

The IFAE/UAB Raman LIDAR for the CTA-North

AtmoHEAD 2018

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for the CTA Consortium (see www.cta-observatory.org)

The IFAE/UAB Raman LIDAR

Outline

- CTA atmospheric characterization of observed fields-of-view
- The design of the IFAE/UAB Raman LIDAR
- First commissioning results
- Future plans

Atmospheric characterization of the observed line-of-sight ved by CTA

Requirements for systematic uncertainties on energy scale, due to atmospheric effects, are very ambitious

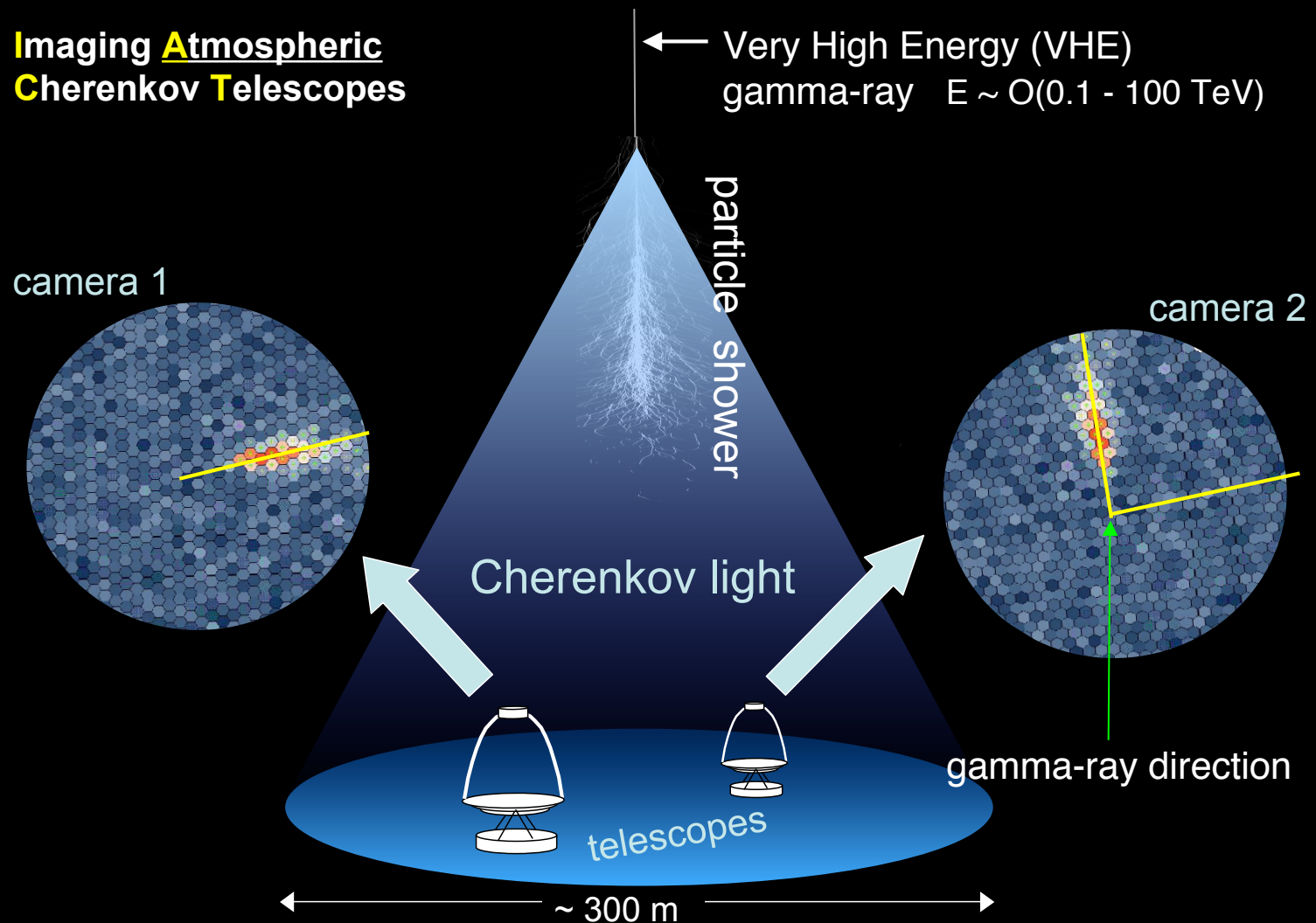
Part	currently achieved	goal for CTA	comments
Simulation codes	5%	1-2%	MC working group
Simplifications in MC	2%	2%	
Cherenkov light creation	5%	2%	mainly molecular profile
Ozone absorption	3%	1%	Potential vorticity, spectrometer
Molecular extinction	2%	1%	Radio sondes and GDAS
Cirrus layers extinction	5-10%	1-2%	Raman LIDARs and FRAM
Boundary layer extinction	5-10%	1-2%	Raman LIDARs and FRAM
Scattered Cherenkov light	<1%	<2%	

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Introduction to CTA

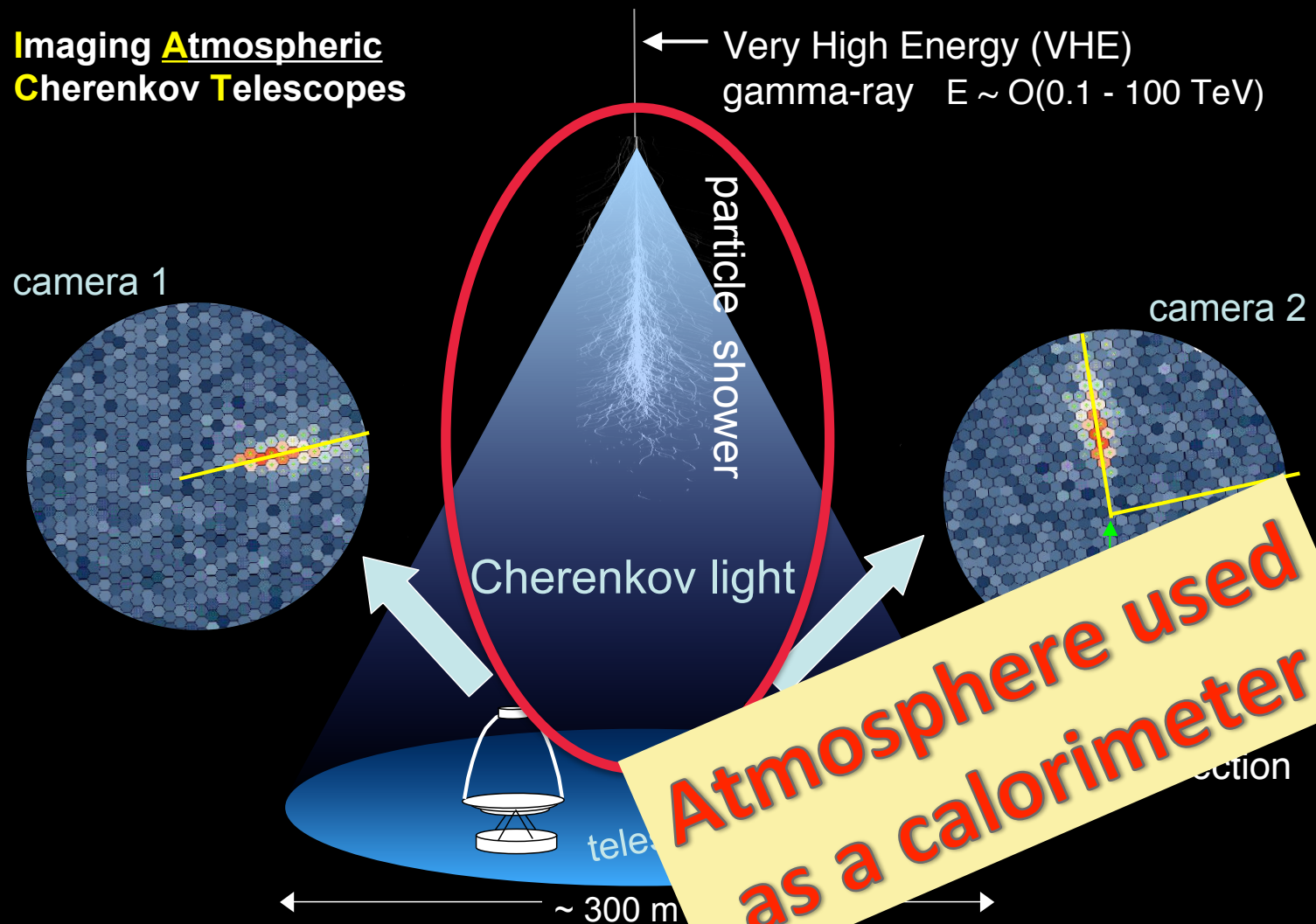
Imaging Atmospheric Cherenkov Telescopes



Introduction to CTA



Imaging Atmospheric Cherenkov Telescopes

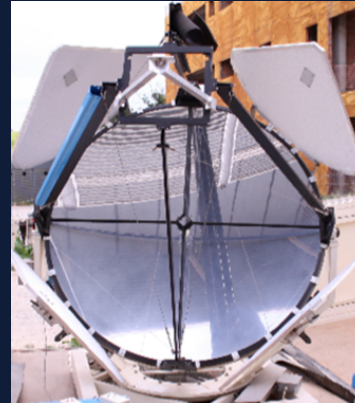


Characterization of observed field-of-view



Need to continuously characterize:

1. The profile from ground to 25 km distance
 - GDAS/ECMWF (see P. Munar, Monday)
 - Raman LIDARs
2. The extension of clouds across the FOV of 10° , determination of time slots with equal atm. conditions
 - FRAM (see P. Janecek, Tuesday)
3. For cross-checks:
 - The Cherenkov Transparency Coefficient (see S. Stefanik, Tuesday)



IFAE/UAB LIDAR



FRAM



LUPM LIDAR

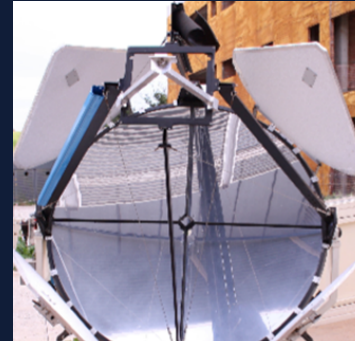
CEILAP LIDAR

Characterization of observed field-of-view



Need to continuously characterize:

1. The profile from ground to 25 km distance
 - GDAS/ECMWF
 - Polar LIDARs



2. The conditions across different time slots

Instruments and analysis algorithms determine:

- Intervals of stable atmospheric conditions
- Corrections for the Instrument Response Functions
- Associated statistical and syst. uncertainties

