

LISA OPTICAL METROLOGY CHALLENGES

Daniel Penkert^{1,2}, Oliver Gerberding^{1,2}, Gerhard Heinzel^{1,2} and the AEI LISA Team

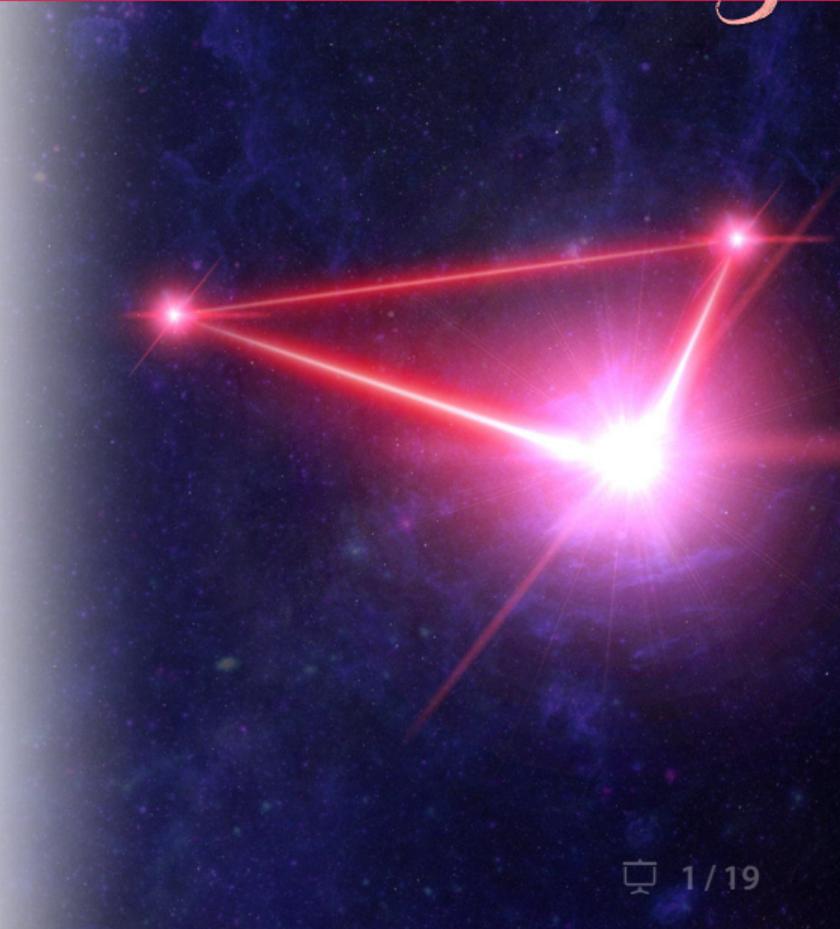
¹ Max Planck Institute for Gravitational Physics (Albert Einstein Institute) ² Leibniz Universität Hannover ... Gravitational-Waves Science & Technology Symposium, March 1-2, 2018





Outline of this talk

1. A Brief Introduction to LISA





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2. The Hexagon Interferometer





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3. The Three-Backlink Experiment





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4. Other Activities





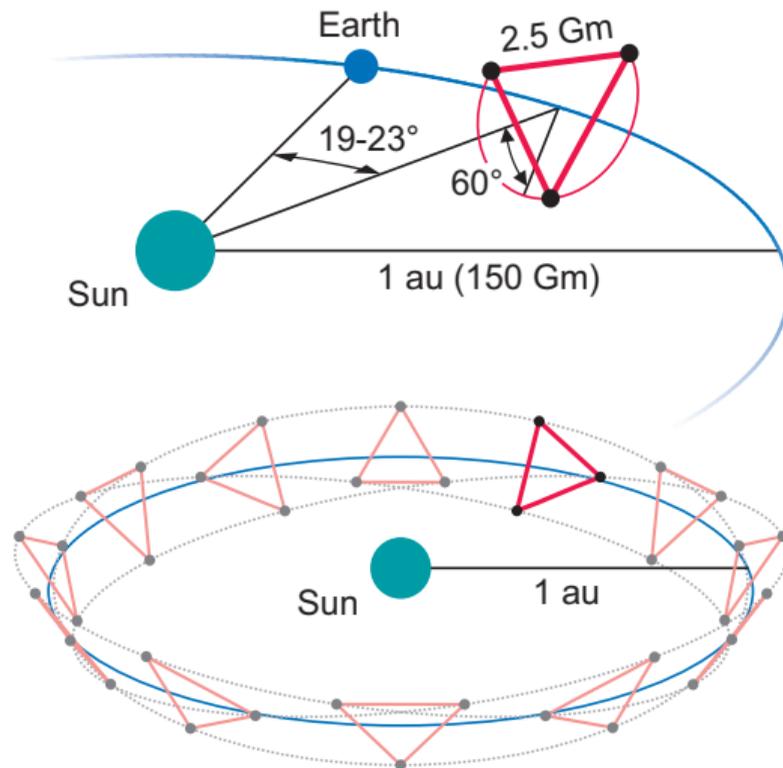
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1. A Brief Introduction to LISA
2. The Hexagon Interferometer
3. The Three-Backlink Experiment
4. Other Activities
5. Summery





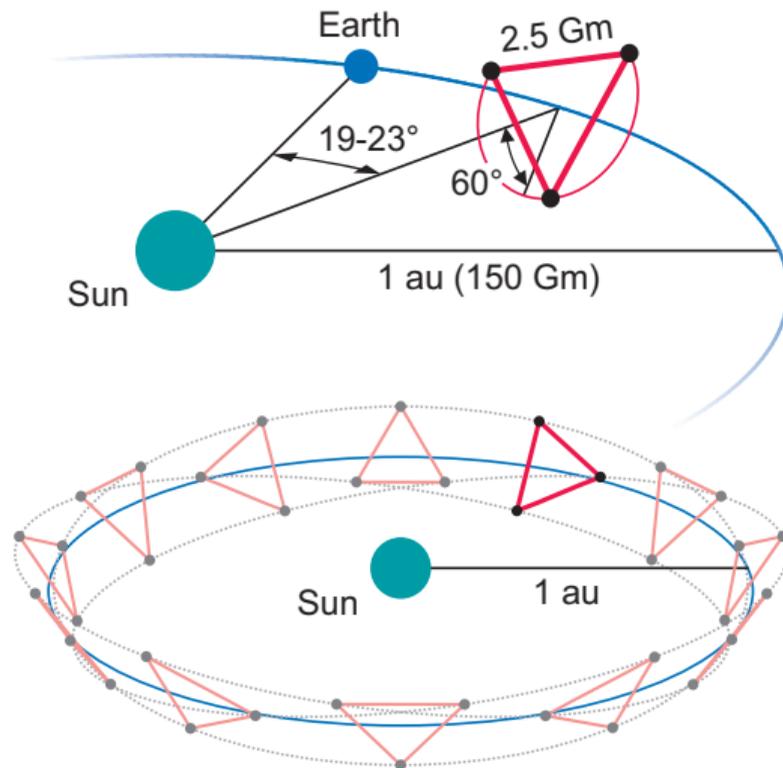
Laser Interferometer Space Antenna





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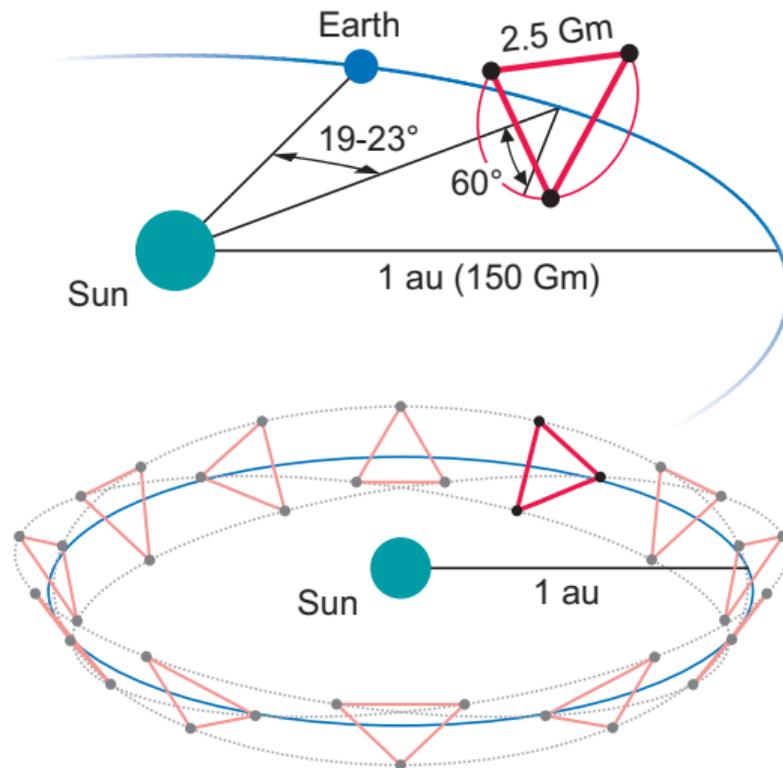
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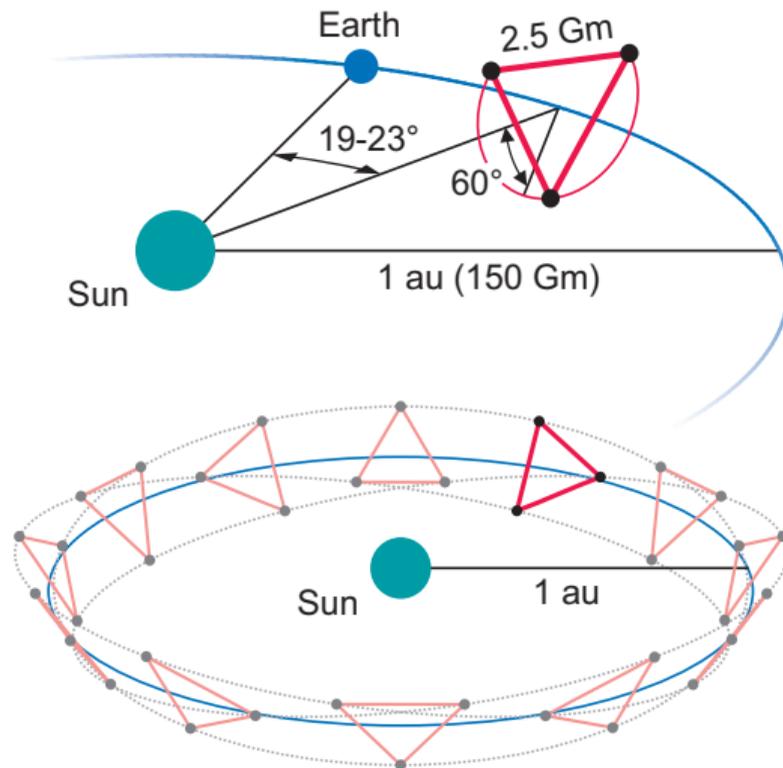
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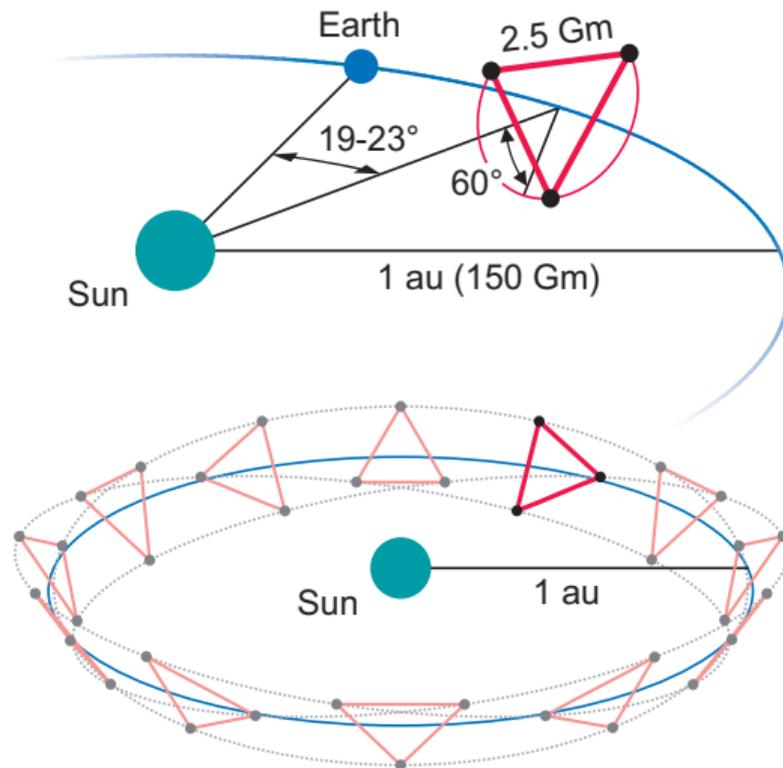
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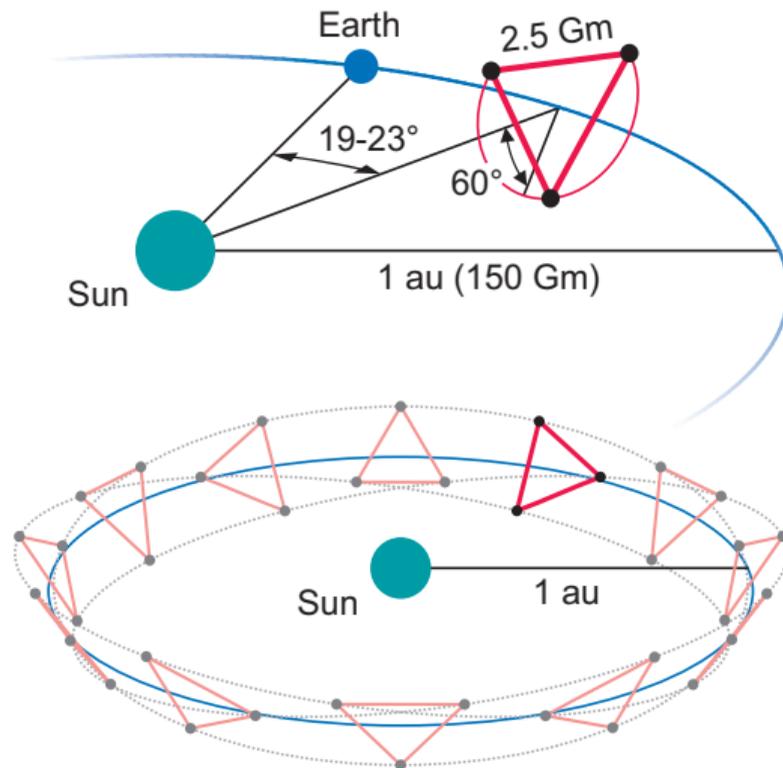
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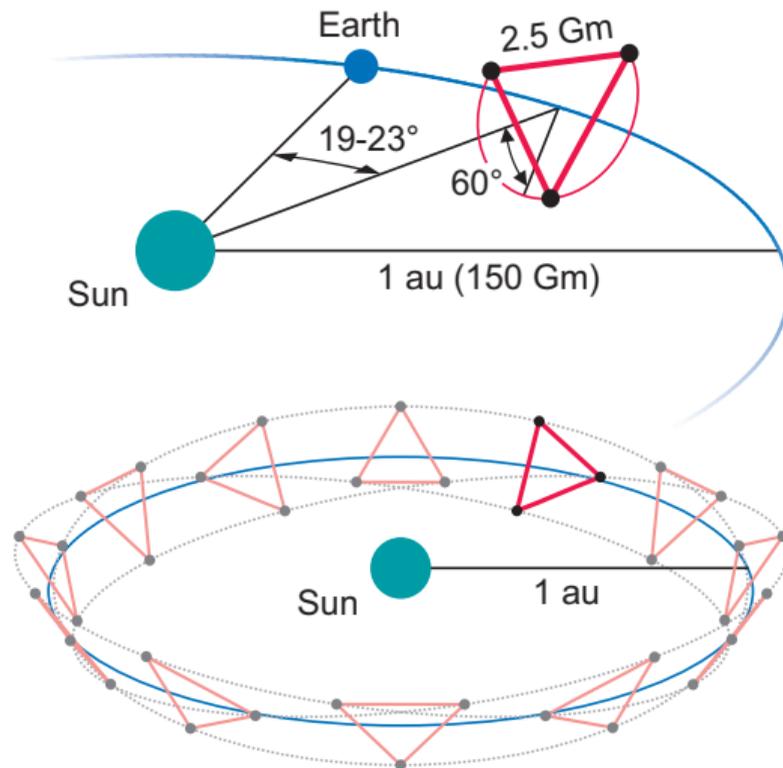
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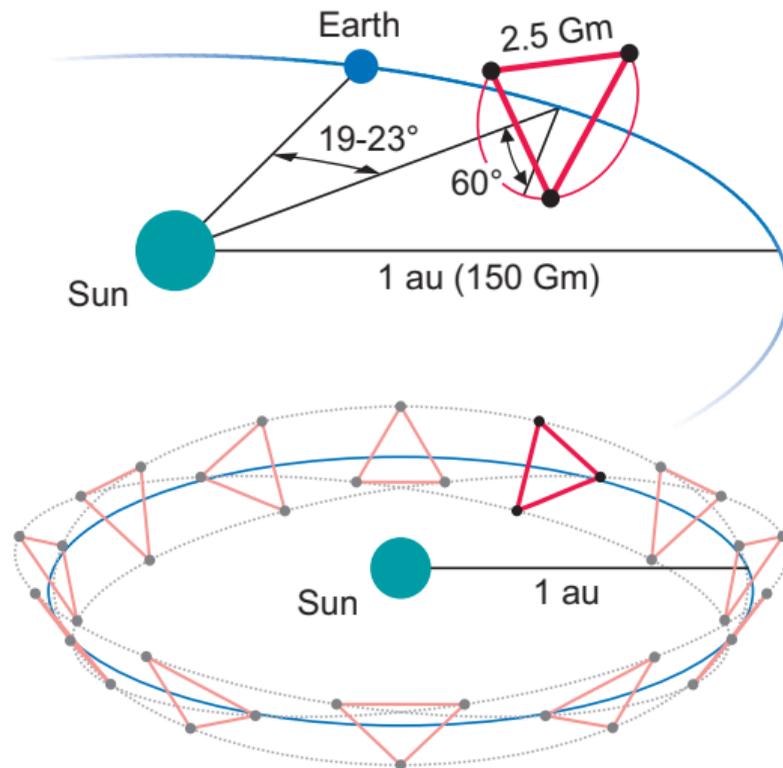
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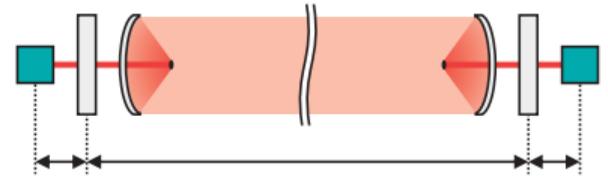
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- optical readout of differential path lengths between free-floating test masses (TMs)





