

# Recent progress on extended Skyrme functionals (and it works!)

A. Pastore<sup>1</sup>, P. Becker<sup>1</sup>, D. Davesne<sup>2</sup>, J. Navarro<sup>3</sup>

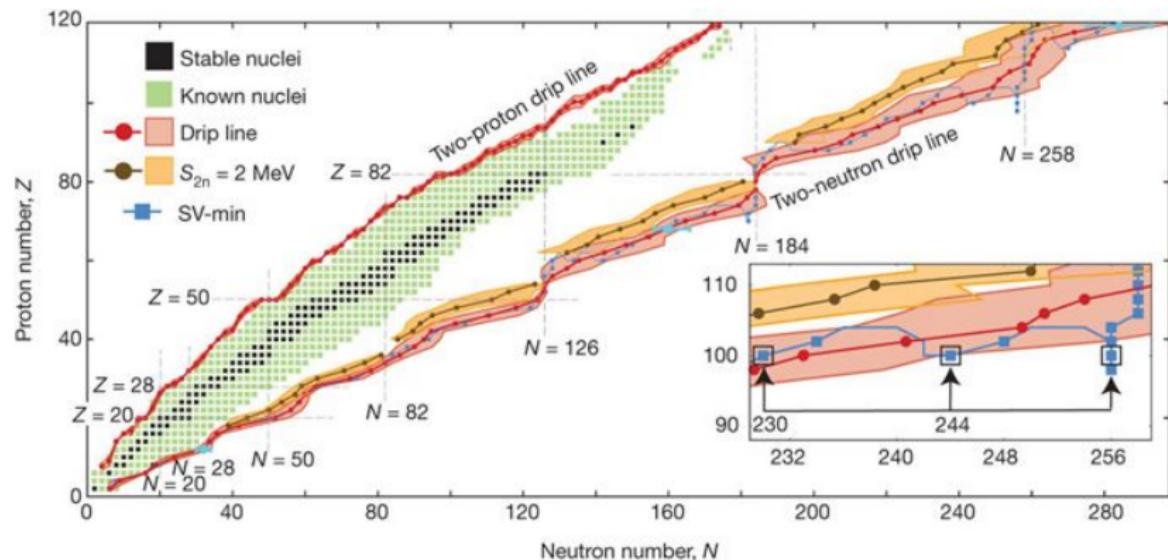
<sup>1</sup> University of York

<sup>2</sup>Université de Lyon, F-69003 Lyon, France

<sup>3</sup>IFIC (CSIC University of Valencia), Apdo. Postal 22085, E-46071 Valencia

September 6, 2018

# Introduction: the nuclear chart



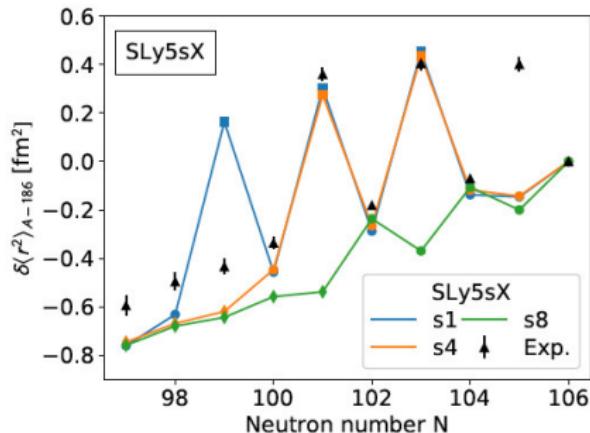
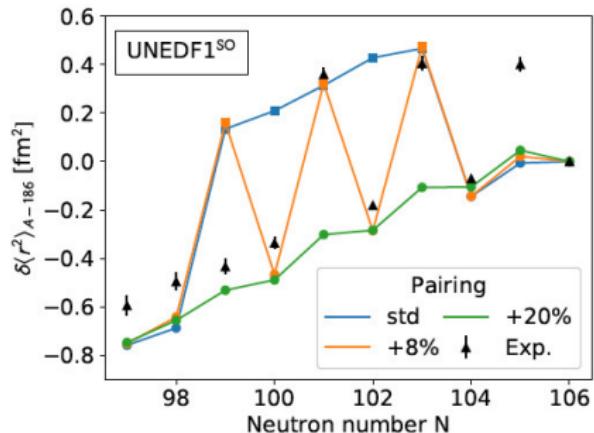
## Nuclear chart

Atomic nucleus is the best example for *strange quantum phenomena*

[J. Erler et al. (2012); Nature, 486(7404), 509.]

# A simple model: Skyrme Nuclear-DFT

Global description of nuclear properties with  $\approx 10$  parameters.



