## Conventional quarkonia: few experimental ideas

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Bound states in strongly coupled systems GGI, Firenze, 02/12/2018

## Disclaimer

$\rightarrow$ I am part of the Belle/Belle II collaborations
$\rightarrow$ I've mostly done bottomonium physics in my life
$\rightarrow$ This talk contains lots of personal opinions

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I believe that conventional bottomonia are (experimentally) more important than conventional charmonia right now:
$\rightarrow$ More states
$\rightarrow$ More transitions to be explored
$\rightarrow$ More almost unexplored topics
$\rightarrow$ Limited time and chances to actually do this physics

## Charmonium: experimental tools

Charmonium is experimentally easy and accessible
$\rightarrow$ Direct production in $\mathrm{e}^{+} \mathrm{e}^{-}$collisions $\underset{\text { B }}{ }$ B $\rightarrow$ Production in $\mathrm{B} \rightarrow \mathrm{K} \mathrm{c} \mathrm{\bar{c}} \mathrm{LHCb} \underset{\mathrm{LB}}{ }$
$\rightarrow$ Photon-photon scattering $\gamma \gamma^{*} \rightarrow(c \bar{c})$

$\rightarrow$ Double Charmonium $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow(\mathrm{cc})(\mathrm{c} \overline{\mathrm{c}})$

$\rightarrow$ Prompt production

$\rightarrow$ Direct production in $\bar{p} \bar{\square}$ (???)
$\rightarrow$ A future super-tau-charm factory (???)


Bottom line: Charmonium will still be fully covered in the next 15 yrs.

## Bottomonium: experimental tools

## Bottomonium is much less accessible

$\rightarrow$ Direct production in $\mathrm{e}^{+} \mathrm{e}^{-}$collisions

$\rightarrow$ Prompt production



Bottom line: after Belle II, only the LHC experiments will cover bottomonia with strong limitations

## Production is not everything

## $\mathrm{e}^{+} \mathrm{e}^{-}$machines

$\rightarrow$ Triggers are quite open
$\rightarrow$ High efficiency / Sensitive to very low momentum
$\rightarrow$ Unique measurements (double charmonium, $\gamma \gamma^{*} \rightarrow \mathrm{cc}$ )
$\rightarrow$ Initial states is always a $1^{-1}$ quarkonium or a B meson
$\rightarrow$ CM energy is a limiting factor



## Production is not everything

## Hadronic machines

$\rightarrow$ Produce any quantum numbers
$\rightarrow$ CM energy is not an issue
$\rightarrow$ Unique measurements (double Y , polarization, cross sections...)
$\rightarrow$ Triggers are a limiting factor
$\rightarrow$ No inclusive analysis: only $\mu \mu, \mathrm{pp} \ldots$
$\rightarrow$ (No soft photons, no neutral mesons ...)


## Production is not everything

## Hadronic machines

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$\rightarrow$ No inclusive analysis: only $\mu \mu$, pp...
$\rightarrow$ (No soft photons, no neutral mesons ...)
Rule of thumb for bottomonia:
$\rightarrow$ Yes to $\pi \pi / \gamma+\mu \mu$ final states

$$
\begin{aligned}
& \mathrm{Y}(\mathrm{nS}) \rightarrow \pi \pi \mathrm{Y}(1 \mathrm{~S}) \\
& \chi_{\mathrm{b}}(\mathrm{nP}) \rightarrow \gamma \mathrm{Y}(1 \mathrm{~S})
\end{aligned}
$$

$\rightarrow$ No to multi-hadrons final states

Transitions pattern circa year 2005


Transitions pattern year 2016


$$
Y(n S) \rightarrow \gamma \eta_{b}(1 S)
$$

## Is Belle II going to take bottomonium data?

The collaboration considers $\mathrm{Y}(3 \mathrm{~S}, 5 \mathrm{~S}, 6 \mathrm{~S})$ runs as part of its physics program from the very beginning
$\rightarrow$ Still, the competition with LHCb on CPV is tough
$\rightarrow$ Nothing comes for free: a document for the Y(3S) run should be submitted by Feb. 2019




## The ground states parameters

## The $\eta_{c}$ mass conundrum



All the fits are performed using a Breit-Wigner shape

## The $\eta_{c}$ mass conundrum

Non-M1 naive average

M1 naive average

$$
\frac{d \Gamma(\omega)}{d \omega}=\frac{4}{3} \alpha \frac{e_{\mathrm{c}}^{2}}{m_{\mathrm{c}}^{2}} \omega^{3}|M|^{2} \mathrm{BW}(\omega) \longrightarrow \frac{d \Gamma(\omega)}{d \omega} \sim \omega^{3} f(\omega) \mathrm{BW}(\omega)
$$

What is the proper theoretical shaper for this dumping factor?

