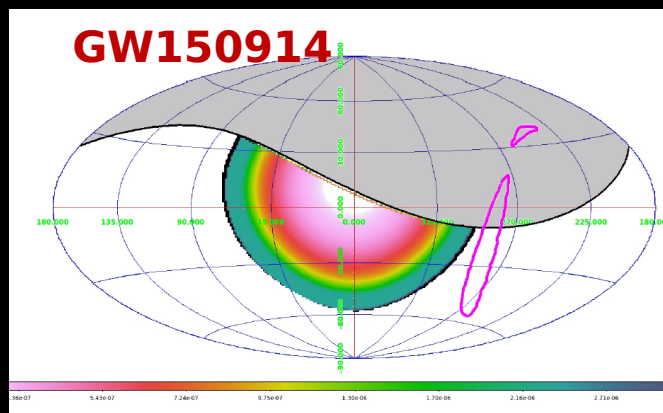
## AGILE and GW astrophysics

- very fast reaction to external GW trigger
- new processing pipelines
- great potential for fast discovery of gamma-ray transients associated with NS-NS, NS-BH and BH-BH (if any) coalescences
- AGILE GW-Team monitoring shifts (24/7) during the O2 GW LIGO-Virgo observing run.

# AGILE search for gamma-ray counterparts of GW events

GW ID	AGILE GCN #s	% coverage of 90% c.l. contour	NEAREST EXP.	Comments on Prompt and papers
150914	----	0 %	+330	<b>Prompt just missed;</b> Tavani et al. 2016
<b>151226</b>	<b>----</b>	<b>30 %</b>	<b>0</b>	<b>Partially covered; ---</b>
<b>170104</b>	<b>20375,20395</b>	<b>36 %</b>	<b>0</b>	<b>Partially covered GRID, covered by MCAL;</b> Verrecchia et al.2017a
<b>170608</b>	<b>21224,21228</b>	<b>40 %</b>	<b>0</b>	<b>Partially covered GRID, covered by MCAL for a few tens of ms; ---</b>
170814	21477, 21482	0 %	+ 500	<b>Not covered (1<sup>st</sup> with Virgo data);</b>
<b>170817</b>	<b>21525,21526, 21562, 21785</b>	<b>0 %</b>	<b>+ 930</b>	<b>OT NOT covered;</b> <b>Verrecchia et al. 2017b</b>

# AGILE prompt maps of main GW events







## First gravitational wave event: GW150914

- $T_0 = 9:50:45$  UT, 14 September, 2015
- Announced by LIGO/Virgo on Feb. 11, 2016
- We learned about the event on Feb. 11, 2016 (no MoU active yet): archival search
- **AGILE publication on the first GW event: Tavani et al., ApJ. 825, L4 (2016)**



# Spinning mode: GW150914 exposure

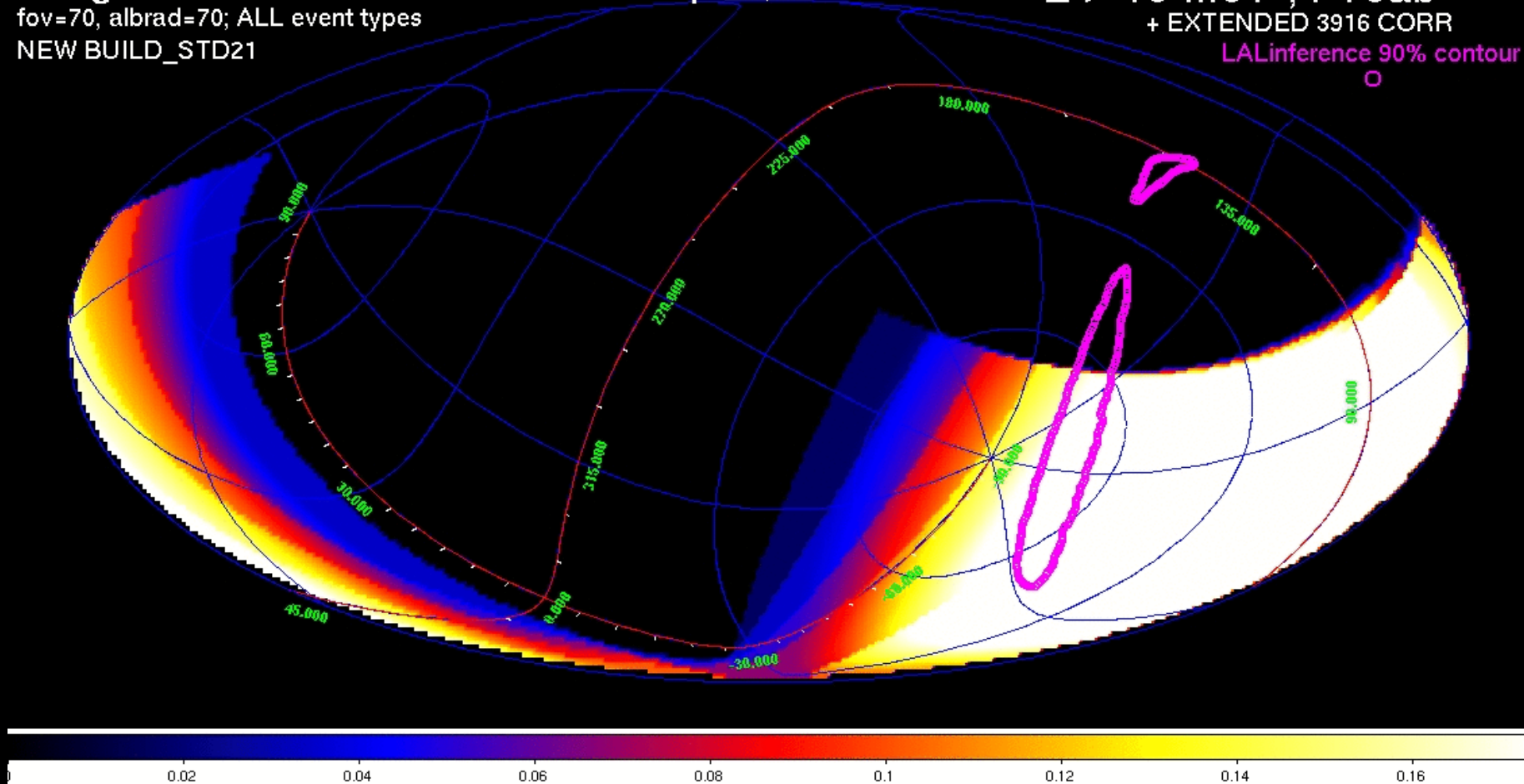
Integration: T0-205s -- T0-105s Sep14,2015

