LNF Autumn Institute II : Testing the Standard Model at low and high energies Frascati, 28 November 2017

focus on LHC, FCC, NA62

Searching for massless Dark Photons

based on :

M.Fabbrichesi, E.Gabrielli, BM, PRL 119 (2017) 031801 E.Gabrielli, BM, M. Raidal, E. Venturini, PRD 94 (2016) 115013 S.Biswas, E.Gabrielli, M.Heikinheimo, BM, PRD 93 (2016) 093001 S.Biswas, E.Gabrielli, M.Heikinheimo, BM, JHEP 1506 (2015) 102 E.Gabrielli, M.Heikinheimo, BM, M.Raidal, PRD 90 (2014) 055032



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a few facts

* expected exp hints of fashionable theory solutions to SM puzzles are being late in showing up * more and more crucial to look at signature-based BSM searches -> boosts discovery potential in a model-independent way Hidden/Dark (SM-uncharged) Sectors can provide new signatures not covered by present searches

four paths to advance in HEP today (ATLAS/CMS):

- ***** by exploring the characteristics of the Higgs sector and confirming/spoiling the SM picture (primary relevance since the Higgs sector is so critical !)
- ***** by searching for new heavy states coupled to the SM acting as a cut-off for the SM [possibly solving the naturalness issues and/or non-SM phenomena (dark matter, ...)]
- ★ by exploring ∧ >> o(1TeV) indirect effects through high-accuracy studies of SM x-sections/distributions and searches for rare processes (EFT parametrization)

by looking for new DARK states (uncoupled to the SM at tree level) either in production or/and heavy-state (t,H...) decays (elusive signatures, maybe long-lived p.les)

four paths to advance in HEP today (ATLAS/CMS):









* every single method is of fundamental importance to make progress !

- * what's peculiar to a massless DP's
- * Hidden Sectors with unbroken extra U(1)
 - predicting massless DP's
- * Higgs decays into massless DP's
- ***** new Higgs signatures from DP's at colliders
- * FCNC mediated by DP's
- * massless DP's in kaon decays

* Outlook

Dark Photons (DP) from extra U(1)'s

- Hidden Sectors can contain light or massless gauge bosons mediating long-range forces between Dark particles
- previous pheno studies mainly involving "massive" DP's
- a massive DP interacts with SM matter via "kinetic mixing" with SM hypercharge U(1)y gauge boson :

$$B_{\mu
u}~=~\partial_{\mu}B_{
u}-\partial_{
u}B_{\mu}$$
 [U(1) gauge invariant]

4D interaction between field-strengths of two different U(1) allowed \rightarrow

mixing param.

$$\mathcal{L}_{mix} = \chi B_{\mu\nu} C^{\mu\nu}$$

→ a massive DP couples to SM particles with strength $-\chi eQ_{el}$

→ quite a few exp bounds on that by now !

DP's may have a relevant role in Cosmology and Astrophysics

D.N. Spergel, P.J. Steinhardt, PRL 84 (2000) M.Vogelsberger, J.Zavala, A.Loeb, Mon.Not. Roy Astron 423 (2012) L.G. Van den Aarssen, T. Bringmann, C. Pfrommer, PRL 109 (2012) S. Tulin, H.B. Yu, K.M. Zurek, PRD 87 (2013)

possible role in galaxy formation and dynamics:

- >may solve the small-scale structure formation problems
- can explain the dark discs of galaxies

in Cosmology

J.Fan, A.Katz, L.Randall, M.Reece, PRL 110 (2013)

in Astro-particle Physics :

>may induce Sommerfeld enhancement of DM annihilation cross section

(from PAMELA-Fermi-AMS2 positron anomaly) N.Arkani-Hamed, D.P. Finkbeiner, T.R. Slatyer, N.Weiner. PRD 79 (2009)

may assist DM annihilations for the required magnitude making asymmetric DM scenarios viable

K.M. Zurek, Phys Rept. 537 (2014)

strong astrophysical and collider bounds on massive DP (Z')!

visible massive DP's decays

arXiv:1412.1485



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strategies of massive-DP (A') searches

