

The ICARUS experiment

F. Tortorici

INFN Sezione di Catania
on behalf of the ICARUS Collaboration*:

*Argonne National Laboratory (ANL), USA
Brookhaven National Laboratory (BNL), USA
CERN, Geneva, Switzerland
Colorado State University, USA
Fermi National Laboratory (FNAL), USA
INFN Sez. di Catania and University, Catania, Italy
INFN GSSI, L'Aquila, Italy
INFN LNGS, Assergi (AQ), Italy
INFN Sez. di Milano Bicocca, Milano, Italy
INFN Sez. di Napoli, Napoli, Italy
INFN Sez. di Padova and University, Padova, Italy
INFN Sez. di Pavia and University, Pavia, Italy
Los Alamos National Laboratory (LANL), USA
Pittsburgh University, USA
SLAC, Stanford, CA, USA
Texas University, Arlington, USA*

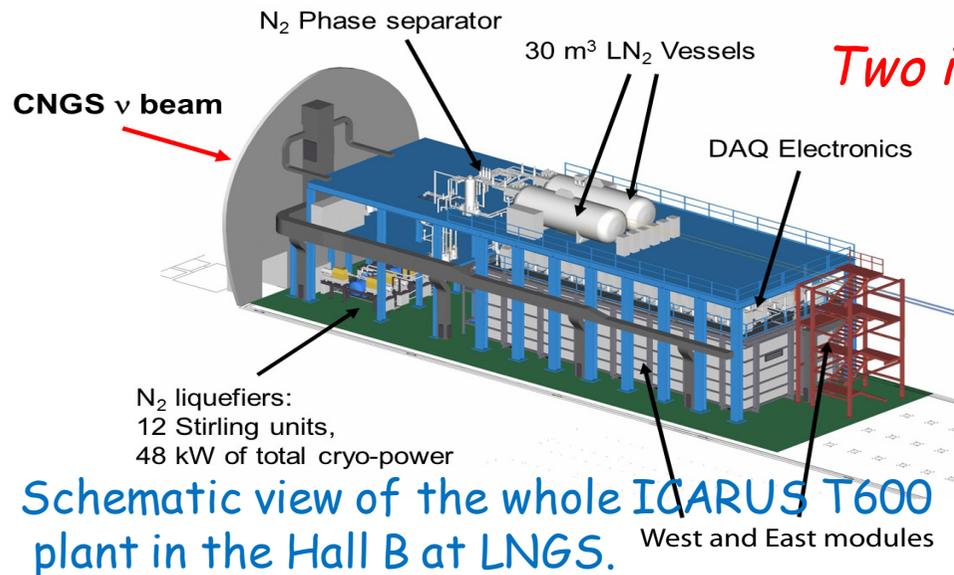
** Spokesman: C. Rubbia , GSSI*

Outline

- ✓ ICARUS LAr-TPC technology:
ICARUS T600 performance
and results @ LNGS.*
- ✓ Generalities on the ICARUS
T600 overhauling .*
- ✓ Search for sterile neutrinos.
at FNAL: the Short Baseline
Neutrino Experiment.*
- ✓ T600 current status.*
- ✓ Conclusions.*

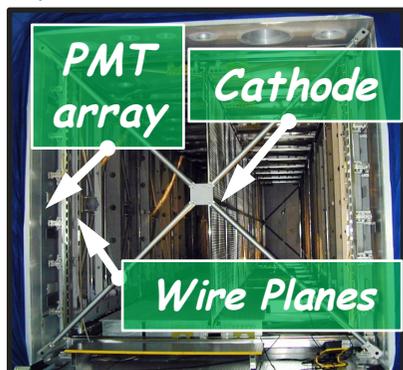
ICARUS T600: the first large Liquid Argon TPC (760 t of LAr)

- ICARUS-T600 LAr TPC is a high granularity uniform self-triggering detector with 3D imaging and calorimetric capabilities, ideal for ν physics. It allows to accurately reconstruct a wide variety of ionizing events with complex topology.
- Exposed to CNGS beam, ICARUS concluded in 2013 a very successful 3 years run at Gran Sasso INFN underground lab, collecting 8.6×10^{19} pot event statistics, with a detector live time $>93\%$, and cosmic ray events.



Two identical modules: 476 t total active mass:

- 2 TPC's per module, with a common central cathode: $E_{\text{Drift}} = 0.5 \text{ kV/cm}$, $v_{\text{Drift}} \sim 1.6 \text{ mm}/\mu\text{s}$, 1.5 m drift length;
- 3 "non-destructive" readout wire planes per TPC, ≈ 54000 wires at $0^\circ, \pm 60^\circ$ w.r.t. horizontal: Induction 1, Induction 2 and Collection views;
- Ionization charge continuously read ($0.4 \mu\text{s}$ sampling time);
- 74 8" PMT's, coated with TPB wls, for t_0 , timing and triggering.

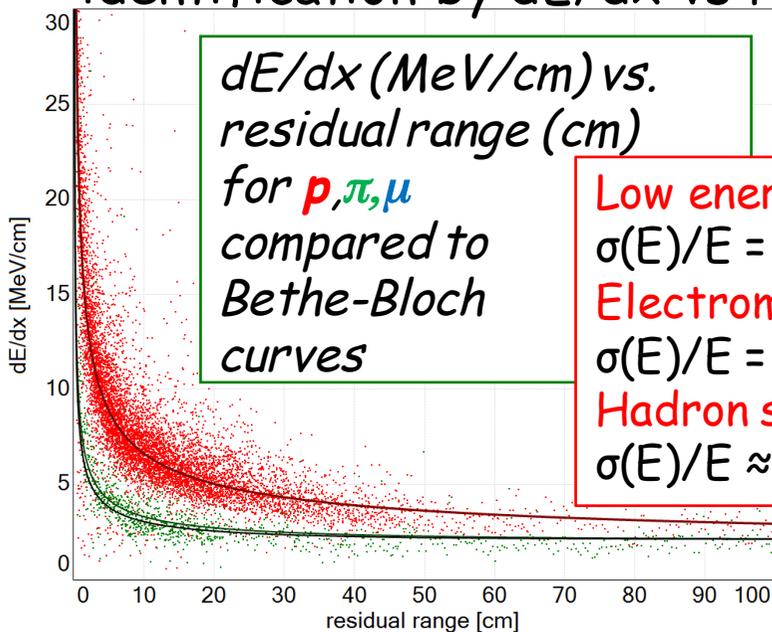


$\approx 2 \cdot 10^4$ ionization electrons per MeV

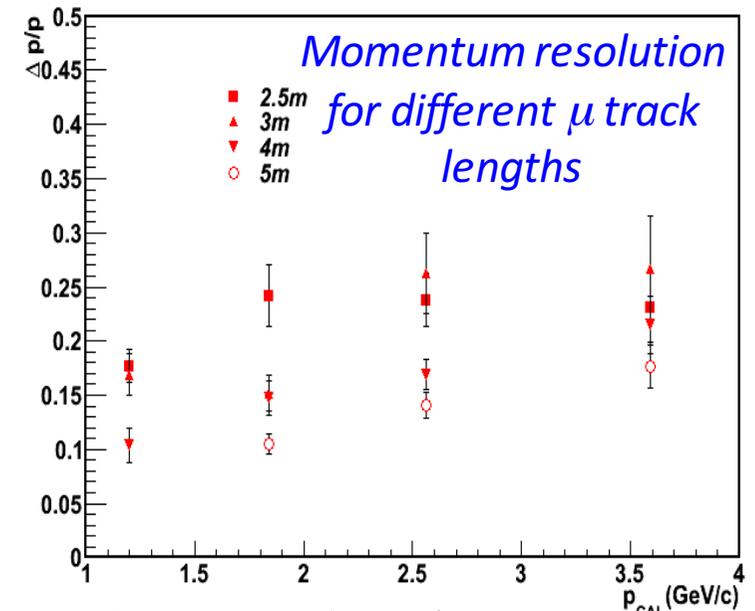
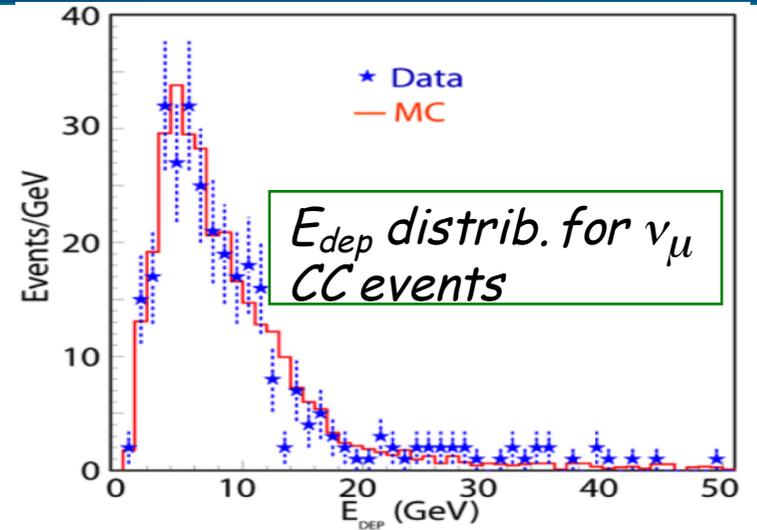
Abundant $\lambda = 128 \text{ nm}$ light ($\sim 10^4 \gamma / \text{MeV}$)

ICARUS LAr-TPC performance (CNGS ν 's and cosmics)

- **Tracking device:** precise 3D event topology, $\sim 1 \text{ mm}^3$ resolution for any ionizing particle;
- **Global calorimeter:** full sampling homogeneous calorimeter; total energy reconstructed by charge integration with excellent accuracy for contained events; momentum of non contained μ by Multiple Coulomb Scattering (MCS) with $\Delta p/p \sim 15\%$;
- **Measurement of local energy deposition dE/dx :** remarkable e/γ separation ($0.02 X_0$ sampling, $X_0 = 14 \text{ cm}$ and a powerful particle identification by dE/dx vs range):



Low energy electrons:
 $\sigma(E)/E = 11\%/\sqrt{E(\text{MeV})} + 2\%$
Electromagnetic showers:
 $\sigma(E)/E = 3\%/\sqrt{E(\text{GeV})}$
Hadron showers:
 $\sigma(E)/E \approx 30\%/\sqrt{E(\text{GeV})}$