3. For cross-checks:
 - The Cherenkov Transparency Coefficient

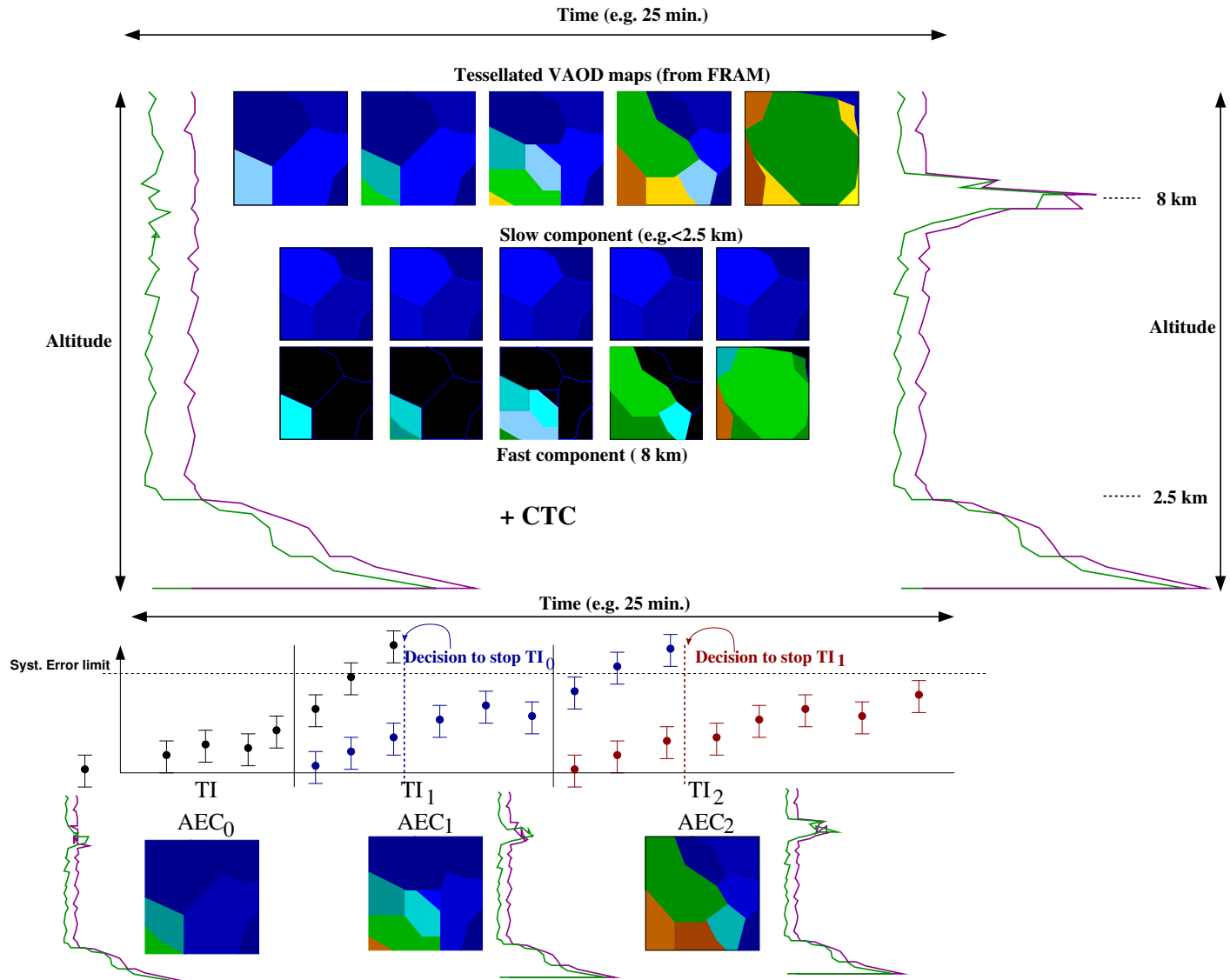


CEILAP LIDAR



LUPM LIDAR

Correction of the instrument IRF

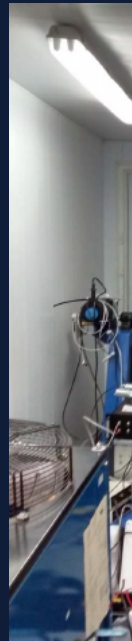


Requirements for a CTA Raman LIDAR



Need to continuously characterize:

1. The profile from ground to 25 km distance
 - Raman LIDARs
2. The extension of clouds across the FOV of 10° , determination of time slots under equal conditions
 - FRAM
3. For cross-checks:
 - The Cherenkov Transparency Coefficient



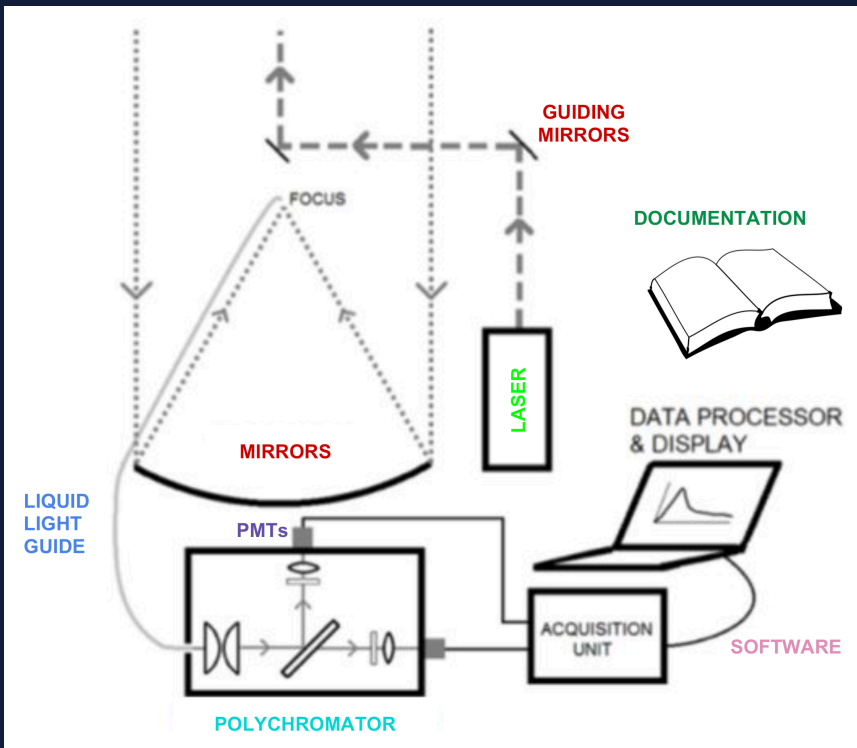
- Full characterization of the atmosphere up to 25 km distance, within 1 minute with Raman capabilities.
- Full sky coverage up to 60° zenith angle.
- At least 2 lines within sensitive window of the CTA photo-detectors
- Characterization of ground-layer aerosols
- Low distance to full overlap range
- Inclusion to CTA internal communication
- OPC-UA compatible communication, standard CTA states
- Low maintenance and failure rates
- Safe operation

Chosen solution for the IFAE/UAB design



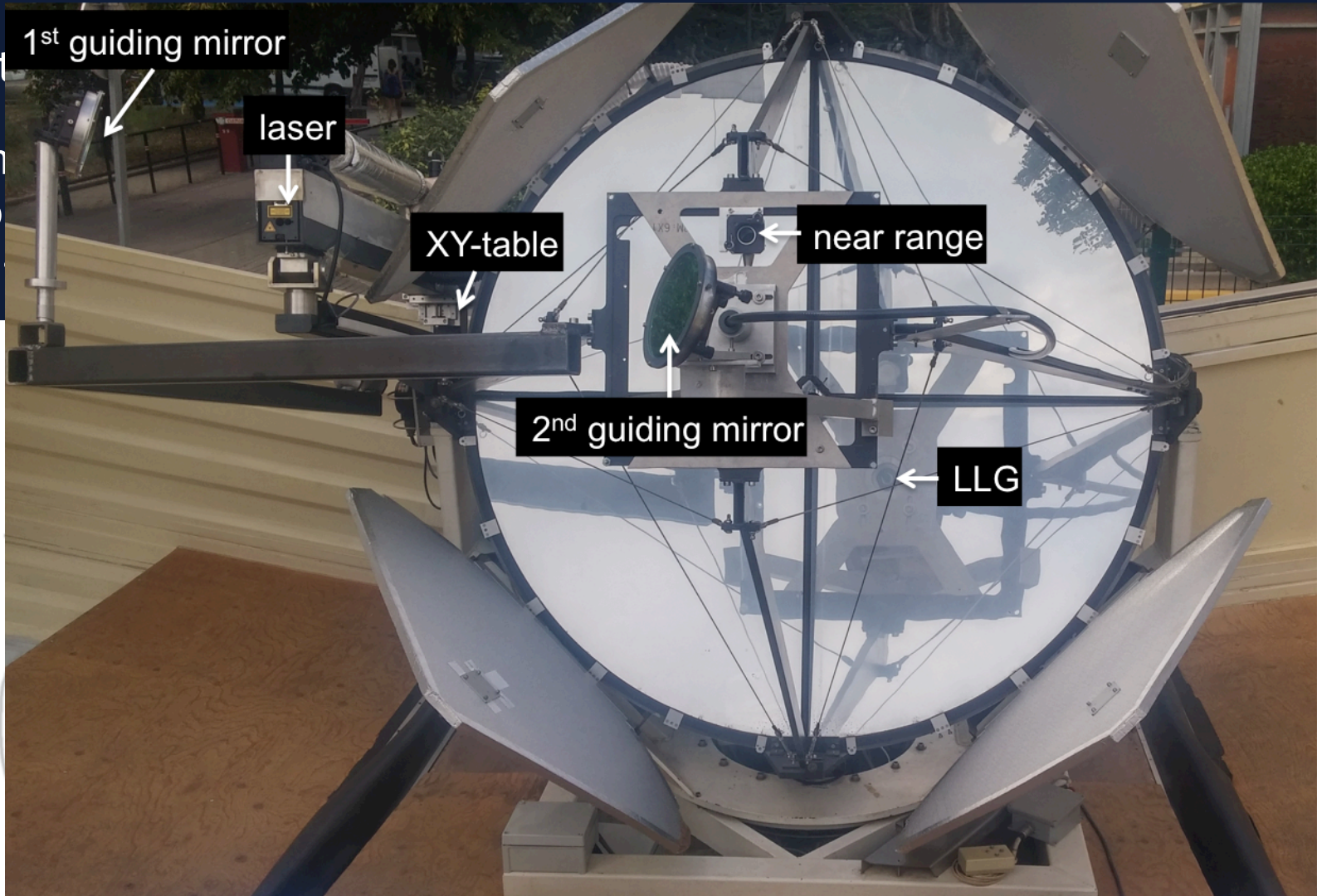
Need to continuously characterize:

1. The profile from ground to 25 km distance
 - Raman LIDARs



- Recycle former CLUE telescope and container **1.8 m (!) parabolic mirror**
- **2 + 2 configuration:**
 - **355 nm + 532 nm** elastic lines
 - **387 nm + 607 nm** N₂ Raman lines
- Laser: Brilliant Compact Q-Switched Nd:YAG 1064nm + frequency tripler head
 - Rate: 20 Hz
 - Pulse: 5 ns
 - Beam divergence: 0.5 mrad
- Dichroic guiding mirrors → coaxial
- Polychromator in-house
- Acquisition: standard LICEL units
- Dedicated near-range optics and readout

Chosen solution for the IFAE/UAB design



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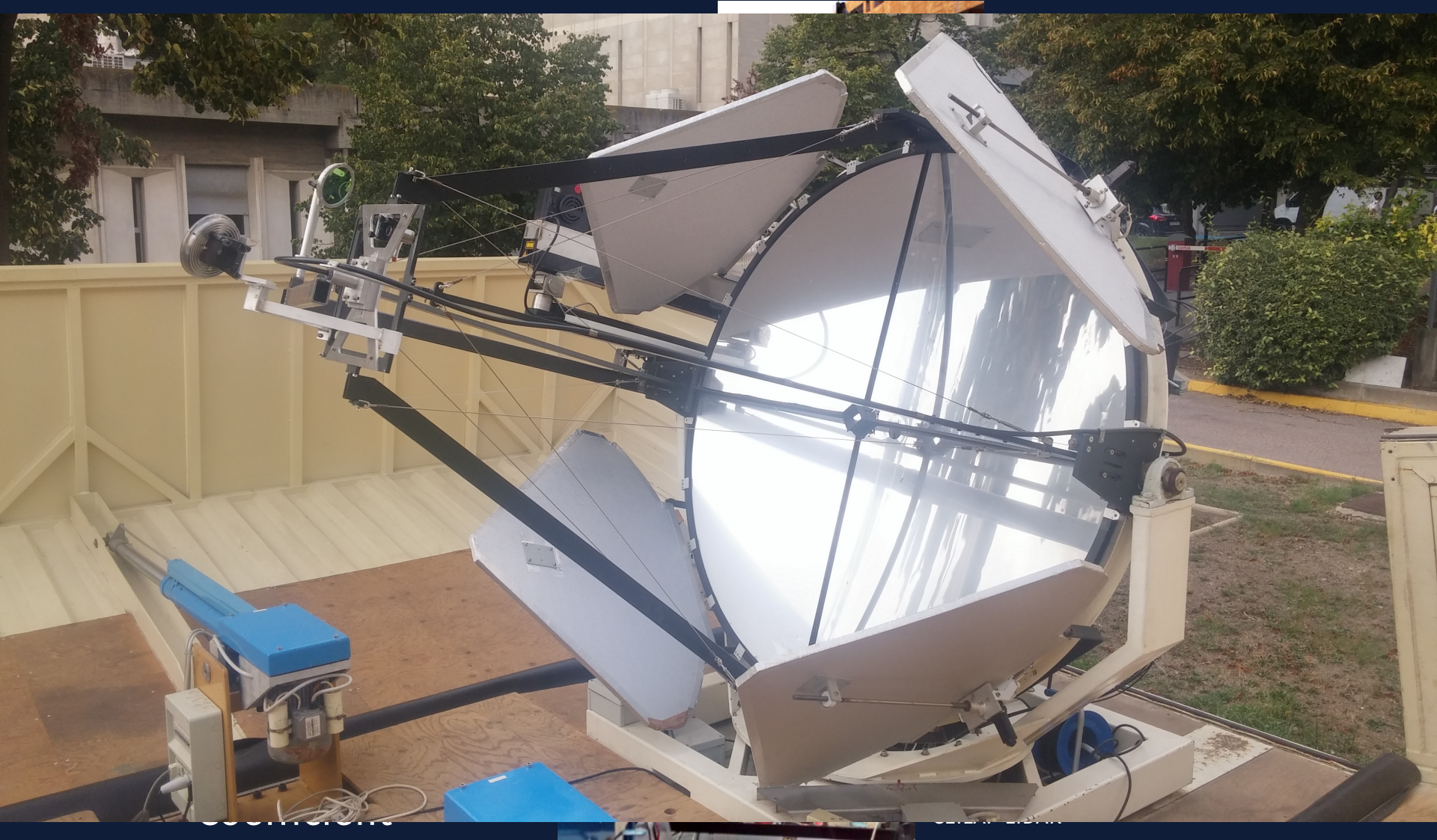
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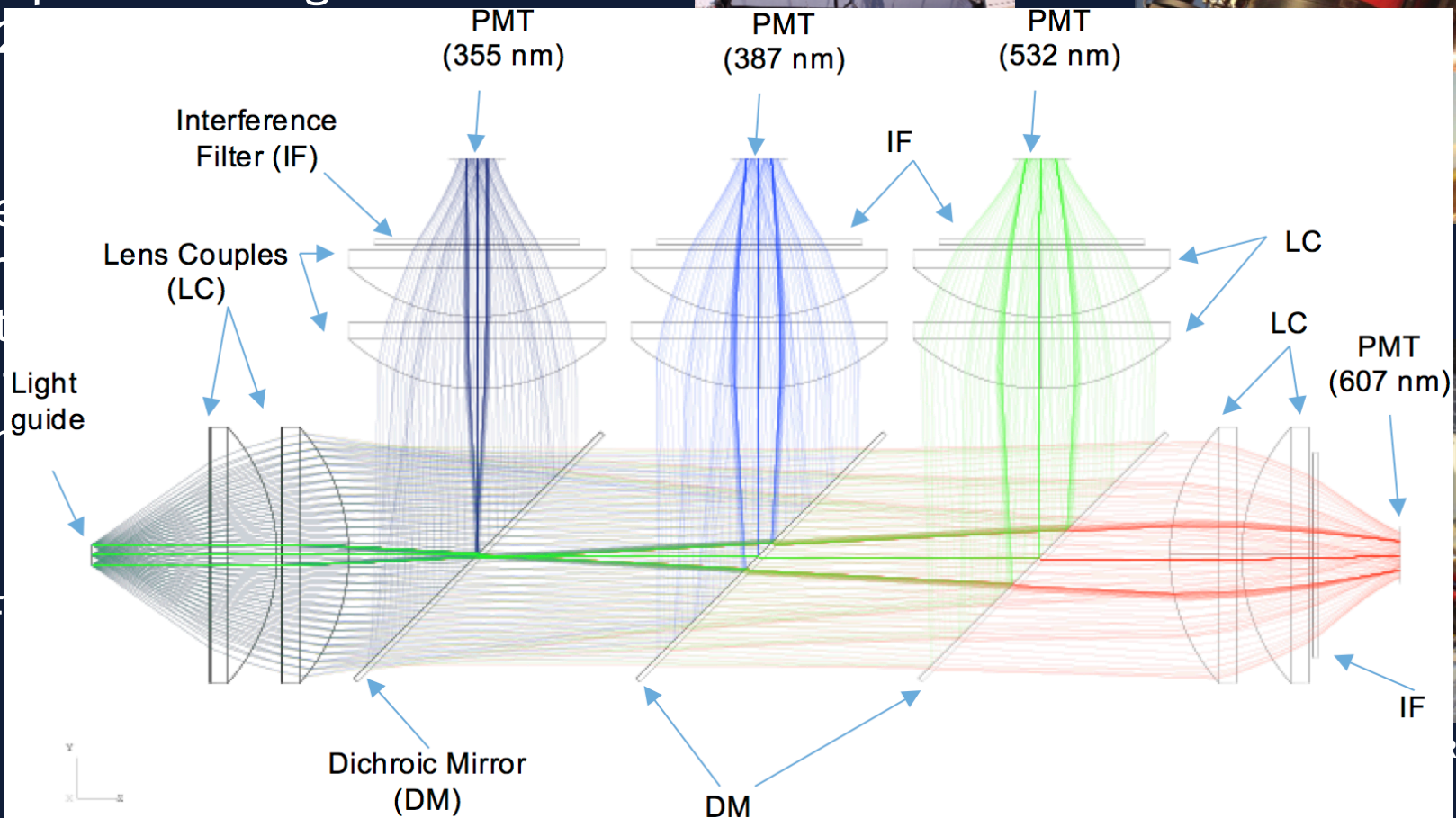
Chosen solution for the IFAE/UAB design



Chosen solution for the IFAE/UAB design Polychromator



Large aperture telescope and PSF of 6 mm
require 100 mm \varnothing optics in polychromator

