

# Measurement of the $^{12}\text{C}(p,\gamma)^{13}\text{N}$ S-factor in inverse kinematics

Stefan Reinicke<sup>1,2</sup>, Shavkat Akhmadaliev<sup>1</sup>, Daniel Bemmerer<sup>1</sup>,  
Marcel Grieger<sup>1,2</sup>, Felix Ludwig<sup>1,2</sup>, Stefan Schulz<sup>1,2</sup>, Ronald Schwengner<sup>1</sup>,  
Klaus Stöckel<sup>1,2</sup>, Marcell Takács<sup>1,2</sup>, Louis Wagner<sup>1,2</sup>, and Kai Zuber<sup>2</sup>

<sup>1</sup> Helmholtz-Zentrum Dresden-Rossendorf (HZDR)

<sup>2</sup> Technische Universität Dresden

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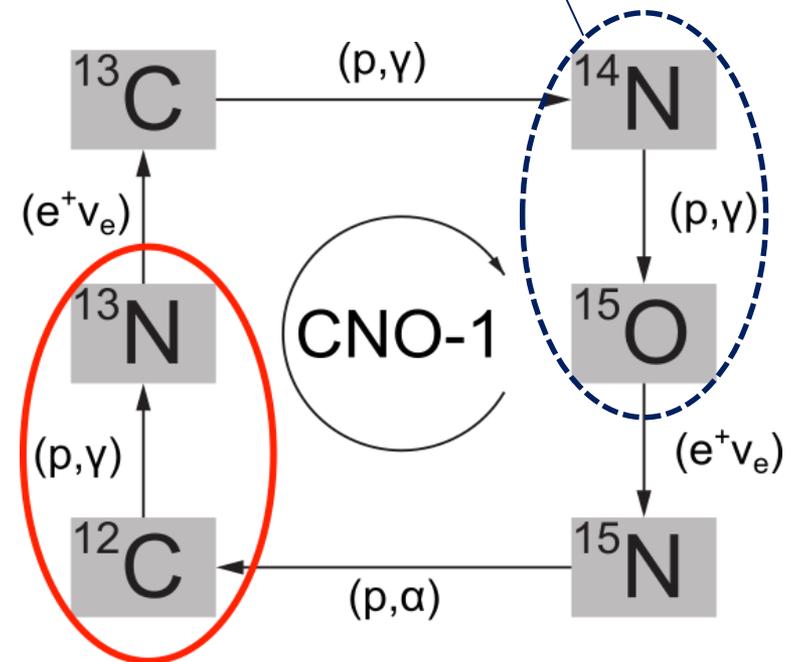
Bologna/Italy, 06.09.2018



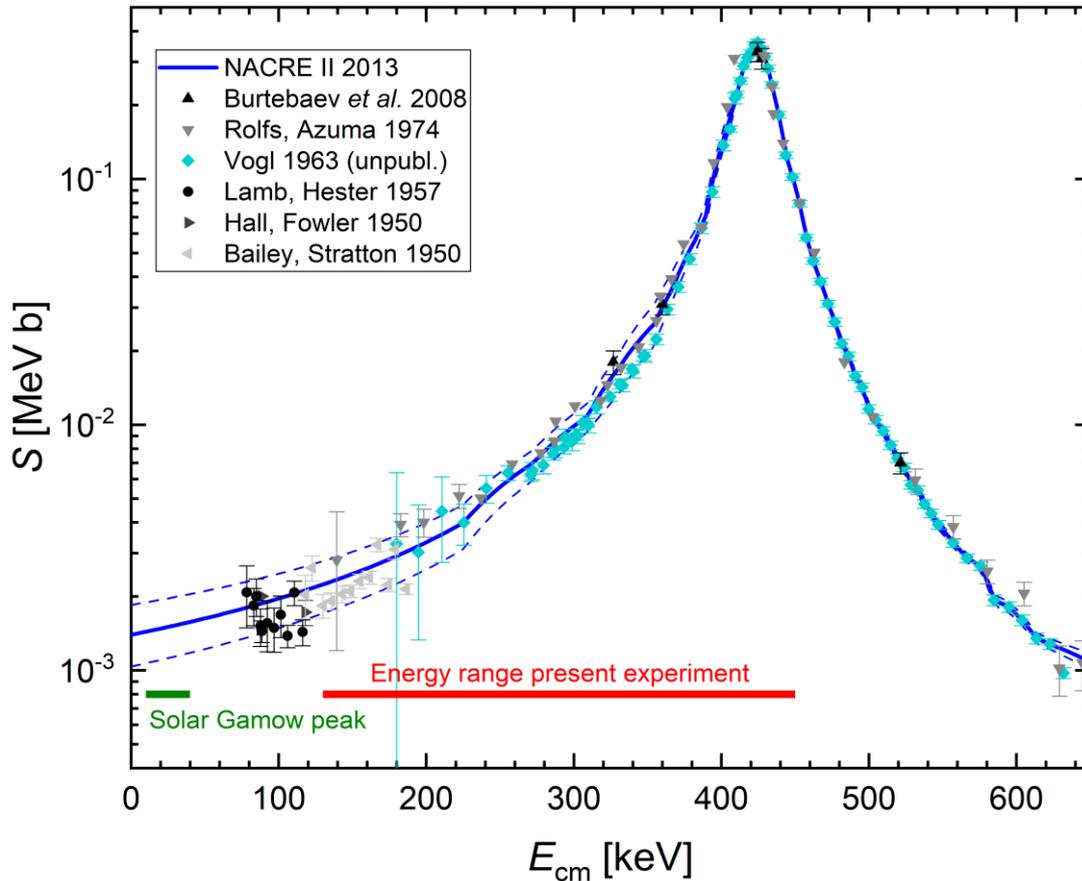
# Motivation

- Part of an effort to restudy the CNO cycle
- $^{12}\text{C}(p,\gamma)^{13}\text{N}$  dominates CNO rate in early development phase of stars and in outer parts of solar core
- Responsible for  $^{13}\text{C}$  production in red giant stars
- Radial emission profile of solar neutrinos from  $\beta^+$ -decay depends on  $^{12}\text{C}(p,\gamma)^{13}\text{N}$  rate

