

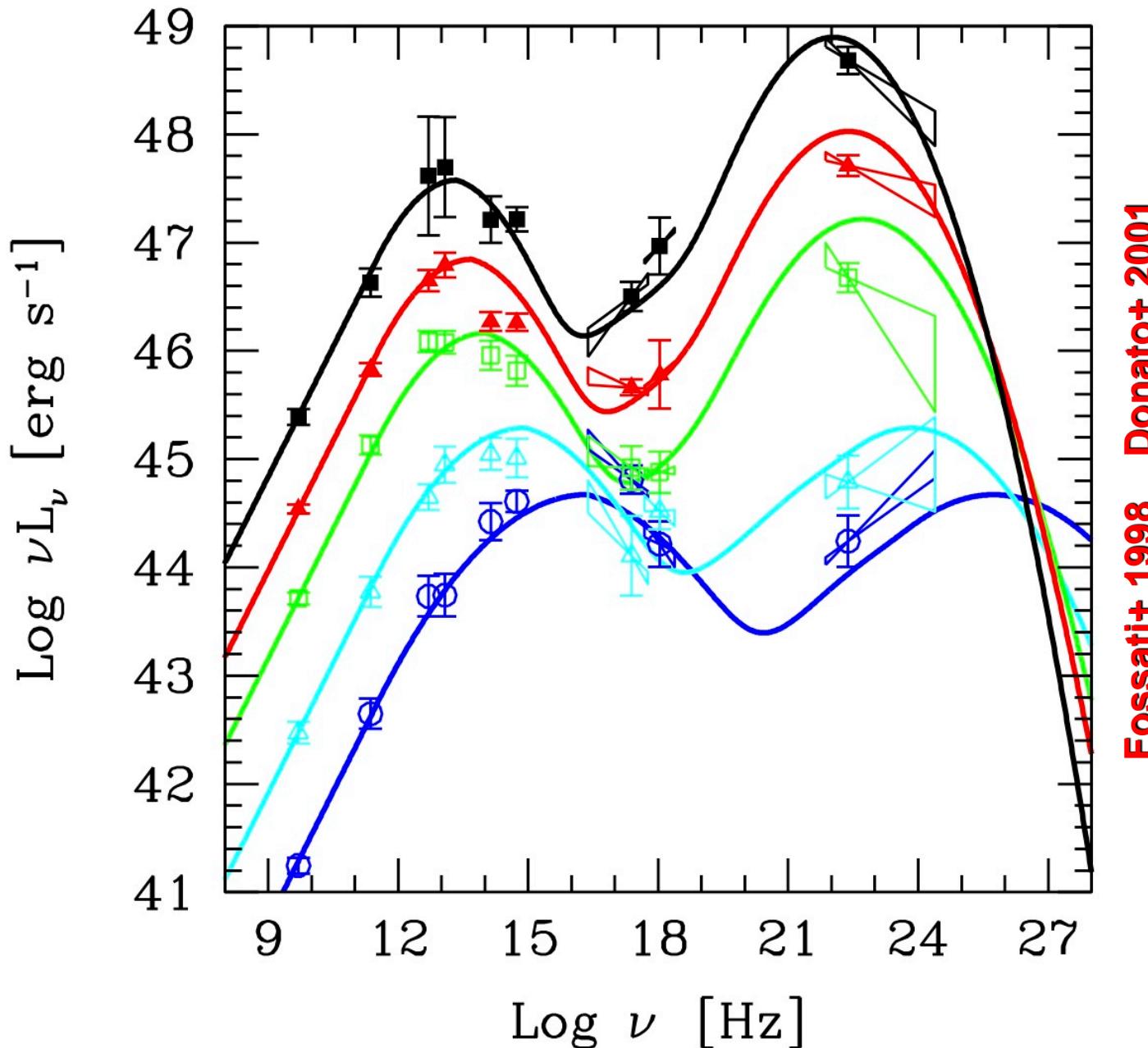
# The blazar sequence is alive

Gabriele Ghisellini

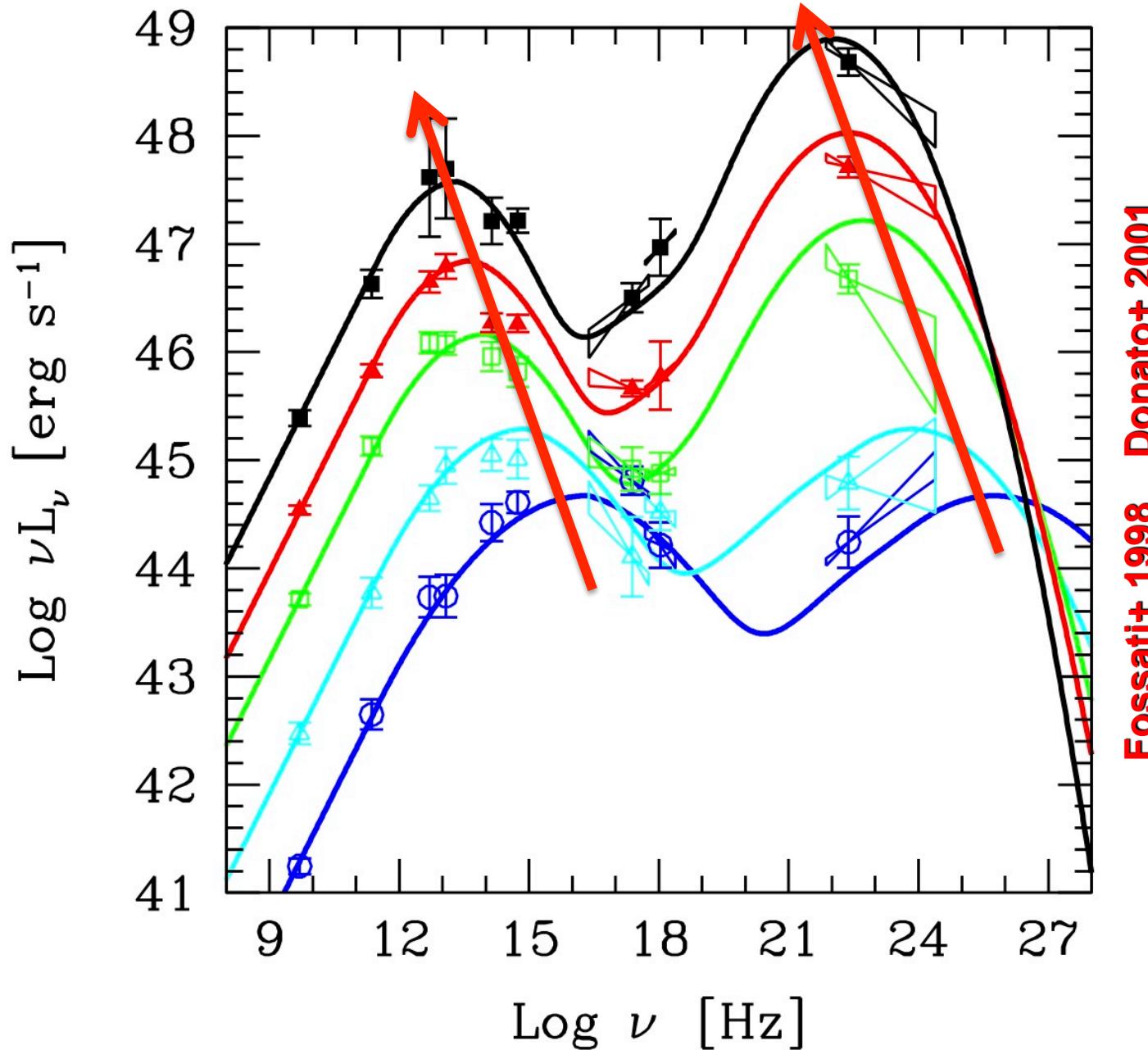
INAF - Osservatorio di Brera

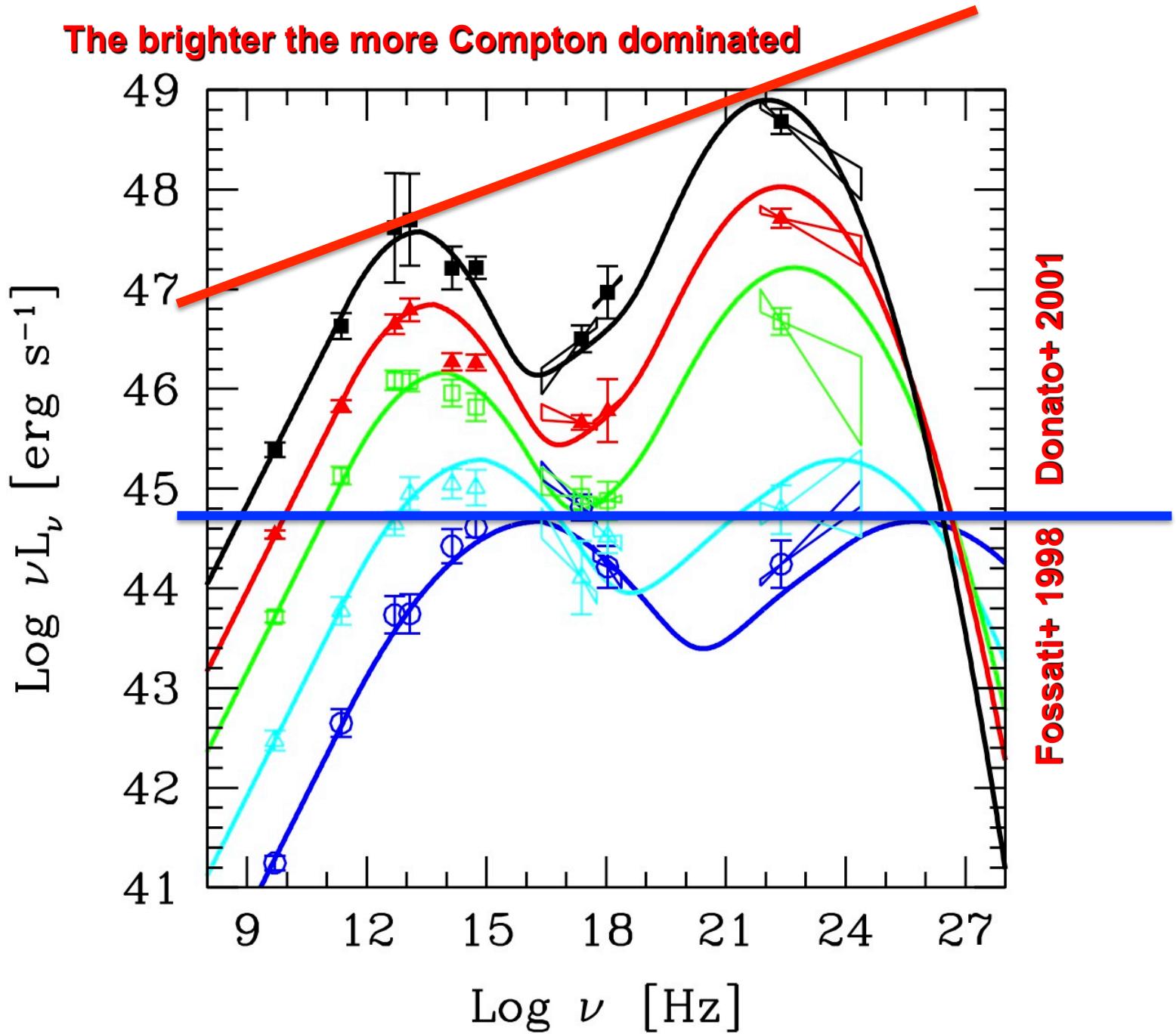


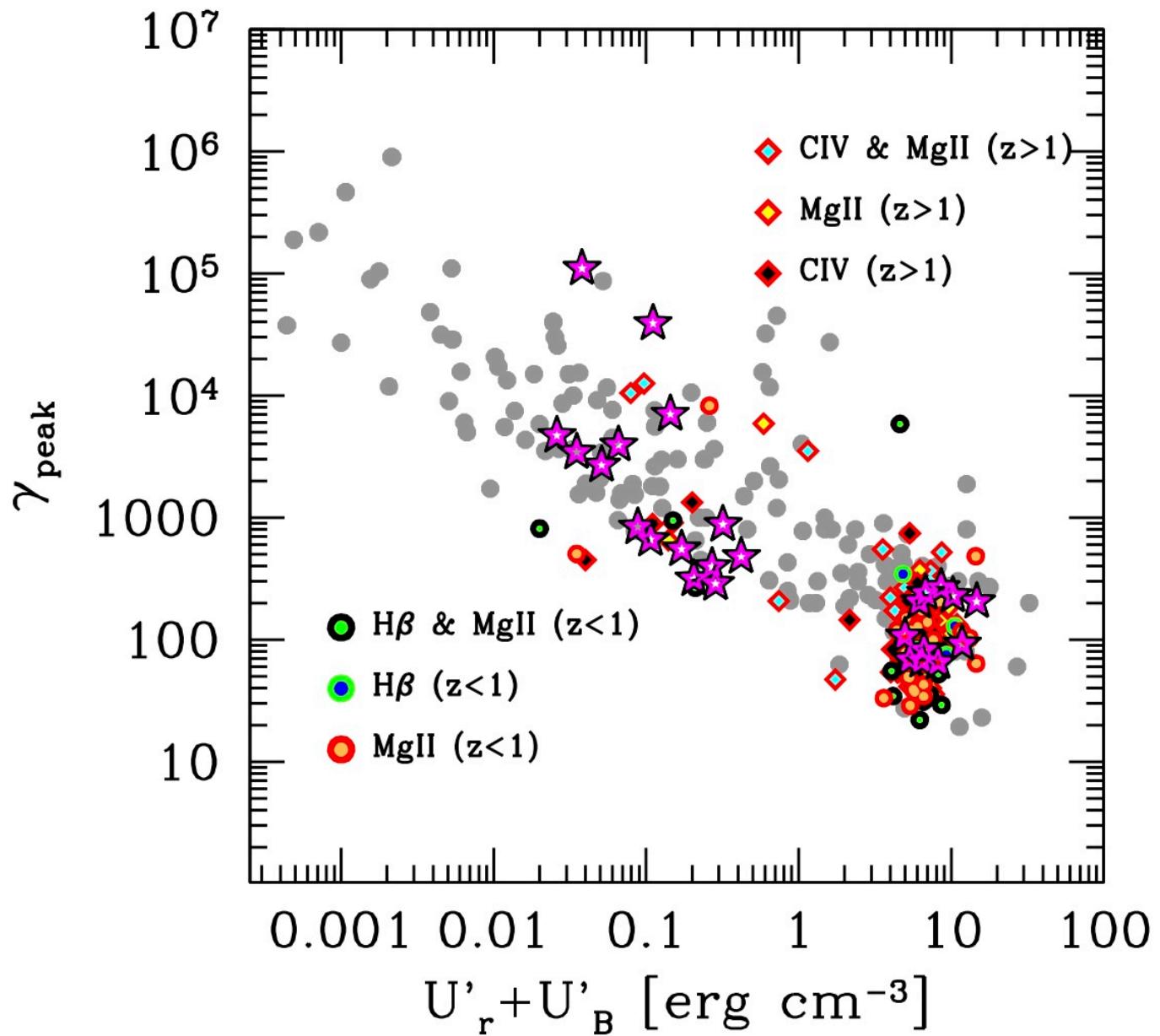
**The blazar sequence 1.0: 126 blazars, only 33 detected in  $\gamma$**