## The $\eta_{c}$ width conundrum



## What do I understand from this?

$\rightarrow$ NNLO is still not enough
$\rightarrow$ Is NRQCD converging fast enough?
$\rightarrow$ The problem is not in the experimental resolution

## Bottom line:

$\rightarrow$ New data may be useful, but the problem with the $\eta_{c}$ seems to be in the theoretical models rather than in the lack of good data

## The $\eta_{b}(1 S)$ case

No specific decay of the $\boldsymbol{\eta}_{b}(1 S)$ has been observed so far
$\rightarrow$ No (known) way to perform an exclusive reconstruction
$\rightarrow$ Can be studied only at B-factories looking at the
photon spectrum in $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow$ hadrons © $\mathrm{Y}(2 \mathrm{~S}, 3 \mathrm{~S})$




## The $\eta_{b}(1 S)$ case

...or by looking at the photon spectrum
in $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow(\pi \pi, \eta) \gamma+$ hadrons @ $\mathrm{Y}(4 \mathrm{~S}, 5 \mathrm{~S})$

PRL 109232002



## The $\eta_{b}(1 S)$ case


$\rightarrow$ No treatment of the dumping factor
$\rightarrow$ No full reconstruction!

## The $\eta_{b}(1 S)$ case


$\rightarrow$ The analysis with conversions somehow further confuses the situation

## The $\eta_{b}$ (1S) case



Quite some room for experimental improvements

## The $\eta_{b}(1 S)$ at Belle II

Luminosities and number of events

| Experiment | $\Upsilon(1 S)$ | $\Upsilon(2 S)$ | $\Upsilon(3 S)$ | $\Upsilon(4 S)$ | $\Upsilon(5 S)$ | $\Upsilon(6 S)$ | $\frac{\Upsilon(n S)}{\Upsilon(4 S)}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| CLEO | $1.2(21)$ | $1.2(10)$ | $1.2(5)$ | $16(17.1)$ | $0.1(0.4)$ | - | $23 \%$ |
| BaBar | - | $14(99)$ | $30(122)$ | $433(471)$ | $\mathrm{R}_{b}$ scan | $\mathrm{R}_{b}$ scan | $11 \%$ |
| Belle | $6(102)$ | $25(158)$ | $3(12)$ | $711(772)$ | $121(36)$ | 5.5 | $23 \%$ |
| BelleII | - | - | $300(1200)$ | $5 \times 10^{4}\left(5.4 \times 10^{4}\right)$ | $1000(300)$ | $100+400($ scan $)$ | $3.6 \%$ |

$$
\begin{aligned}
& \mathrm{Y}(3 S) \rightarrow \gamma \eta_{b}(1 S): \sim 200 k \text { evts ( } \sim 5000 \text { with conversion!) } \\
& \mathrm{Y}(3 S) \rightarrow \pi \pi \mathrm{Y}(2 S) \rightarrow \pi \pi \gamma \eta_{b}(1 \mathrm{~S}): 3000 \text { evts, no ISR background } \\
& \mathrm{Y}(4 \mathrm{~S}) \rightarrow \eta h_{b}(1 \mathrm{P}) \rightarrow(\gamma \gamma) \gamma \eta_{b}(1 \mathrm{~S}): 2.5 \text { Million events } \\
& \mathrm{Y}(5 \mathrm{~S}) \rightarrow \pi \pi \mathrm{h}_{\mathrm{b}}(\mathrm{nP}) \rightarrow \pi \pi \gamma \eta_{b}(1 \mathrm{~S}, 2 \mathrm{~S}): 125 \mathrm{k} \text { each }
\end{aligned}
$$



## Ground state exclusive decays

## Phys. Rev. Lett. 119, 252001 (2017)




With $50 \mathrm{ab}^{-1}$ of $\mathrm{Y}(4 \mathrm{~S})$ Belle II can measure $\eta_{\mathrm{b}}(1 \mathrm{~S}) \rightarrow \gamma \gamma$ with $\sim 20 \%$ uncertainty

## Di-pion transitions



What we are missing:
$\mathrm{Y}(5 \mathrm{~S}) \rightarrow \eta \mathrm{hb}(1 \mathrm{P})$
$Y(5 S) \rightarrow \eta \mathrm{hb}(2 \mathrm{P}) \quad$ Belle II
$\mathrm{Y}(4 \mathrm{~S}) \rightarrow \pi \pi \mathrm{hb}(1 \mathrm{P})$
$\chi_{b}(3 \mathrm{P}) \rightarrow \omega \mathrm{Y}(1 \mathrm{~S}) \quad$ Belle II (?)
$\chi_{b}(2 \mathrm{P}) \rightarrow \eta \eta_{b}(1 \mathrm{~S})$ Belle II (?)
$\begin{array}{ll}Y(1 D) \rightarrow \eta Y(1 S) & \text { Belle II } \\ Y(3 S) \rightarrow \eta\end{array}$
$\mathrm{Y}(3 \mathrm{~S}) \rightarrow \boldsymbol{\eta} \mathrm{Y}(1 \mathrm{~S}) \quad$ Belle II
All the hadronic transitions from
Spin singlets states are unknown
$\rightarrow$ Any theoretical prediction is welcome!