Wagner+, Phys. Rev. C 97, 015801 (2018)



# State of the art



## Aim of this work

- Measurement in wide energy range between 130 and 450 keV
- Improve data in low energy region
- Use of inverse kinematics (no previous data with this method)
  - Different systematics
  - Lower beam induced background
- Low natural background due to thick lead shielding and active muon veto

Astrophysical S factor:  $\sigma(E) = S(E) \frac{1}{E} e^{-2\pi\eta}$

# Setup

## Accelerator and ion beam

- Rossendorf 3 MV Tandetron accelerator
- $^{12}\text{C}^{2+}$  beam on target: 5 – 20  $\mu\text{A}$



## Targets

- $\text{TiH}_2$  targets with Ta backing
  - 3 hydrated, 200 nm  $\text{TiH}_2$
  - 2 hydrated, 100 nm  $\text{TiH}_2$
  - 1 H-implanted, 100 nm  $\text{TiH}_2$
- Mounted with  $55^\circ$  angle towards beam axis



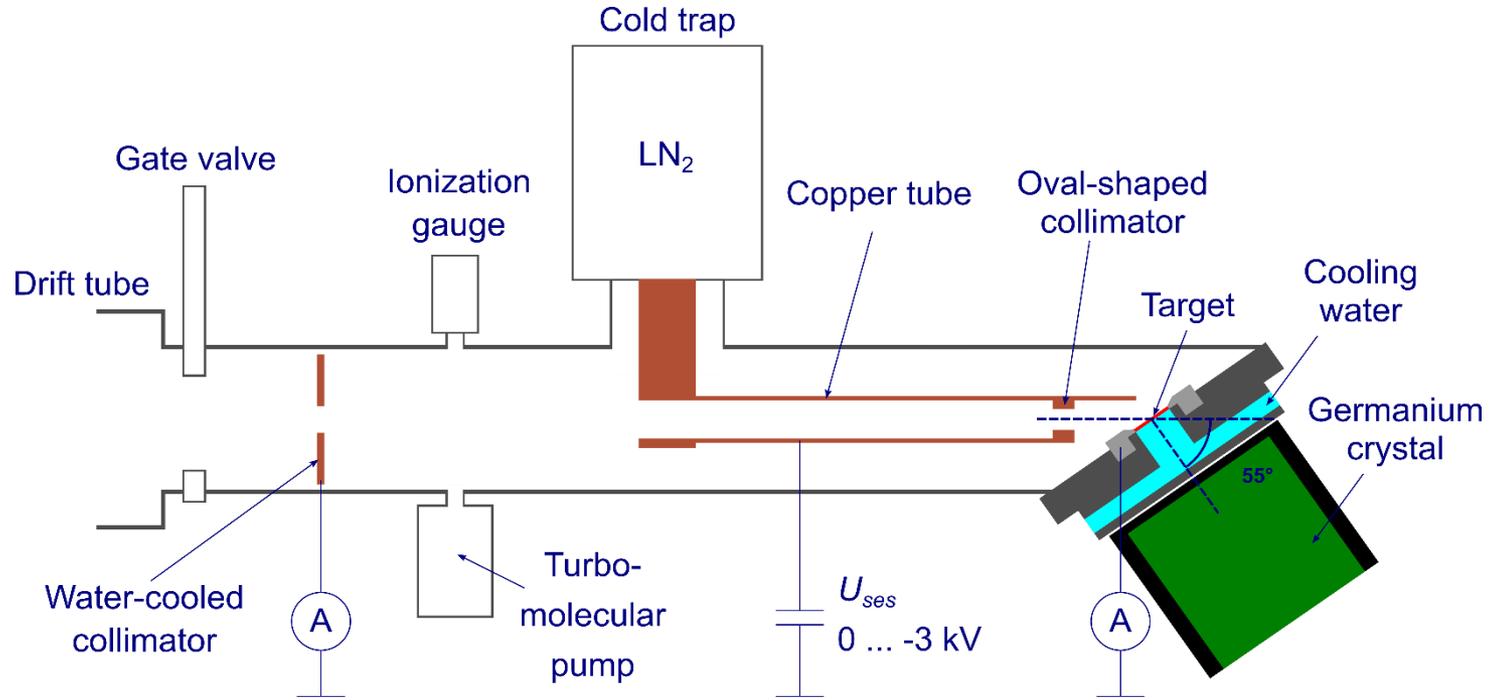
## Detector and shielding

- HPGe detector with 90% relative efficiency at  $55^\circ$
- 12 cm thick lead castle
- 5 cm thick scintillators for active muon veto