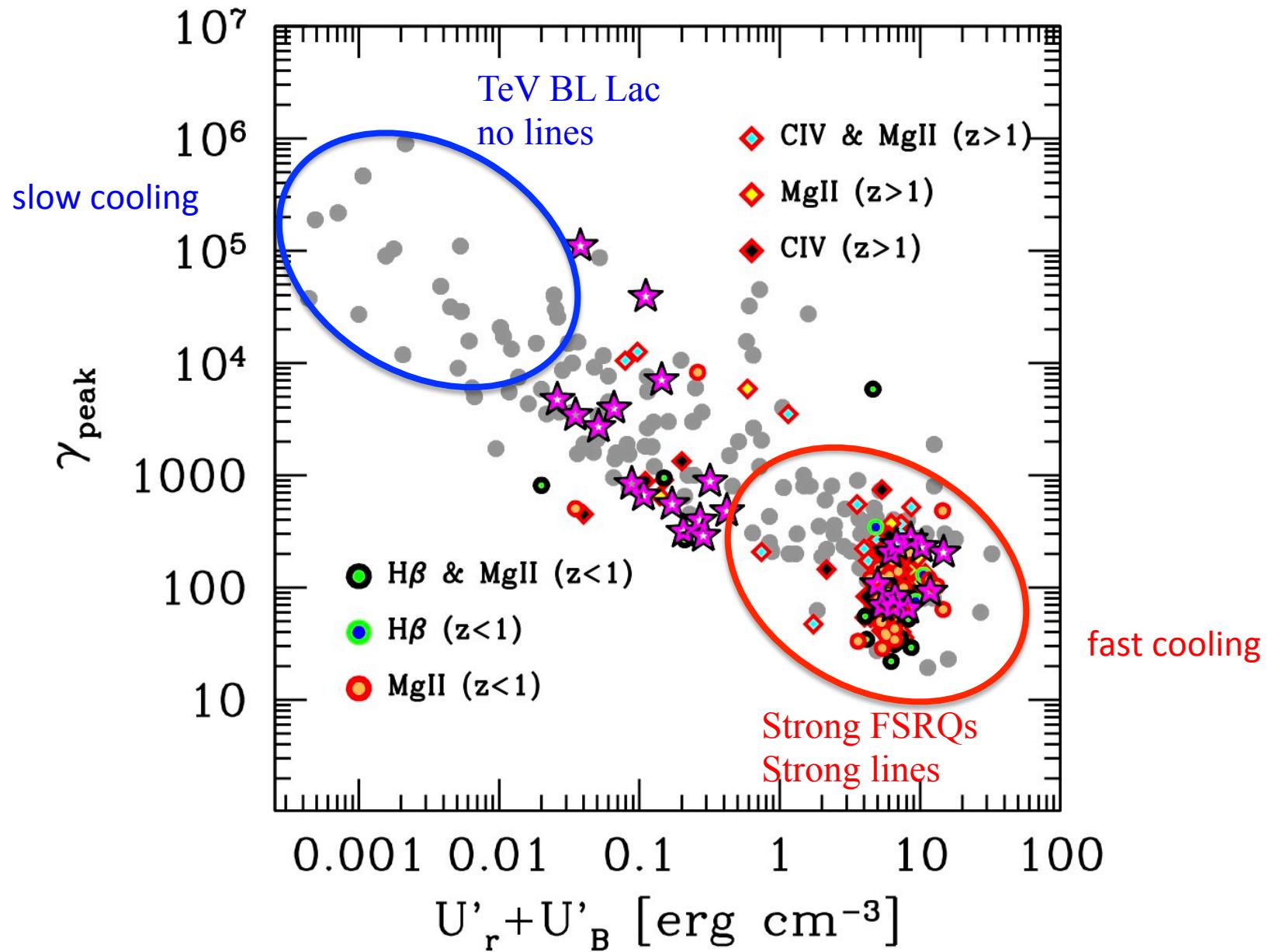


## The brighter the redder

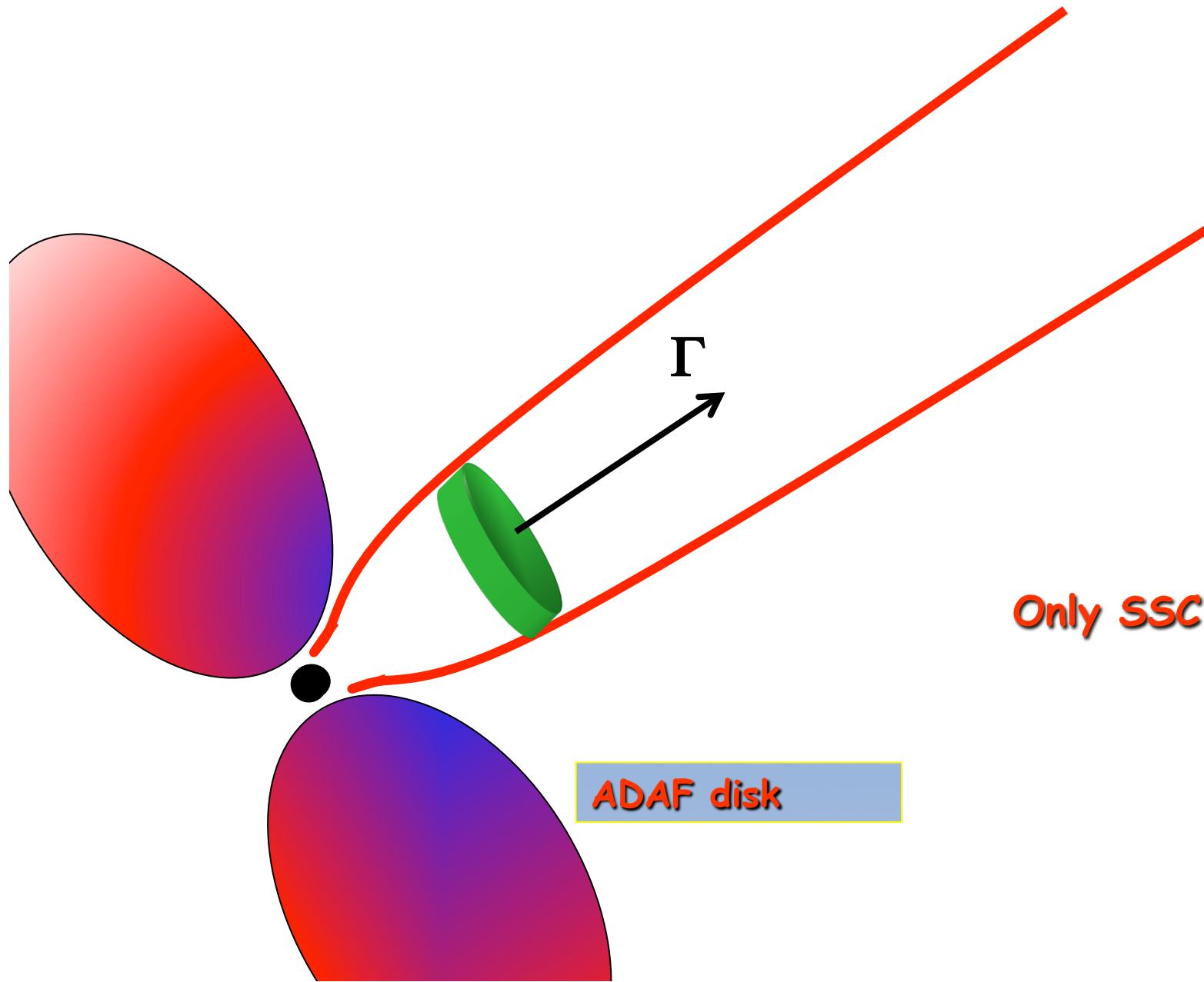




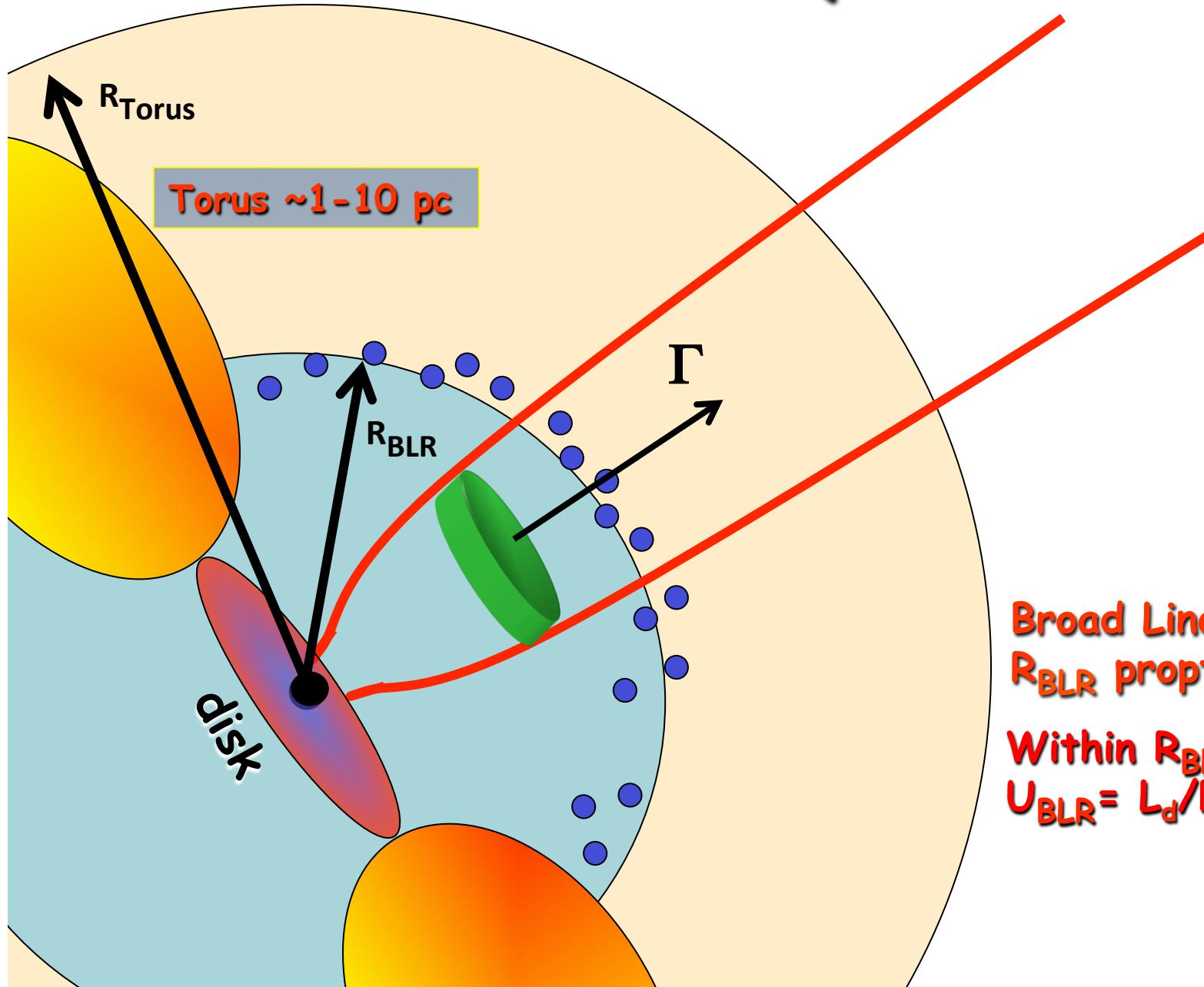




# BL Lacs



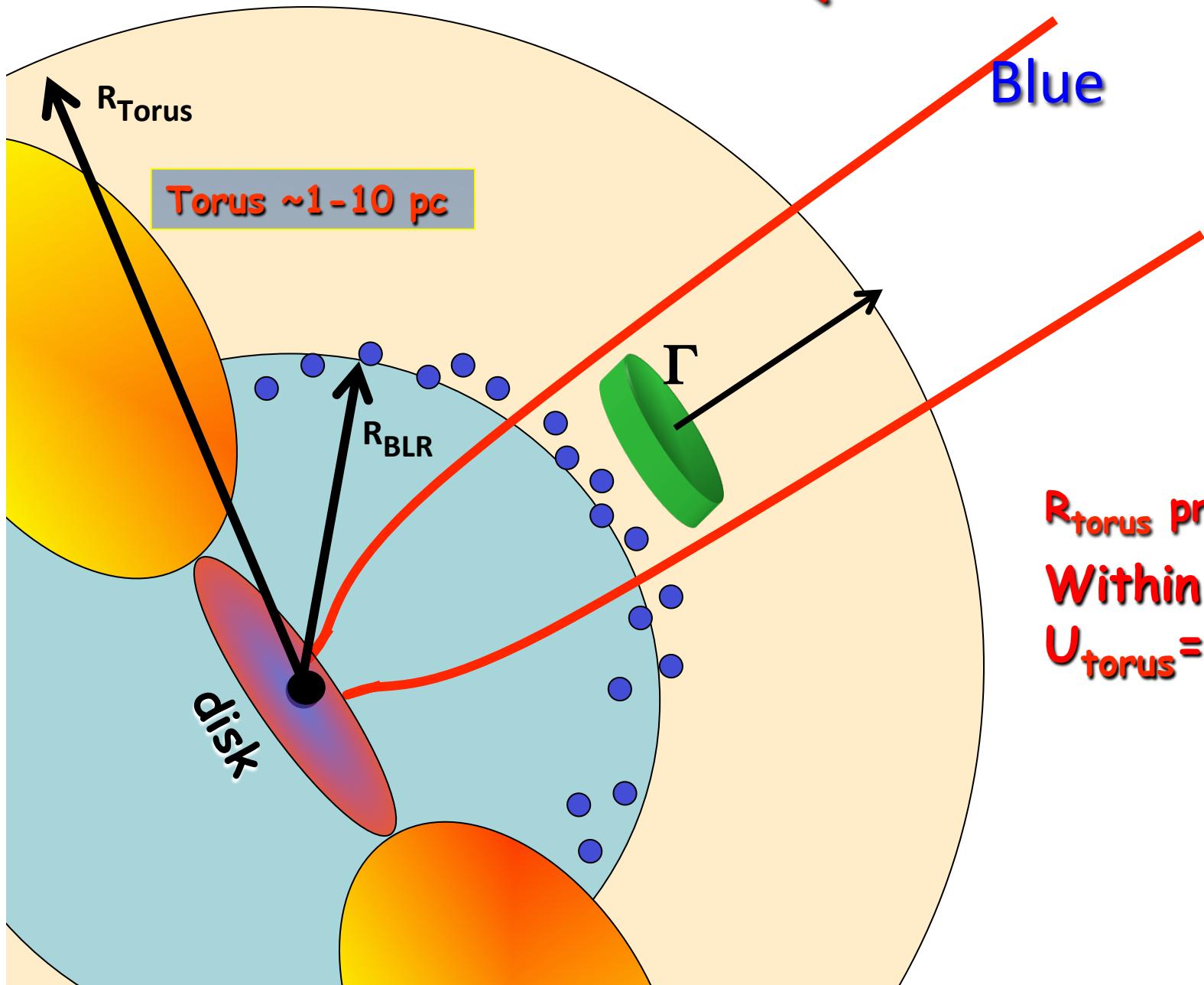
# FSRQs



Broad Line Region:  
 $R_{BLR} \propto L_d^{1/2}$

Within  $R_{BLR}$   
 $U_{BLR} = L_d / R^2 c = \text{const}$

# FSRQs



$R_{\text{Torus}} \propto L_d^{1/2}$

Within  $R_{\text{Torus}}$   
 $U_{\text{Torus}} = \text{const}$

## **Criticism:**

**Giommi, Menna, Padovani 1999**

**Perlman + 2001**

**Padovani+ 2003**

**Caccianiga & Marcha 2004**

**Anton & Browne 2005**

**Giommi+ 2005**

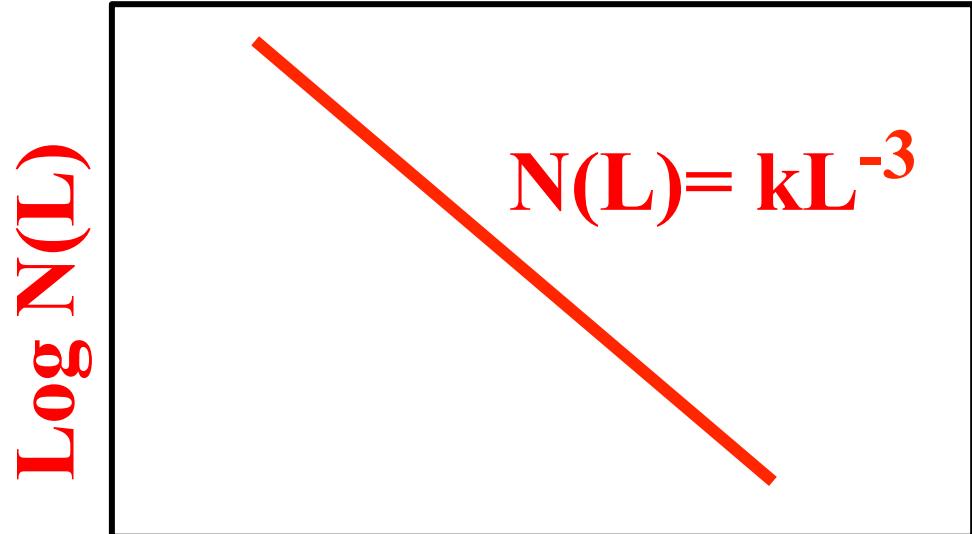
**Nieppola, Tornikoski, Valtoja 2006**

**Padovani, Giommi, Rau 2012**

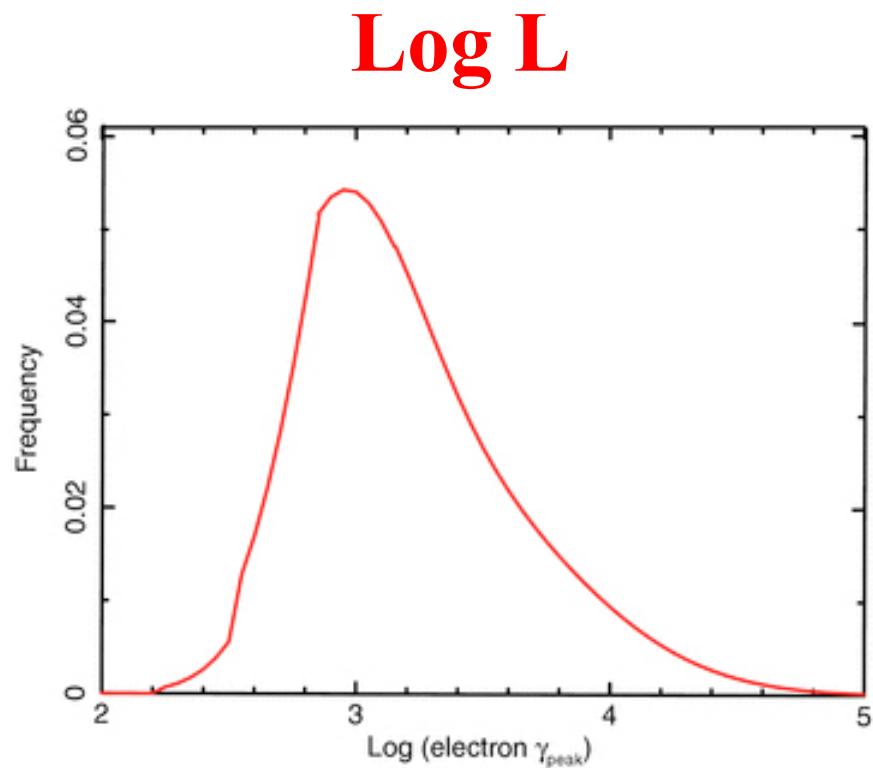
**For review Padovani 2007 and Ghisellini & Tavecchio 2008**

