

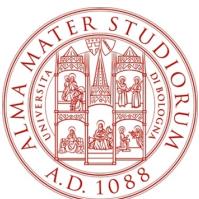
Eu N  
P C

$\beta\beta$

# Results from the CUORE experiment



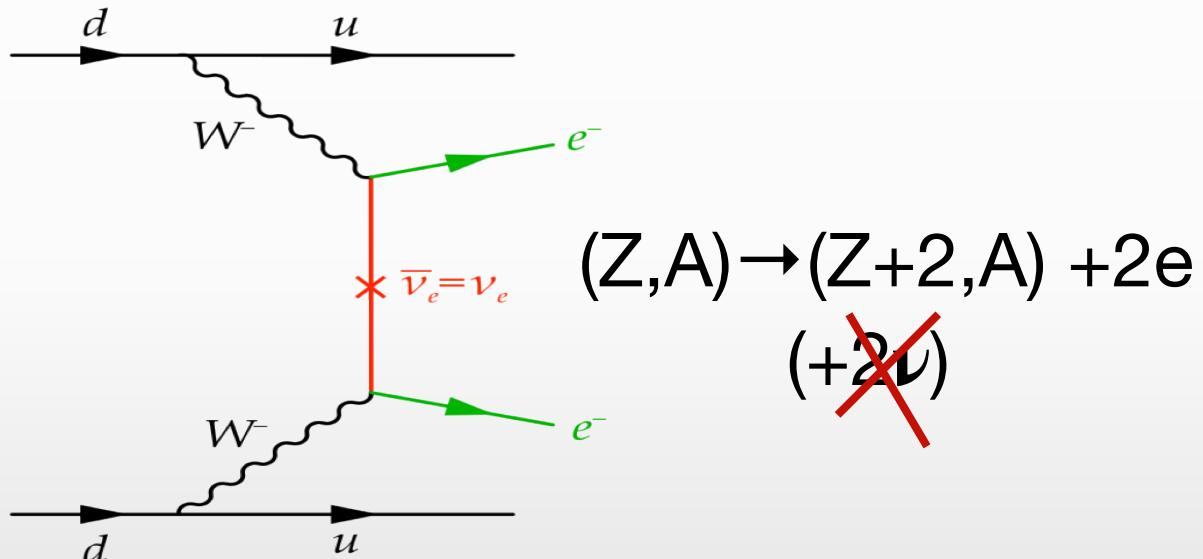
Niccolò Moggi - Univ. and INFN Bologna  
on behalf the CUORE Collaboration  
Bologna – September 2018





# $0\nu\beta\beta$ decay and signature

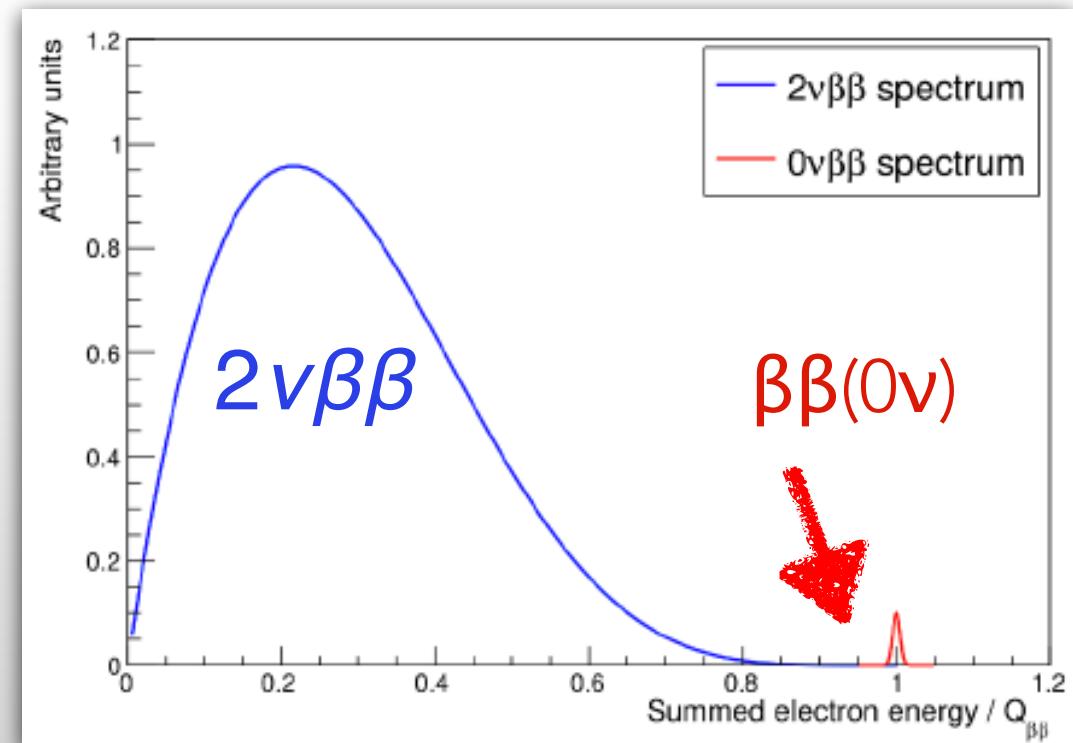
CUORE primary goal is the search  
for  $0\nu\beta\beta$  decay:  $^{130}\text{Te} \rightarrow ^{130}\text{Xe} + 2e^-$



- SM forbidden
- $\Delta L = 2$
- If observed  $\Rightarrow \nu$  is a Majorana particle

Observable: line at Q-value

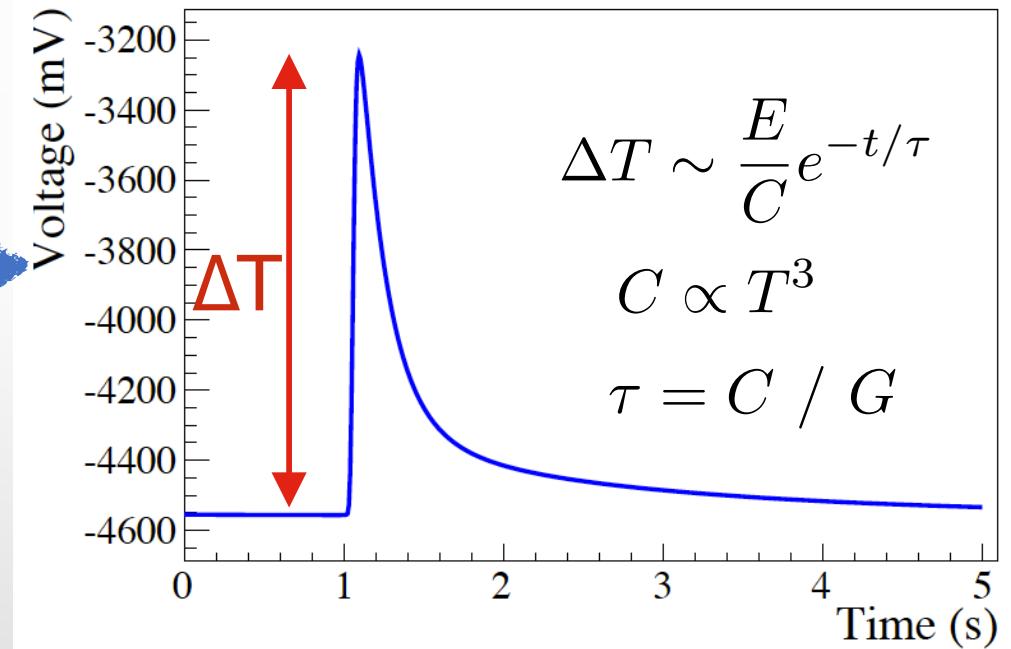
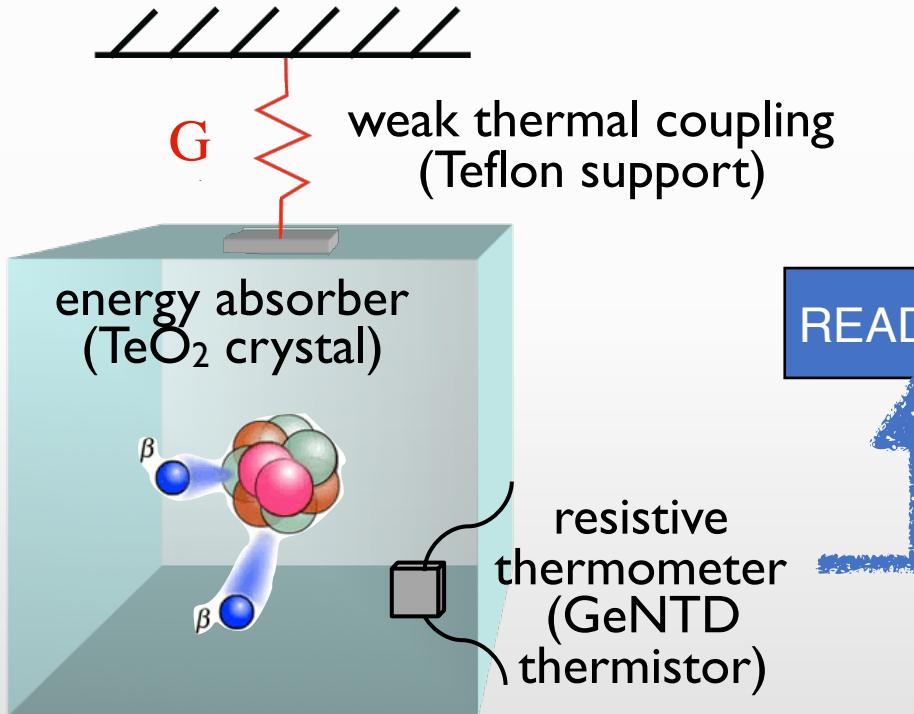
- measure  $E_{\beta\beta} = E_{\beta_1} + E_{\beta_2}$
- smeared by energy resolution
- $Q(\text{Te}) \approx 2528 \text{ keV}$





