Open Questions in Fundamental Physics

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2ND JENNIFER SUMMER SCHOOL

ICTP, TRIESTE

31 JUL 2018

Open Questions: 1900 to Today

Open Questions from Data

Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Open Questions from Data

Dark Matter

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Are we answering the wrong question?

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Uranus or Mercury?



Urbain le Verrier

- **1845** his computations of **Uranus** orbit revealed an anomaly predicted existence of <u>Neptune</u>
 - **1846** Neptune discovered!
- **1859** his computations of **Mercury** orbit revealed an anomaly predicted existence of <u>Vulcanus</u>
 - 1916 General Relativity!

(another important lesson: the power of computing the n-th digit)

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1916 General Relativity!

(another important lesson: the power of computing the n-th digit)

Modified Newtonian Dynamics (MOND) 1983 Mordehai Milgrom: from F = m a to $\begin{cases} F = m a & a \gtrsim a_0 \approx 10^{-10} \text{ m/s}^2 \\ F = m a^2/a_0 & a \ll a_0 \end{cases}$

Just a phenomenological relation, but allowed to fit motion of stars&gas in galaxies! Important: back then, neither Bullet Cluster nor CMB spectrum observed yet

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Challenges to Modifying Gravity

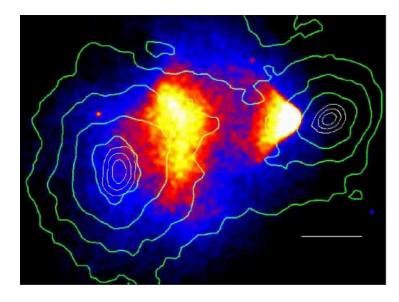
SM + GR + 2 fields, 3 free parameters, 1 arbitrary function
 should be consistent and reproduce confirmed predictions of General Relativity (energy and angular momentum conservation, relativity, gravitational lensing,...)

Challenges to Modifying Gravity

Theory should be consistent and reproduce confirmed predictions of General Relativity (energy and angular momentum conservation, relativity, gravitational lensing,...)
 2004 "Tensor-Vector-Scalar gravity" Bekenstein astro-ph/0403694

SM + GR + 2 fields, 3 free parameters, 1 arbitrary function

Clusters



A DIRECT EMPIRICAL PROOF OF THE EXISTENCE OF DARK MATTER *

Douglas Clowe¹, Maruša Bradač², Anthony H. Gonzalez³, Maxim Markevitch^{4,5}, Scott W. Randall⁴, Christine Jones⁴, and Dennis Zaritsky¹

ApJ Letters in press

ABSTRACT

We present new weak lensing observations of 1E0657-558 (z = 0.296), a unique cluster merger, that enable a direct detection of dark matter, independent of assumptions regarding the nature of the gravitational force law. Due to the collision of two clusters, the dissipationless stellar component and the fluid-like X-ray emitting plasma are spatially segregated. By using both wide-field ground based images and HST/ACS images of the cluster cores, we create gravitational lensing maps which show that the gravitational potential does not trace the plasma distribution, the dominant baryonic mass component, but rather approximately traces the distribution of galaxies. An 8σ significance spatial offset of the center of the total mass from the center of the baryonic mass peaks cannot be explained with an alteration of the gravitational force law, and thus proves that the majority of the matter in the system is unseen.

CMB Substituting DM with Modified Gravity fails to explain CMB [see Skordis 0903.3602 for a review] Only way-out found <u>so far</u>: add matter, that we do not see today (Dark Matter is back!)

No "no-go theorems" here, but explanation why most people work within GR + Dark Matter

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Back to General Relativity + Dark Matter