**Recently: “Simplified blazar scenario”**

**Giommi, Padovani et al. 2012: sequence is apparent, being the result of selection effects**

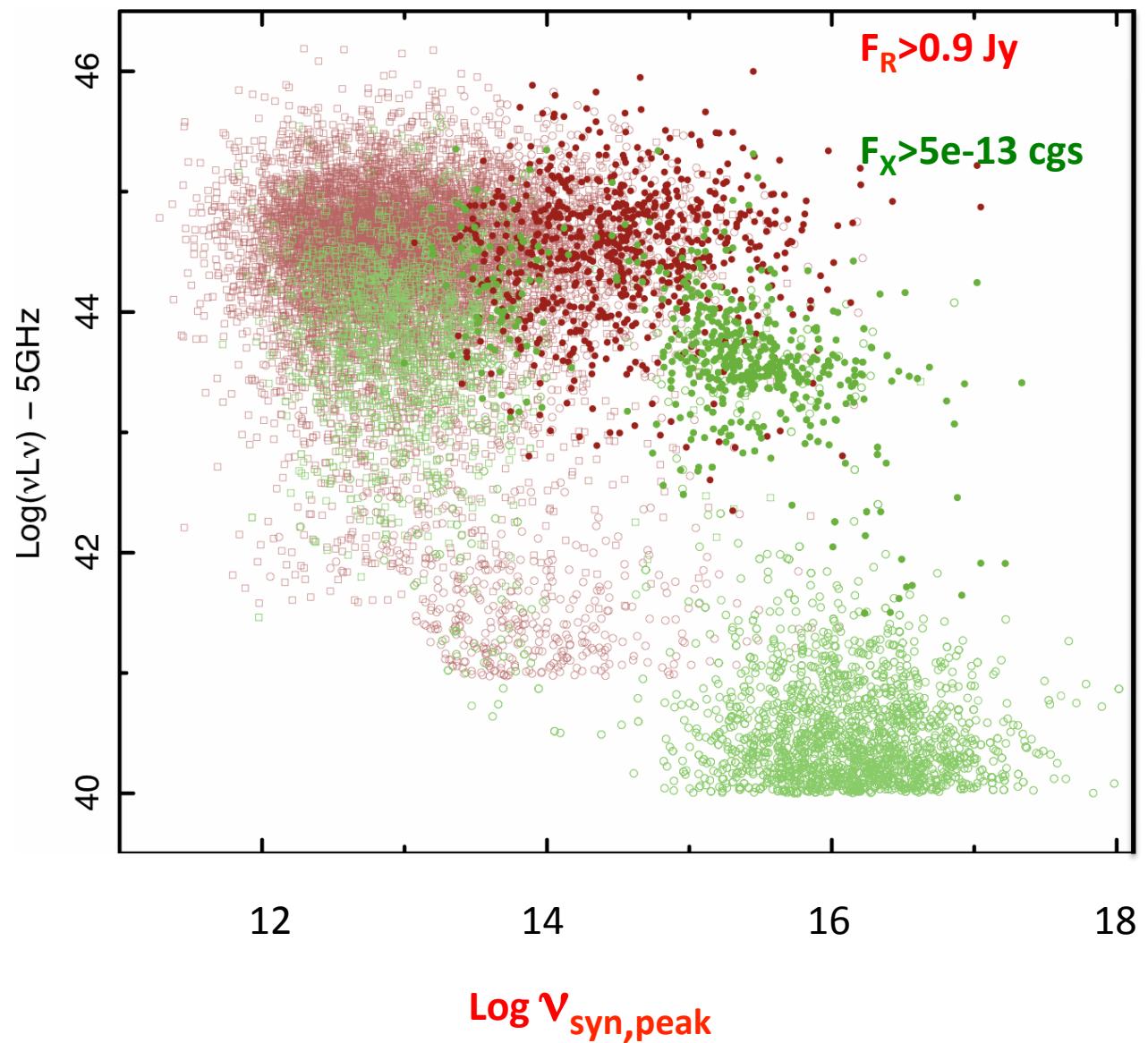


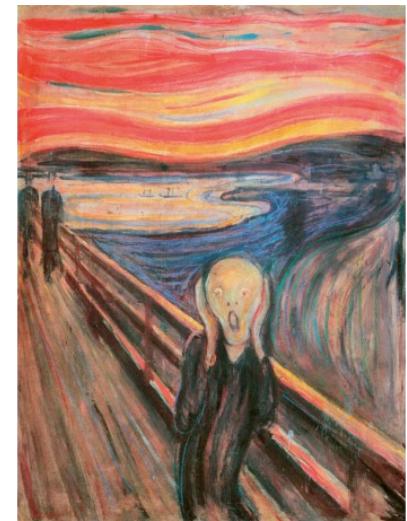
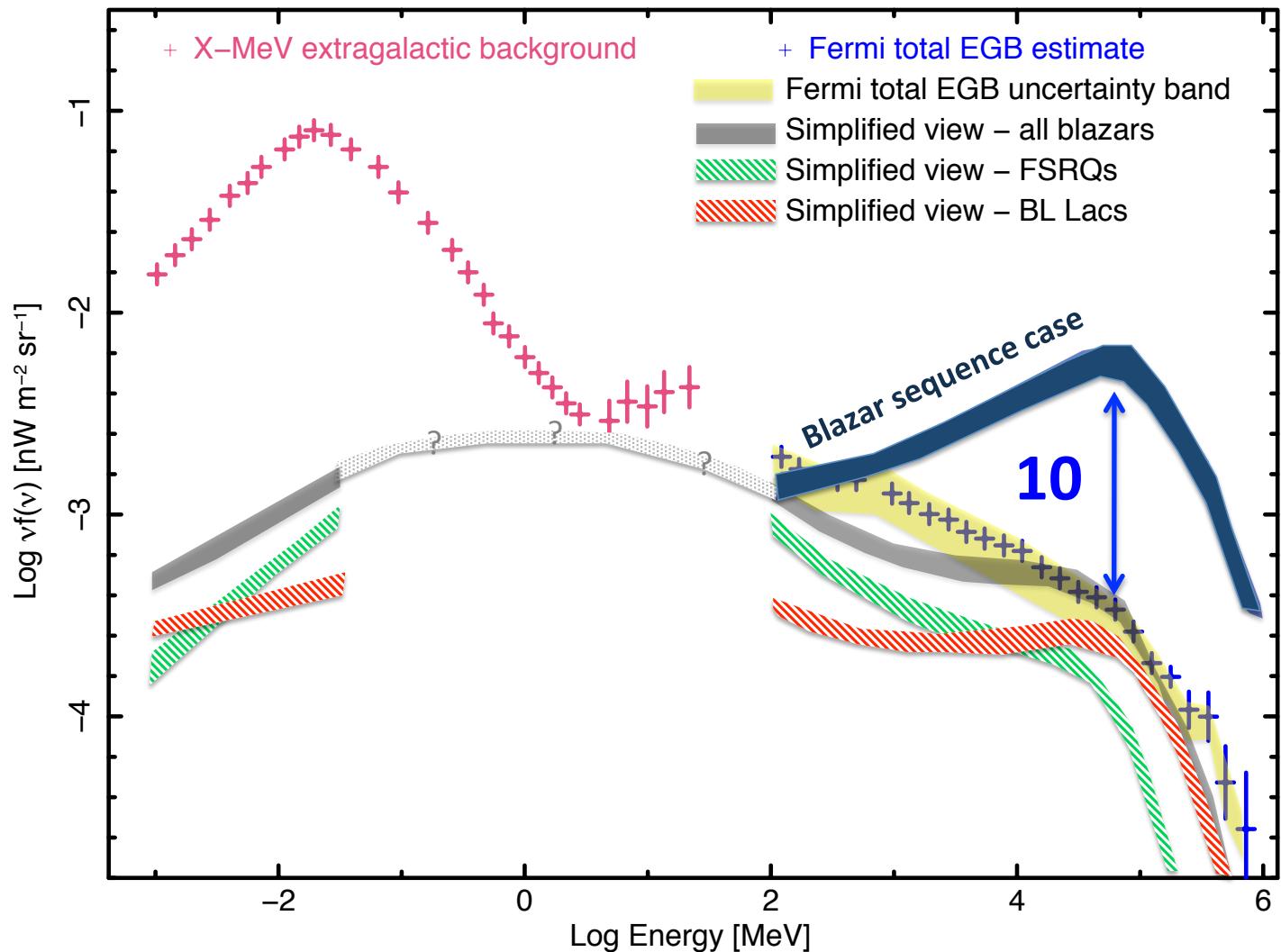
Radio Luminosity  
function @ 41 GHz