## Odd-even staggering in Hg nuclei

A sudden change in the potential energy surface (oblate/prolate)

[B. Marsh, . . . , J. Dobaczewski, AP et al.; Nature Physics (accepted)]

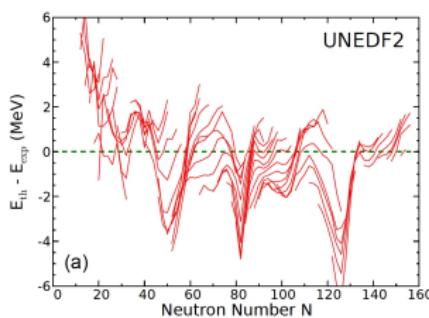
[S. Sels, . . . , J. Dobaczewski, AP et al. PRC (in preparation)]

# How to determine the coupling constants?

We impose a fitting protocol (observables and pseudo-observables)

- IM properties (*i.e.*  $E/A, K_\infty, m^*, \dots$ )
- Ground state of some nuclei (*i.e.*  $^{40}\text{Ca}, ^{48}\text{Ca}, ^{208}\text{Pb}, \dots$ )
- Charge radii
- Spin orbit splitting
- ...

[M . Kortelainen et al. Phys. Rev. C89, 054314 (2014) ]



Observable	UNEDF0	UNEDF1	UNEDF2	No.
$E$	1.428	1.912	1.950	555
$E (A < 80)$	2.092	2.566	2.475	113
$E (A \geq 80)$	1.200	1.705	1.792	442
$S_{2n}$	0.758	0.752	0.843	500
$S_{2n} (A < 80)$	1.447	1.161	1.243	99
$S_{2n} (A \geq 80)$	0.446	0.609	0.711	401
$S_{2p}$	0.862	0.791	0.778	477
$S_{2p} (A < 80)$	1.496	1.264	1.309	96
$S_{2p} (A \geq 80)$	0.605	0.618	0.572	381
$\tilde{\Delta}_n^{(3)}$	0.355	0.358	0.285	442
$\tilde{\Delta}_n^{(3)} (A < 80)$	0.401	0.388	0.327	89
$\tilde{\Delta}_n^{(3)} (A \geq 80)$	0.342	0.350	0.273	353
$\tilde{\Delta}_p^{(3)}$	0.258	0.261	0.276	395
$\tilde{\Delta}_p^{(3)} (A < 80)$	0.346	0.304	0.472	83
$\tilde{\Delta}_p^{(3)} (A \geq 80)$	0.229	0.248	0.194	312
$R_p$	0.017	0.017	0.018	49
$R_p (A < 80)$	0.022	0.019	0.020	16
$R_p (A \geq 80)$	0.013	0.015	0.017	33

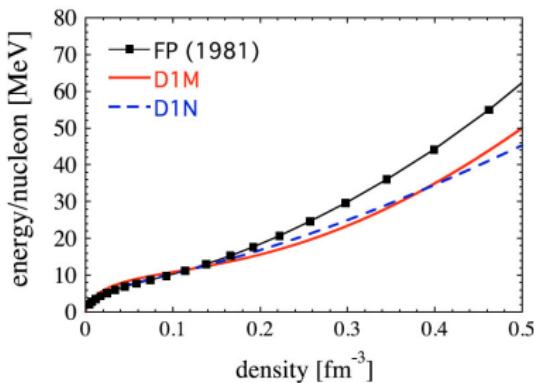
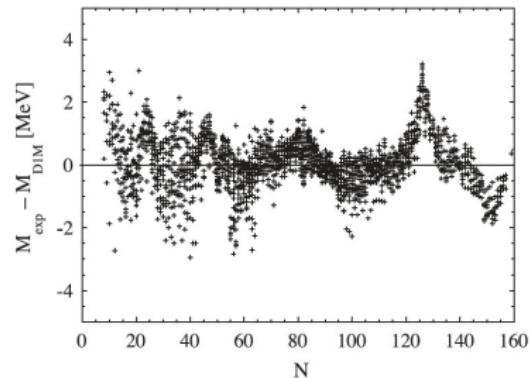
## Result

Skyrme functional can not further improved: need to go *beyond*. How?

# Option 1: multi-reference calculations

Change many-body method!

Gogny D1M → parameters adjusted using 5D collective Hamiltonian



[S. Goriely et al. Physical Review letters, 102(24), 24250 ]

## Result

Major improvement on masses  $\sigma_{\text{mass}} \approx 0.8$  MeV.... but still not perfect!

## Option 2... back to the Skyrme idea...

$$t_{12} = \delta(\mathbf{r}_1 - \mathbf{r}_2) t(\mathbf{k}', \mathbf{k}) \quad (3)$$

where  $\mathbf{k}$  is the operator corresponding to the relative wave-number,

$$\mathbf{k} = \frac{1}{2}i(\nabla_1 - \nabla_2); \quad (4)$$

placed on the *right* of the delta-function  $\mathbf{k}'$  denotes the same operator placed on the *left*; this form was used in an earlier discussion<sup>8)</sup> of the spin-orbit potential. An ordinary static (Wigner) interaction would then be described by a  $t$  dependent only on  $\mathbf{k}' - \mathbf{k}$ .