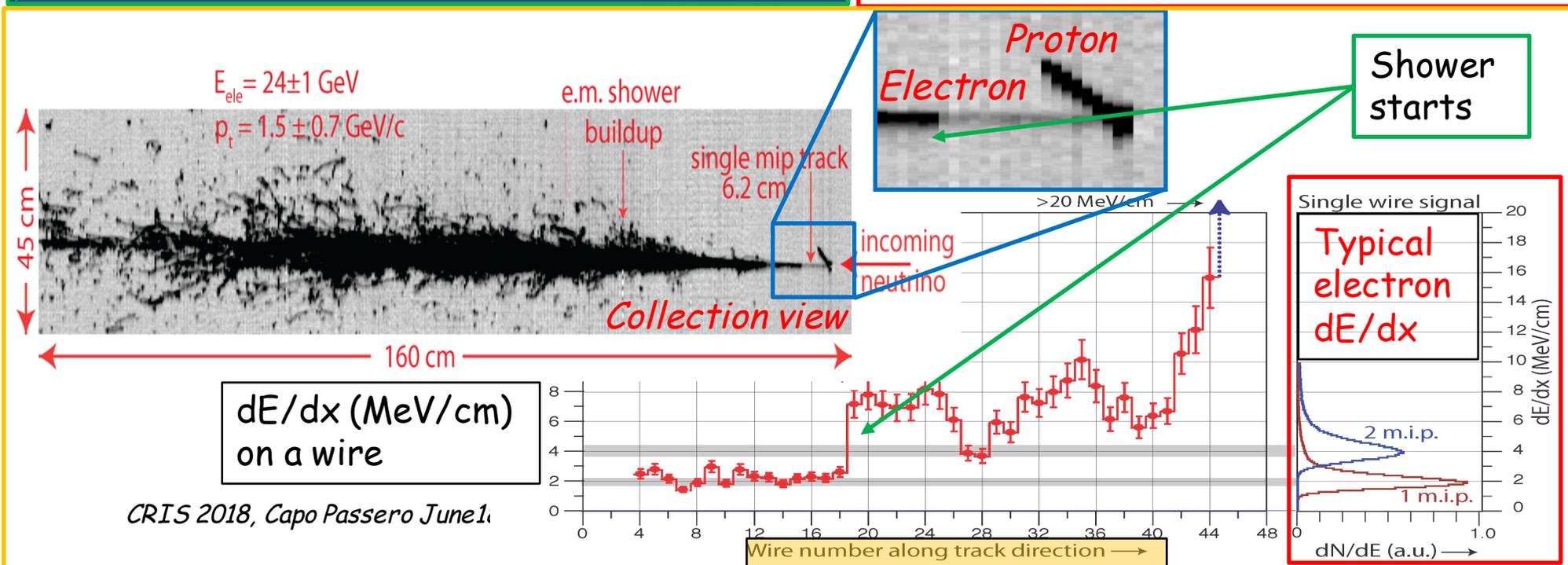
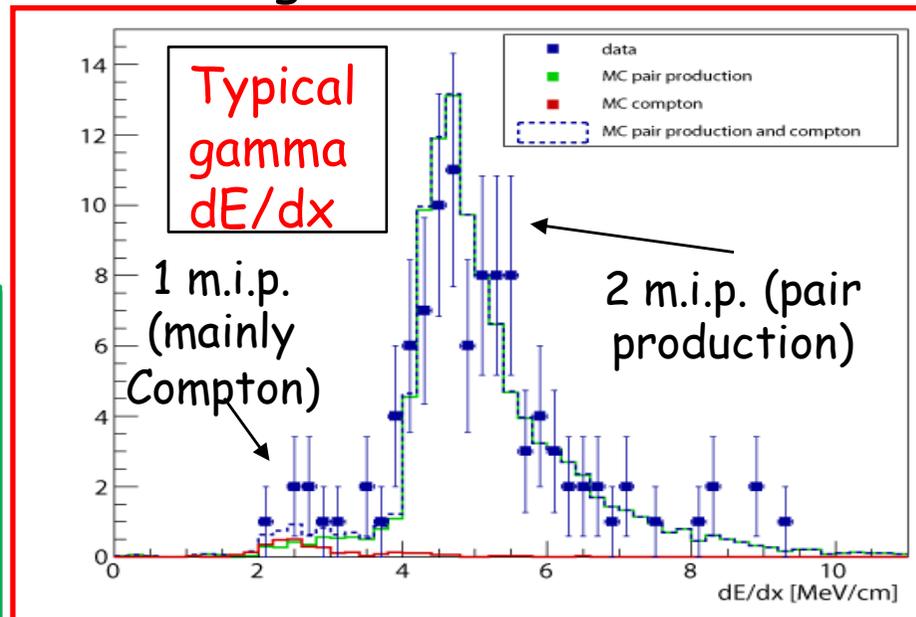
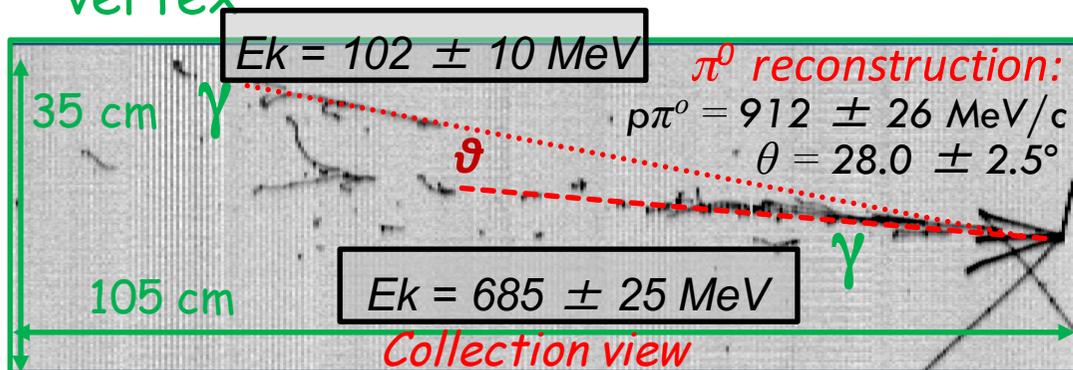


Measured with stopping μ comparing p_{MCS} with calorimetric measure

ν_e CC identification in CNGS beam: Electron/gamma separation

Three "handles" to separate e/ γ and reject NC background:

- reconstruction of π^0 invariant mass
- dE/dx : single vs. double m.i.p.
- γ conversion separated from primary vertex

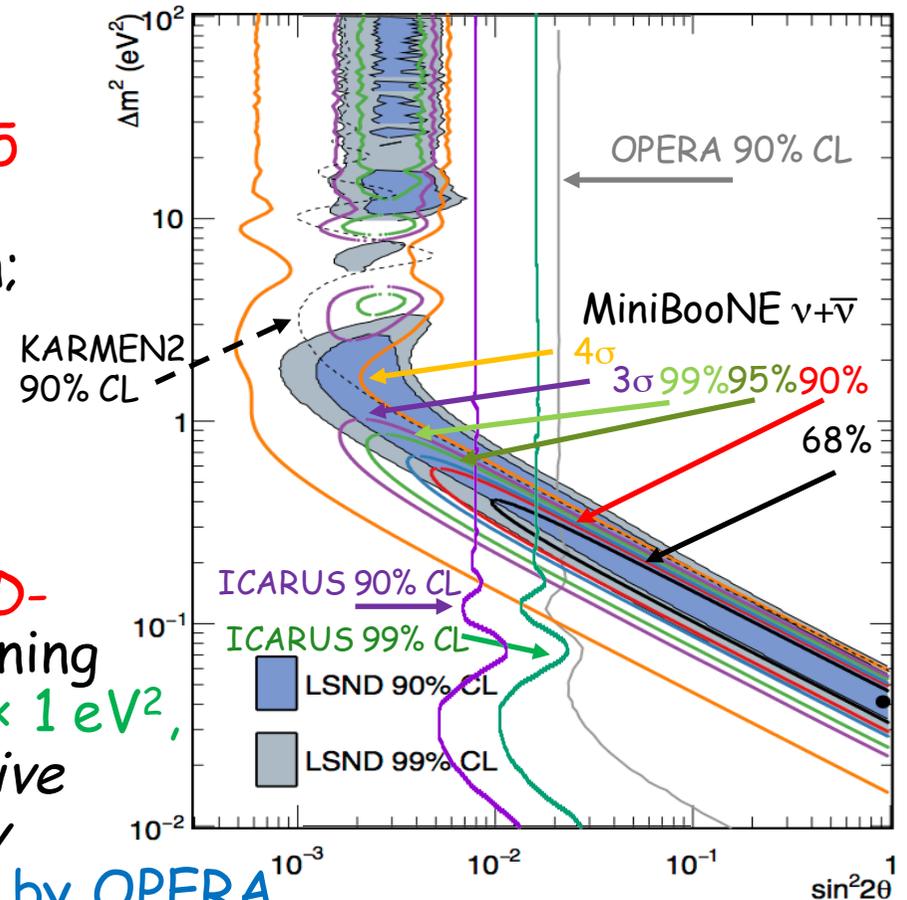


CRIS 2018, Capo Passero June 1.