# The bolometric technique

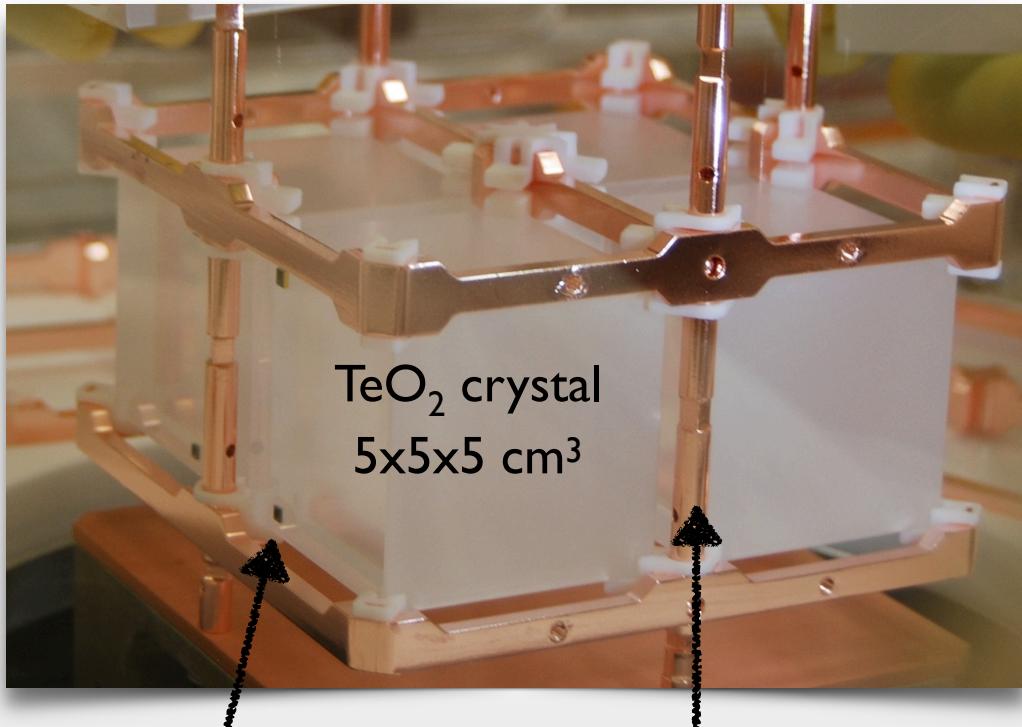
The detector: 1) contains the isotope 2) absorbs the energy 3) increase its temperature  
heat sink (Cu holder)



- ▶ 2e- mostly contained in the bulk
- ▶ excellent efficiency & energy resolution
- ▶ hardly discriminate signal from bkg

Working temperature:  
@ ~10 mK

# CUORE bolometers: TeO<sub>2</sub> crystals



NTD-Ge  
thermistor

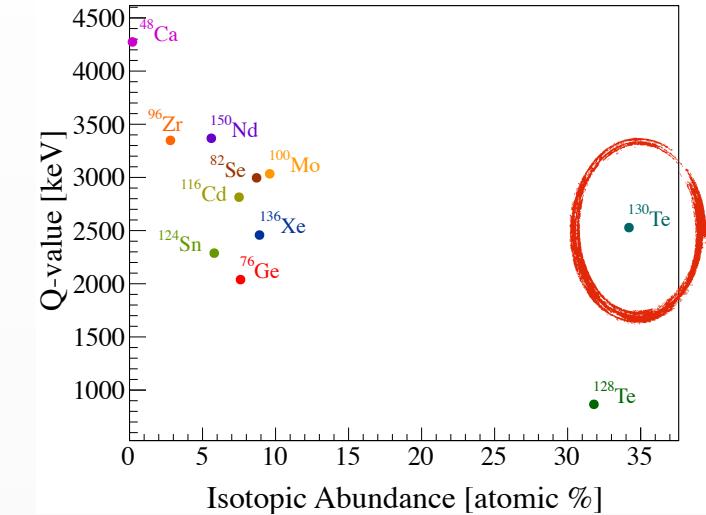
copper  
frame

With:  
 $E \sim 1 \text{ MeV}$   
 $C \sim 10^{-9} \text{ J/K} @ 10 \text{ mK}$

$$\left. \Delta T \approx 0.1 \text{ mK/MeV} \right]$$

Choice of TeO<sub>2</sub>:

- ▶  $^{130}\text{Te}$  abundance = 34%
- ▶  $Q = 2527.5 \text{ keV}$  above most of natural radioactivity
- ▶ high quality crystals and large mass scale possible



The sensitivity may be expressed as:

$$\left( T_{1/2}^{0\nu} \right)^{sens} \propto i.a. \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{b \cdot \Delta E}}$$

↓ Isotopic abundance  
 ↑ Efficiency  
 ↓ Mass  
 ↓ Exposure time  
 ↓ Bkg. rate  
 ↓ Energy resolution

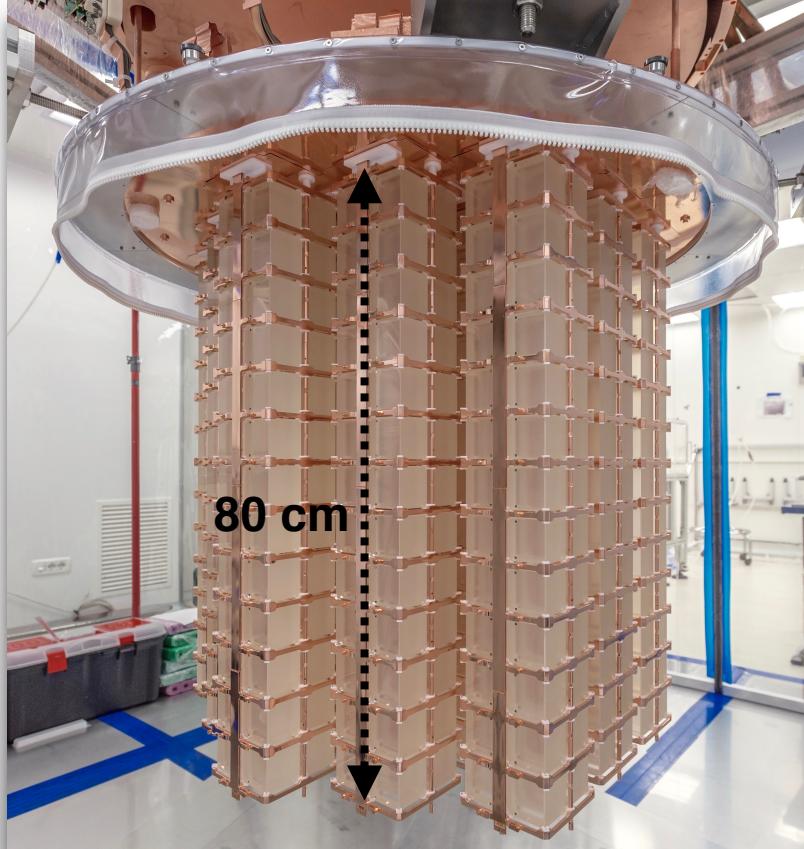


# Cryogenic Underground Observatory for Rare Events

1- detector

2- cryostat

3- background mitigation + passive shielding



- ▶ 988 crystals 5x5x5 cm,  
closely packed
- ▶ arranged in 19 towers  
of 13 floors each
- ▶ 742 kg (206 kg of  $^{130}\text{Te}$ )
- ▶ background goal: 0.01  
counts/(kg keV yr)
- ▶ energy resolution goal:  
5 keV FWHM in the  
ROI