Laser Interferometer Space Antenna (cont.)

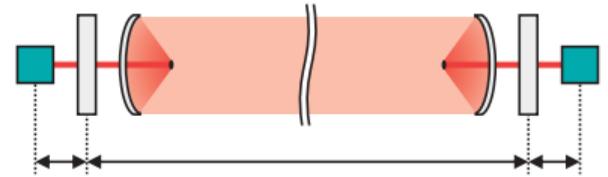
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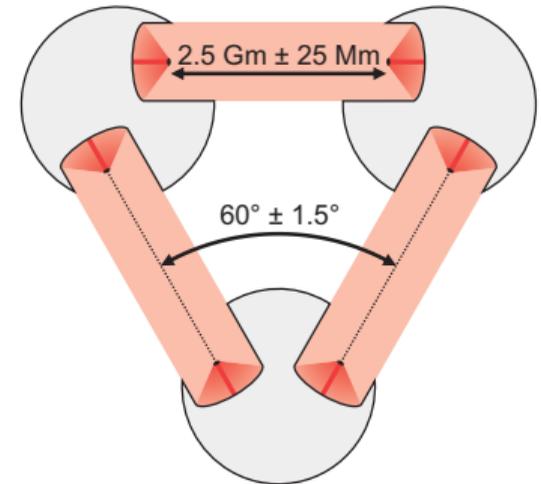
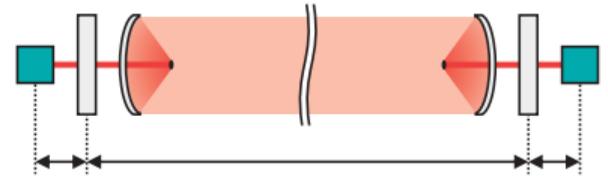
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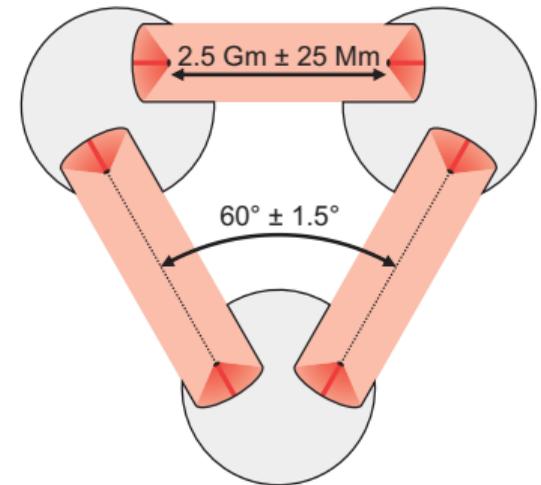
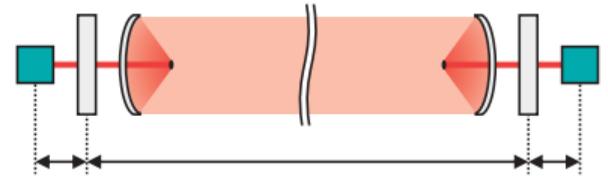
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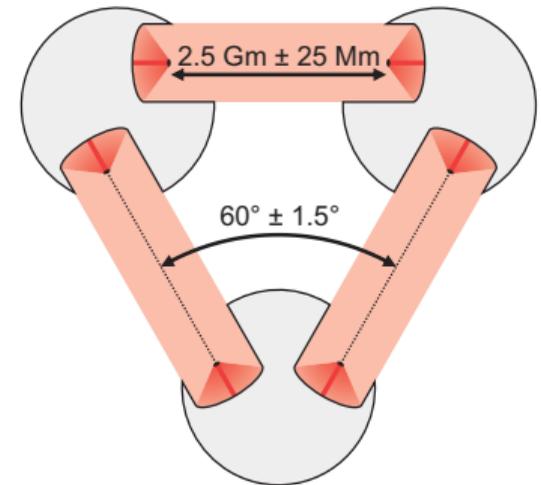
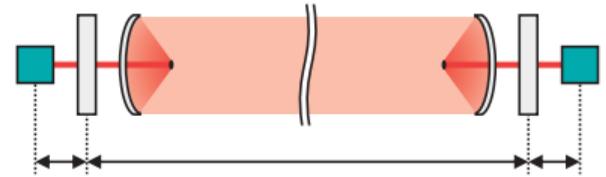
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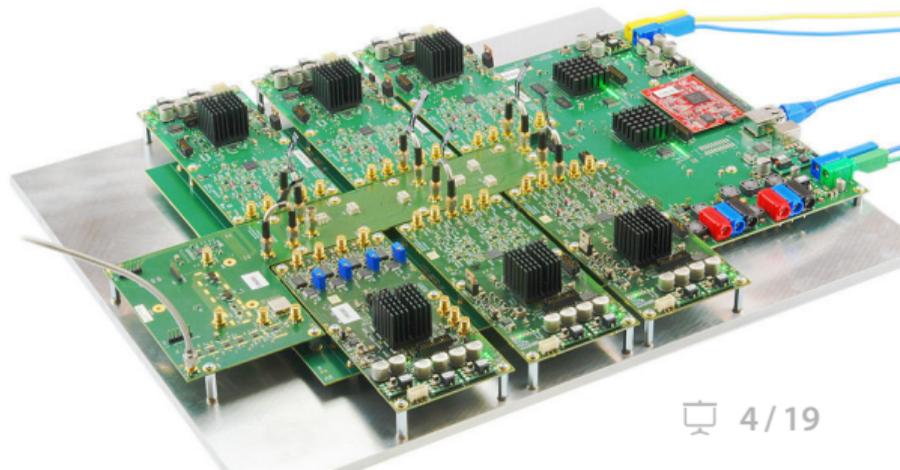
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- many aspects of inter-S/C interferometry will be tested within the GRACE-FO mission (scheduled for launch on April 27, 2018)





Primary beat-note readout and phase extraction

Image: FPGA-based LISA phasemeter prototype
(co-developed in the scope of an ESA technology
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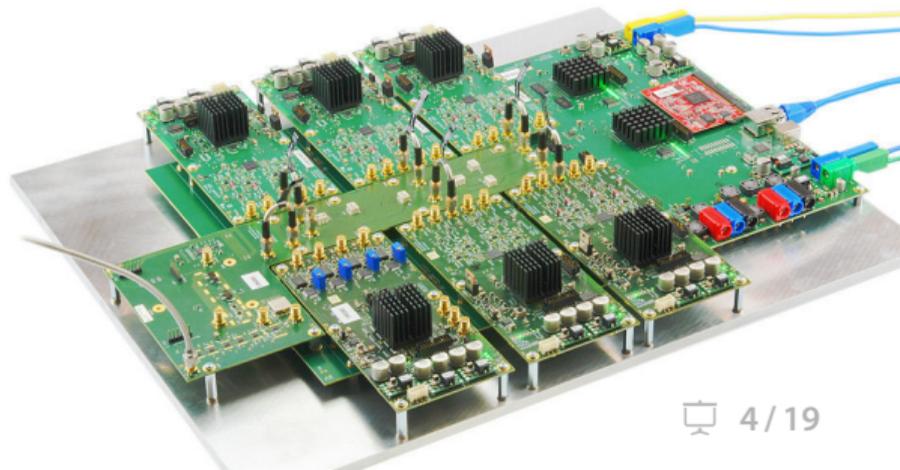




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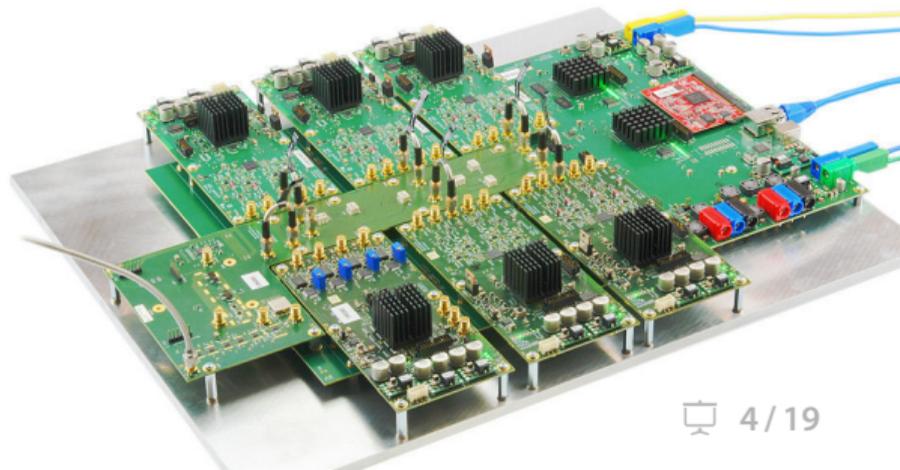




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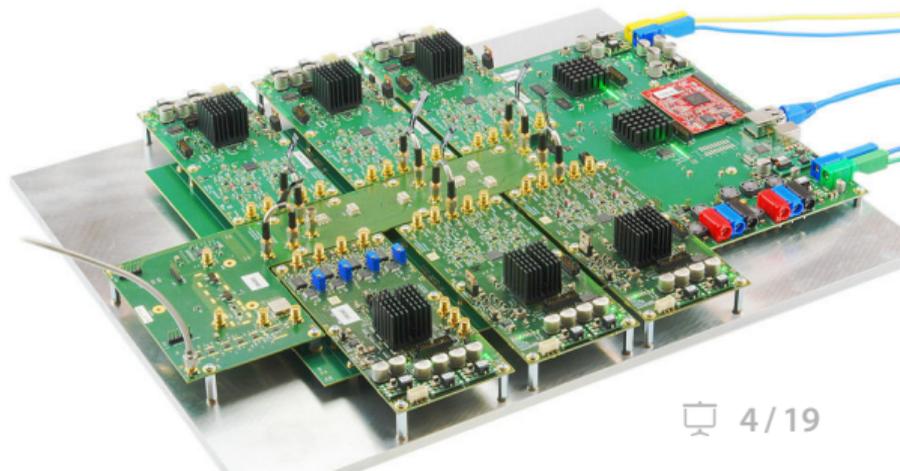




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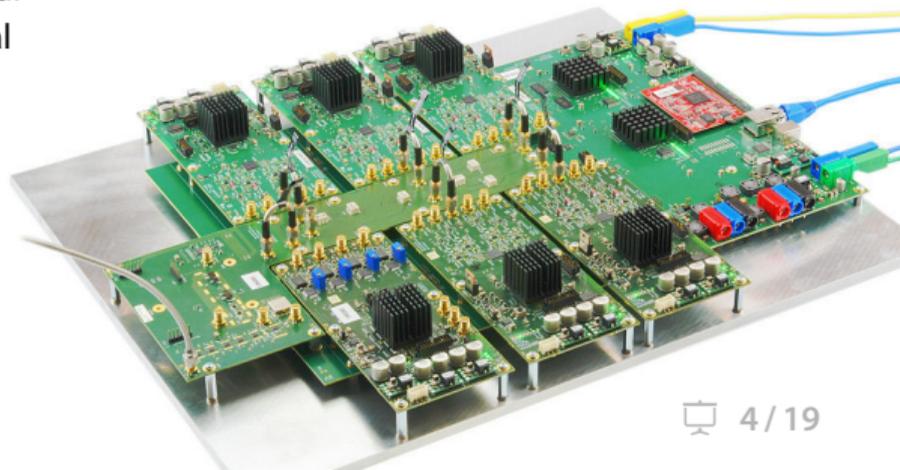




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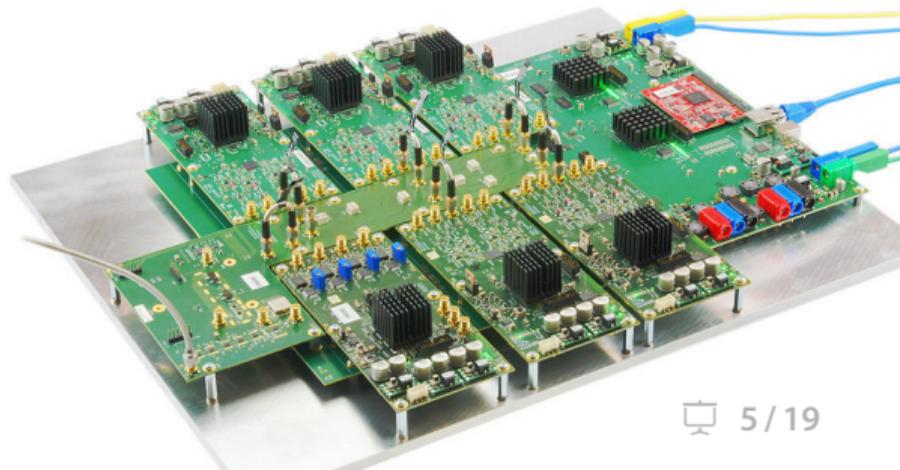
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- cascaded arm interferometry approach \Rightarrow no optical cancellation of laser frequency noise in the sectional measurements \Rightarrow phase readout with high dynamic range (up to 10 orders of magnitude)

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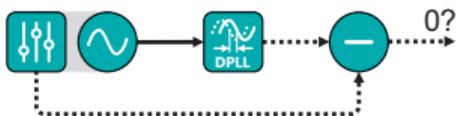


Problem: How can we test a phase readout system for this unusually high level of precision?