Distribution of  $\gamma_{\text{peak}}$

Giommi, Padovani+ 2012

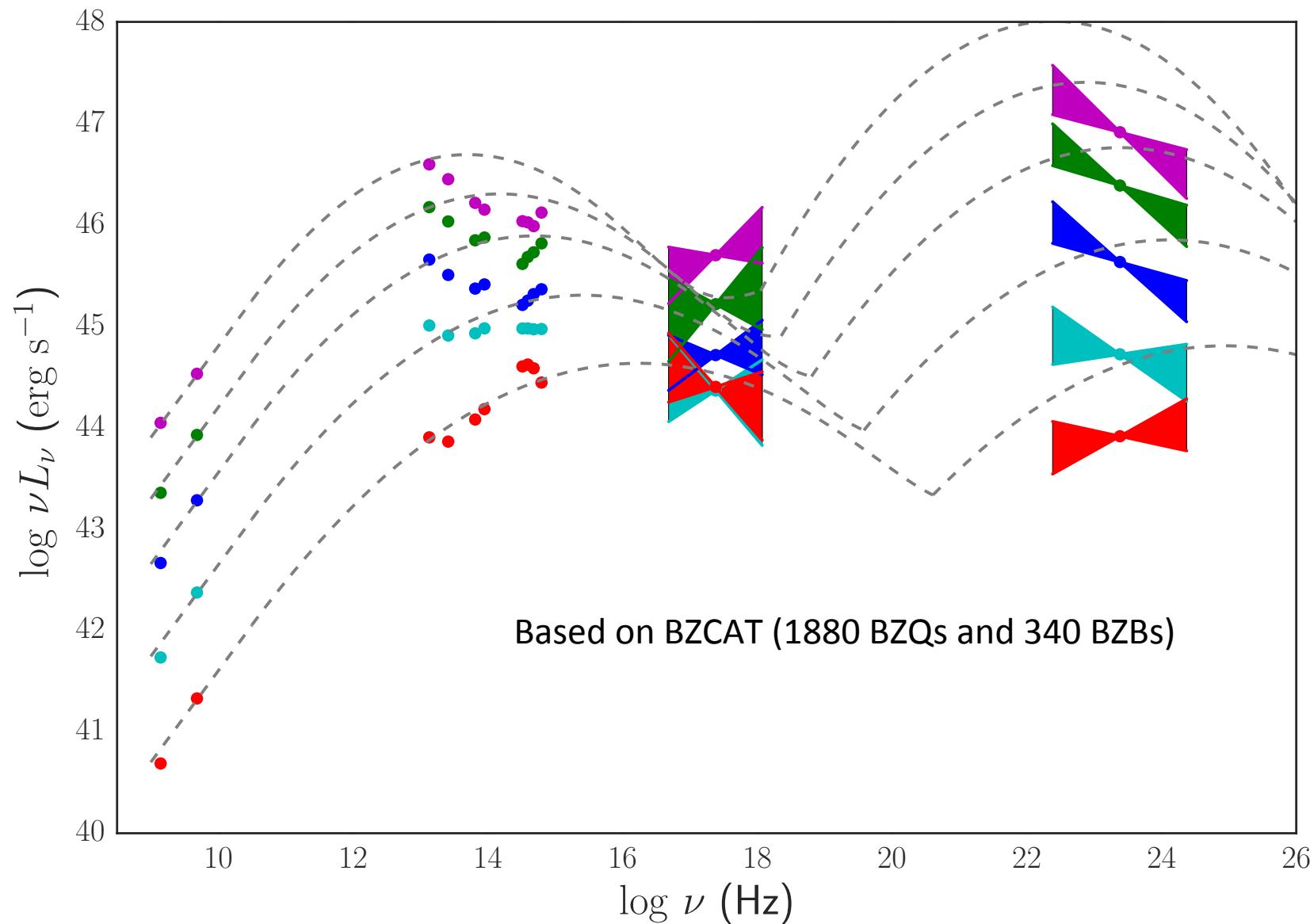




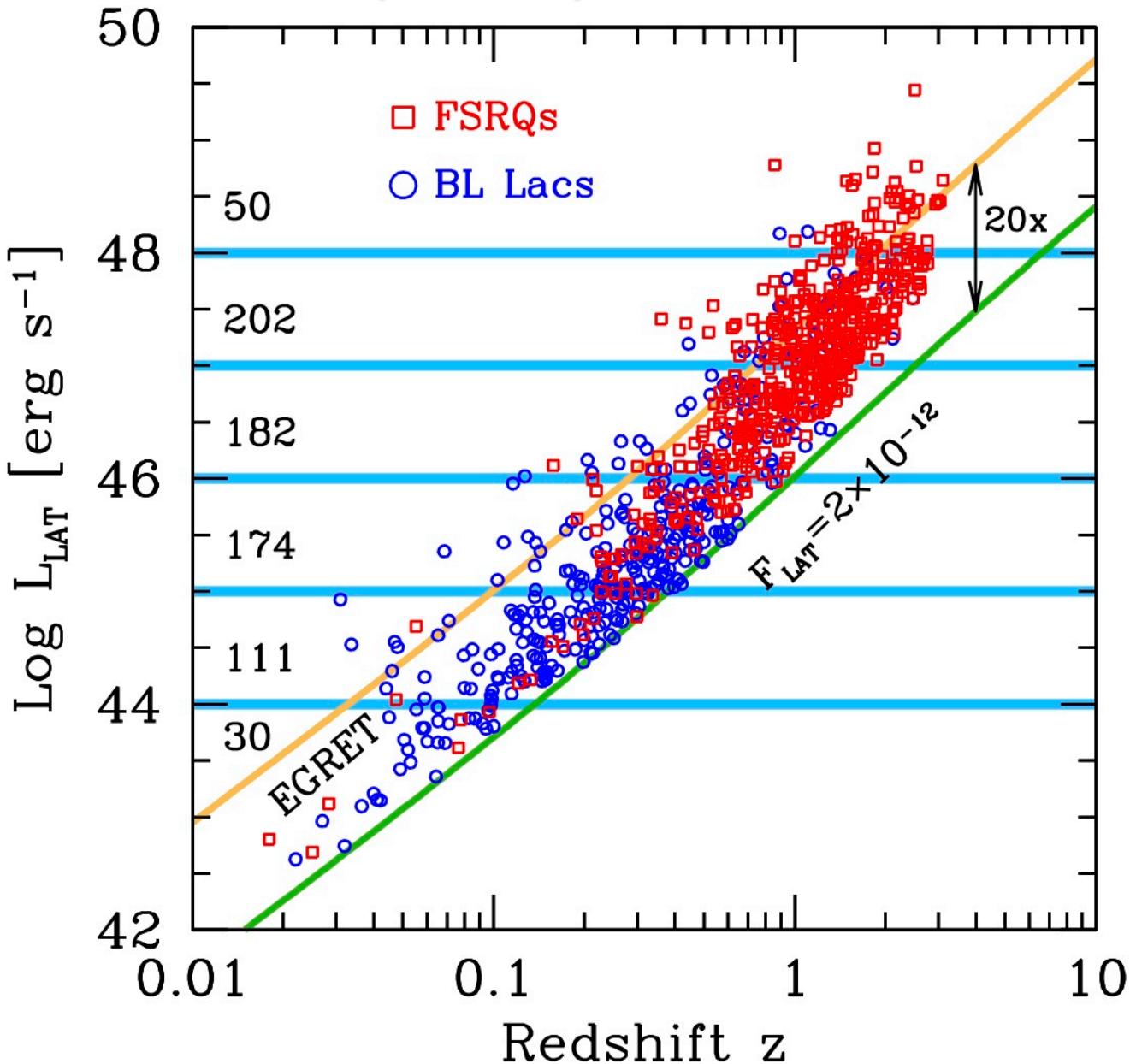
Munch

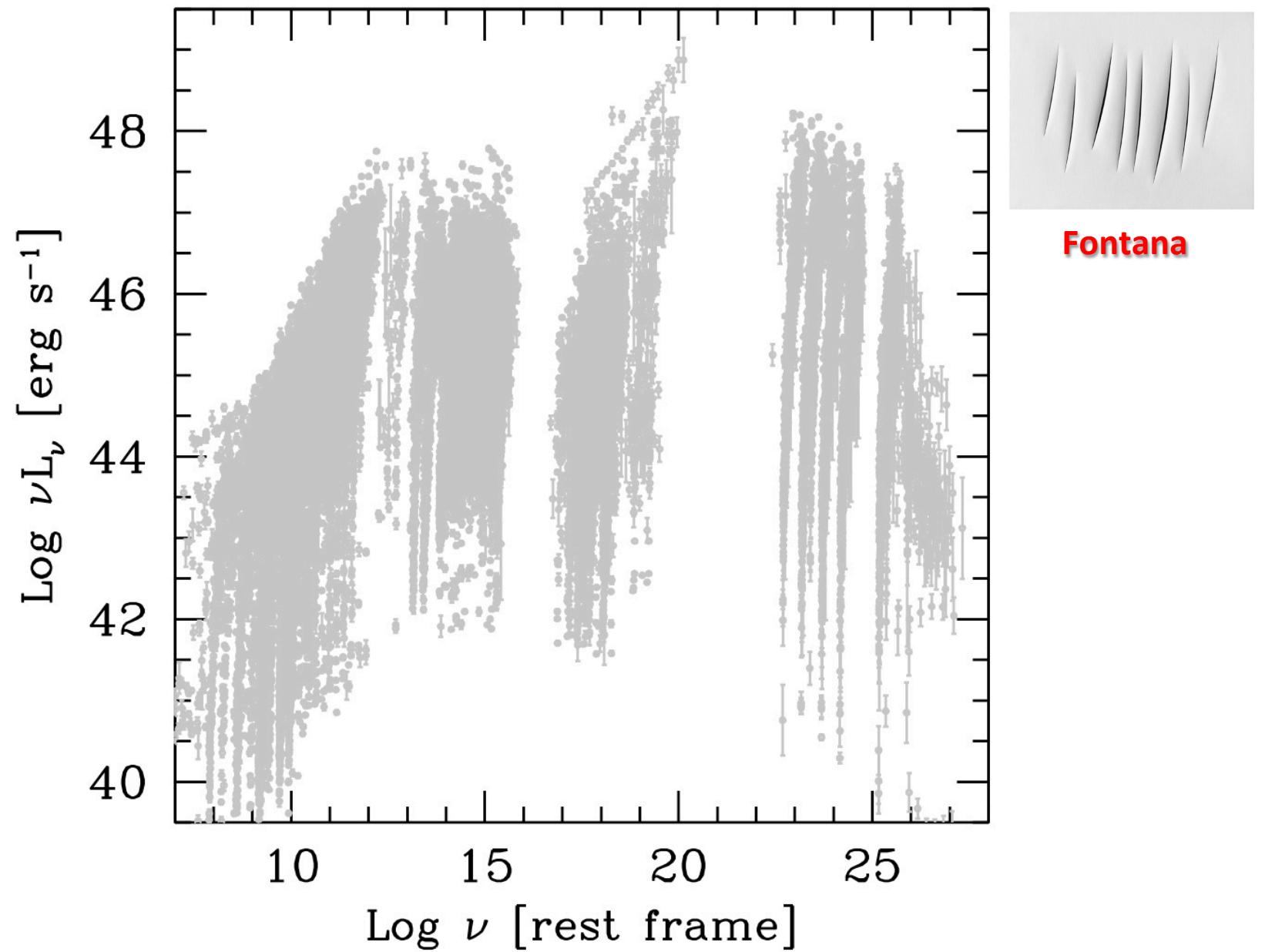
Giommi & Padovani 2015

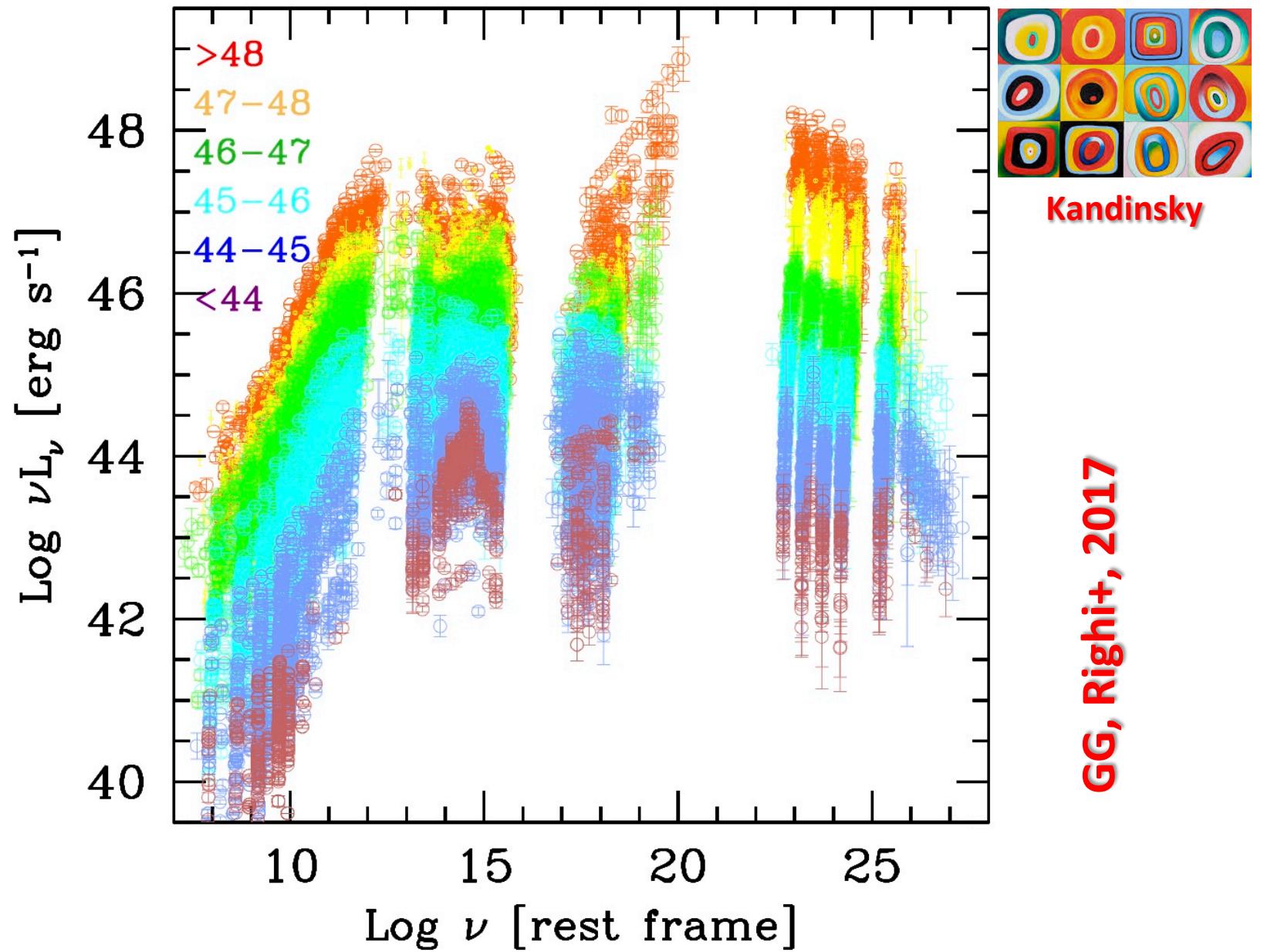
Mao, Urry, Massaro et al. 2016, ApJS 224, 26



