Production methods for precision experiments with protonium



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Quest:

Tests of fundamental symmetries in comparative studies of neutral matter and antimatter.

- Antihydrogen... spectroscopy: CPT symmetry test on 1S-2S of 2·10⁻¹² [1] probing of the WEP of antimatter: Indirectly known to hold at 10⁻⁷ g Limit on neutral *matter:* relative 10⁻⁹ δg/g [2]; absolute 10⁻⁵ δg/g [3]; WEP 10⁻¹⁴ [4]
- **Protonium**... measure g, WEP in purely baryonic system, spectroscopy to probe the residual strong force on threshold
- Antiprotonic deuterium...

isospin dependence on g, spectroscopy

• Muonic antihydrogen...

spectroscopy, Lamb shift due to finite size of \overline{p} Proton-charge-radius-puzzle observed in μp [5]

- [1] M. Ahmadi et al., Nature 557, 71 (2018)
- [2] H. Mueller et al., Nature 463, 926 (2010)
- [3] R. F. C. Vessot et al., *Phys. Rev. Lett.* **45**, 2081 (1980)
- [4] S. Baessler et al., Phys. Rev. Lett. 83, 3583 (1999)
- [5] A. Antognini et al., Science **339**, 417 (2013), r_{E} =0.84087(39) fm, 7 σ different than 2010CODATA

Orders of magnitude

Coulomb bound system $V_{coulomb} = - Ze^2/r$

Binding energies $B_n = -(m_r c^2 \alpha^2) (Z^2/2n^2)$

Radii

$$a_n = (hc/m_r c^2 \alpha) (n^2/Z)$$

| dimension | | m [MeV/c²] | B _{n=1} [keV] | r _{Bohr} [fm] |
|----------------|----------------------|---------------|---------------------------|---------------------------|
| 'atomic' dim. | <mark>e</mark> ⁻p | 0.511 | 0.0136 | 0.5·10 ⁵ |
| 1 | μ⁻p | 105 | 2.5 | 279 |
| | π⁻p | 140 | 3.2 | 216 |
| \checkmark | <mark>₽</mark> p | 938 | 12.5 | 57 |
| 'nuclear' dim. | <r<sub>p></r<sub> | | | 0.8 |
| | | | | |



Pn properties

Pn energies & length scales to H:







E. Klempt et al., Physics Rep. 363, 116 (2002)

Pn binding energy, hadronic shift of levels



pΗ

meV



D. Gotta et al., Nucl. Phys. A 660, 289 (1999)

Pn properties

Pn energies & length scales to H:







Reid potential corresponding to residual strong force



E. Klempt et al., Physics Rep. 363, 116 (2002)

Cold Pn production via H⁻ in a Penning trap

Spectroscopy and g-measurement benefit from colder Pn

Pn 'traditionally' produced in eV range by injecting \overline{p} into liquid H

 \rightarrow Penning trap can produce Pn with < 0.1 eV via resonant-charge-exchange

p/H/H⁻ source

e.g. >1 μA H⁻ RF cw source: T. Kalvas et al., AIP Conf. **1655**, 030015 (2015)



M. Hori & J. Walz, Prog. of Partic. Nucl. Phys. 72, 206 (2013)

Penning - Malmberg trap

10 keV \bar{p} are trapped and electron-cooled to < 0.1 eV

R. Brusa et al. (AEgIS), J. Phys.: Conf Ser. 791, 012014 (2017)
M. Ahmadi et al. (ALPHA), Nature Com. 8, 681 (2017)
C. Amole et al. (ASACUSA), Nature Com. 5, 3955 (2014)
G. Gabrielse et al. (ATRAP), Phys. Rev. Lett. 108, 113002 (2012)



Pn production schemes via H⁻



Challenge: Direct H sources can be broad in energy (1 to 2 eV)

Pn production schemes via H^- , H_2



[1] ~100 Pn@<1 eV N. Zurlo et al. (ATHENA), Hyp. Int. 172, 97-105 (2006)

Two-step Pn production schemes via H



Formation cross section of Pn via H



Pn production in Rydberg states... two benefits



Orbital momentum L

Pn production in Rydberg states... two benefits





J. S. Cohen and N. T. Padial, Phys. Rev. A 41, 7 (1990)





 $<\tau_{Pn30(from H1s)}>$ ~ 2 μs

Pn in state 30s, by J. Baez (2015)

R. S. Hayano, Nucl. Phys. A655, 318 (1999)

$H^{-}(1s) \rightarrow H(Rydberg)$

What is the best way to excite H⁻ to Rydberg H ?