The LAr-TPC technology and ICARUS-T600

- ICARUS run at LNGS allowed to reach several physics/technical result demonstrating the maturity of the LAr-TPC technology:

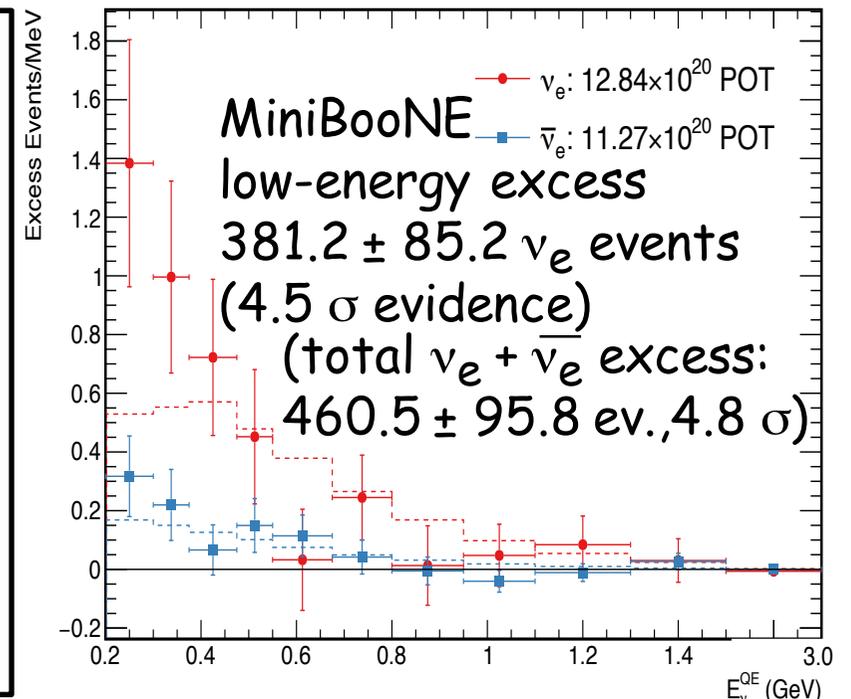
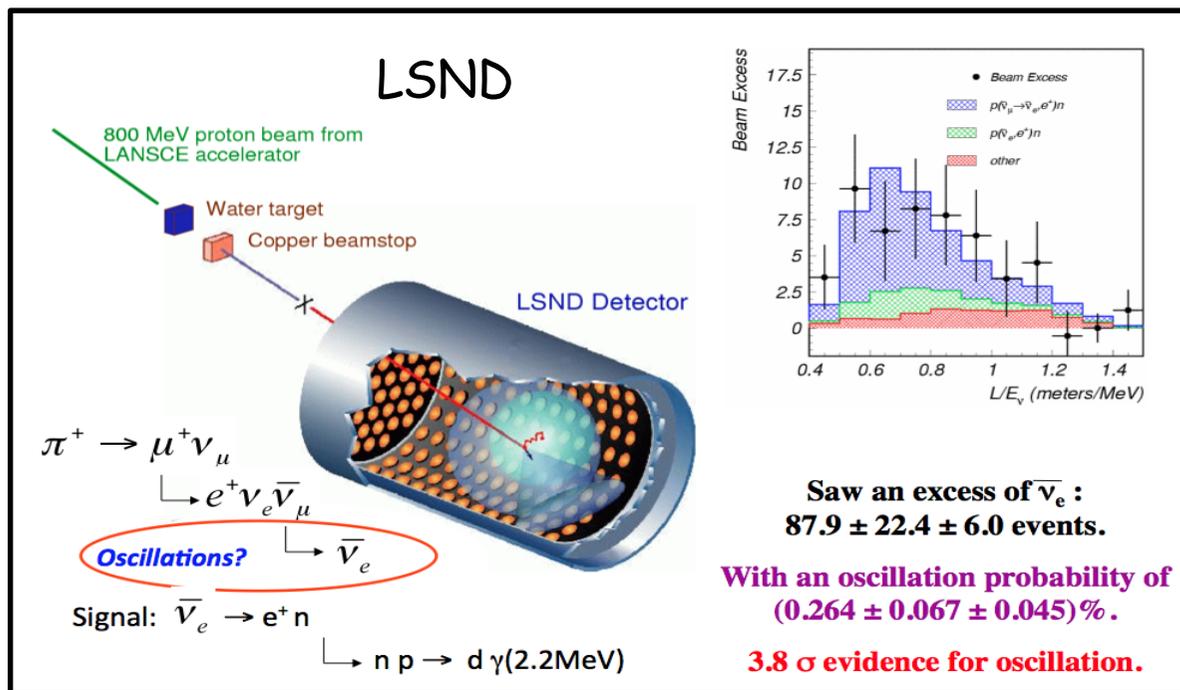
- *An exceptionally low level ~ 20 p.p.t. $[O_2]$ eq. of electronegative impurities in LAr;* the measured e^- lifetime $\tau_{ele} > 15$ ms ensured few m long drift path of ionization e^- signal without attenuation;
- *Demonstrated detector performance,* especially in ν_e identification and π^0 background rejection in $\nu_\mu - \nu_e$ study to unprecedented level;
- *Performed a sensitive search for LSND-like anomaly with CNGS beam,* constraining LSND window to narrow region at $\Delta m^2 < 1 \text{ eV}^2$, $\sin^2 2\theta \sim 0.005$ where all positive/negative experimental results can be coherently accommodated at 90% C.L., confirmed by OPERA.



- These results marked a milestone for LAr-TPC technology with a large impact on neutrino and astro-particle physics projects: SBN short baseline neutrino program at FNAL with 3 LAr-TPC's (SBND, MicroBooNE and ICARUS) and the multi-kt DUNE LAr-TPC.

“Sterile neutrino puzzle” 1/2

- Anomalies have been collected in last years in neutrino sector, despite the well-established 3-flavour mixing picture within Standard Model:
 - appearance of ν_e from ν_μ beams in accelerator experiments (LSND 3.8σ , MiniBooNE 4.8σ ($\nu_e + \text{anti-}\nu_e$));
 - disappearance of anti- ν_e , hinted by near-by nuclear reactor experiments (formerly, ratio observed/predicted event rate $R = 0.943 \pm 0.023$, Mention et. al, 2011);
 - disappearance of ν_e , hinted by solar ν experiments during their calibration with Mega-Curie sources (SAGE, GALLEX, $R = 0.84 \pm 0.05$).



CRIS 2018, Capo Passero June 18-22th

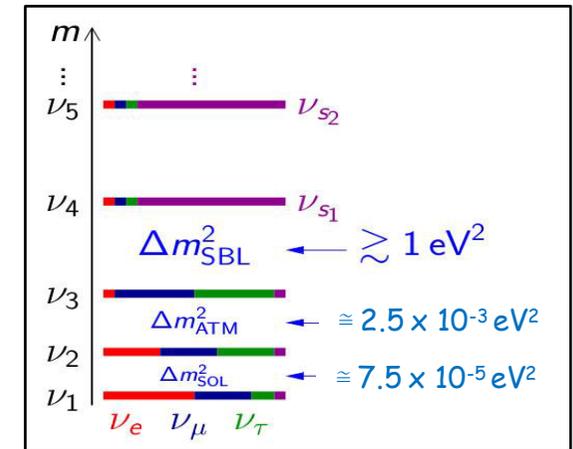
1805.12028 [hep/ex], neutrino2018

Slide#: 7

“Sterile neutrino puzzle” 2/2

- Results **hint to a new “sterile” flavor**, described by $\Delta m^2 \sim eV^2$ and small mixing angle, driving oscillations at short distance:
 - ICARUS indicates $\Delta m^2_{new} \leq 1 eV^2$, small mixing;
 - Planck data and Big Bang cosmology point to at most one further flavor with $m_{new} < 0.24 eV$;
 - **No evidence for ν_μ disappearance** in MINOS and IceCube in 0.32-20 TeV;
 - Recent reactor data are intriguing but inconclusive (global analysis prefers sterile ν hypothesis at $\sim 3\sigma$ with $\Delta m^2 \sim 1.3 eV^2$).

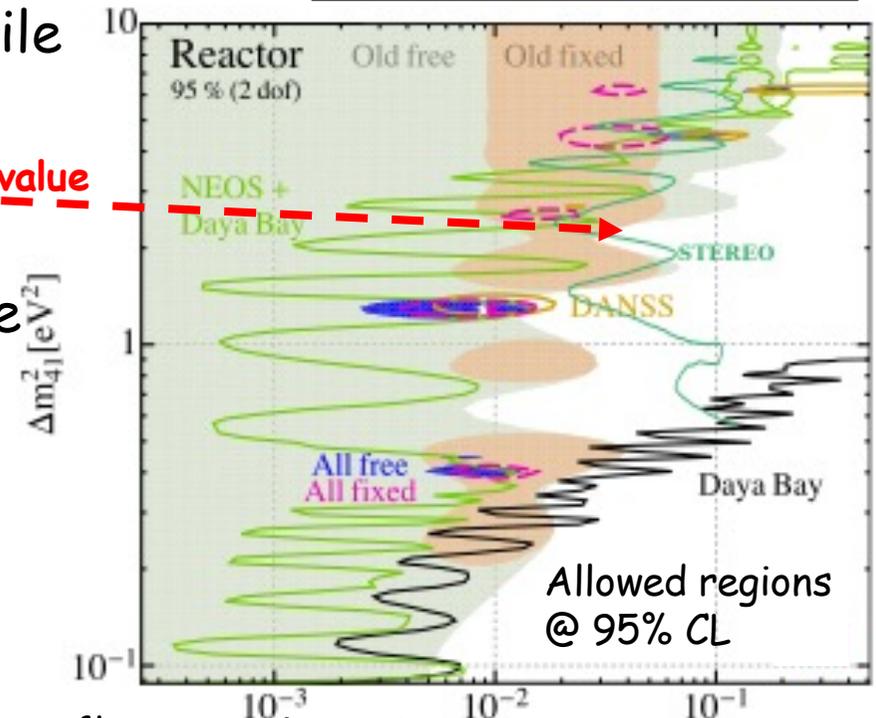
SBL = Short base line
 ATM = Atmospheric ν
 SOL = Solar neutrino



Reactor Antineutrino Anomaly best fit value
 $\Delta m^2 \sim 2.4 eV^2, \sin^2(2\theta) \sim 0.14$

