

Producing realistic atmospheres for radio simulations of EAS using

GDAS



AtmoHEAD2018

Pragati Mitra, A. Bonardi, S. Buitink, A. Corstanje, H. Falcke, B.M. Hare, J.R. Hörandel, K. Mulrey, A. Nelles, J.P. Rachen, L. Rossetto, P. Schellart, O. Scholten, S. ter Veen, S. Thoudam, T. Huege, T.N.G. Trinh, T. Winchen



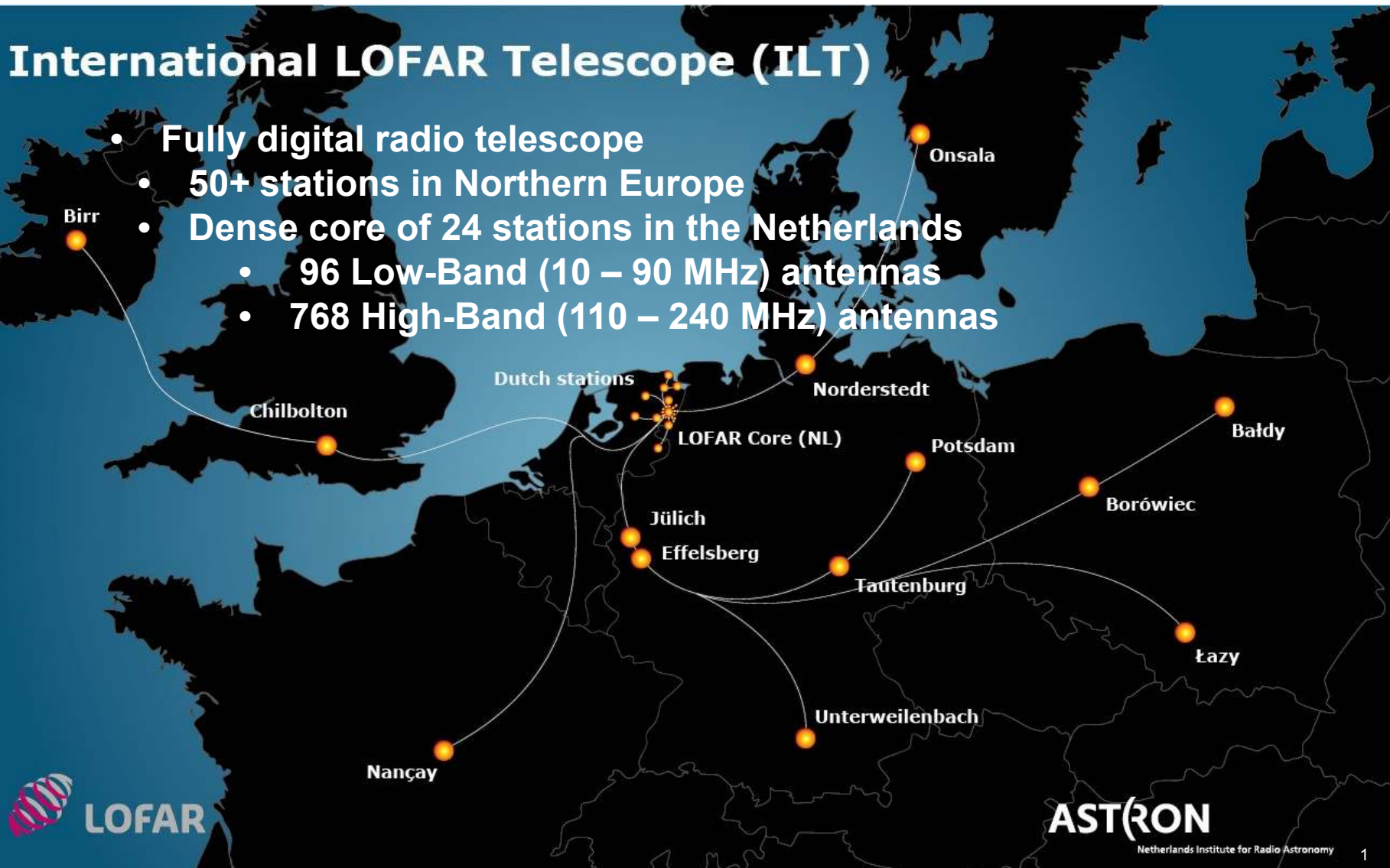
European Research Council
Excellence in European Research



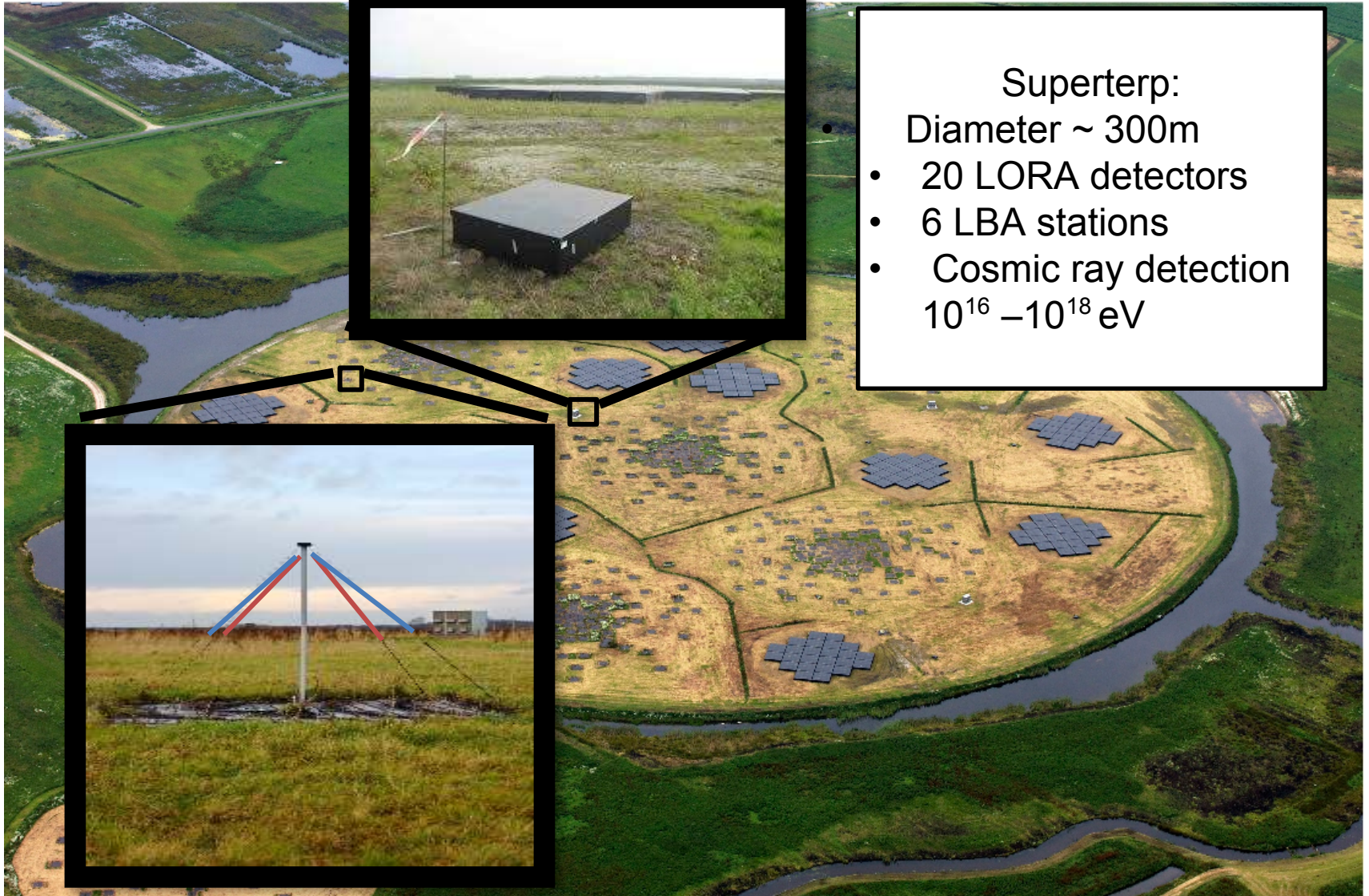
LOW Frequency ARay

International LOFAR Telescope (ILT)

- Fully digital radio telescope
- 50+ stations in Northern Europe
- Dense core of 24 stations in the Netherlands
 - 96 Low-Band (10 – 90 MHz) antennas
 - 768 High-Band (110 – 240 MHz) antennas