# Setup (continued)

## Target chamber and detector

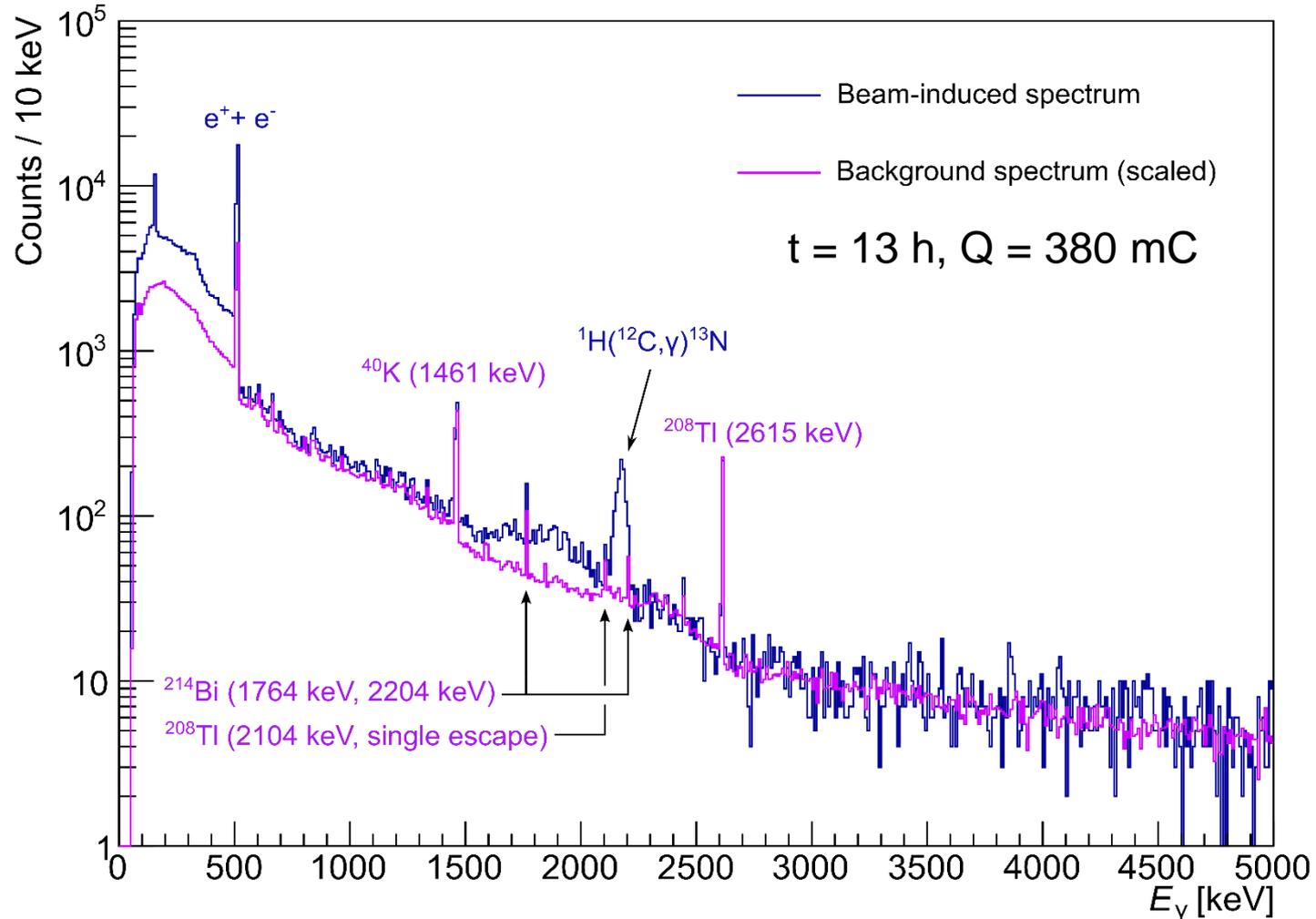


## Data acquisition

- Analog chain
  - ORTEC 671, ORTEC 919E (MAESTRO, histogramming)
- Digital chain
  - CAEN N1728b 100 Ms/s digitizer (TNT2, list mode data)