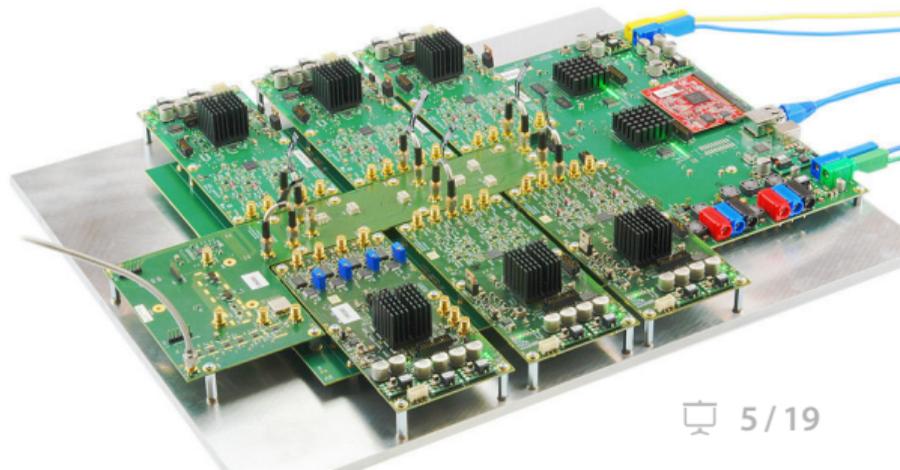




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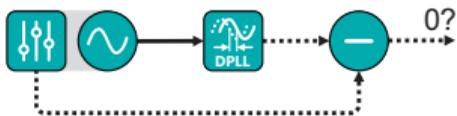


Absolute signal test: requires a signal generator with an extremely high phase fidelity (not readily available)

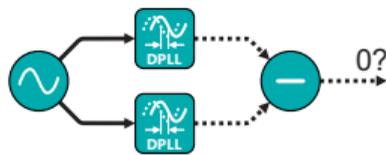




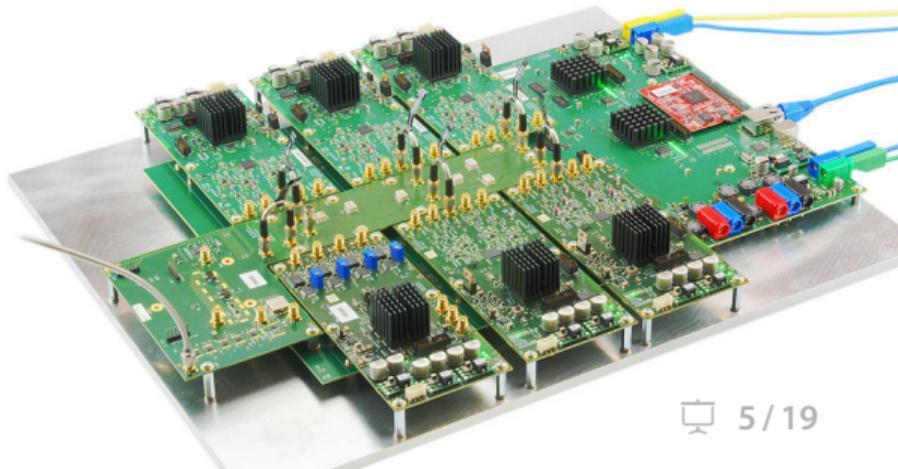
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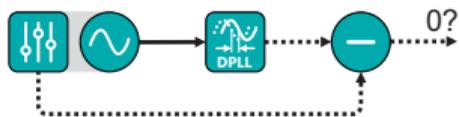


Split signal test: oblivious to common-mode effects like phase nonlinearity or signal-triggered cycle slips

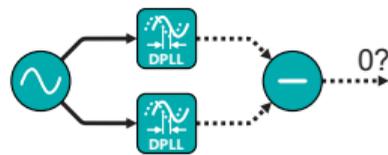




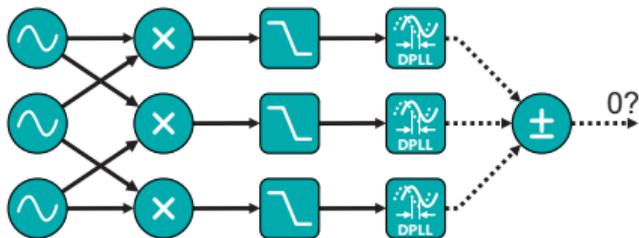
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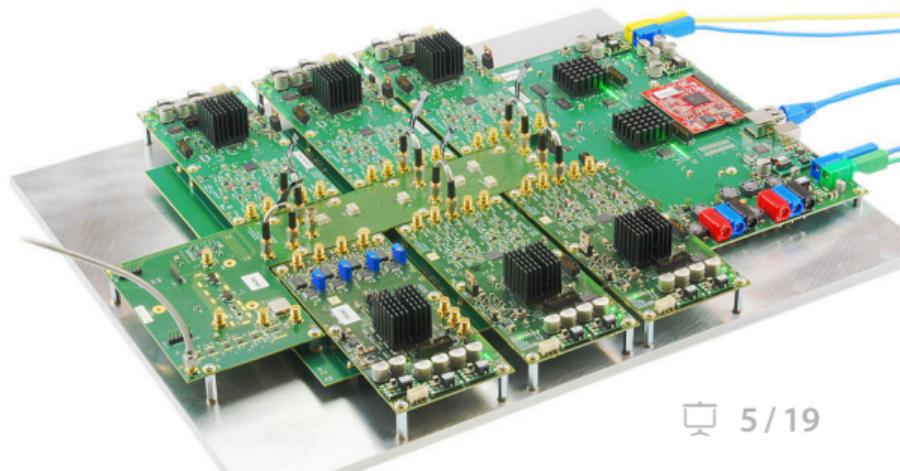
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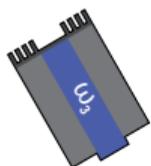


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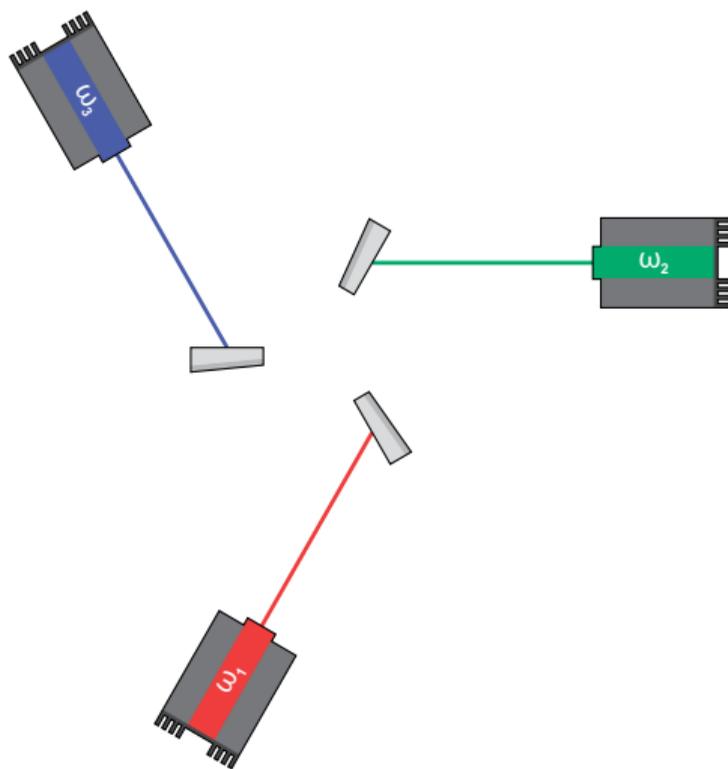


Three-signal test: generates three signals, the sum of which should vanish if measured correctly \Rightarrow Hexagon