LOFAR Superterp

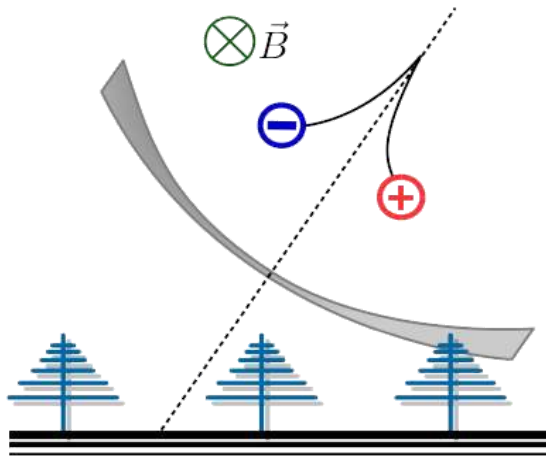


Superterp:
Diameter ~ 300m

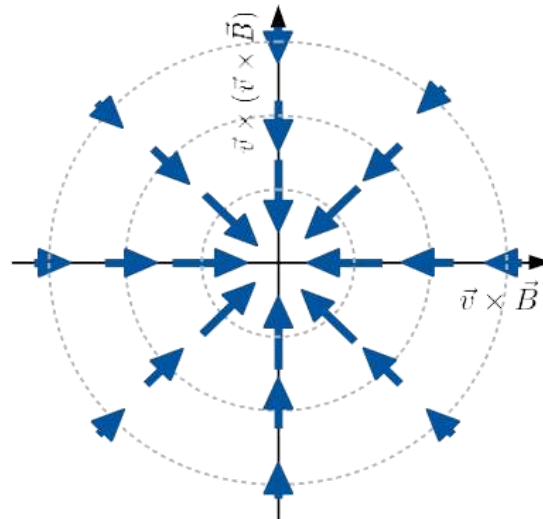
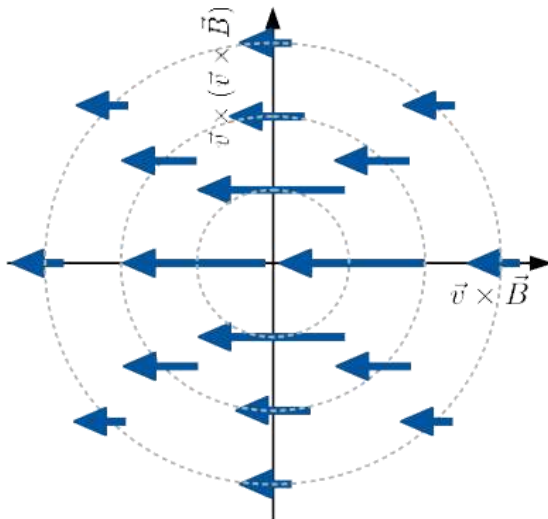
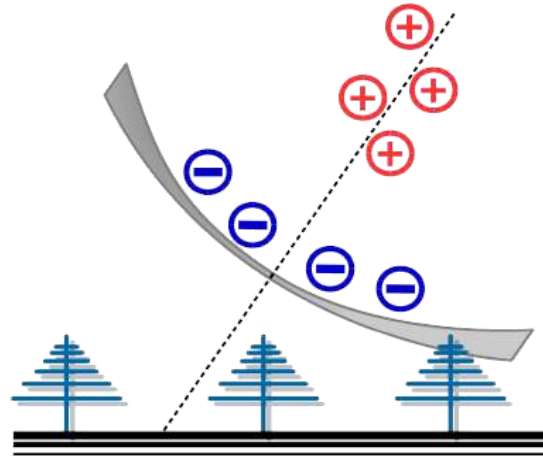
- 20 LORA detectors
- 6 LBA stations
- Cosmic ray detection
 $10^{16} - 10^{18}$ eV