## Di-pion transitions: a family picture

PRD76 072001 (2007)

0.280 .290 .30 .310 .320
$M_{\pi \pi}=\sqrt{q^{2}}\left(\mathrm{GeV} / \mathrm{c}^{2}\right)$


PoS (hadron2017) 040


Phys.Rev. D91 (2015) no.7, 072003


## Di-pion transitions

Exotic stats contribute to the hadronic and radiative transitions from narrow quarkonia
Y.H. Chen et al, PRD93 (2016) 034030


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## To-do list

Full amplitude analysis of
$\mathrm{Y}(3 \mathrm{~S}) \rightarrow \pi \pi \mathrm{Y}(1 \mathrm{~S}), \mathrm{Y}(2 \mathrm{~S})$


$\chi_{b}(2 \mathrm{P}) \rightarrow \pi \pi \chi_{b}(1 \mathrm{P})$

LHCb and CMS could contribute here!

Two (or three) ideas to exploit the $Y(n S)$ annihilations

## $Y(n S)$ annihilations

$\mathrm{Y}(\mathrm{nS})$ annihilations into hadrons are quite peculiar and not very well know

## 1) Baryon and strangeness enhancement

PRD76 012005 (2007)

CLEO absolute yields

| Particle Type | $(\mathrm{ggg}) /(q \bar{q})[$ Data] |
| :---: | :---: |
| $\Lambda$ | $(873600 \pm 1400) /(107300 \pm 600)$ |
| $p$ | $(1399800 \pm 1200) /(295900 \pm 500)$ |
| $\bar{p}$ | $(1359500 \pm 1200) /(285400 \pm 500)$ |
| $\phi$ | $(227900 \pm 1600) /(48300 \pm 800)$ |
| $f_{2}(1270)$ | $(193000 \pm 4000) /(66500 \pm 1800)$ |

$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{Y}(1 \mathrm{~S}, 2 \mathrm{~S}, 3 \mathrm{~S})$ is a
"low momentum hyperon factory"


## $Y(n S)$ annihilations

$\mathrm{Y}(\mathrm{nS})$ annihilations into hadrons are quite peculiar and not very well know

## 1) Baryon and strangeness enhancement

PRD76 012005 (2007)

## 1) Production of nuclei

Phys.Rev. D89 (2014) no.11, 111102

| Process | Rate |
| :--- | :--- |
| $\mathcal{B}(\Upsilon(3 S) \rightarrow \bar{d} X)$ | $\left(2.33 \pm 0.15_{-0.28}^{+0.31}\right) \times 10^{-5}$ |
| $\mathcal{B}(\Upsilon(2 S) \rightarrow \bar{d} X)$ | $\left(2.64 \pm 0.11_{-0.21}^{+0.26}\right) \times 10^{-5}$ |
| $\mathcal{B}(\Upsilon(1 S) \rightarrow \bar{d} X)$ | $\left(2.81 \pm 0.49_{-0.24}^{+0.20}\right) \times 10^{-5}$ |
| $\sigma\left(e^{+} e^{-} \rightarrow \bar{d} X\right)[\sqrt{s} \approx 10.58 \mathrm{GeV}]$ | $\left(9.63 \pm 0.41_{-1.01}^{+1.17}\right) \mathrm{fb}$ |
| $\frac{\sigma\left(e^{+} e^{-} \rightarrow \bar{d} X\right)}{\sigma\left(e^{+} e^{-} \rightarrow \text { Hadrons }\right)}$ | $\left(3.01 \pm 0.13_{-0.31}^{+0.37}\right) \times 10^{-6}$ |

Anti-deuteron is 10 times more abundant in $\mathrm{Y}(\mathrm{nS}) \rightarrow$ ggg than in $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{q} \overline{\mathrm{q}}$ at the same energy

## Idea nr. 1: Y(nS) for exotic charmonia

Lots of observation of exotica, but quite few completely independent confirmations
$\rightarrow$ Only X(3872) has been seen in prompt production (in $\overline{\mathrm{p}}$ and pp collisions)

Based on Phys. Rev. D 93, 112013 [Belle]


## Idea nr. 1: Y(nS) for exotic charmonia

A tentative comparison between Belle and CMS.


Belle II prospects with $300 \mathrm{fb}^{-1}$ :
$\rightarrow$ 3-5 $\times$ sensitivity in inclusive production from $Y(3 S)$

$$
\mathrm{B}[\mathrm{Y}(\mathrm{nS}) \rightarrow \mathrm{X}(3872)+\mathrm{had}] / \mathrm{B}\left[\mathrm{Y}(\mathrm{nS}) \rightarrow \psi^{\prime}+\mathrm{had}\right]>7 \%
$$

$\rightarrow 10-15 \times$ sensitivity in double charmonium

## Idea nr. 1: Y(nS) for exotic charmonia

BaBar measured a reasonably high production of $D^{*}$ from $\mathrm{Y}(1 \mathrm{~S})$ annihilations
$\mathrm{B}\left[\mathrm{Y}(\mathrm{nS}) \rightarrow \mathrm{D}^{*}+\mathrm{X}\right]=2.5 \%$

Belle II will have:
$\rightarrow \sim 10 x$ the data
$\rightarrow$ Better efficiency at low momenta


FIG. 3: Reconstruction efficiency for the decay chain $\Upsilon(2 S) \rightarrow \pi^{+} \pi^{-} \Upsilon(1 S), \Upsilon(1 S) \rightarrow D^{* \pm} X$ as a function of the scaled $D^{* \pm}$ momentum $\mathrm{x}_{\mathrm{p}}$.