fov=70, albrad=70; ALL event types  
NEW BUILD\_STD21

$E > 10$  MeV ; FT3ab

+ EXTENDED 3916 CORR

LALInference 90% contour



# AGILE just missed the T0 !

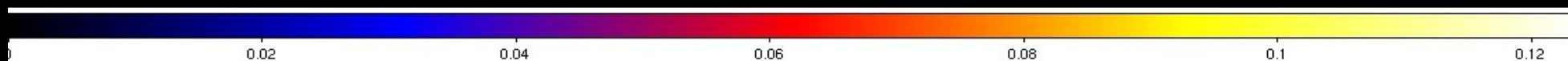
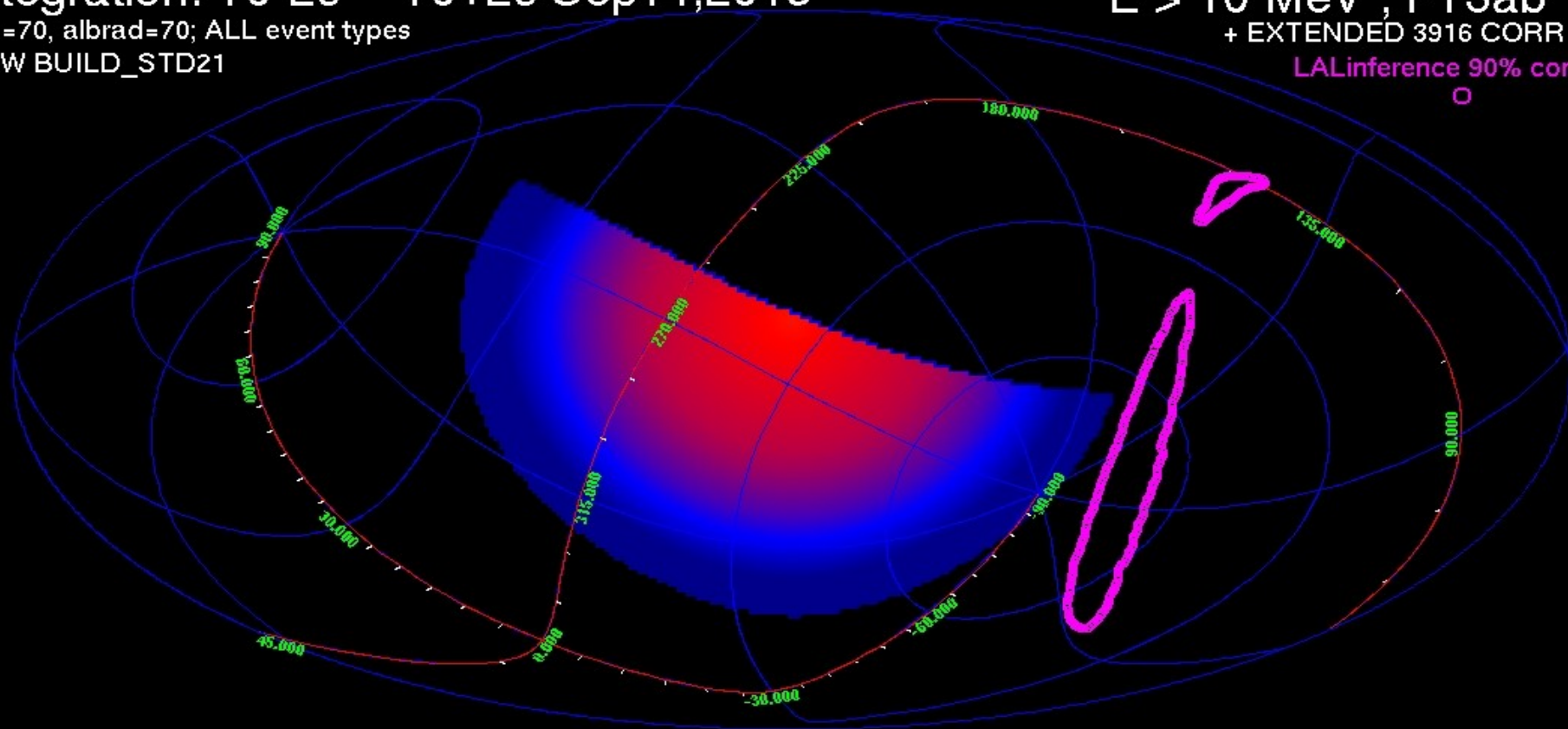
Integration: T0-2s -- T0+2s Sep14,2015

```
fov=70, albrad=70; ALL event types
NEW BUILD_STD21
```