It is generally believed that the most important part of the two-body interaction can be represented by a contact potential, i.e. by constant  $t(\mathbf{k}', \mathbf{k})$ ; this suggests an expansion in powers of  $\mathbf{k}'$  and  $\mathbf{k}$ . If this expansion

### Why 2nd order only?

We can remove the approximation of Skyrme and make the expansion beyond 2nd order

# Taylor expansion of Gogny interaction

We can make a Taylor expansion of a finite range interaction

$$v_G^c(\mathbf{k}, \mathbf{k}') = \sum_{n=1}^2 \left[ W^n + B^{(n)} P_\sigma - H^{(n)} P_\tau - M^{(n)} P_\sigma P_\tau \right] \pi^{3/2} (\mu_n^{c,G})^3 e^{-(\mathbf{k}-\mathbf{k}')^2/4(\mu_n^{c,G})^2},$$

Mapping to N3LO by a formal Taylor expansion of a scalar function  $F(\mathbf{k} - \mathbf{k}')$

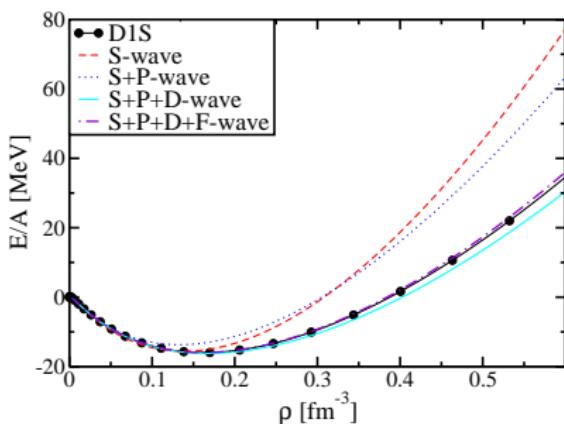
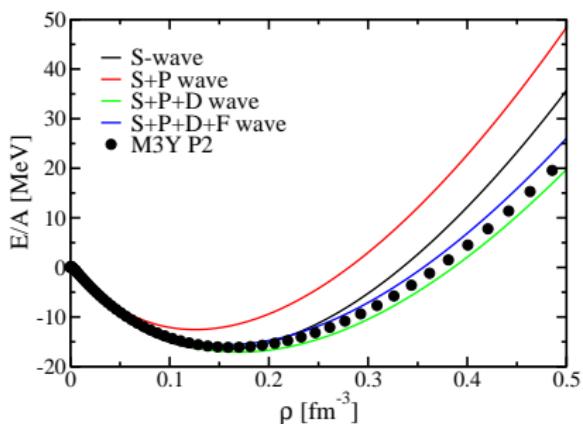
$$\begin{aligned} F(\mathbf{k} - \mathbf{k}') &= C_0 + C_2(\mathbf{k} - \mathbf{k}')^2 + C_4(\mathbf{k} - \mathbf{k}')^4 + \dots \\ &= C_0 + C_2 \left[ \mathbf{k}^2 + \mathbf{k}'^2 - 2\mathbf{k}' \cdot \mathbf{k} \right] \\ &\quad + C_4 \left[ (\mathbf{k}^2 + \mathbf{k}'^2)^2 + 4(\mathbf{k}' \cdot \mathbf{k})^2 - 4(\mathbf{k}' \cdot \mathbf{k})(\mathbf{k}^2 + \mathbf{k}'^2) \right] + \dots \end{aligned}$$

Skyrme N2LO: Skyrme has more freedom than Gogny!!!

$$\begin{aligned} V_{\text{N2LO}}^c &= t_0(1 + x_0 P_\sigma) + \frac{1}{2} t_1(1 + x_1 P_\sigma)(\mathbf{k}^2 + \mathbf{k}'^2) + t_2(1 + x_2 P_\sigma)(\mathbf{k} \cdot \mathbf{k}') \\ &\quad + \frac{1}{4} t_1^{(4)}(1 + x_1^{(4)} P_\sigma) \left[ (\mathbf{k}^2 + \mathbf{k}'^2)^2 + 4(\mathbf{k}' \cdot \mathbf{k})^2 \right] \\ &\quad + t_2^{(4)}(1 + x_2^{(4)} P_\sigma)(\mathbf{k}' \cdot \mathbf{k})(\mathbf{k}^2 + \mathbf{k}'^2) \end{aligned}$$

# Partial wave expansion

A finite range interaction can be decomposed into an infinite sum of partial waves

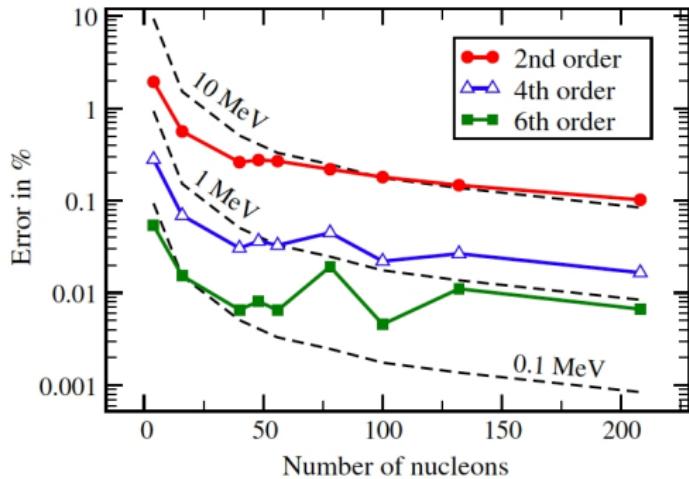


Fast convergence!