Radio emission from air showers

Geomagnetic



Charge excess

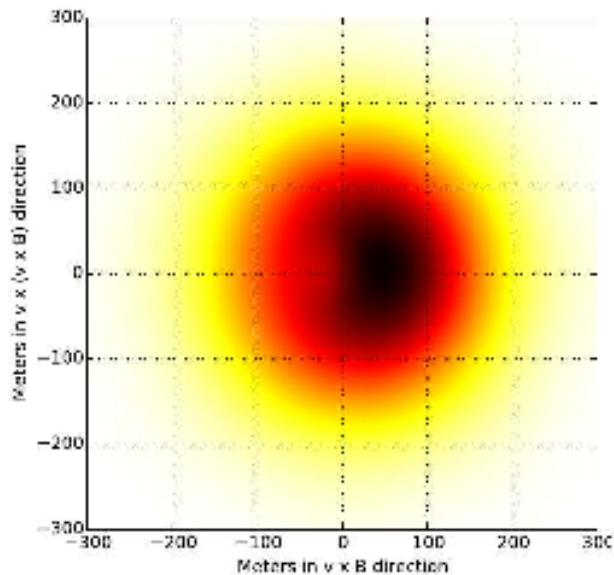


Combined effect results in asymmetry in radio profile

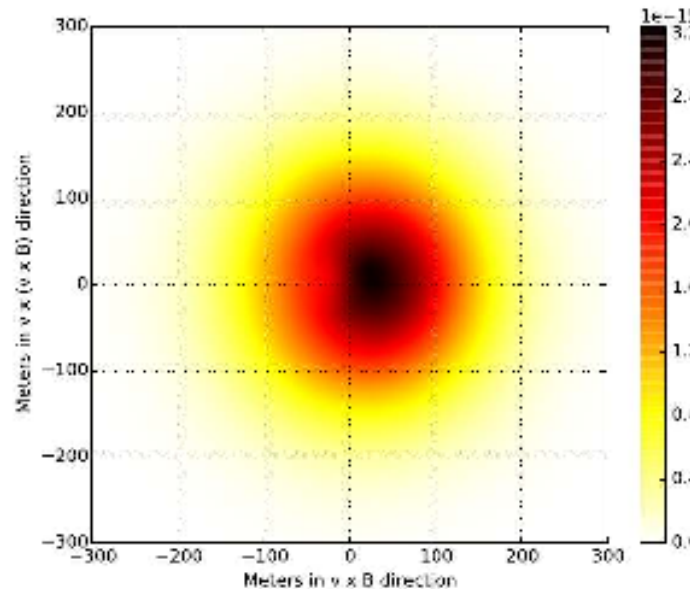
Radio footprint

Size depends on X_{\max} : deeper shower results in smaller footprint

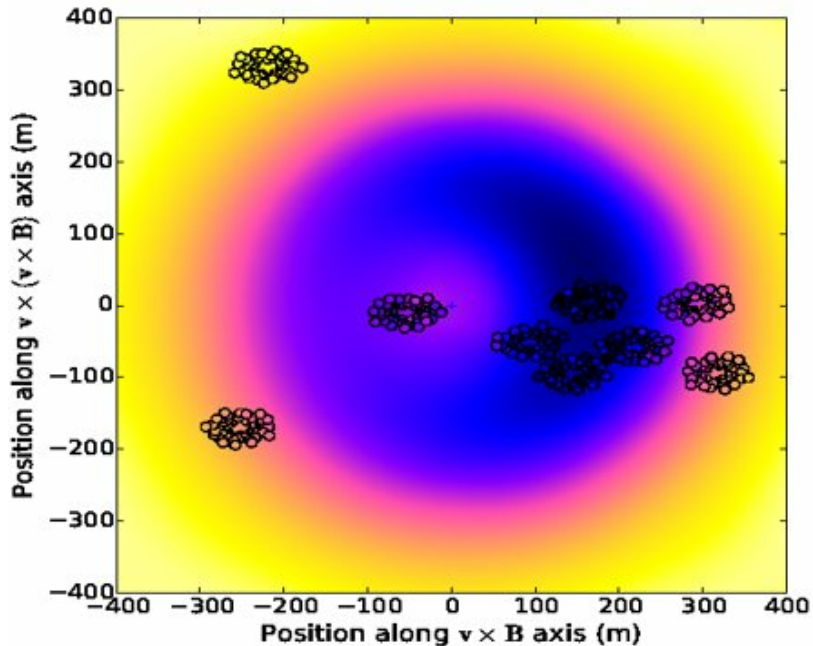
$X_{\max} = 630 \text{ g/cm}^2$



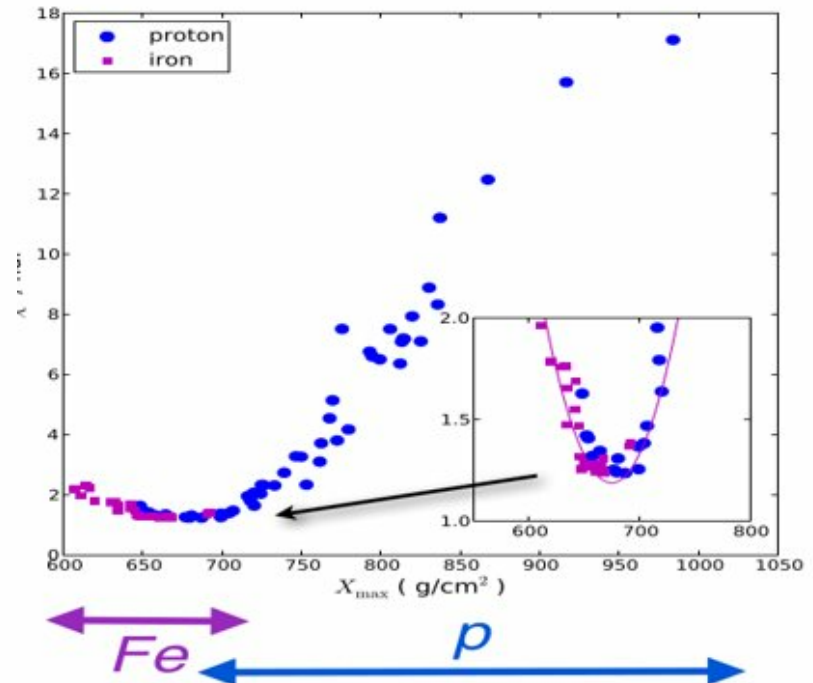
$X_{\max} = 700 \text{ g/cm}^2$



Xmax reconstruction

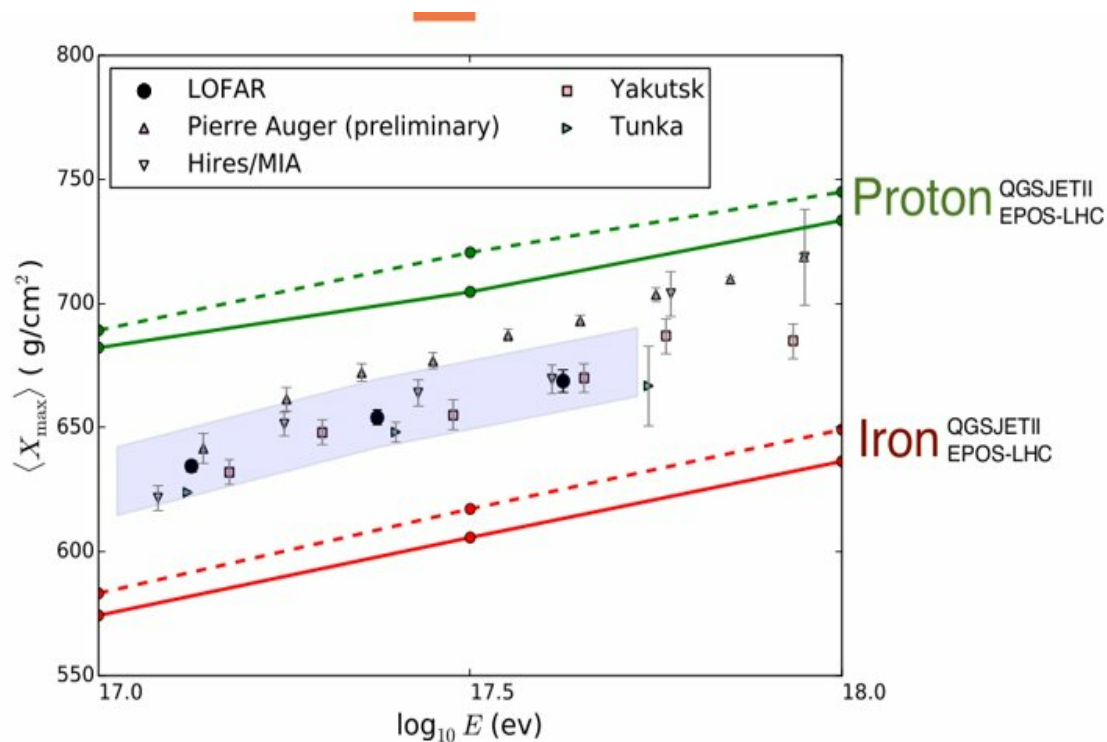


Measured antenna power fitted to a 2D simulated radio map



- Dedicated simulation set for each event.
- Minimum of fitted parabola → reconstructed Xmax

Systematics on Xmax



S.B et al Nature 513,70,2016

- Zenith angle dependence
~ 16 g/cm²
- Hadronic interaction
model ~ 5 g/cm²
- **Atmospheric effects**

Effect of atmosphere on Xmax

- Monte Carlo simulation codes for air shower like CORSIKA/ CoREAS uses default US stdA or averaged profiles.
- The variation of refractive index in the atmosphere important for radio detection.
- Refractivity $N=(n-1) \cdot 10^6$ depends on pressure, temperature and relative humidity in radio frequencies -

$$N = 77.6890 \frac{p_d}{T} + 71.2952 \frac{p_w}{T} + 375463 \frac{p_w}{T^2}$$

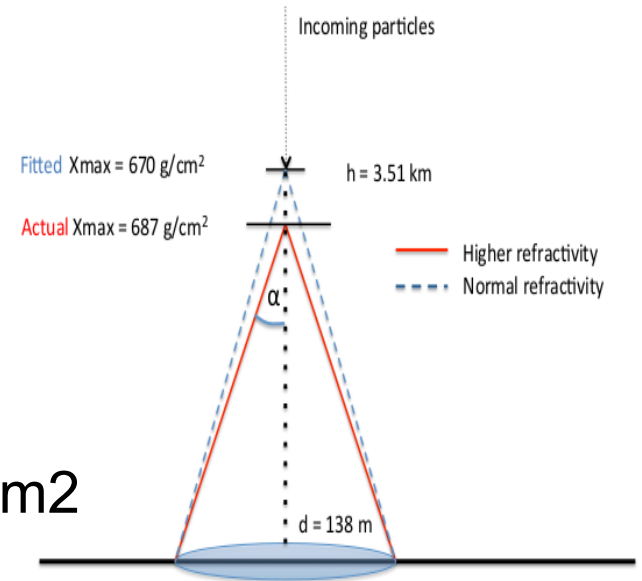
[after J.Rueger]

- Effect of humidity in N important in radio regime

Effect of refractive index on Radio Footprint

Toy Model

- Radio pulses compressed in time ~ Cherenkov like emission.
- Opening angle scales : **$\text{Cos}\alpha = 1/\beta n$**
- 10% higher N underestimates X_{max} 17 g/cm²

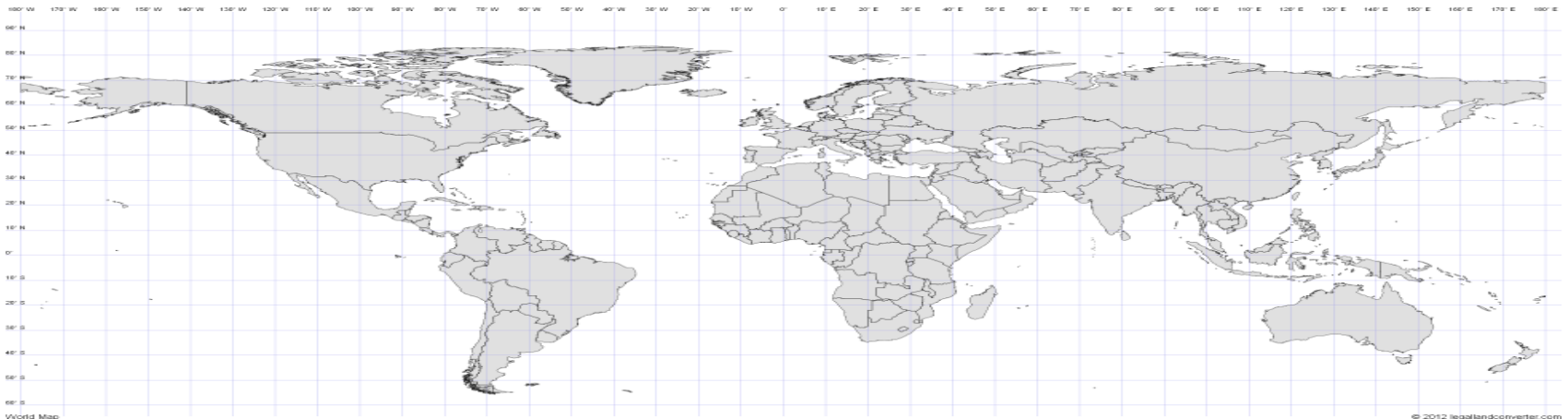


A. Corstanje et al
arXiv:1701.07338v1

Limitations: considers only local RI.
Integrated RI important for propagation effects.

Towards realistic profile: GDAS

- Global Data Assimilation System:
database of atmospheric data used for weather forecasting.
- $1^\circ \times 1^\circ$, 3 hour grid.
- 23 constant pressure level data.
- Altitude profiles of temperature, pressure, humidity.
- Calculated the density and refractive index for use in the simulations.

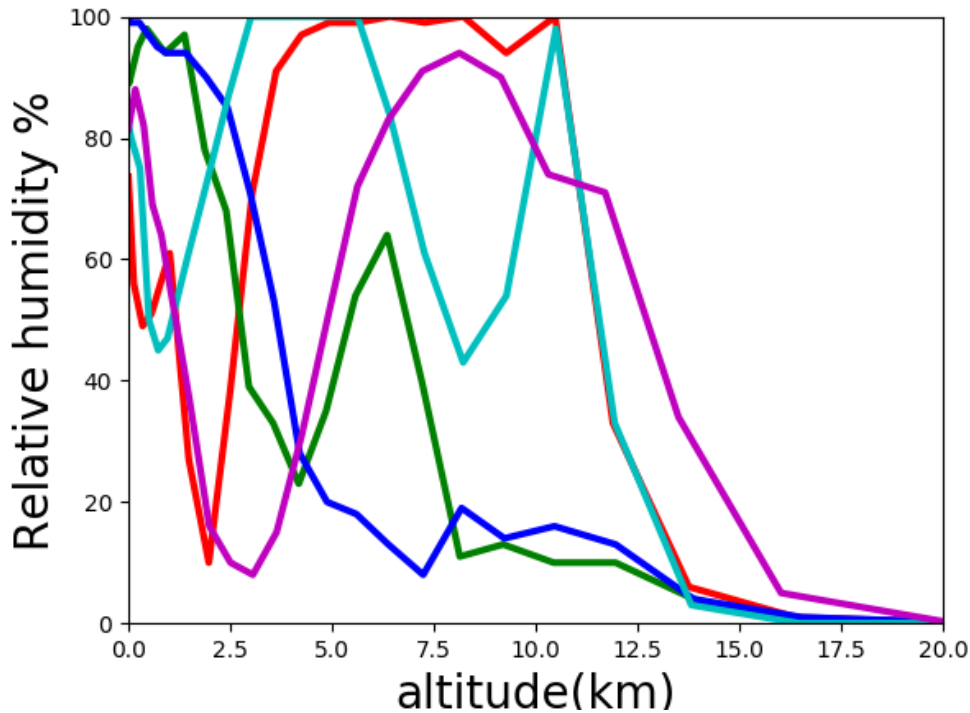


Towards realistic profile: GDAS

- Global Data Assimilation System:
database of atmospheric data used for weather forecasting.
- $1^\circ \times 1^\circ$, 3 hour grid.
- 23 constant pressure level data.
- Altitude profiles of temperature, pressure, humidity.
- Calculated the density and refractive index for use in the simulations.