Tension between appearance/disappearance results

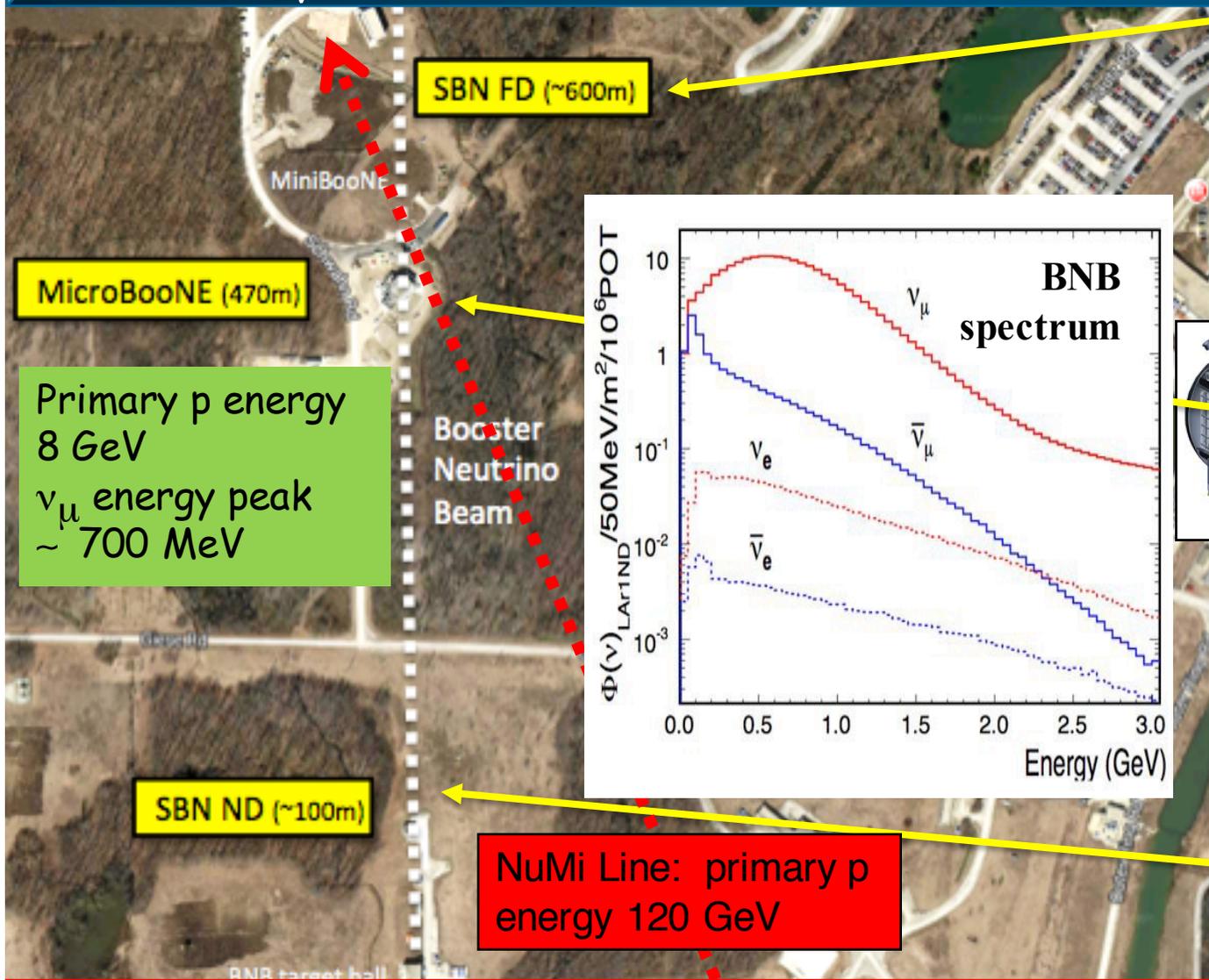
THE EXPERIMENTAL SCENARIO CALLS FOR A DEFINITIVE CLARIFICATION!



SBN 0.8 GeV ν FNAL Booster: 3 shallow-depth LAr-TPCs as definitive answer to sterile ν puzzle
 $L/E_\nu \sim 600 \text{ m} / 800 \text{ MeV} \sim 1 \text{ m/MeV}$



ICARUS T600 (476 t active mass) will operate beneath SBN Far Site Building

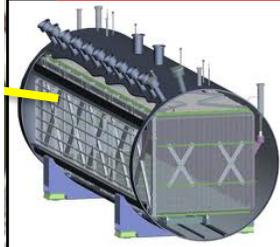
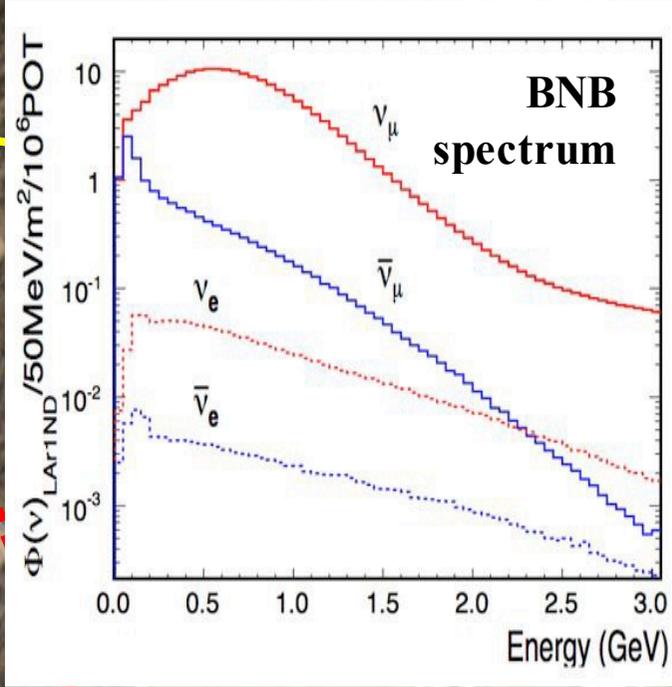


MicroBooNE (470m)

Primary p energy 8 GeV
 ν_μ energy peak $\sim 700 \text{ MeV}$

SBN FD (~600m)

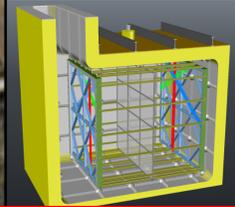
Booster Neutrino Beam



MicroBooNE 89 t active mass

SBN ND (~100m)

NuMI Line: primary p energy 120 GeV

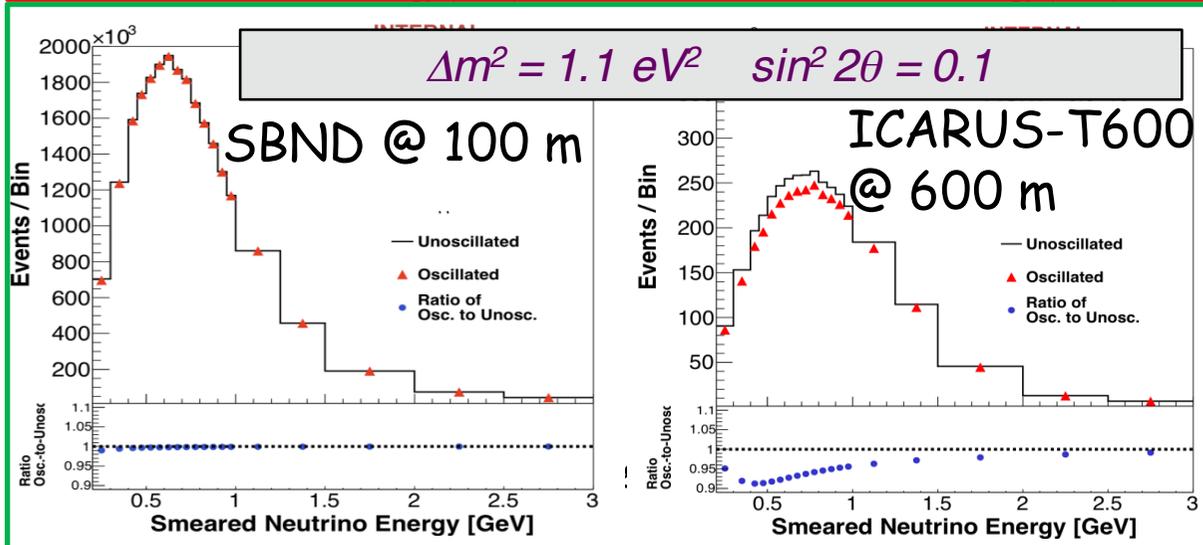
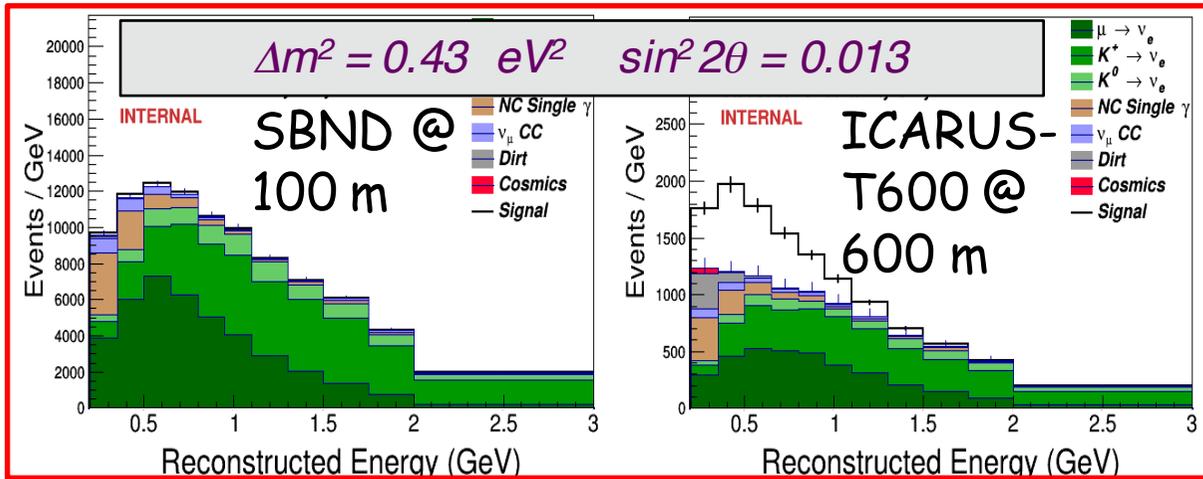


SBND 82 t active mass

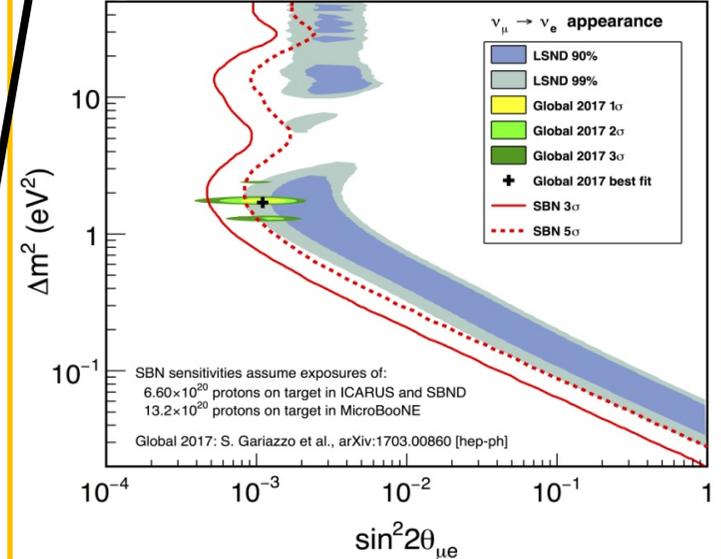
ICARUS T600 will collect also $\sim 2 \text{ GeV } \nu_e$ NuMI Off-Axis: an asset for next LBNF-DUNE (ν cross-section in LAr measurements)