***** Bremsstrahlung: $e^{-}Z \rightarrow e^{-}ZA'$, e- incident on a nuclear target (also $p Z \rightarrow p ZA'$) ***** Annihilation: $e+e- \rightarrow \gamma A'$ (favored for invisible A' decays) ***** Meson decays: Dalitz decays, $\pi_0/\eta/\eta \rightarrow \gamma A'$, and rare meson decays such as $K \rightarrow \pi A', \varphi \rightarrow nA', and D \rightarrow D A'$ (A' mass reach limited by parent meson mass) ***** Drell-Yan: $qq^- \rightarrow A' \rightarrow |+|-$ (or h+h-)

strategies of massive DP detection

***** Bump hunt in visible final-state invariant mass: $A' \rightarrow |+|- \text{ or } A' \rightarrow h+h-$ ***** Bump hunt in missing-mass in $e+e- \rightarrow \gamma A'$ or meson decay production channels * Vertex detection in A' \rightarrow I+I- ; $\mathcal{L}_{kin.mix.} = \frac{1}{2} \epsilon F^{\mu\nu} F'_{\mu\nu}$ A' decay length scales with $1/(\epsilon^2 m_{A'})$, \rightarrow searches for displaced vertices in visible decay modes probe the very low- ε regions of parameter space.

A' search strategies vs (ϵ^2 , $m_{A'}$)



present constraints (gray dashed area)



many ongoing and proposed experiments !

if $U(1)_F$ unbroken no such constraints !



massless DP's then interact with the SM sector only through higher-dimensional (→ suppressed by 1/M^{D-4}) interactions via messenger (if any) exchange !

→ potentially large DP couplings $\bar{\alpha}$ in the Hidden Sector (HS) allowed !

evading most of present exp bounds on massive DP's !

Explaining Yukawa hierarchy via HS and extra $U(1)_F$

Hidden Sectors (HS) possibly explaining
Flavor hierarchy + Dark Matter
Gabrie

Gabrielli, Raidal, arXiv:1310.1090

- ➢ Yukawa's are not fundamental constants
 but effective low-energy couplings
 (→ scalar messengers transfer radiatively Flavor and
 Chiral Symm. Breaking from HS fermions to SM fermions
 giving Yukawa couplings at one-loop)
- ▷ predict extra unbroken $U(1)_F$ → massless DP's
- ▶ for integer-q(dark fermions) sequence : $M_{D_f} \sim \exp(-\frac{\kappa}{q_{D_f}^2 \bar{\alpha}})$ → exponential hierarchy in M(Dark fermions)
 → exponential hierarchy in radiative Y(SM fermions)
 ▶ Dark fermions as dark-matter candidates

heavy scalar messengers $S_{L,R}$ sector

SL,R

heavy scalar messengers
 (squark/slepton-like)
 connecting SM states
 with HS states

	Fields	Spin	$SU(2)_L$	$U(1)_Y$	$SU(3)_c$	$U(1)_F$
	$\hat{S}_L^{D_i}$	0	1/2	1/3	3	- q_{D_i}
messengers	$\hat{S}_L^{U_i}$	0	1/2	1/3	3	- q_{U_i}
(Scalars)	$S_R^{D_i}$	0	0	-2/3	3	- q_{D_i}
	$S_R^{U_i}$	0	0	4/3	3	- q_{U_i}
Dark Sector	Q^{D_i}	1/2	0	0	0	q_{D_i}
(Fermions+	Q^{U_i}	1/2	0	0	0	$q_{U_{i}}$
singlet Scalar)	S_0	0	0	0	0	0

DP production mechanism ?

massless Dark Photon signature

if produced in collisions :
→ stable + noninteracting
→ neutrino-like signature

Higgs as a "source" of Dark Photons



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→ plenty of new signatures at colliders involving stable dark photons exploration just started !