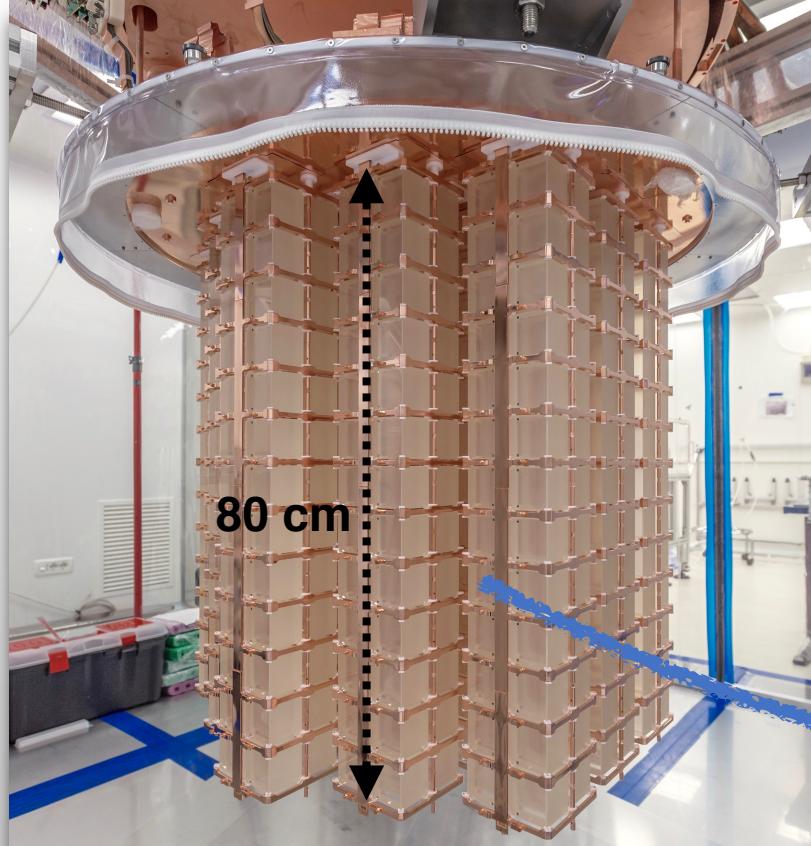


# Cryogenic Underground Observatory for Rare Events

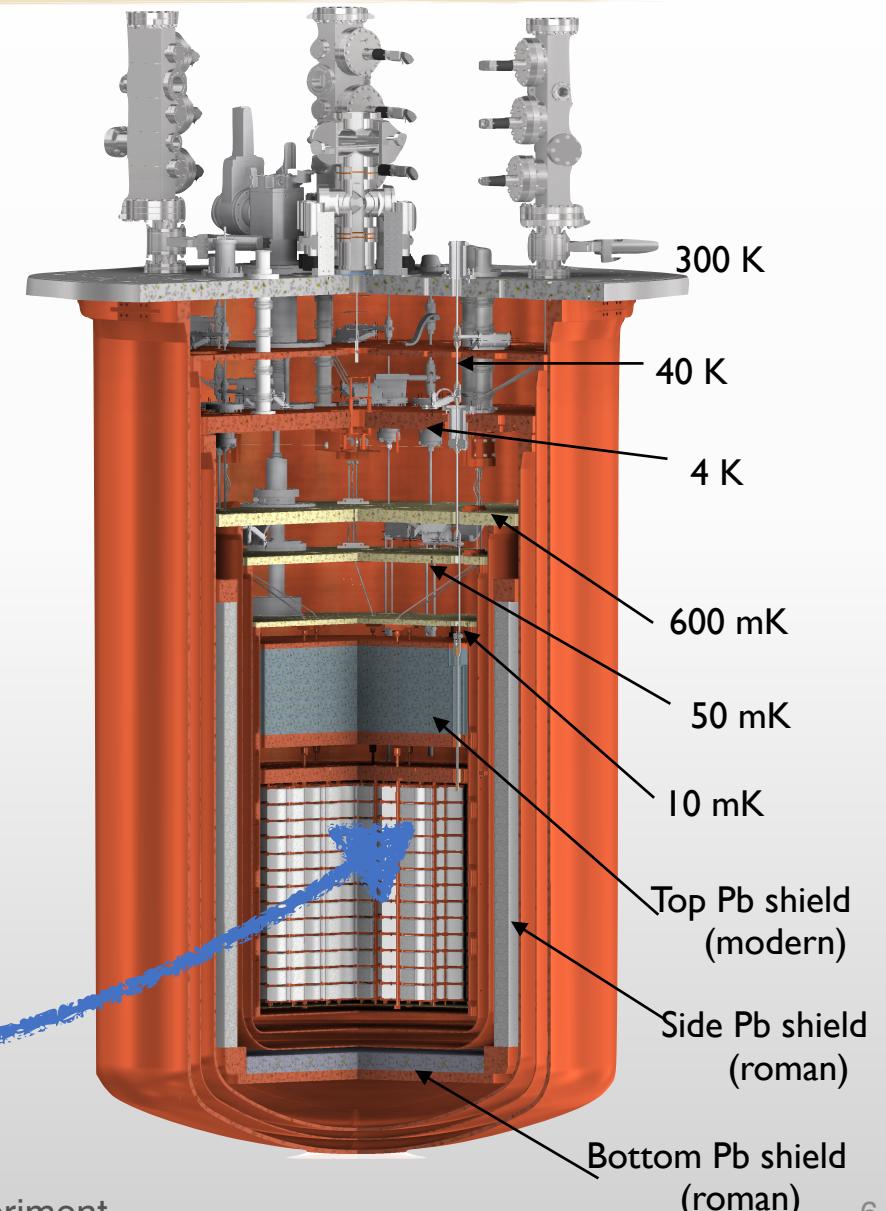
**1- detector**

**2- cryostat**

**3- background mitigation + passive shielding**



- ▶ 988 crystals 5x5x5 cm, closely packed
- ▶ arranged in 19 towers of 13 floors each
- ▶ 742 kg (206 kg of <sup>130</sup>Te)
- ▶ background goal: 0.01 counts/(kg keV yr)
- ▶ energy resolution goal: 5 keV FWHM in the ROI



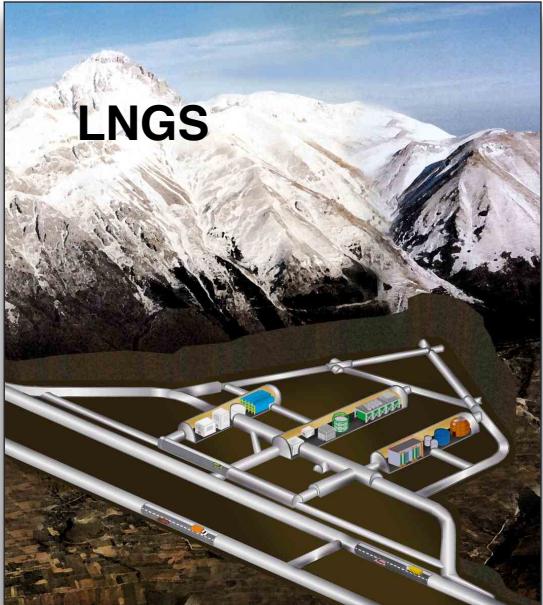


# Cryogenic Underground Observatory for Rare Events

1- detector

2- cryostat

**3- background mitigation + passive shielding**



- 3600 m.w.e. deep equivalent
- 1400 m rock overburden cuts muon flux by  $\sim 10^6$

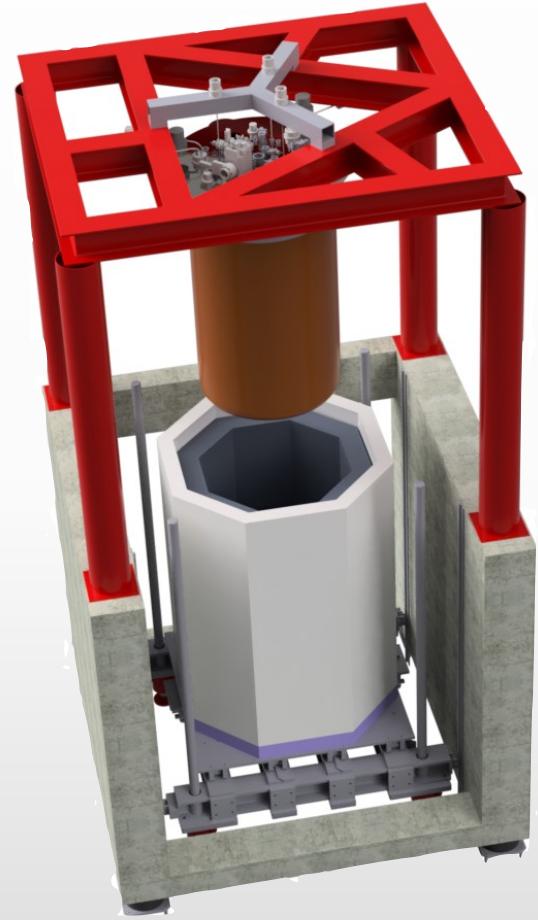


- tower assembly and storage in N<sub>2</sub> atmosphere @LNGS
- ✓ material selection
- ✓ crystal surface polishing
- ✓ special cleaning process for Cu and PTFE parts



Internal shields:

- 6,.5 tons ancient roman lead shied
- ✓ Cu + modern lead



External shields:

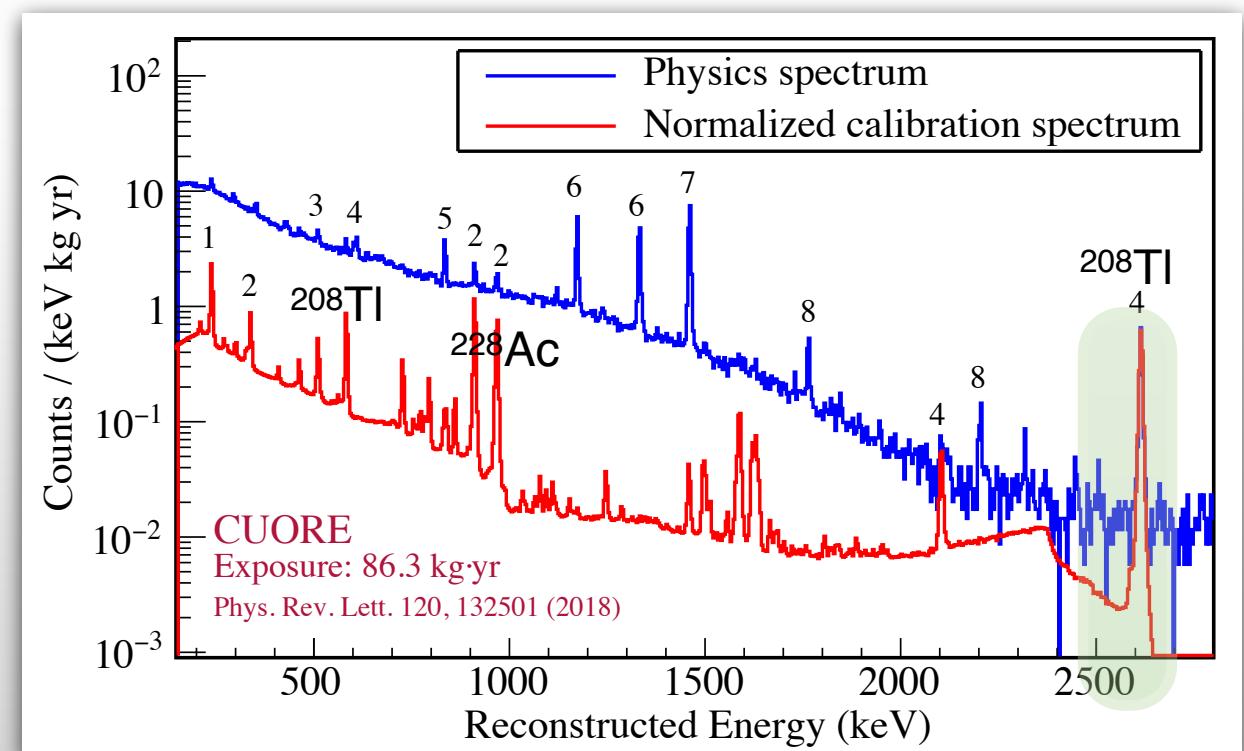
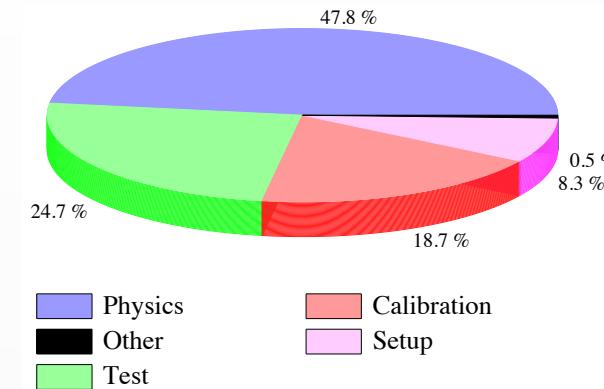
- 25 cm external Pb ( $\gamma$ )
- 18 cm polyethylene + 2 cm H<sub>3</sub>BO<sub>3</sub> (neutrons)