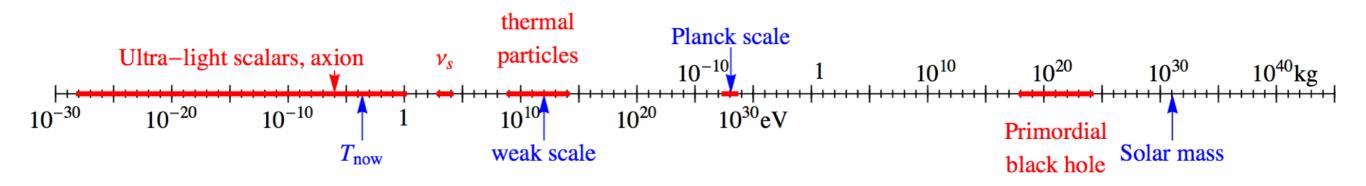
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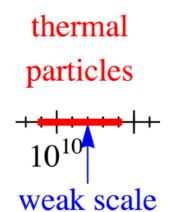
Cooking a DM model: ingredients

| "Particle" properties | - feels Gravity | / | |
|-------------------------|----------------------------------|-----------------------------------|--------------------|
| | - CMB (& not s | spoil BBN,) = <mark>non</mark> - | baryonic |
| | - Invisible now | <pre>v = almost electricall</pre> | y neutral |
| | stable enoug | gh | |
| | | | |
| "Historical" properties | How much? | $\Omega_{\rm DM} \simeq 0.26$ | (Planck satellite) |
| | Since when? | enough before CN | ſΒ |
| | How fast? | Non-relativistic | |

Rules of the game + creativity =



Possibly the "least conventional" candidate



Solar mass

 10^{20}

Primordial

black hole

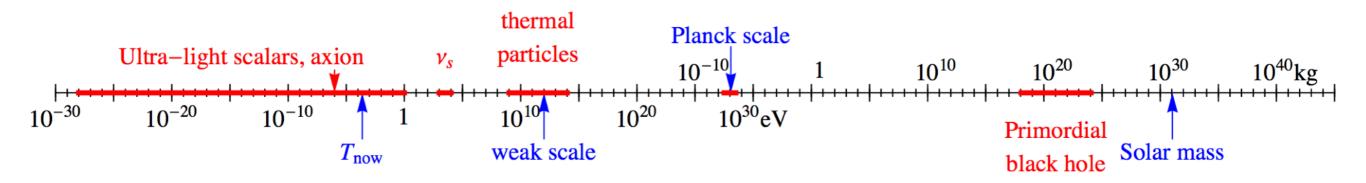
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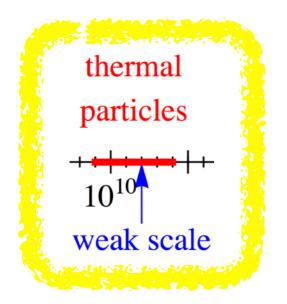
Apologies for not explaining all the rest...

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Rules of the game + creativity =





Weakly Interacting Massive Particles

Weak = SM weak force, DM charged under $SU(2)_w \times U(1)_Y$

(or = whatever interaction with the SM, provided $\alpha = 10^{-3} - 10^{-1}$)