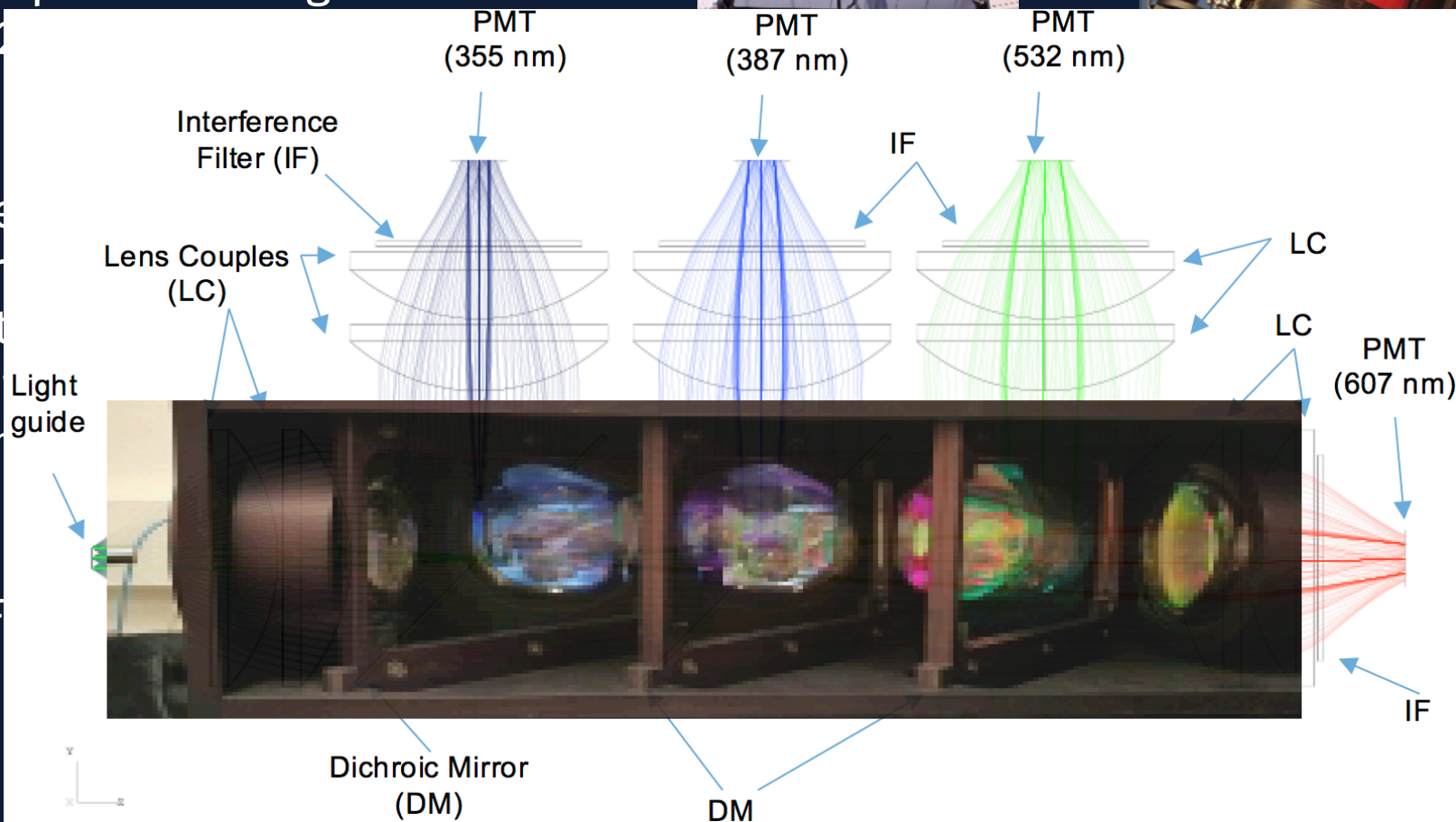


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Chosen solution for the IFAE/UAB design Polychromator



Polychromator was designed by CNR Padova and built by at IFAE workshop



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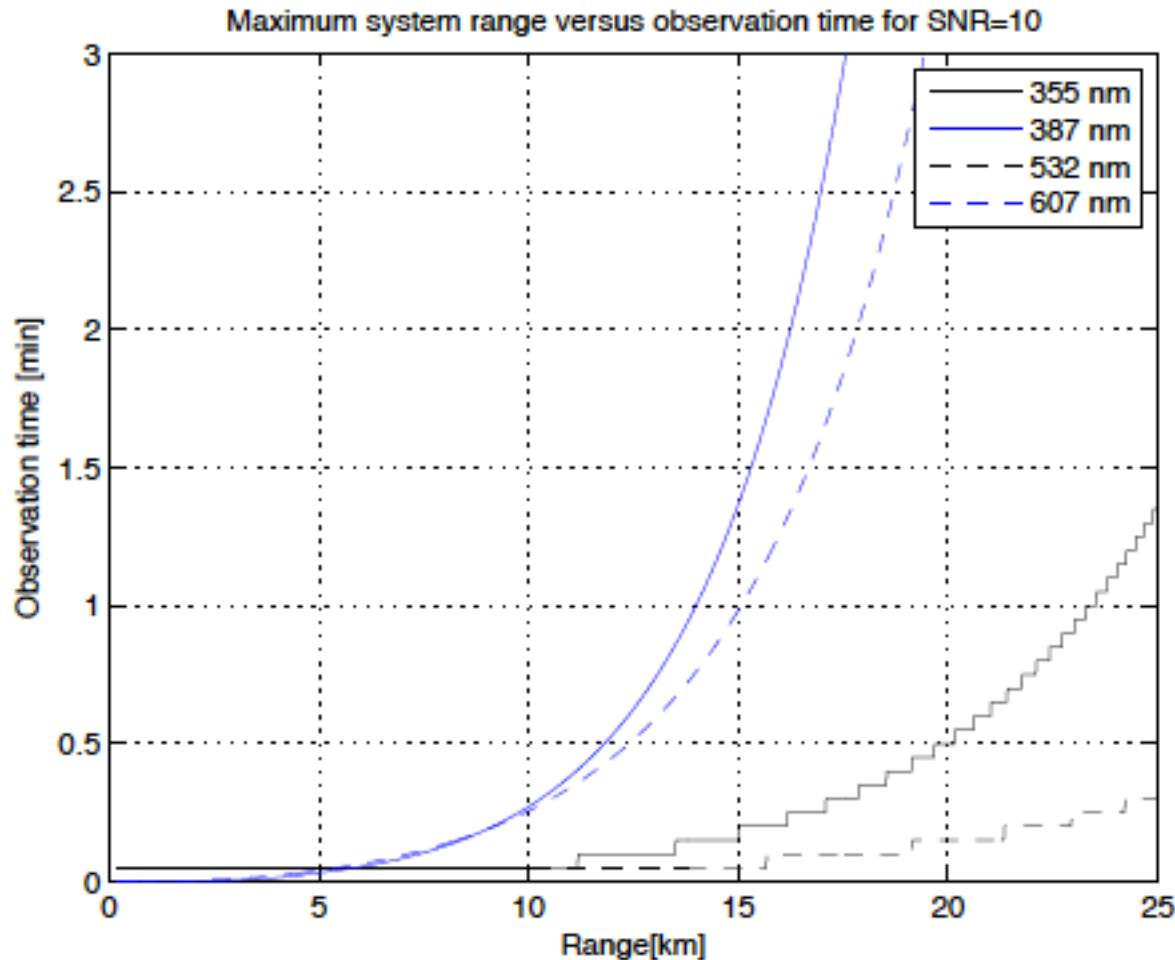
Chosen solution for the IFAE/UAB design Polychromator



**Light leakage from any elastic to any Raman
channel $< 2 \times 10^{-7}$ (!)**



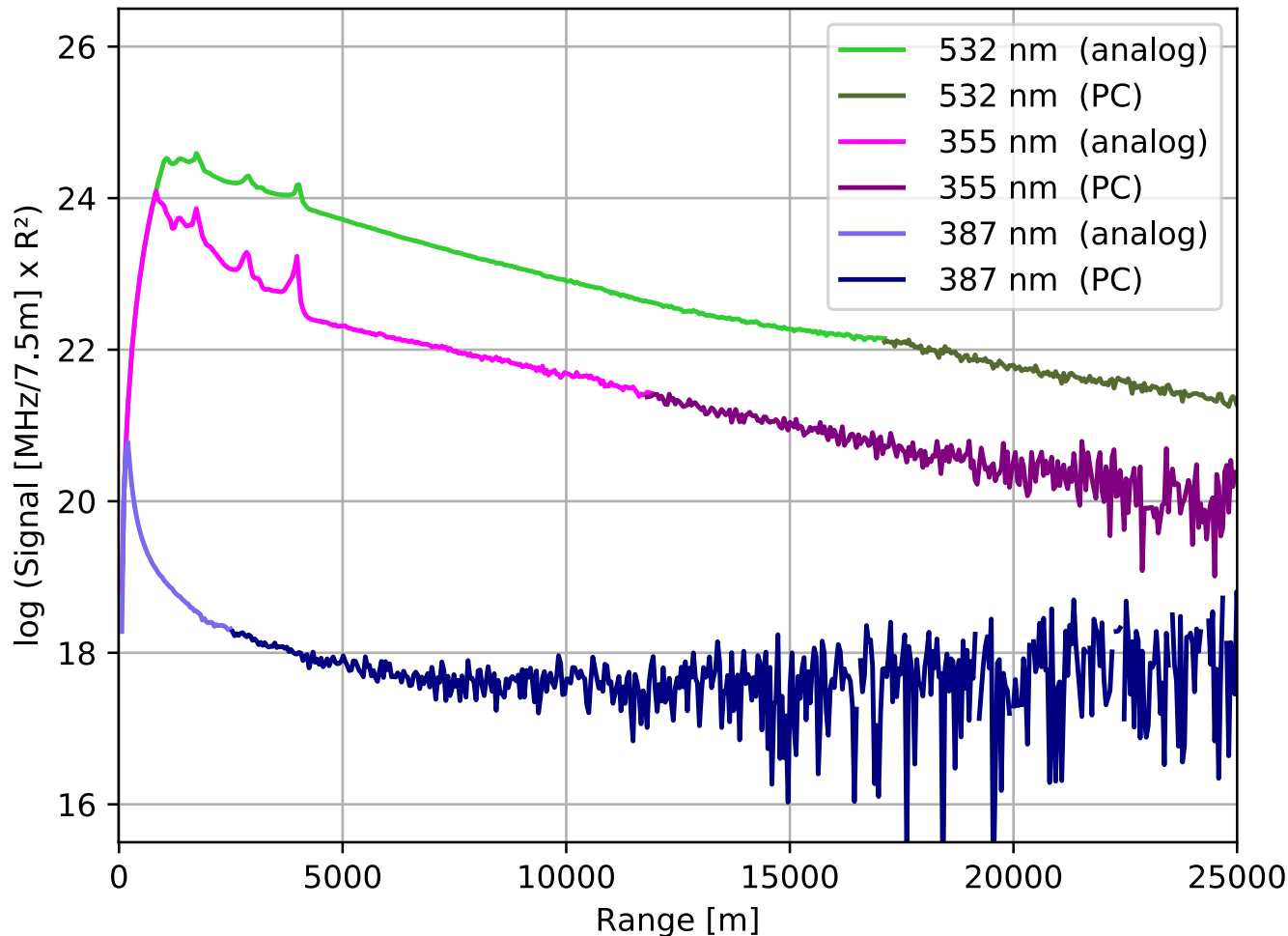
Link budget study for sensitivity



- Sensitivity is limited by the two Raman lines
- Should reach 15 km within less than one minute