# CUORE data-taking 2017

- ▶ In January 2017 first pulse
- ▶ during summer 2017 two physics-data periods:
  - ▶ exposure: 86.3 kg yr of TeO<sub>2</sub> (37.6 + 48.7 kg yr)
  - ▶ <sup>130</sup>Te exposure: 24.0 kg yr
  - ▶ each dataset bracketed by calibrations
  - ▶ 984/988 active channels
  - ▶ 92% of channels passing analysis cuts
  - ▶ Signal efficiency ~80%  
( $75.7 \pm 3.0$  and  $83.0 \pm 2.6$ )

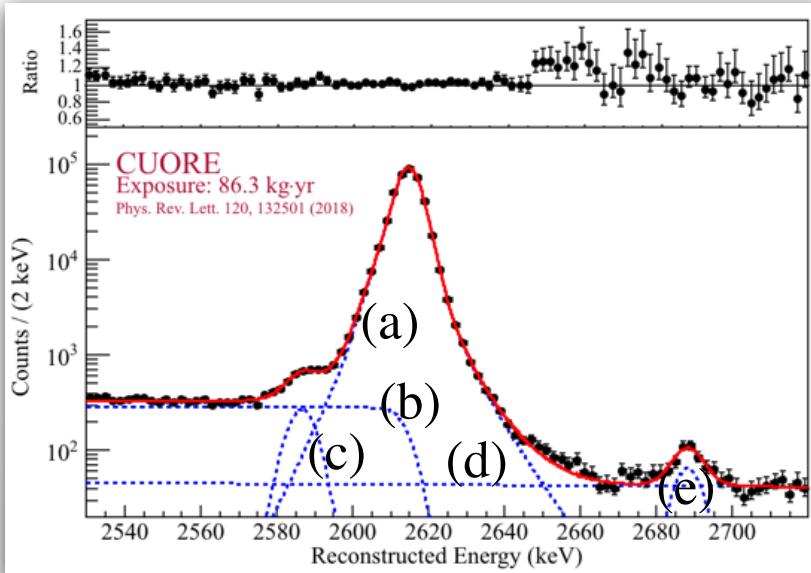
Phys. Rev. Lett. 120, 132501 (2018)



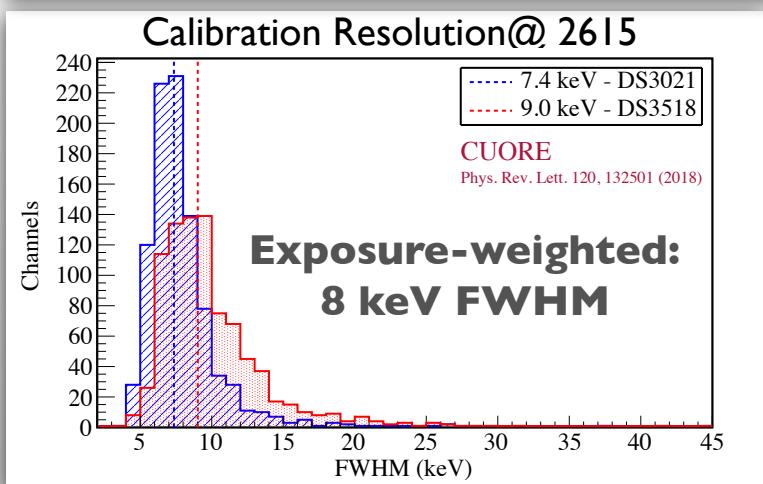


# Detector performance (energy resolution)

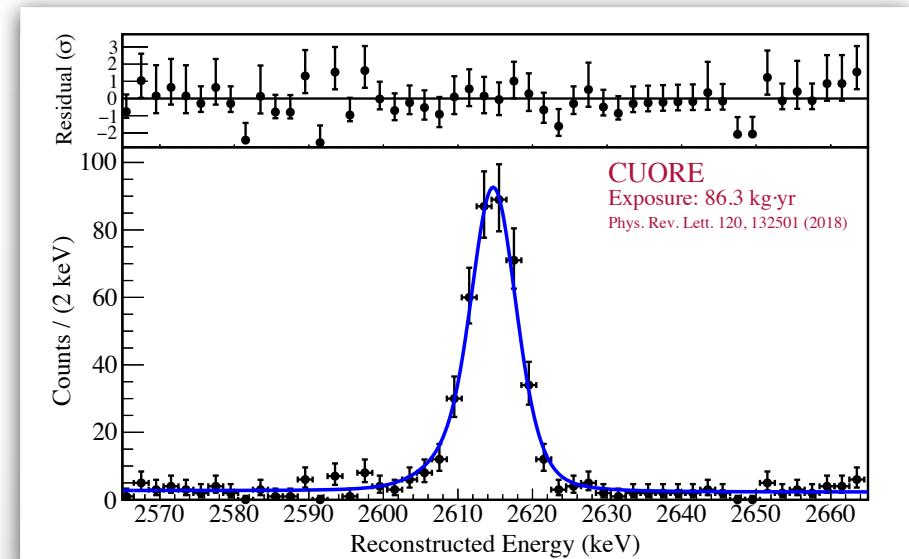
## Calibration data



$^{208}\text{TI}$  2615 keV from  $^{232}\text{Th}$  calibration is well modelled with: (a) Gaussian photopeak, (b) Compton, (c) X-ray escape 30 KeV, (d) linear bkg, (e) 2687 keV = 2615 + 583  $\gamma(^{232}\text{Th}) - 511\gamma(e^+e^-)$   
For calibration purpose a fit is done per bolometer-dataset.



## Physics data



A scaling factor is obtained from (six) spectral lines and its value extrapolated at the Q-value

Physics data resolution@ Q-value  
Dataset 1:  $(8.3 \pm 0.4)$  keV FWHM  
Dataset 2:  $(7.4 \pm 0.7)$  keV FWHM  
Exposure-weighted:  $(7.7 \pm 0.5)$  keV FWHM



# Event selection

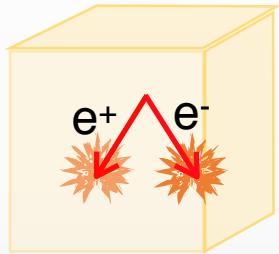
## I. Removal of low-quality data (~1% of the total live time)



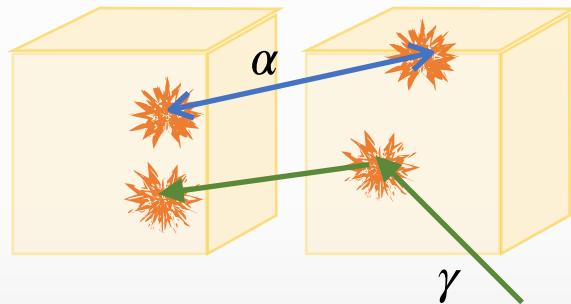
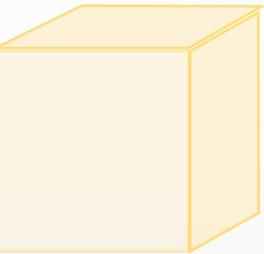
# Event selection

1. Removal of low-quality data ( $\sim 1\%$  of the total live time)

2. Select multiplicity=1 (M1) events

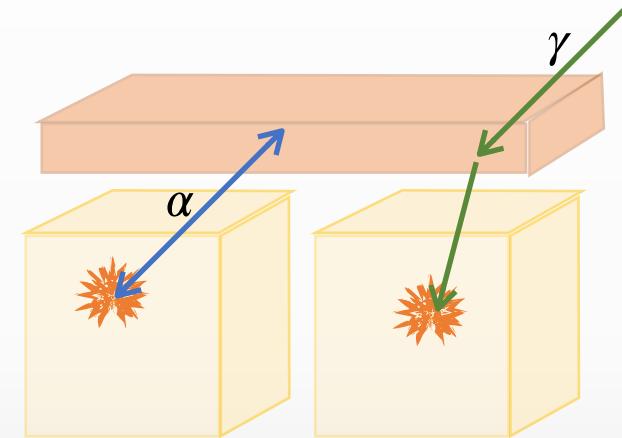


M1: signal like



M2: rejected background

Simultaneous energy deposit  
in two crystals is likely due  
to background events:  
we require no coincidences in  
a 10s time window