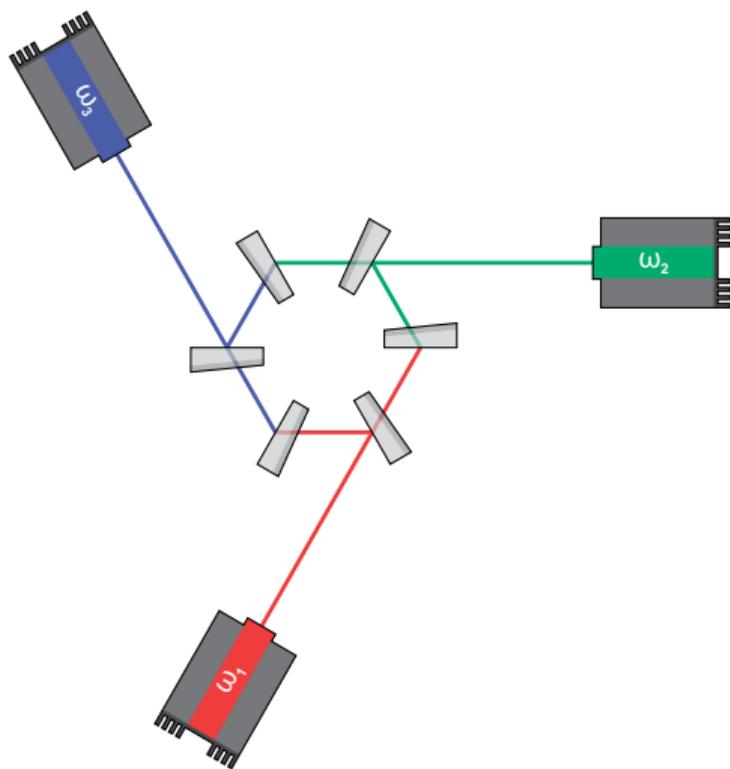




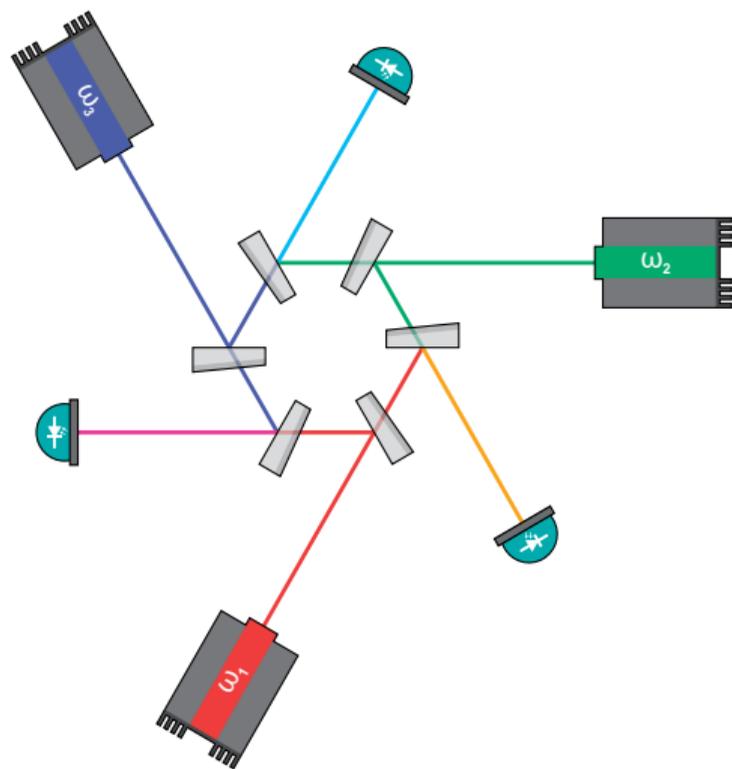
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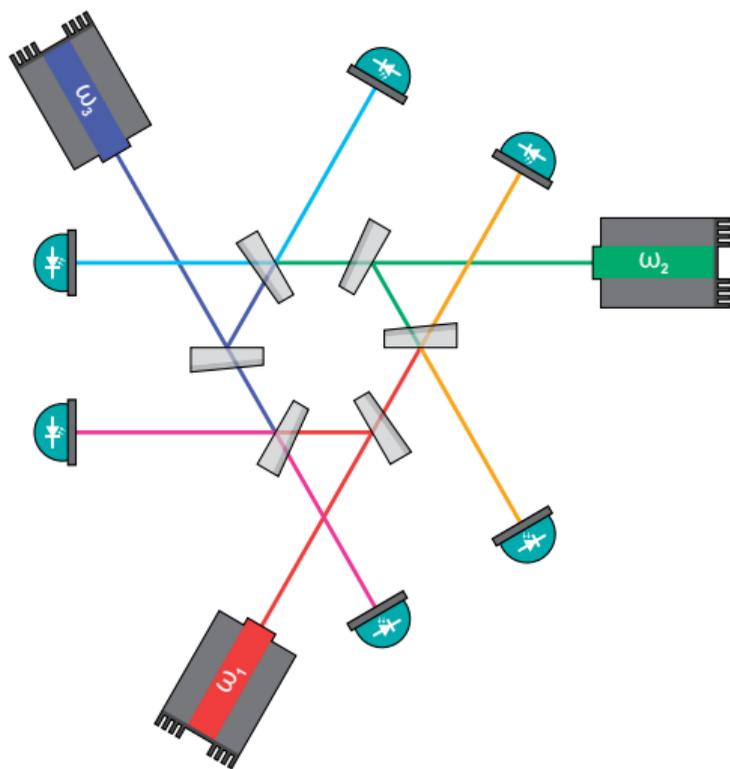
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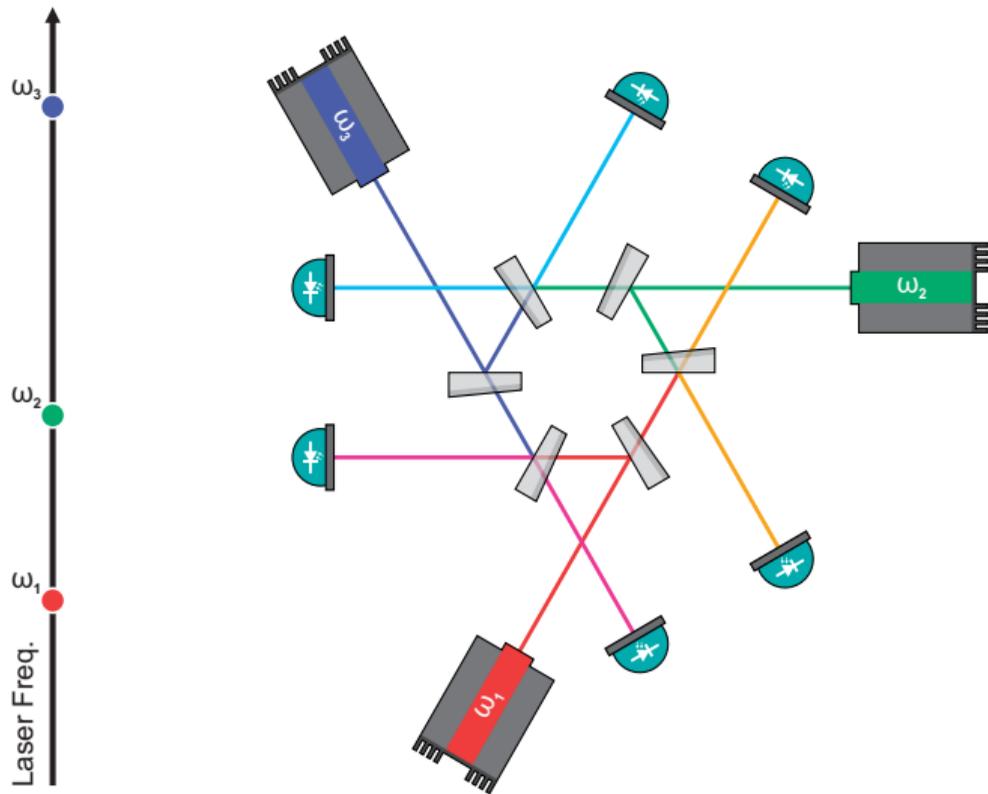
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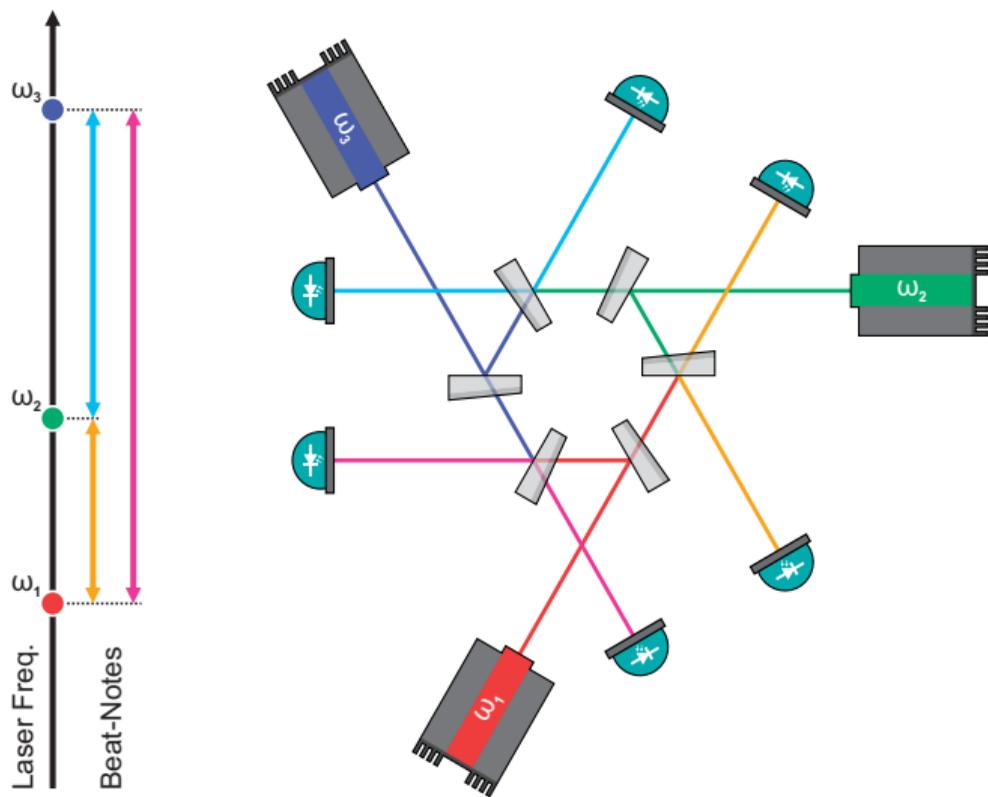
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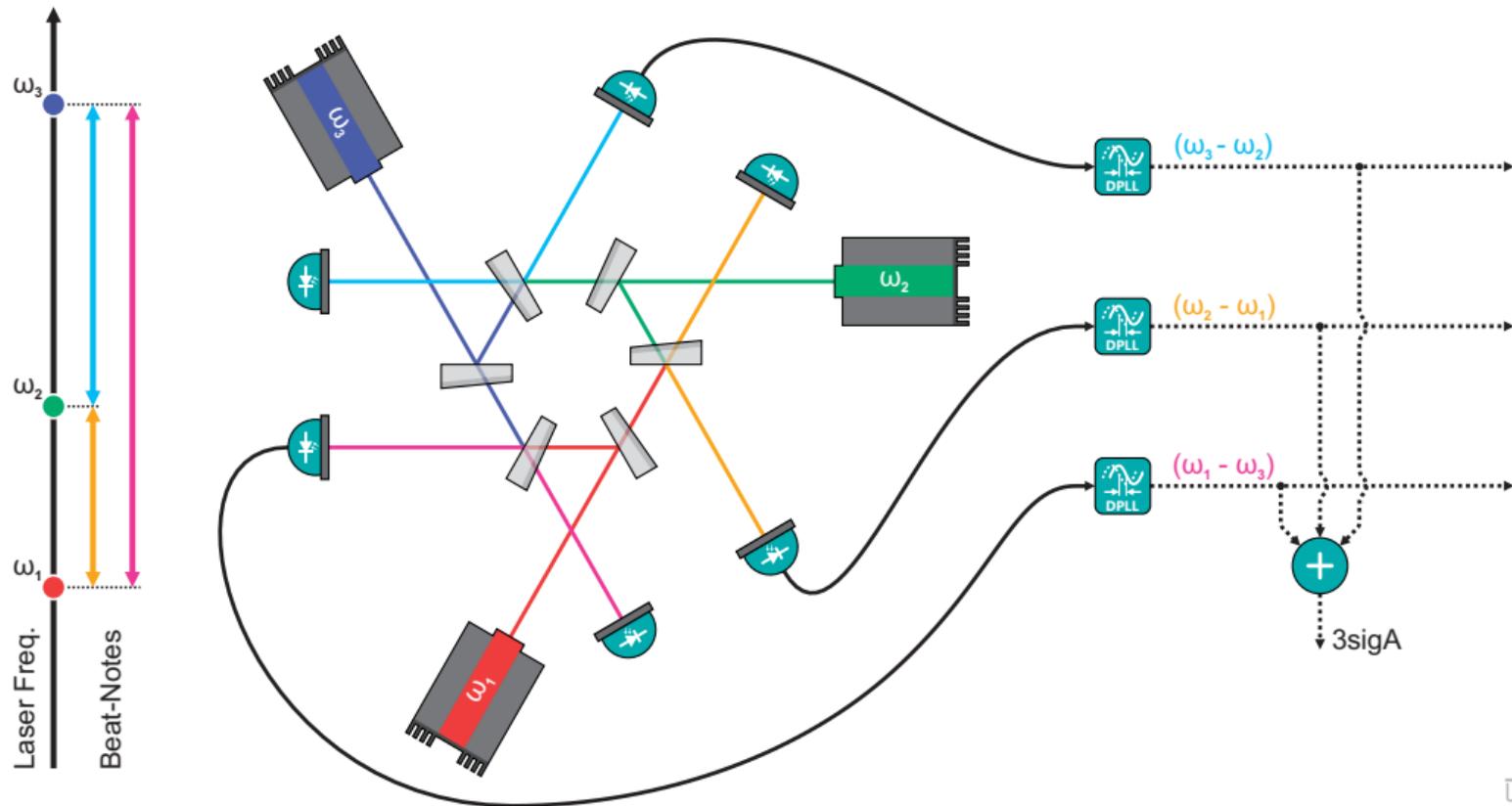
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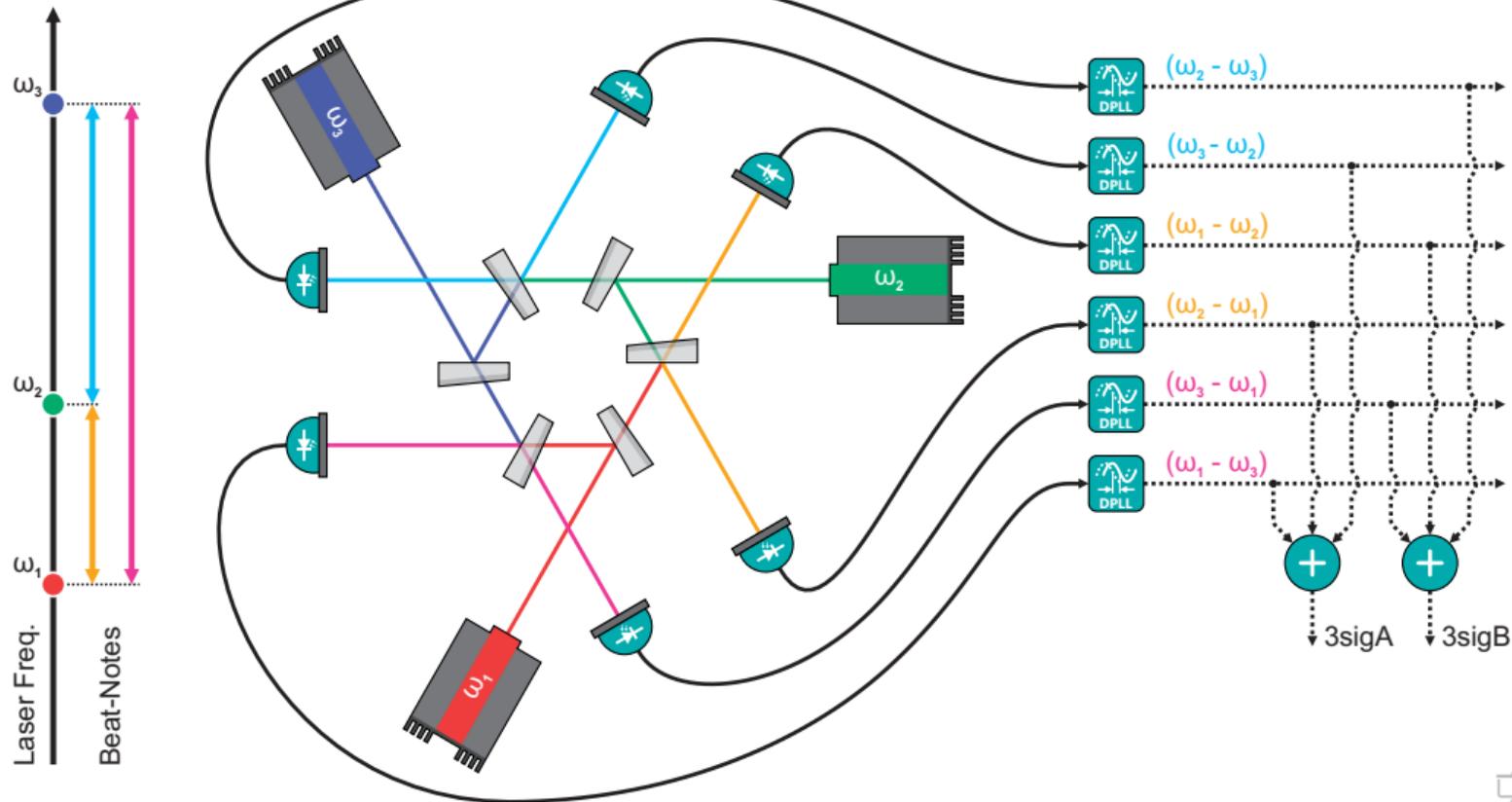
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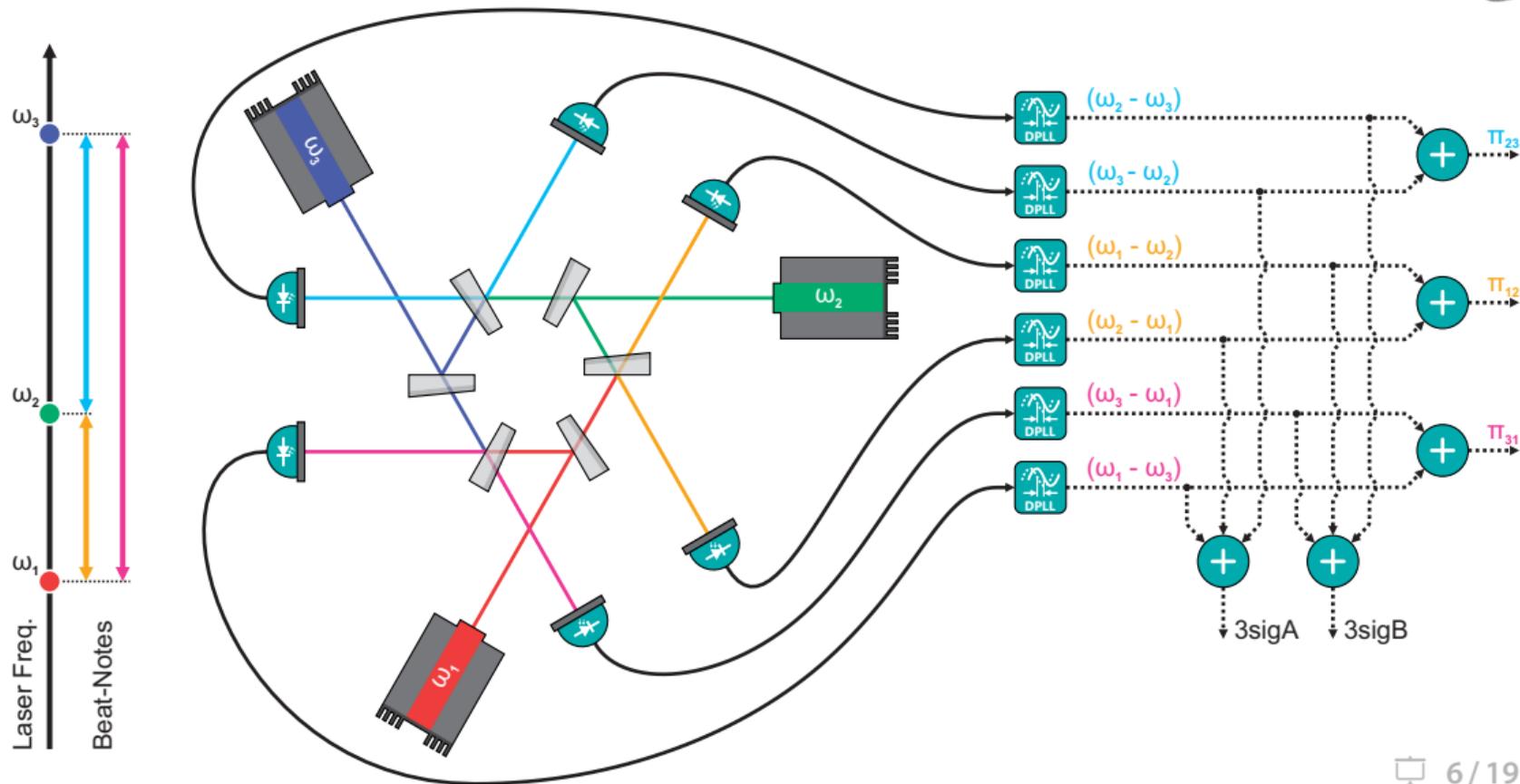
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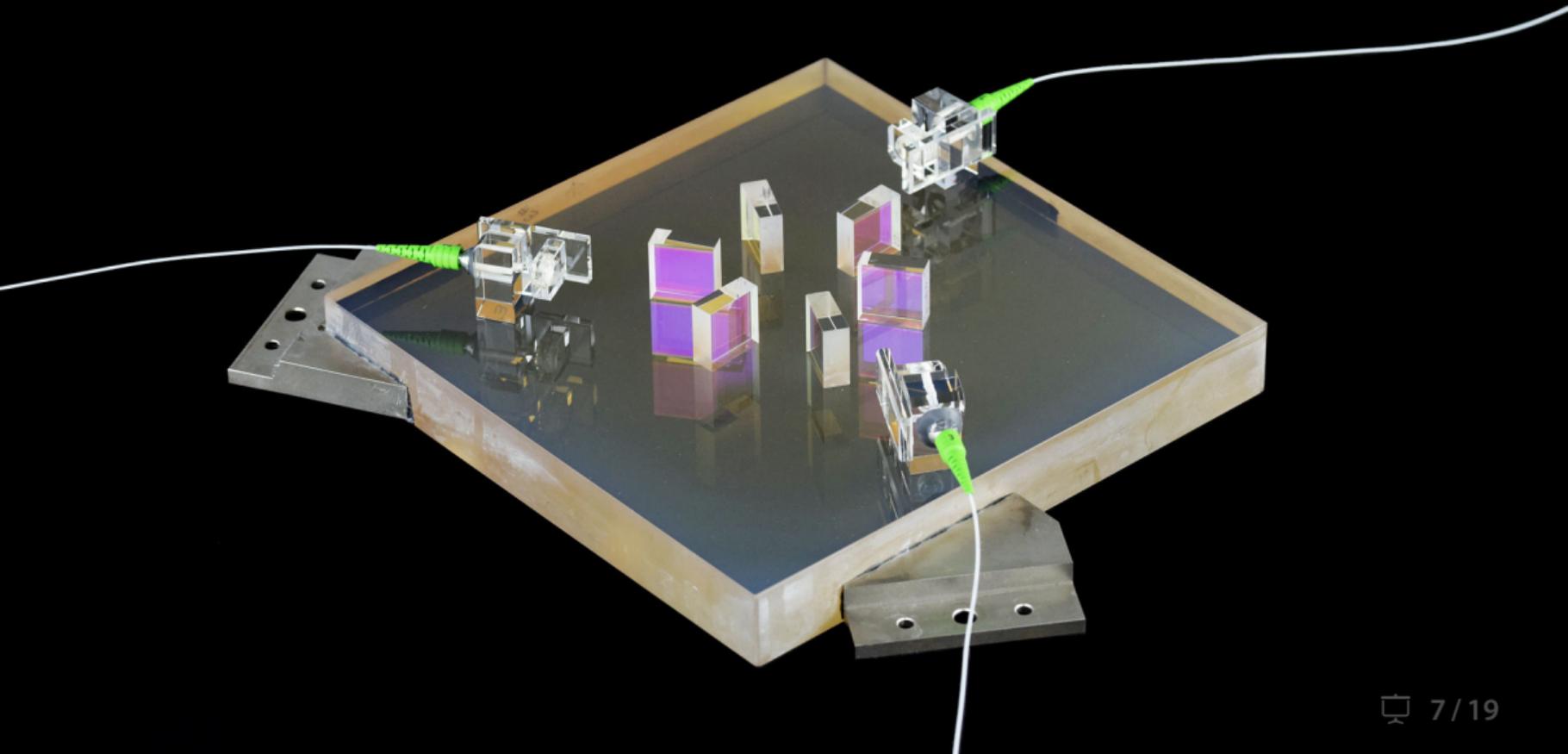
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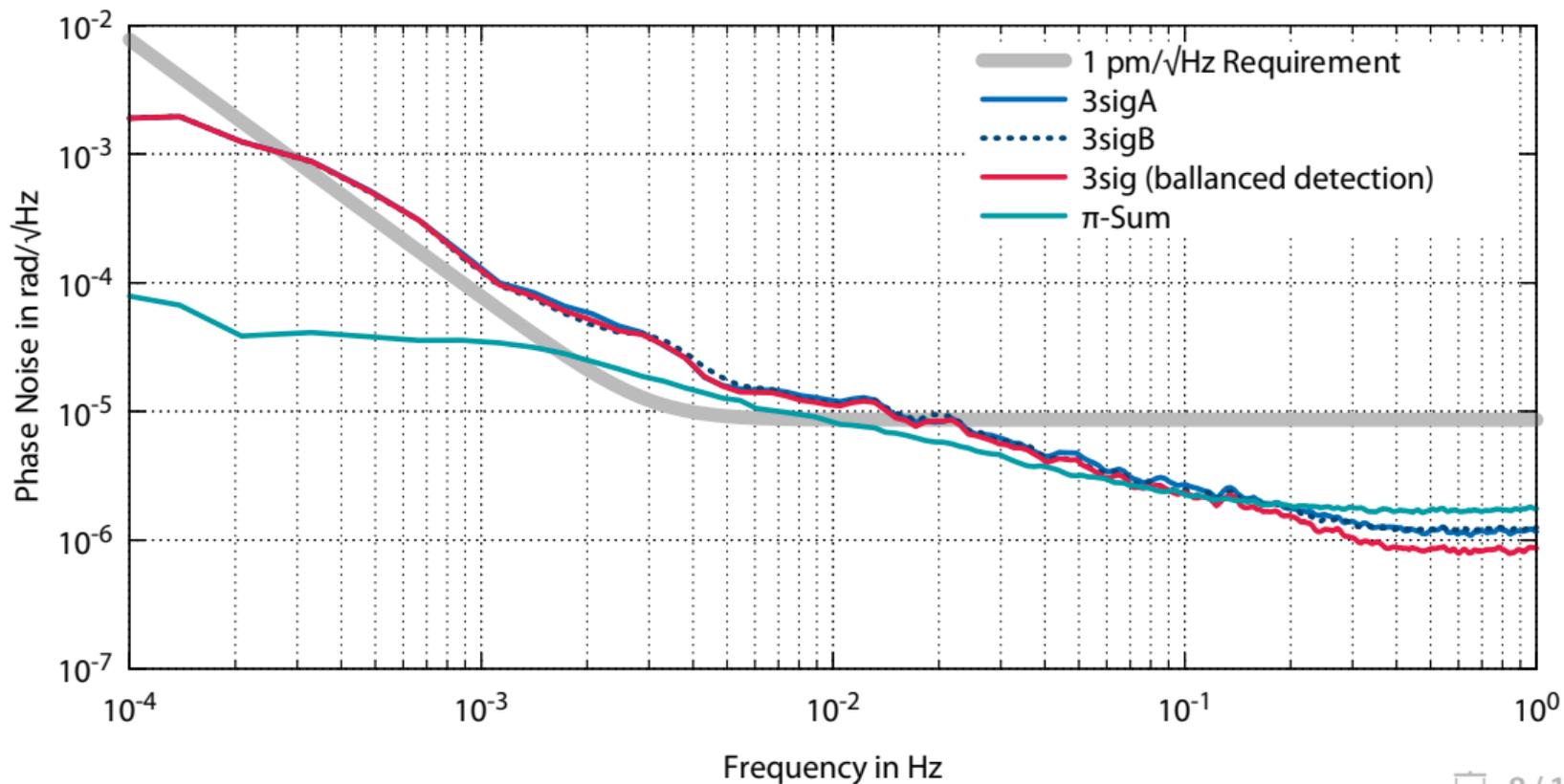
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Hexagon – Ultra-stable Implementation