“First light” results

500 shots @ 10 Hz



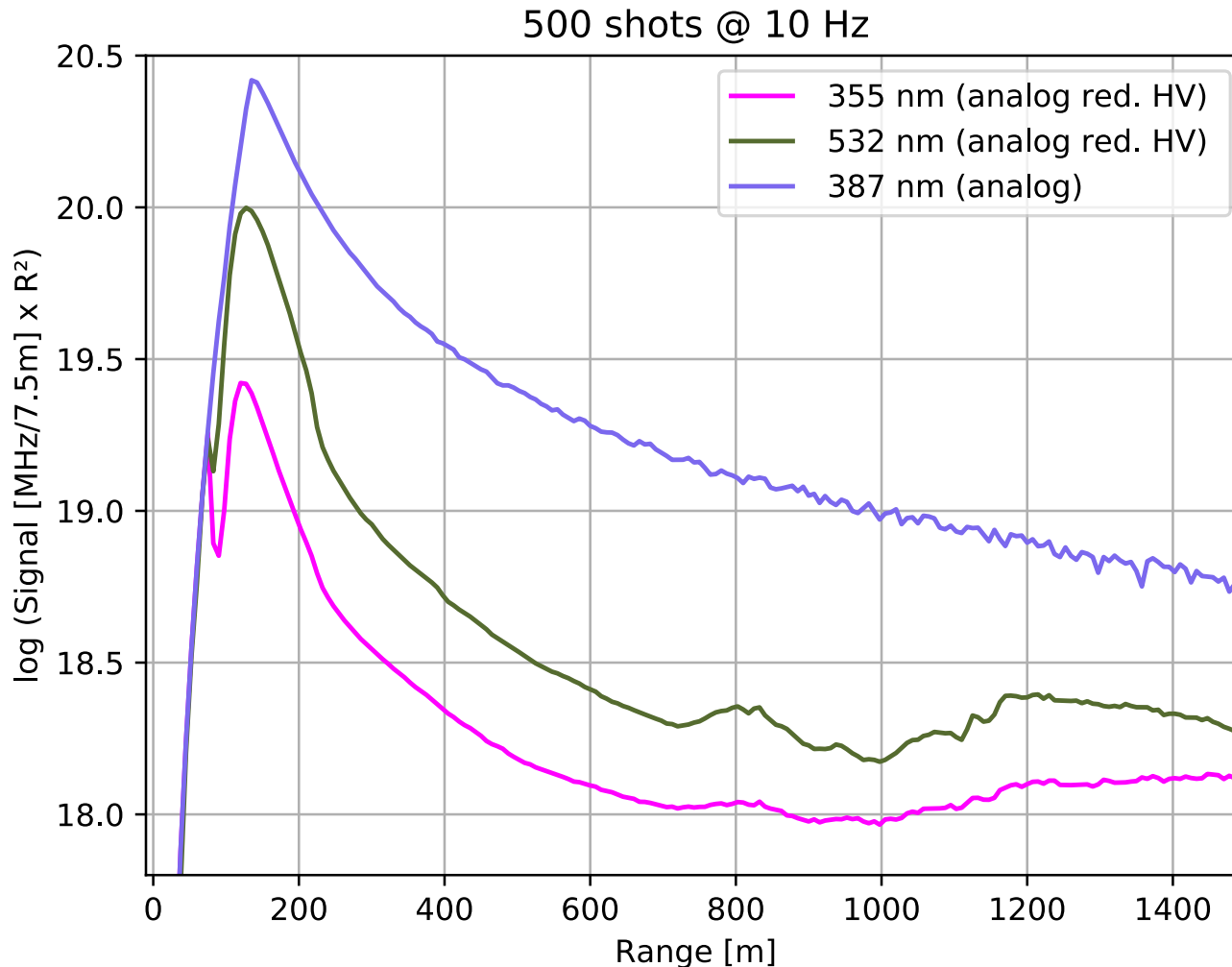
Currently reach within 50 seconds and 10 Hz trigger rate:

- 40 km with elastic lines
- 10 km with 387 nm Raman lines

Sensitivity can be further improved by:

- Raising laser frequency to 20 Hz
- Cleaning the primary mirror

“First light” results



Full overlap is reached at ≈ 150 m:

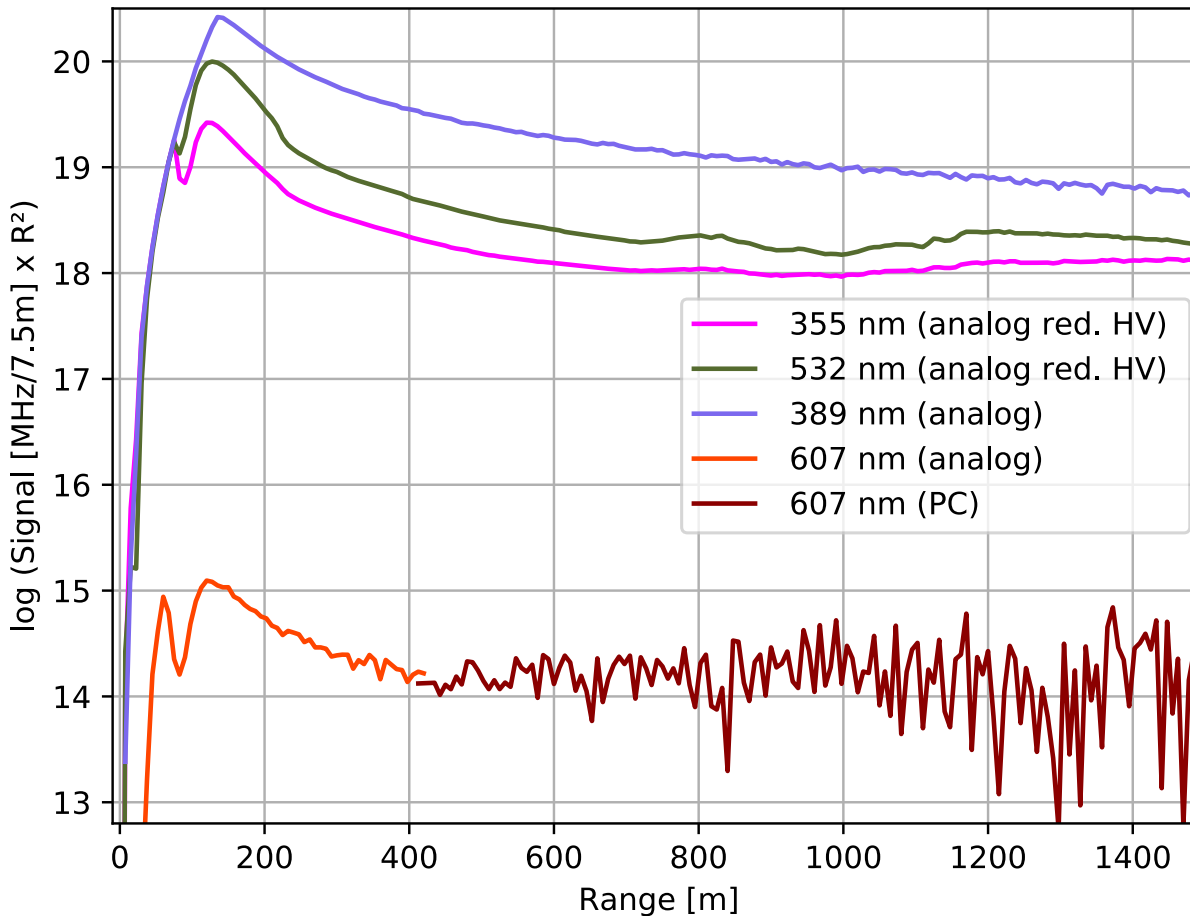
- for elastic lines with reduced gain
- for all Raman lines

Sensitivity to the near field can be further improved by:

- Dedicated near-range optics and readout (in preparation)

“First light” results

500 shots @ 10 Hz



Full overlap is reached at ≈ 150 m:

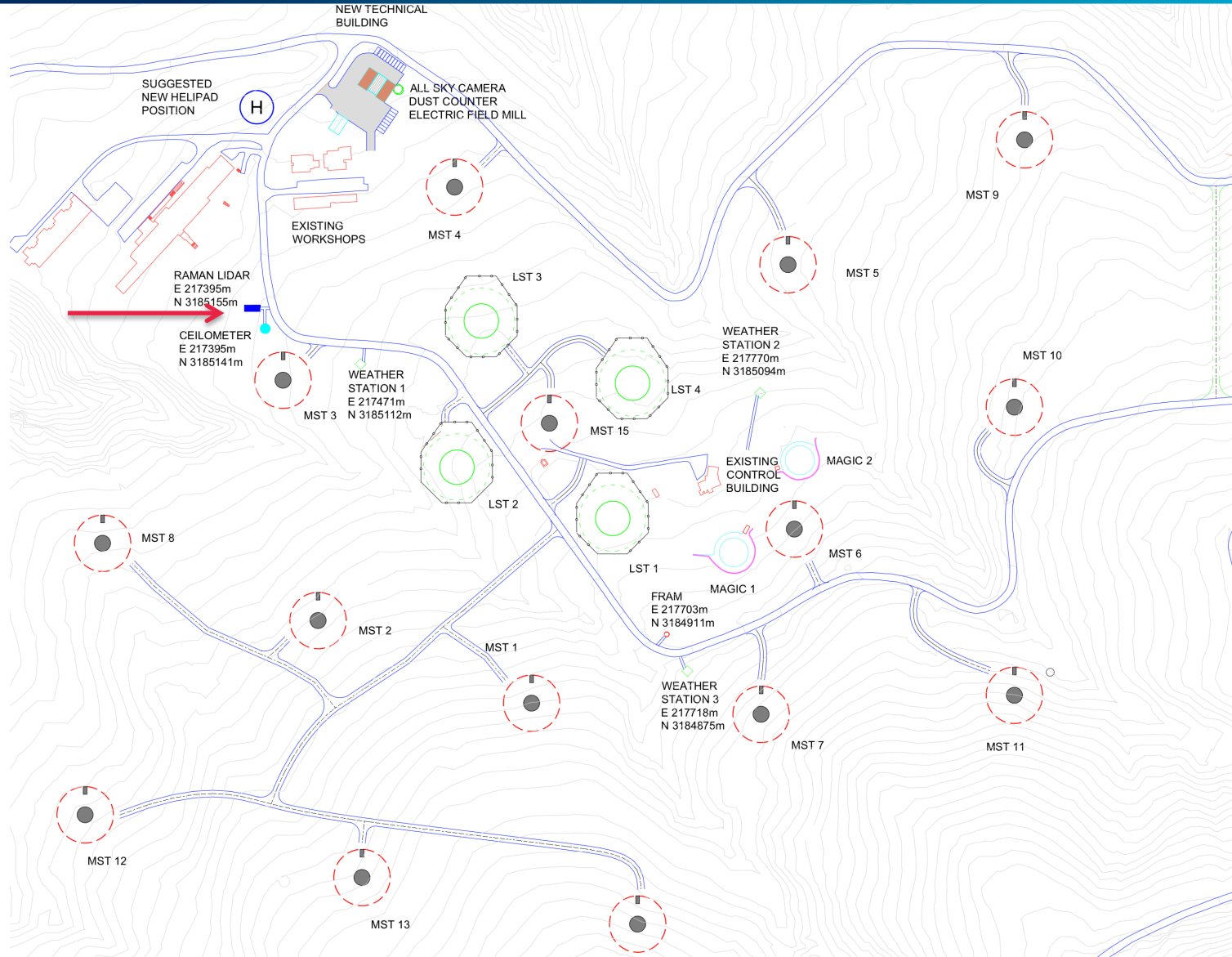
- for elastic lines with reduced gain
- for all Raman lines

Sensitivity to the near field can be further improved by:

- Dedicated near-range optics and readout (in preparation)

Very low quantum efficiency of 607 nm PMT still!

