

${}^7\text{Be}(n,p)$ cross section measurement for the Cosmological Lithium Problem at the n_TOF facility at CERN

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Big Bang Nucleosynthesis (BBN) theory predicts the abundances of the light elements D, ${}^3\text{He}$, ${}^4\text{He}$ and ${}^7\text{Li}$ produced in the early universe. The primordial abundances of D, ${}^3\text{He}$ and ${}^4\text{He}$ inferred from observational data are in good agreement with predictions, however, the BBN theory overestimates the primordial ${}^7\text{Li}$ abundance by about a factor of three with respect to the observations in metal poor halo stars [1]. This discrepancy is known as “Cosmological Lithium Problem”(CLiP). Since primordial ${}^7\text{Li}$ is produced mainly by the decay of ${}^7\text{Be}$, reducing the amount of ${}^7\text{Be}$ surviving the BBN phase, reduces the primordial ${}^7\text{Li}$. The two principal reactions responsible of the destruction of ${}^7\text{Be}$ via neutron reactions are: the ${}^7\text{Be}(n,p){}^7\text{Li}$, providing 97% destruction of ${}^7\text{Be}$ and the ${}^7\text{Be}(n,\alpha){}^4\text{He}$, responsible of 2.5%. The (n, α) reaction has already been studied at the n_TOF facility at CERN, where its cross section has been found too low to solve the CLiP[2]. Various measurements have excluded also a significant effect on the CLiP of charged particle induced reactions on ${}^7\text{Be}$, so the only possibility left to find a Nuclear physics solution to the problem is the (n,p) reaction. Despite the importance of this reaction in BBN, there is a lack of cross section data. Taking advantage of the innovative features of the second experimental area at n_TOF facility at CERN[3][4], e.g. the very high instantaneous flux, the wide energy range and the low background conditions, an accurate measurement of ${}^7\text{Be}(n,p){}^7\text{Li}$ cross section has been recently performed at n_TOF with a pure ${}^7\text{Be}$ target produced by implantation of a ${}^7\text{Be}$ beam at ISOLDE. The experimental procedure, the set-up used in the measurement and the results will be presented in this talk.

References

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- [4] C.Weiss et al., NIM A {799}(2015)90

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