Including D-wave we reproduce the same physics at saturation

# Density Matrix Expansion (DME)

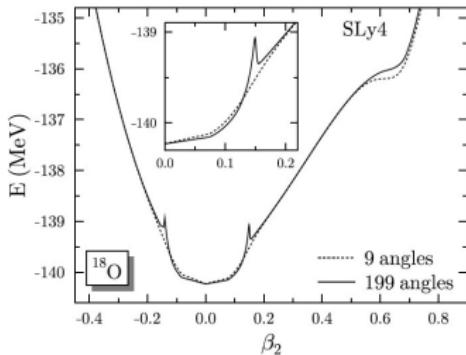
[G. Carlsson and J. Dobaczewski; Phys. Rev. Lett. (2013) ]



## Results

DME converges quite well at 4th order!!

# Role of density-dependent term



[ D. Lacroix et al; Physical Review C 79.4 (2009): 044318.]

- Density dependent (DD) term are used in *all* current pseudo-potential
- Multi-reference is OK for interactions with no DD terms:
  - ① SV → 2-body force ( $m^* = 0.4$ )
  - ② SLYMR0 → 2- and 3-body force [ J. Sadoudi et al. Phys. Rev. C 88.6 (2013): 064326.]

## Important result

To have  $m^* \approx 0.7$  (or more) we need a 3-body term (or density dependent one!)

[ D. Lacroix et al; Physical Review C 79.4 (2009): 044318.]

# The N $\ell$ LO Skyrme interaction

$$\begin{aligned} \mathcal{V}_{Ex}(\mathbf{r}_1, \mathbf{r}_2) = & t_0 (1 + x_0 P_\sigma) + \frac{1}{6} t_3 (1 + x_3 P_\sigma) \rho^\alpha(R) \\ & + \frac{1}{2} t_1 (1 + x_1 P_\sigma) \left[ \mathbf{k}^2 + \mathbf{k}'^2 \right] + t_2 (1 + x_2 P_\sigma) \mathbf{k}' \cdot \mathbf{k} \\ & + \frac{1}{4} t_1^{(4)} (1 + x_1^{(4)} P_\sigma) \left[ (\mathbf{k}^2 + \mathbf{k}'^2)^2 + 4(\mathbf{k}' \cdot \mathbf{k})^2 \right] \\ & + t_2^{(4)} (1 + x_2^{(4)} P_\sigma) (\mathbf{k}' \cdot \mathbf{k})(\mathbf{k}^2 + \mathbf{k}'^2) \\ & + \frac{1}{2} t_1^{(6)} (1 + x_1^{(6)} P_\sigma) (\mathbf{k}'^2 + \mathbf{k}^2) \left[ (\mathbf{k}'^2 + \mathbf{k}^2)^2 + 12(\mathbf{k}' \cdot \mathbf{k})^2 \right] \\ & + t_2^{(6)} (1 + x_2^{(6)} P_\sigma) (\mathbf{k}' \cdot \mathbf{k}) \left[ 3(\mathbf{k}'^2 + \mathbf{k}^2)^2 + 4(\mathbf{k}' \cdot \mathbf{k})^2 \right]. \end{aligned} \quad \left. \begin{array}{l} \text{Skyrme NLO} \\ \text{Skyrme N2LO} \\ \text{Skyrme N3LO} \end{array} \right\}$$

D and F partial waves included

Gauge invariant

18 parameters,  
8 new

Also includes

- a spin-orbit term  $W_0$
- tensor terms (4 new)

# N2LO in finite nuclei

## A 4th order equation in spherical symmetry

$$\epsilon R = A_4 R^{(4)} + A_3 R^{(3)} + \color{blue}{A_{2R} R^{(2)}} + A_{1R} R' + \color{blue}{A_{0R} R} \\ + \frac{\ell(\ell+1)}{r^2} \left[ A_{2C} R^{(2)} + A_{1C} R' + \color{blue}{A_{0C} R} + \frac{\ell(\ell+1)}{r^2} \color{red}{A_{0CC} R} \right] \\ + C_{jls} \left[ W_{2R} R^{(2)} + W_{1R} R' + \color{blue}{W_{0R} R} + \frac{\ell(\ell+1)}{r^2} \color{red}{W_{0C} R} \right]$$