Location of the ORM



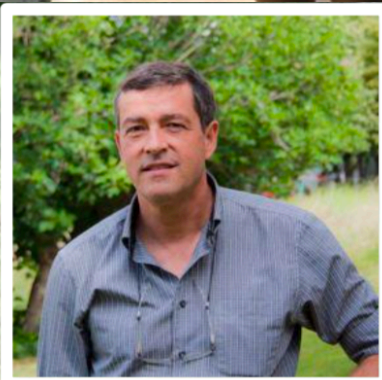
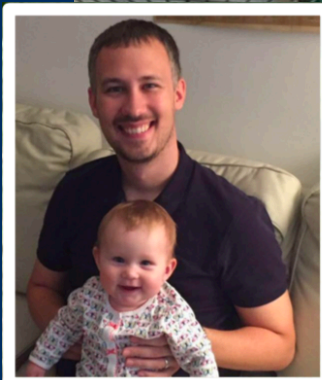
Conclusions

- During operation of CTA need to characterize the line-of-sight, across a 10° field-of-view.
- Requirements on systematic uncertainties are very ambitious, in order to make CTA a precision tool.
- Large aperture Raman LIDAR with powerful laser
- Basic design works now !
- Commissioning work still ongoing (part of it will be carried out at La Palma in 2019)



cherenkov
telescope
array

Thank you!

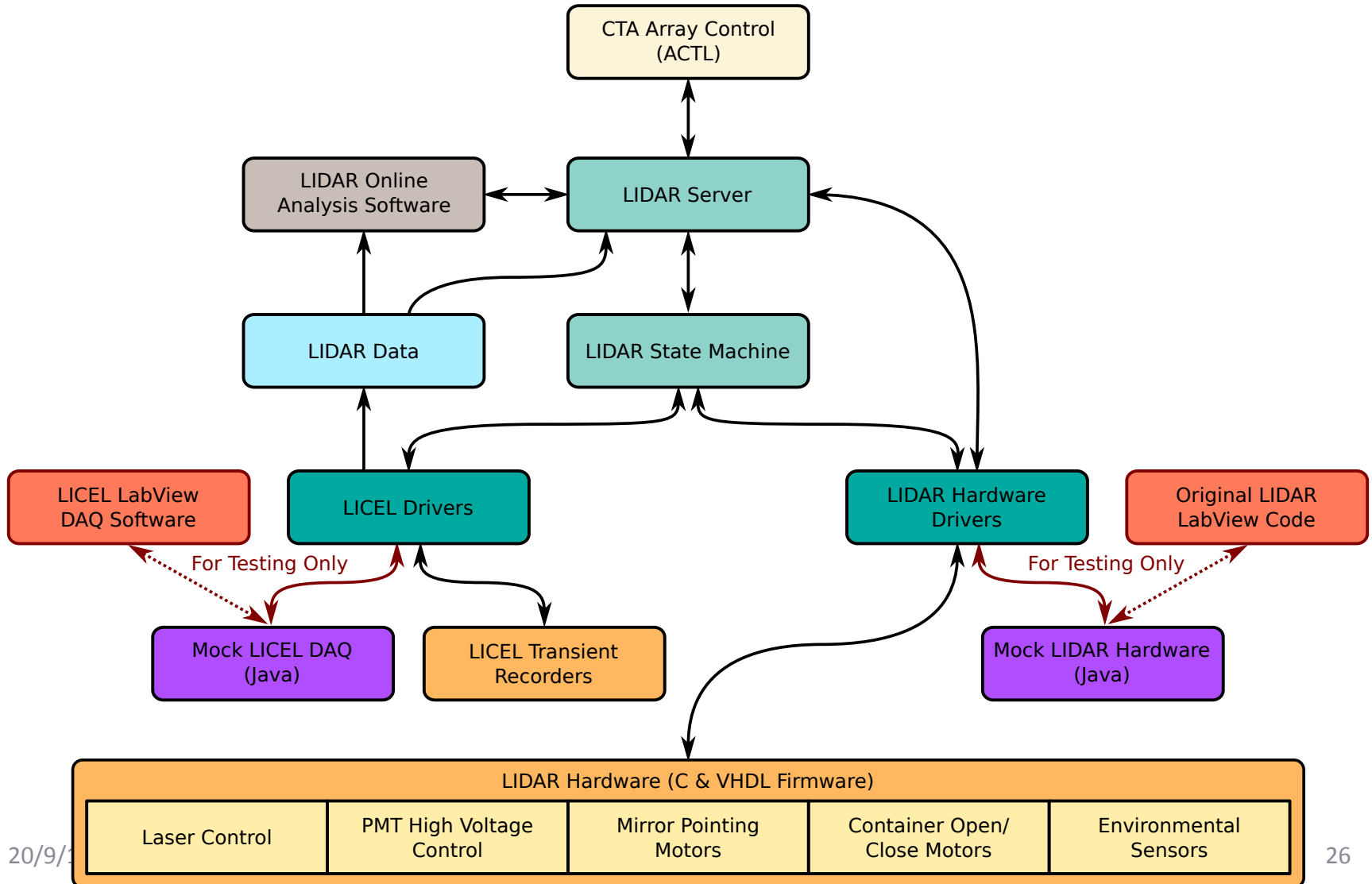




cherenkov
telescope
array

Backup

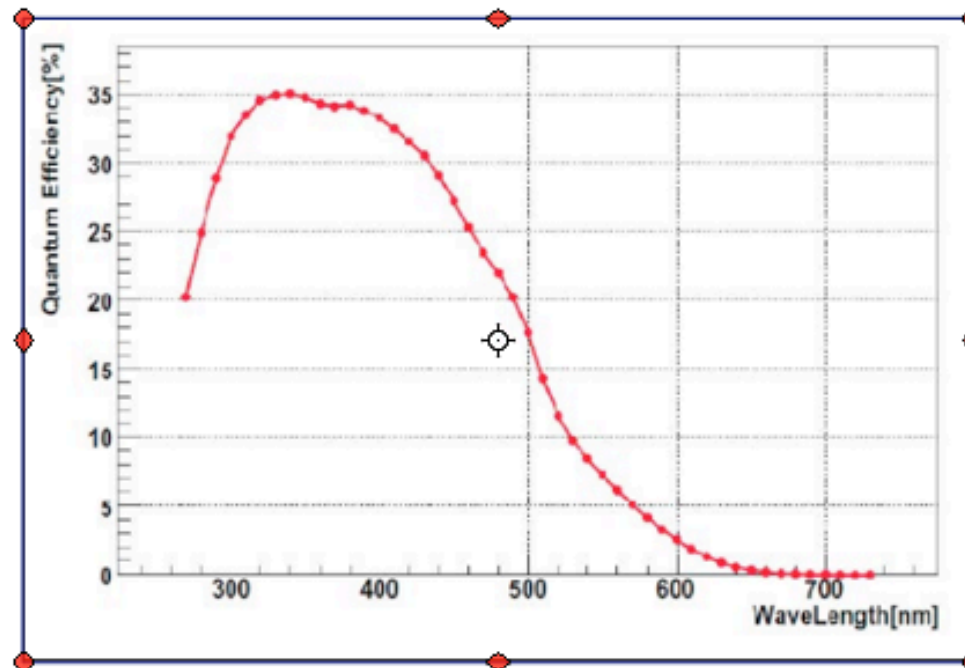
Architecture of control software



PMT characterization

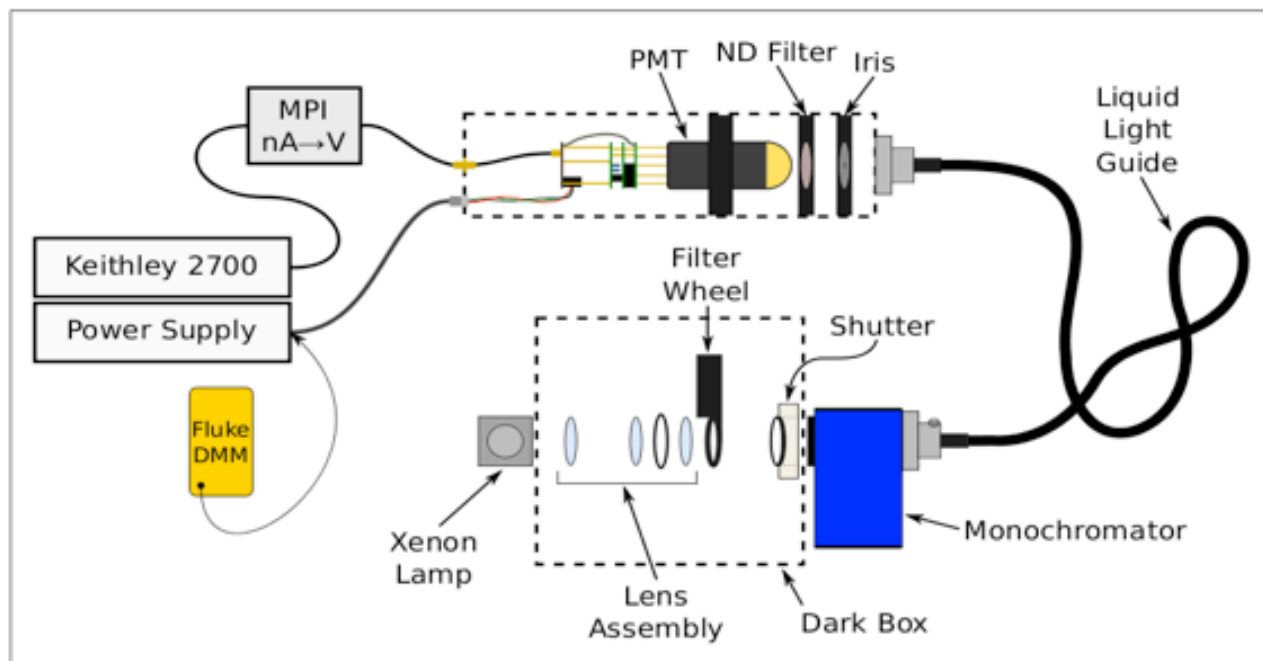
- 4 PMTs needed in the polychromator: -2 for the elastic channels (355nm and 532nm)
-2 for the Raman channels (387nm and 607nm)
- 4 Hamamatsu R11920-100 high quantum efficiency PMTs available:
ZQ6623, ZQ5819, ZQ6627, ZQ6622

$$C = QE \times HV\text{-dependent-gain}$$



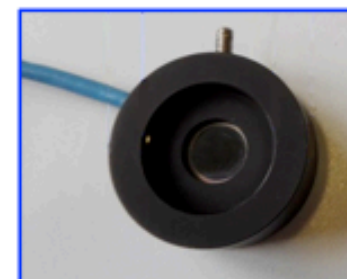
Orito, R. 2011, International Cosmic Ray Conference, 9, 170