mono-photon resonant signature

Higgs non-decoupling effects (just as in SM) can enhance BR

> Gabrielli,Heikinheimo, BM, Raidal, arXiv:1405.5196

Higgs momentum balanced by a massless invisible system

Biswas, Gabrielli, Heikinheimo, BM, arXiv:1503.05836

▶ in H production



	resonant mono-p	hoton signat	ture at	8 Te	V
	(A_1) 5	$0 \text{ GeV} < p_T^{\gamma} < 63 \text{ GeV}$	$V(A_2)$ 60 Ge	$V < p_T^{\gamma} < $	63 GeV
0	$a \rightarrow H \rightarrow \bar{\gamma}\gamma$		<u>σ (fb)</u>	$\sigma \times A_1$	$\sigma \times A_2$
9	9 / 11 / //	$\underline{\qquad} Signal BR_{H \to \gamma}$	$_{\gamma \gamma} = 1\%$	65	34
$F_{miss} \sim F_{v} \sim m_{H}/2$	γj		715	65	
		$\gamma Z ightarrow \gamma \iota$	$ u\overline{ u}$	157	27
		$jZ ightarrow j\nu$	$ u \overline{ u} $	63	11
M_{r}	$T = \sqrt{2p_T^{\gamma} E_T (1 - \cos \Delta \phi)}$	$W \to e i$	ソ	22	0
1000	Total background		957	103	
	1000	$S/\sqrt{S+B}$ (BR _H)	9.1	13.0	
> 500 W		$S/\sqrt{S+B}$ (BR _H -	4.6	6.9	
M_T [fb/(100 50 7	$BR_{H}^{ar{\gamma}\gamma}\!=\!5\%$	(8TeV/2	20fb ⁻¹)
do/d]	γZ		model-ind easureme	depend ent of	lent BR _{DP} !
	60 ⁻⁷⁰ 80 90 100	110 120 130			
(Bari	(parton-level analysis) M_T [GeV]	LNF, 28 November 2017			21

resonant mono-photon signature at 14TeV



σ (fb)	$\sigma \times A \ [8 {\rm TeV}]$	$\sigma \times A \ [14 \mathrm{TeV}]$
$H \rightarrow \gamma \bar{\gamma} (BR_{\gamma \bar{\gamma}} = 1\%)$	44	101
γj	63	202
new $jj \rightarrow \gamma j$	59	432
$e \rightarrow \gamma$	55	93
$W(\rightarrow \ell \nu)\gamma$	58	123
$Z(\rightarrow\nu\nu)\gamma$	102	174
total background	337	1024

MadGraph5_aMC@NLO + PYTHIA (bckgr) ALPGEN + PYTHIA (H signal)

 γj bckgr modeled on data at 8 TeV (CMS, arXiv:1507.00359 [hep-ex] (PLB))

Biswas, Gabrielli, Heikinheimo, BM, arXiv:1603.01377 (PRD)

(includes parton-shower)

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model-independent bounds @ LHC 14 TeV

(including shower effects)

arXiv:1603.01377 (PRD)

$$gg \to H \to \bar{\gamma}\gamma$$
 vs $VV \to H \to \bar{\gamma}\gamma$

$\mathrm{BR}_{\gamma\bar{\gamma}}$ (%)	$L = 100 \text{fb}^{-1}$		$\mathrm{L}{=}300\mathrm{fb}^{-1}$		$L=3ab^{-1}$	
Significance	2σ	5σ	2σ	5σ	2σ	5σ
$\mathrm{BR}_{\gamma \bar{\gamma}}(\mathrm{VBF})$	0.76	1.9	0.43	1.1	0.14	0.34
$\mathrm{BR}_{\gamma \bar{\gamma}}\left(ggF ight)$	0.064	0.16	0.037	0.092	0.012	0.029

gg fusion sensitive down to BR_{DP} ~ 10⁻⁴-10⁻³ (VBF ~10 times worse ...) Biswas, Gabrielli, Heikinheimo, BM,

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new Higgs signatures at e⁺e⁻ colliders from stable dark photons

▶ in H production :

$$e^+e^- \to H \ \bar{\gamma} \to b\bar{b} \ \bar{\gamma}$$



p_{Higgs} balanced by a massless invisible system

Biswas, Gabrielli, Heikinheimo, BM, arXiv:1503.05836 (JHEP)

in H decays :



 $e^+e^- \rightarrow ZH \rightarrow Z \gamma \bar{\gamma}$

(photon + E_{miss}) resonant signature

Biswas, Gabrielli, Heikinheimo, BM, arXiv:1703.00402 (PRD)

 $e^+e^- \to H \ \bar{\gamma} \to bb \ \bar{\gamma}$



FCC-ee (~ ILC) : $\int S = 240$ GeV with $\int L \sim 10$ ab⁻¹

we assume: b-tagging efficiency of 80% fake b-jet rejection factor 1/100



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 $e^+e^- \to ZH \to \mu^+\mu^- \gamma\bar{\gamma} \ (q\bar{q} \ \gamma\bar{\gamma})$