Particle properties: massive

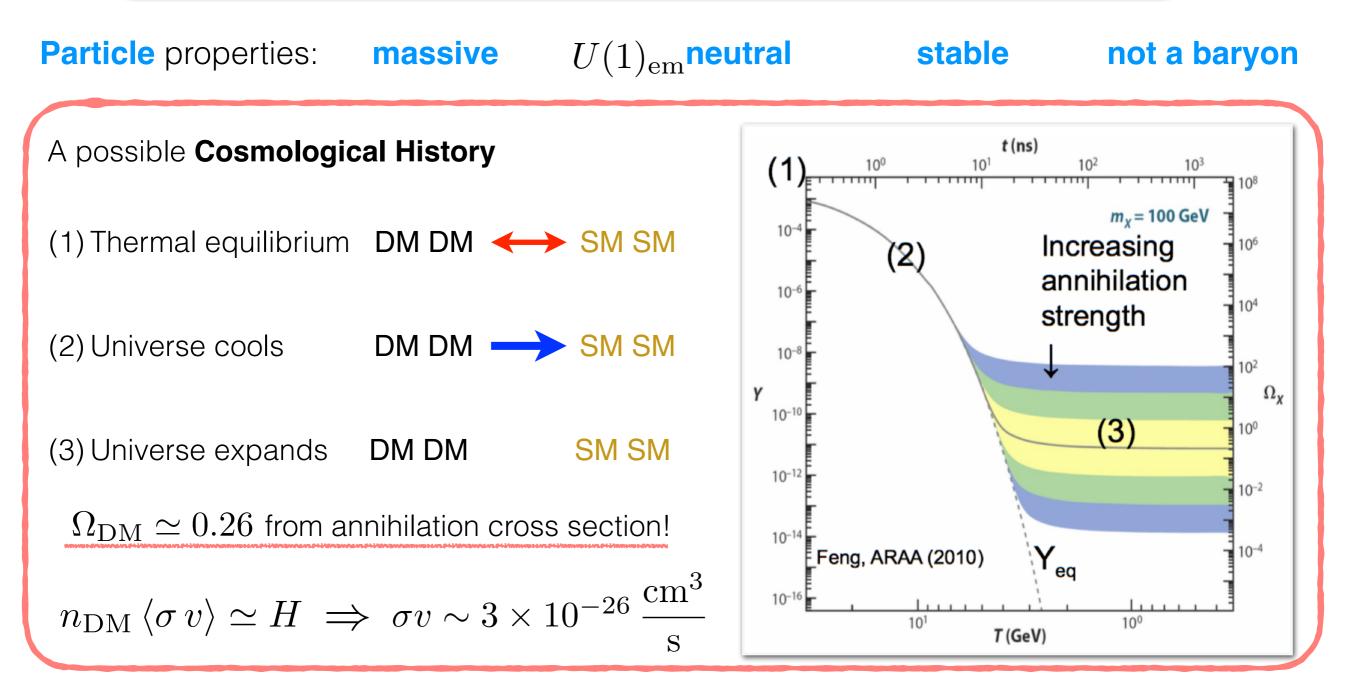
 $U(1)_{\rm em}$ neutral

stable not a baryon

Weakly Interacting Massive Particles

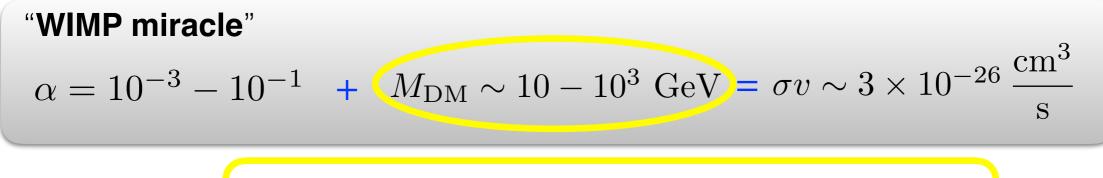
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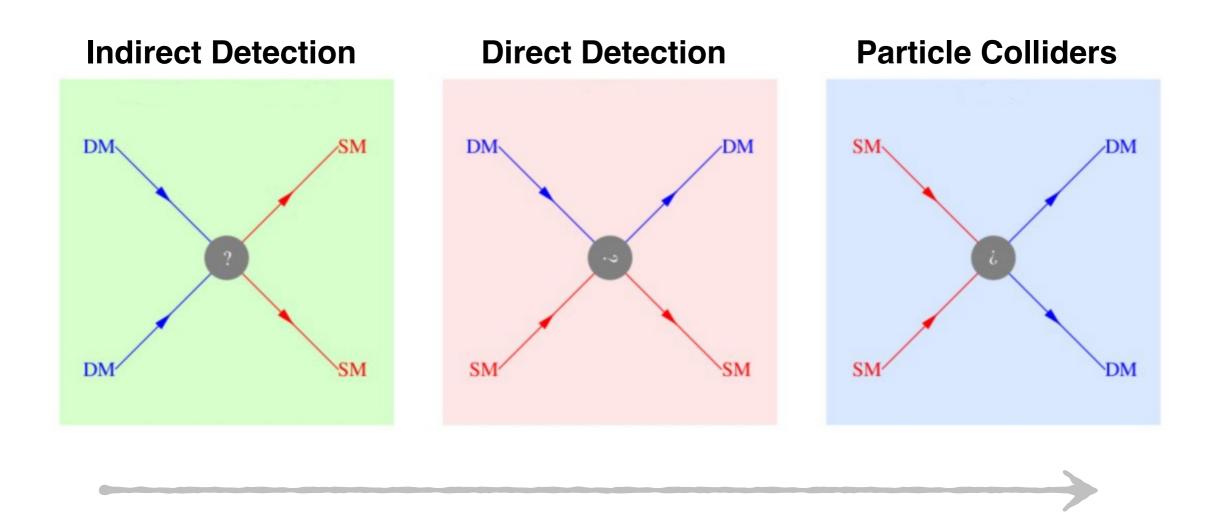


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How to see WIMPs?