M1: backgrounds

Dangerous backgrounds  
mimicking energy deposit in  
one crystal

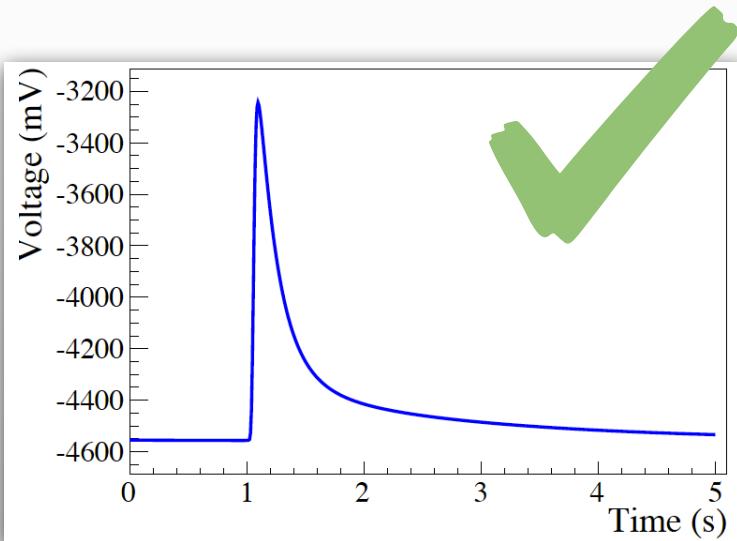


# Event selection

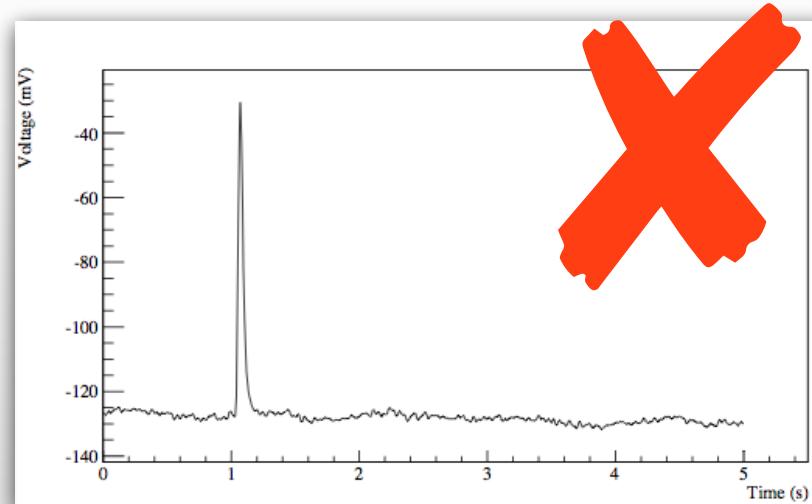
1. Removal of low-quality data (~1% of the total live time)

2. Select multiplicity=1 (M1) events

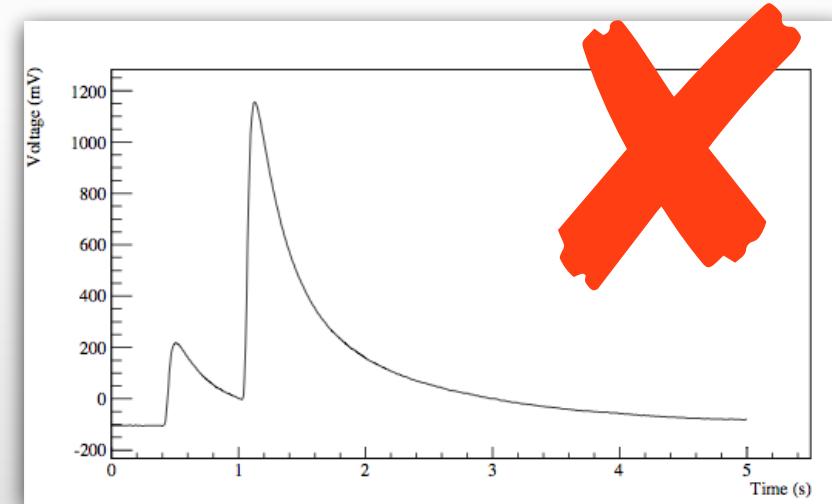
3. Pulse shape analysis



Signal



Noise



Pile-up



# Event selection

1. Removal of low-quality data ( $\sim 1\%$  of the total live time)

2. Select multiplicity=1 (M1) events

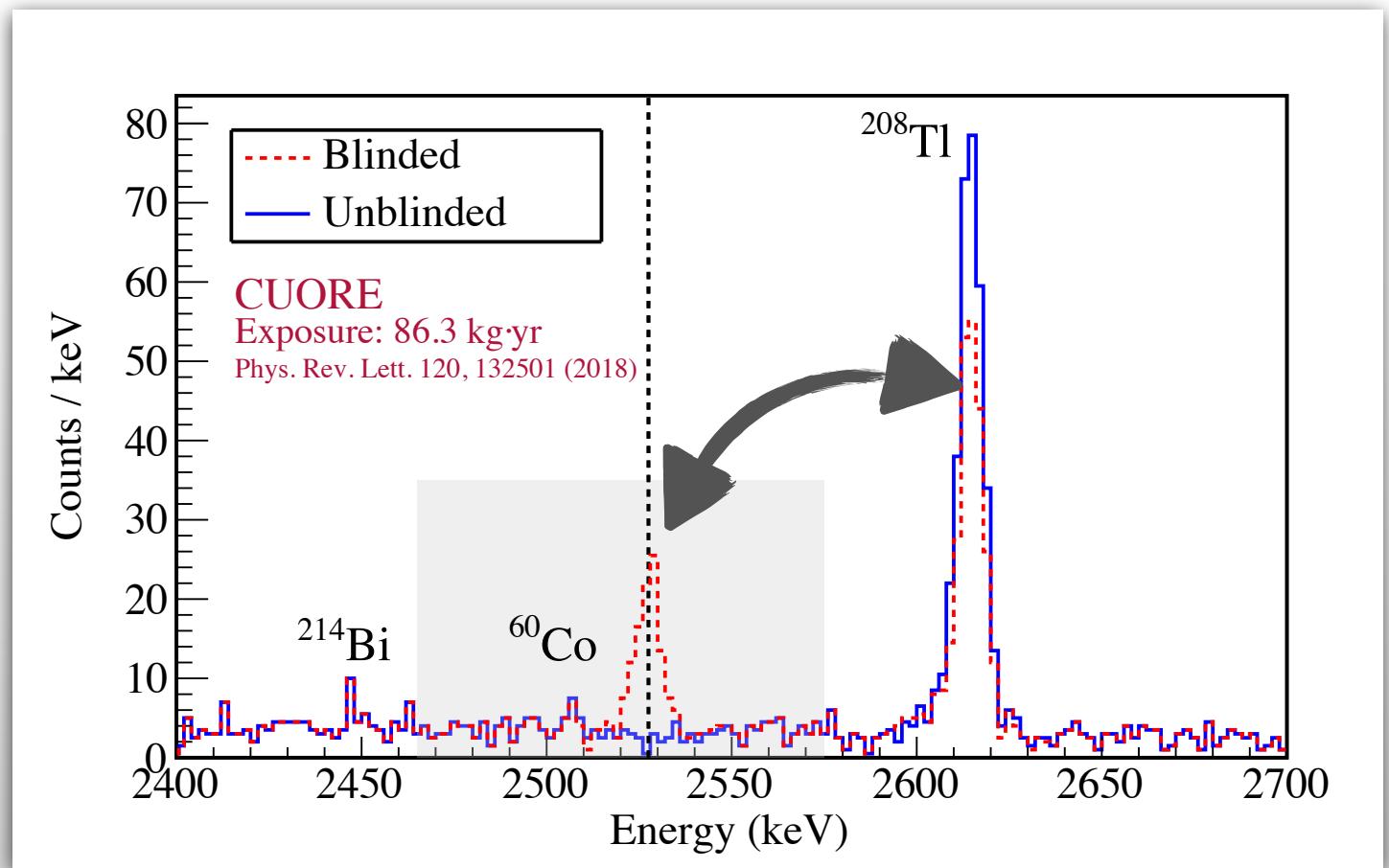
3. Pulse shape analysis

4. Blinding

Swap an (unknown) fraction of events from  $^{208}\text{Tl}$  peak to the Q-value region.

Unblinding only when the full analysis procedure is fixed

155 events are left in the ROI



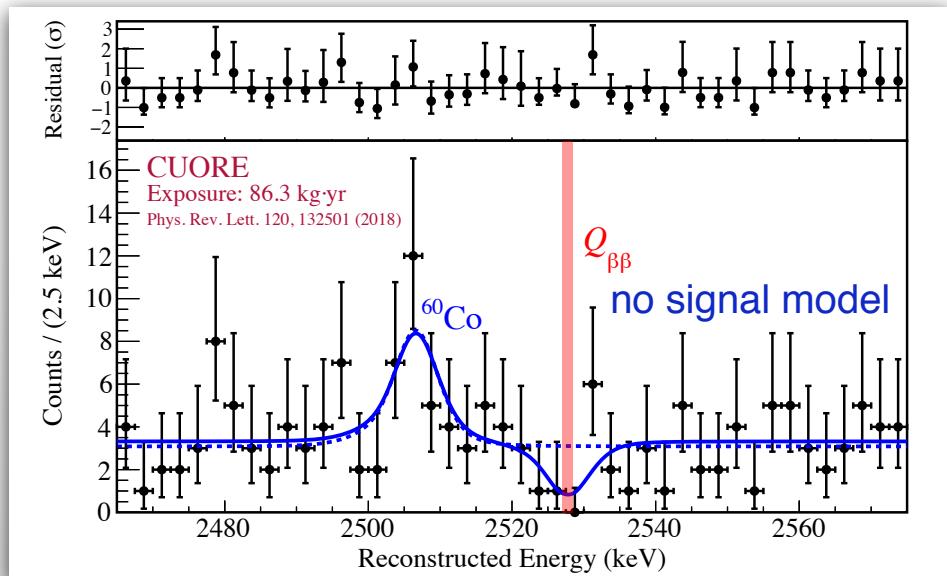


# Fit and results

Phys. Rev. Lett. 120, 132501 (2018)

Extended maximum likelihood fit in 2465-2575 keV:

- ${}^{60}\text{Co}$  peak 2506 keV (floating position)
- peak at  $Q_{\beta\beta}$  (fixed position, floating rate)
- flat background (dataset dependent)



Background Index in counts/(kg keV yr):

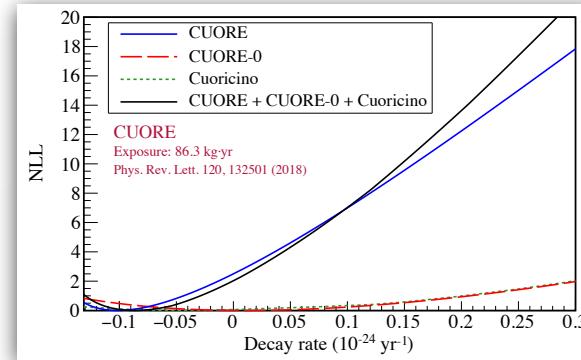
1st dataset :  $(1.49^{+0.18}_{-0.17}) \times 10^{-2}$