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hadronic channel

$e^+e^- \to ZH \to qq \ \gamma \bar{\gamma}$

$\int S = 240 \text{ GeV} \quad \int L \sim 10 \text{ ab-1}$



Biswas, Gabrielli, Heikinheimo, BM, arXiv:1703.00402 (PRD)

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 $e^+e^- \to ZH \to \mu^+\mu^- \gamma\bar{\gamma} \ (q\bar{q} \ \gamma\bar{\gamma})$

 $\int S = 240 \, GeV \, \int L \sim 10 \, ab-1$

hadronic Z most sensitive channel :

5σ sensitivity for **BR**_{DP} ~ 2x10⁻⁴ **!!!**



testing BSM via FCNC top interactions



***** LHC more and more a top factory \rightarrow great opportunity ! order of magnitude × improvement at HL-LHC huge gain at future colliders !

top FCNC's mediated by massless Dark Photons



also : $b \to s \, \bar{\gamma}$

vs $b \rightarrow s \gamma$

 $\mathbf{S}_{\mathbf{m}}^{\mathrm{f}_{\mathrm{i}}}$ $\mathbf{f}_{\mathbf{R}}^{\mathbf{J}}$ $\mathbf{Q}^{\mathbf{f}_i}$

 $t \to (c, u) \bar{\gamma}$

new heavy states in loops contribute with same flavor matrix (but different U(1) charges) to FCNC decays into photon and dark photon

$$BR(t \to (c, u) \bar{\gamma}) = \frac{\bar{\alpha}}{\alpha} \left(\frac{q_3^U f_2(x_3^U, \xi_U)}{e_U \bar{f}_2(x_3^U, \xi_U)} \right)^2 BR(t \to (c, u) \gamma)$$

LHC (present bounds):

 $\mathrm{BR}^{(t \to u \gamma)}(t \to u \, \bar{\gamma}) \ < \ 1.8 \times 10^{-2} \left(\frac{\bar{\alpha}}{0.1} \right)$ $BR^{exp}(t \rightarrow u \gamma) < 1.3 \times 10^{-4}$ $BR^{(t\to c\gamma)}(t\to c\,\bar{\gamma}) < 2.3\times 10^{-1}\left(\frac{\bar{\alpha}}{0.1}\right)$ $\mathrm{BR}^{\mathrm{exp}}(t \to c \gamma) < 1.7 \times 10^{-3}$

> but imposing vacuum-stability and dark-matter bounds gives BR($t
> ightarrow q \, ar{\gamma}$) < 10^{-4}

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further upper bounds from $f
ightarrow f' \gamma$ constraints

$$BR^{exp}(\bar{B} \to X_S \gamma) = (3.43 \pm 0.21 \pm 0.07) \times 10^{-4}$$
$$BR^{(b \to s\gamma)}(b \to s\bar{\gamma}) < 8.5 \times 10^{-3} \left(\frac{\bar{\alpha}}{0.1}\right)$$

$$BR(\tau^- \to e^- \gamma) < 3.3 \times 10^{-8}$$
$$BR(\tau^- \to \mu^- \gamma) < 4.4 \times 10^{-8}$$

$$BR^{(\tau \to \mu \gamma)}(\tau \to \mu \bar{\gamma}) < 5.9 \times 10^{-6} \left(\frac{\alpha}{0.1}\right)$$
$$BR^{(\tau \to e\gamma)}(\tau \to e \bar{\gamma}) < 1.1 \times 10^{-5} \left(\frac{\bar{\alpha}}{0.1}\right)$$

Gabrielli, BM, Raidal, Venturini



\rightarrow at the LHC new FCNC signatures in BOTH top decay AND top production



signature in Kaon physics

Fabbrichesi, Gabrielli, BM, *Phys. Rev. Lett.* **119** (2017) 031801 [arXiv:1705.03470]

massless dark photon

unbroken U(1) symmetry









$$K^+ \to \pi^+ \bar{\gamma}$$





simplified model of dark sector



$\begin{array}{ll} \mbox{matrix element} & \hat{M} \equiv \langle \bar{\gamma} \ \pi^+ \pi^0 | \ \mathcal{H}_{eff}^{\Delta S=1} \ | K^+ \rangle \\ \\ \mathcal{H}_{eff}^{\Delta S=1} = \frac{e_D}{64 \ \pi^2} \ \hat{\xi} \ \hat{Q} & \hat{Q} = (\bar{s} \ \sigma^{\mu\nu} \ d) \ \bar{F}_{\mu\nu} \\ \\ \end{array}$ $\begin{array}{l} \mbox{matching scale} \rightarrow \mbox{mass of lightest-messenger and dark-fermion} \end{array}$

atching scale → mass of lightest-messenger and dark-fermion (assumed degenerate)