Sensitivity and Sterile neutrino @ SBN

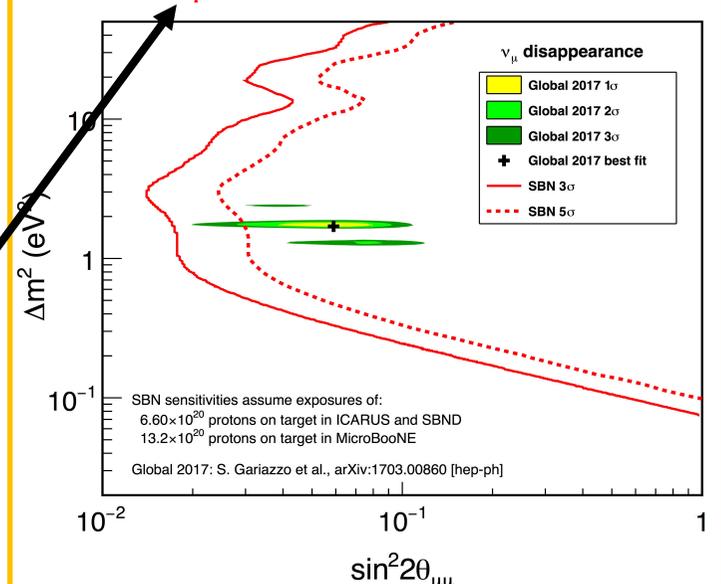
- SBN can clarify the issue by exploiting similar LAr-TPCs at different distances from the target
 - SBND will give the "initial" BNB flux/composition
 - ICARUS, as far detector, shall characterize the ν oscillation parameters.



ν_e appearance: LSND 99% CL region covered at 5σ level

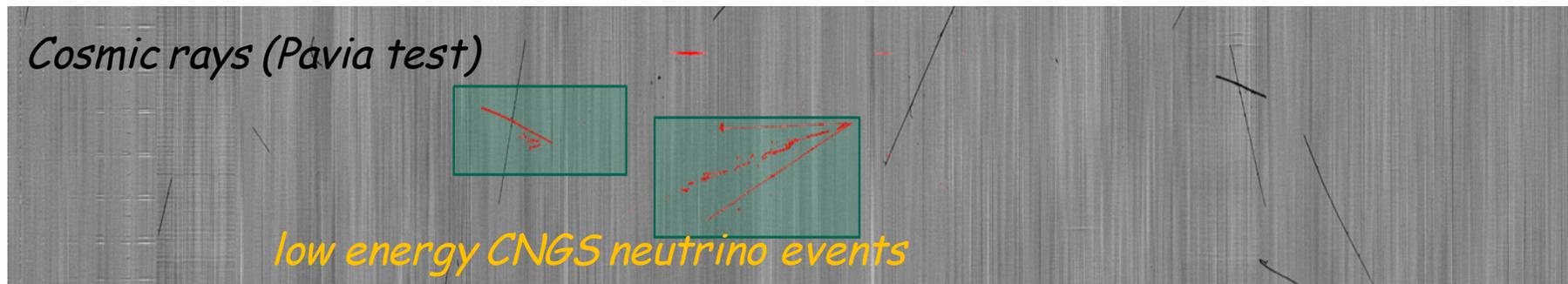


Both plots: 6.6×10^{20} pot (3 years)
 $3-5\sigma$ ν_μ disapp. SBN sensitiv.

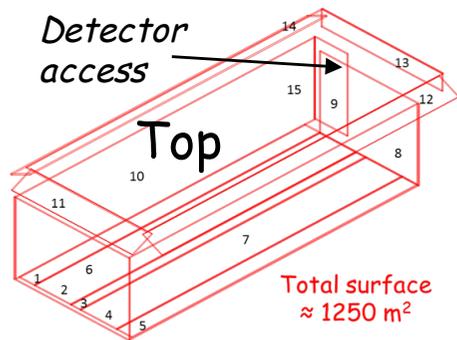


Taking data @ shallow depth: Cosmic Ray Tagger is mandatory

- ICARUS at FNAL is facing a more challenging experimental condition than at LNGS, requiring the recognition of ν interactions amongst 11 KHz of cosmics.
- **A 3 m concrete overburden** will remove contribution from charged hadrons/ γ 's.
- Moreover $\sim 11 \mu$ tracks will occur per triggering event in 1 ms TPC drift readout: associated γ 's represent a serious background source for ν_e search since e 's produced via Compton scatt./ pair prod. can mimic a genuine ν_e CC.



- **Rejecting cosmic background**, i.e. reconstructing the triggering events, requires to precisely know the time of each track in the TPC image:
 - A much improved **light detection system**, with $\sim ns$ time resolution;
 - An external **cosmic ray tagger (CRT)** to detect incoming particles and measure their direction of propagation by time-of-flight:
 - ✓ Scintillating bars surrounding T600 (aim: 98% coverage) equipped with optical fibers to convey light to SiPM arrays.
 - ✓ Top coverage under INFN/ CERN responsibility. FNAL is recovering modules by MINOS/Double Chooz for side/bottom.



ICARUS T600 Overhauling at CERN (WA104/NP01)

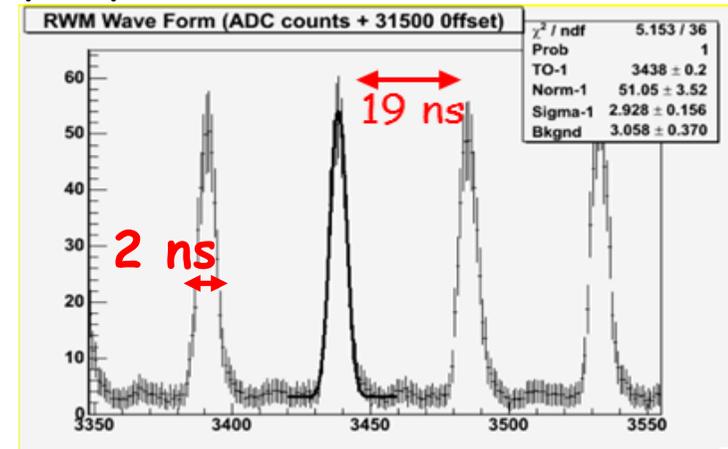
- ICARUS T600 detector **underwent an intensive overhauling at CERN** in 2015/17 in the framework of CERN Neutrino Platform (**WA104/NP01 project**) before being shipped to FNAL:
 - **New cold vessels**, purely passive insulation;
 - **Renovated cryogenic / LAr purification** equipment;
 - **Flattening of TPC cathode**: few mm planarity;
 - **Upgrade of light collection system**;
 - **New higher performance TPC read-out electronics** with both analogue/digital parts integrated in a single board in direct communication with ad-hoc signal feedthrough flanges.



The light collection system 1/2

In ICARUS, light collection is used to:

- Identify precisely the time of occurrence (T_0) of each interaction;
- Identify the event topology for fast selection purposes;
- Generate a trigger signal to enable the event read-out by combining:
 - Pattern/majority of hit PMT signals
 - BNB/NuMI bunched beam spill
 - Veto from CRT

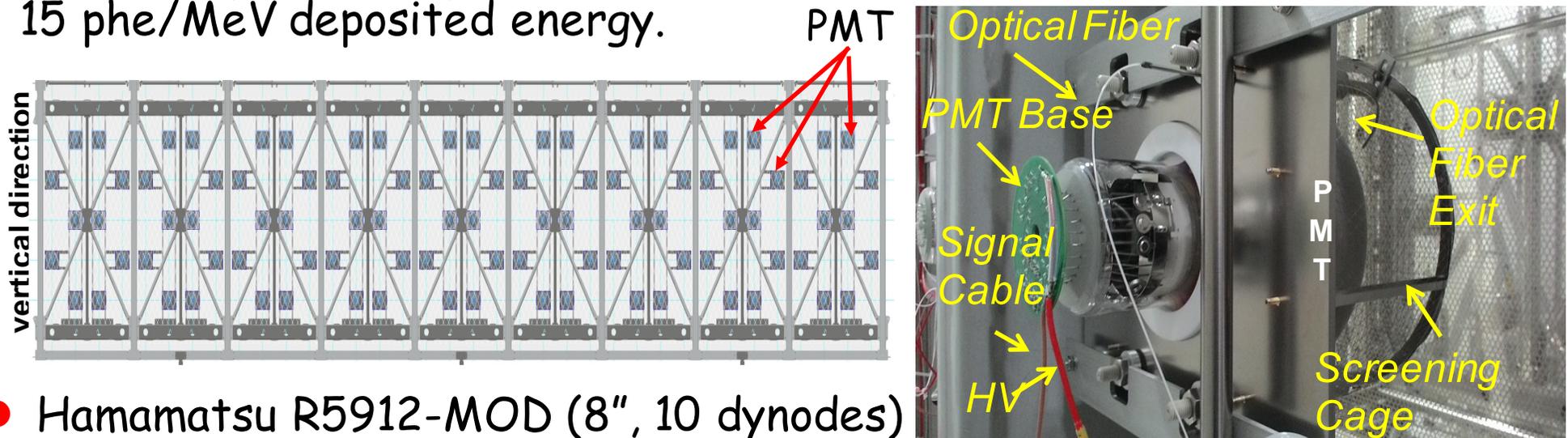


The light collection system is based on 360 PMT's, 90/chamber, to have:

1. High detection coverage, to be sensitive to the lowest-expected neutrino energy deposition in the TPC (approximately 100 MeV), also using the light fast-component only;
2. High detection granularity (\sim mm), longitudinal resolution is better than 0.5 m (effective Q.E. = 5%).
3. Fast response time/ high time resolution (\approx 1 ns), with a PMT timing calibration provided by a laser system (Hamamatsu PLP10, $\lambda \sim$ 450 nm, FWHM < 100 ps, peak power \sim 400 mW) + 50 μ m optical fiber. *Slide# : 13*

The light collection system 2/2

- **90 PMT's per TPC layout:** 5% cathode coverage area, allowing to collect 15 phe/MeV deposited energy.