$E > 10 \text{ MeV}$  ; FT3ab

+ EXTENDED 3916 CORR

LAInference 90% contour



## Spinning mode: one single rotation

Integration: T0-120s -- T0+300s Sep14,2015

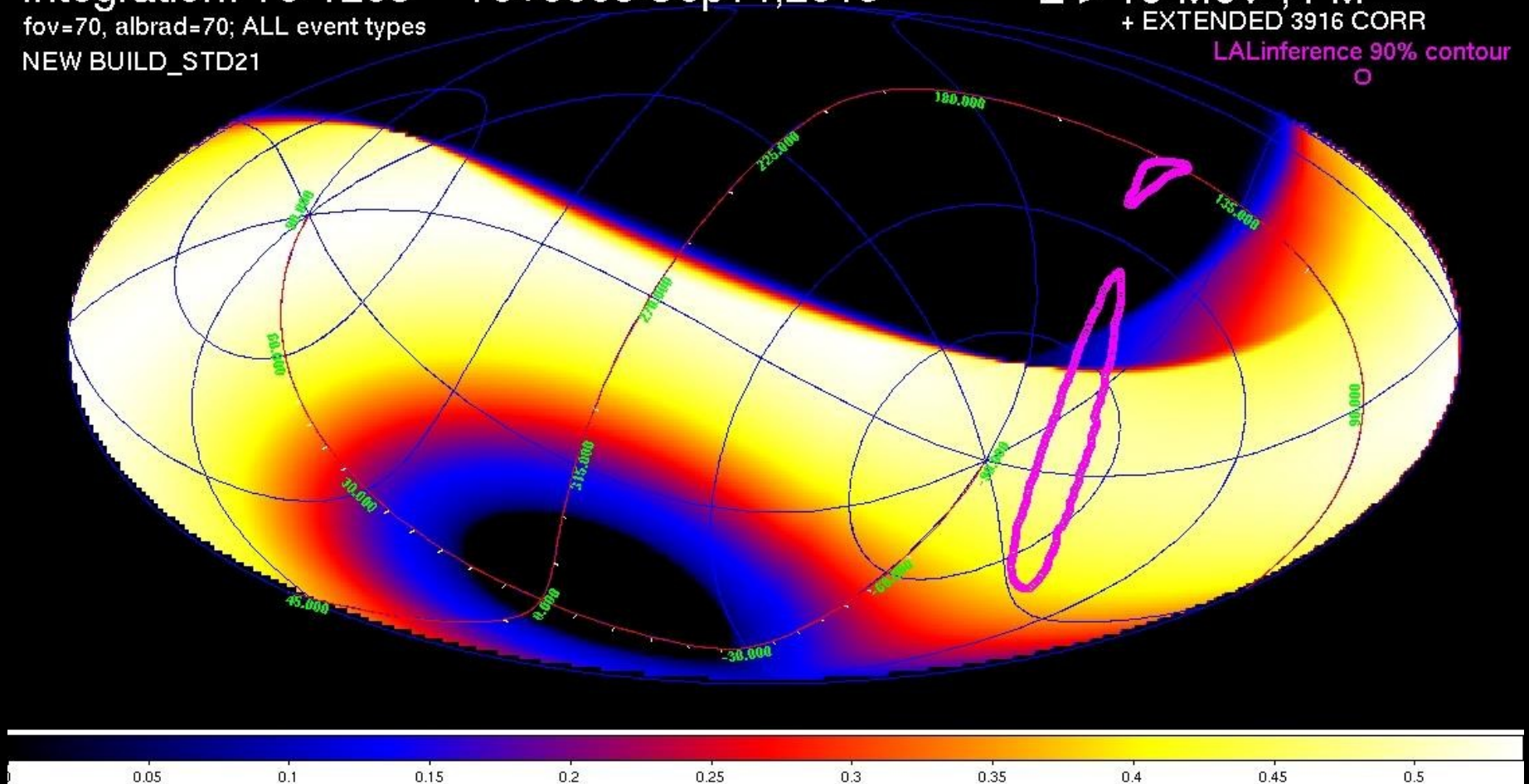
fov=70, albrad=70; ALL event types

NEW BUILD\_STD21

**$E > 10 \text{ MeV}$  ; FM**

+ EXTENDED' 3916 CORR

LALInference 90% contour

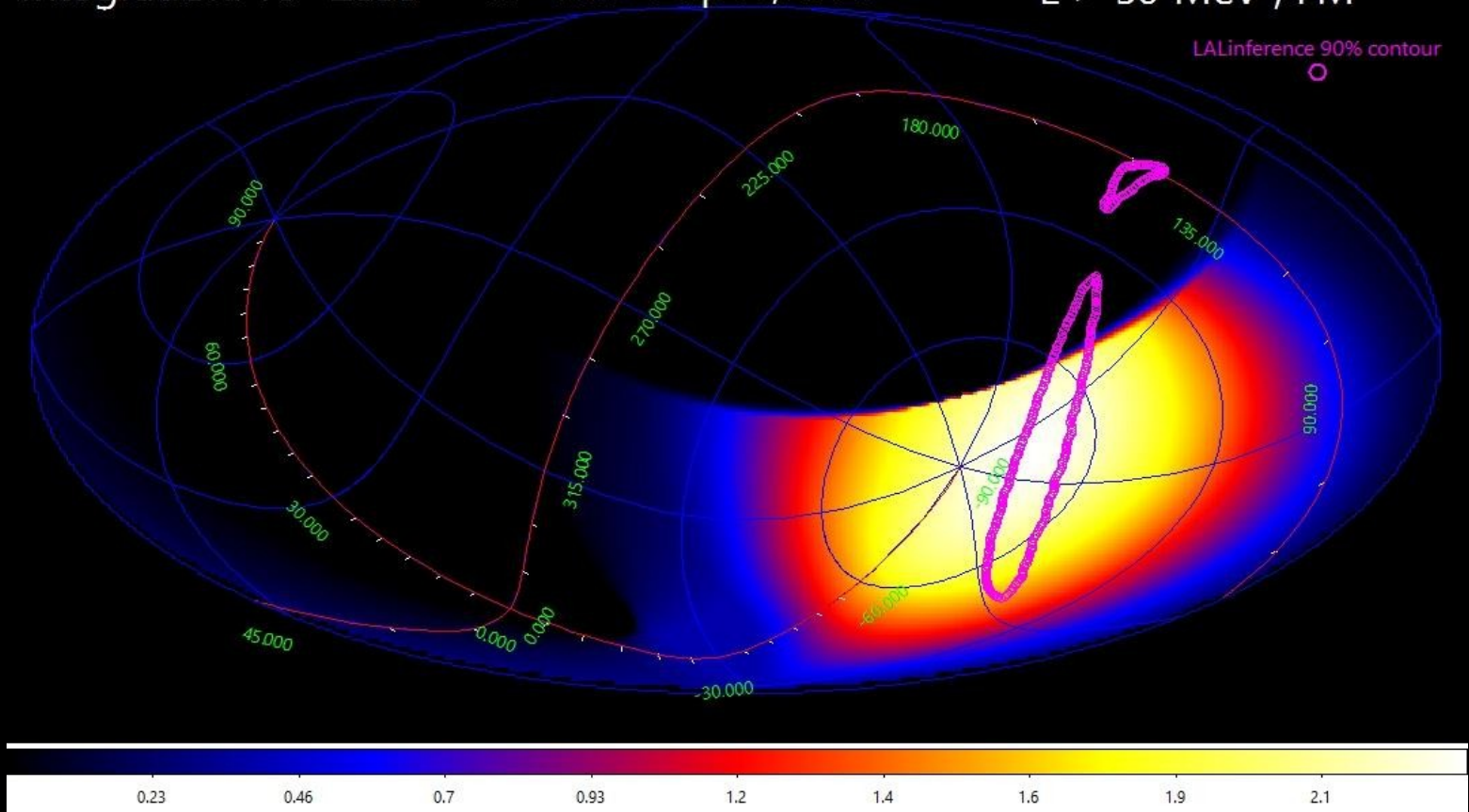


# GW150914 analysis (No MoU)

## AGILE exposure at $T_0 + 330$ (+/- 50) s

Integration:  $T_0 + 283$  s --  $T_0 + 383$  s Sep14, 2015

$E > 50$  MeV ; FM



**2-sigma upper limit ( $E > 50$  MeV) =  $1.9 \times 10^{-8}$  erg cm<sup>-2</sup> s<sup>-1</sup>**



# **GW150914 as test case for dedicated AGILE GW analyses**

- **Developed dedicated MCAL&GRID data analysis tasks**
- **Tested on archival search**
- **Pre-cursor/delayed GRID event search on short (100s) up to long (days) time scales: «grb-mode», maximum likelihood**
- **MCAL&GRID automatic data analysis reaction pipeline on-trigger**
  - **Improvement of the on-board triggering capabilities: need for larger time coverage**

# Participation to the O2 run

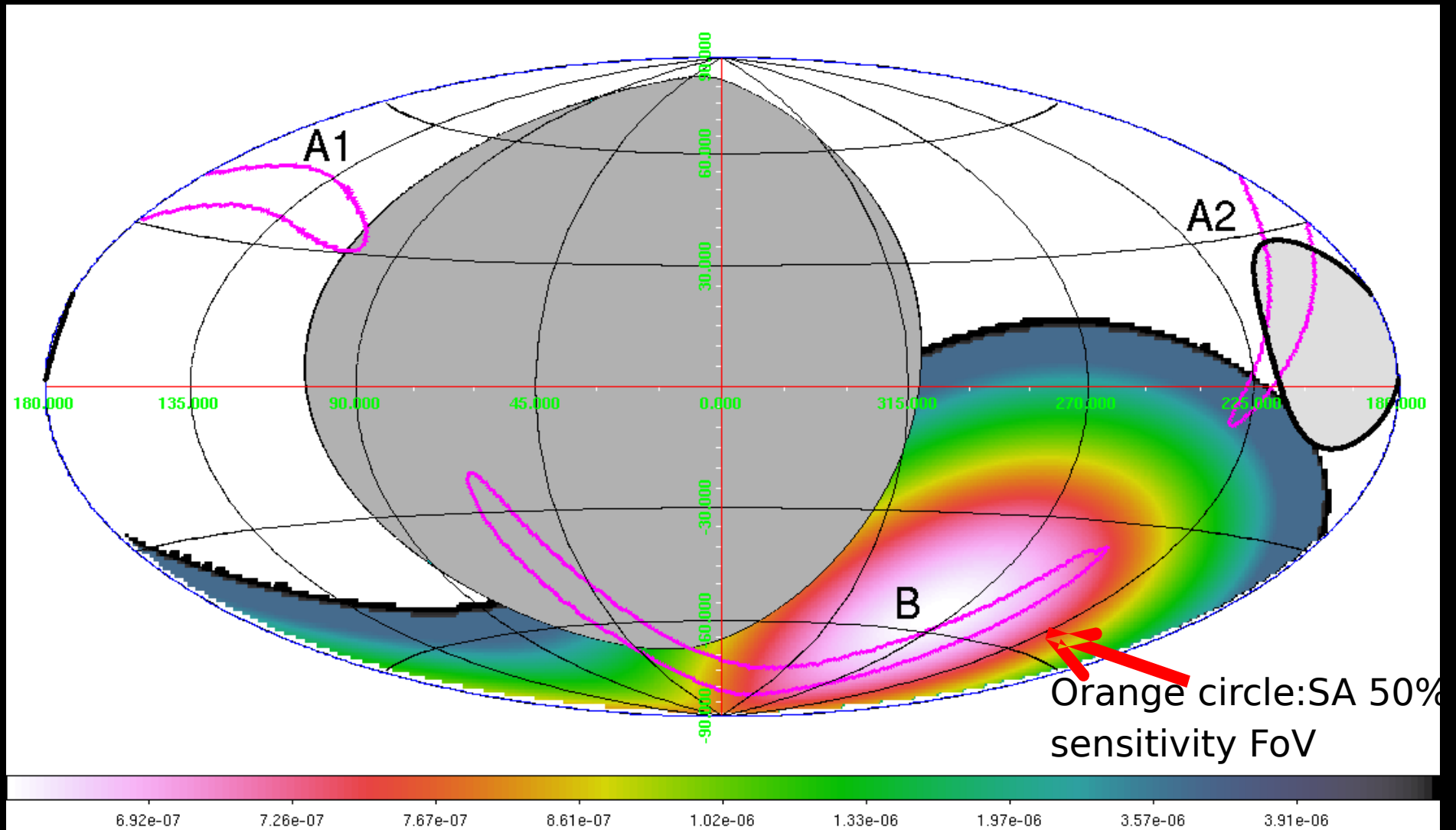
- AGILE observed all the GW events
- Main scientific results:
  - the 1st O2 event, **GW170104** : prompt partial spatial coverage in  $\gamma$ -rays ( $E > 30$  MeV), U.L.; MCAL light curve covering of the  $T_0$ .
  - the famous **GW170817** event, first with an identified e.m. counterpart: prompt occulted by Earth; nearest in time gamma-ray U.L.; precursor/delayed emission scan.

# **GW170104**

**$T_0 = 10:11:58:59$  UT, 4 January, 2017**

- **90% c.l. contour covered for 36% at event time**
- **Two LVC-GCN, by MCAL and GRID published**