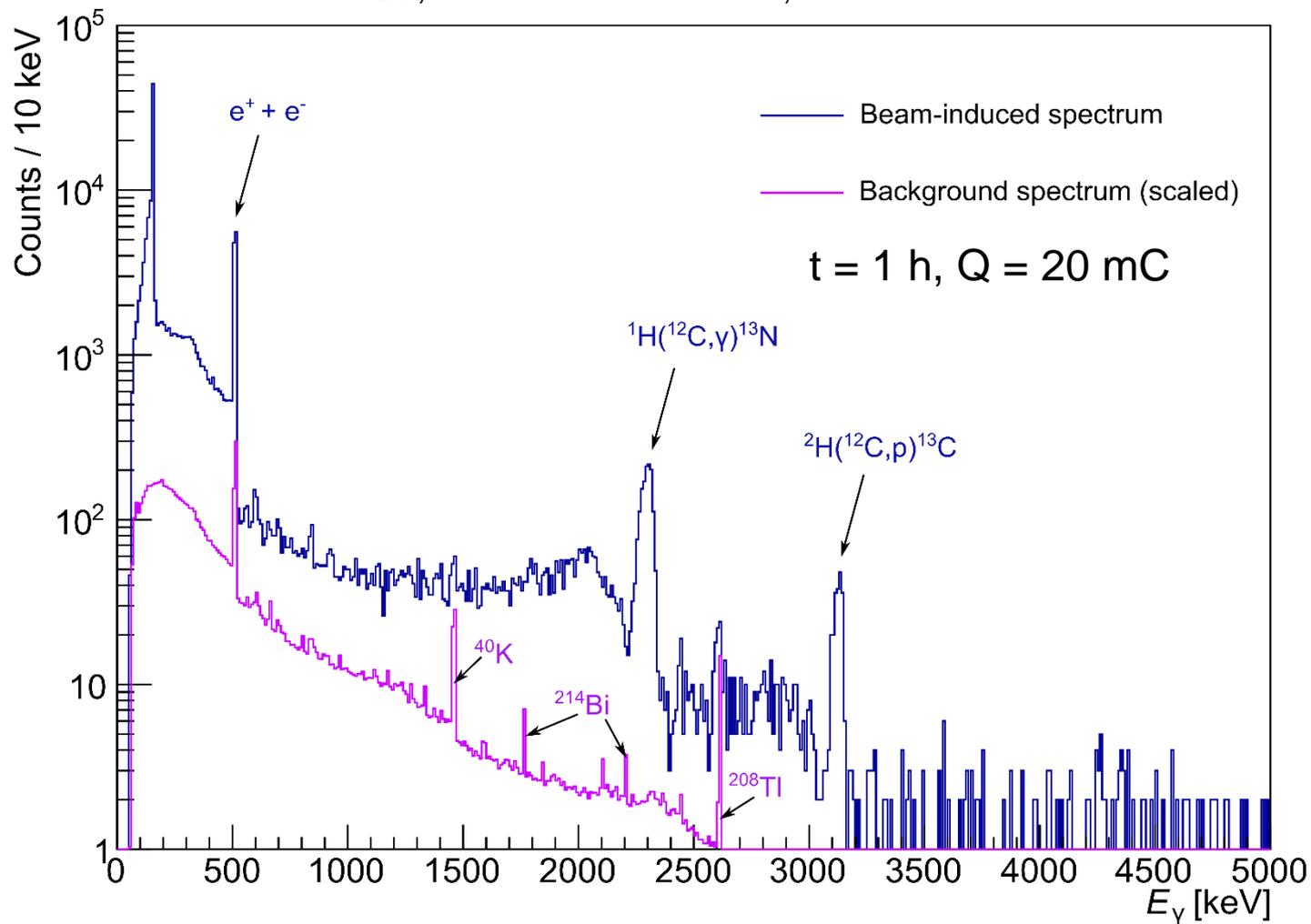
# Gamma-ray spectra

$$E_{12\text{C},0} = 3.0 \text{ MeV} \quad (E_{\text{cm},0} = 230 \text{ keV})$$

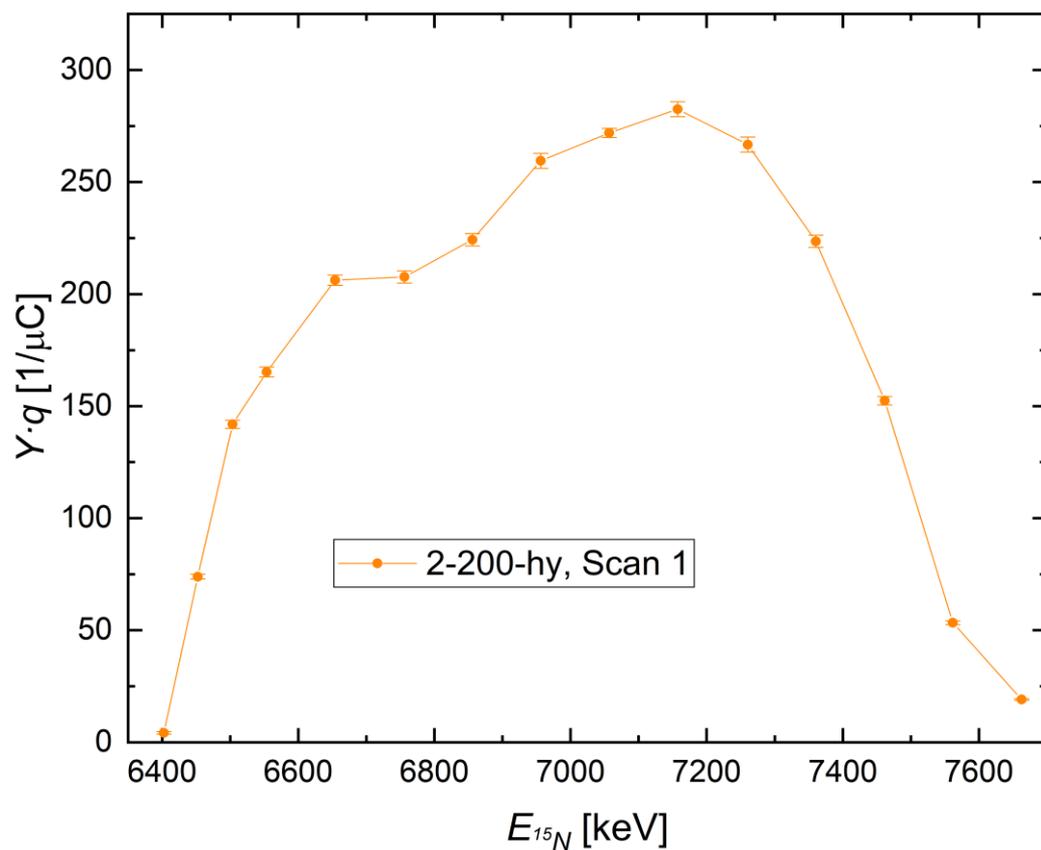


# Gamma-ray spectra

$$E_{12\text{C},0} = 4.5 \text{ MeV} \quad (E_{\text{cm},0} = 350 \text{ keV})$$



# Nuclear Resonant Reaction Analysis (NRRA)



Target #2, 200 nm, hydrated