## Idea nr. 1: Y(nS) for exotic charmonia

## BaBar measured a reasonably high production of $\mathrm{D}^{*}$ from $\mathrm{Y}(1 \mathrm{~S})$ annihilations

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$$

- Production at Colliders speaks against extended objects;
- using Pythia to estimate the probability to find a D-Dbar pair in the relevant phase space, factors of $10^{-2}$ with respect to the X (3872) cross section measured by CDF ( $\sim 30 \mathrm{nb}$ ) are found.
L. Maiani's talk

$\rightarrow$ We can aim for associated DD* and (maybe) DD* correlations
$\rightarrow$ And if we actually observe also the X(3872)...


## Idea nr. 2: Bottomonium for astrophysics

$\overline{\mathrm{d}}$ detection in cosmic rays is considered since long a probe

## for low or intermediate mass WIMPs

$\rightarrow$ it's kinematically easier to produce a d from $\chi \chi$ annihilation than from SM processes


Nuclear uncertainties
$\rightarrow \overline{\mathrm{p}}$ and $\overline{\mathrm{n}}$ production rates rel. uncertainty $\sim 10$
$\rightarrow \overline{\mathrm{d}}$ production model rel. uncertainty $\sim \sim 50-200$


Astrophysical uncertainties
$\rightarrow$ Galactic density profile

$$
\text { rel. uncertainty } \sim 20
$$

$\rightarrow$ Transport models

$$
\text { rel. uncertainty } \sim 500
$$

## Idea nr. 2: Bottomonium for astrophysics

$\rightarrow$ Anti-deuteron production is described by p-n coalescence Aramaki et al. Phys. Rept. 618 (2016) 1-37 models tuned on the HEP data

$$
\frac{\mathrm{d} N_{\bar{d}}}{\mathrm{~d} T_{\bar{d}}}=\frac{\sum_{d}^{3}}{6} \frac{m_{\bar{d}}}{m_{\bar{n}} m_{\bar{p}}} \frac{1}{\sqrt{T_{\bar{d}}^{2}+2 m_{\bar{d}} T_{\bar{d}}}} \frac{\mathrm{~d} N_{\bar{n}}}{\mathrm{~d} T_{\bar{n}}} \frac{\mathrm{~d} N_{\bar{p}}}{\mathrm{~d} T_{\bar{p}}}
$$

$\rightarrow$ Most recent data are from Alice (large final state, MC-driven correction)
$\rightarrow$ Strong need to further constrain the d production model (new AMS-02 data are coming, few $\overline{\mathrm{He}} 3$ could have been observed )


## Idea nr. 2: Bottomonium for astrophysics

Use the Belle II data to investigate the basic mechanism for d production
$\rightarrow$ No final state interaction (complementarity with Alice)
$\rightarrow$ Better particle identification than Belle and BaBar
$\rightarrow$ Collect $\sim 30000 \overline{\mathrm{~d}}$, with dedicated tracking and PID
$\rightarrow$ Is coalescence really the whole story?

Need for theoretical models!
$\rightarrow$ d production models are made for HIC!


## Idea nr. 3: Hyperon-Hyperon interactions

Two results from Belle:

Near-threshold enhancement in exclusive Y annihilations


No sign of weakly bound H-dibaryon


Rough extrapolation for $300 \mathrm{fb}^{-1} \mathrm{Y}(3 \mathrm{~S})$ $\sim 60$ Million events with one $\Lambda$ or $\bar{\Lambda}$ ~3 Million events with one $\Lambda \bar{\Lambda}$ pair
$\rightarrow$ High statistics study near threshold enhancement
$\rightarrow$ Stable H di-baryon in missing mass
$\rightarrow$ Extract the $\Lambda \Lambda$ potential from correlations?
$\rightarrow$ Need for theoretical input on the $\Lambda \Lambda$ correlations in a small volume

## Conclusions

This talk was extremely incomplete
$\rightarrow$ Lots of topics has been neglected
The next years can represent our last chance to fully investigate the bottomonium spectrum
$\rightarrow$ Again, lots of topics has been neglected
Hadronic annihilations are a bridge between sectors of low energy QCD that we should exploit

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## Backup

## The uniqueness of quarkonia: the $X$ (3915) saga

Belle 2017: New analysis of $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \mathrm{J} / \psi \mathrm{D}^{0} \overline{\mathrm{D}}^{0}: \mathbf{X}(\mathbf{3 8 6 0})$


X(3915)
X(3860)

1) Dominant decay to $D^{0} \bar{D}^{0}$
2) $\mathrm{Be} 80-120 \mathrm{MeV}$ below $\chi_{\mathrm{cJ2}}(2 \mathrm{P})$
3) $\mathcal{B}\left(\chi_{c 0}^{\prime} \rightarrow \omega J / \psi\right)<7.8 \%$.


