Multiwavelength Cross-Correlations and Flaring Activity in Bright Blazars

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Half Century of Blazars and Beyond Turin 2018



Introduction

5y[10-7 ph What's the relation between different emission regions? X-ray (2-10 keV) 11-11 513 d [k]m] Optical, UV (R. V, B, U, W2 bands) How do the flaring properties of one wavelength relate to another? PD [%] Optical polarization degree (V, no filter) VPA Optical polarization ang (V no filter) Where and how are γ -rays produced in blazars jets? Near Infrared (Ks. K. H. J bands)

2009 2008 y-ray (>200 MeV) iio [JV] 15, 37, 230 GHz 800 [MJD - 54.000]

Abdo et al. 2009

3C279

Sample

0.76-m Katzman Automatic Imaging Telescope (KAIT)

40-m Owens Valley radio observatory at 15 GHz (**OVRO**)

Fermi γ-ray space telescope 30-d binned light curves

~8 year long light curves

145 blazars: 93 BL Lacs, 45 FSRQs, 5 Unclassified

(75 LSPs, 28 ISPs, 24 HSPs, 18 No Info.)





Tools

Time-lags: Discrete correlation function (DCF, Edelson & Krolic 1988) Significance of DCF: Source randomization method (Cohen et al. 2014)

Flare identification: Bayesian blocks (Scargle et al 2013)





Flaring statistics

For the optical—γ-ray and optical—radio correlated sources 78% of radio and 86% of γ-ray flares are associated with an optical counterpart.

~60% of radio and γ-ray flares are associated for sources with a radio—γ-ray correlation.

Associated flares have higher amplitudes than orphan flares for the optical—y-ray and optical—radio correlated sources

No difference in associated and orphan flares in γ-ray for the radio—γ-ray correlated sources.



J2253+1608

J0237+2848



Probing the optical to y-ray correlation

All sources from KAIT +Steward observatory + SMARTS +3-day,10-day binned γ-ray light curves



Summary

Explored the time-lags between optical—radio—y-rays for a large number of sources

Radio is typically lagging both optical and y-rays, exceptions exist!

There is a strong connection between optical and y-rays suggesting IC scattering as the dominant mechanism

Associated flares tend to have higher amplitudes than orphan flares.

Additional analysis is underway!

Thank you!

Additional slides

Time-lags: 30d versus 3d light curves



Block analysis: 30d versus 7d light curves



Are non-zero lags really non-zero lags?