Hexagon – Performance Results



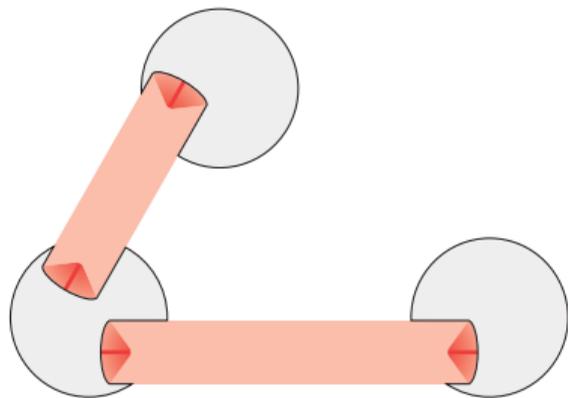


Post-processing & Auxiliary Functionality



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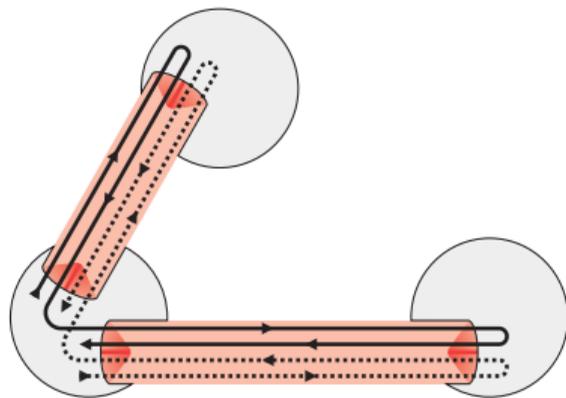
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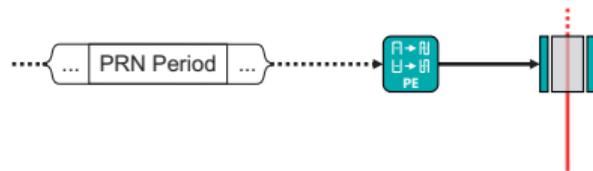
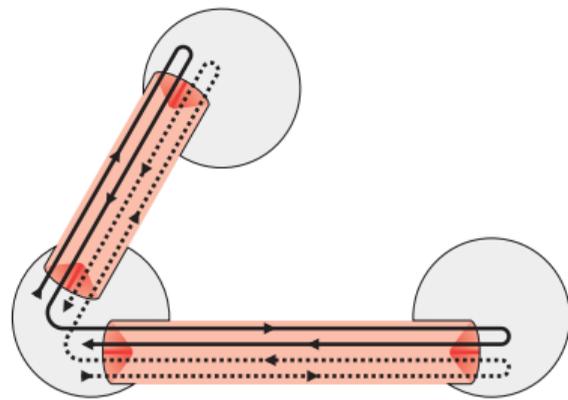
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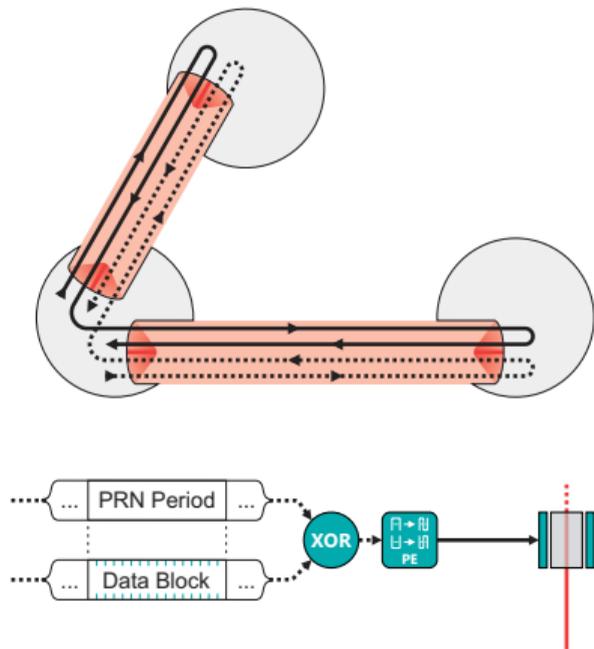
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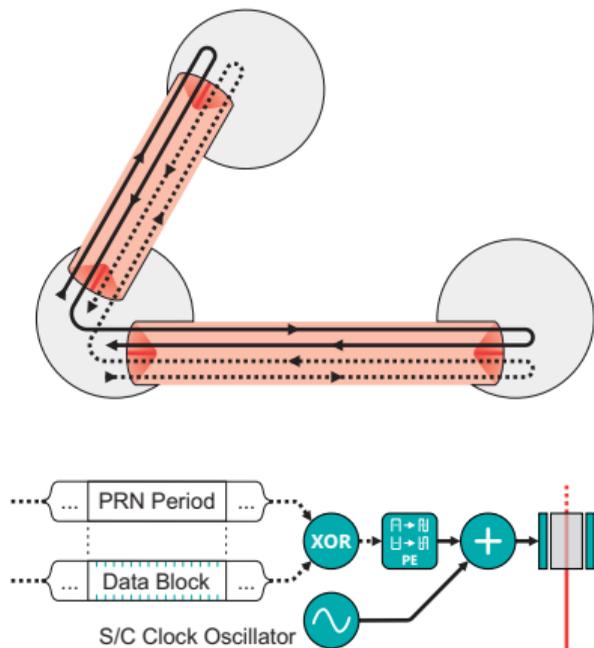
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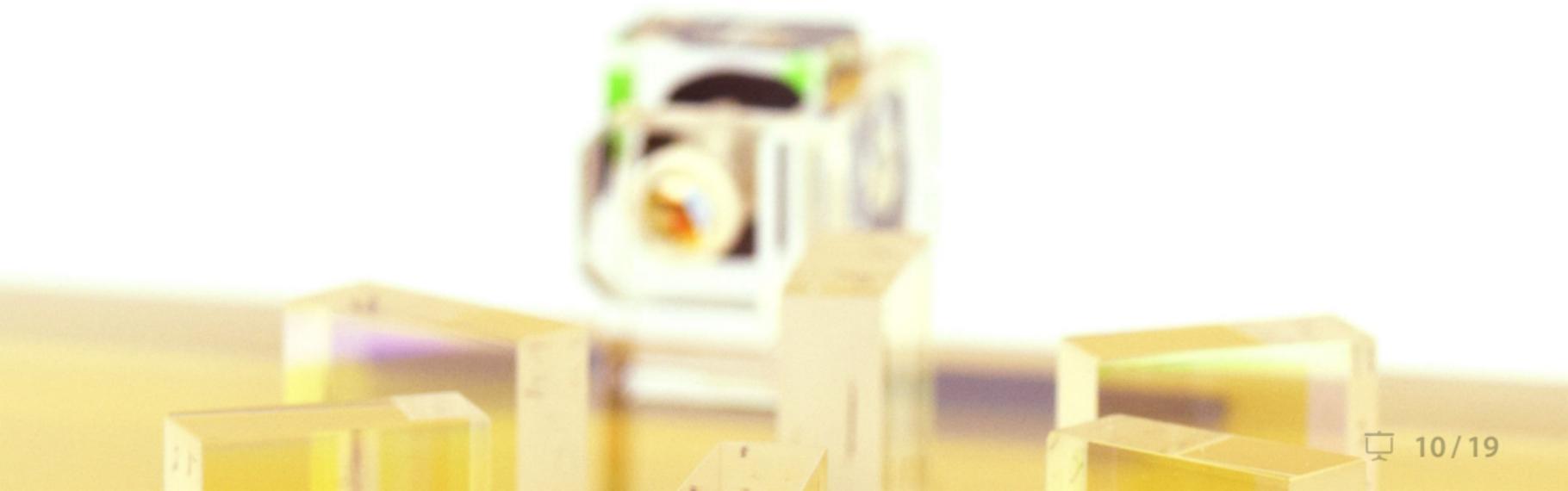
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- each S/C running on its own separate clock \Rightarrow clock-tone transfer via GHz sidebands on the laser links \Rightarrow dynamic re-sampling in post-processing (often considered part of TDI)





Step-by-step simplification of LISA





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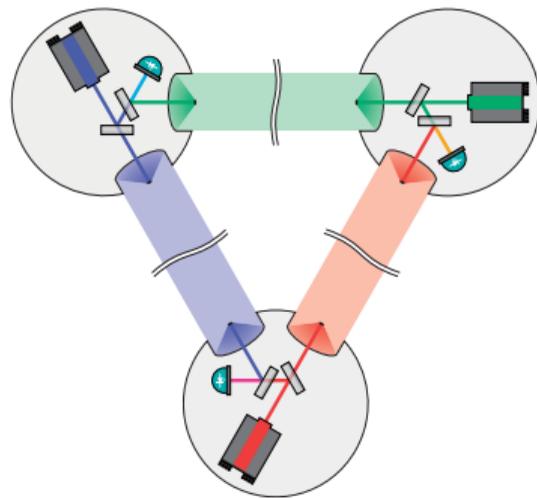
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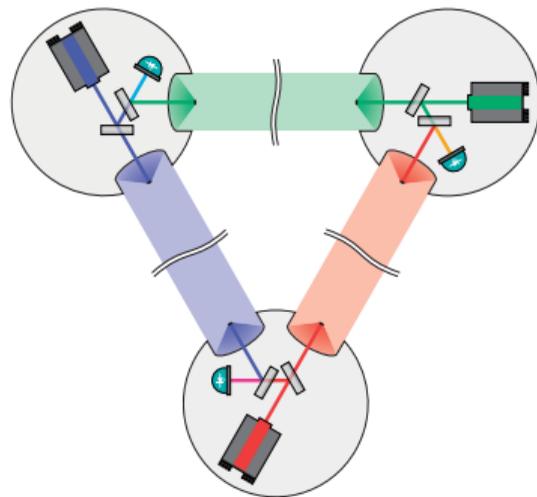
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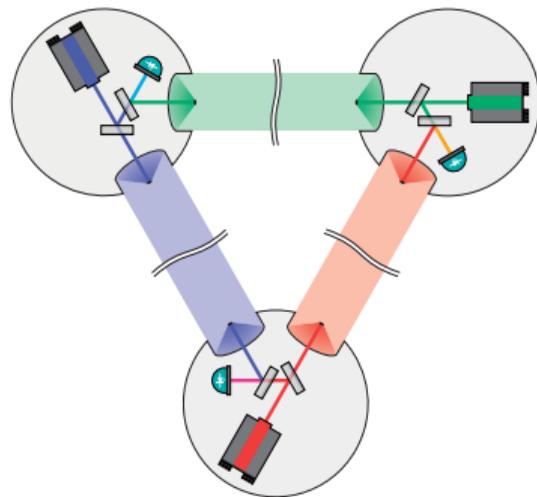
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- remove telescopes and shrink constellation down to fit on a single ultra-stable optical bench



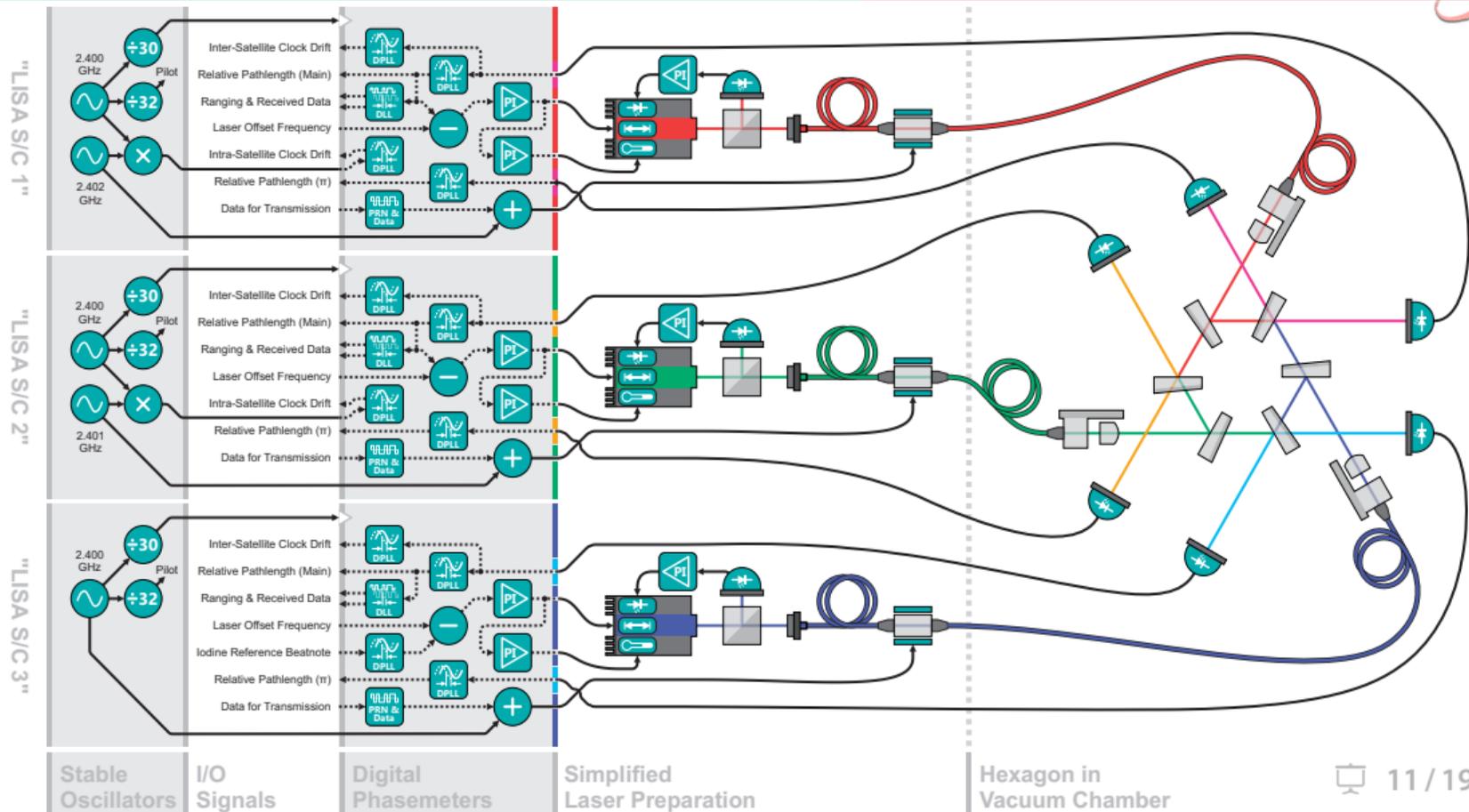


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- remove telescopes and shrink constellation down to fit on a single ultra-stable optical bench \Rightarrow Hexagon