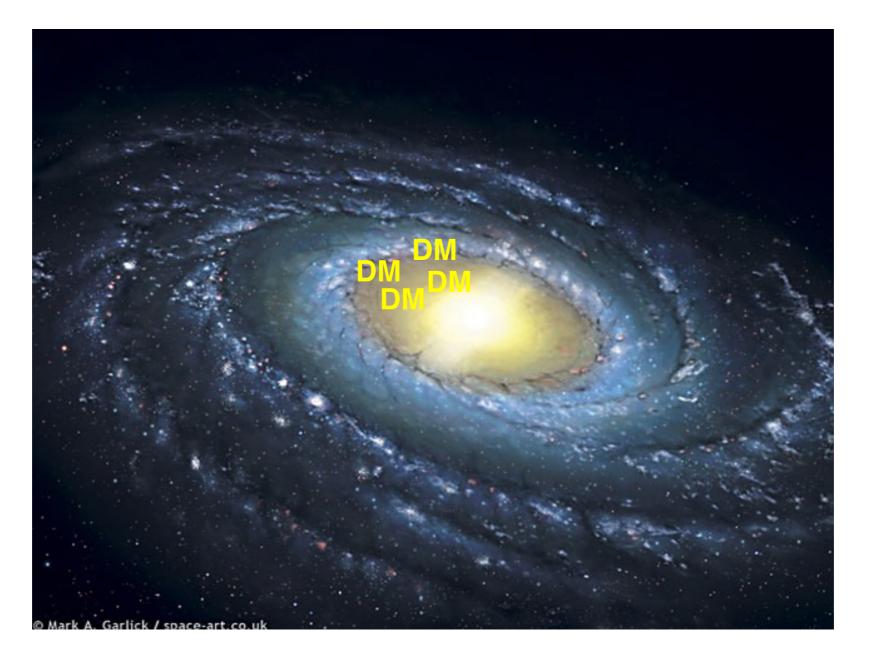


WIMP miracle motivates BSM at the current experiments!



Indirect Detection

Annihilations of DM in the Universe produce $\gamma, e^+, \bar{p}, \nu, \ldots$ to be seen with telescopes!



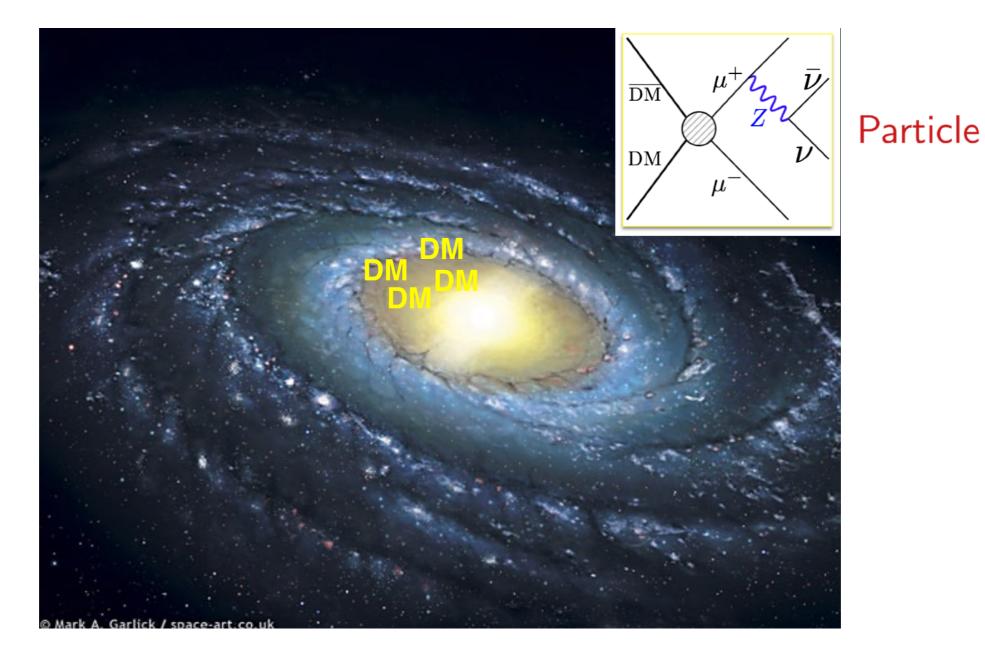
A good primer to this field: "PPPC4DMID", Cirelli+ 1012.4515