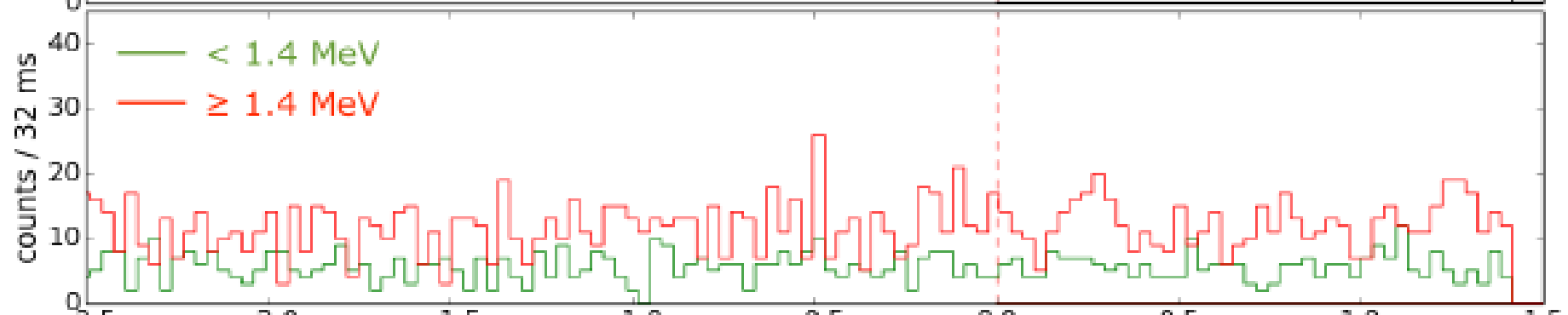
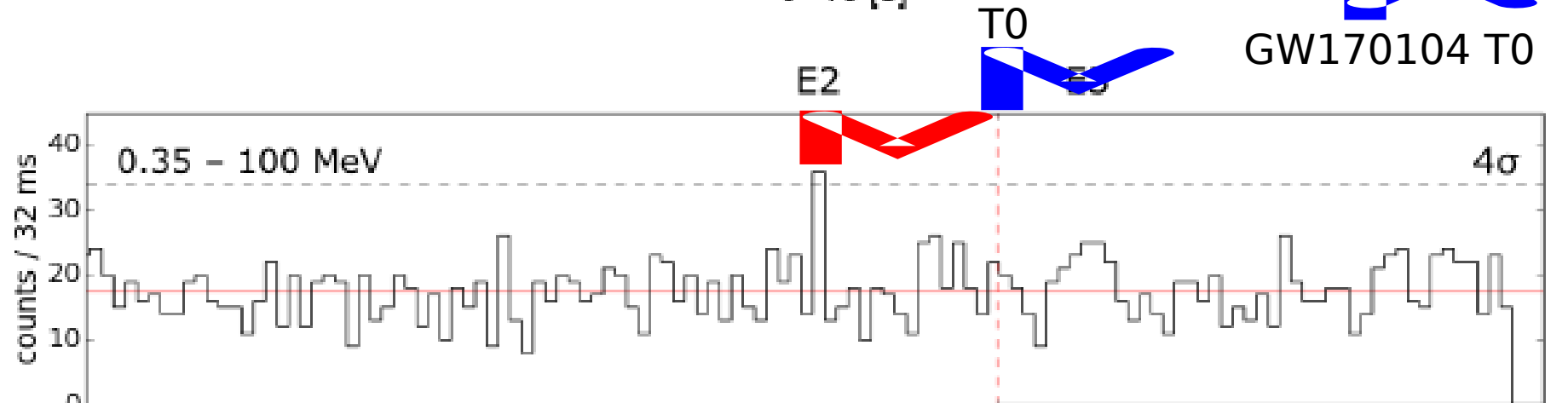
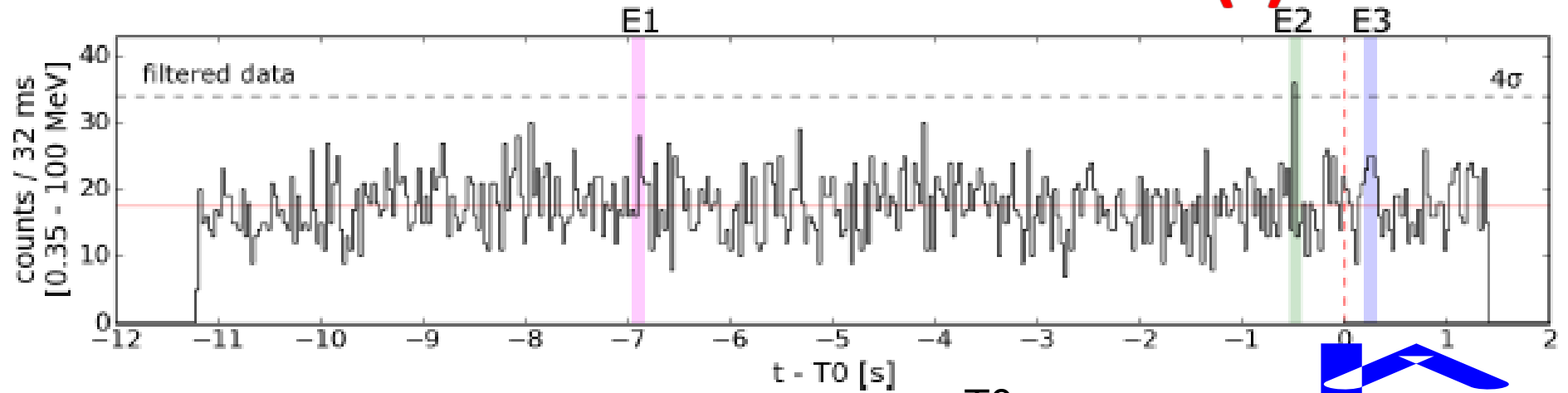
# AGILE gamma-ray exposure at trigger time (-2 / +2 sec)



**3-sigma upper limit ( $E > 50$  MeV) =  $2.9 \times 10^{-8} \text{ erg cm}^{-2} \text{ s}^{-1}$**



# an MCAL candidate event (!)

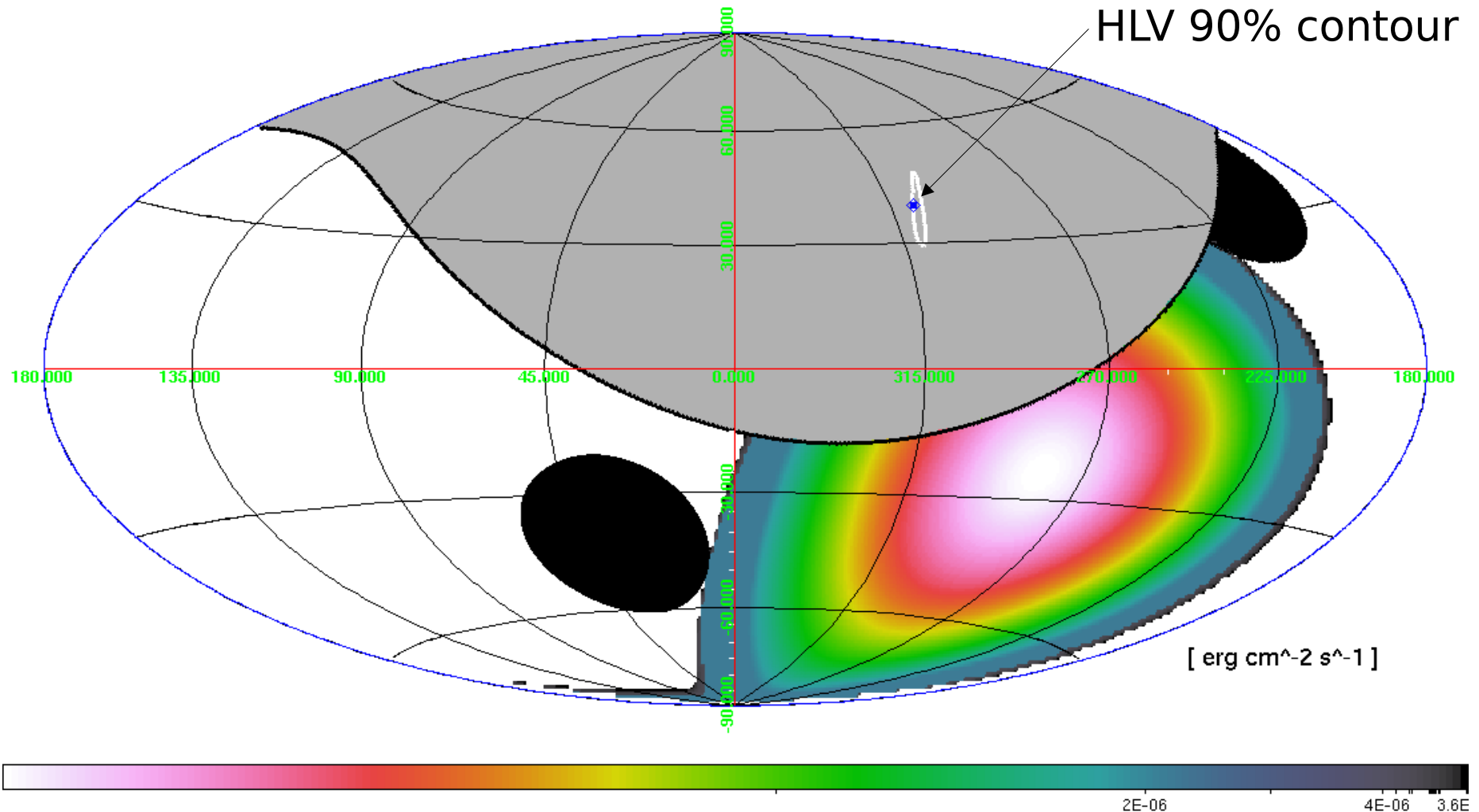


# The event: GW170817

$T_0 = 12:41:06$  UT, 17 August, 2017

- LVC error region OCCULTED by the Earth at  $T_0$
- first  $\gamma$ -ray instrument with exposure starting at  $\sim T_0 + 930s$
- Four LVC-GCN published, 1 on MCAL data analysis and 3 on GRID data, regarding the prompt data analysis and OT follow-up.