14 parameters,  
4 new

Doubly centrifugal term

$$\left( \frac{\ell(\ell+1)}{r^2} \right)^2$$

New spin-orbit  
contributions

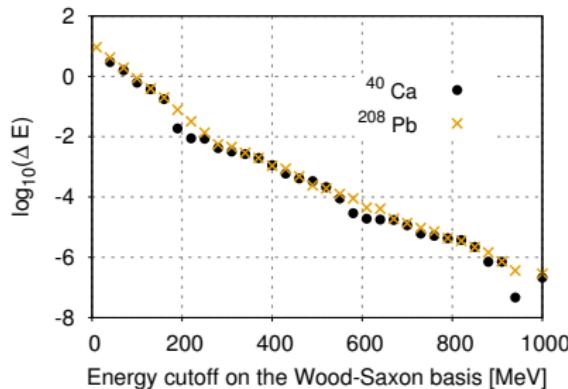
$$W_{2R}, \quad W_{1R}, \quad \frac{\ell(\ell+1)}{r^2} W_{0C}$$

Same behaviour at the  
origin than Skyrme NLO

Numerical solver in coordinate space: fast and accurate!

[ P .Becker et al.; Physical Review C 96.4 (2017): 044330.]

# WHISKY: 2 basis methods HFB code



<sup>208</sup> Pb		
[MeV]	WHISKY	LENTEUR
Total Energy	-1636.106	-1636.105
Kinetic energy	3874.789	3874.795
Field energy	-6209.642	-6209.650
Spin-orbit	-99.081	-99.081
Direct Coulomb	829.143	829.143
Exchange Coulomb	-31.314	-31.314

<sup>208</sup> Pb		
[MeV]	WHISKY	MOCCA
Total Energy	-1539.253	-1539.263
Total energy N2LO	89.278	89.360
$E[(\Delta\rho)^2]$	4.394	4.395
$E[\rho Q]$	37.477	37.488
$E[\tau^2]$	27.212	27.221
$E[\tau_{\mu\nu}\tau_{\mu\nu} - \tau_{\mu\nu}\nabla_\mu\nabla_\nu\rho]$	19.855	19.861
$E[K_{\mu\nu\kappa}K_{\mu\nu\kappa}]$	0.05460	0.05461
$E[J_{\mu\nu}V_{\mu\nu}]$	0.33850	0.33858

# Fitting protocol

Infinite nuclear matter			
$\rho_{sat}$	0.1600	0.001	$\text{fm}^{-3}$
E/A ( $\rho_{sat}$ )	-16.0000	0.2	MeV
$m^*/m$	0.7000	0.02	
$K_\infty$	230.00	10.00	MeV
$J$	32.00	2.00	MeV
EoS PNM			
E/N ( $\rho=0.1$ )	11.88	2.0	MeV
E/N ( $\rho=0.3$ )	35.94	7.0	MeV
E/N ( $\rho=0.35$ )	44.14	9.0	MeV
Stability			
INM(S,M,T)	$\rho_{crit} \geq 0.24$	asymmetric constraint	$\text{fm}^{-3}$
Finite nuclei			
Binding energies			
$^{40}\text{Ca}$	-342.02300	1.5	MeV
$^{48}\text{Ca}$	-415.98300	1.0	MeV
$^{56}\text{Ni}$	-483.95300	1.5	MeV
$^{100}\text{Sn}$	-825.13000	1.5	MeV
$^{132}\text{Sn}$	-1102.67300	1.0	MeV
$^{208}\text{Pb}$	-1635.86100	1.0	MeV
Proton radii			
$^{40}\text{Ca}$	3.38282	0.03	fm
$^{48}\text{Ca}$	3.39070	0.02	fm
$^{56}\text{Ni}$	3.66189	0.03	fm
$^{132}\text{Sn}$	4.64745	0.02	fm
$^{208}\text{Pb}$	5.45007	0.02	fm
Parameter $W_0$			
	120.0	2.0	$\text{MeV fm}^5$

# Introducing SN2LO1

## Valencia-Lyon-York fitting protocol

- Double magic nuclei
- Infinite matter properties
- Finite-size instabilities
- $W_0$  fixed (for the moment)