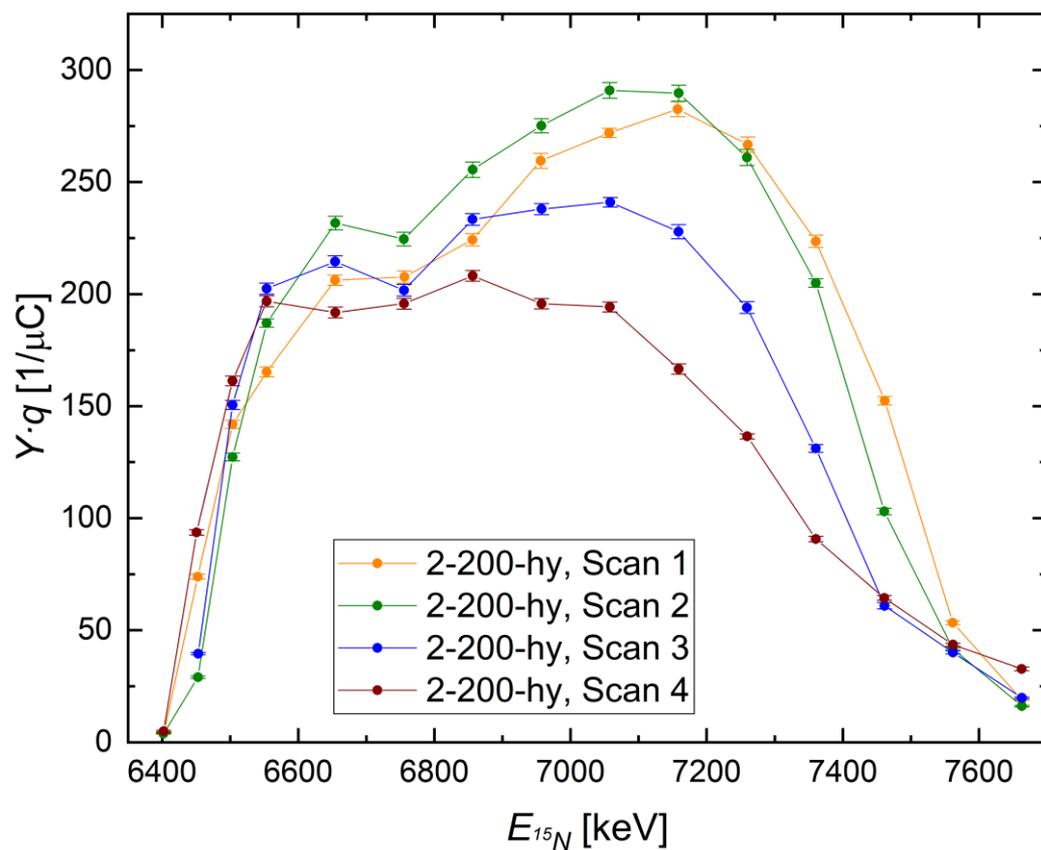
## Target scans

- Use of  ${}^1\text{H}({}^{15}\text{N}, \alpha\gamma){}^{12}\text{C}$  reaction
  - Yield of  $E_\gamma = 4439$  keV
  - $E_{15\text{N}} = 6.4$  MeV narrow resonance

Reinhardt+, Nucl. Inst. Meth. B 381, 58-66 (2016)

- Measured each day and after switching the target