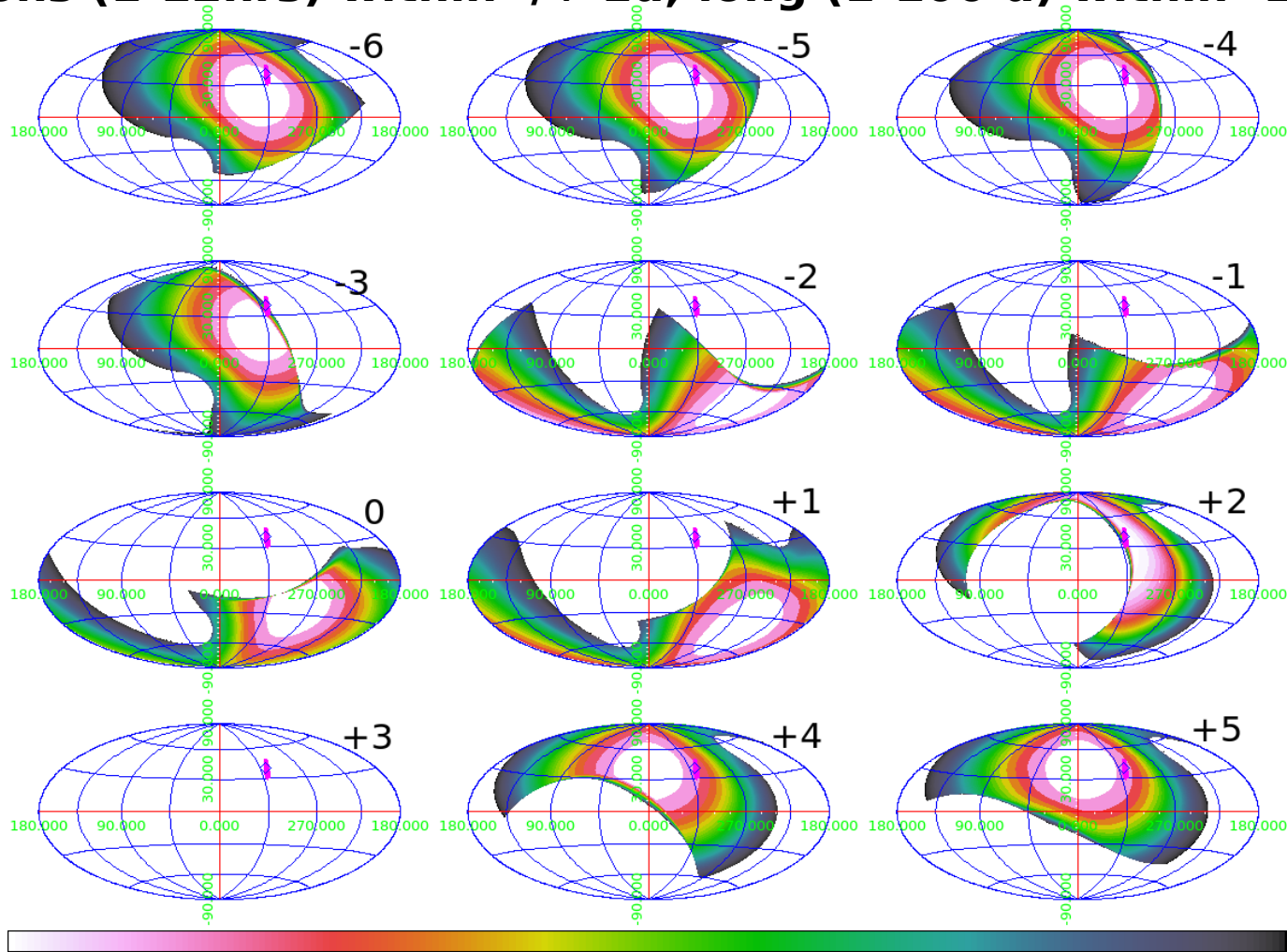
# AGILE exposure at trigger time (-2 / +2 sec) occulted!



**In  $E > 30$  MeV energy band**

# AGILE-GRID precursor/delayed emission search: short time scales (-/+1 hr)

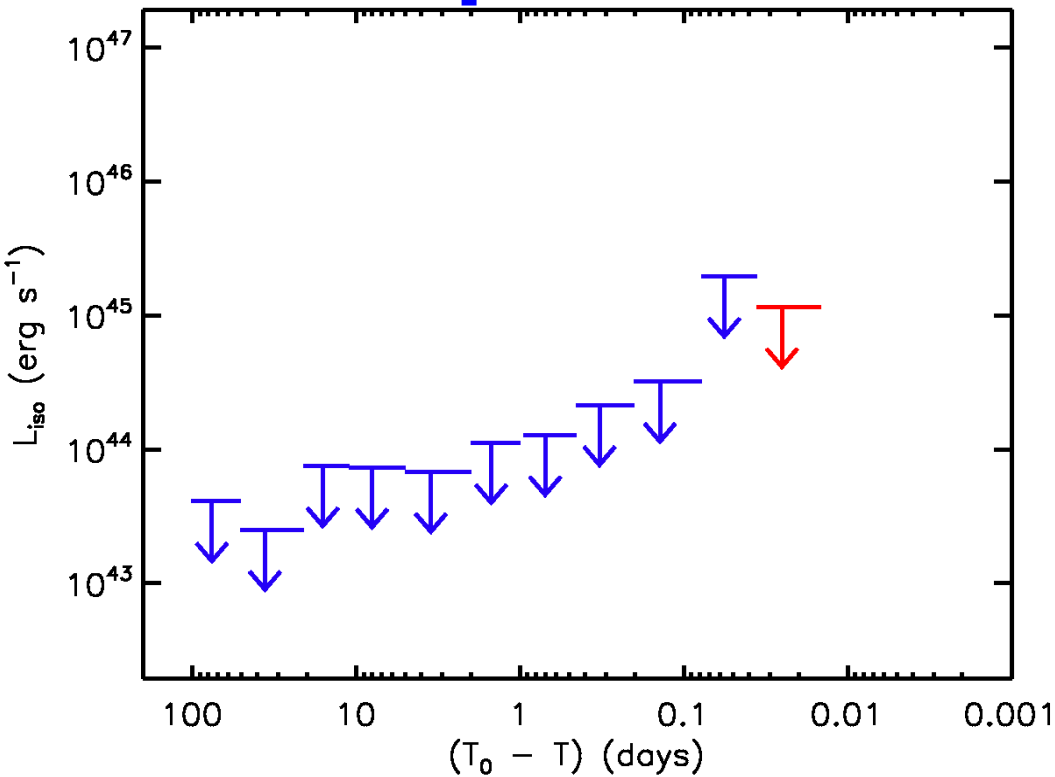
**Evaluation of GRID  $2\sigma$ -upper limit Pre-/Post  $-T_0$ @OT position, in 3 ranges:  
Short, integrations of 150s, within -/+ 1h:  
in «GRB detection mode» ( $E > 30$  MeV); medium  
integrations (1-12hrs) within -/+ 1d; long (1-100 d) within -100/+21d.**