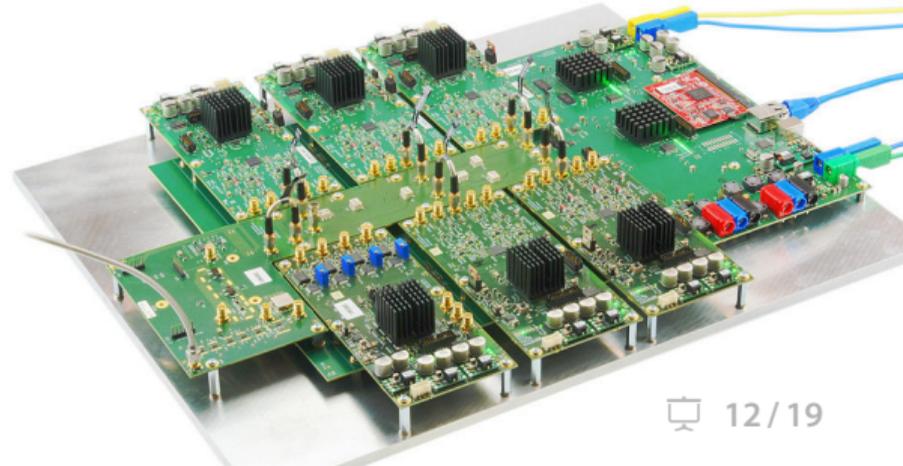
Hexagon – Testing the Full LISA Arm Metrology Chain





- **automatic beat-note acquisition**

FFT-based search for spectral peaks and initialization of digital phase-locked loops (DPLLs)



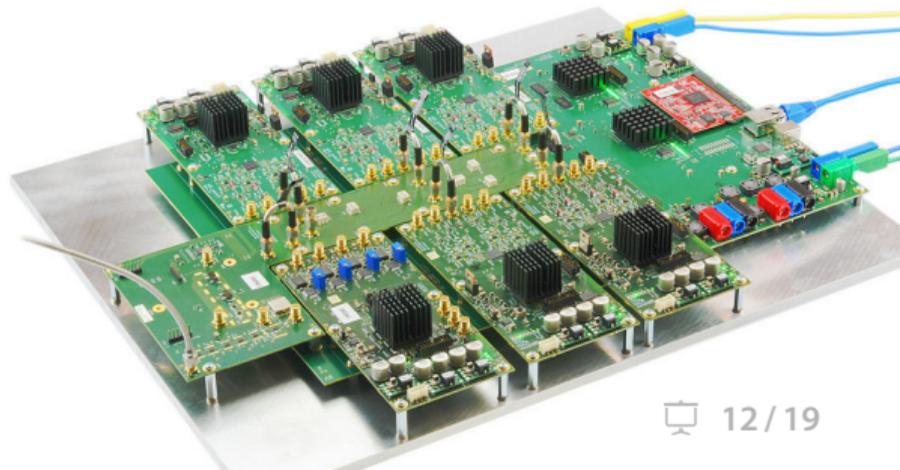


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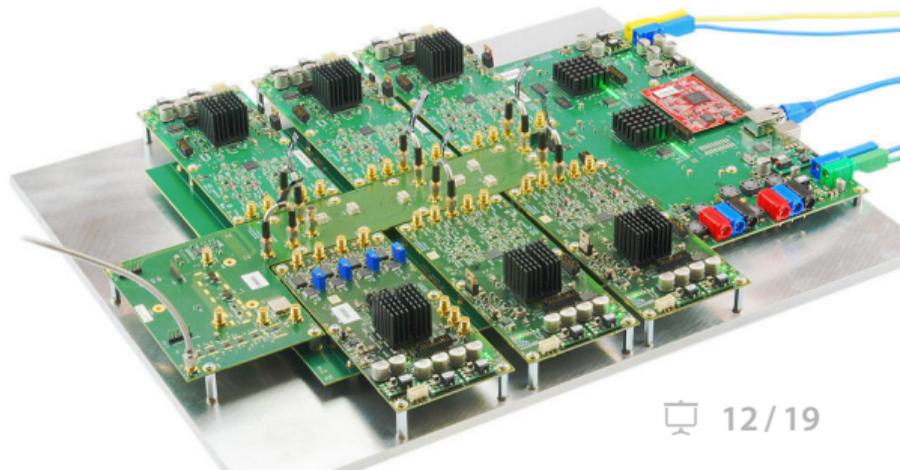
- **laser offset locking**

keep beat-note within phasemeter's bandwidth & beat-note sweeping (\Rightarrow dynamic Doppler shifts)



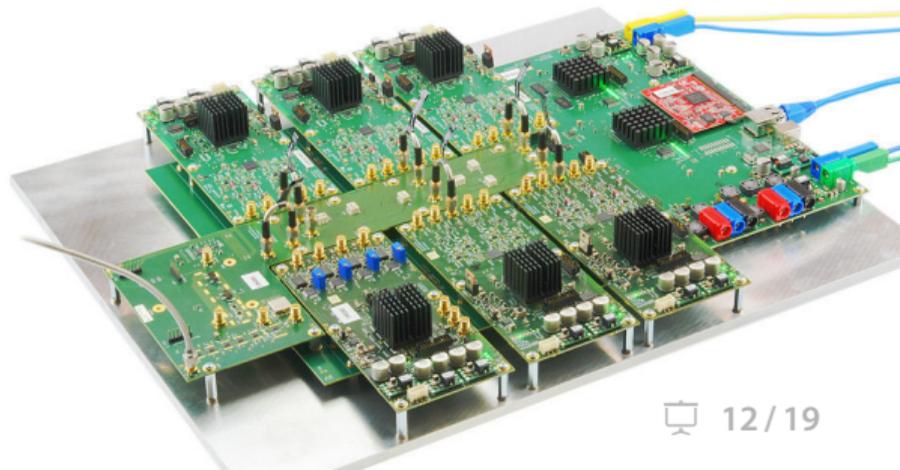


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using sum/differential signals



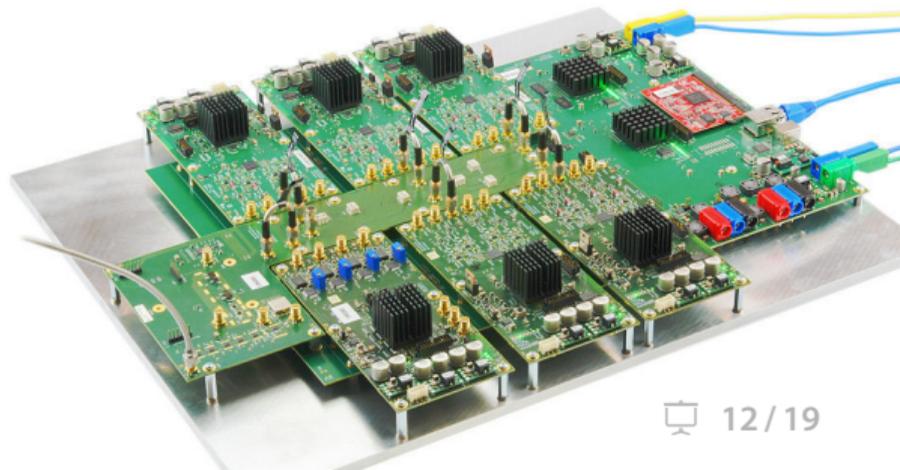


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measure intra-satellite clock jitter



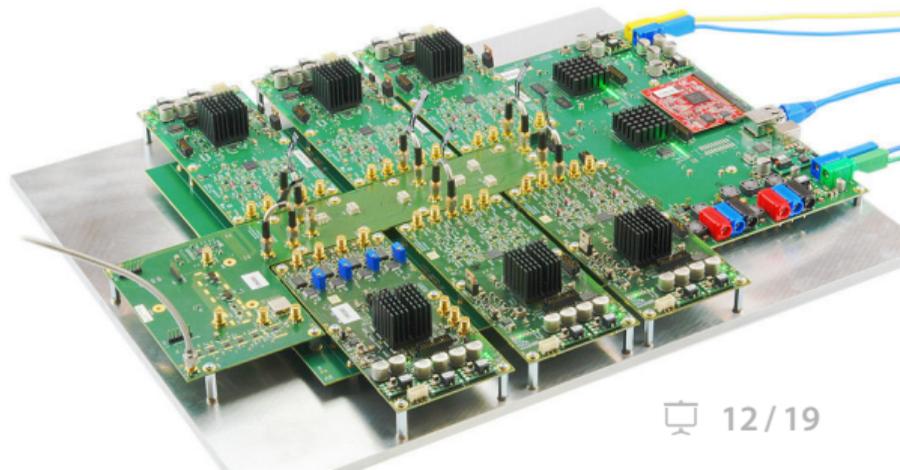


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- **clocktone transfer**
measure inter-satellite clock drifts



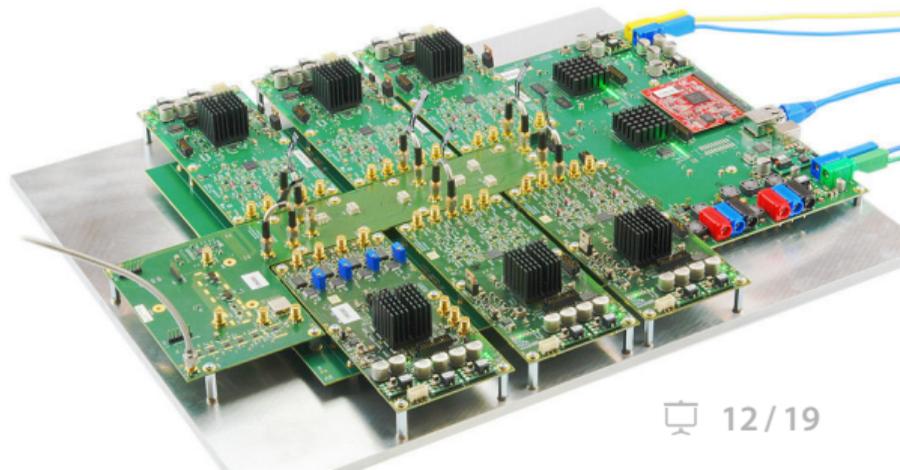


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PRN modulation (TX) & delay locked loop (RX)



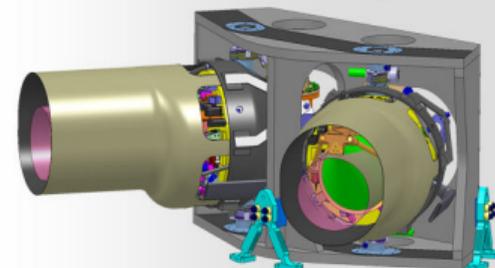
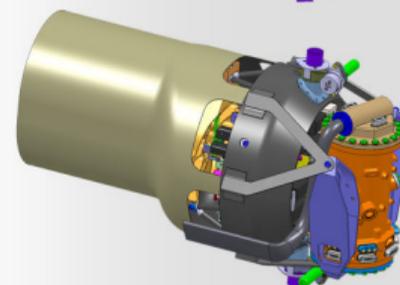
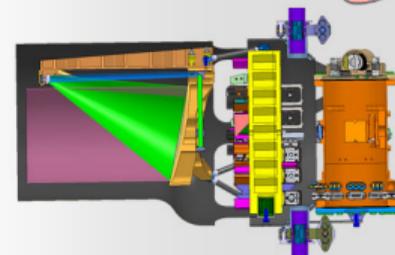


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PRN modulation (TX) & delay locked loop (RX)
- **low bandwidth data communication**
switching sign of chunks of PRN chips





Intra-Spacecraft Motion and Spurious Light

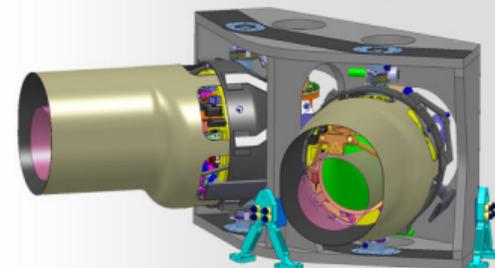
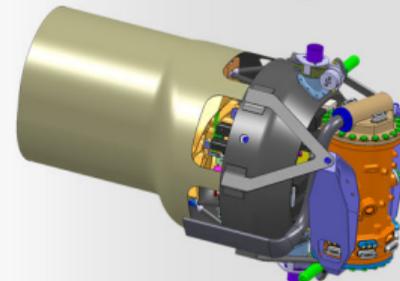
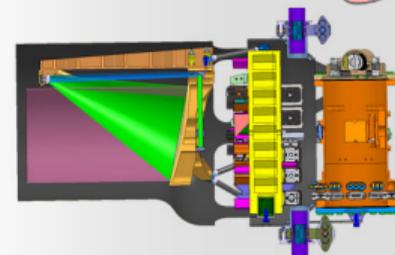


Images: LISA moving optical subassemblies (MOSAs), each comprising telescope, optical bench and test mass chamber; courtesy of Airbus D&S



Intra-Spacecraft Motion and Spurious Light

- angular breathing \Rightarrow variable pointing of telescopes and their rigidly attached dedicated optical benches \Rightarrow inter-bench phase reference distribution system (PRDS a.k.a. "backlink") to subtract laser frequency noise in post-processing (requirement: $2 \text{ pm}/\sqrt{\text{Hz}}$ reciprocal phase noise)

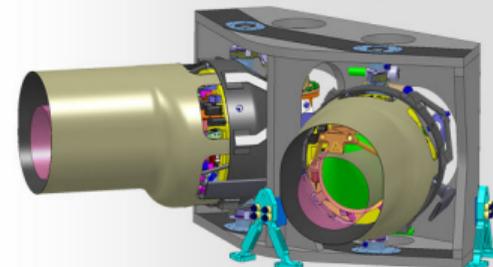
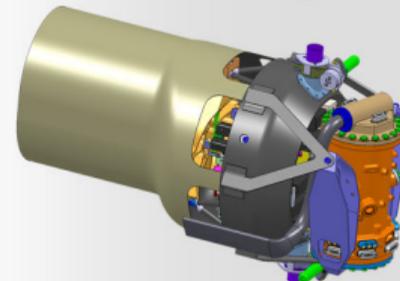
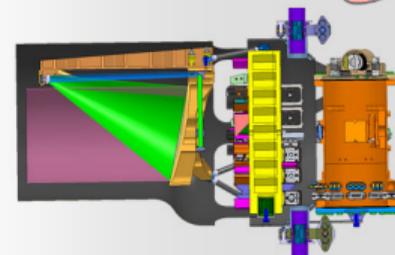