- Hamamatsu R5912-MOD (8", 10 dynodes) are rated for cryogenic temperature, as they feature a cathode with platinum under-layer.
- Each PMT is enclosed in a wire screening cage to prevent induction of PMT pulses on the facing TPC wires
- PMT sand blasted glass windows coated by $\sim 200 \mu\text{g}/\text{cm}^2$ of Tetra-Phenyl-Butadiene (TPB) wavelength shifter to detect the $\lambda = 128 \text{ nm}$ scintillation light in LAr.

A clear cosmic μ 's identification will be provided by genetic algorithms ($\sim 2\%$ expected residual misidentification).

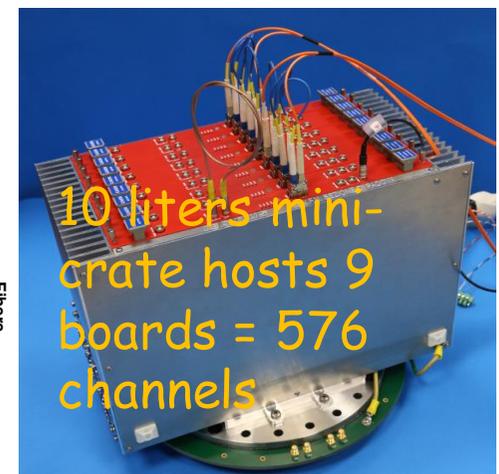
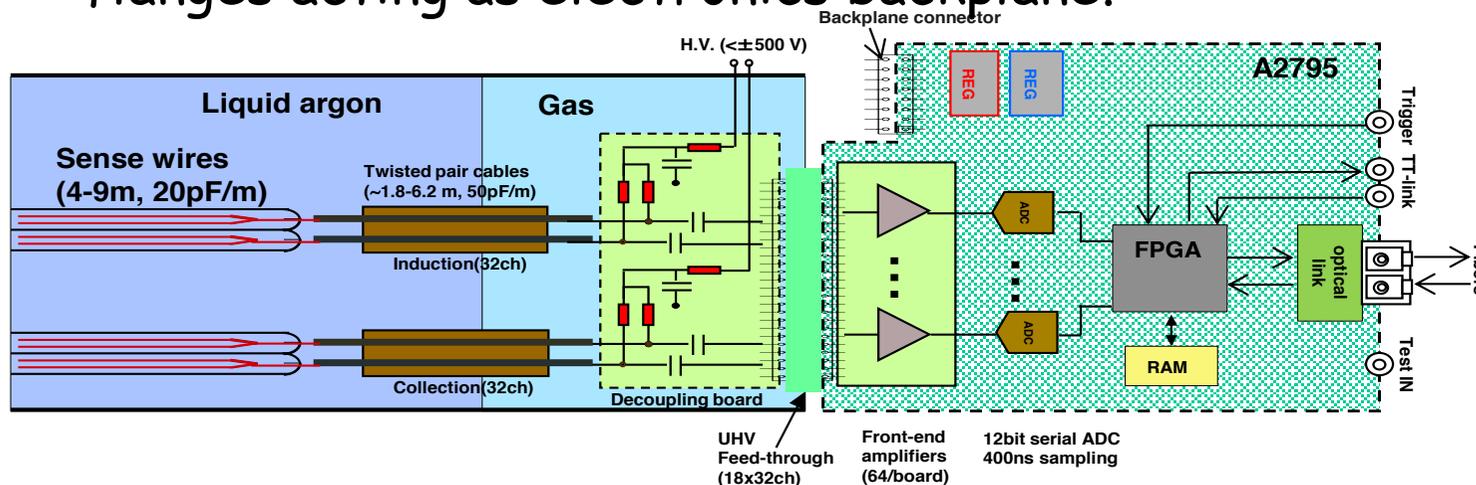
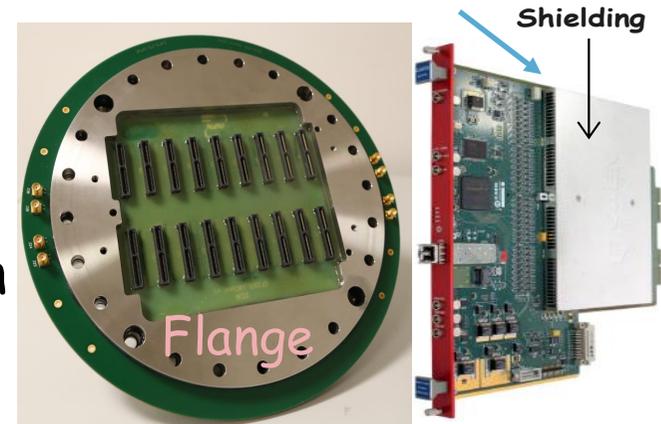
The new TPC read-out electronics

- ICARUS electronics at LNGS was based on analogue low noise "warm" front-end amplifier, a multiplexed 12-bit 2.5 MHz ADC and a digital VME module for local storage, data compression, trigger information:
 - $S/N \sim 9$ in Collection, ~ 0.7 mm single hit resolution, resulting in a precise spatial event reconstr. and μ momentum measurement by MCS.

- Improvements concern:

- Serial 12 bits ADC, one per ch, 400 ns sampling synchronous on the whole detector;
- Serial bus architecture with Gbit/s optical links to increase the bandwidth (10 MHz);
- Both analogue/digital electronics are housed in a single board inserted in a new mini-crate directly installed on ad-hoc signal feedthrough flanges acting as electronics backplane.

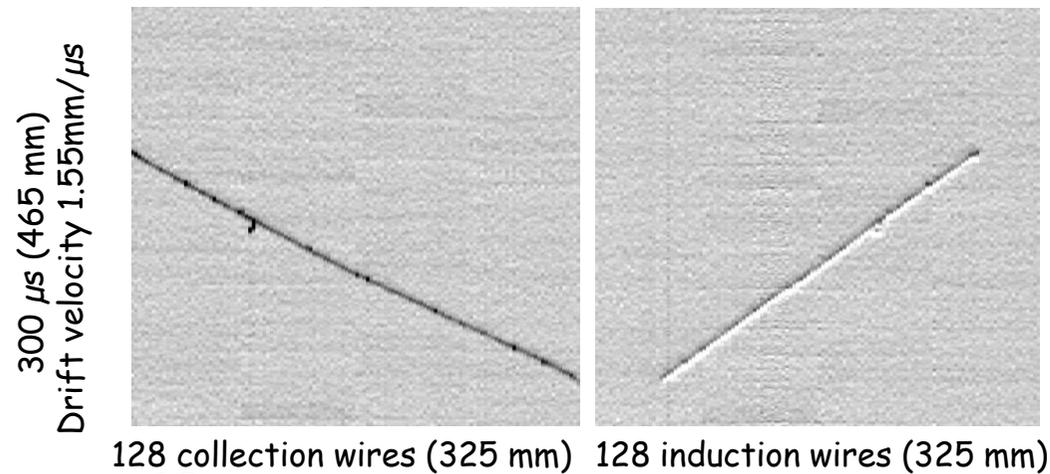
CAEN A2795 board, 64 chs



Improved front-end electronics for T600

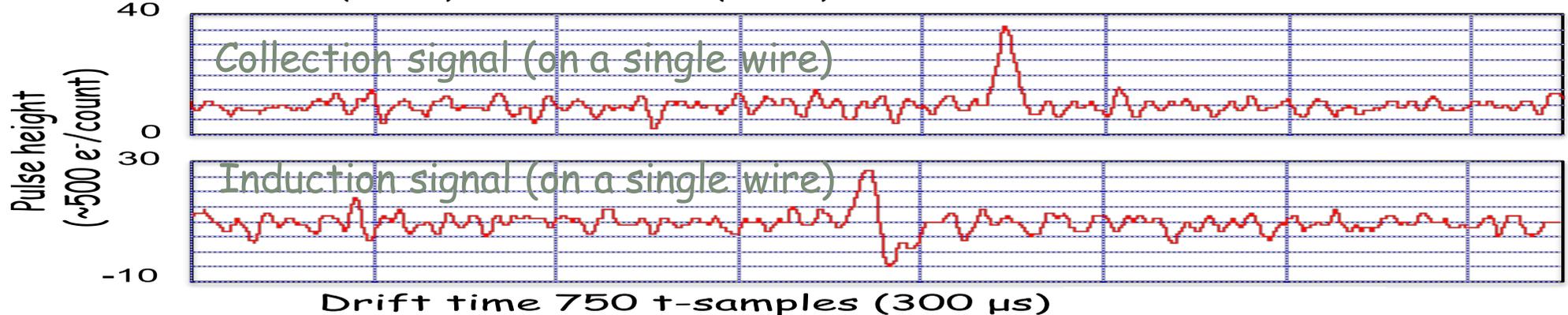
- Adopted improvements in the analogue front-end
 - A faster shaping time $\sim 1.5 \mu\text{s}$ of analogue signals to match electron transit time in wire plane spacing;
 - A drastic reduction of undershoot in the preamp response as well as of the low frequency noise while maintaining a same or better S/N;
 - Same preamplifier for induction and collection planes, so induction view can be used for dE/dx measurement as well.