Experimental setup

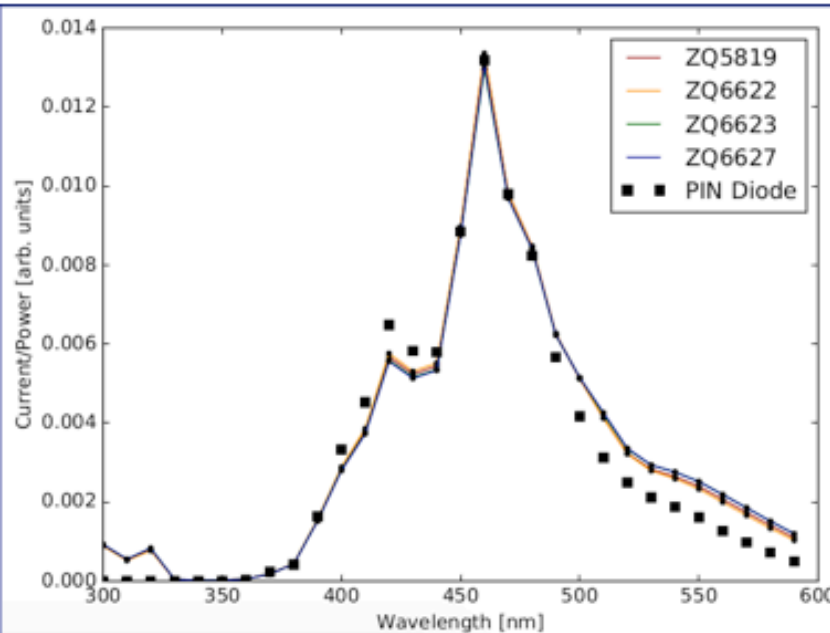


→ use of a calibrated Newport 818-UV
PIN photodiode

→ $V=1200V$ (chosen arbitrarily)



Results

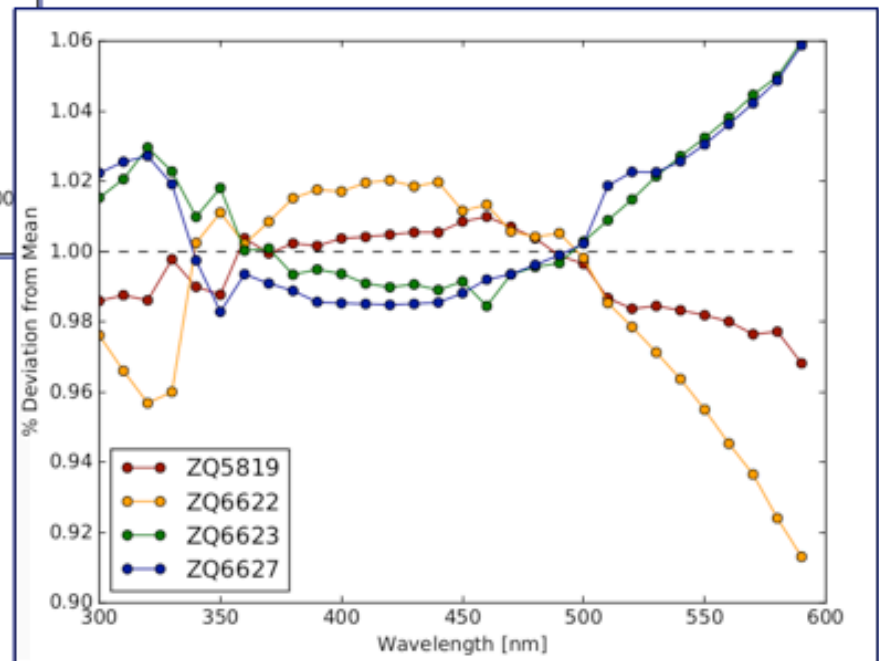


- Increment the wavelength by 10nm (± 2 nm) from 300nm to 600nm
- Measure of the PMT current with the shutter open (closed)
- Subtraction of the background from the PMT measurements
- Comparison with the PIN-diode subtracted-background data (PIN-diode data multiplied by the PMT QE and ND filter transmission)

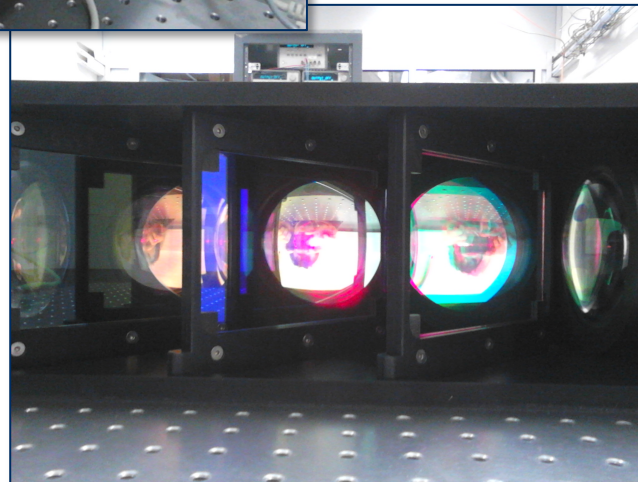
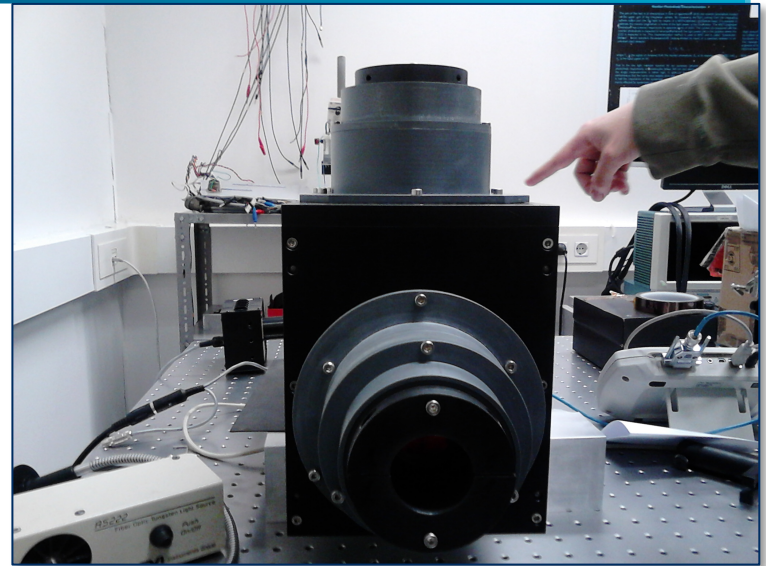
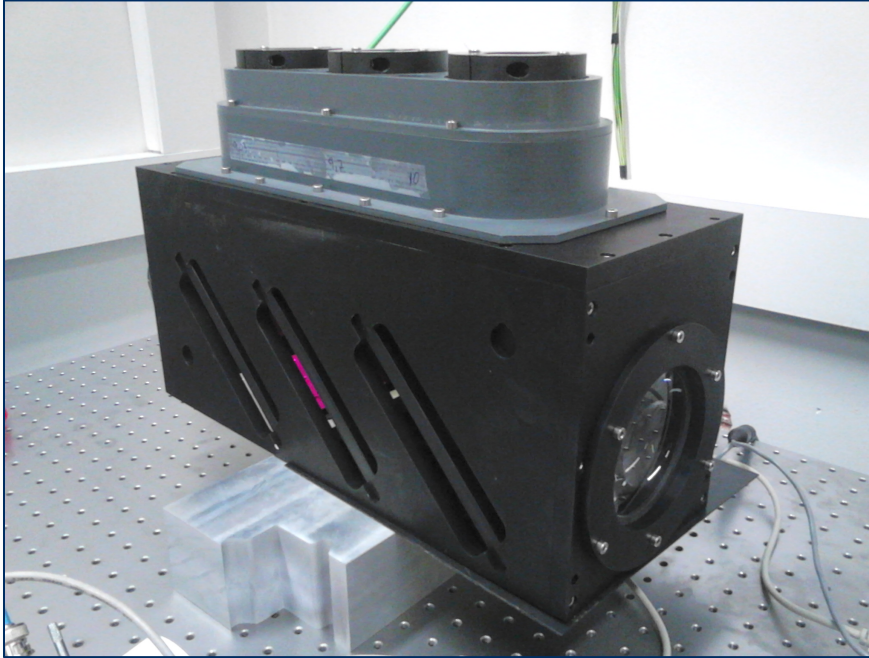
Elastic channels

- 355nm \rightarrow ZQ5819
- 387nm \rightarrow ZQ6622
- 532nm \rightarrow ZQ6623
- 607nm \rightarrow ZQ6627