- Changed target after significant decrease of hydrogen content
  - After 2-3 days for 200 nm targets
  - After 1 day for 100 nm targets

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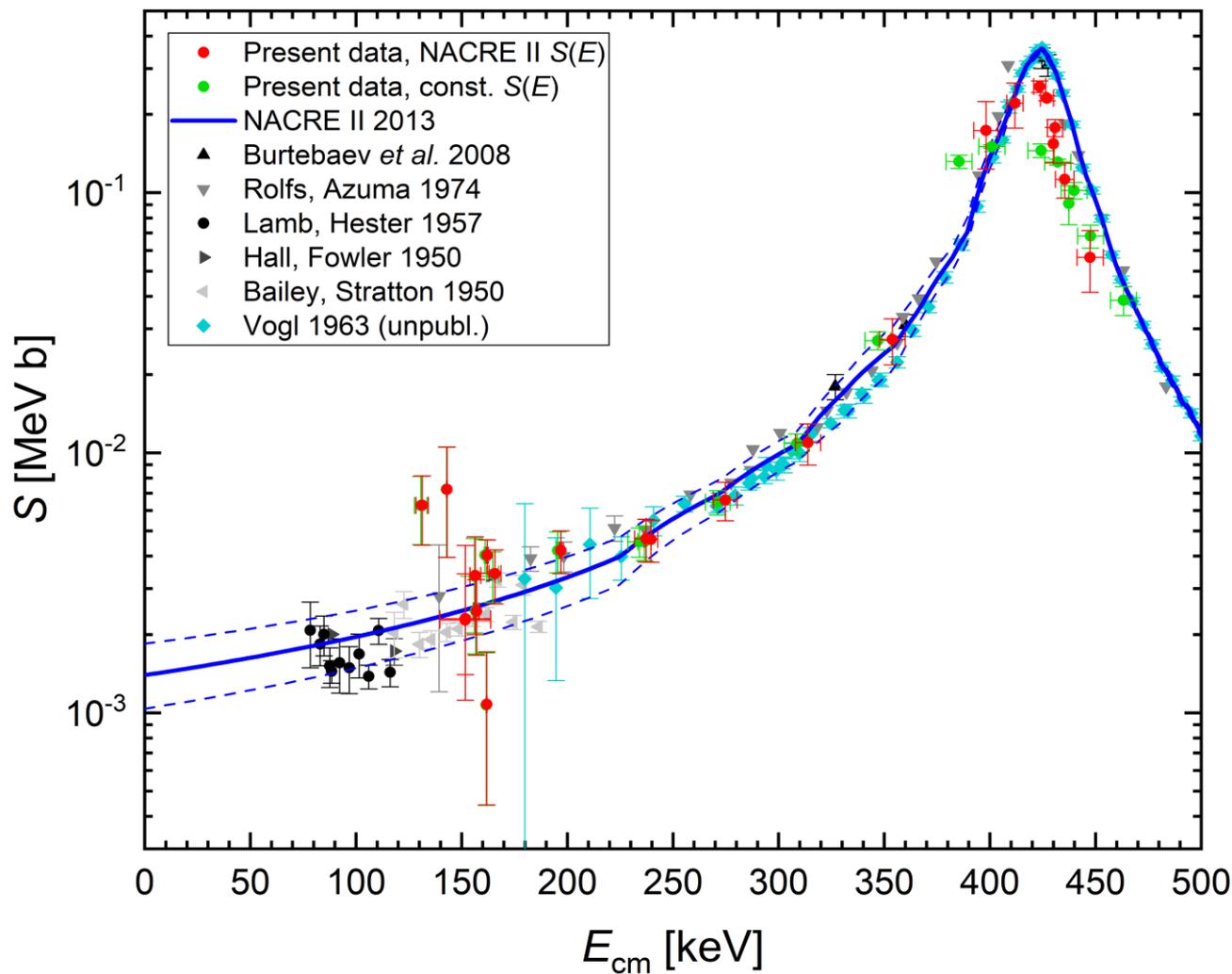
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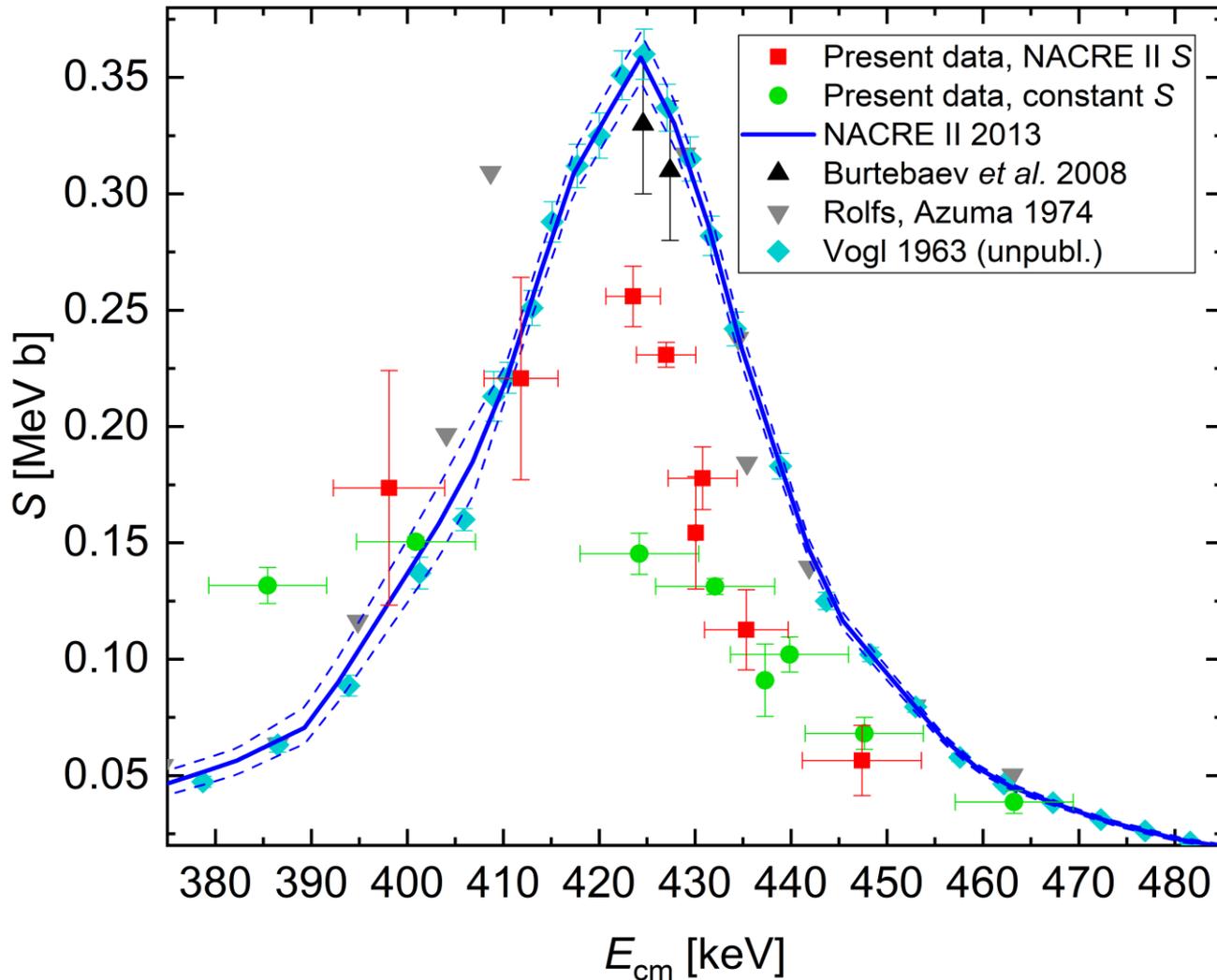
Reinhardt+, Nucl. Inst. Meth. B 381, 58-66 (2016)