A better event reconstruction is then possible



Example of single m.i.p. track:

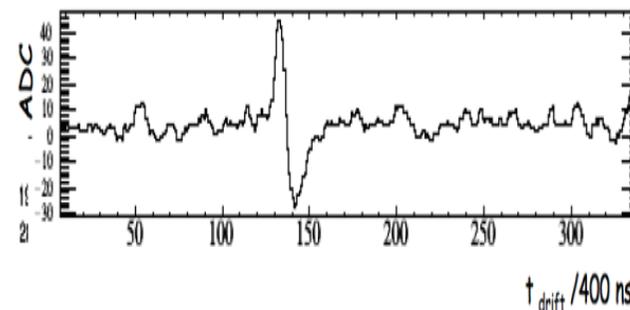
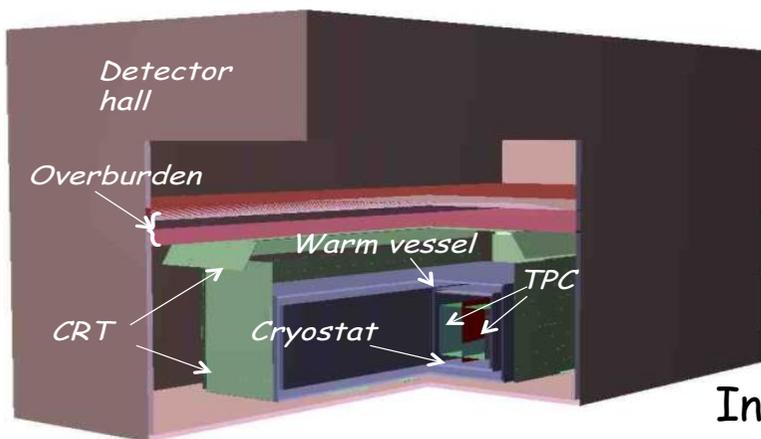
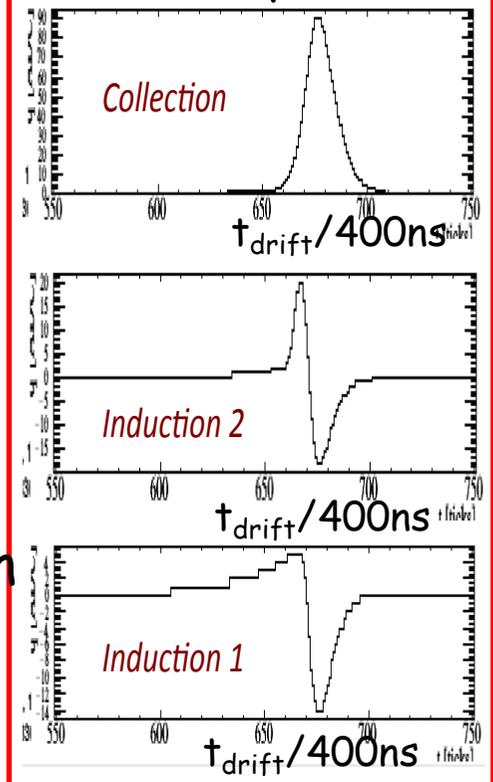
- Same ~ 2 ADC counts ($\sim 1000 e^-$) noise for both Collect. & Induct.;
- Unipolar Collection signal: ~ 25 ADC counts;
- No filter applied to any data.



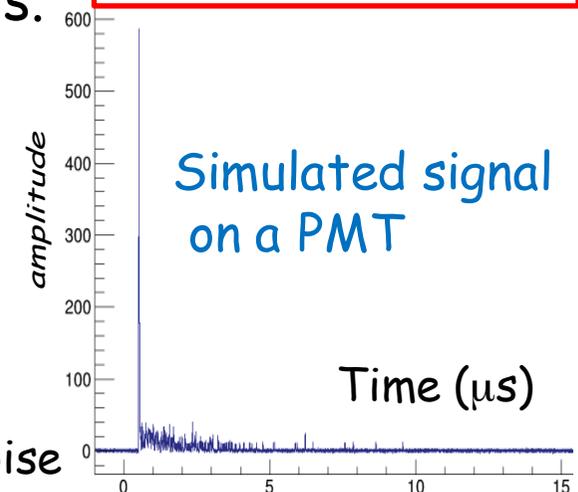
ICARUS @ SBN software status 1/2

- ICARUS adopts a framework (LarSoft) common to SBN experiments, providing tools to simulate, reconstruct/identify events (cosmic μ 's, e.m. showers, neutrinos, ...).
- Experimental geometry setup is described in LarSoft.
- Some reconstruction/analysis tools (particle ID, 3D and shower axis reconstructions, momentum from Multiple Coulomb Scattering...) are inherited from previous LNGS ICARUS software and are being ported to LarSoft.
- **Scintillation light** in LAr is parameterized to simulate PMT signals for any MC event, to study event recognition with genetic algorithms, and for trigger simulation.
- **MC simulations** include new wire electronic response/realistic noise, as well as PMT scintillation light signals.

Typical wire shape for 1 m.i.p. μ



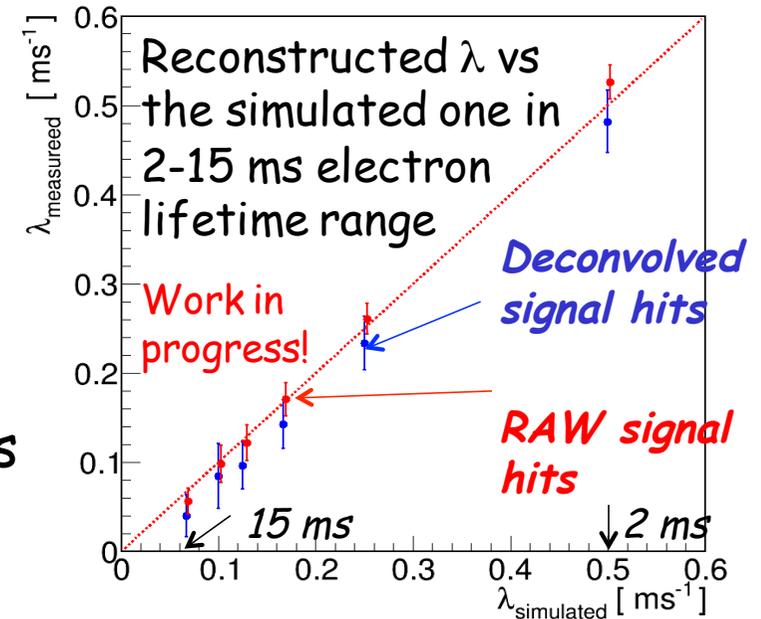
Ind.2 signal simulation, including noise



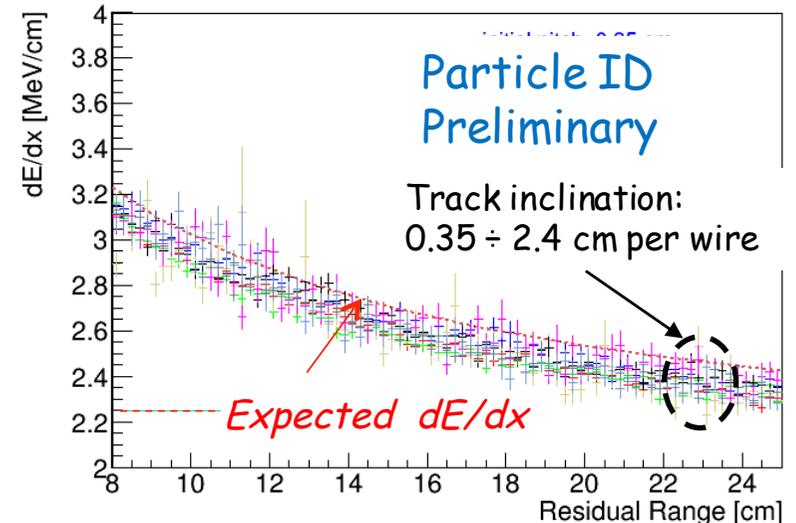
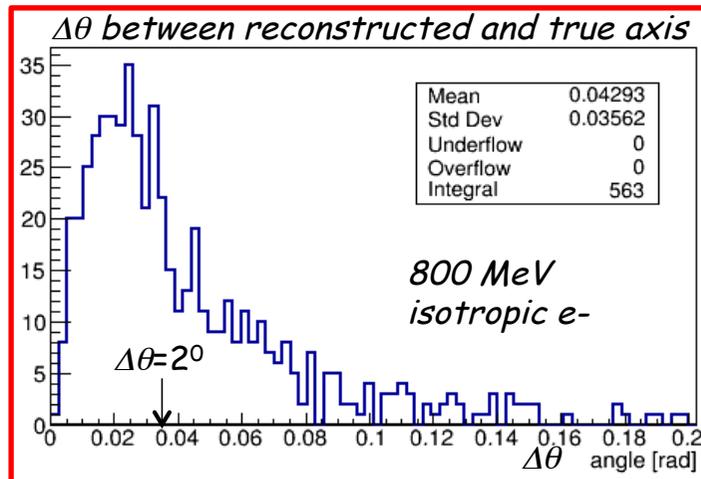
Simulated signal on a PMT

Software status 2/2

- Examples of advanced tools (already ported in Larsoft!):
 - **LAr purity** $\lambda=1/\tau_{ele}$ (τ_{ele} : electron lifetime) measurement from charge attenuation of cosmic μ 's tracks along the drift
 - ➔ Track selection at shallow depth difficult due to crowded events and lower energy μ 's
 - **Particle ID**, based on dE/dx vs range
 - **Electromagnetic shower axis identification**
 - ➔ Provides 3D reconstruction of shower

