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## Measurement of the $^{12}\text{C}(\text{p},\gamma)^{13}\text{N}$ S-factor in inverse kinematics

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The  $^{12}\text{C}(\text{p},\gamma)^{13}\text{N}$  reaction is relevant in several astrophysical scenarios, such as the early stages of the Bethe-Weizsäcker cycle of hydrogen burning and the production of  $^{13}\text{C}$  in stars on the asymptotic giant branch of the Hertzsprung-Russell diagram. Here new data on the  $^{12}\text{C}(\text{p},\gamma)^{13}\text{N}$  astrophysical S-factor at low energy, 0.1-0.5 MeV in the center of mass system are reported from an experiment in inverse kinematics. Titanium hydride targets were irradiated with an intensive  $^{12}\text{C}$  ion beam from the HZDR 3 MV Tandetron accelerator. The emitted  $\gamma$ -rays were detected in a lead shielded high-purity germanium detector also equipped with a cosmic ray veto. For target characterization, Nuclear Resonant Reaction Analysis (NRRA) was used with a 6.4 MeV  $^{15}\text{N}$  beam. The new data will contribute to the understanding of the creation of chemical elements in the precursors of core-collapse supernovae.

### Selected session

Nuclear Astrophysics

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