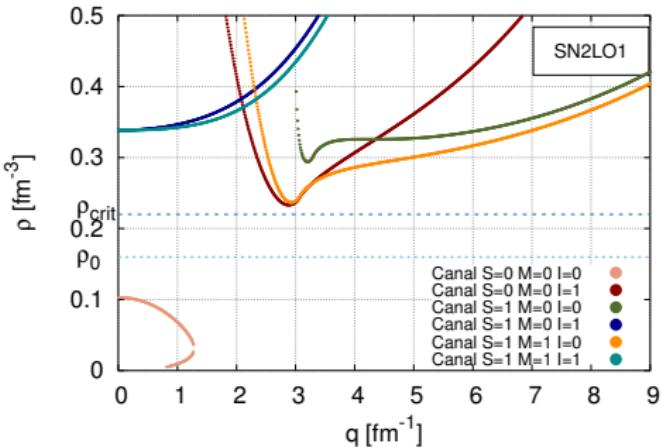
Natural Units			
$C_0^\rho$	-1.06	$C_1^\rho$	0.754
$C_0^\rho[\rho^\alpha]$	13.0	$C_1^\rho[\rho^\alpha]$	-12.1
$C_0^T$	0.892	$C_1^T$	0.00624
$C_0^{\Delta\rho}$	-1.06	$C_1^{\Delta\rho}$	0.382
$C_0^{\nabla J}$	-1.22	$C_1^{\nabla J}$	-0.406
$C_0^T$	-0.0882	$C_1^T$	-0.816
$C_0^{(\Delta\rho)^2}$	-0.115	$C_1^{(\Delta\rho)^2}$	0.0396
$C_0^{M\rho}$	-0.288	$C_1^{M\rho}$	0.143
$C_0^{Ms}$	0.117	$C_1^{Ms}$	-0.0162

## Important remarks

- The coupling constants look *natural* (Grain of salt!!)
- We need further studies on naturalness

[Kortelainen, M., Furnstahl, R. J., Nazarewicz, W., Stoitsov, M. V. (2010). Phys. Rev. C, 82(1), 011304.

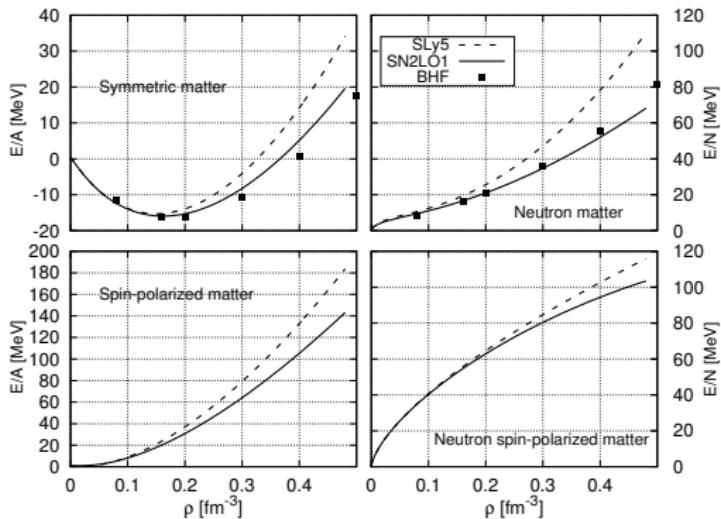
# Finite-size instabilities



## An empirical solution

- Use Linear Response to detect instabilities [Lesinski, T., et al. ; Phys. Rev. C 74.4 (2006): 044315. ] [PA et al. Physics reports 563 (2015): 1-67.] [ De Pace, A., and M. Martini; Physical Review C 94.2 (2016): 024342.]
- Impose  $\rho_{pole} > 1.2 \rho_{sat}$

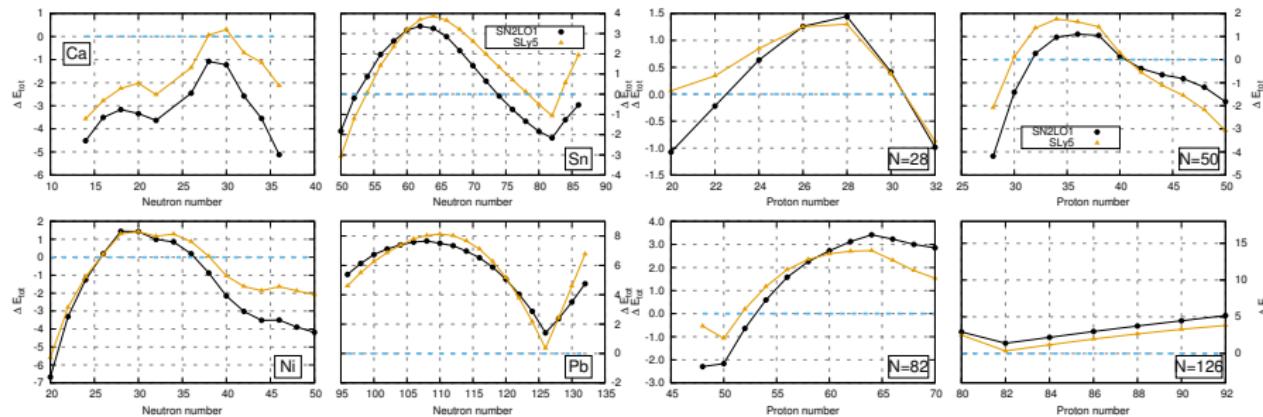
# Infinite nuclear matter properties



	SN2LO1	SLy5
$\rho_0$ [ $\text{fm}^{-3}$ ]	0.162	0.1603
$E/A(\rho_0)$ [MeV]	-15.948	-15.98
$K_\infty$ [MeV]	221.9	229.92
$J$ [MeV]	31.95	32.03
$L$ [MeV]	48.9	48.15
$m^*/m$	0.709	0.696

# Masses

Systematic comparison of binding energies for isotopic (isotonic) chains calculated with SN2LO1 and SLy5 parametrisation



Only even-even nuclei!