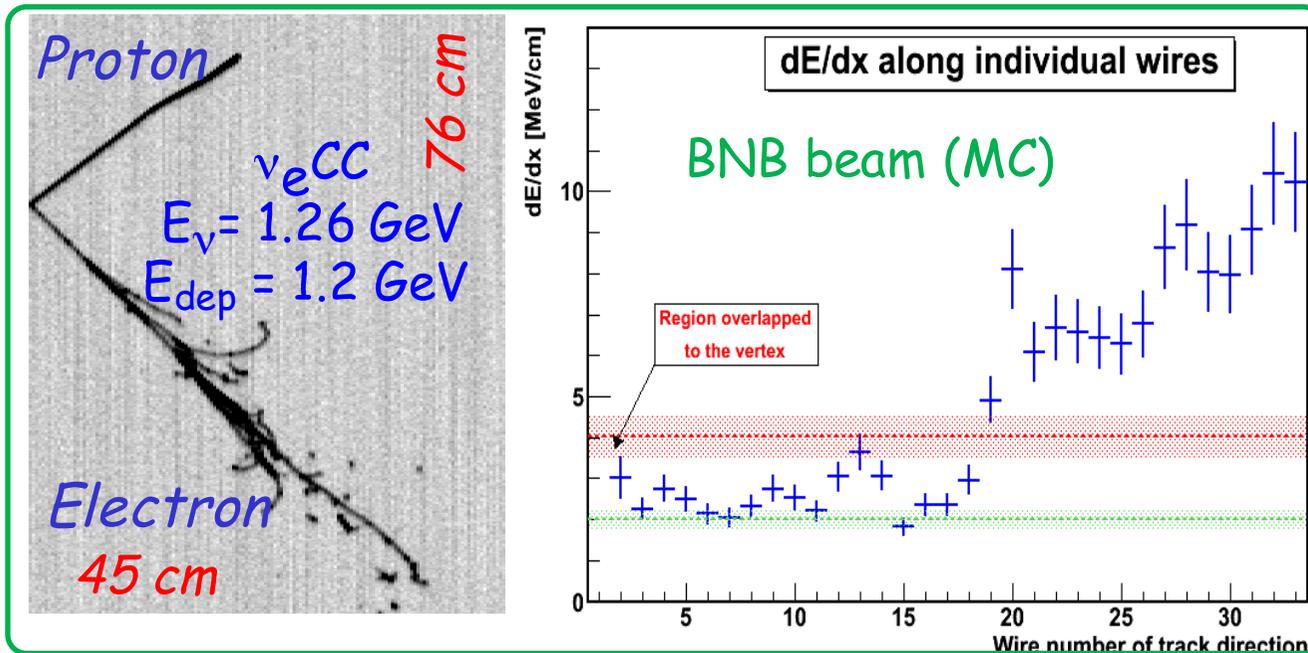


dE/dx vs range for μ 's



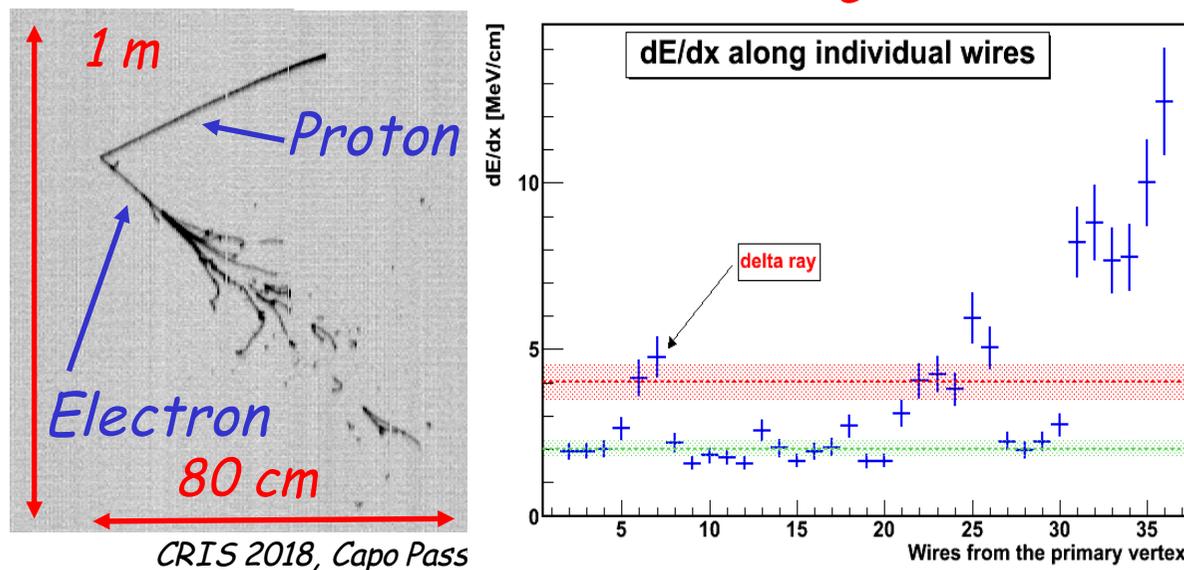
Software is mature enough to realistically simulate events with BNB beam

BNB (MC) and real atmospheric ν_e CC events comparison



- MC SBN ν_e CC interactions (on the left) are very alike to typical atmospheric ν_e CC events @ LNGS (below)
- Similar results hold for ν_μ CC interactions

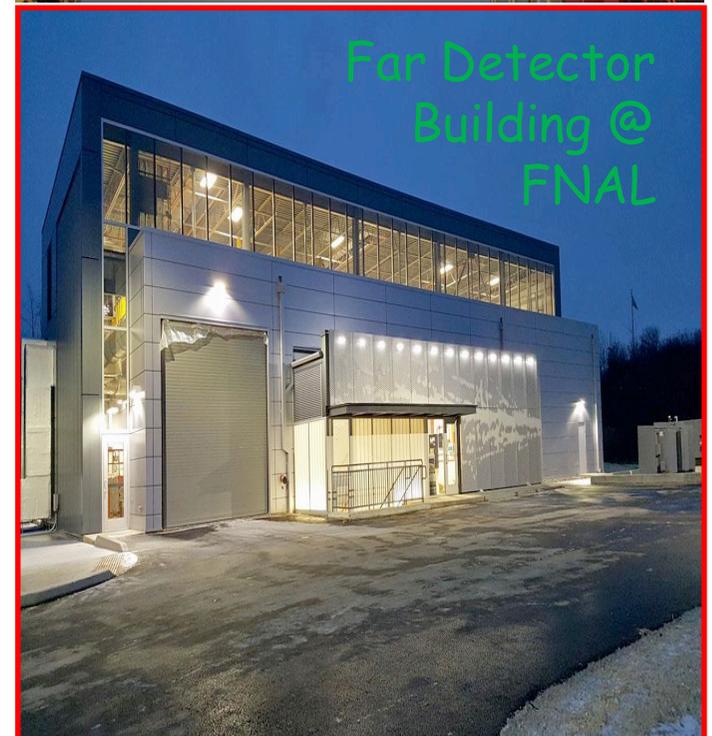
LNGS ν_e ATMOSPHERIC EVENT



- Quasi-elastic ν_e CC $E_{\text{Dep}} = 0.9$ GeV.
- Proton identified by dE/dx.
- Electron identified by single m.i.p. before showering

Current status of ICARUS installation @ FNAL 1/3

- The lower part of the warm vessel had been assembled inside the Far Detector (FD) building at FNAL by summer 2017.
- 14 modules of the bottom CRT (200 m² total area) have been installed by summer 2017
 - Each module (4m x 1.6m x 3.2cm) consists of 2 layers of 32 parallel scintillator strips (5 cm width), read out by a 64-pixel multi-anode PMT
- Assembly of cold shields completed by May 2018
- Before moving cold vessels in the final position (expected by middle July 2018), the main doors of the vessels have to be sealed. A helium leak tests **is being performed** and should be completed by this week (~mid June 2018)
- Installation of detector supports is in progress.



Current status of ICARUS installation @ FNAL 2/3

- Connectivity tests have to be performed as part of the installation of chimneys (cold vessel to outside of thermal insulation connections), by beginning August 2018.
- Top part of cold shield will be installed and tested, followed by installation of top part of warm vessel (August 2018).
- Side CRT (double layer, $\sim 1000\text{m}^2$ total) should start read-out board production in fall
 - Each module ($8\text{m} \times 80.5\text{cm} \times 1\text{cm}$) has 20 parallel scintillator strips, SiPM-based readout.
- From Sept. 2018, activities on top of detector will start (cryo, purification and vacuum systems, ext. cabling, feedthrough flanges with decoupling boards, optical fibers, read-out,...)
- Vacuum pumping should start by Nov 2018 and last until ready to start cool-down (Jan 2019), which is bound to "clearance to start operation" issued by Fermilab.



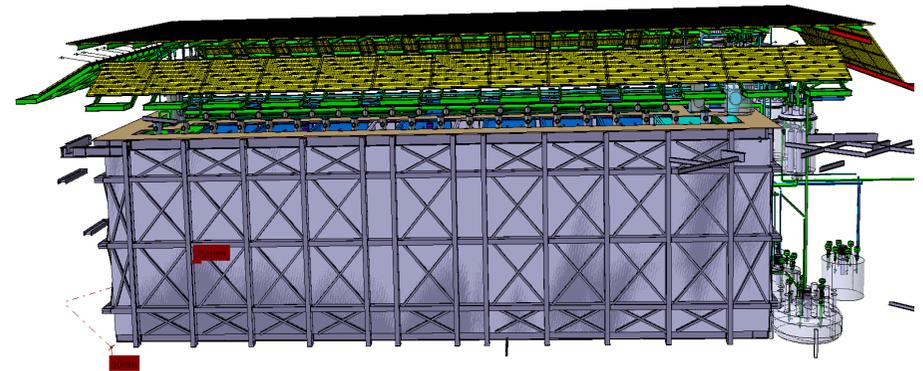
Door welding & vacuum tests



Cold shields installed by mid June

Current status of ICARUS installation @ FNAL 3/3

- Meanwhile:
 - Testing of DAQ, noises, PMTs, Slow controls, etc. should continue.
 - Cryo systems installation and pre-commissioning will be completed.
- Detector **commissioning** consists of three phases:
 - Cryogenic commissioning (2 months in total): Cooling (15 days), Filling (15 days), Purification (1 month), Stabilization (1 month).
 - TPC and PMT system commissioning (2 months in total): HV for the E field drift, PMT's supply, calibrations, DAQ & trigger commissioning.
 - CRT commissioning:
 - ➔ Side+bottom CRT can be installed and commissioned in parallel with the activities for the completion of cryo, TPC and PMT system commissioning.
 - ➔ Top CRT (~400m², 125 modules, 64 scint. strips per module, SiPM-based readout, 2 fibers/strip) "barn-style" installation should start in **2019**.



Conclusions

- LAr-TPC detection technique taken to full maturity with ICARUS-T600.
- ICARUS completed in 2013 a successful continuous 3-year run at LNGS exposed to CNGS neutrinos and cosmic rays, and performed a sensitive search for a potential LSND-like ν_e excess defining a narrow region at $(\Delta m^2, \sin^2 2\theta) \sim (1 \text{ eV}^2, 0.005)$. Result confirmed by OPERA.
- ICARUS underwent a major overhauling at CERN and was transported to FNAL to be exposed to Booster and NuMI neutrinos, to provide a clarification of the sterile neutrino issue, both in appearance and disappearance modes (SBN experiment).
- ICARUS installation in the Far Site building @ FNAL is in progress
 - Done: lower part of warm vessel, 14 modules of bottom CRT in summer 2017, cold shields: assembly in May 2018, installation by mid Jun 2018.
 - Now: Installation of detector supports.
 - Next steps:
 - ➔ **Installation** of: chimneys (Aug 2018), top of cold shield, warm vessel (Aug 2018); activities on top (cryo, purification, vacuum systems...) (from Sep 2018); vacuum pumping should start by Oct-Nov 2018.
 - ➔ Detector **commissioning** while waiting for clearance by FNAL (by Feb 2019) to start cool-down and filling. Then **data taking for physics!**



Thank you!