## The low-energy radiative transitions

The M1 radiative transition $\Upsilon(1 S) \rightarrow \eta_{b}(1 S) \gamma$ is the unique electromagnetic decay of $\Upsilon(1 S)$ state, which has no experimental information until now.


This is going to be impossible unfeasible just a nightmare very challenging...
... This of course doesn't mean that we are never going to try do to this analysis

## Y(3S): precision spectroscopy

Belle II

The components of the $\mathrm{Y}(1 \mathrm{D})$ triplet have not been disentagled yet


Yield per $10^{9}$ $Y(3 S)$ decays

- $2.4 \mathrm{k}^{3} \mathrm{D}_{1}$
- $19 \mathrm{k}^{3} \mathrm{D}_{2}$
- $6.8 \mathrm{k}^{3} \mathrm{D}_{3}$

Godfrey and Moats, PRD 92, 054034 (2015)


Belle II prospects with $300 \mathrm{fb}^{-1}$ :
$\rightarrow$ Separate the components of the 1D triplet
$\rightarrow$ (not shown here) $\eta_{b}(1 S)$ line-shape measurement

## Summary and readiness: bottomonium

## Competition and complementarity

- No other experiment, running or planned, can address the open topics in bottomonium physics
- Belle II is the last chance we have to make further measurements
$->30$ unique papers with less than $1 a b^{-1}$ of data ( $Y(3 S)$ and $Y(6 S)$ only)
CMS energy requirements
- Run at $\mathrm{Y}(3 \mathrm{~S})$ (200 MeV below nominal energy)
- Run at $\mathrm{Y}(6 \mathrm{~S})$ ( 460 MeV above nominal energy)
- Run at $\mathrm{Y}(5 \mathrm{~S})$ ( 300 MeV above nominal energy)

Luminosity
$-0.3 \mathrm{ab}^{-1}$ for $\mathrm{Y}(3 \mathrm{~S})$
$-0.1 a b^{-1}$ for $\mathrm{Y}(6 \mathrm{~S})$

- $1 \mathrm{ab}^{-1}$ for $\mathrm{Y}(5 \mathrm{~S})$ (to be used for Bs physics)
$-0.4 \mathrm{ab}^{-1}$ for scans (+ possibility for $20 \mathrm{fb}^{-1}$ in Phase II)
- These luminosity request are the minimal ones needed to achieve the bottomonium physics program. Any reduction would significantly compromise parts of it, in particular the new physics searches $\mathrm{Y}(3 \mathrm{~S})$


## Triggers:

- Special trigger for $\mathrm{Y}(1 \mathrm{~S}) \rightarrow$ invisible, under development


## Summary and readiness: bottomonium

## A possible run plan

- Only an hypothesis, still to be discussed with the accelerator group
- Devote the end of each year of data taking to non- $\mathrm{Y}(4 \mathrm{~S})$ physics (few weeks / year o average, at most)


The $\mathrm{Y}(3 \mathrm{~S})$ run would require (not including the time needed to change energy and assuming no changes in the luminosity):
$\rightarrow$ few months at $0.5 \times 10^{35} \mathrm{~cm}^{-2} \mathrm{~s}^{-1} \quad(\sim$ May 2018)
$\rightarrow$ few weeks at $3 \times 10^{35} \mathrm{~cm}^{-2} \mathrm{~s}^{-1} \quad(\sim$ May 2020)
$\rightarrow$ few days at $8 \times 10^{35} \mathrm{~cm}^{-2} \mathrm{~s}^{-1} \quad(\sim$ May 2022)
$\mathrm{Y}(3 \mathrm{~S})$ data should be preferably taken at low luminosity, to fully exploit the di-pion trigger for $\mathrm{Y}(1 \mathrm{~S}) \rightarrow$ invisible

The $Y(6 S)$ scan would require $\sim 2$ months independently from the luminosity ( 40 points, $10 \mathrm{fb}^{-1}$ each)

The $\mathrm{Y}(6 \mathrm{~S})$ on-resonance would require few days. Possibly split it in $10 \mathrm{fb}^{-1}$ at the very beginning of phase III and the rest afterwards ?

## Charmonium at Belle II

## Competition and complementarity

- LHCb and BESIII run a parallel program in charmonium physics
- Competition for the vector states (BESIII) and for the $\mathrm{B} \rightarrow$ (cc) K (LHCb)
- Unique topics: double charmonium (cross section, absolute BF, spectroscopy),

$$
\gamma \rightarrow \mathrm{c} \bar{c} \quad \text { (form factors, spectroscopy) }
$$

## CMS energy requirements

- The charmonium physics program is part of the $\mathrm{Y}(4 \mathrm{~S})$ physics program
- Double charmonium, $\gamma \gamma$ fusion and the ISR program can take place at any energy


## Luminosity

- No tight requirement for $\gamma \gamma \rightarrow \mathrm{cc}$, precise results starting from $10 \mathrm{ab}^{-1}$
- As much as possible for double charmonium
- Crucial for ISR and $B \rightarrow \bar{c} \bar{K}$. Running 6 month/year would pose us significantly beyond BESIII and LHCb.