Gives tools to compute signals given a DM model, you can try playing with it!

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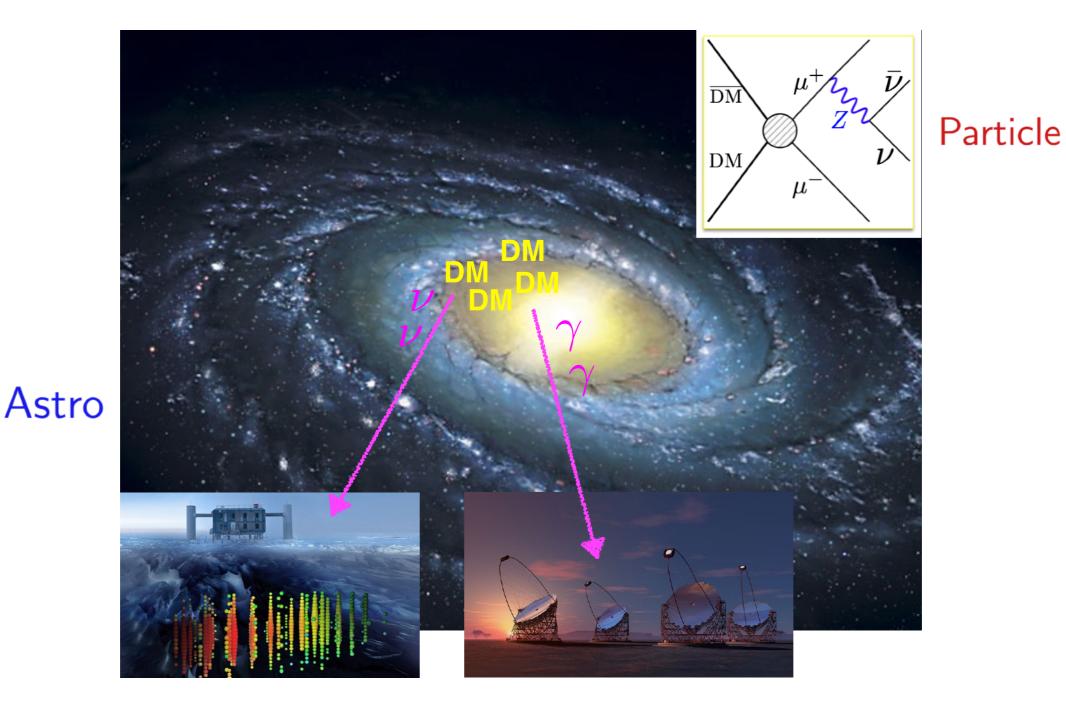
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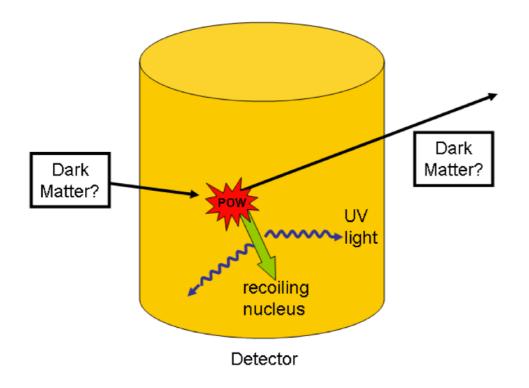
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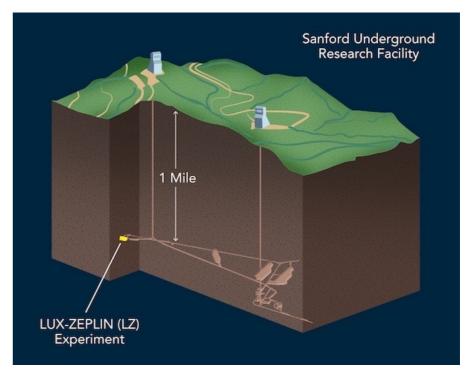