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chiral quark model

(quarks are coupled to hadrons by an effective interaction so that matrix elements can be evaluated by loop diagrams)



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<u>constraints</u>:

from extra contributions to KºKº mixing

astrophysics

star cooling by emission of dark photons

 $K^+ \to \pi^+ \pi^0 \bar{\gamma}_{\rm 300}$ **BR** contour $BR(K^+ \to \pi^+ \pi^0 \bar{\gamma}) \lesssim 1.6 \times 10^{-7}$ 10¹³ K⁺ at NA62 soon with hermetic photon coverage and good missing-mass resolution (under consideration...)



TABLE I. In the first two rows, relevant operators are numbered according to the notation in [20,21]. The matrix elements $\langle K^0 | Q_i | \bar{K}^0 \rangle$ (in the vacuum insertion approximation for the renormalized operators Q_i at the low energy scale $\mu = 2$ GeV) are given in the third row multiplied by the respective bag factors $B_i(\mu)$ [21] evaluated at same scale, with $X_K(\mu) = (m_K/(m_d(\mu) + m_s(\mu)))^2$. The fourth row gives the Wilson coefficients at the matching scale (the common factor at the matching being $C^2 = \xi^2/(16\pi^2\Lambda^2)$ [9], where $\xi = g_L g_R/2$). Following [21], we take $m_d(\mu) = 7$ MeV, $m_s(\mu) = 125$ MeV, $m_K = 497$ MeV, $f_K = 160$ MeV, and $B_{1,2,3,4,5}(\mu) = 0.60, 0.66, 1.05, 1.03, 0.73$, respectively.

Q_1, \tilde{Q}_1	Q_2, \tilde{Q}_2	Q_3, \tilde{Q}_3	Q_4	Q_5
$\bar{d}^{\alpha}_{L}\gamma_{\mu}s^{\alpha}_{L}\bar{d}^{\beta}_{L}\gamma_{\mu}s^{\beta}_{L}, (L \leftrightarrow R)$	$\bar{d}^{\alpha}_{R}s^{\alpha}_{L}\bar{d}^{\beta}_{R}s^{\beta}_{L}, (L \leftrightarrow R)$	$\bar{d}^{\alpha}_{R}s^{\beta}_{L}\bar{d}^{\beta}_{R}s^{\alpha}_{L}, (L \leftrightarrow R)$	$ar{d}^{lpha}_R s^{lpha}_L ar{d}^{eta}_L s^{eta}_R$	$\bar{d}^{lpha}_R s^{eta}_L \bar{d}^{eta}_L s^{lpha}_R$
$1/3m_K f_K^2 B_1(\mu)$	$-5/2X_K m_K f_K^2 B_3(\mu)$	$1/24X_K m_K f_K^2 B_3(\mu)$	$1/4X_K m_K f_K^2 B_4(\mu)$	$1/12X_K m_K f_K^2 B_5(\mu)$
$-1/24C^{2}$	0	$1/12C^{2}$	$1/6C^{2}$	$1/6C^2$

2.

$$\hat{M} = \frac{M(z_1, z_2)}{m_K^3} \varepsilon_{\mu\nu\rho\sigma} q_1^{\nu} q_2^{\rho} k^{\sigma} \varepsilon^{\mu}(k)$$

^{3.} <u>main sources of uncertainty</u>

—Chiral perturbation theory higher order corrections
 —Chiral-quark-model M-parameter dependence
 —RGE corrections

Outlook

- * massless DP's theoretically appealing
 - (evading most of present exp bounds on massive DP's !)
- * Higgs boson as the SM portal to DP's
 - ***** new effective vertices for DP's
 - (from Hidden Sectors explaining Flavor Hierarchy + Dark Matter)
- ***** rich phenomenological implications @ LHC and ee colliders
- * new class of FCNC signatures from top, b, c, s, tau, mu decays into a massless DP

* very distinctive \rightarrow bounds expected to be limited just by statistics ! * look for $K^+ \rightarrow \pi^+ \pi^0 \bar{\gamma}$ at NA62 !

***implications for astro-part/cosmology** (mostly yet to work out !)