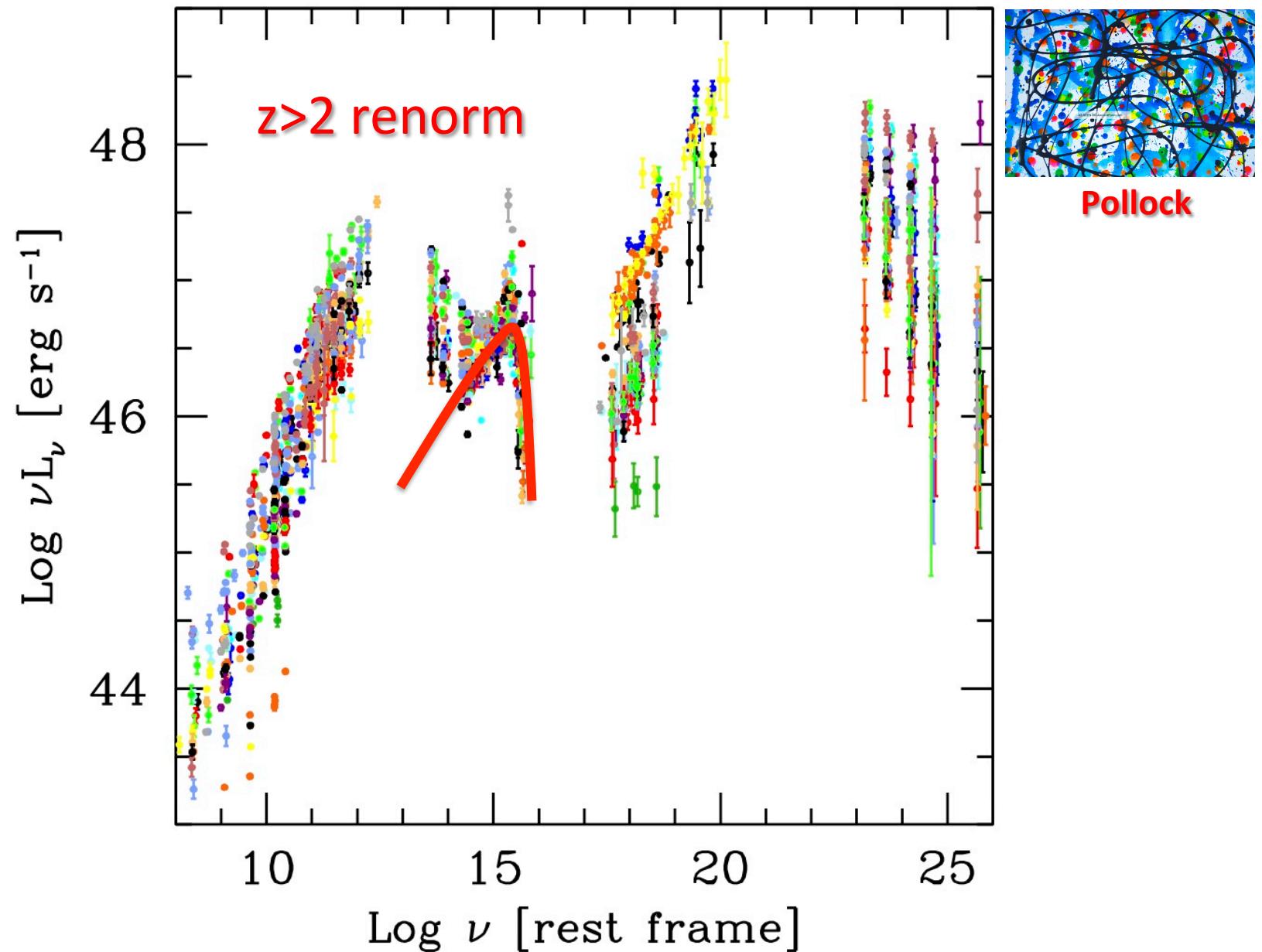
### 3LAC, "clean", with z: 749 blazars

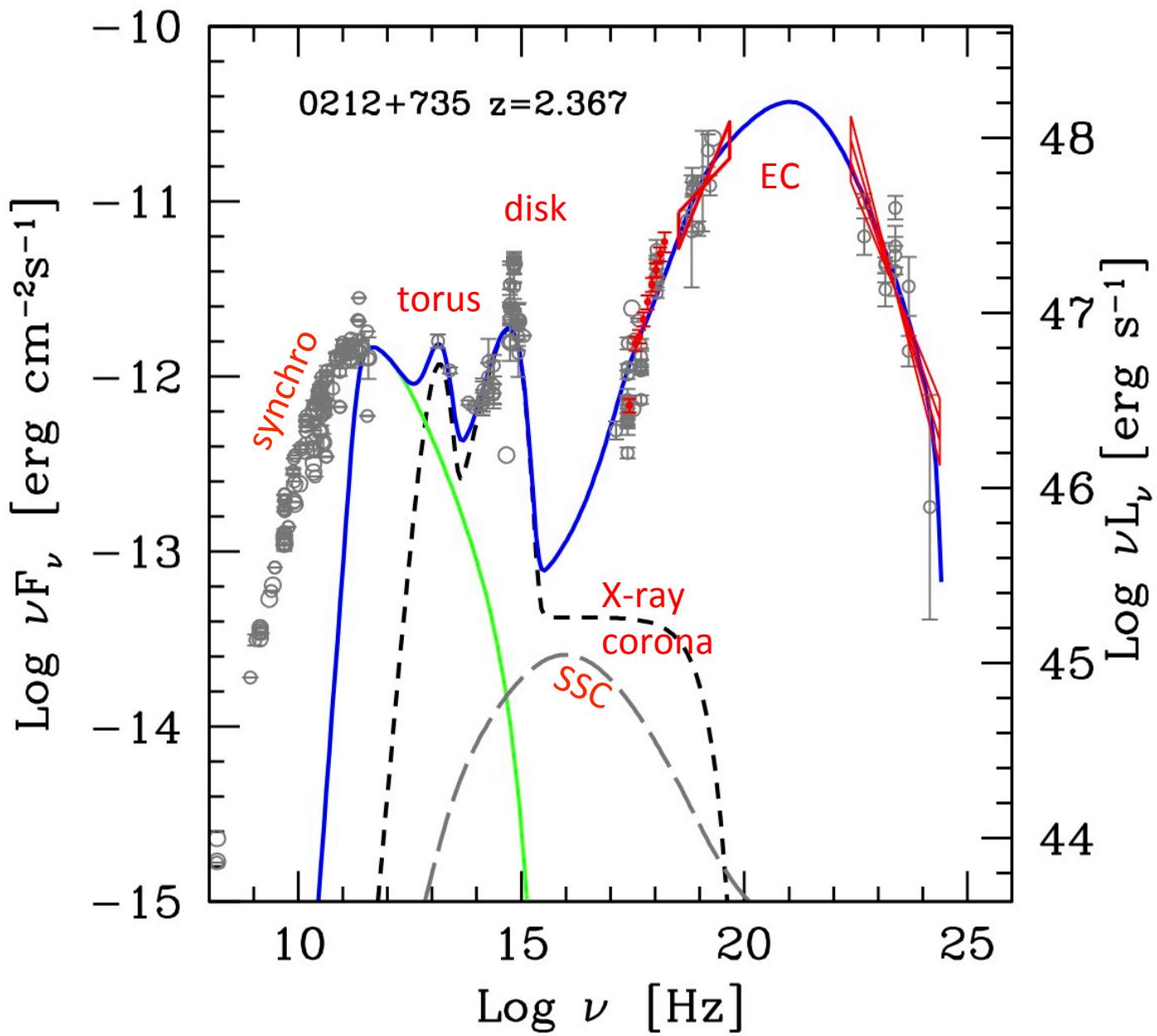


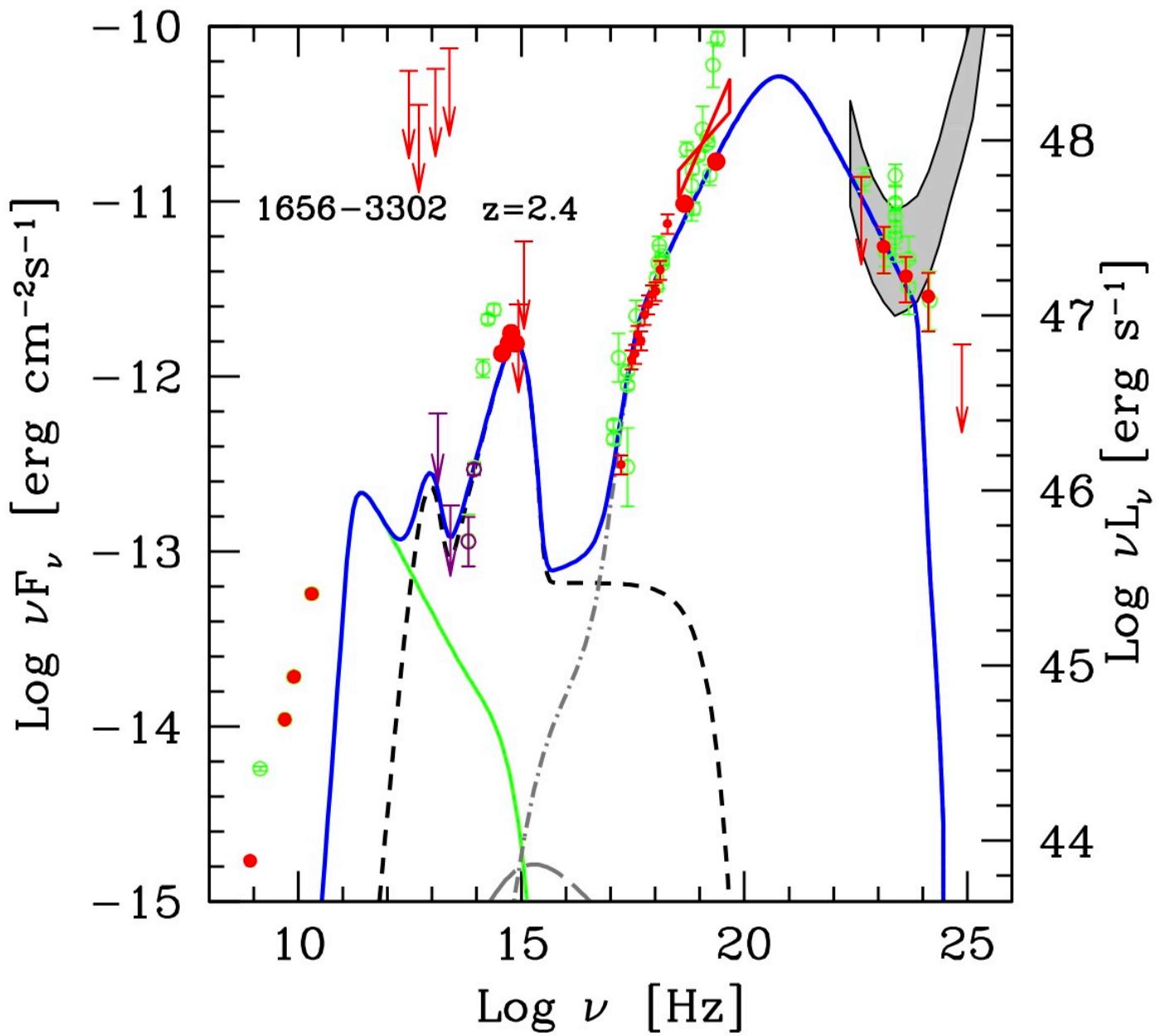




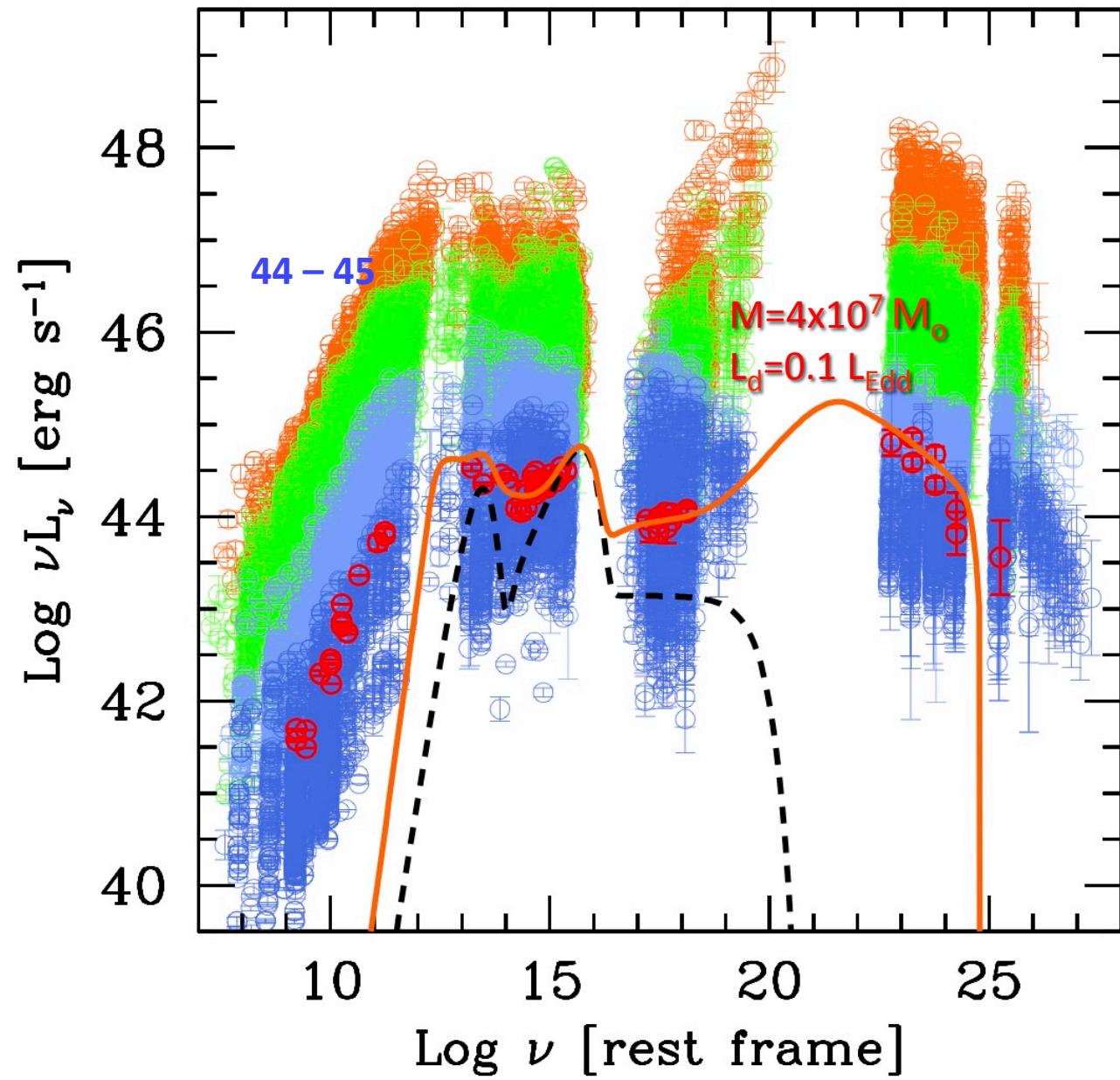
**In powerful blazars at  $z > 2$  the accretion disk is well visible. Also the torus is visible.**