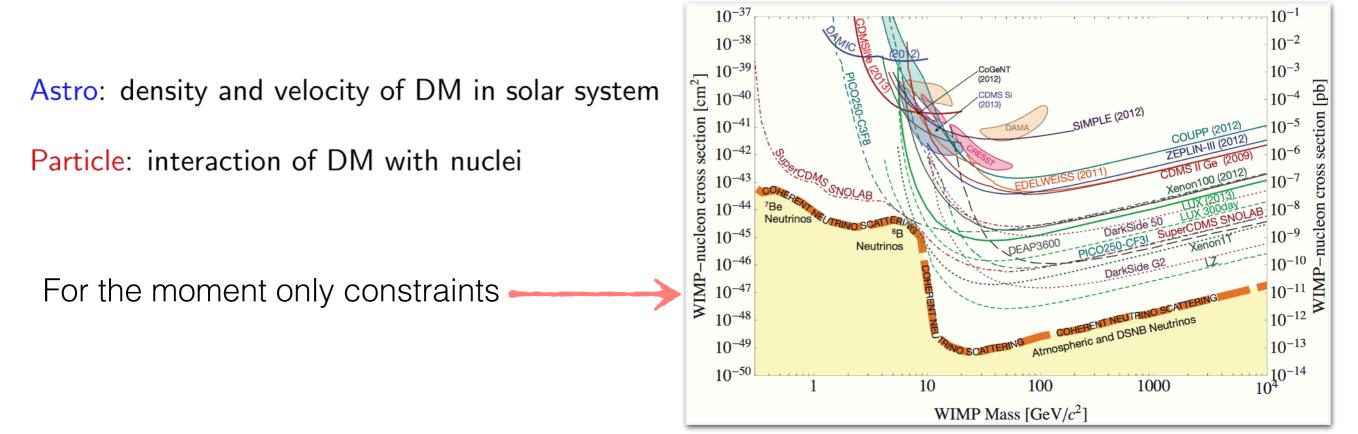
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Direct Detection







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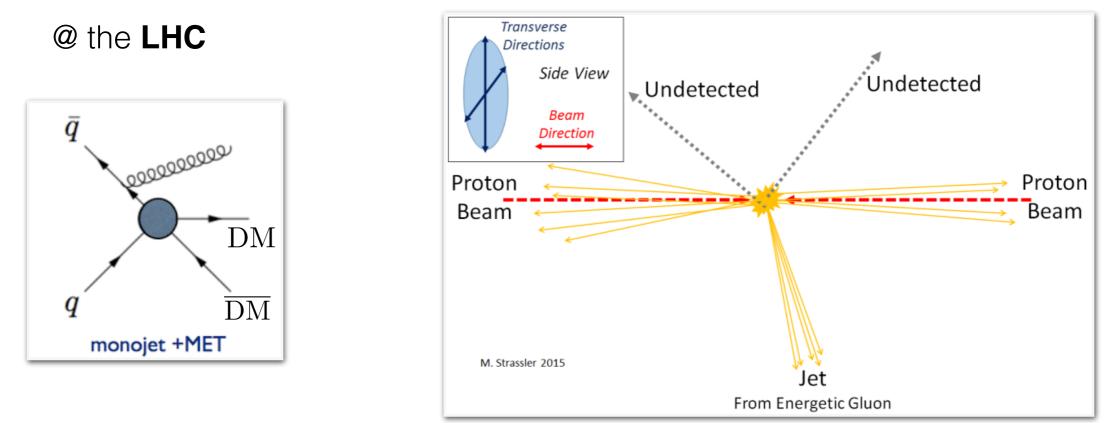
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WIMPs at Colliders

DM (weakly) coupled to the SM \longrightarrow It is produced it in particle collisions!