## Triggers:

- No need for specific triggers, all the final states have several charged tracks


## Software:

- ISR generators (PHOKARA, KKMC, BABAYAGA...) are part of the generator package
- No need for dedicated analysis or fitting tools


## Charmonium in ISR

| Golden Channels | $E_{c . m .}(\mathrm{GeV})$ | Statistical error (\%) | Related $X Y Z$ states |  |
| :---: | :---: | :---: | :---: | :---: |
| $\pi^{+} \pi^{-} J / \psi$ | 4.23 | $7.5(3.0)$ | $Y(4008), Y(4260), Z_{c}(3900)$ |  |
| $\pi^{+} \pi^{-} \psi(2 S)$ | 4.36 | $12(5.0)$ | $Y(4260), Y(4360), Y(4660), Z_{c}(4050)$ |  |
| $K^{+} K^{-} J / \psi$ | 4.53 | $15(6.5)$ | $Z_{c s}$ |  |
| $\pi^{+} \pi^{-} h_{c}$ | 4.23 | $15(6.5)$ | $Y(4220), Y(4390), Z_{c}(4020), Z_{c}(4025)$ |  |
| $\omega \chi_{c 0}$ | 4.23 | $35(15)$ | $Y(4220)$ |  |
|  |  |  |  |  |




## Belle II prospects:

$\rightarrow$ At $50 \mathrm{ab}^{-1}$, Belle II would match BESIII on a wider spectrum
$\rightarrow$ Line-shape of the $\mathrm{Y}(4260)$
$\rightarrow$ Strange partner of the $\mathrm{Z}(3900)$ in $\mathrm{KKJ} / \psi$
$\rightarrow$ Cross sections of exclusive $(\mathrm{c} \overline{\mathrm{c}})+$ hadrons

## Bottomonium at Belle II

Current samples in $\mathrm{fb}^{-1}$ (millions of events), and the proposal for Belle II

| Experiment | $\Upsilon(1 S)$ | $\Upsilon(2 S)$ | $\Upsilon(3 S)$ | $\Upsilon(4 S)$ | $\Upsilon(5 S)$ | $\Upsilon(6 S)$ | $\frac{\Upsilon(n S)}{\Upsilon(4 S)}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
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- Narrow states spectroscopy (Y(1D), $\left.\chi_{b}(\mathrm{nP}) \ldots\right)$


## - Exotica as virtual contributions to transitions

- Precision NRQCD test
- New Physics (DM / light higgs)
- Missing hadronic and radiative transitions
- Baryon physics (inc. correlations)
- Anti-nuclei production (with DM applications)
- Gluon fragmentation
- Inclusive charmonium production and $\mathrm{D} \overline{\mathrm{D}}$ correlations
- Bs physics
- Exotica discovery
- Precision Zb mass measurement
- Missing hadronic and radiative transitions
- Light meson spectroscopy in transitions


## Accelerator requirements for bottomonium

Beam Transport Final Focus



## $Y(6 S)$ on-resonance run: conventional

$\rightarrow Y(5 S)-Y(6 S)$ are portals to the missing narrow states
$\rightarrow \mathrm{Y}(5 \mathrm{~S}) \rightarrow \eta \mathrm{Y}(1 \mathrm{D})$ is the largest $\mathrm{Y}(5 \mathrm{~S})$, single-meson transition
$\rightarrow$ The conventional spectrum gets contributions from the couple channel effect (again, light quarks...)


Mod. Phys. Lett. A 32, 1750025 (2017)

| Name | $L$ | $S$ | $J^{P C}$ | Emitted hadrons [Threshold, $\left.\mathrm{GeV} / c^{2}\right]$ |
| :---: | :---: | :---: | :---: | :--- |
| $\eta_{b}(3 S)$ | 0 | 0 | $0^{-+}$ | $\omega[11.12], \phi[11.36]$ |
| $h_{b}(3 P)$ | 1 | 0 | $1^{+-}$ | $\pi^{+} \pi^{-}[10.82], \eta[11.09], \eta^{\prime}[11.50]$ |
| $\eta_{b 2}(1 D)$ | 2 | 0 | $2^{-+}$ | $\omega[10.93], \phi[11.17]$ |
| $\eta_{b 2}(2 D)$ | 2 | 0 | $2^{-+}$ | $\omega[11.23], \phi[11.47]$ |
| $\Upsilon_{J}(2 D)$ | 2 | 1 | $(1,2,3)^{--}$ | $\pi^{+} \pi^{-}[10.73], \eta[11.00], \eta^{\prime}[11.41]$ |
| $h_{b 3}(1 F)$ | 3 | 0 | $3^{+-}$ | $\pi^{+} \pi^{-}[10.63], \eta[10.90], \eta^{\prime}[11.31]$ |
| $\chi_{b J}(1 F)$ | 3 | 1 | $(2,3,4)^{++}$ | $\omega[11.14], \phi[11.38]$ |
| $\eta_{b 4}(1 G)$ | 4 | 0 | $4^{-+}$ | $\omega[11.31], \phi[11.55]$ |
| $\Upsilon_{J}(1 G)$ | 4 | 1 | $(3,4,5)^{--}$ | $\pi^{+} \pi^{-}[10.81], \eta[11.08], \eta^{\prime}[11.49]$ |