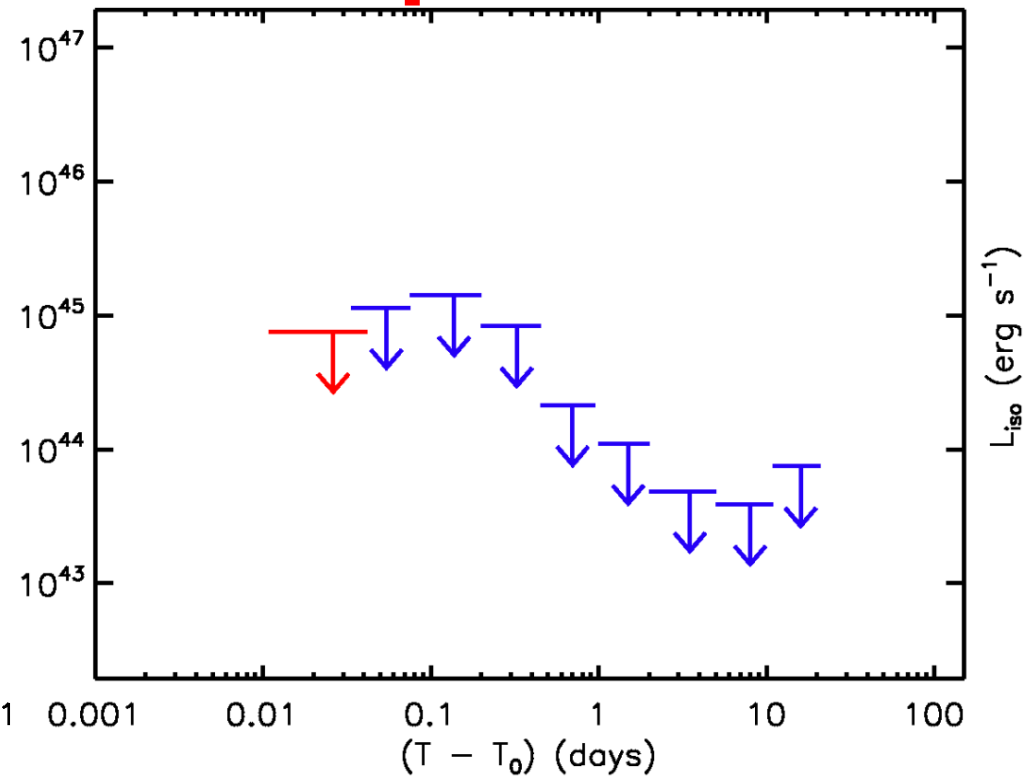


# AGILE-GRID precursor/delayed emission search: medium/long time scales (-/+1 , -101/+21 days)

**pre-**

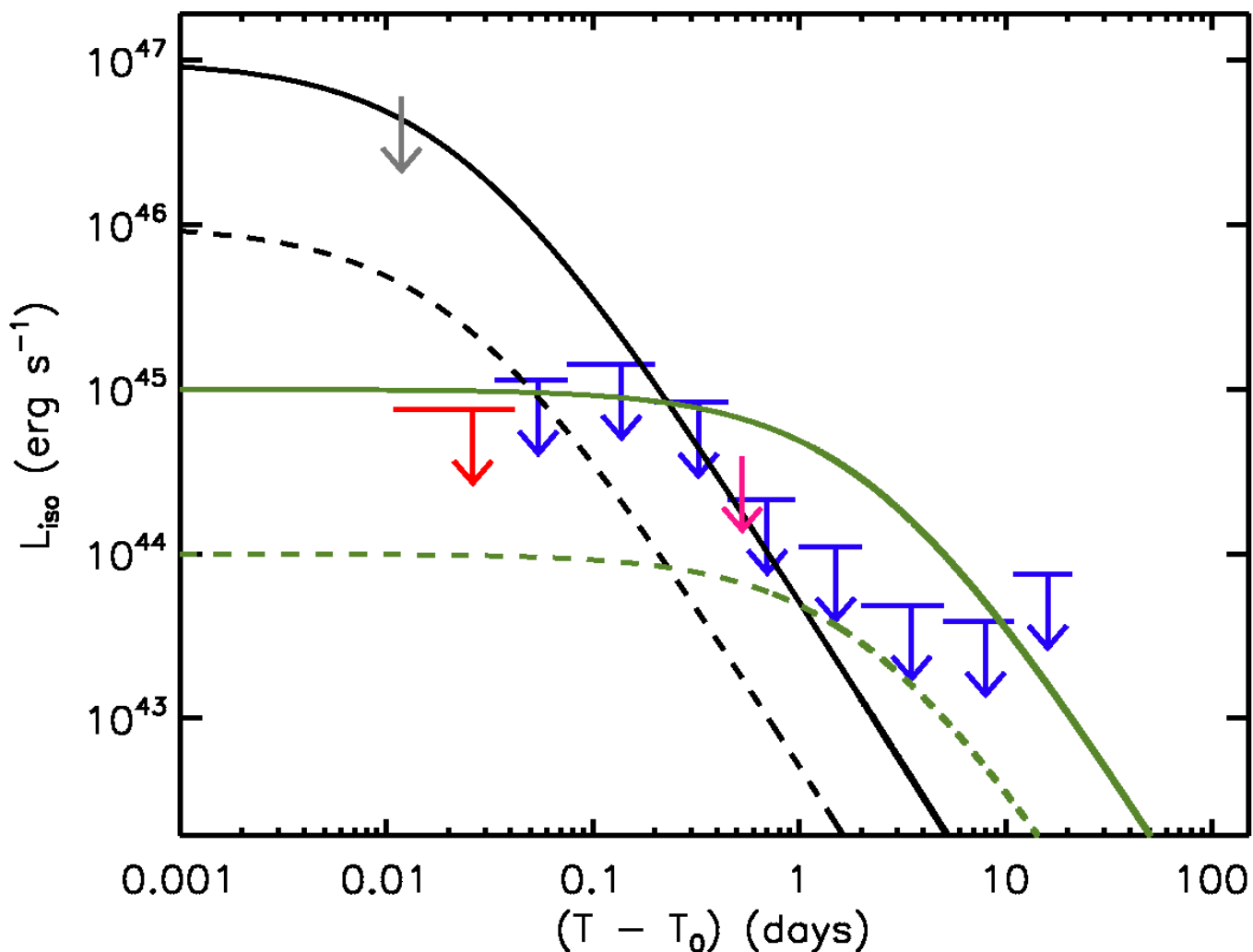


**post-**



# AGILE limits on magnetar emission:

High-energy emission from a magnetar remnant left by NS-NS coalescence model:



Radiation efficiency

$\eta = 0.01 \rightarrow$  solid

$\eta = 0.001 \rightarrow$  dashed

$B = 10^{14} \text{ G}$  (green)

$B = 10^{15} \text{ G}$  (black)

Magenta: SA U.L.

Gray: MCAL U.L.

If prompt data  
available, 170817A  
significant SA  
detection possible





## Summary 2016-2017:

- AGILE in the MoU since Nov 2016 promptly reacted to all GW candidate events communicated by LIGO-Virgo in O2 **with reaction time of 2-3 hrs** (including manual refined validation)
- 1 possible AGILE-MCAL gamma-ray transient candidate found as counterpart of GW170104 (Verrecchia et al., ApJL 847, 2017)
- **AGILE and GW170817: first  $\gamma$ -ray instrument with exposure on the localization region starting at  $\sim T_0 + 930s$  (Verrecchia et al., ApJL 850, 2017)**
- AGILE observations provided among the fastest response and the most significant upper limits above 30 MeV **to many GW events detected up to now!!**

# ))) Prospects for LIGO-Virgo O3 run

- Open Public Alerts!
  - **Automatic LVC public alerts issued within Minutes w/o human vetting.**
  - **Specific «scientific MoU» for particular topics**
- Ranges of likely rates (from LVC):
  - > **BBH: at least a few per month, maybe more**
  - > **BNS: 1-10, possibly up to ~1 per month**
  - > **NSBH: ~1 during O3, but uncertain.**
- Timeline: joint LIGO-Virgo 1-month engineering Run on **~Jan 2019; O3 Start: Feb 2019.**



# Summary

- **AGILE has unique capabilities: hard X/ $\gamma$ -ray coverage**
- **Much improved performance with enhanced MCAL and SA**
- **Ready for automatic detection of short-weak events**
- **Capable very soon to issue alerts to LVC for AGILE-detected (new channel)**
- **Important role in the astronomy of gravitational waves**

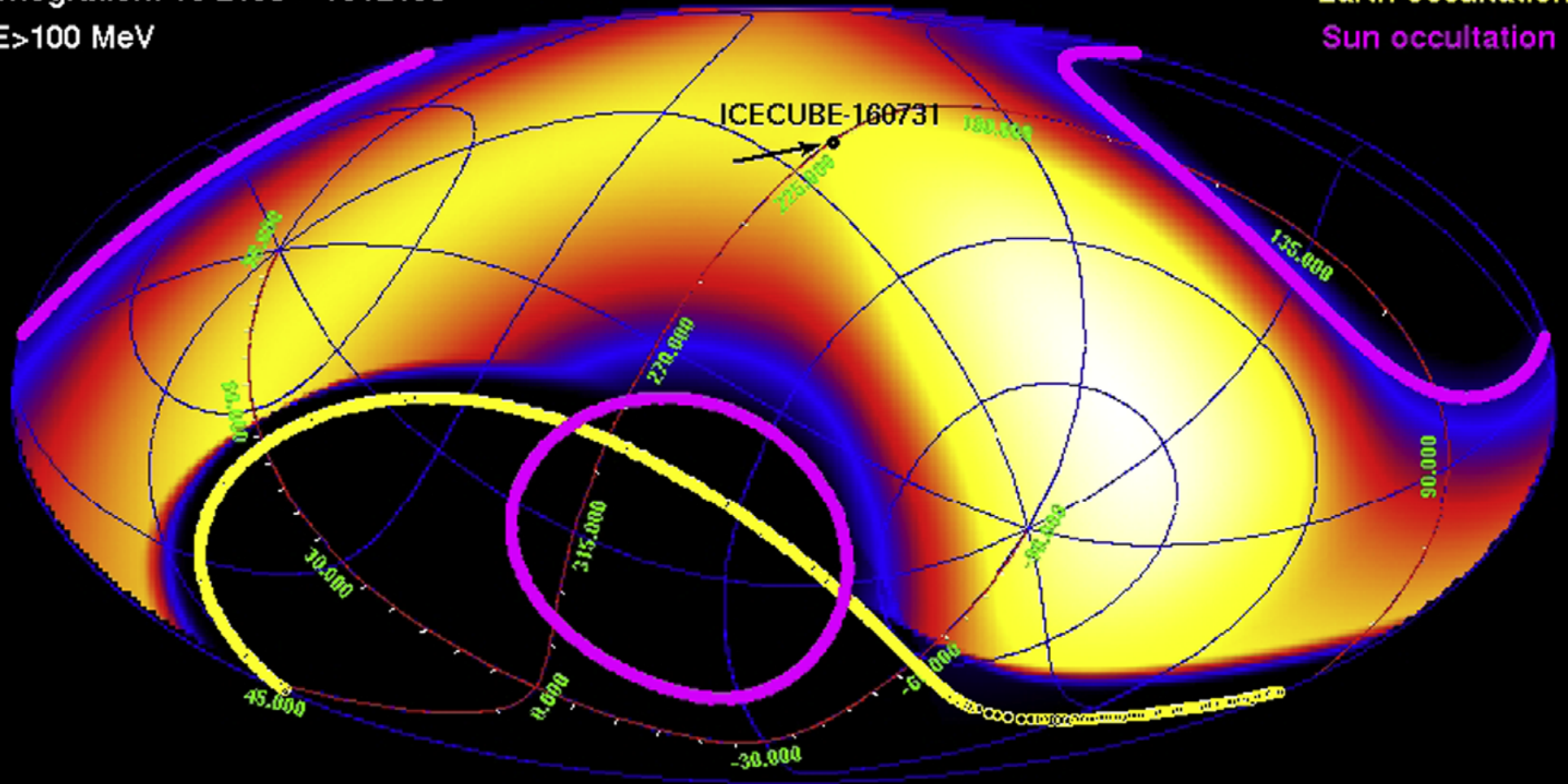
# AGILE and Neutrinos

**Integration: T0-210s -- T0+210s**

**E>100 MeV**

## Earth occultation

## Sun occultation





## Last (but not least): AGILE and neutrinos

- **AGILE and IC-160731**: Gamma-ray emission  $\sim T_0 - 1d$  announced by AGILE in ATel #9265 and further investigated in **ApJ 846 (Lucarelli et al. 2017)**
- **AGILE and IC-170922**: Gamma-ray emission **observed by both Fermi-LAT and AGILE** (ATel #10791 and **ATel #10801**) from the direction of the BL Lac blazar TXS 0506+056
- In progress:

### Automatic AGILE QL detections

Lucarelli F. – 16th AGILE Workshop

- AGILE *Quick Look* detection system searches for gamma-ray transients above 100 MeV over predefined 2-day maps
- Blind search for ground using standard AGILE **STAY TUNED!** (finder).
- Each candidate transient is then evaluated using the standard AGILE maximum likelihood (ML) algorithm.

Systematically search in the AGILE QL database for transient gamma-ray detections *spatially and temporally* consistent with the IceCube neutrino events announced since 2016.

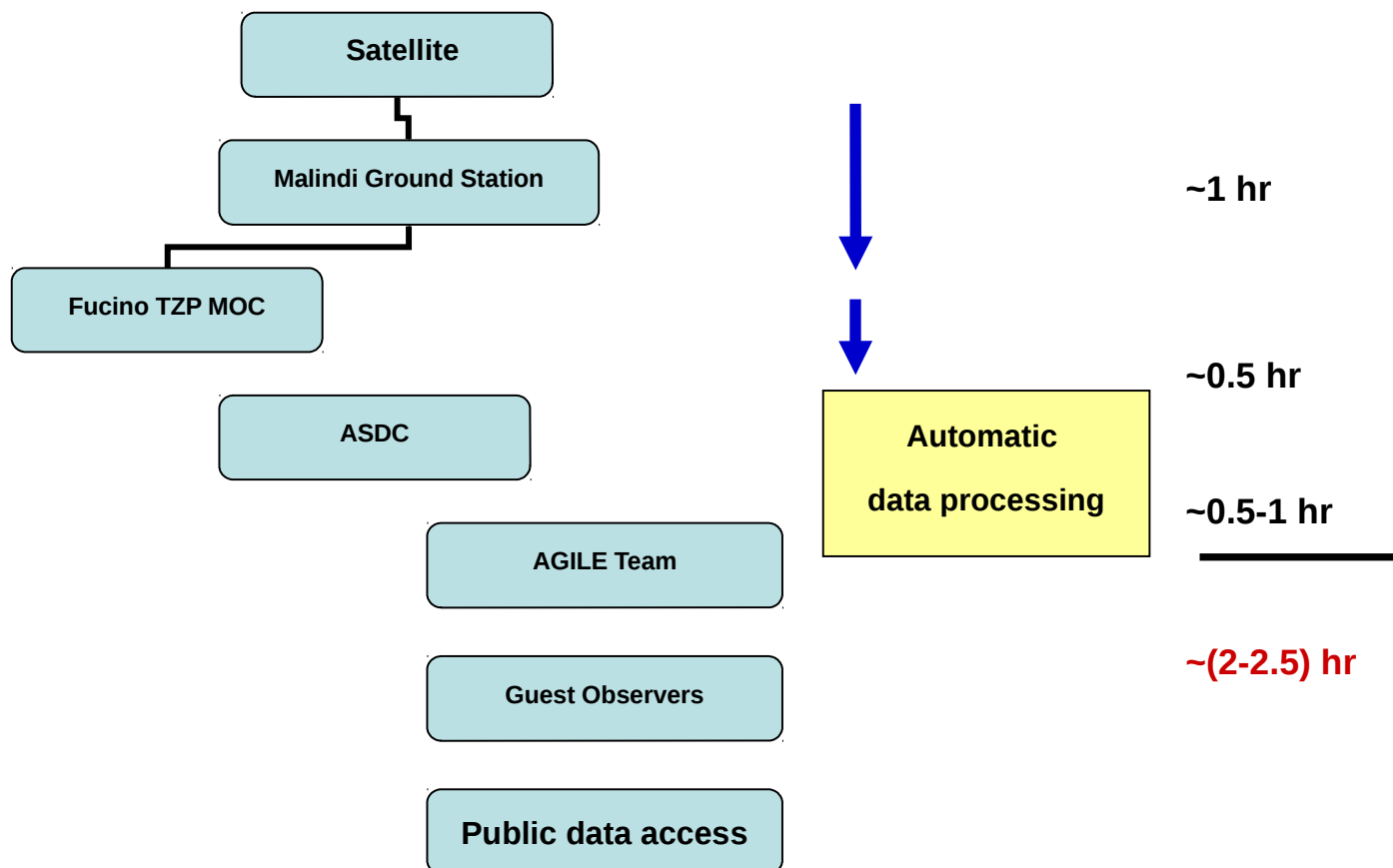
**F. Lucarelli** talk at the  
16<sup>th</sup> AGILE Workshop