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# Assumed energy dependence of $S$



# Assumed energy dependence of $S$ (continued)



Resonance energy?

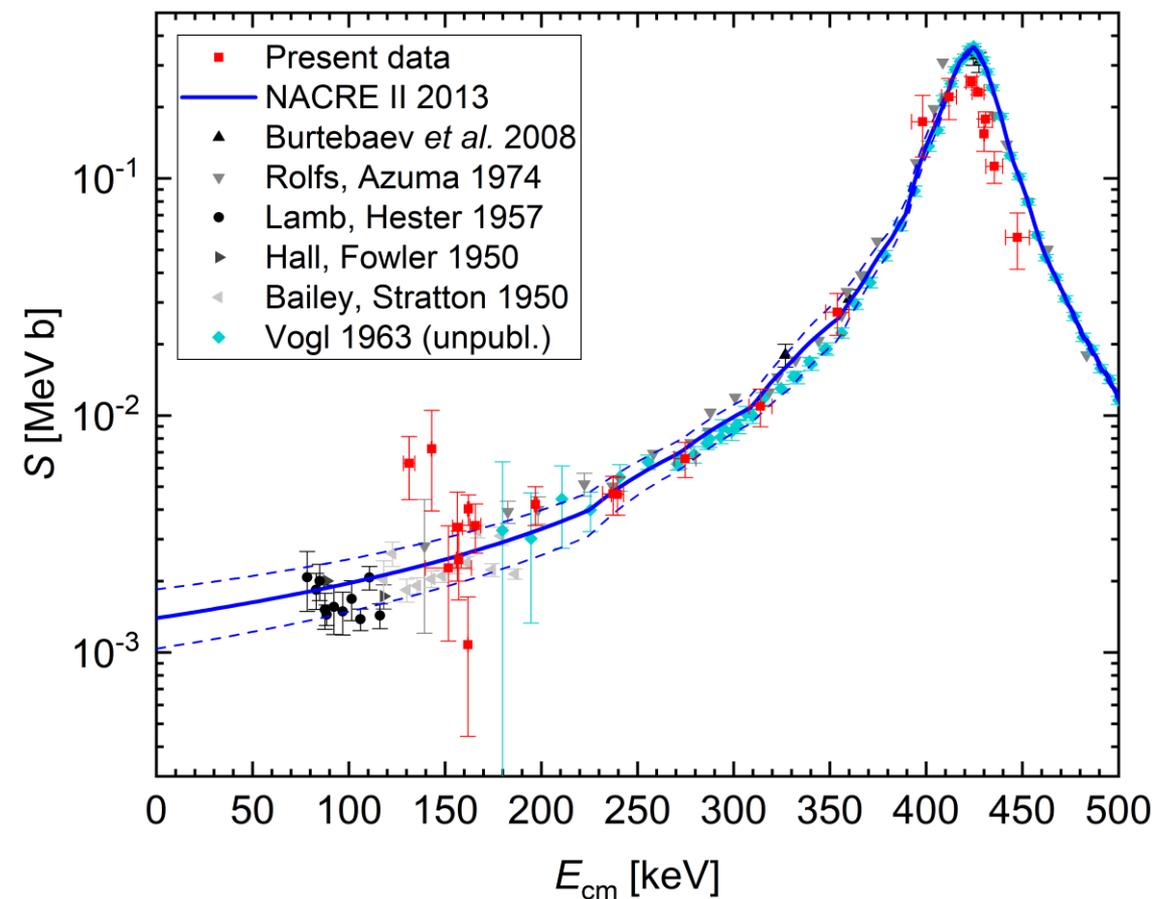
ENSDF

$E_x = (2364.9 \pm 0.6) \text{ keV}$

Width: 31.7 keV

	$E_R$ [keV]
Fowler+ (1949)	$420.7 \pm 1.8$
Hunt+ (1953)	$421.4 \pm 0.5$
Vogl (1963)	424.9
Rolfs+ (1974)	$421.6 \pm 0.9$
Blatt+ (1974)	421.5
Hinds+ (1992)	423.2

# Conclusion



130 – 200 keV:

- Present data on average about 20 % higher than NACRE II (consistent within error bars)
- Lower limit given by cosmic ray induced BG

230 – 415 keV:

- Present data consistent with NACRE II fit

420 – 450 keV:

- Present data significantly lower

## Summary

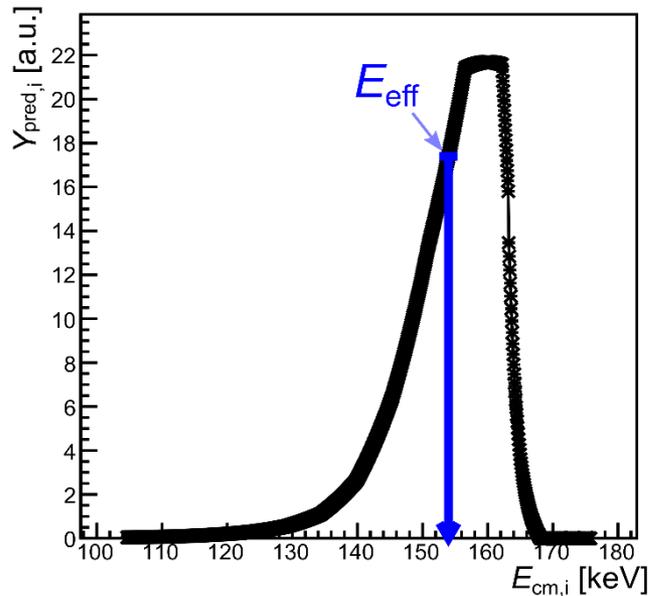
- Measurement of  $^{12}\text{C}(p,\gamma)^{13}\text{N}$  S-Factor in inverse kinematics at  $E_{cm} = 130 - 450$  keV
- New data in low energy region about 20% higher than NACRE II
- New data between 420 – 450 keV significantly below fit
  - Possible problem with resonance energy

## Outlook

- Use of new data for extrapolation towards astrophysical energies
- Check resonance energy
- Extension towards lower energies
  - Use of underground accelerator to reduce laboratory background
    - LUNA Gran Sasso → direct kinematics
    - Felsenkeller Dresden → inverse kinematics

# Backup

# Determination of S factor



## Predicted yield $Y_{\text{pred}}$

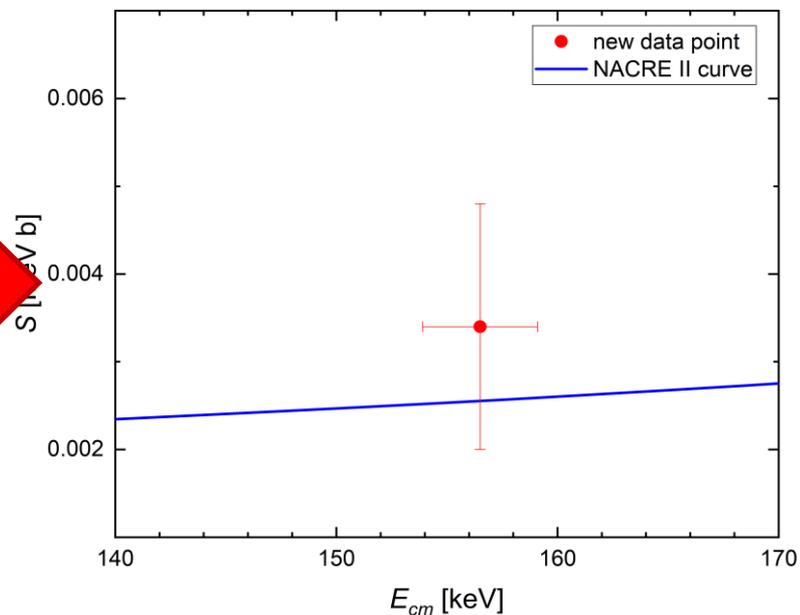
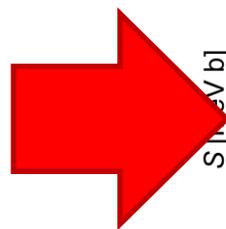
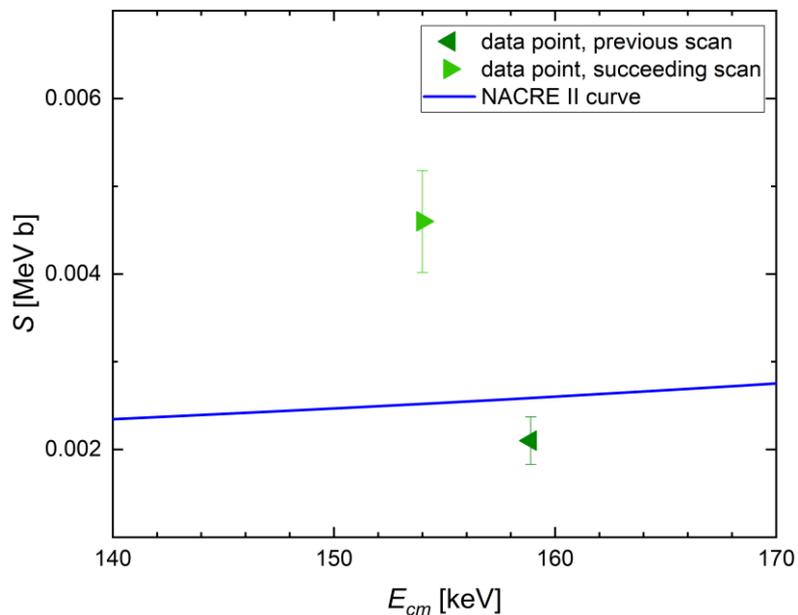
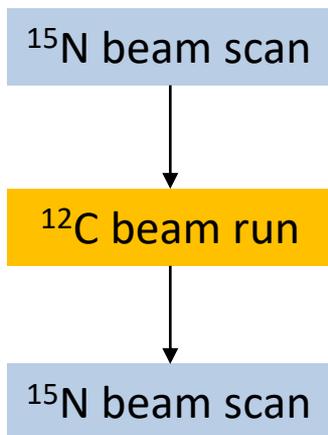
- Measured target profile divided into 1000 thin slices
- Calculation of energy  $E_{\text{cm},i}$  in each slice using stopping power from SRIM
- Assumed energy dependence  $S_{\text{pred}}(E)$  from NACRE II curve