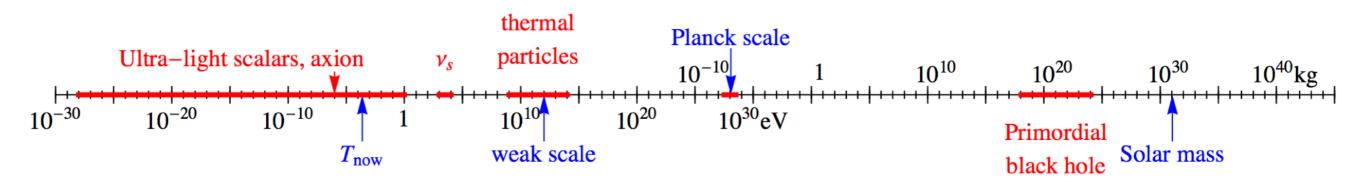
So: look for SM objects (photon, gluon, W, Z, ...) recoiling against "nothing"

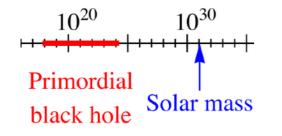
aka mono-SM+ Missing Transverse Energy (MET)



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Rules of the game + creativity =





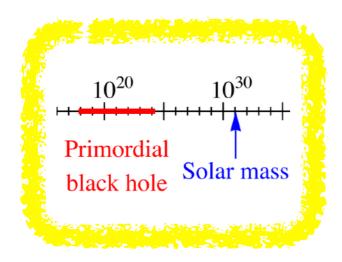
thermal particles

 10^{10} weak scale

WIMPs not discovered so far

(though some "hints" in direct and indirect detection exist, you can ask offline)

Rules of the game + creativity =



Primordial Black Holes (?)

- Gravity (and nothing else: non baryonic + electrically neutral)
- Stable enough (provided they did not evaporate)

. . . .

- Mow to have them? How to have them at CMB??? See e.g. Anne Green 1403.1198
 - Large density perturbations from inflation
 - Cosmic strings loops

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What masses can those mechanisms produce? ~ anything (as far as I understand)

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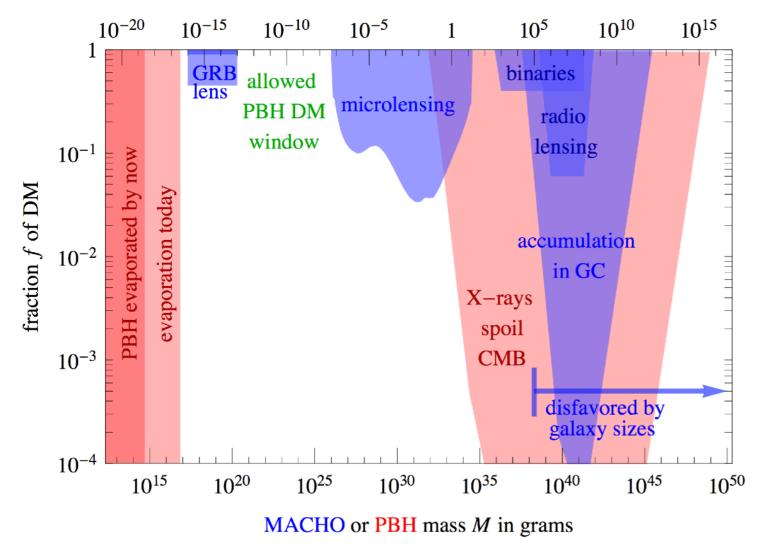
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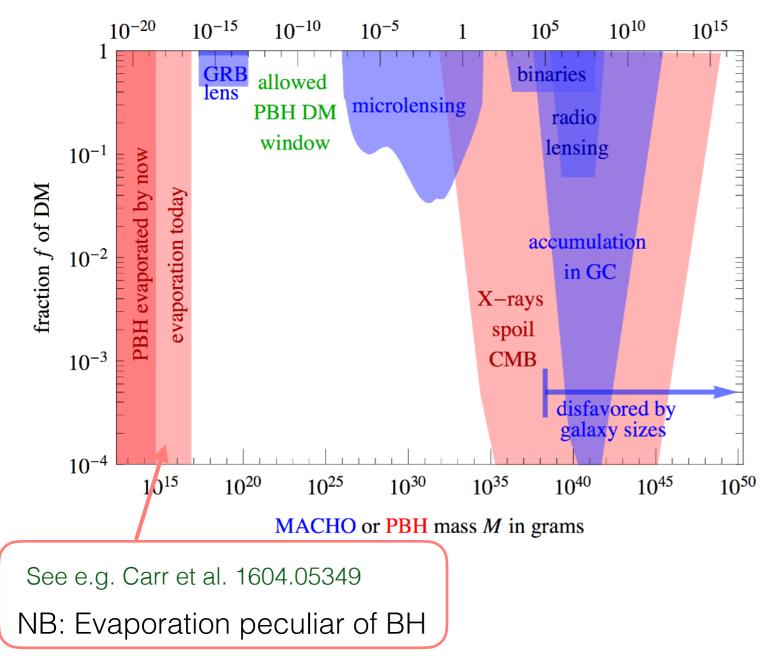
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Still we have gravity to probe them!