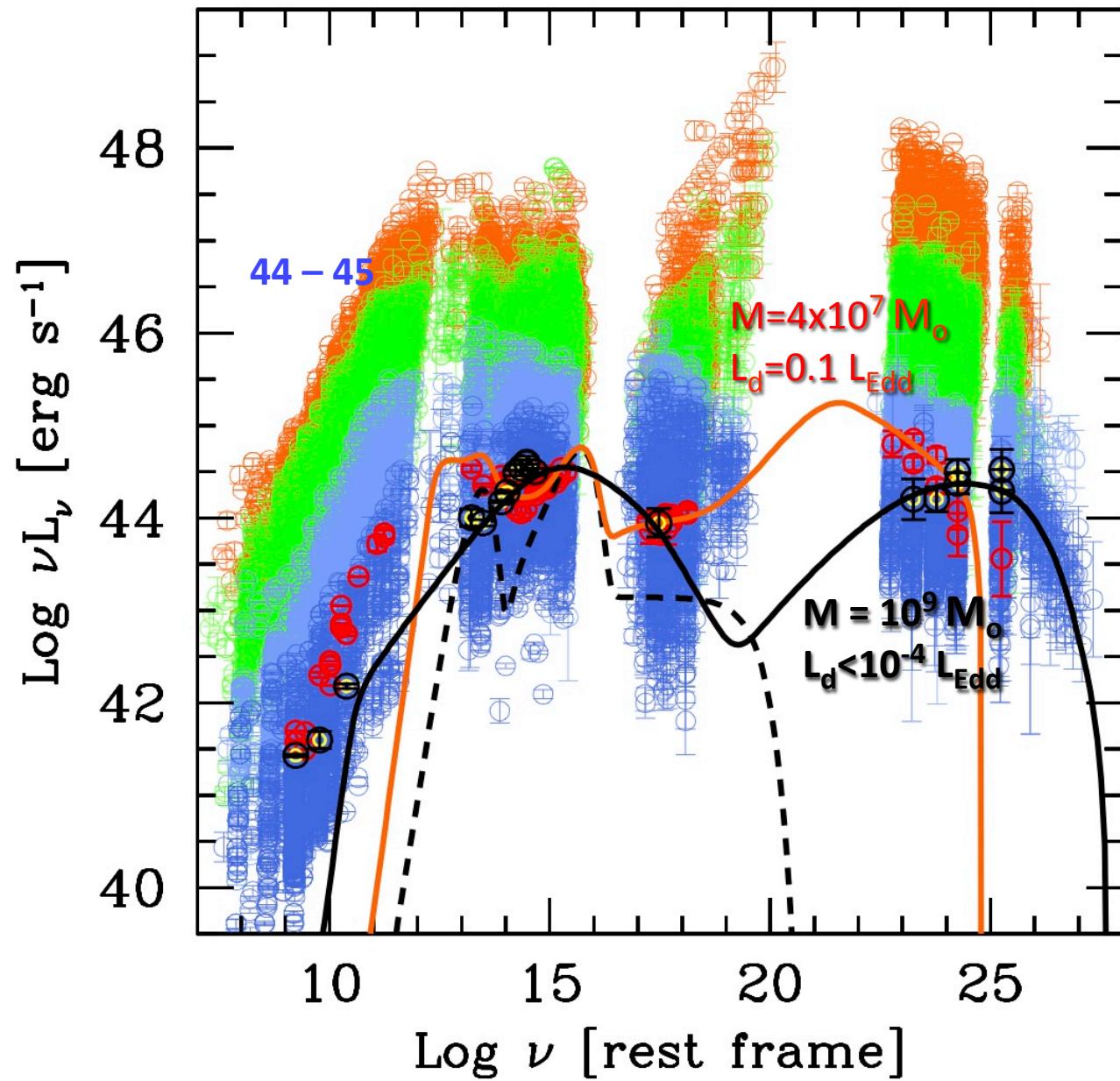




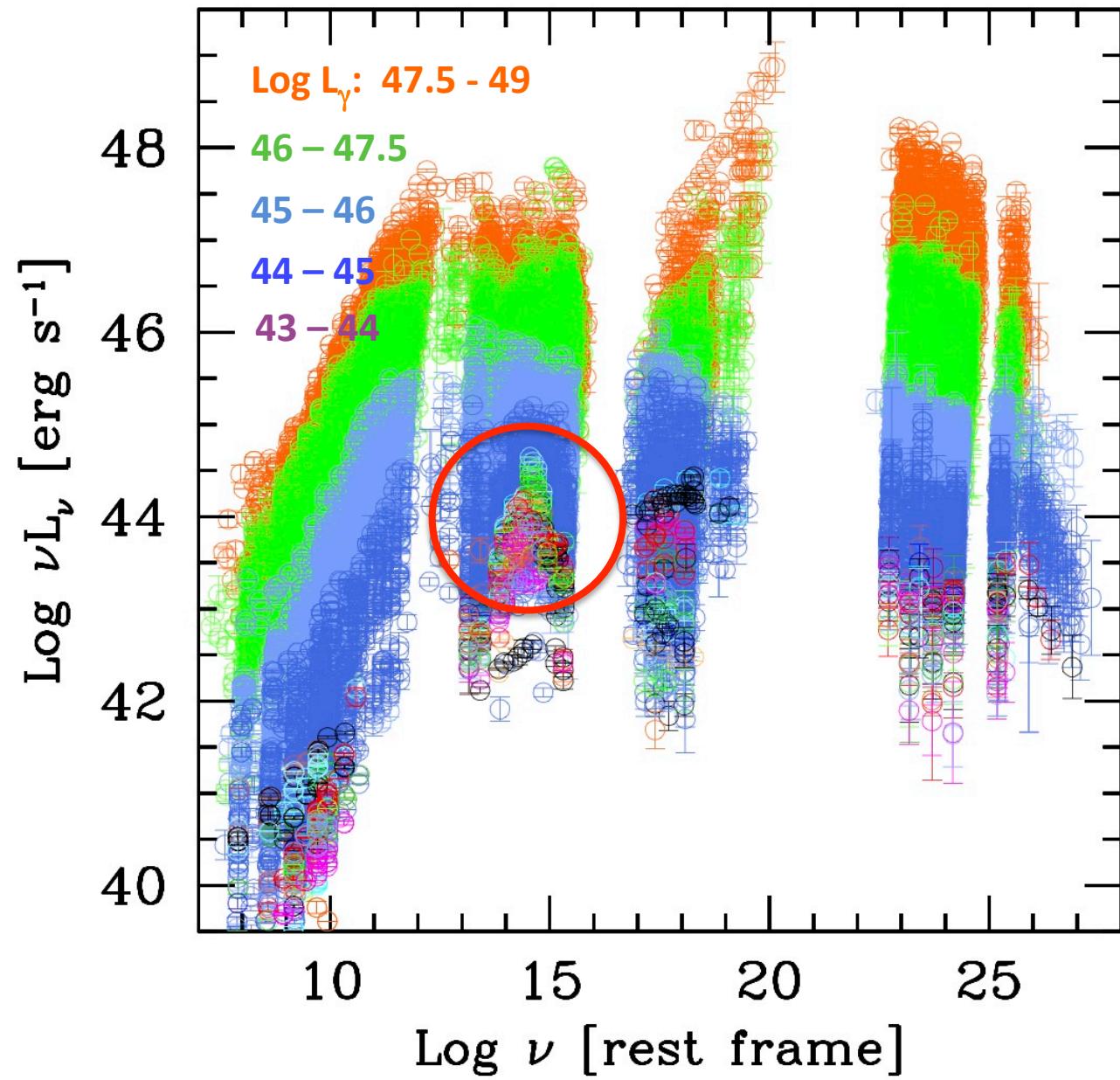


# Intermediate luminosities: a mess

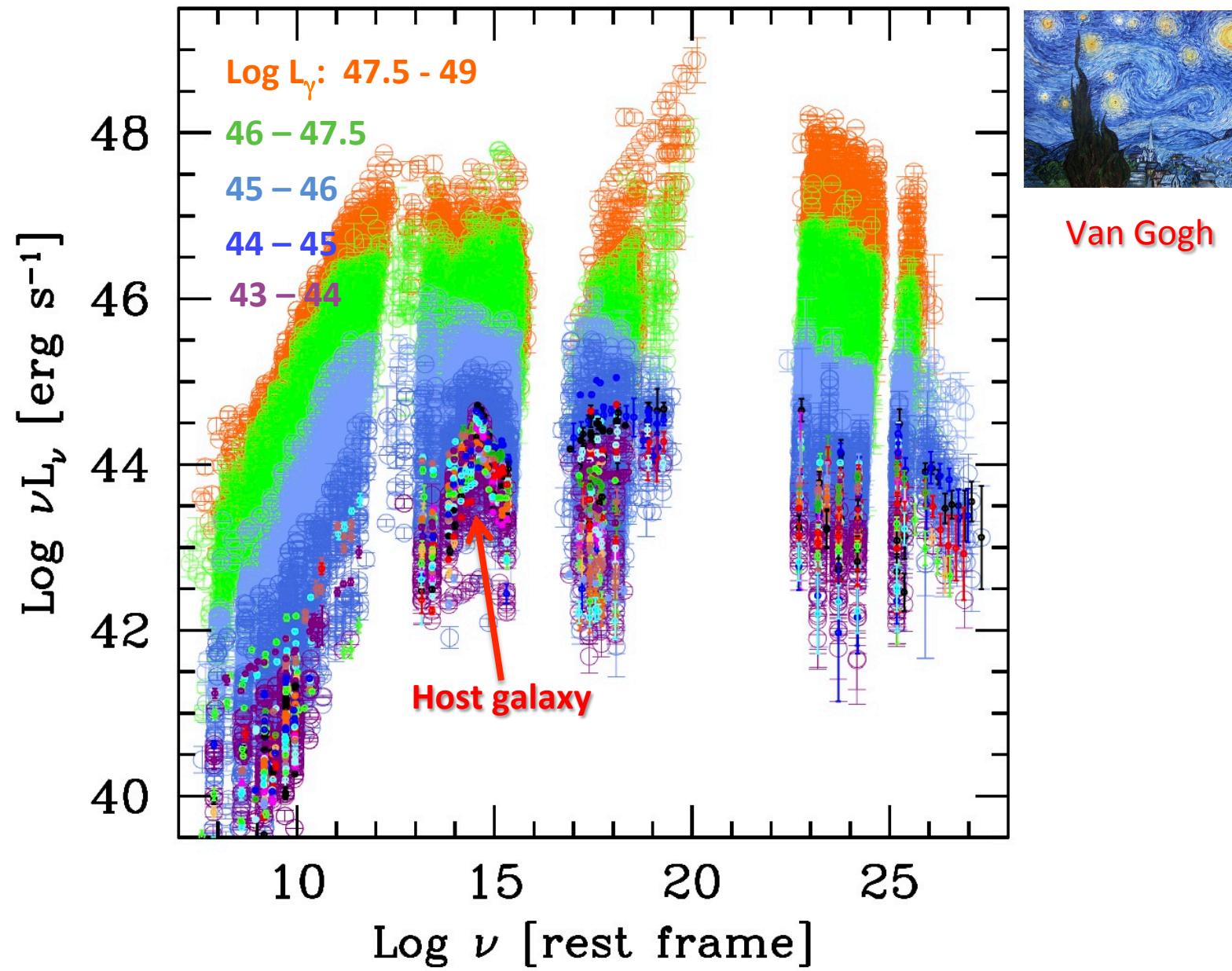


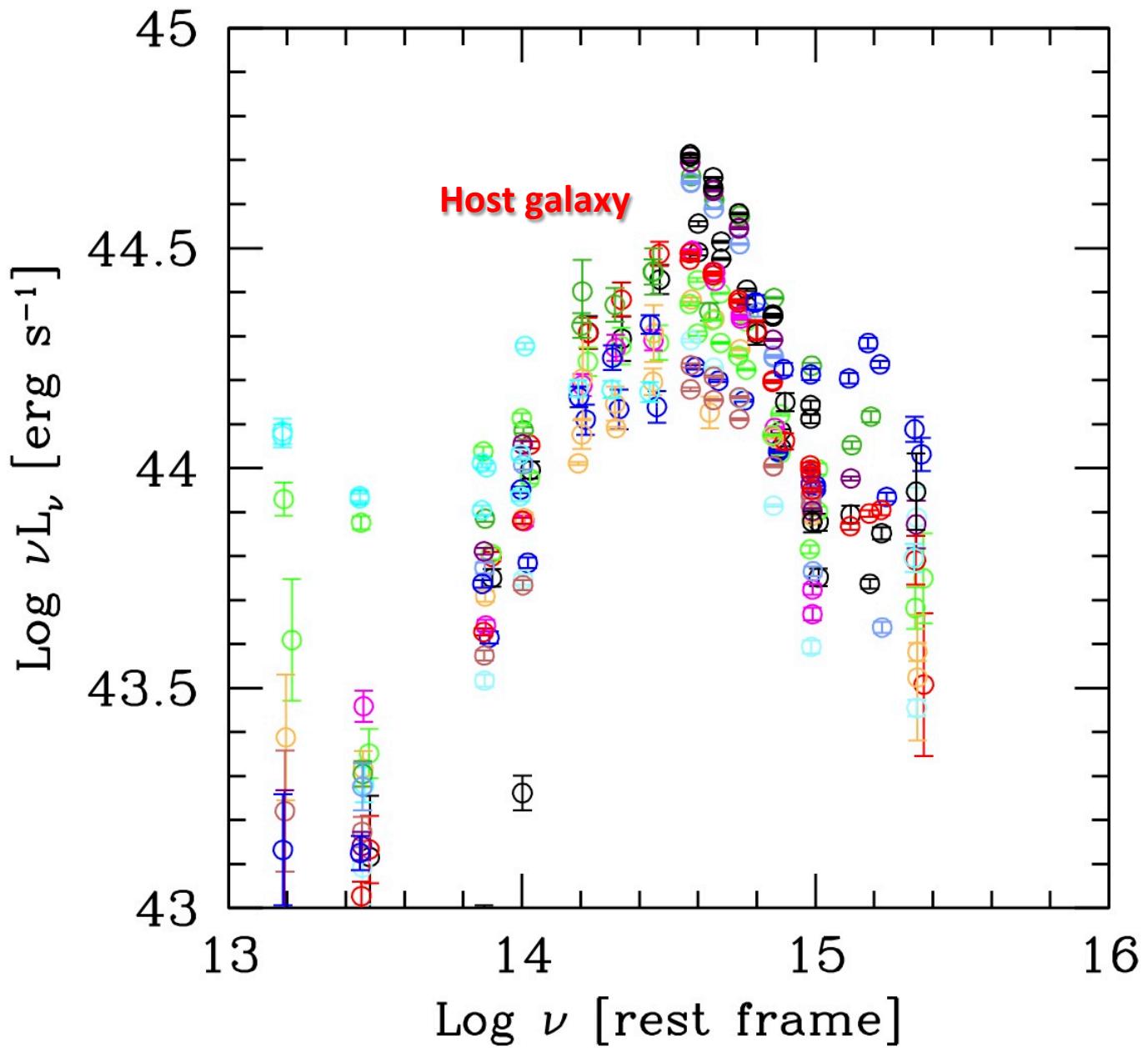


**Low luminosities:  
the host galaxy**

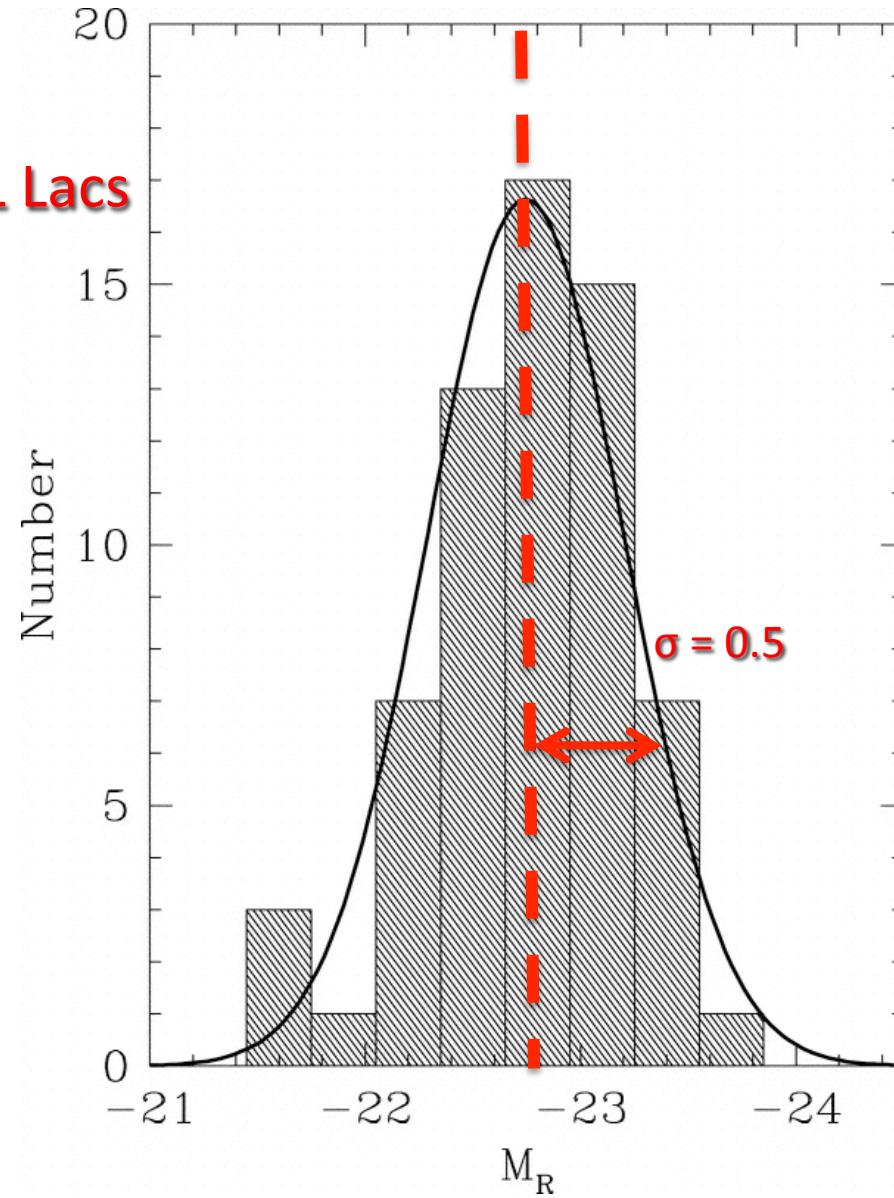


Van Gogh



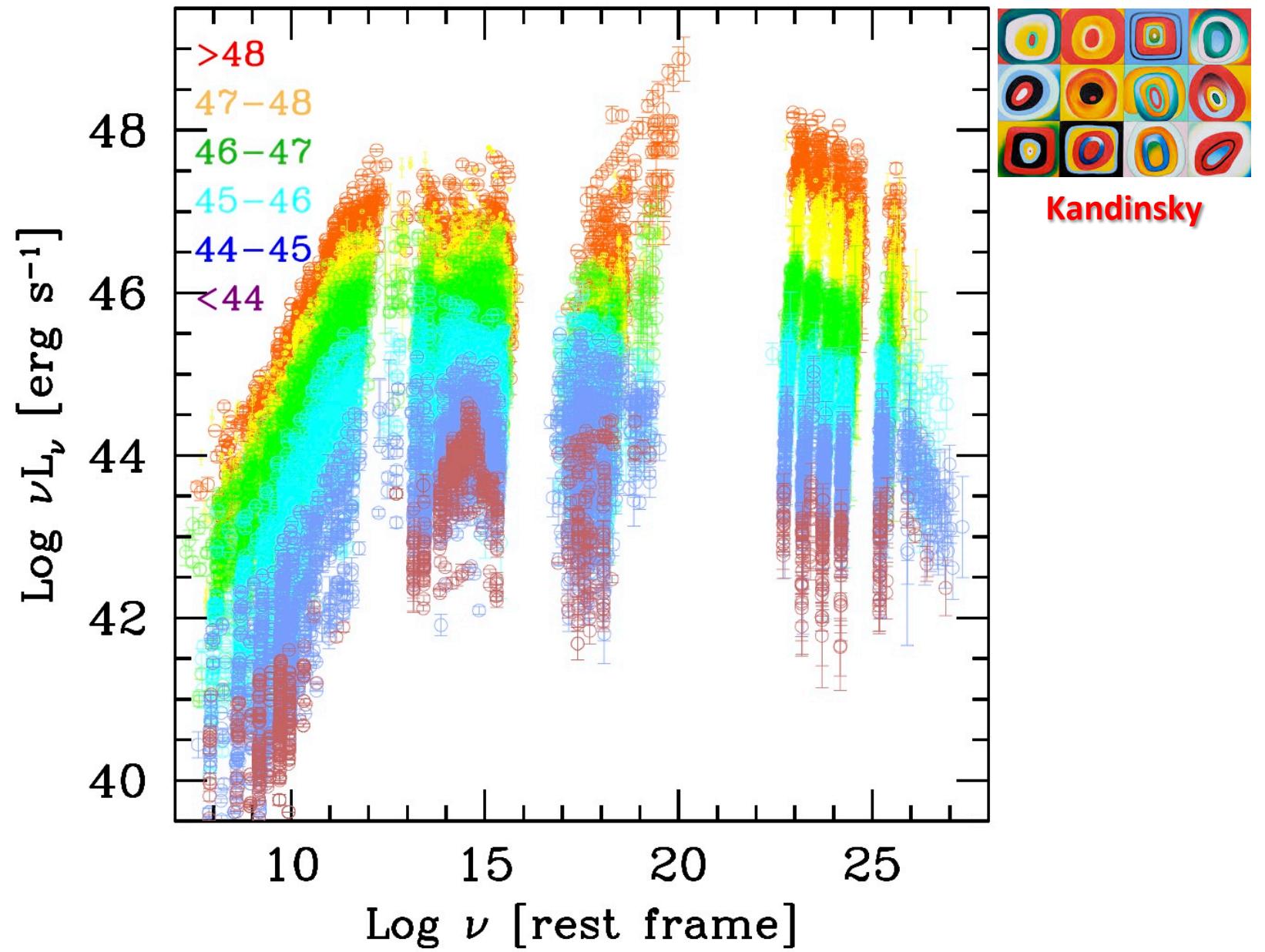


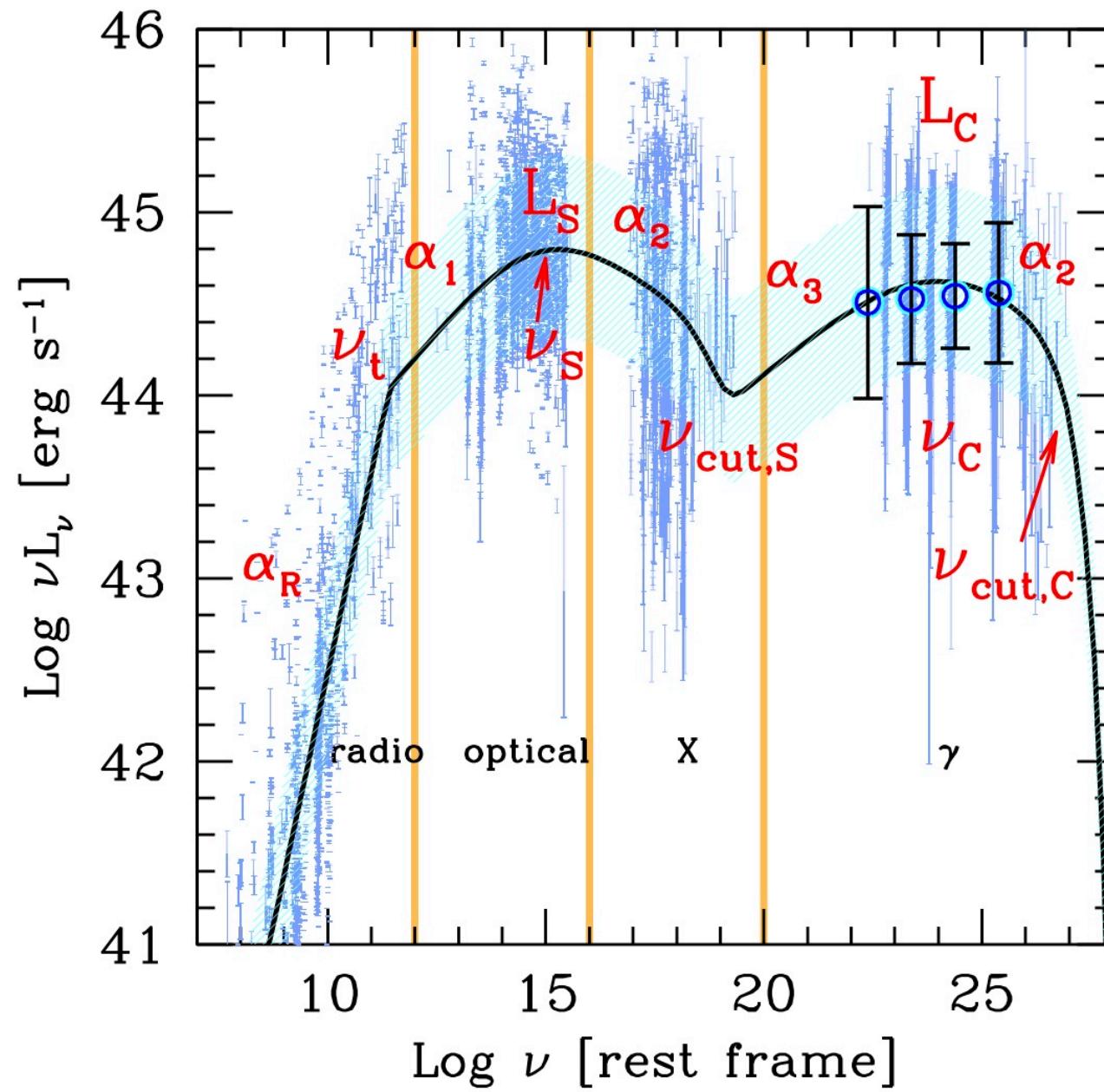
## Host galaxies of BL Lacs

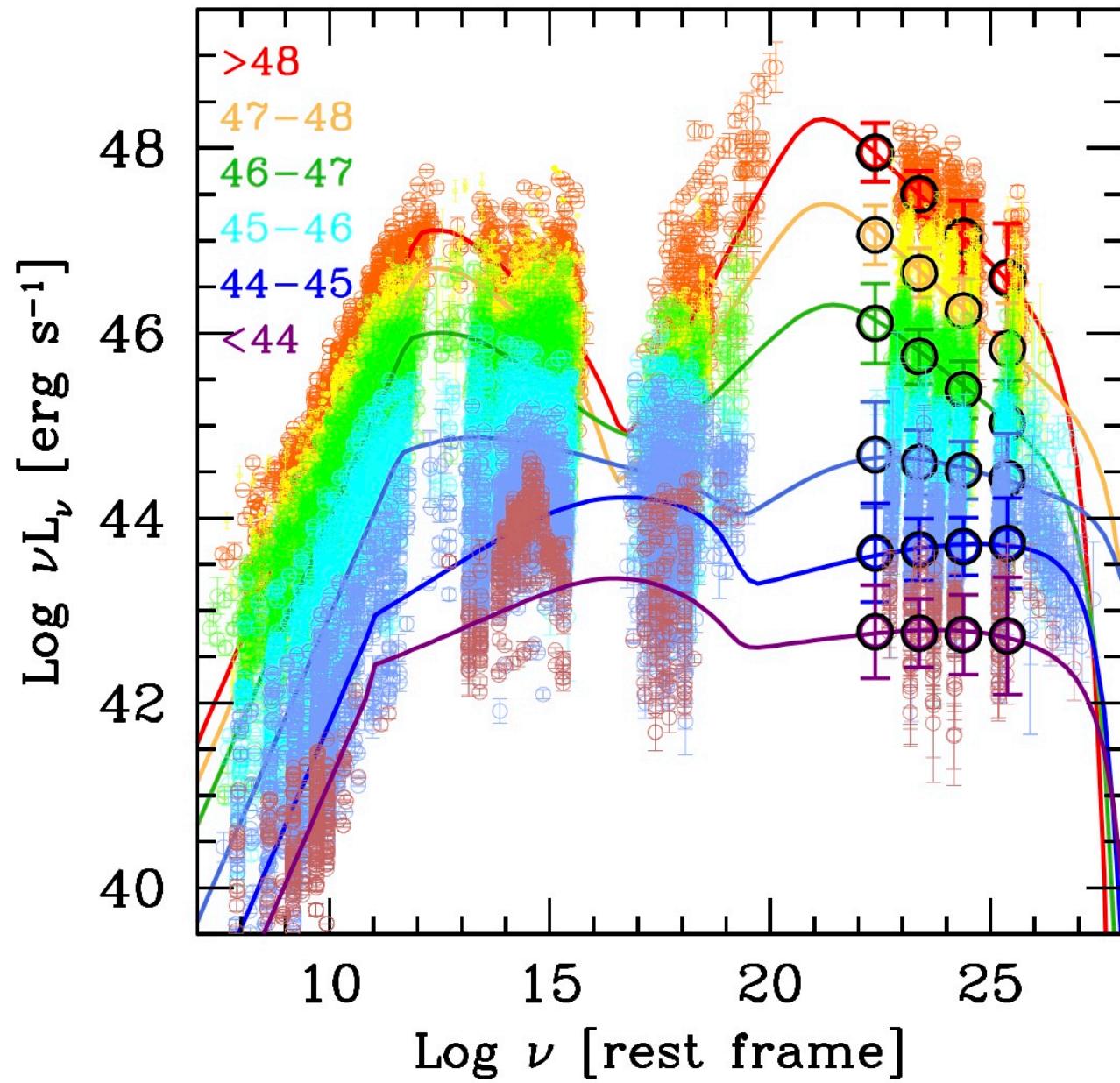