## Effective Energy $E_{\text{eff}}$

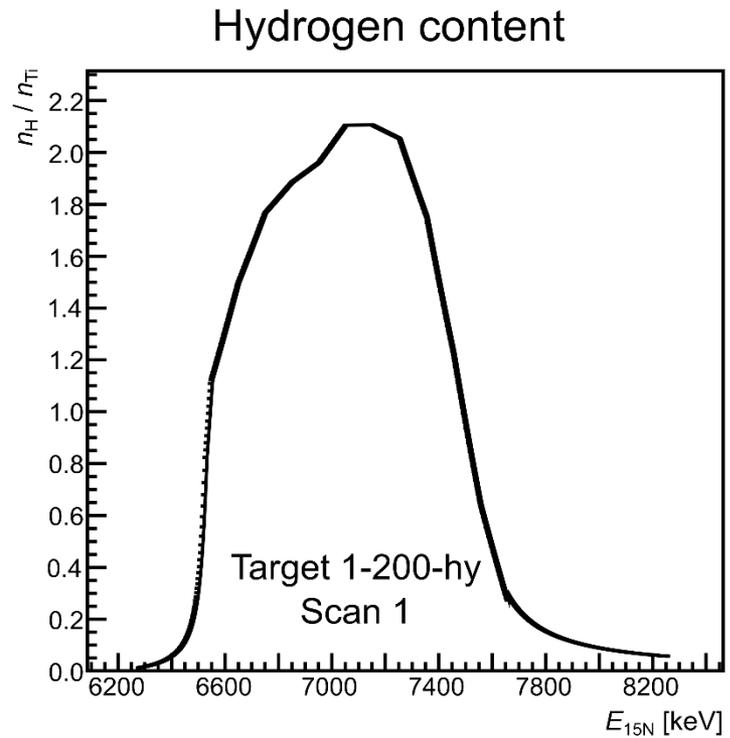
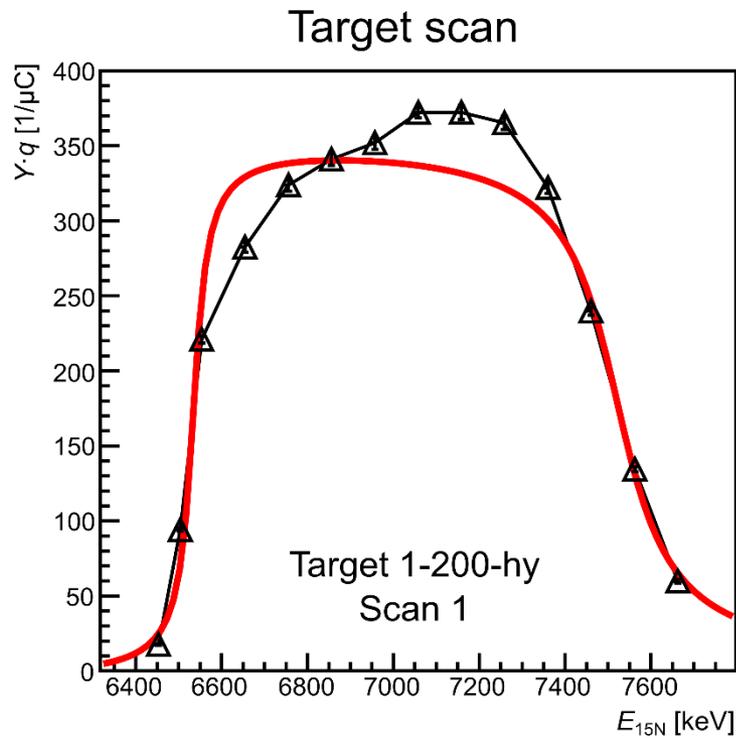
- Mean value of the energy of each slice weighted with its yield contribution

$$S_{\text{exp}}(E_{\text{eff}}) = \frac{Y_{\text{exp}}}{Y_{\text{pred}}} S_{\text{pred}}(E_{\text{eff}})$$

# Determination of S factor (continued)



# Hydrogen content



# Uncertainties

S-Factor:

Quantity	Stat. uncert.	Sys. uncert.
Counts	1.0 – 54 %	
Change of H content	0.7 – 39 %	
Stopping (SRIM)		3 – 8 %
Fit of target scan		0.1 – 3 %
Efficiency		1.6 %
Charge		1 %
total	2.3 – 59 %	3.5 – 8.7 %

Effective energy:

Quantity	Stat. uncert.	Sys. uncert.
Change of H content	0.0 – 6.2 keV	
Accelerator voltage	0.1 keV	
Stopping (SRIM)		0.9 – 2.9 keV
Fit of target scan		0.1 – 1.3 keV
Energy calibration		0.2 keV
total	0.1 – 6.2 keV	0.9 – 3.2 keV