SN2LO1 gives slightly worst results than SLy5.

# Error analysis

	SN2LO1	SN2LO2	SLy5*	Last fit	%	
INM Properties	1.40	0.85	0.07	1.17		
Neutron matter	0.19	0.21	0.05	0.2		
Stability	5.04	6.72	5.32	3.09		
Nuclei	236.82	57.44	64.20	18.36	- 71 %	✓
Proton radii	8.81	13.44	17.31	8.57	- 51 %	✓
$W_0$ Parameter	1.10	0.66	0.02	1.22		
Total $\chi^2$	8.77	2.77	3.03	1.15	- 62 %	✓

What new constraints ?

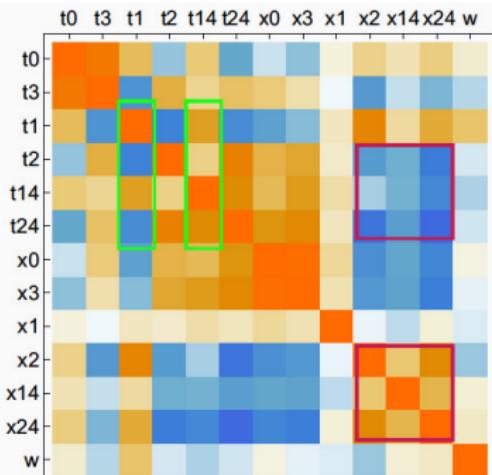
- Nuclear channels
- Landau parameters
- s.p. levels
- New nuclei, maybe with pairing or deformation

## Results

- *Special* observables for higher order terms
- New statistical tools  
(Bootstrap/Gaussian Process Emulator)
- New interesting terms  
 $C_1 l(l+1) + C_2 l^2(l+1)^2 \rightarrow$   
Candidate to change position  
high  $l$  states

## Covariance matrix

- Different correlation structure
- Net terms NOT compatible with 0



## Work in progress

- 1.00 Errors :
- $t0 = 2477.48 \pm 5.88$
  - $t3 = 13660.68 \pm 54.95$
  - $t1 = 507.81 \pm 6.64$
  - $t2 = -421.48 \pm 23.07$
  - $t14 = -29.68 \pm 4.67$
  - $t24 = 4.69 \pm 0.63$
  - $x0 = 0.77 \pm 0.01$
  - $x3 = 1.17 \pm 0.02$
  - $x1 = 0.0014 \pm 0.0035$
  - $x2 = -0.92 \pm 0.01$
  - $x14 = 0.353 \pm 0.006$
  - $x24 = 2.14 \pm 0.06$
  - $w = 122.30 \pm 6.72$

# Summary and Conclusions

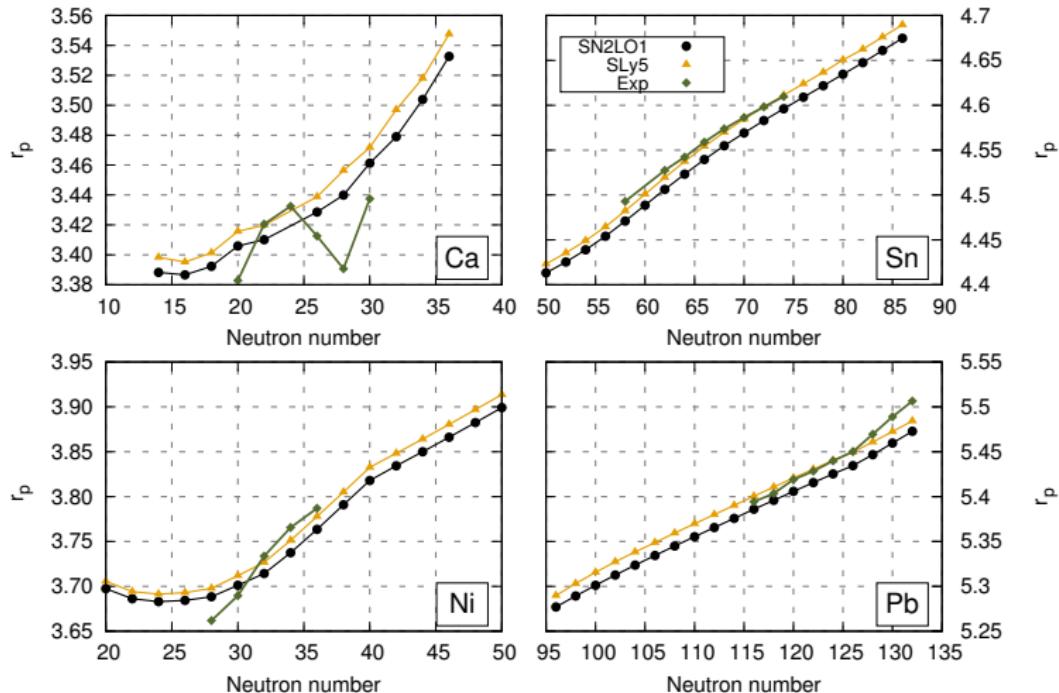
- We have produced the first N2LO parametrisation
- Very difficult fit higher order:
  - use of RPA in IM to avoid instabilities
  - need *ad-hoc* observables
  - understand the role of new fields/densities
  - N3LO is also ready spherical/deformed
- New numerical code