Sbarufatti, Treves & Falomo, 2005

altogether

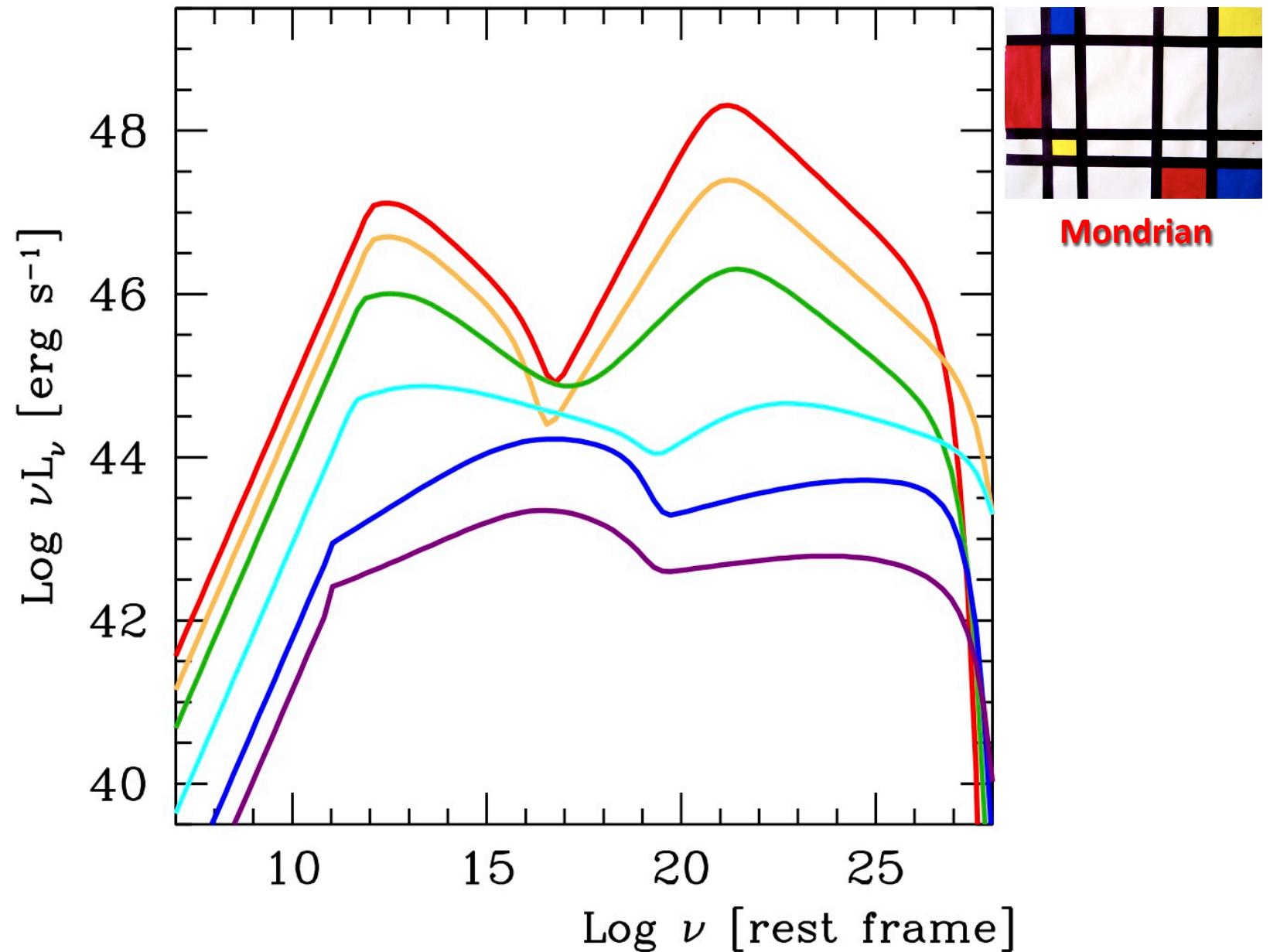


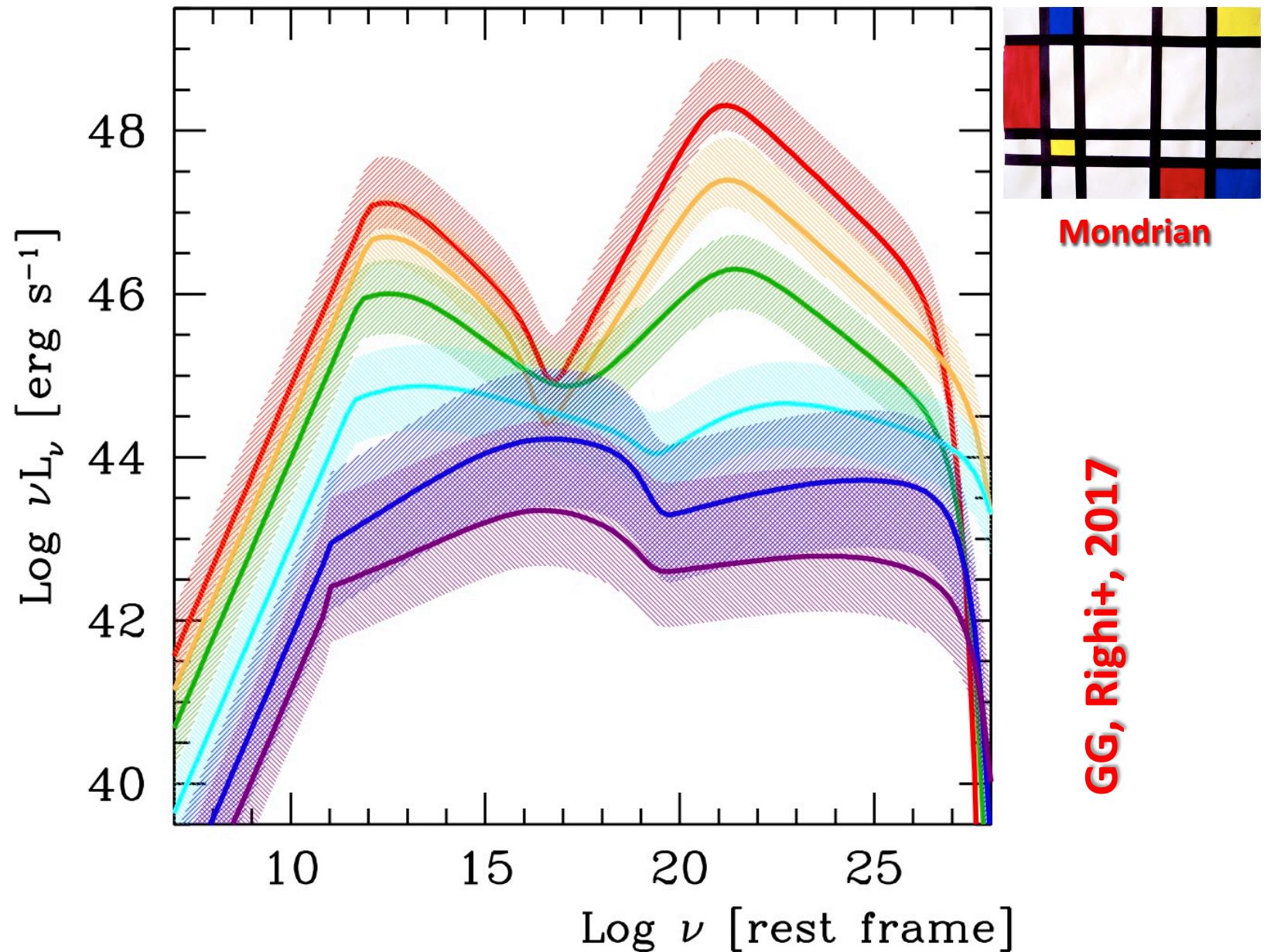


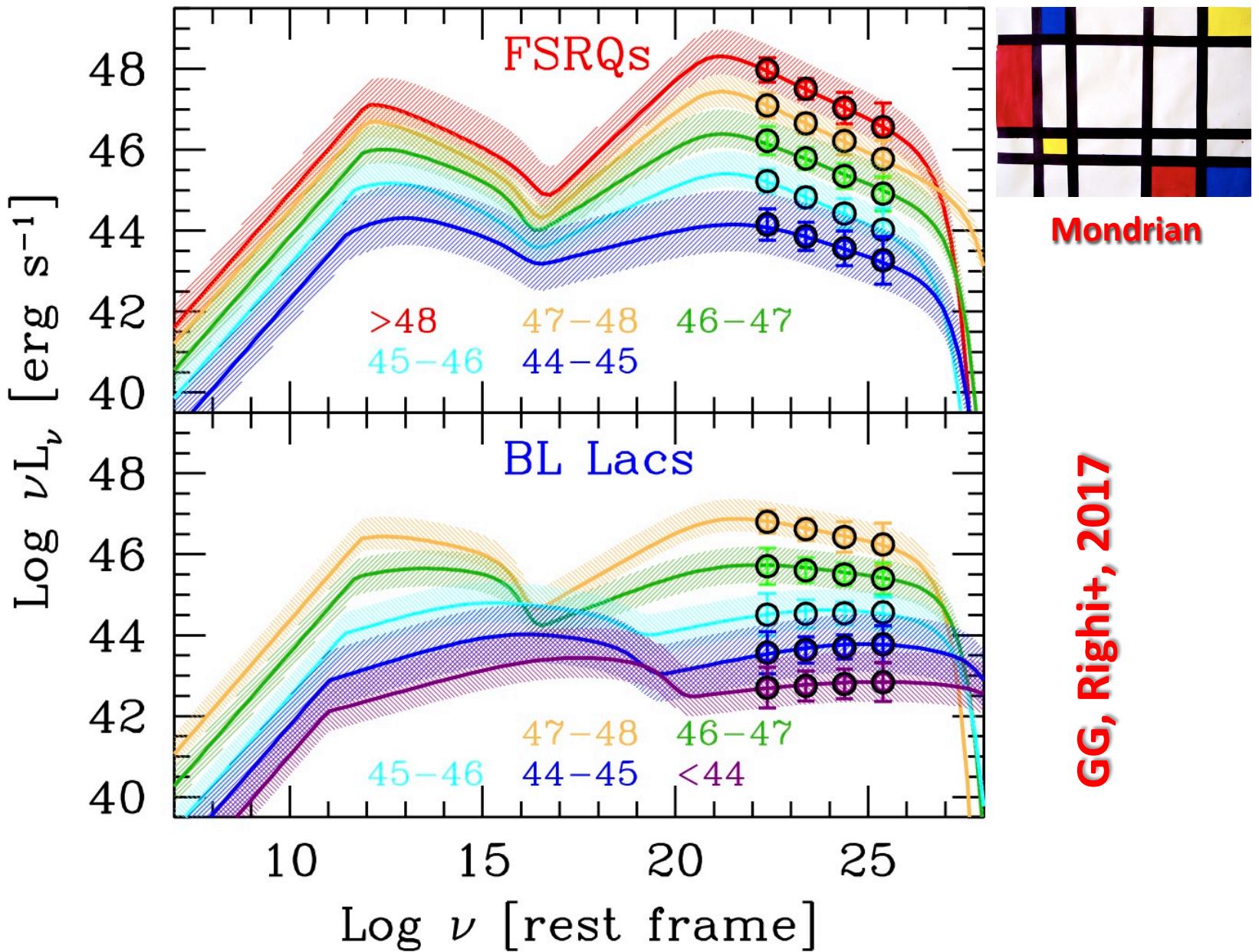


Matisse

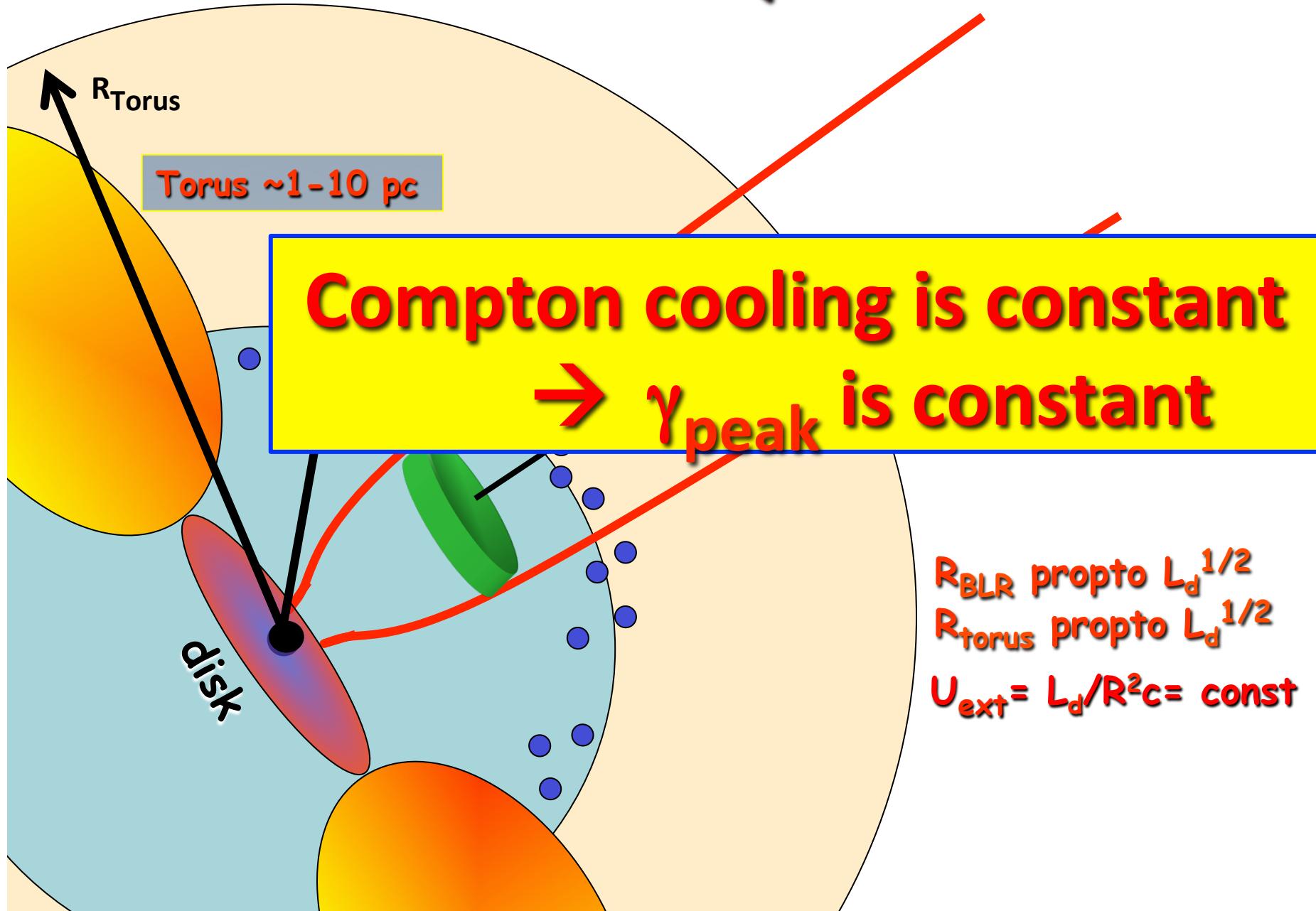
GG, Righi+, 2017

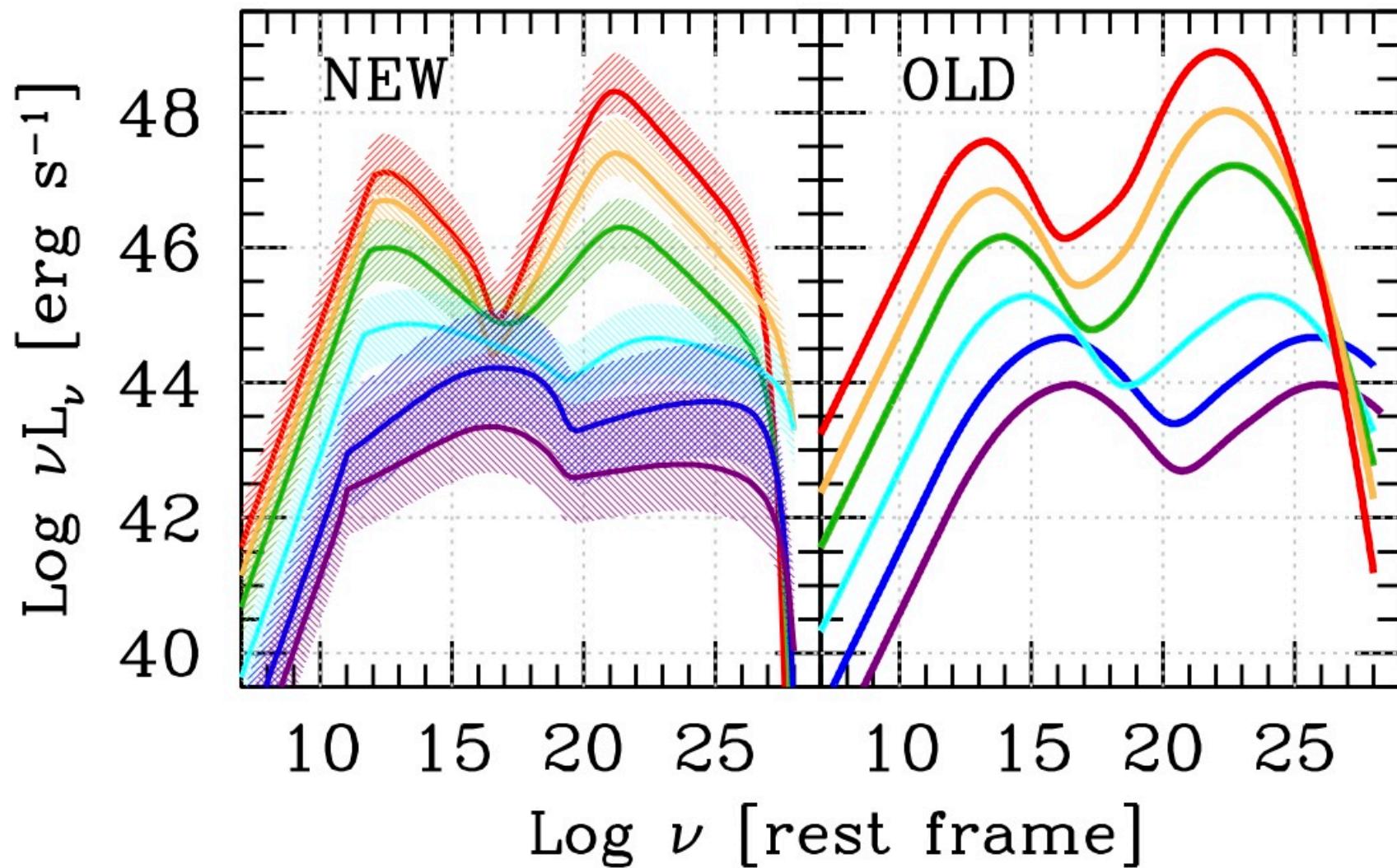




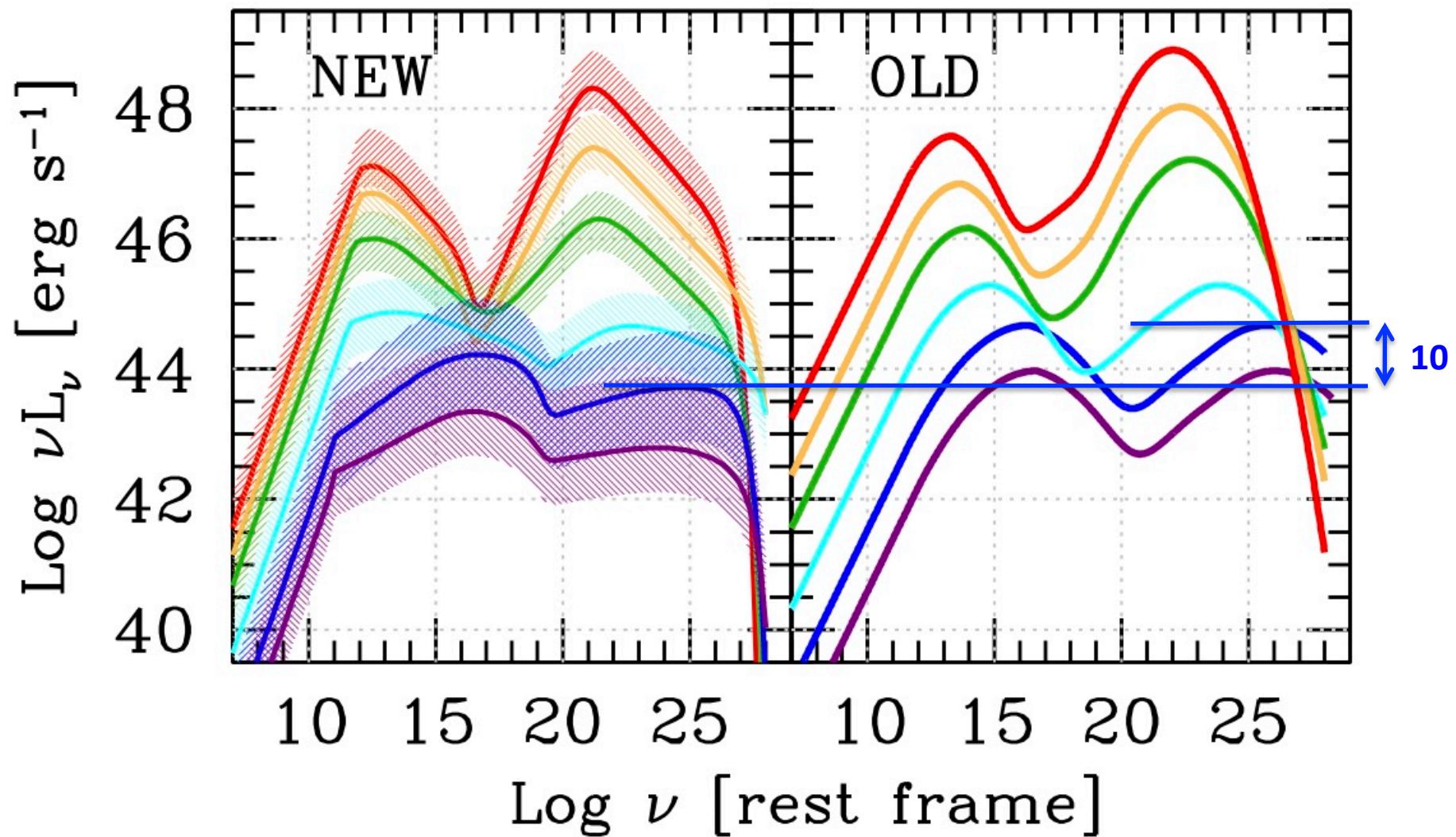


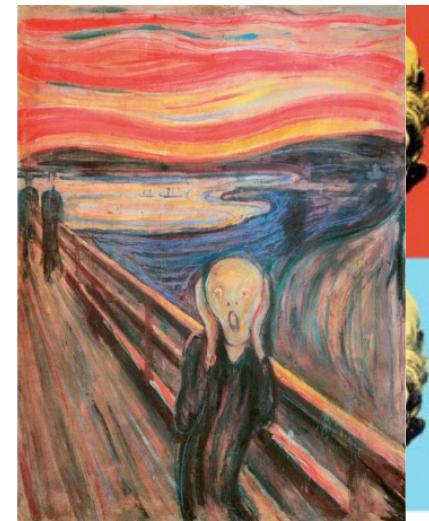
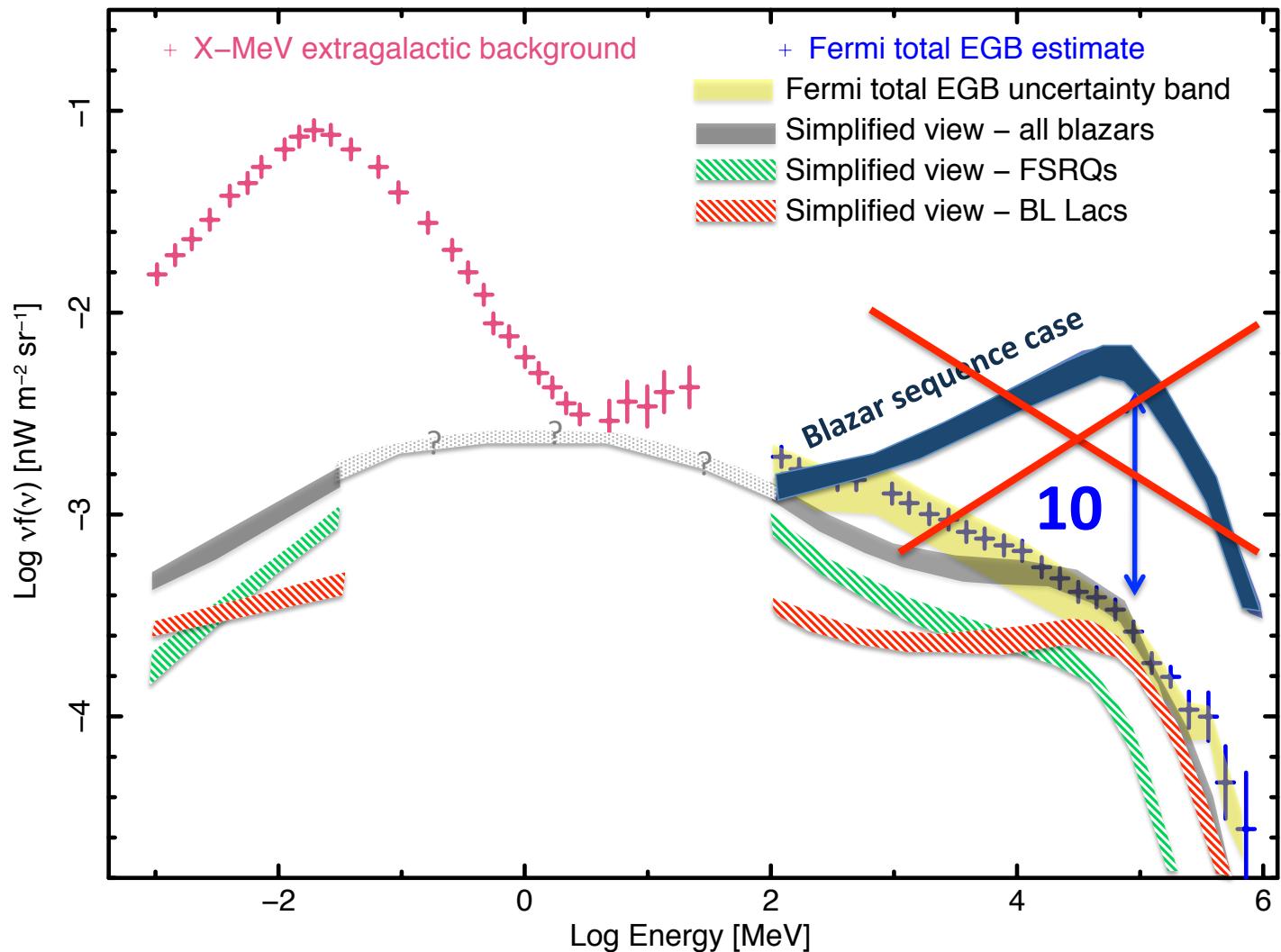