## Belle II goals:

$\rightarrow$ Search for new, predicted, resonances
$\rightarrow$ Use both single transitions and double cascades
$\rightarrow$ Fill the remaining spectrum to measure the effects of the coupled channels contributions

## Y(5S): Zb masses

The measurement of the Zb masses is foundamental to determine their nature: are they above or below the $\mathrm{B}^{(*)} \mathrm{B}^{*}$ thresholds?
$\rightarrow$ Equivalent to the $\mathrm{X}(3872)$ mass problem: above or below the open threshold?


Current best estimate of the Zb location with respect to the thresholds:

$$
\begin{aligned}
\varepsilon_{B}\left(Z_{b}\right) & =\left(0.60_{-0.49}^{+1.40} \pm i 0.02_{-0.01}^{+0.02}\right) \mathrm{MeV} \\
\varepsilon_{B}\left(Z_{b}^{\prime}\right) & =\left(0.97_{-0.68}^{+1.42} \pm i 0.84_{-0.34}^{+0.22}\right) \mathrm{MeV}
\end{aligned}
$$

## Belle II Goals:

$\rightarrow$ Determine if the Zb are located above or below the open flavour threshold using $1 \mathrm{ab}^{-1}$ of $\mathrm{Y}(5 \mathrm{~S})$

## $Y(5 S)-Y(6 S)$ scan



$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \pi^{+} \pi^{-} \mathrm{h}_{\mathrm{b}}(1 \mathrm{P})$
$\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \pi^{+} \pi^{-} \mathrm{h}_{\mathrm{b}}(1 \mathrm{P})$

## Belle II scan goal:

$\rightarrow$ Investigate the presence of a broad resonance at 10.750 GeV
$\rightarrow 10 \mathrm{MeV}$ wide steps, $10 \mathrm{fb}^{-1}$ each ( $10 \times$ Belle scan)
$\rightarrow Y(5 S)$ and $Y(6 S)$ line-shapes in $R, R_{Y \pi \pi}$ and $R_{h \pi \pi}$
$\rightarrow \mathrm{Rb}$ decomposition ( $\mathrm{BB}, \mathrm{BB}^{*}, \mathrm{~B}^{*} \mathrm{~B}^{*}, \mathrm{BB}^{*} \pi, \mathrm{~B}^{*} \mathrm{~B}^{*} \pi$, $\mathrm{BsBs} \ldots$ )
$\rightarrow$ Overall goal: settle the nature of the $\mathrm{Y}(5 \mathrm{~S})$

## $Y(3 S):$ rare $\chi_{b}$ decays

 $\chi_{b}(2 \mathrm{P}) \rightarrow \tau \tau$ is sensitive to the presence of a CP-even light Higgs (as $\mathrm{B} \rightarrow \tau \tau, \mathrm{B} \rightarrow \tau \vee \ldots$ )$$
\left.\begin{array}{l}
\mathrm{BR}^{H}\left(\chi_{b 0}(1 P) \rightarrow \tau \tau\right)=3.1 \times 10^{-13} \\
\mathrm{BR}^{H}\left(\chi_{b 0}(2 P) \rightarrow \tau \tau\right)=(1.9 \pm 0.5) \times 10^{-12}
\end{array}\right\} \times\left[1+\frac{M_{H_{125}}^{2} \tan ^{2} \beta}{M_{\mathrm{new}}^{2}-M_{\chi_{b 0}}^{2}}\right]^{2}
$$

Will only need $\left(M_{H_{125}} / M_{H_{\text {new }}}\right) \tan \beta \sim 30$ for $\mathcal{O}(100)$ signal events in $\gamma(3 S) \rightarrow \gamma \chi_{b 0}(2 P) \rightarrow \gamma \tau \tau$

## Results: $\Upsilon(3 S)$



## Belle II goals:

$\rightarrow \chi_{\mathrm{b} 0}(2 \mathrm{P}, 1 \mathrm{P}) \rightarrow \gamma \tau \tau$ inclusive
$\rightarrow \chi_{\mathrm{b} 0}(2 \mathrm{P}, 1 \mathrm{P}) \rightarrow \gamma \tau \tau$ per exclusive final state
$\rightarrow$ MC studies ongoing