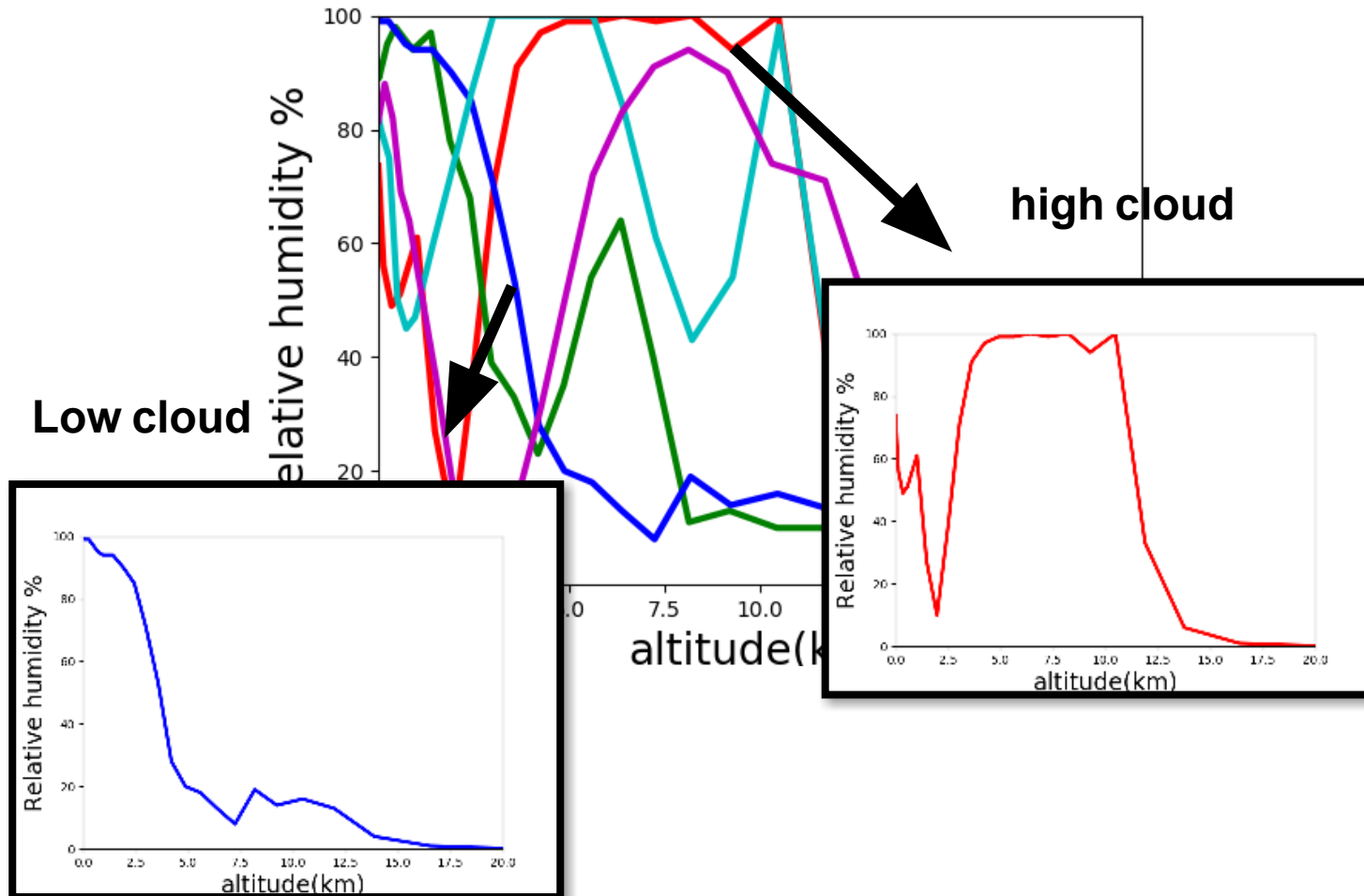


GDAS Atmospheric Profiles



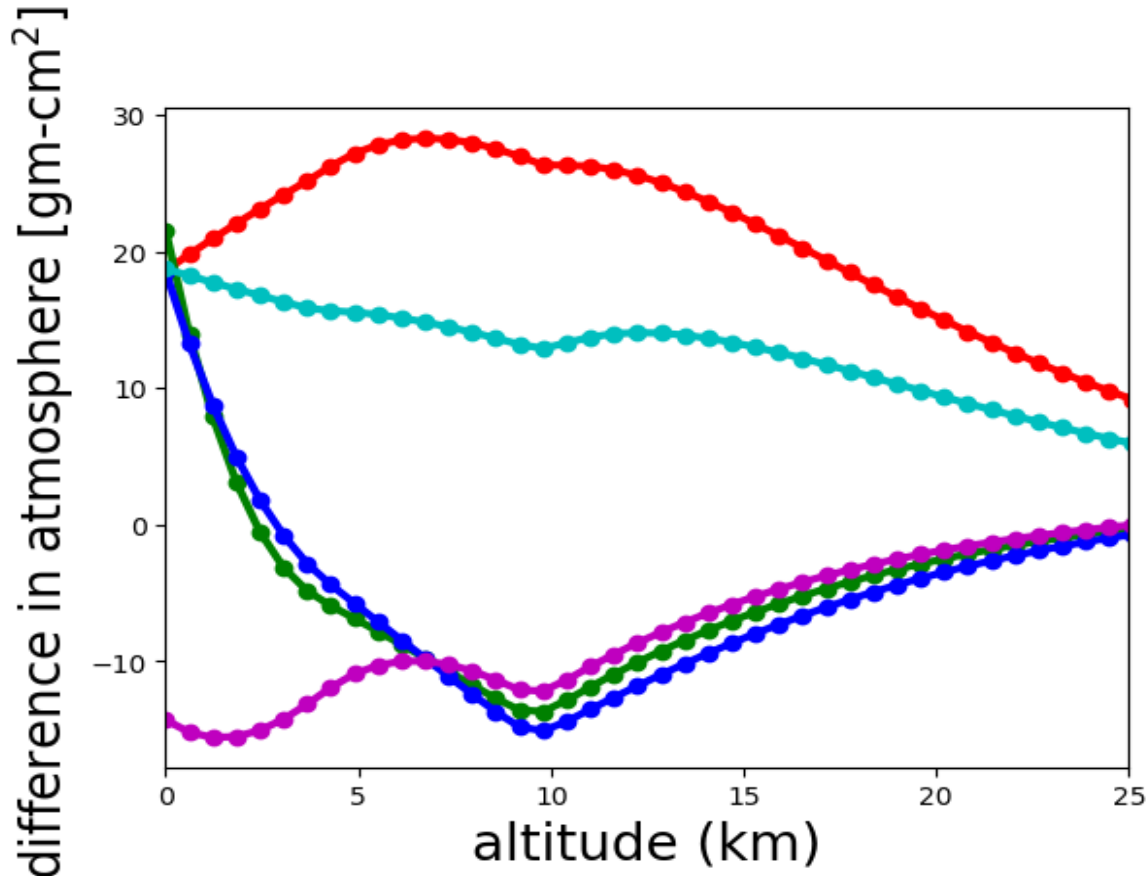
- Profile for 5 events.
- Huge variance at LOFAR site.

GDAS Atmospheric Profiles



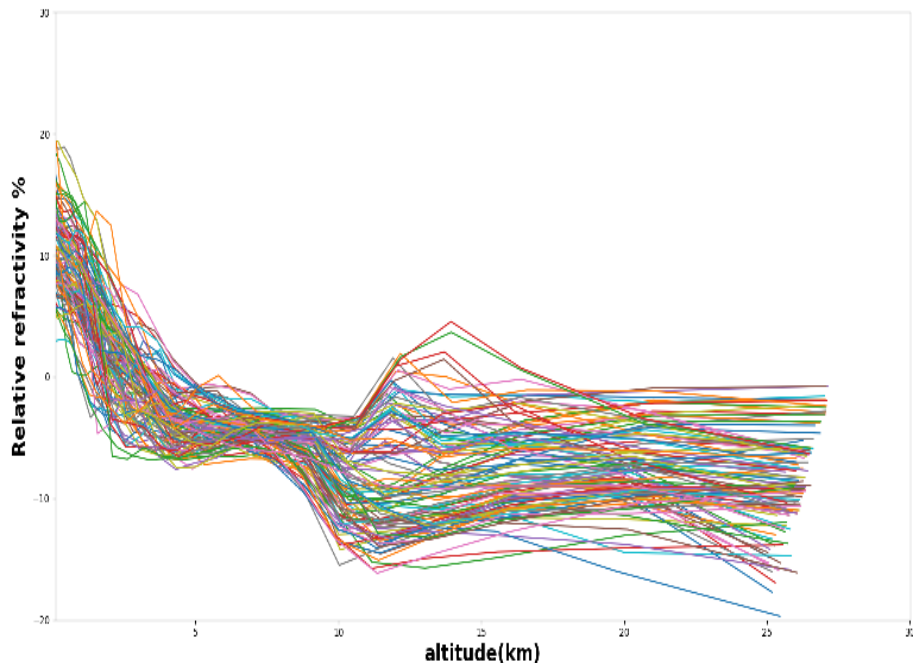
GDAS Profile: Atmosphere

GDAS- Us stdA atmosphere.



- Events with same ground pressure could see different atmosphere over height.
- Full atmospheric description over correction to ground pressure.