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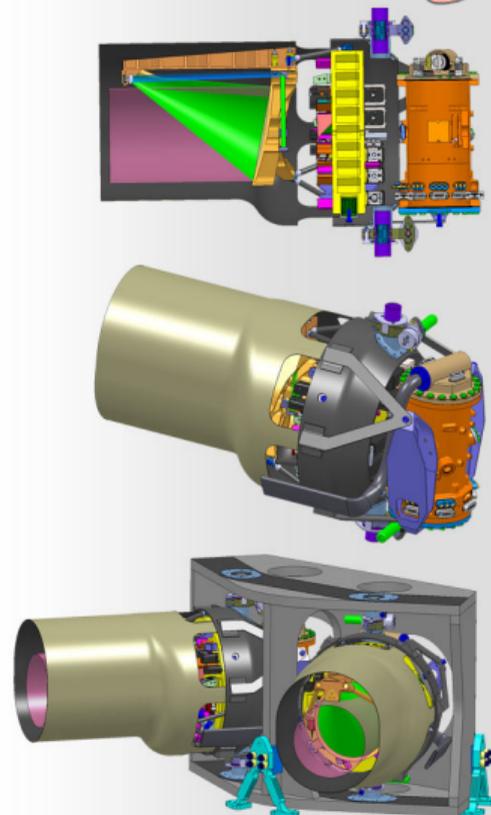
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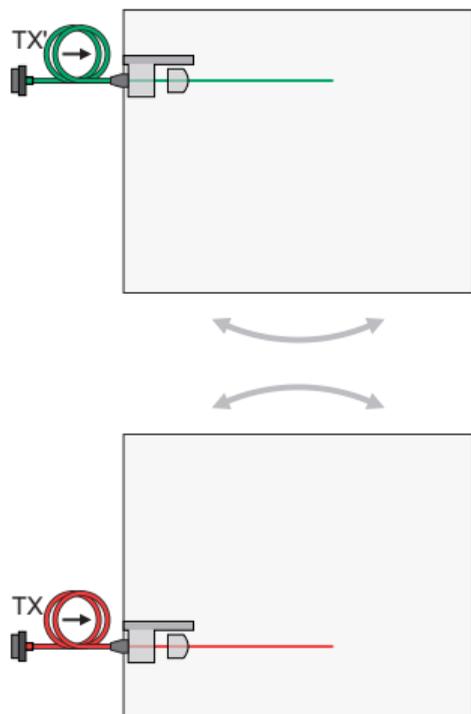
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- fiber back scatter within PM fibers has previously been measured to be $\approx 4 \text{ ppm}/\text{m}$, but what are the effects of cosmic/solar radiation? \Rightarrow separate experiment using gamma and neutron radiation [AEI & Fraunhofer INT]

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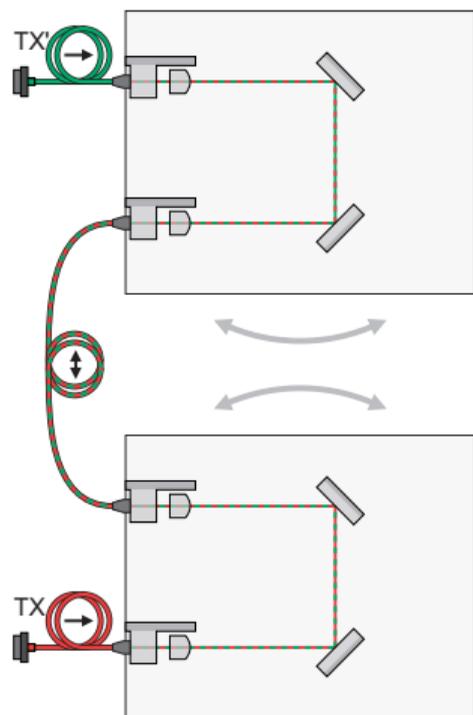
Classic Attenuated Fiber BL



- objective: each spacecraft houses two TX laser sources (one for each arm, indicated in red and green) which need to be interconnected by a flexible optical backlink



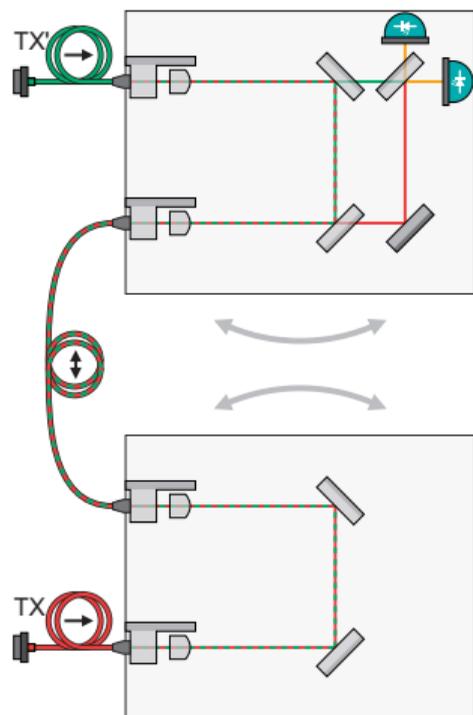
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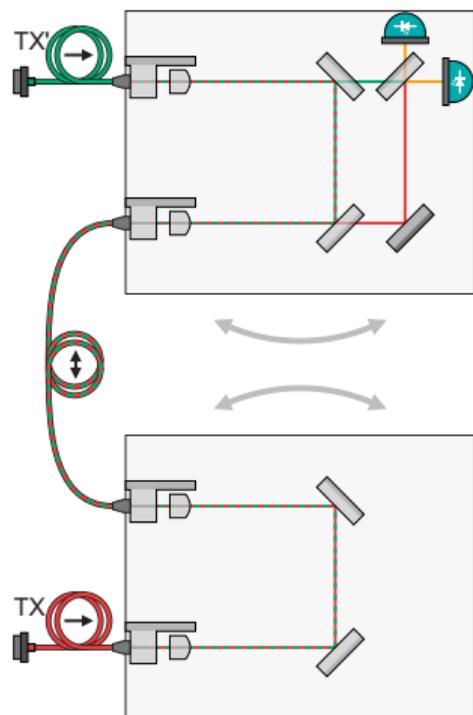
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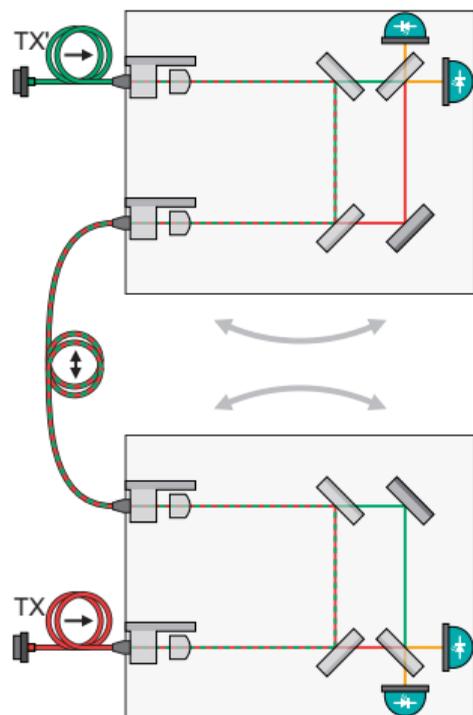
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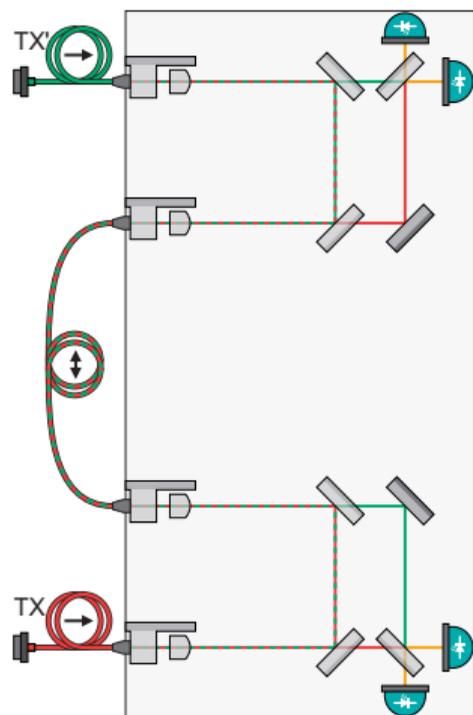
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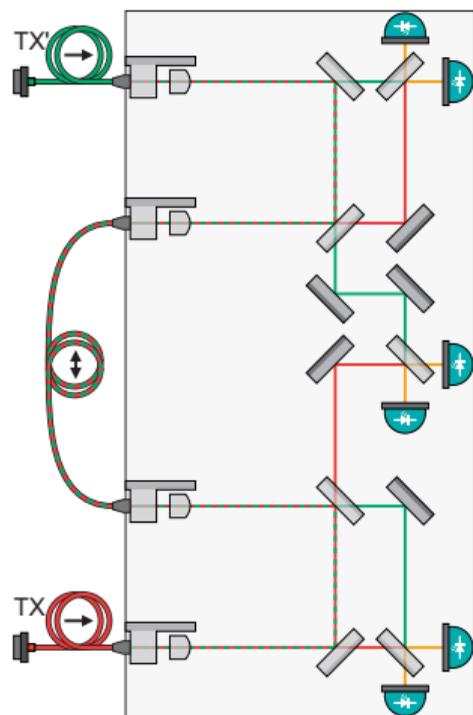
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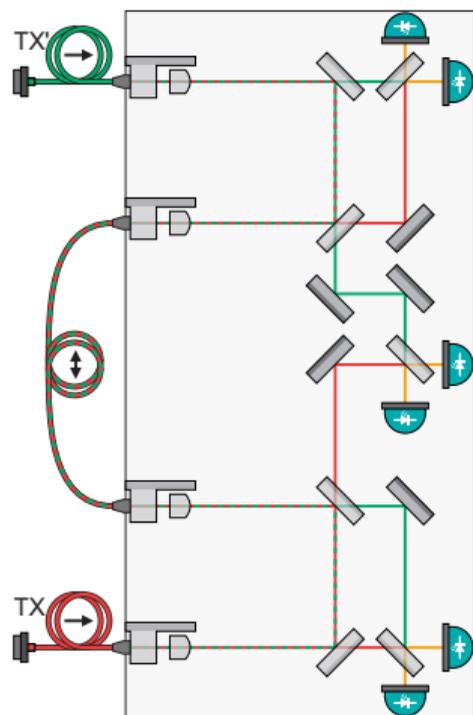
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- lab test: place setup on a common optical bench and add reference interferometer (middle) \Rightarrow classic fiber backlink experiment



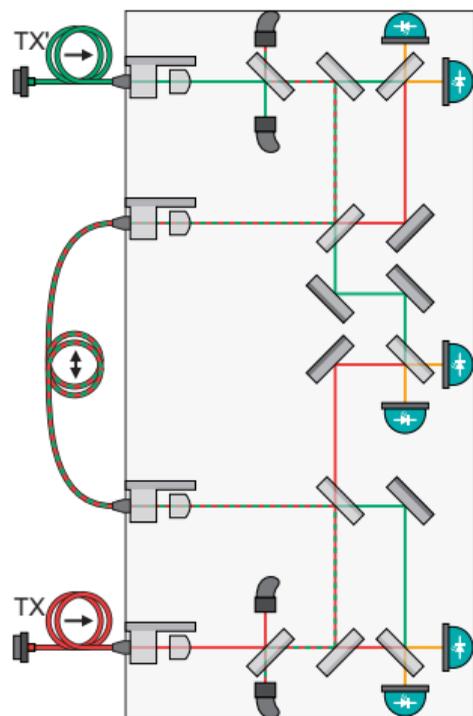
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- not sufficient: back-reflected beams wherever light was coupled into a fiber coupler \Rightarrow spurious beat-notes in our interferometers



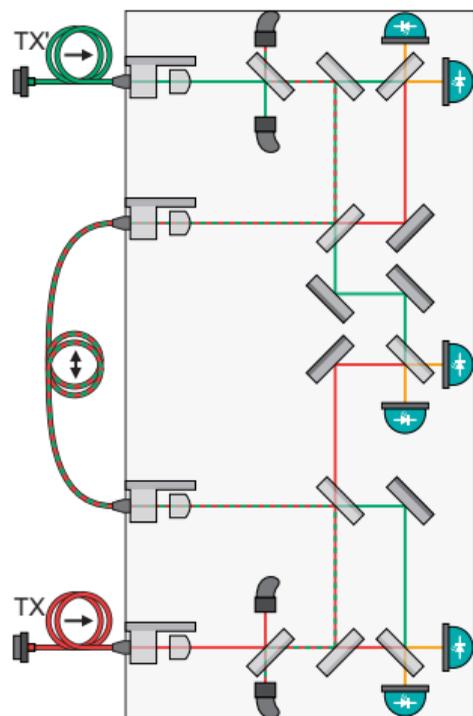
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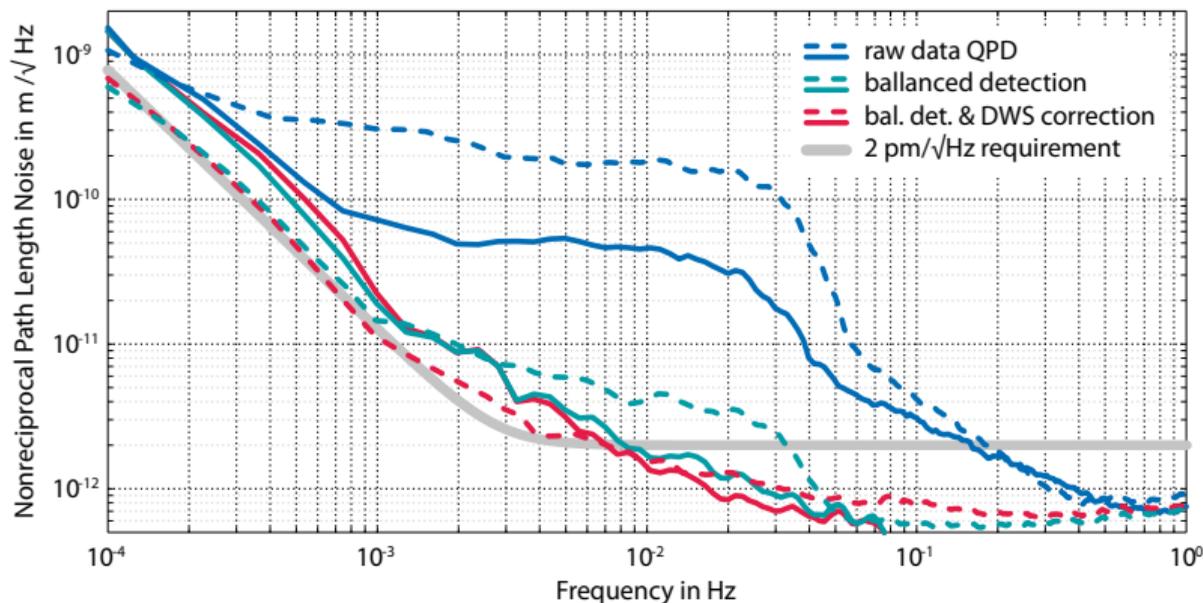
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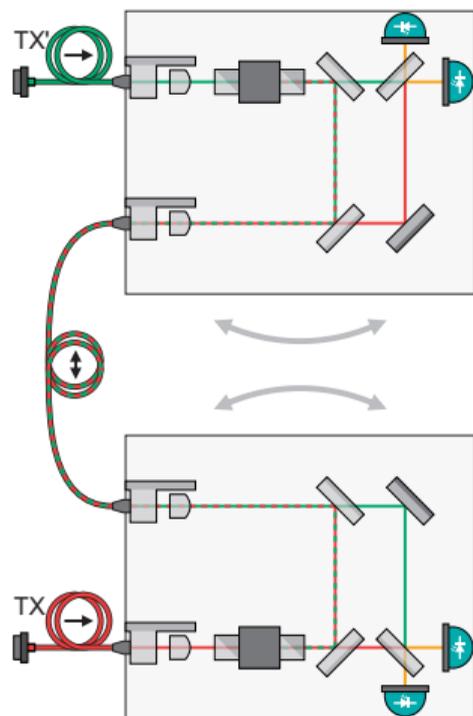
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- attenuation stages after TX fiber couplers \Rightarrow back-reflected beams attenuated twice as much as TX beams
- no such solution exists for the backlink fiber couplers (backlink light and spurious beams would see the same attenuation)



- attenuation stages (solid lines) and balanced detection sufficient for required fiber phase reciprocity
- balanced detection requires more photoreceivers and phasemeter channels
- DWS correction (classic backlink used commercial fiber couplers instead of monolithic ones)