2nd dataset :  $(1.35^{+0.20}_{-0.18}) \times 10^{-2}$

**Exposure weighted :  $0.014 \pm 0.2$**

Best fit decay rate:

$$\Gamma_{0\nu} = [-1.0^{+0.4}_{-0.3}(\text{stat}) \pm 0.1(\text{syst})] \times 10^{-25} \text{ yr}^{-1}$$



No evidence for  $0\nu\beta\beta$

Decay rate limit (90% C.L. incl. syst.) =  $0.51 \times 10^{-25} \text{ yr}^{-1}$

$T_{1/2}$  (90% C.L. including syst.) =  $1.3 \times 10^{25} \text{ yr}$

We combine CUORE results with Cuoricino (19.75 kg yr) and CUORE-0 (9.8 kg yr) data and obtain:

**$T_{1/2}$  (90% C.L.) >  $1.5 \times 10^{25} \text{ yr}$**



# Interpretation of results

**$T_{1/2}$  (90% C.L.) >  $1.5 \times 10^{25}$  yr**

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G^{0\nu}(Q, Z) \left|M_{nucl}^{0\nu}\right|^2 \frac{\langle m_{\beta\beta}^2 \rangle}{m_e^2}$$

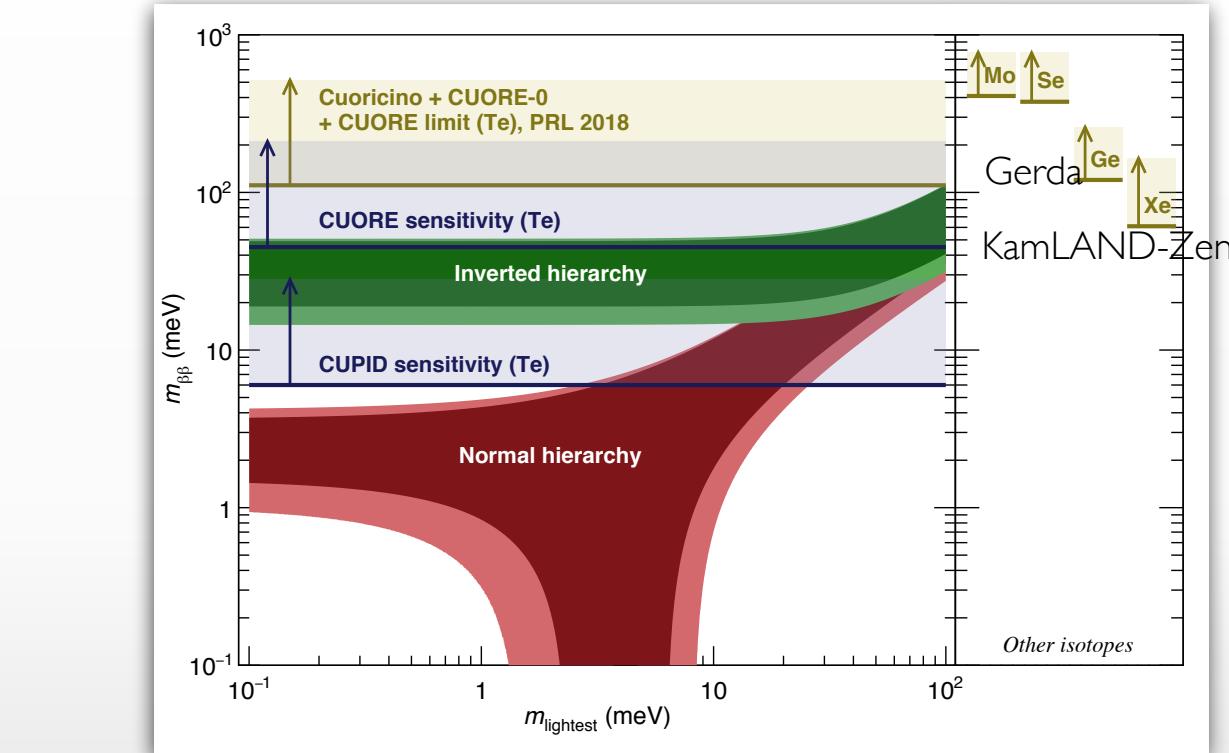
**$m_{\beta\beta} < 110 - 520$  meV (90% C.L.)**

## Future sensitivity

Expected in 5 yr lifetime:  $T_{1/2} = 9 \times 10^{25}$  yr

$m_{\beta\beta} < 60 - 165$  meV

assuming bkg = 0.01 c/(kg keV  $\gamma$ )  
and  $\Delta E = 5$  keV FWHM



Assuming  
 $g_A=1.27$   
and  
N.M.E.

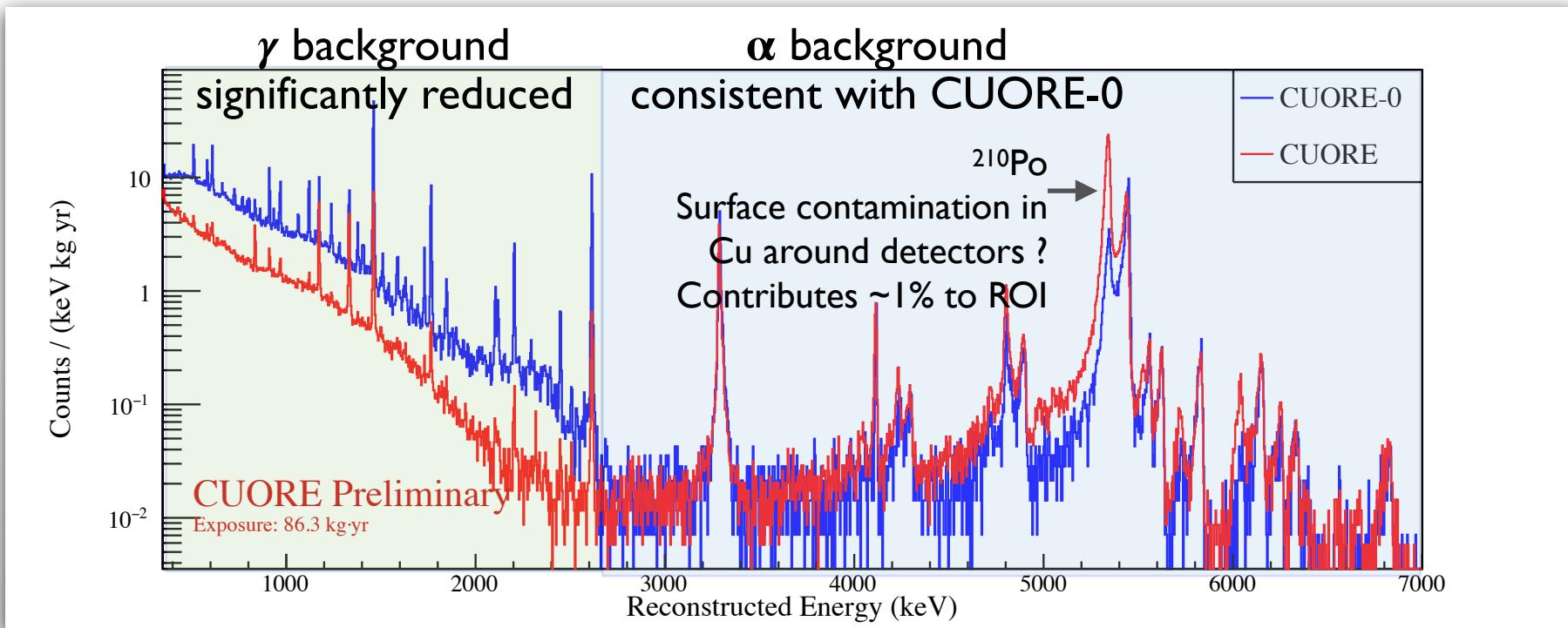
- JHEP02 (2013) 025
- Nucl. Phys. A 818, 139 (2009), ISM
- Phys. Rev. C 87, 045501 (2013), QRPA
- Phys. Rev. C 87, 064302 (2014), QRPA
- Phys. Rev. C 91, 034304 (2015), IBM-2
- Phys. Rev. C 91, 024613 (2015), pnQRPA
- Phys. Rev. C 91, 024309 (2015), SM
- Phys. Rev. C 91, 024316 (2015)
- Phys. Rev. Lett. 105, 252503 (2010), EDF
- Phys. Rev. Lett. 111, 142501 (2013), EDF

- $^{130}\text{Te}$ :  $1.5 \times 10^{25}$  yr, this analysis
- $^{76}\text{Ge}$ :  $8.0 \times 10^{25}$  yr, PRL 120, 132503 (2018)
- $^{136}\text{Xe}$ :  $1.1 \times 10^{26}$  yr, PRL 117, 082503 (2016)
- $^{100}\text{Mo}$ :  $1.1 \times 10^{24}$  yr, PR D89, 111101 (2014)
- $^{82}\text{Se}$ :  $2.4 \times 10^{24}$  yr, PRL 120, 232502 (2018)



# CUORE background

Compare to CUORE-0 (first CUORE tower tested in older cryostat)



background  
consistent with  
expectations

Origin of background:

- ▶ contamination of cryostat, shields and detector parts:
  - $^{232}\text{Th}$ ,  $^{238}\text{U}$ ,  $^{40}\text{K}$  natural contaminations
  - environmental  $^{222}\text{Rn}$  resulting in  $^{210}\text{Pb}$  surface implantation
- ▶ environmental muons and neutrons

Dominating background in ROI:

- ▶  $\alpha$  from surface of crystal and copper close-parts



# CUORE background model

## I. MC model

- ▶ Simulate the contaminations coming from different origin (= detector location + source type) using a detailed Geant4 MC simulation
- ▶ ~60 independent parameters representing various contaminations that could contribute to the CUORE background model
- ▶ detailed detector geometry description (cryostat, shieds, detector parts)