Raman channels

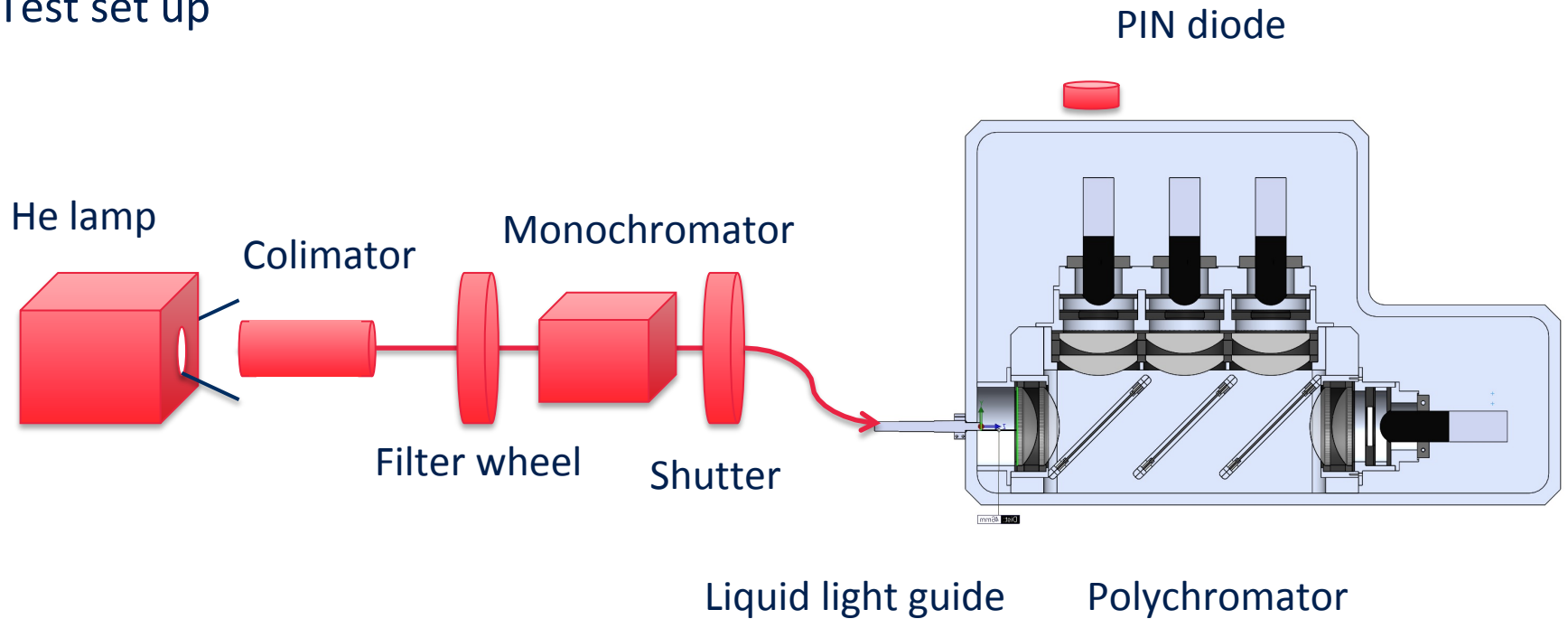


Polychromator



Characterization of the polychromator response

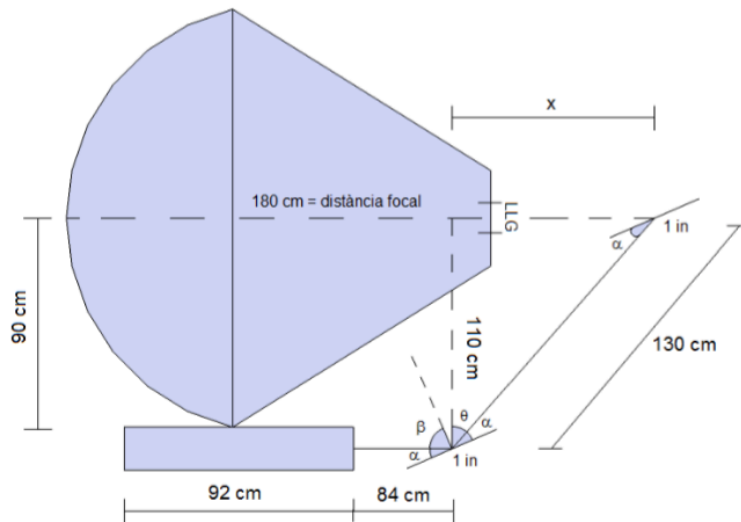
Test set up



Status prototype

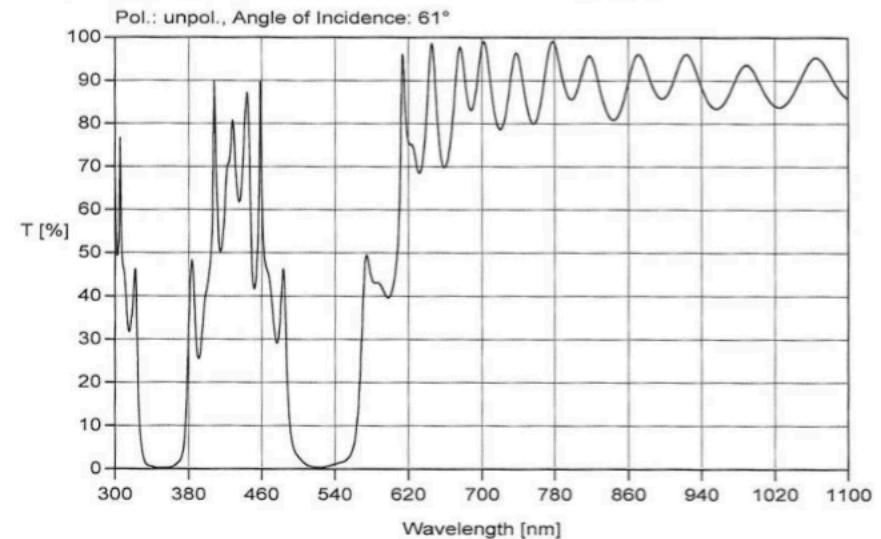
Hardware

- ✓ • Mirror characterized
- ✓ • Liquid light guide characterized
- ✓ • Guiding mirrors (dichroic)



Design of the guiding mirrors

Ru(355+532)>97%, Tu(1064)>80%, AOI=61° high power



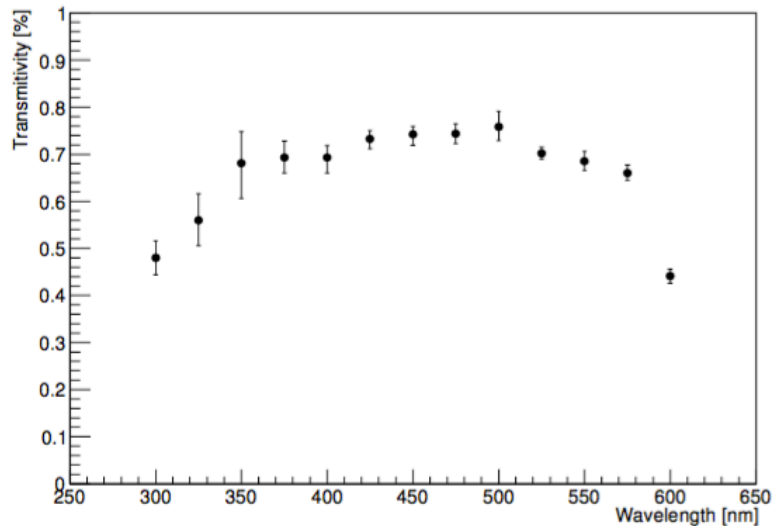
Transmission vs. wavelength

(see Bc thesis Eudald Font Pladevall)

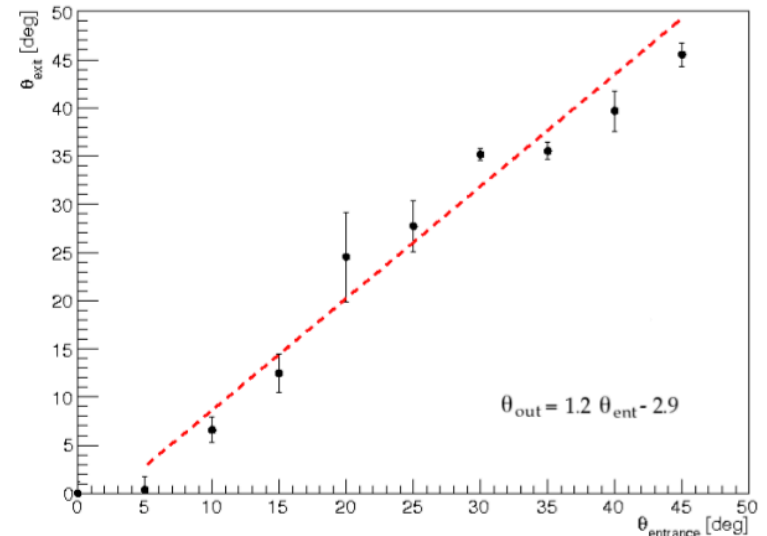
Status prototype

Hardware

- ✓ • Mirror characterized
- ✓ • Liquid light guide characterized



Transmission vs. wavelength



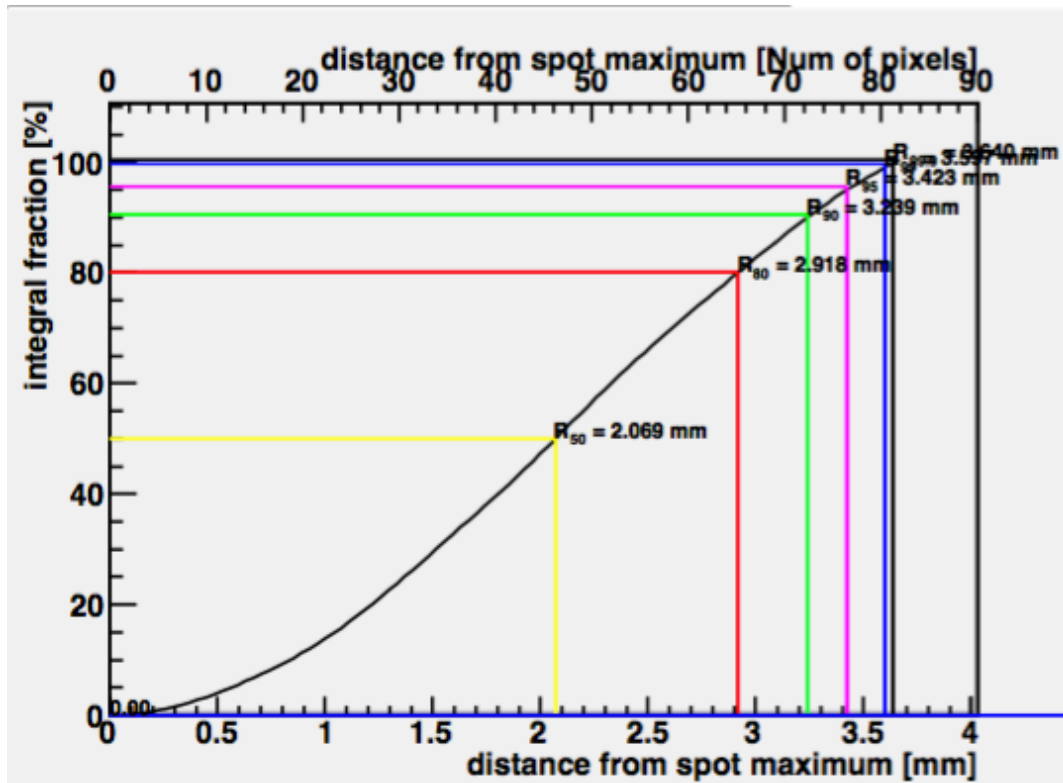
Output angle vs. input angle

(see PhD thesis Alicia López Oramas)

Status prototype

Hardware

- ✓ • Mirror characterized (using different methods)



Fraction of focused light falling into a circle with radius x from an artificial “star” located at 65 m from the telescope.

(Reflectivity was 64% at 350 nm at that time).

(see PhD thesis Alicia López Oramas)