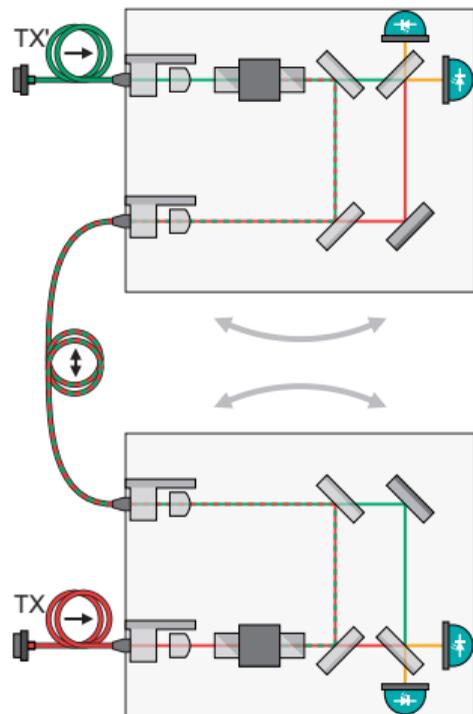


Classic Fiber Equivalent BL

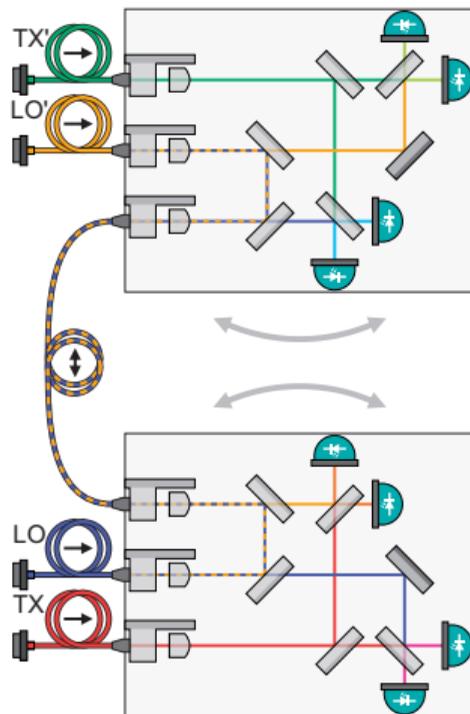




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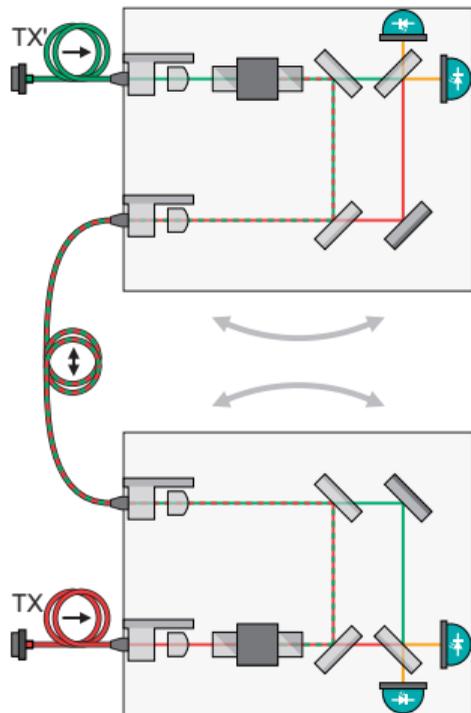


Frequency-shifted Fiber BL

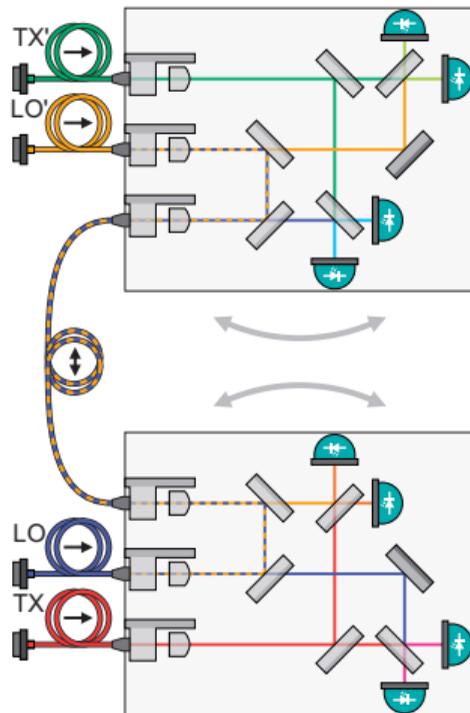




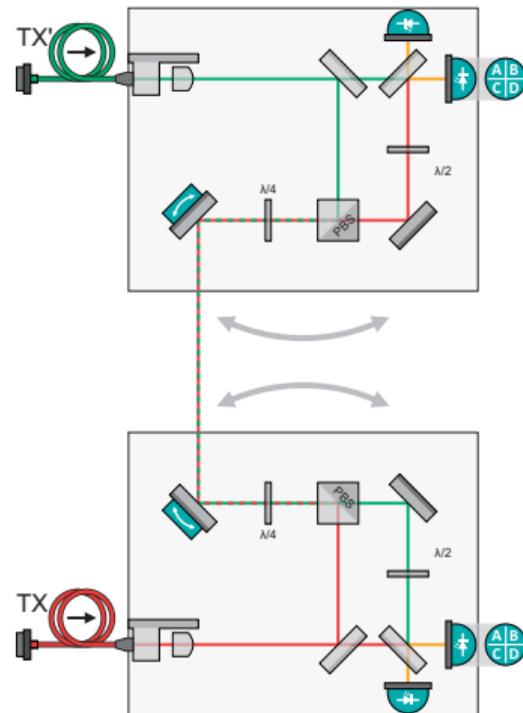
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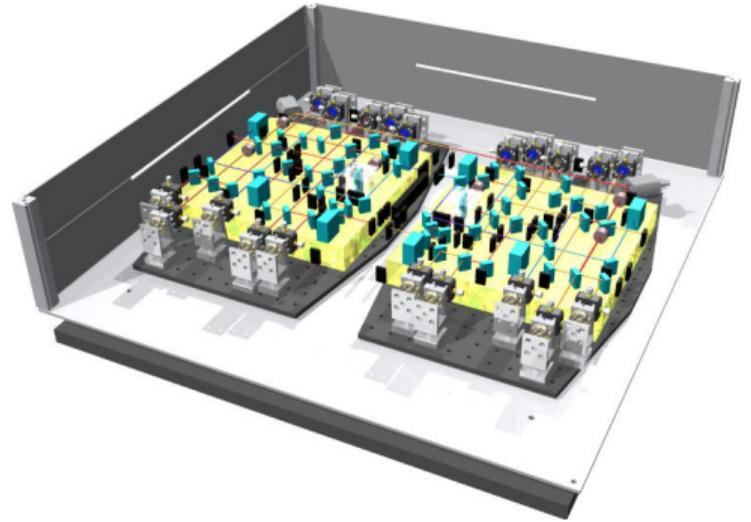
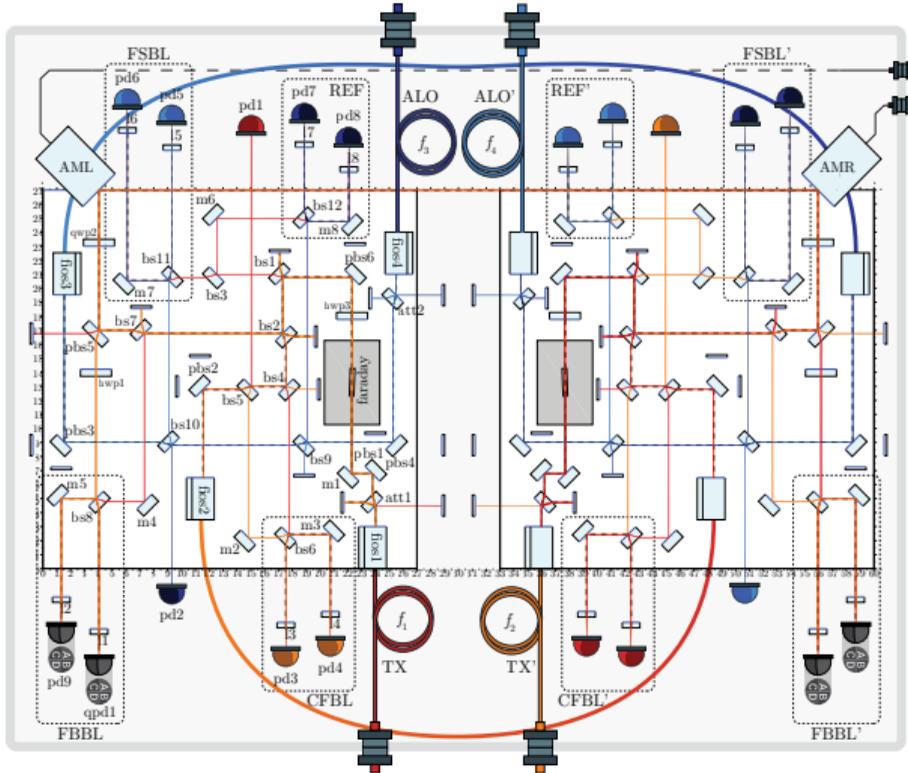
Frequency-shifted Fiber BL



Pol-encoded Free Beam BL



Three-Backlink Experiment

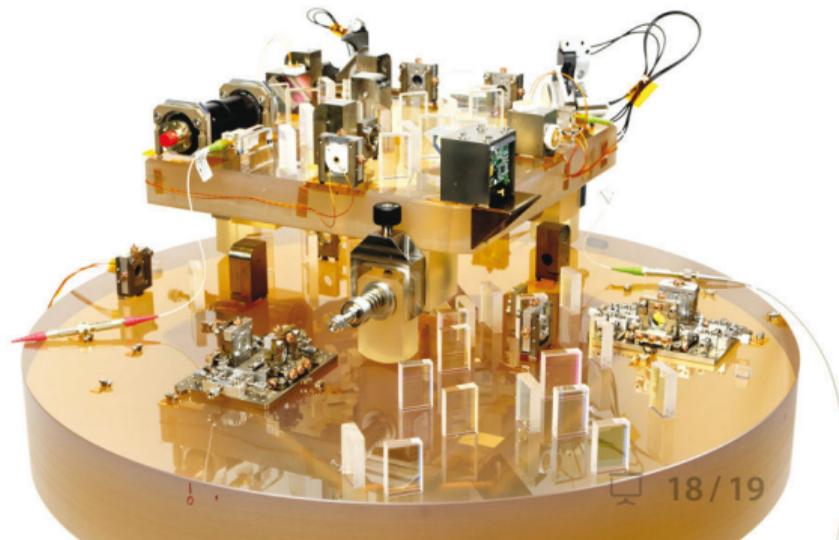


Paper: KS Isleif et al., Towards the LISA Backlink, CQG 2018



Other Activities

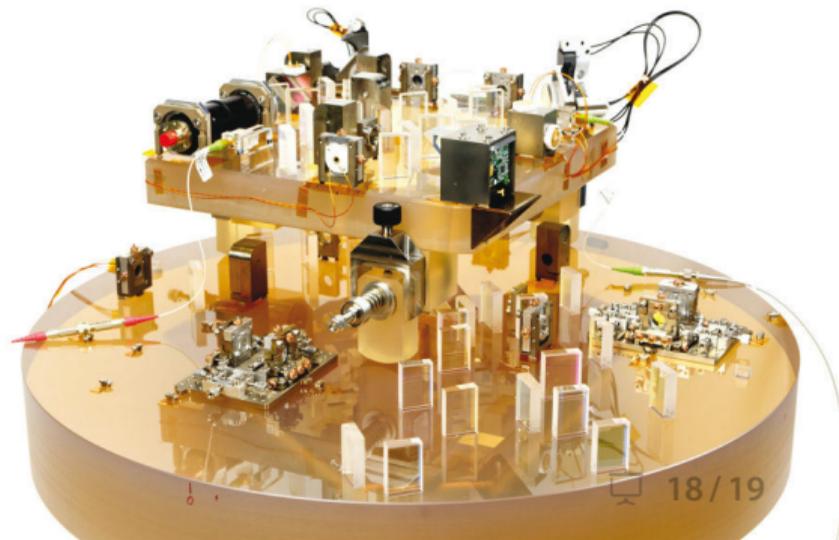
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Other Activities

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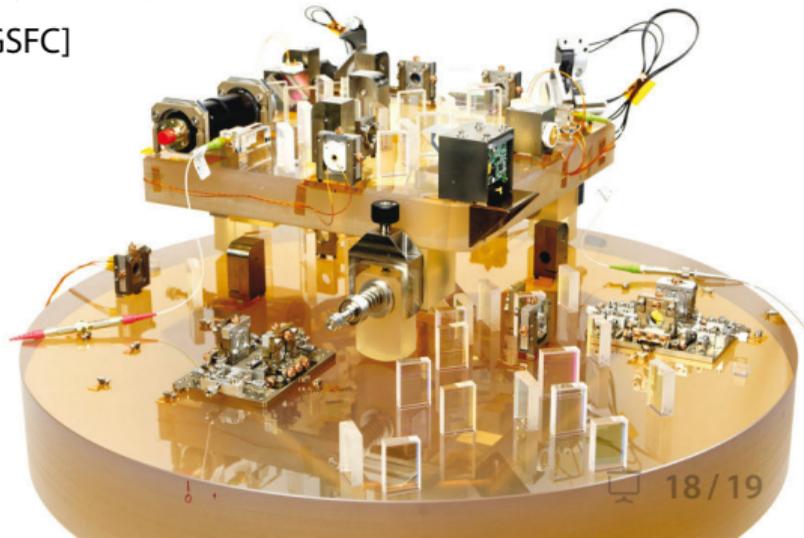
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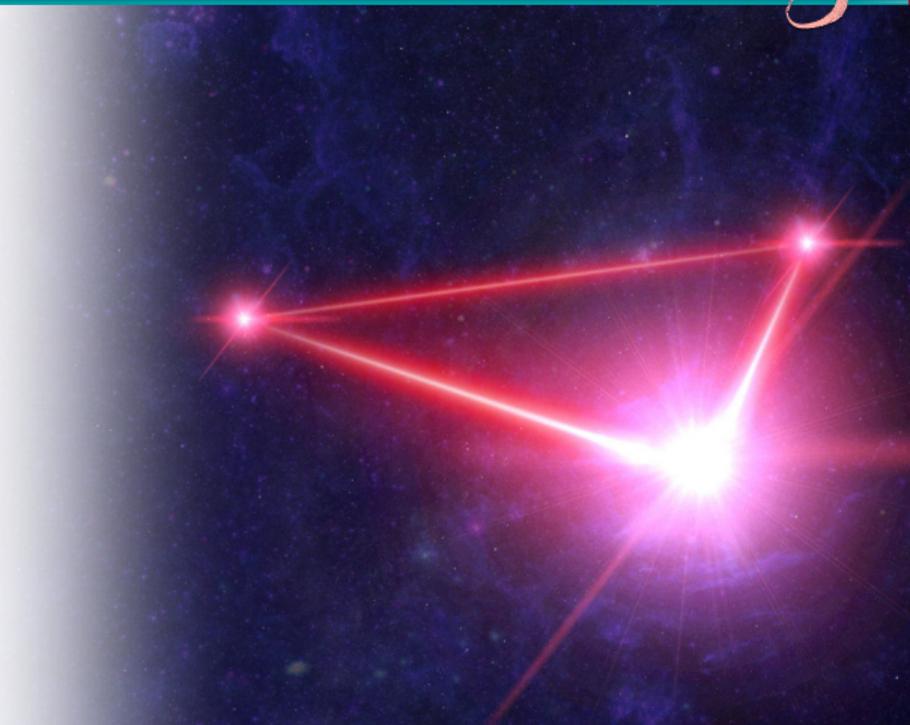
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- and much more: laser development, frequency pre-stabilisation, laser link acquisition, signal pre-processing, TDI, LISA simulator, ...





- **Laser Interferometer Space Antenna**
space-borne gravitational wave observatory for sources inaccessible to ground-based detectors
- **The Hexagon Interferometer**
an optical three-signal testbed for LISA's phase readout system but capable of testing the full LISA arm metrology chain in the future
- **The Three-Backlink Experiment**
direct comparison between several alternative implementations of a LISA backlink, providing an optical phase reference between the two moving optical benches within one S/C



Thank you very much!