The next step...

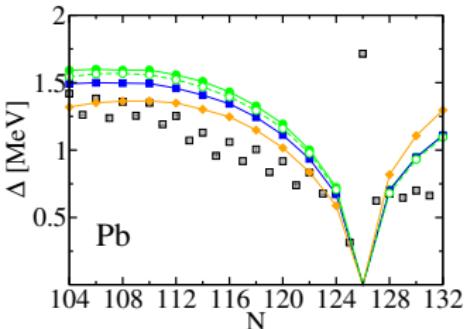
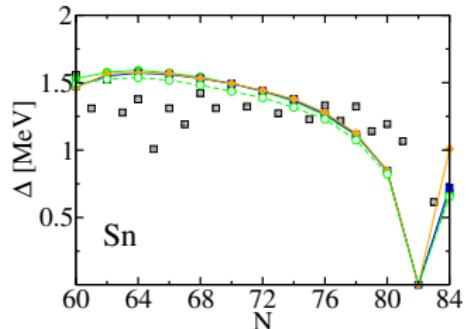
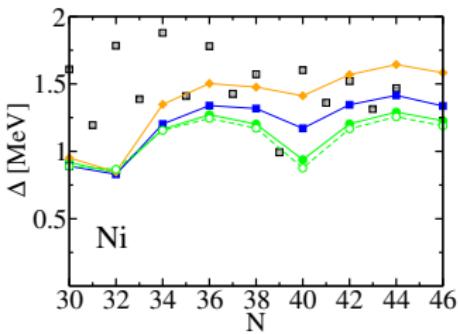
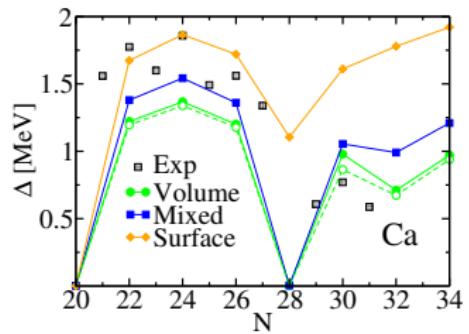
- Need new fitting protocol beyond modified Saclay-Lyon → ideas for  $\chi^2$ ???
- Improving super-heavy nuclei spectroscopy
- Exploring excited states (joint PhD student York/Lyon/CEA)

THANK YOU

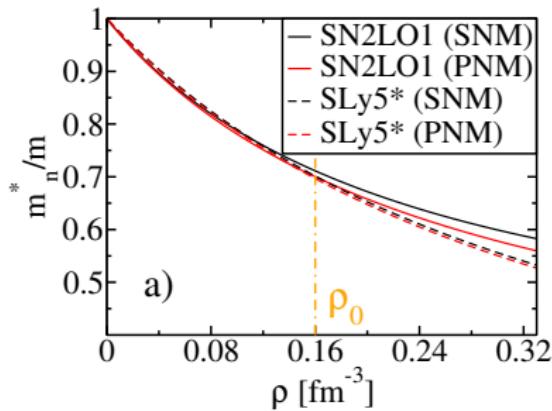
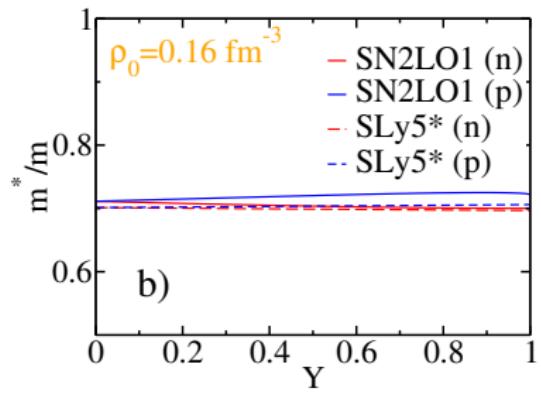
# Charge radii



# Pairing gaps



# Effective mass



Better to compare with SLy5\*

SN2LO1 → **same** fitting protocol SLy5\* (and same results!)

[Pastore, A., Davesne, D., Bennaceur, K., Meyer, J., and Hellemans, V. (2013); Phys. Script., 2013(T154), 014014. ]