 $H^{-} + \gamma \rightarrow H + e^{-}$

 $H + \overline{p} \rightarrow Pn + e^{-}$

One-photon options:

1) 113 nm (photodetachment on Feshbach resonance) + Rydberg laser

2) 1652 nm photodet. on threshold + 121 nm + Rydberg

3) 1064 nm photodetachment + 121 nm + Rydberg laser

Two-photon options:

- 4) 224 nm (on Feshbach res.) + Rydberg
- 5) 1652 nm + 242 nm + Rydberg



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H. C. Bryant et al., Phys. Rev. Let. 38, 5 (1977)

One-photon options:

1064 nm photodetatchment + 121 nm + 366 nm Rydberg laser



H⁻(1s) – H(1s) – H(2p) – H(30d)

Cross sections are known for H:

- σ bound-bound
- σ bound-continuum
- $\sigma_{\text{H- to H}} \!= 4 \!\cdot\! 10^{-17} \text{ cm}^2$

5-level optical rate equations using parameters: $w_{lasers} = 1 \text{ cm}^2$,

t [ns]

 $P_1 = 9 \text{ mJ/1 ns}, P_2 = 20 \mu \text{J/1 ns}, P_3 = 11 \text{ mJ/1 ns}$



Fermi's Golden Rule: A. Burgees, Astrophys. J. 141, 1588 (1965)

$H^{-}(1s) - H(2s) - H(10p)$

Two-photon options with SCRAP using LICS [1]: 224 nm (Pump) + 462 nm (Stokes) + 379 nm Rydberg laser



[1] Similar scheme as in e.g.: L. P. Yatsenko et al. Phys. Rev. A 71, 033418 (2005)

H⁻(1s) – H(2s) – H(n10p)



Calculation for field free solution:

S.... Stark shifts Ω ... Rabi frequencies Γ ...ionization rates

The ionization rates are for field free (B=0) Estimated with loss-free Feshbach transition

Table 2.1 lists the values for the ionization rates and Stark shifts in units of 1/ns for intensities in GW/cm^{-2} for the relevant levels of Fig.3:

| $\Gamma_{2,P}(t) = 2.91 \ I_P(t)$ | $\Gamma_{2,S}(t) = 0$ | $\Gamma_{2,R}(t) = 0$ |
|-----------------------------------|---|-----------------------------------|
| $\Gamma_{3,P}(t) = 0.18 I_P(t)$ | $\Gamma_{3,S}(t) = 1.98 \ I_S(t)$ | $\Gamma_{3,R}(t) = 1.04 I_R(t)$ |
| $\Gamma_{4,P}(t) = 1.29 I_P(t)$ | $\Gamma_{4,S}(t) = 0.14 I_S(t)$ | $\Gamma_{4,R}(t) = 1.58 I_R(t)$ |
| $S_{1,P}(t) = -1.89 I_P(t)$ | $S_{1,S}(t) = -1.35 I_S(t)$ | $S_{1,R}(t) = -1.19 I_R(t)$ |
| $S_{2,P}(t) = 4.78 I_P(t)$ | $S_{2,S}(t) = 19.89 I_S(t)$ | $S_{2,R}(t) = 23.32 I_R(t)$ |
| $S_{3,P}(t) = 3.89 I_P(t)$ | $S_{3,S}(t) = 12.31 I_S(t)$ | $S_{3,R}(t) = 7.84 I_R(t)$ |
| $S_{4,P}(t) = 1352 I_P(t)$ | $S_{4,S}(t) = 5789 \ I_S(t)$ | $S_{4,R}(t) = 3875 I_R(t)$ |
| $\Omega_P(t) = 4.63 \ I_P(t)$ | $\Omega_{SP}(t) = 20.28\sqrt{I_P(t)I_S(t)}$ | $\Omega_R(t) = 1442\sqrt{I_R(t)}$ |
| q=0.5 | | |

Optimization for the laser parameters, table 2.1 shows the intensities in GW/cm^2





Transfer efficiency vs laser intensities



Pn production in a Penning trap

MCMC approach

Start with H^{-}/\bar{p} plasma in trap at thermal equilibrium (density n, temperature T). Interaction with lasers at t=0



Created number of H*

Pn production via resonant charge-exchange. Pn emission pattern determined by B, n, T



Pn formation – contour scan for Rydberg H(10p)



Outlook for gravity measurement of Pn

Idea:

g-measurement in an axial rotatable Talbot-Lau atom interferometer



de Broglie wavelength
$$\lambda = \frac{h}{m v}$$

 $\alpha = 0^{\circ} - 30^{\circ}$ Talbot-Lau layout (is sensitive to 0 and 0.5 g):



R. Colella et al., *Phys. Rev. Lett.* **34**, 1472 (1975) M. Oberthaler, *Heidelberg*

H. Mueller et al., *Nature*, **463**, 926 (2010)S. Baessler et al., *Phys. Rev. Lett.*, **83**, 3583 (1999)

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Thank you very much for your attention!