## 2. CUORE data

- ▶ 86.3 kg yr of TeO<sub>2</sub> (same as 0νBB analysis)
- ▶ split in M1, M2 and M2sum
  - ▶ M1 more sensitive to signal events
  - ▶ M2, M2sum constrain subsets of backgrounds
- ▶ split M1 between inner and outer layers (2 crystals thick)
  - ▶ outer more sensible to backgrounds originated outside the detector

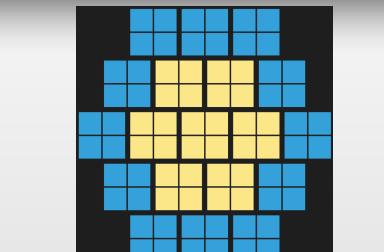
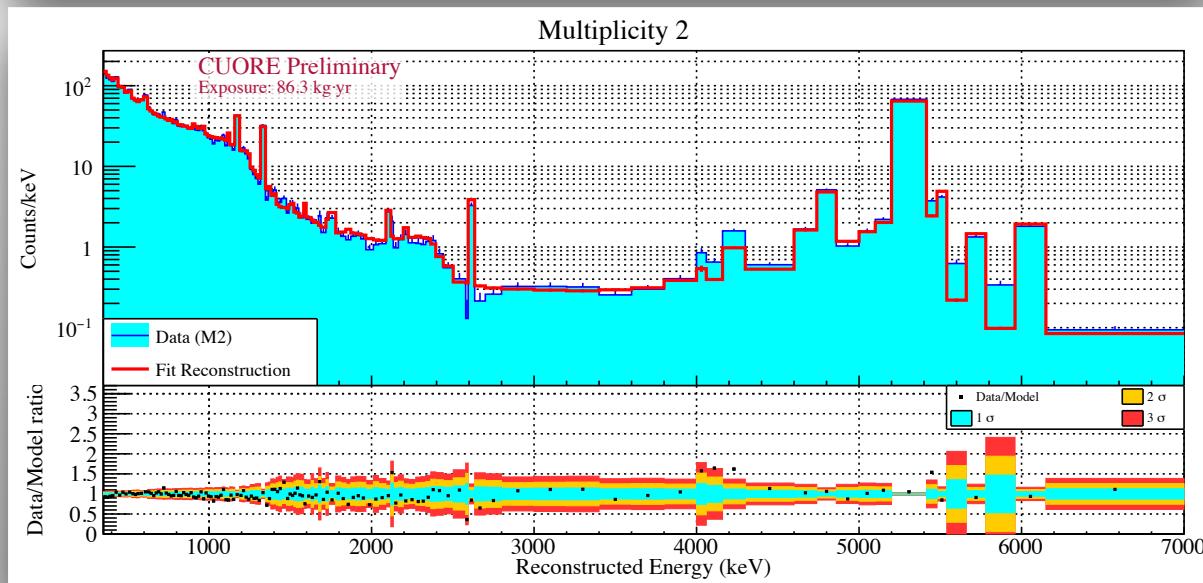
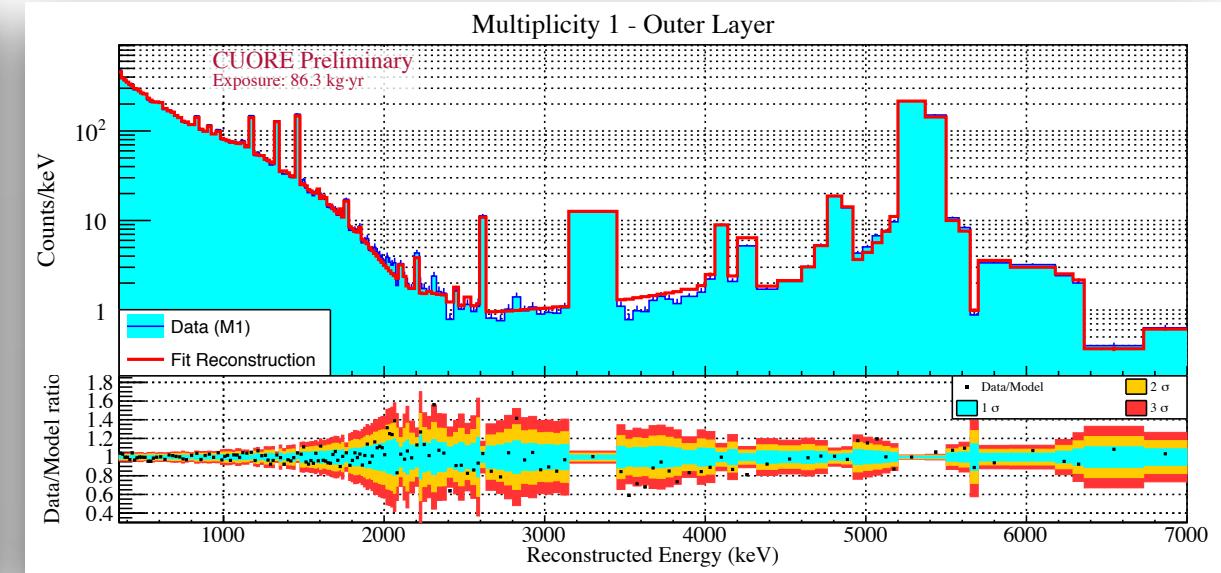
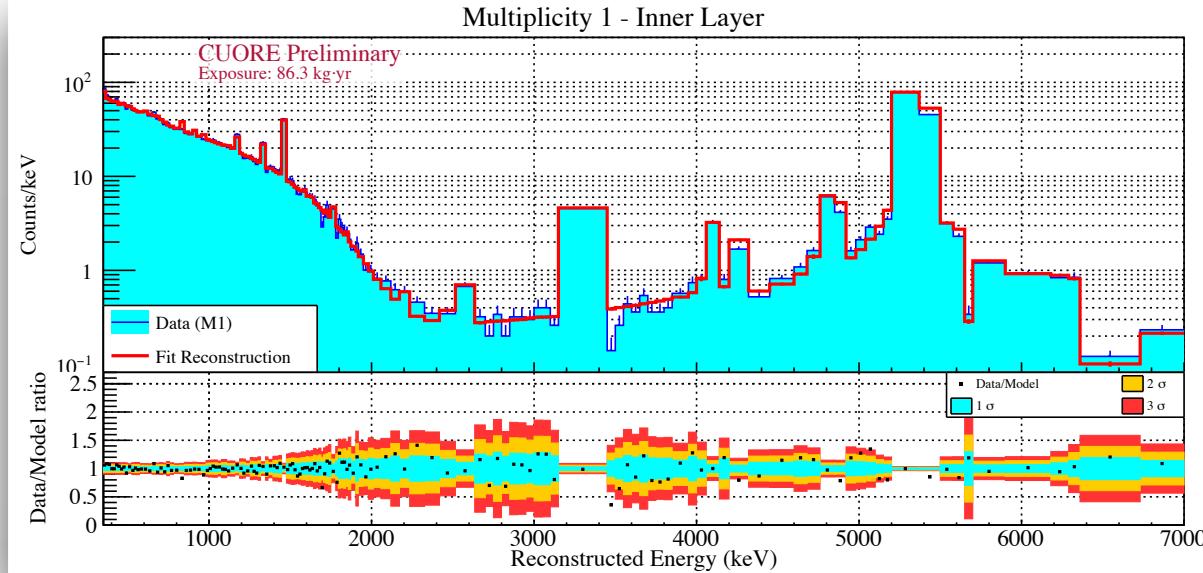
Volume	Type	Components
TeO <sub>2</sub>	Bulk	$2\nu\beta\beta$ , <sup>210</sup> Pb, <sup>232</sup> Th, <sup>228</sup> Ra- <sup>208</sup> Pb, <sup>238</sup> U- <sup>230</sup> Th, <sup>230</sup> Th, <sup>226</sup> Ra- <sup>210</sup> Pb, <sup>40</sup> K, <sup>60</sup> Co, <sup>125</sup> Sb, <sup>190</sup> Pt
TeO <sub>2</sub>	Surface (0.01 μm)	<sup>232</sup> Th, <sup>228</sup> Ra- <sup>208</sup> Pb, <sup>238</sup> U- <sup>230</sup> Th, <sup>226</sup> Ra- <sup>210</sup> Pb, <sup>210</sup> Pb
TeO <sub>2</sub>	Surface (1 μm)	<sup>210</sup> Pb
TeO <sub>2</sub>	Surface (10 μm)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>40</sup> K, <sup>60</sup> Co, <sup>54</sup> Mn
CuNOSV	Surface (0.01 μm)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Surface (1 μm)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
CuNOSV	Surface (10 μm)	<sup>210</sup> Pb, <sup>232</sup> Th, <sup>238</sup> U
Roman lead	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>108m</sup> Ag
Top lead	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>210</sup> Bi
Ext. lead	Bulk	<sup>210</sup> Bi
CuOFE	Bulk	<sup>232</sup> Th, <sup>238</sup> U, <sup>60</sup> Co
External	-	Cosmic muons

## 3. Fit

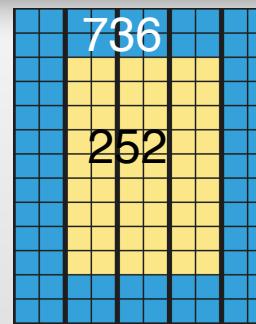
- ▶ perform a Bayesian fit of our model to the data with flat priors on all parameters (except muons which come from a cosmogenic analysis)
- ▶ simultaneous fit on M1, M1inner, M2, M2sum spectra to determine the source/location activities (MC normalisation)