# **Supplementary Slides**



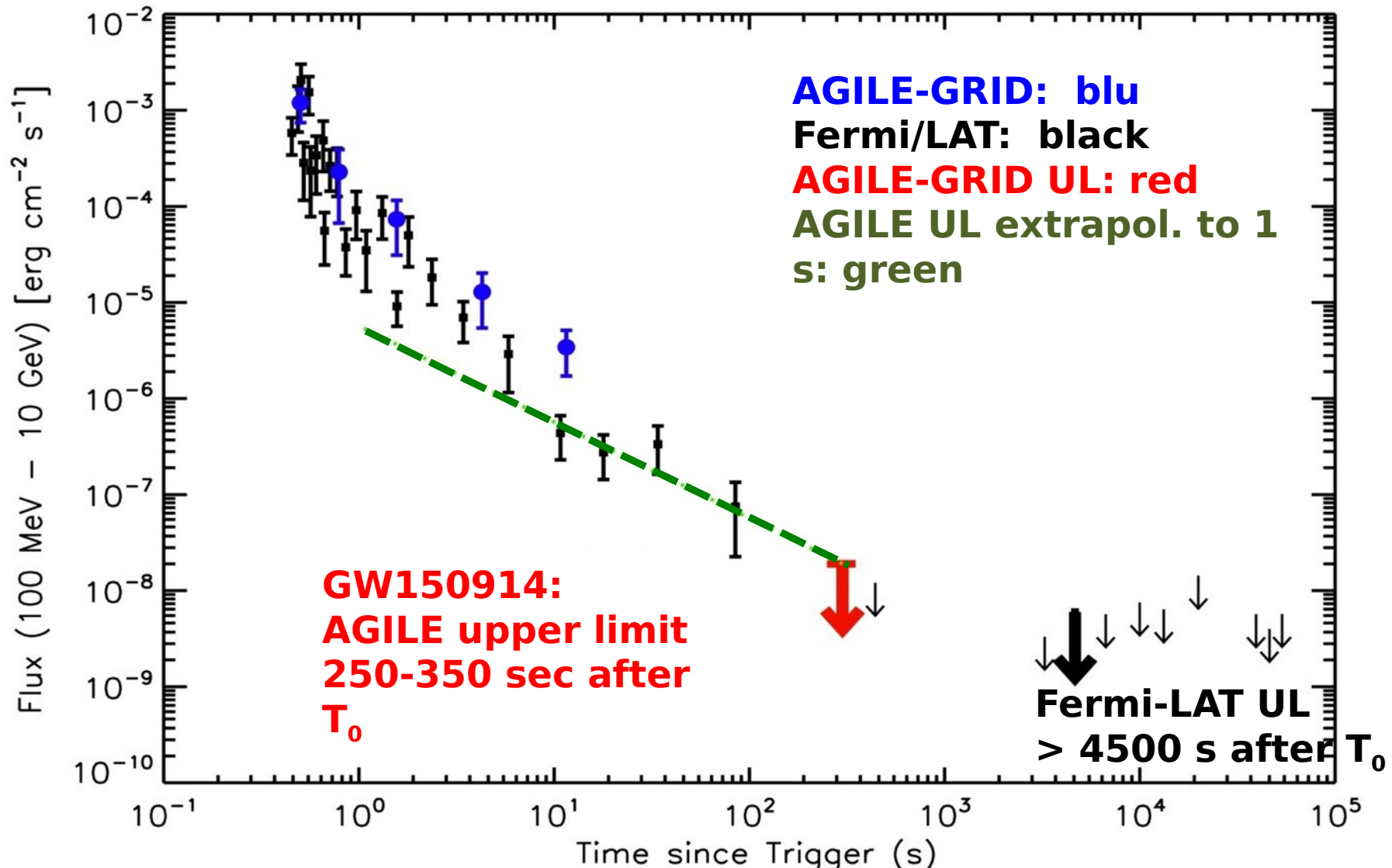


## AGILE: “very fast” Ground Segment



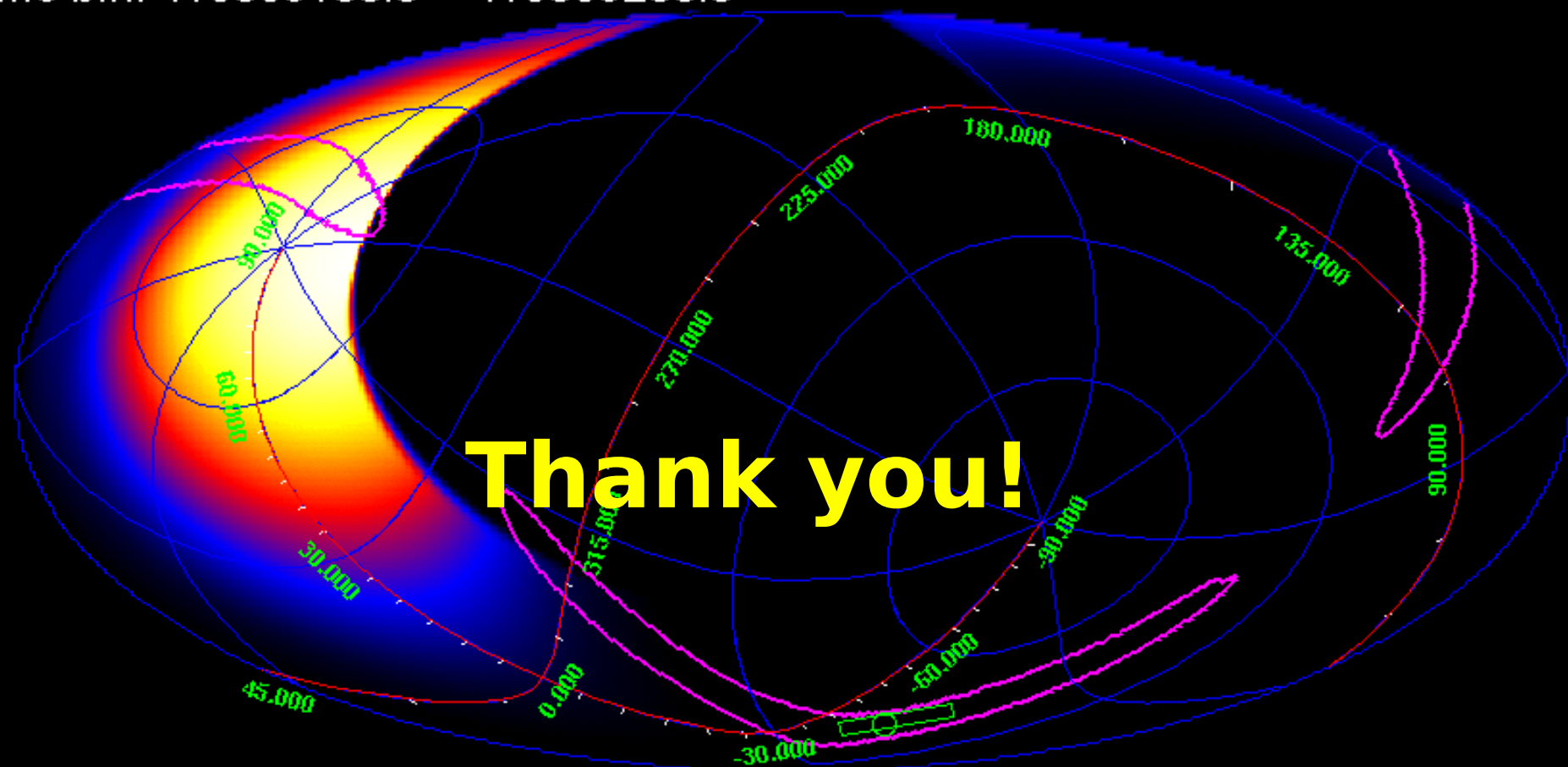
**Now even faster: ~ 25 min latency. Optimized for GW counterpart hunt!**  
**Record for a gamma-ray mission! App AGILEScience for mobile dev**

# AGILE and Fermi-LAT upper limits in the GRB 090510 light curve used as a template for GW events (scaled at $z = 0.1$ )



Bin central time: -300s  
Time bin: 410609168.6 -- 410609268.6

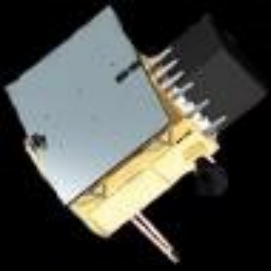
Trigger date: 2017-01-04



by F.Verrecchia on behalf of AGILE Team

fovrad=70, albrad=75

0.22 0.44 0.66 0.88 1.1 1.3 1.5 1.8 2

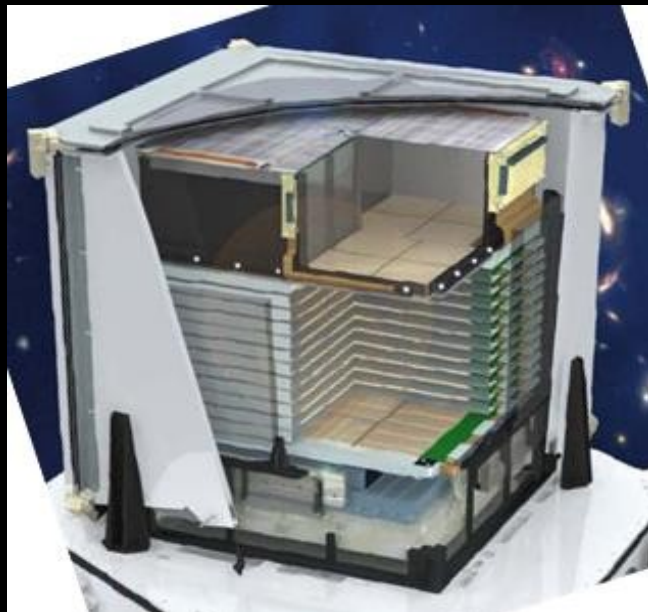


# AGILE characteristics

*Science Data Center*

**AGILE is unique combination of X-ray and gamma-ray detectors for GW searches**

**Two co-aligned detectors in hard X-rays (20-60 keV;super-A) and  $\gamma$ -rays (30 MeV-10GeV; GRID) +MCAL (0.4-100 MeV)**



**ANTICOINCIDENCE**

**HARD X-RAY IMAGER  
SUPER-AGILE (SA)**

**Energy Range: 18–60 keV**

**SILICON TRACKER**

**GAMMA-RAY IMAGER (GRID)**

**Energy Range: 30 MeV – 30GeV**

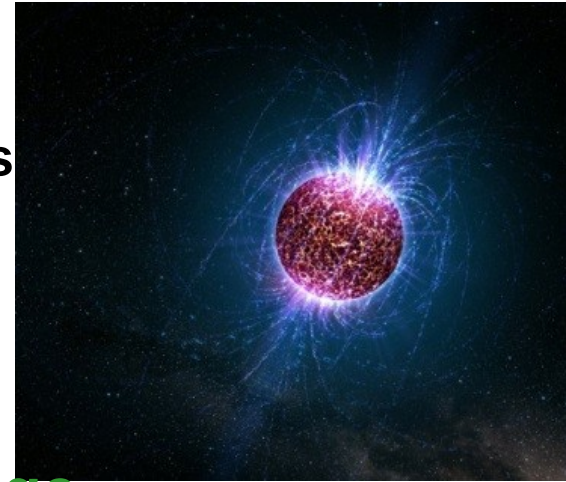
**(MINI) CALORIMETER**

**Energy Range: 0.3–100 MeV**



## AGILE limits on magnetar emission:

AGILE UL set important constraints in the early phases to **exclude** a highly magnetized magnetar for the remnant of **GW170817- GRB170817A**



### Future GW hunt:

- **AGILE fast and unique hard X/ $\gamma$ -ray coverage**
- **Improved performance with NEW MCAL pipeline developed for “sub-threshold events” btw 4  $\div$  5 sigma pre-trial significance**
- **Alerts will be issued to LVC for AGILE-detected events (new channel)**
- **Can play an important role in the new astronomy of gravitational waves. Waiting for Ligo-Virgo O3 run!**