# FSRQs





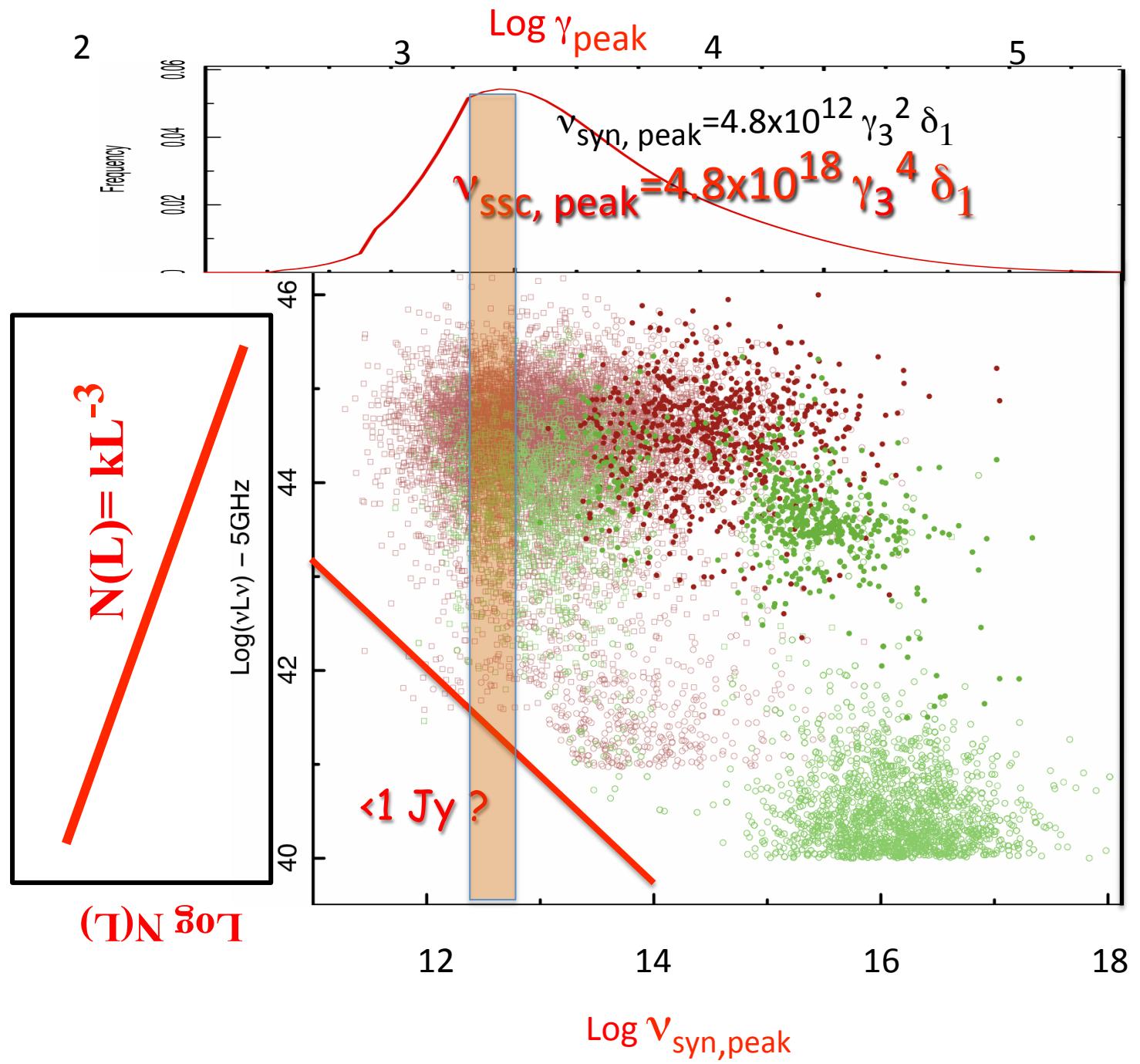
**Redder when brighter**

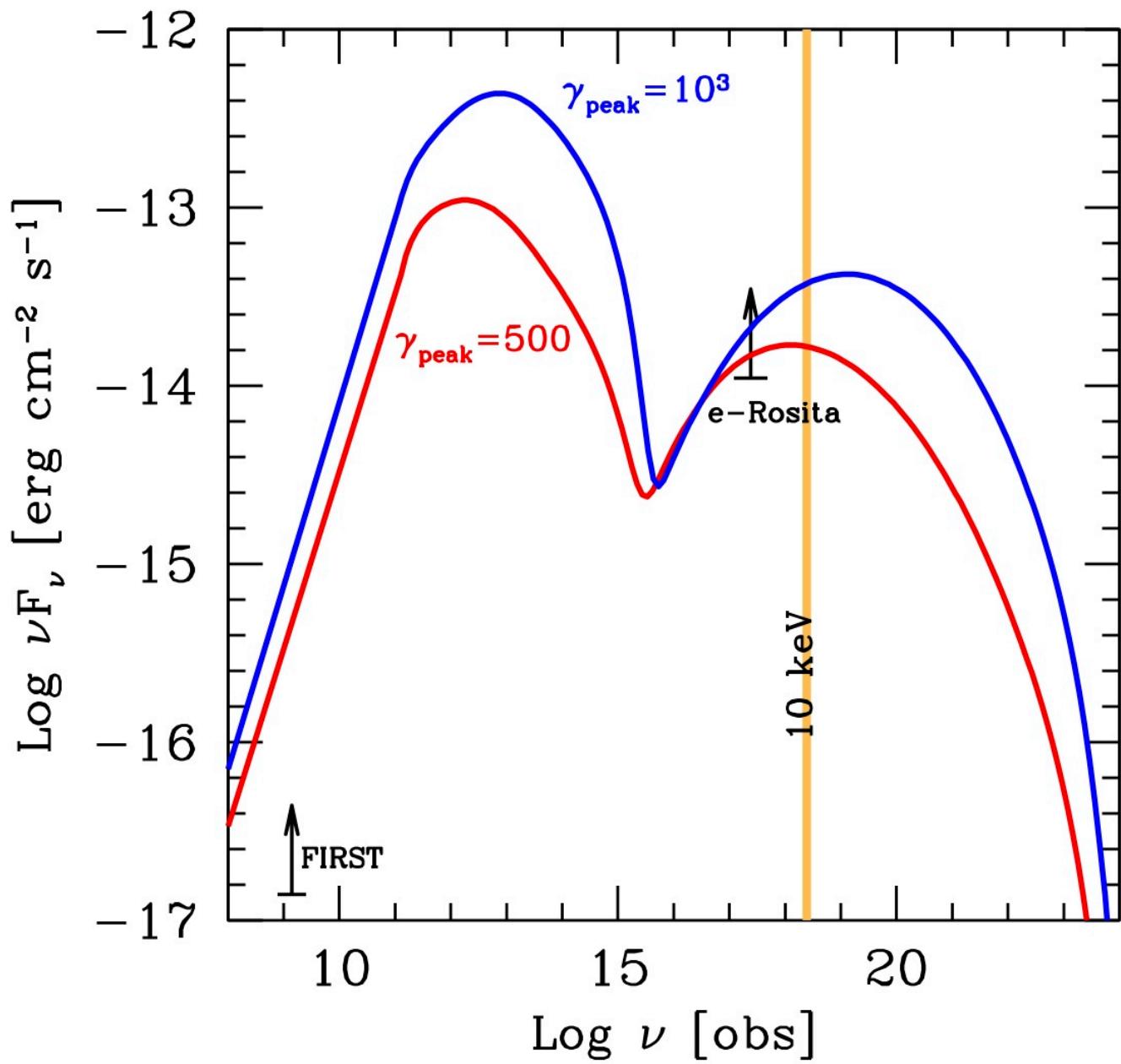




Munch

Giommi & Padovani 2015





# Conclusions

- ◆ The blazar sequence describes well the trend of the SED of the observed blazars
- ◆ It has a physical, simple, explanation
- ◆ BL Lacs and FSRQs behave differently
- ◆ This supports radiative cooling as an explanation for the sequence