# CUORE background fit



inner and outer layers  
top view



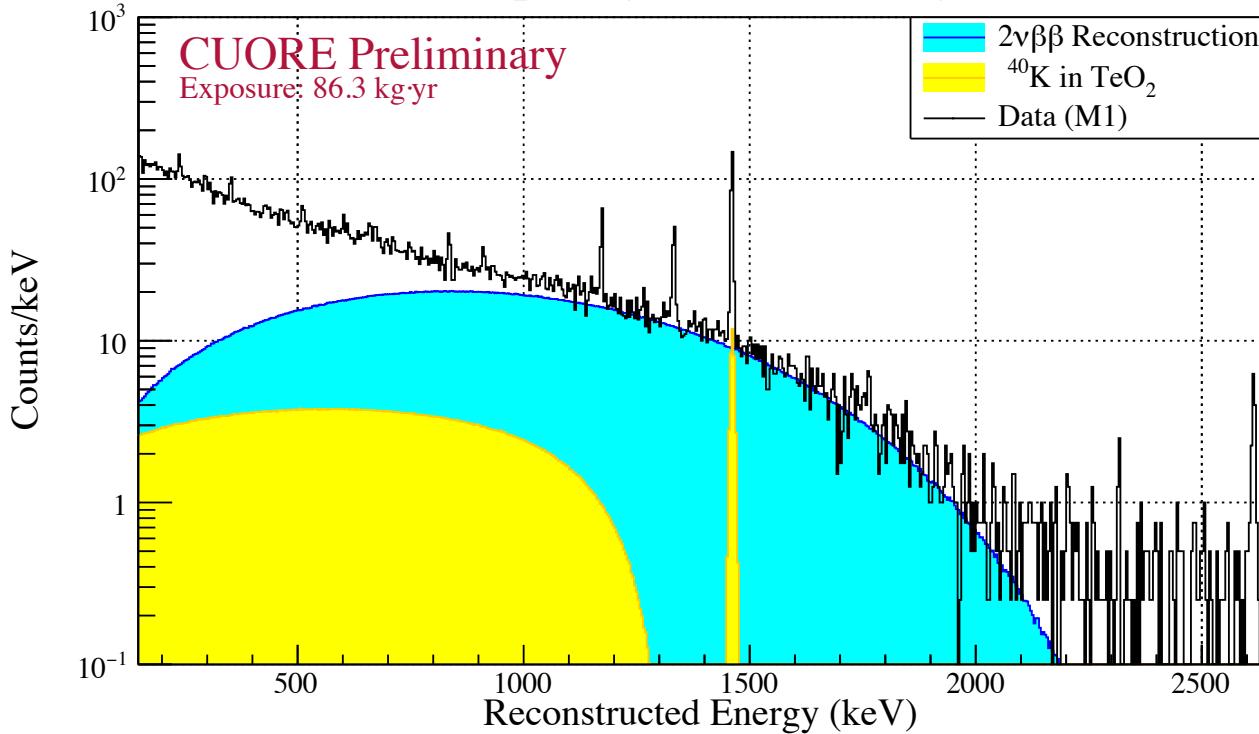
side  
view

**The main features of the CUORE observed spectrum are well reproduced**



# $2\nu\beta\beta$ measurement

Multiplicity 1 -- Inner Layer



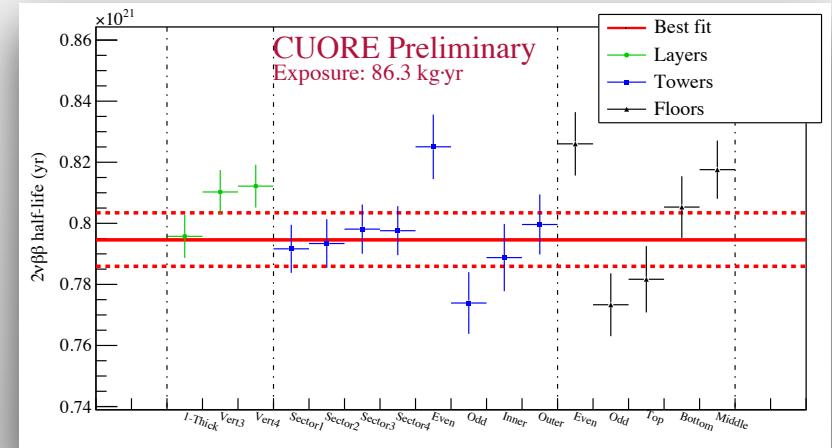
CUORE:  
 $T_{1/2} = [7.9 \pm 0.1(\text{stat.}) \pm 0.2(\text{syst.})] \times 10^{20} \text{ y}$

..consistent with CUORE-0 (2016):

$$T_{1/2} = [8.2 \pm 0.2 \text{ (stat.)} \pm 0.6 \text{ (syst.)}] \times 10^{20} \text{ y}$$

$2\nu\beta\beta$  is the dominant component of the MI background spectrum in the region  $\sim[1-2]$  MeV (compare  $\sim 20\%$  in CUORE-0)

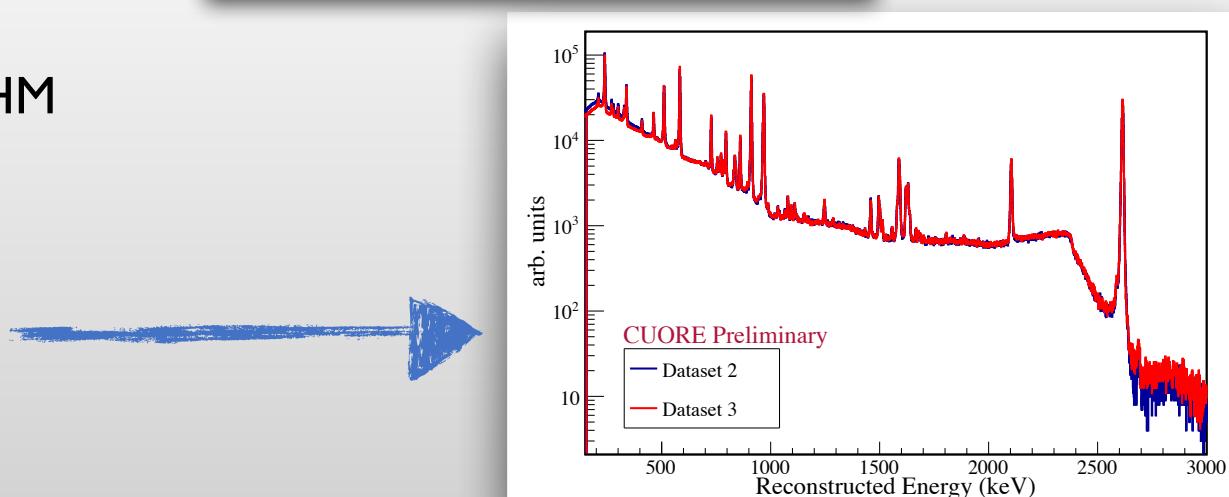
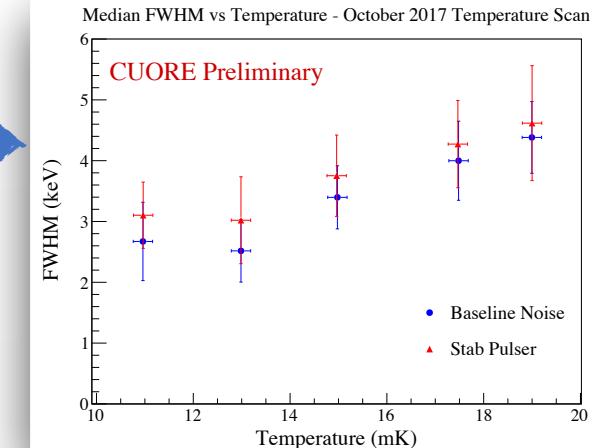
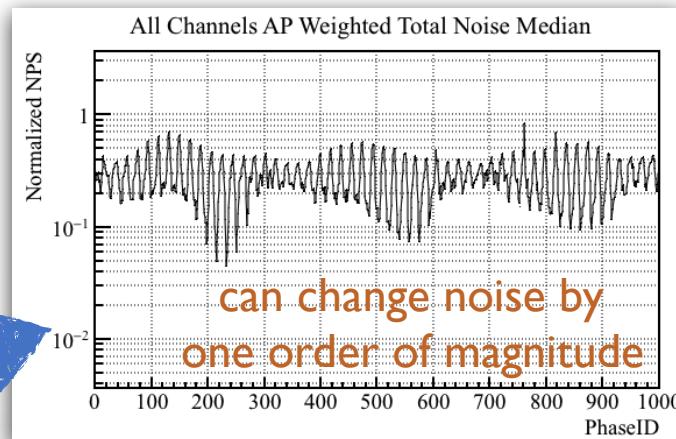
The largest systematic uncertainty due to the geometrical splitting of data (inner/outer) and is originated by the uncertainty on the origin of localised contaminations.





# CUORE current status

- ▶ Oct-Dec 2017: scan of detector performance vs temperature:
  - ▶ new operating temperature at 11 mK
- ▶ Jan-March 2018: warm up to 100 K
  - ▶ upgrade gate valves in DCS
- ▶ March 2018:
  - ▶ return to base temperature (11 mK)
  - ▶ perform Pulse-Tube phase-scan (vary PT phase shifts and select configuration that minimises the average noise)
- ▶ April 2018:
  - ▶ calibration shows  $\Delta E = 7.6$  keV FWHM
- ▶ May 2018:
  - ▶ **stable physics data taking**
  - ▶ **93% channels passing cuts**
  - ▶ **doubled+ 2017 data**
  - ▶ **analysis ongoing**





# Conclusions

- ▶ CUORE is the first  $0\nu\beta\beta$  cryogenic experiment at ton-scale
- ▶ Most stringent limit on the  $0\nu\beta\beta$  half-life of  $^{130}\text{Te}$  to date
- ▶ Most precise measurement of the  $2\nu\beta\beta$  half-life of  $^{130}\text{Te}$  to date
- ▶ We have restarted physics data taking
  - working to optimise background and energy resolution
- ▶ CUORE will continue to be one of the most sensitive searches for  $0\nu\beta\beta$  over the coming years
- ▶ Many more potential physics searches:
  - low energy searches: Dark Matter, axions
  - other  $\beta\beta$  decays and decays to excited states,  $\beta^+/\text{E.C.}$  decays

**THANK YOU !**