## $Y(1 S) \rightarrow$ invisible

$\mathrm{Y}(1 \mathrm{~S}) \rightarrow$ invisible is well calculable in the SM

$$
\begin{gathered}
\frac{B R(Y(1 S) \rightarrow v \bar{v})}{B R\left(Y(1 S) \rightarrow e^{+} e^{-}\right)}=\frac{27 G^{2} M_{Y(1 S)}^{4}}{64 \pi^{2} \alpha^{2}}\left(-1+\frac{4}{3} \sin ^{2} \theta_{W}\right)^{2}=4.14 \times 10^{-4} \\
B R(Y(1 S) \rightarrow v \bar{v}) \sim 9.9 \times 10^{-6}
\end{gathered}
$$

Non-SM contributions from $\mathrm{Y}(1 \mathrm{~S}) \rightarrow \chi \chi$



BaBar, Phys. Rev. Lett. 103, 251801 (2009)


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\end{gathered}
$$

Non-SM contributions from $\mathrm{Y}(1 \mathrm{~S}) \rightarrow \chi \chi$

Belle: Phys.Rev.Lett. 98 (2007) 132001

| Source | $(\%)$ |
| :--- | :---: |
| Track selection | 5.6 |
| $\pi^{0}$ veto | 2.4 |
| Fisher discriminant | 6.1 |
| Other selection requirements | 1.1 |
| $\Upsilon(3 S) \rightarrow \pi^{+} \pi^{-} \Upsilon(1 S)$ | $7.6 \quad 4 \%$ in BaBar |
| Trigger efficiency | 8.7 |
| Fit bias | 0.2 |
| Statistics of control sample | 1.4 |
| $\mathcal{B}\left(\Upsilon \rightarrow \mu^{+} \mu^{-}\right)$ | 2.0 |
| Total | 14.7 |

## Belle II prospects

$\rightarrow 10 \times$ dataset $\mathrm{w} /$ respect to BaBar
$\rightarrow$ Sensitivity $\sim 1 \times 10^{-4}$ on the BF
$\rightarrow$ Reduce the systematic with precision measurement of the pp and gg transitions
$\rightarrow$ Trigger is crucial: capability to trigger on $2 p+$ missing energy depends on the BG levels and luminosity

## Charmonium from B decay

3
Belle II
$\mathrm{B} \rightarrow \mathrm{K}(\mathrm{cc} \overline{\mathrm{c}}) \rightarrow \mathrm{K}($ hadrons, hadrons $+\mu \mu, \mathrm{n} \gamma+\mu \mu)$
$\rightarrow$ Competitive in neutral transitions $\left(\mathrm{Xcc}_{\bar{c}} \rightarrow \eta, \pi^{0}, \omega \mathrm{~J} / \psi\right)$
$\rightarrow$ Competitive for finals states with large multiplicities ( $h_{c}$ and $\eta_{c}$ )
$\rightarrow$ Unique opportunity for inclusive measurements

## Belle II prospects:

$\rightarrow$ Discover the $\eta_{c 2}(2 D)$, last narrow charmonium missing in $B \rightarrow K \gamma h_{c}$ $\rightarrow$ Comprehensive study of $\mathrm{B} \rightarrow \mathrm{K} D \overline{\mathrm{D}}, \mathrm{KD} \overline{\mathrm{D}}^{*}, \mathrm{KD}{ }^{*} \overline{\mathrm{D}}^{*}, \mathrm{~K} \overline{\mathrm{D}} \mathrm{D}^{* *}, \mathrm{~K} \overline{\mathrm{D}}^{*} \mathrm{D}^{* *}$



## Y(3S): $\pi \pi$ scattering length

## Q-value for $\mathrm{Y}(3 \mathrm{~S}) \rightarrow \pi \pi$

At low energy the $\pi \pi$ interaction is described by two scattering lengths who vanish in the chiral limit:
$a_{0}^{0}=\frac{7 M_{\pi}^{2}}{32 \pi F_{\pi}^{2}}+\mathcal{O}\left(m_{q}^{2}\right) \quad a_{0}^{2}=-\frac{M_{\pi}^{2}}{16 \pi F_{\pi}^{2}}+\mathcal{O}\left(m_{q}^{2}\right)$

## Weinberg, PRL17,616(1966)

Using ChPT, theory predicts:

$$
a_{0}^{0}-a_{0}^{2}=0.265 \pm 0.004
$$

Colangelo, et al, PLB488,261(2000)

$\mathrm{Y}(2 \mathrm{~S})$ is only 50 MeV
Liu et al,EPJC73, 2284 (2013)




## The $\eta_{c}$ width conundrum



## What do I understand from this?

$\rightarrow$ NNLO is still not enough
$\rightarrow$ Is NRQCD converging fast enough?
$\rightarrow$ The problem is not in the experimental resolution

A funny coincidence: what happens if we take the measurements done with M1 naive fit?

1986
2000F BES
2003
BES

$$
\begin{aligned}
& J / \psi \rightarrow \gamma \mathrm{X}, \psi(2 S) \rightarrow \gamma \mathrm{X} \\
& J / \psi \rightarrow \gamma \eta_{c} \text { and } \psi(2 S) \rightarrow \gamma \eta_{c} \\
& J / \psi \rightarrow \gamma \eta_{c}
\end{aligned}
$$