GDAS Profile: Refractivity

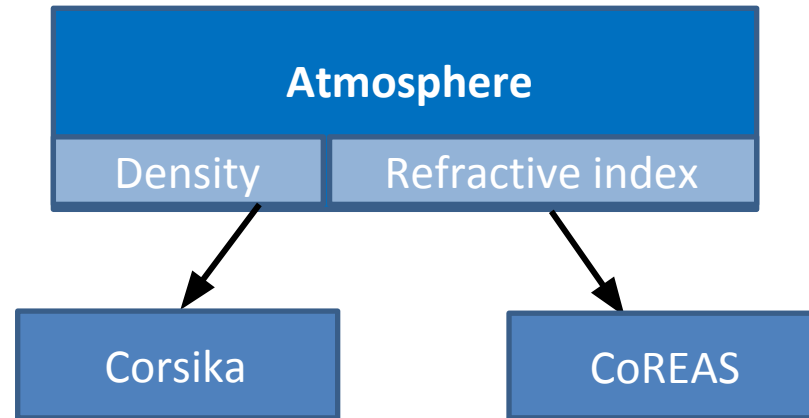


Relative refractivity:

$$(N_{\text{gdas}} - N_{\text{us}}) / N_{\text{us}}$$

- profile for 100 different CR events recorded at LOFAR.
- 3-5 % variation at 5-8 km ~ region of shower maximum.
- Integrated relative refractive index ~ 7-10 % between the same region.

Implementation of GDAS to CORSIKA/CoREAS



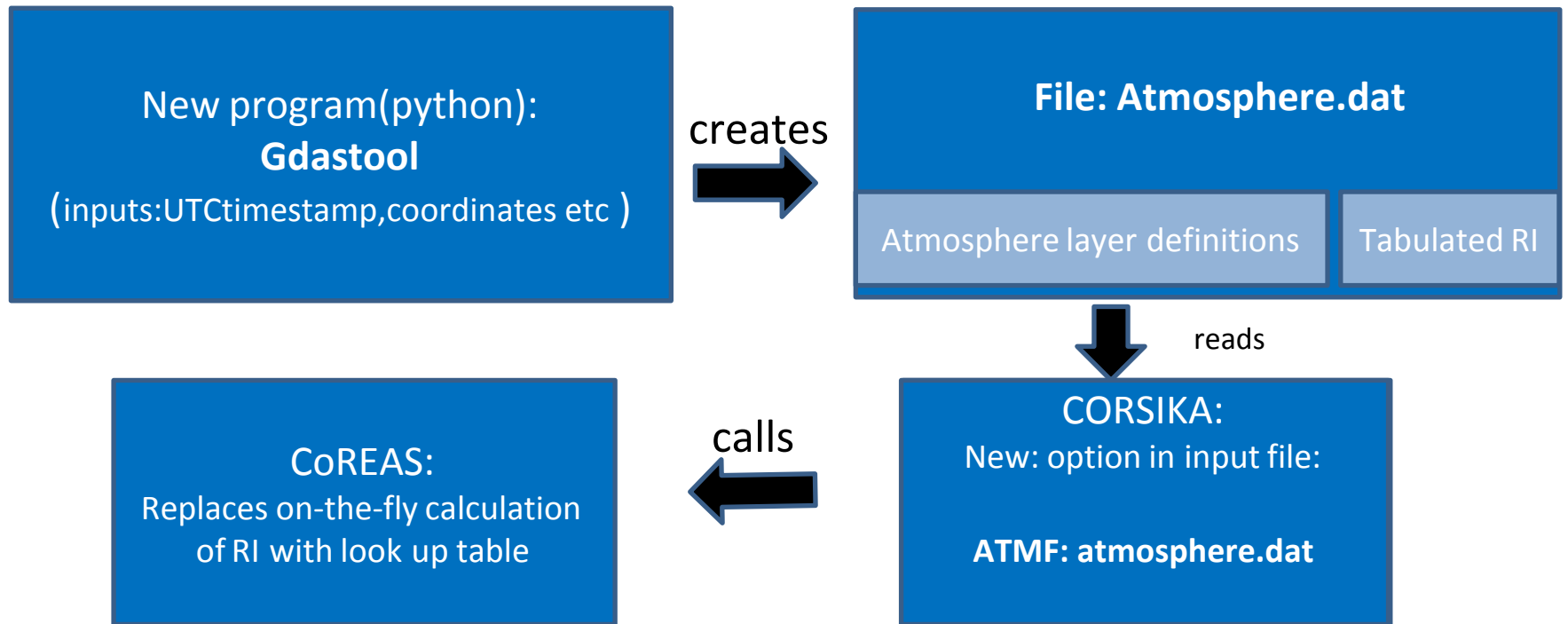
CORSIKA

- Fit layered atmosphere model to GDAS data.
- Feed fit parameters unto simulation.

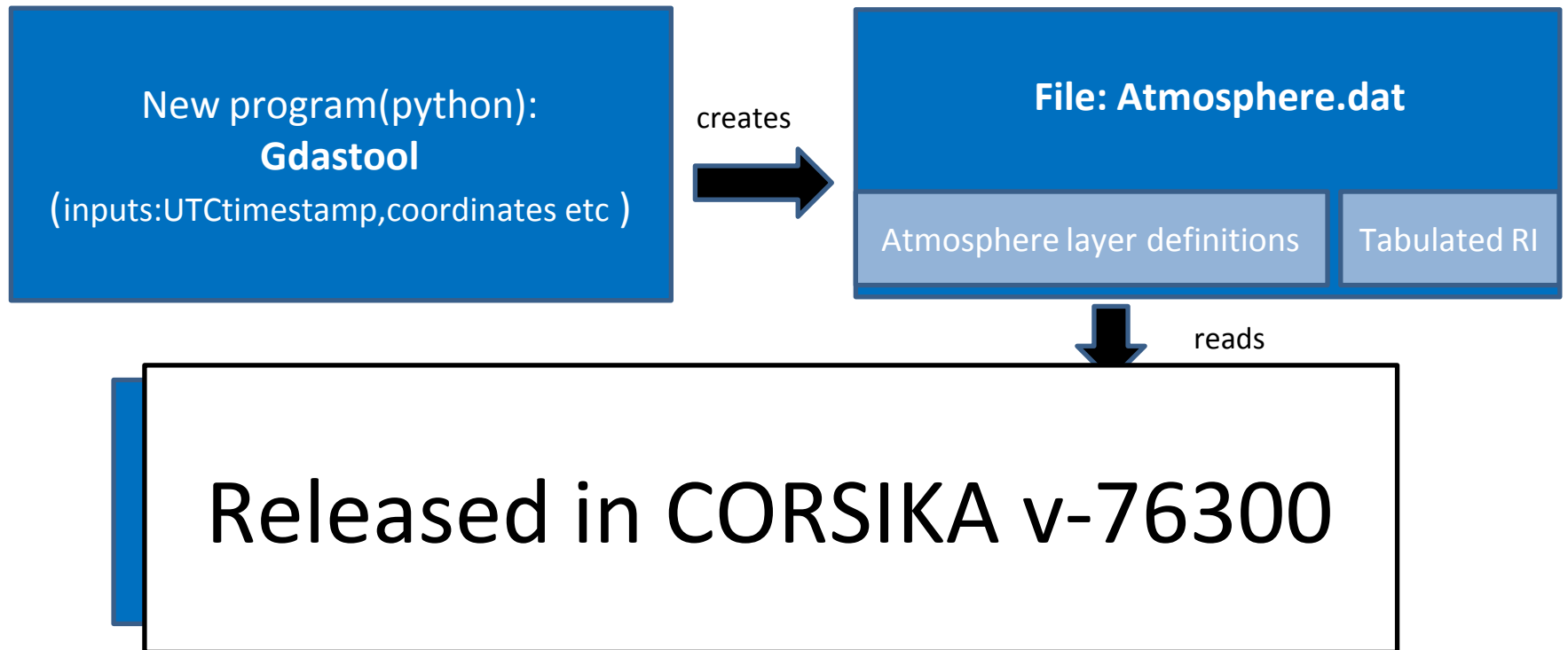
Coreas

- Replace calculation of RI with look up table.
- Read RI data from file.

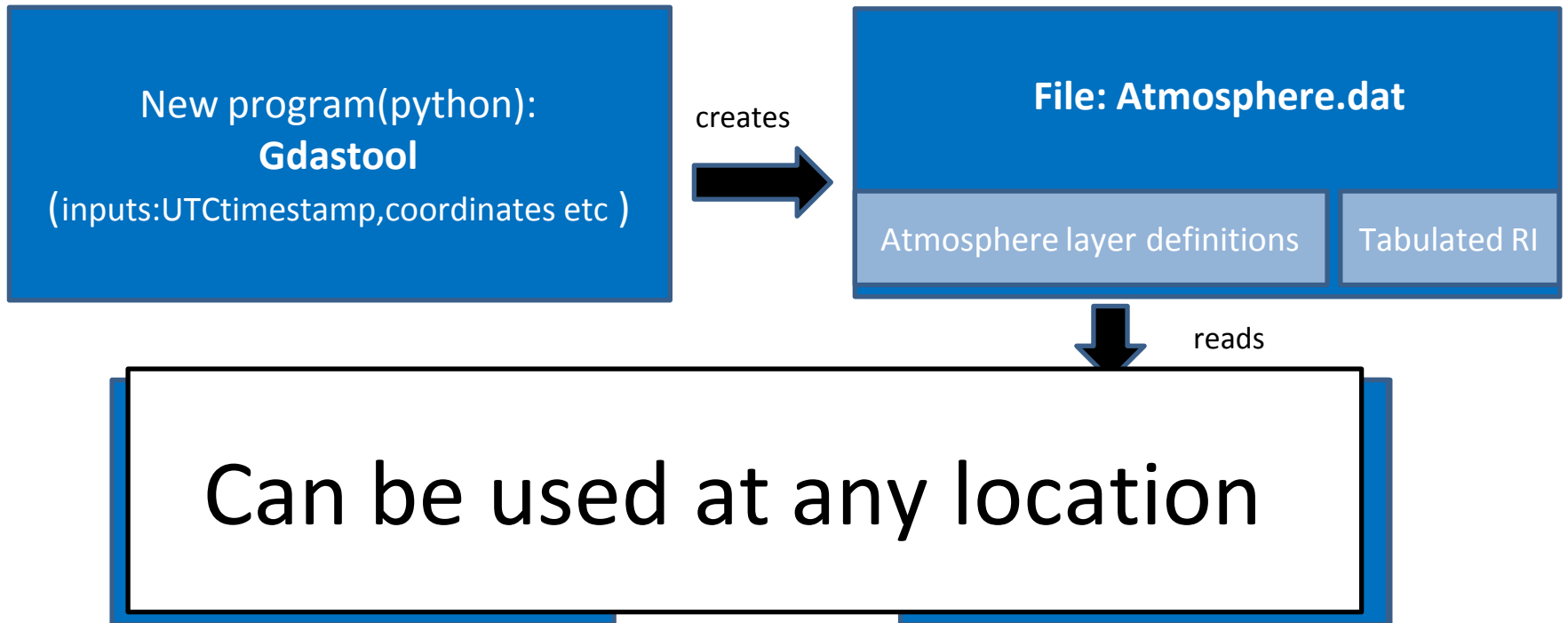
Implementation of GDAS to CORSIKA/CoREAS



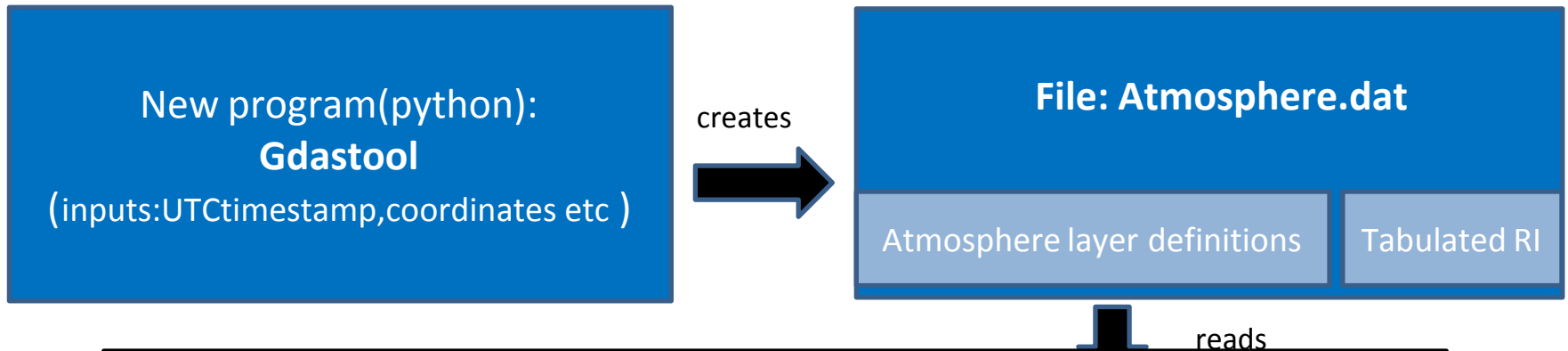
Implementation of GDAS to CORSIKA/CoREAS



Implementation of GDAS to CORSIKA/CoREAS



Implementation of GDAS to CORSIKA/CoREAS



Execution: `gdastool [-h] [-t UTCTIMESTAMP]
[-o OUTPUT]
(--observatory {lofar,aera} | -c COORDINATES
COORDINATES)`

Fitting Procedure

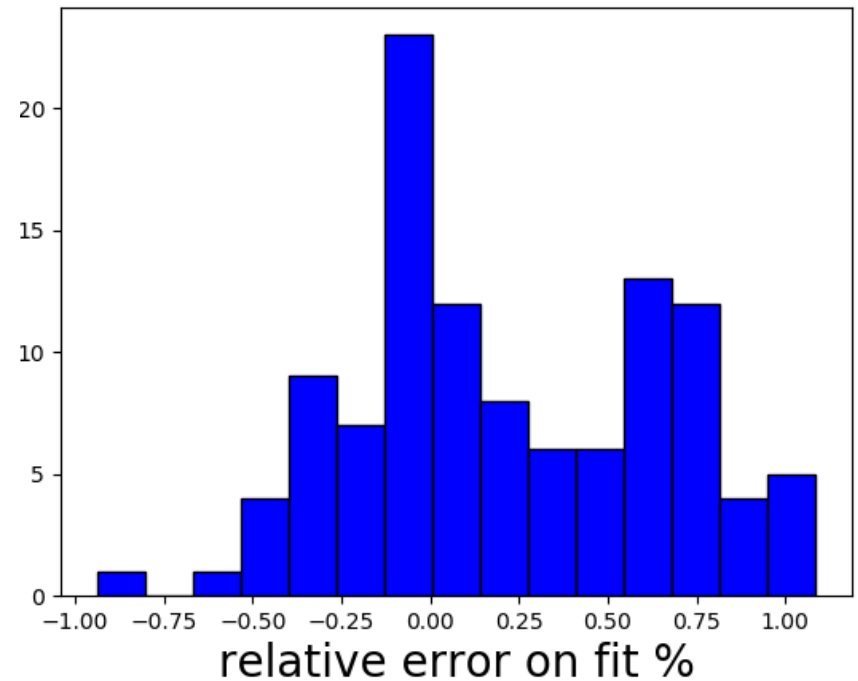
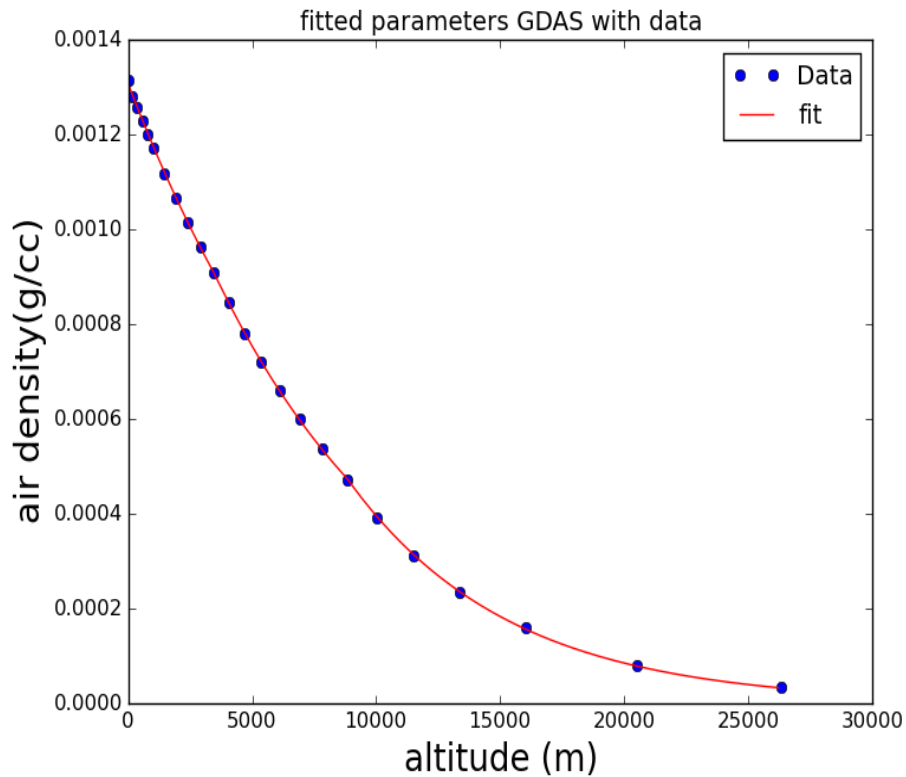
- 5 layer atmospheric model

$$X(h) = a + b \exp(-h/c)$$

$$\rho(h) = b/c \exp(-h/c)$$

- Boundary condition: $X(h), \rho(h)$ should be continuous at layer boundary
- Choosing boundary layers-24 GDAS points divided into 3 layers
- Fit density profile for b, c
- Analytically solve for a

gdastool : Output Example



Relative error: Δ density/
density data

Simulation with LOFAR data

- New simulation with LOFAR events.
- 123 for this analysis
- New GDAS atmosphere .
- Xmax reconstruction with radio only fitting
- Compare between event sets:

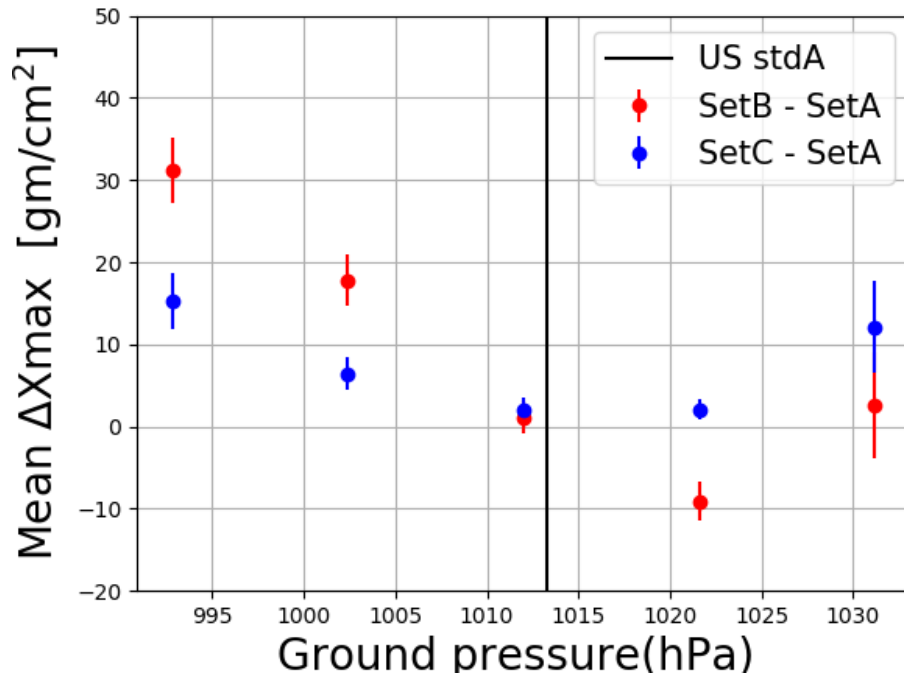
Set- A
GDAS atm + CORSIKA-
new

Set- B
US stdA + CORSIKA-old

Set-C
US stdA +atm
correction+ CORSIKA-
old

CORSIKA-old: v 74385
CORSIKA-new: v 76300
Offset = g/cm2

Mean Xmax vs ground pressure



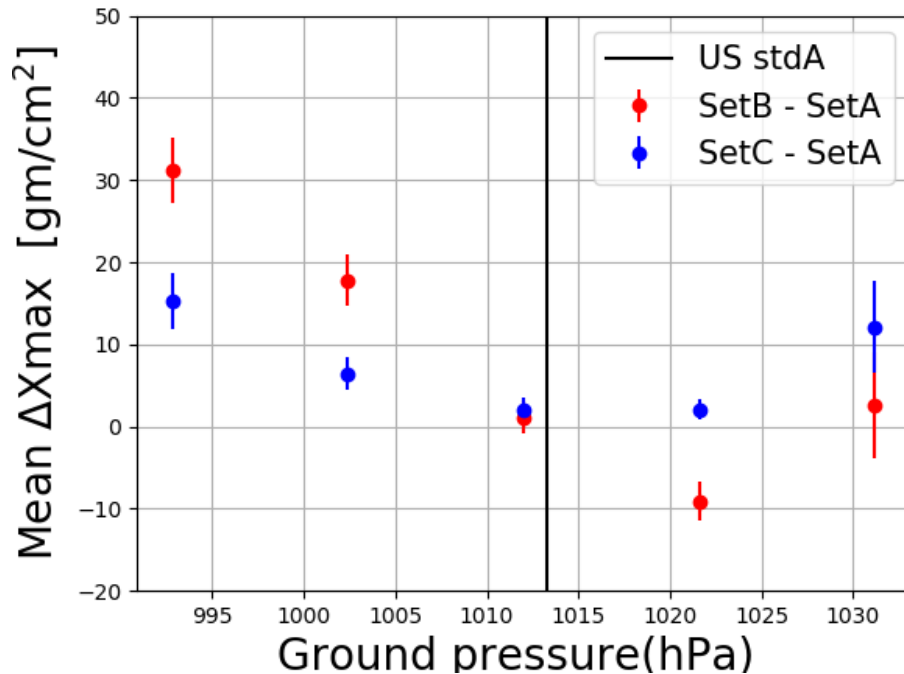
Difference in mean Xmax between

Set- A
GDAS atm +
CORSIKA-new

Set- B
US stdA +
CORSIKA-old

Set-C
US stdA +atm
correction+
CORSIKA-old

Mean Xmax vs ground pressure



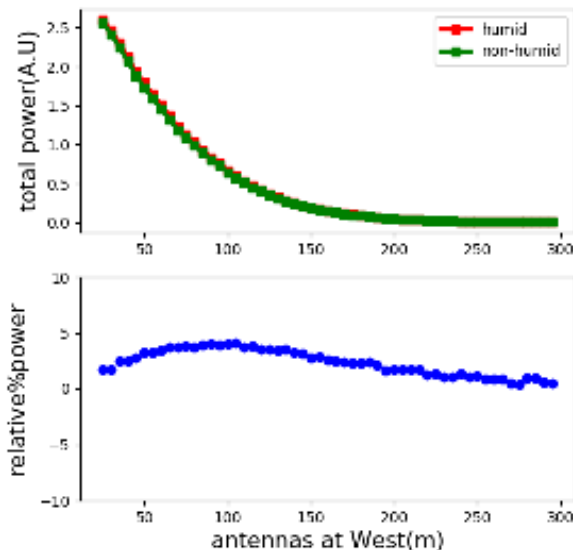
- Correction for pressure is important.
- Linear correction is not sufficient at lower pressure.
- full GDAS-based simulations required.

Effects of humidity on radio LDF

- 2 simulation sets: GDAS **N** profile with and without humidity
- Density profile same \longrightarrow same X_{\max}

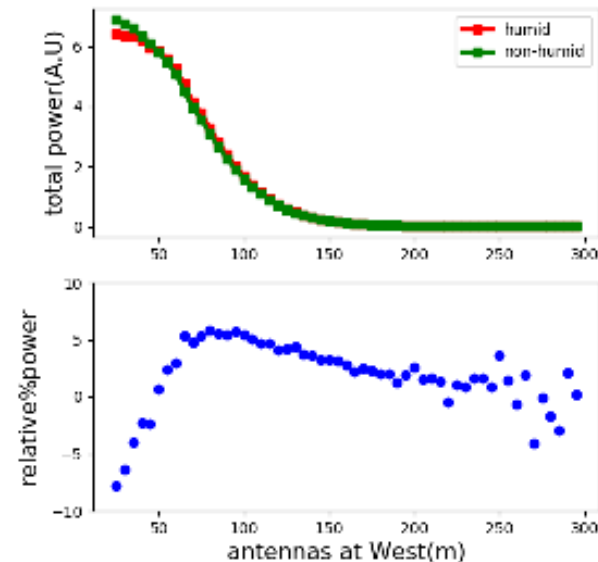
10^{17} eV proton shower, zenith 0, $X_{\max}=604$ g/cm²

30-80 MHz (LOFAR)



- Too small effect to be visible in data

50-350MHz (SKA-low)



- Cherenkov like emissions more prominent sensitive to **N**
- Higher effects $\sim 10\%$

Conclusions

- GDAS atmosphere included in **CORSIKA-v76300**
- A Stand-alone script "**gdastool**" provides local real time atmospheric profiles (density + RI) for any location.
- LOFAR data simulated with GDAS atmosphere.
- Full GDAS atmospheric profile is required over Linear correction to US stdA.

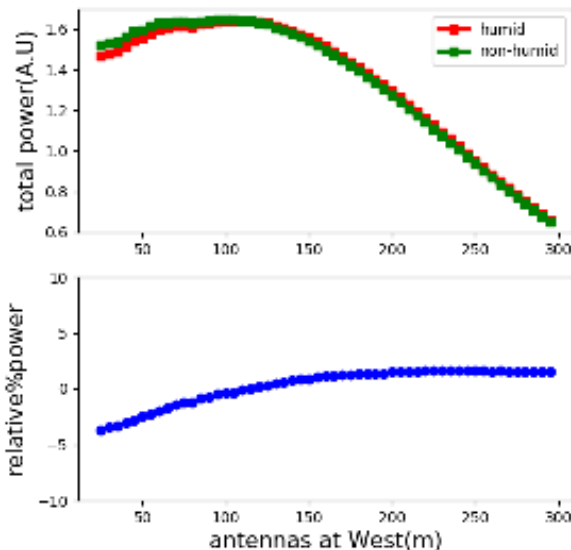
Back ups

Effects of humidity on radio LDF

- 2 simulation sets: GDAS **N** profile with and without humidity
- Density profile same \longrightarrow same X_{\max}

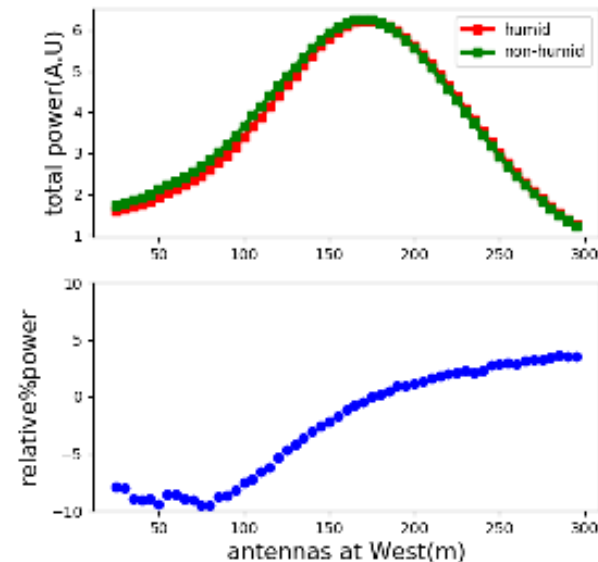
10^{17} eV proton shower, zenith 45, $X_{\max}=593$ g/cm²

30-80 MHz (LOFAR)



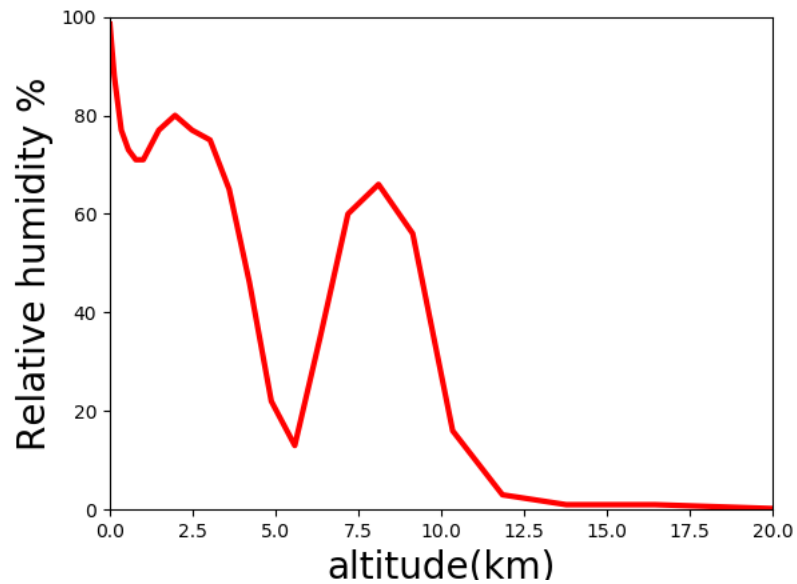
- Too small effect to be visible in data

50-350 MHz (SKA-low)

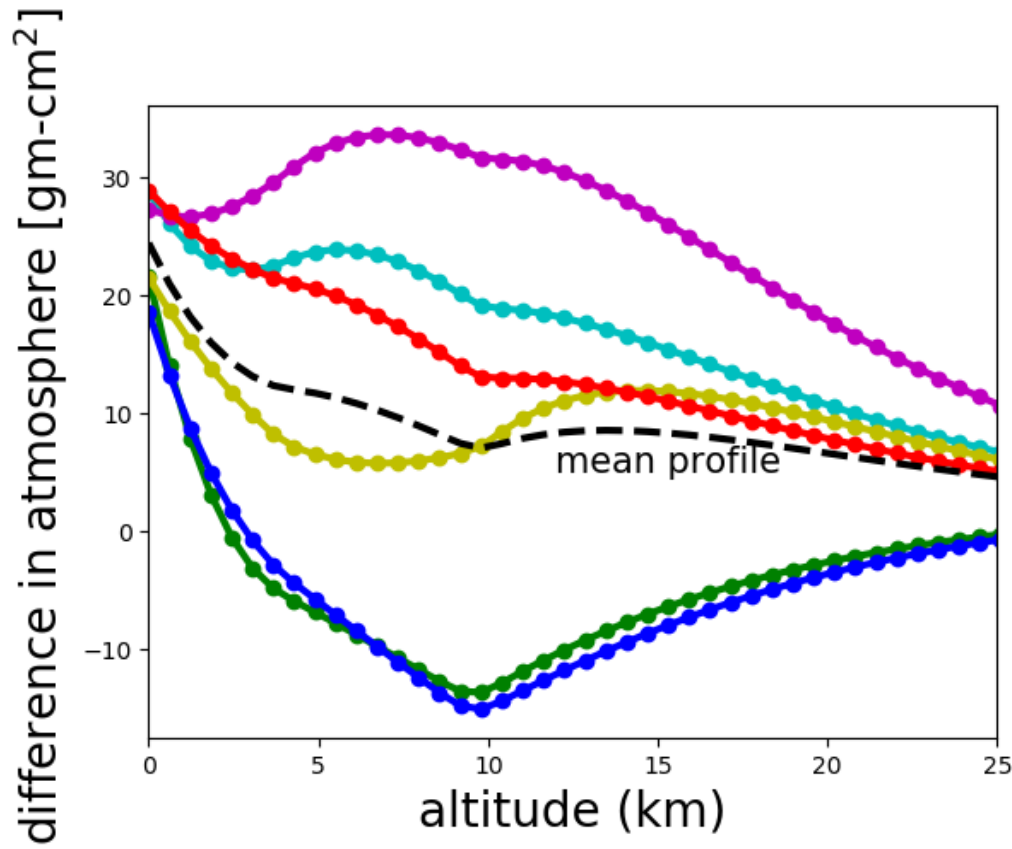


- Cherenkov like emissions more prominent sensitive to **N**
- Higher effects $\sim 10\%$

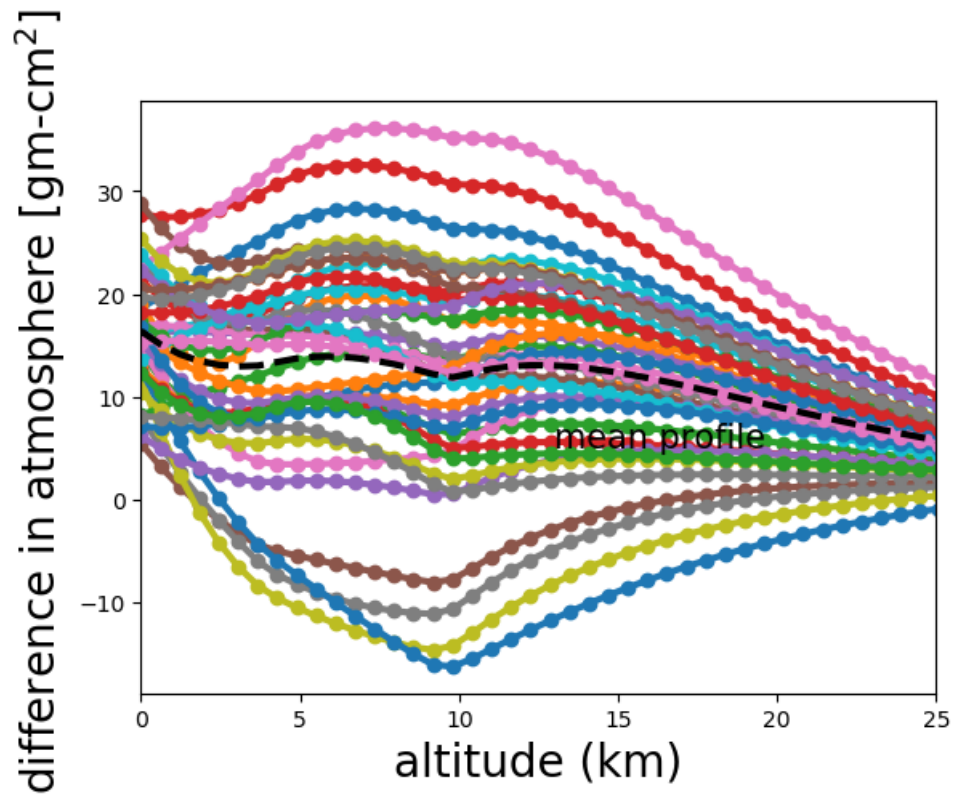
Humidity profile:event



Atmosphere profile: last Pressure bin(high P)



Atmosphere profile: 2nd last Pressure bin



Bias Test

Set	Mean Xmax Humidity=low	Mean Xmax Humidity=high
New	661.8	659.9
Old corrected old	666.3	663.8
	663.4	665.3

Set	Zenith=low	Zenith=high
New	655.6	666.6
Old corrected old	659.6	670.5
	659.5	669.2

Humidity Effects: From MC studies

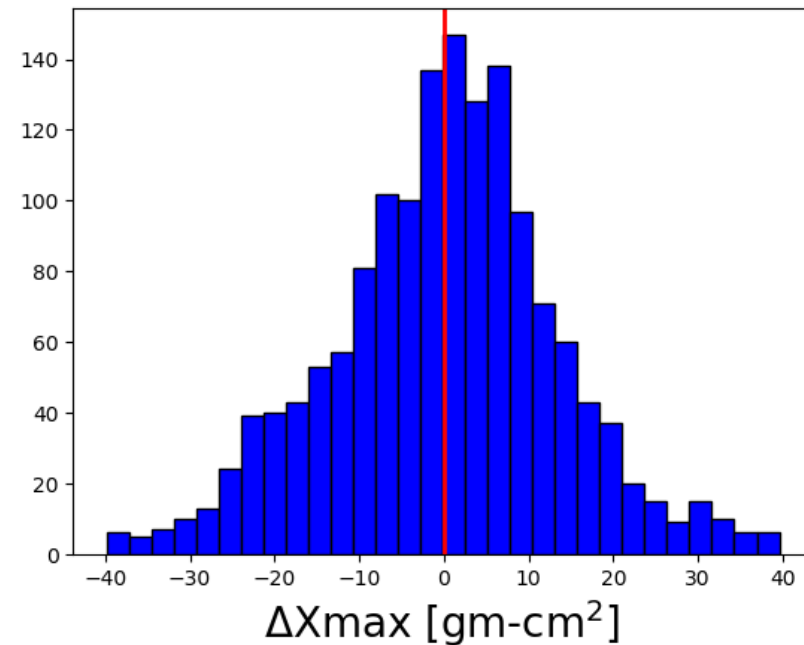
Set1

- New CORSIKA/CoREAS
- No humidity

Set2

- New CORSIKA/CoREAS
- With humidity

- Take one shower from **Set2** as "fake data"
- Fit showers from **Set1** to it.
- Difference $X_{\text{max}_{\text{reco}}} - X_{\text{max}_{\text{real}}}$: humidity effect.
- 50 event sets; each \sim 50-100 simulations.



Humidity Effects: From MC studies

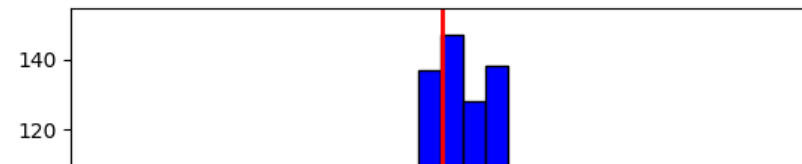
Set1

- New CORSIKA/CoREAS
- No humidity

Set2

- New CORSIKA/CoREAS
- With humidity

- Take one shower from **Set2** as "fake data"
- Fit showers from **Set1** to it.
- Difference $X_{\text{max}_{\text{reco}}} - X_{\text{max}_{\text{real}}}$: humidity effect.
- 50 event sets; each ~ 50 -100 simulations.



- no visible effect of humidity
- Further investigation