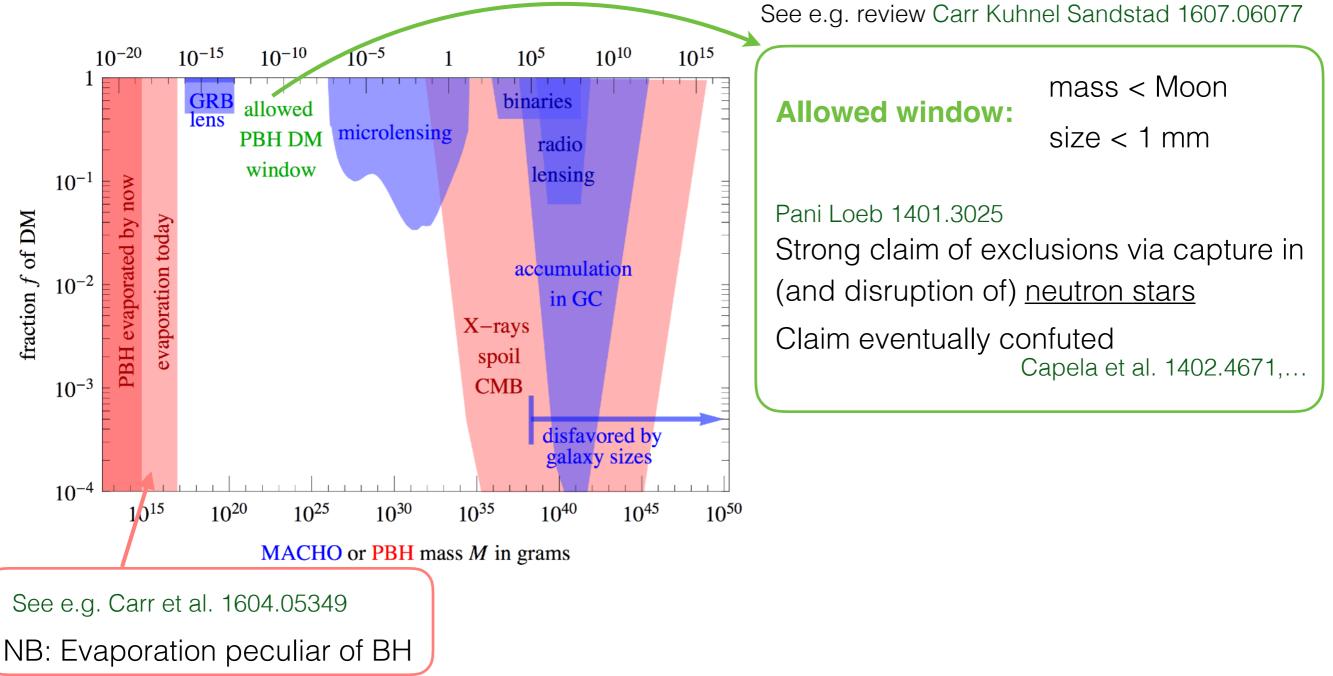
See e.g. review Carr Kuhnel Sandstad 1607.06077

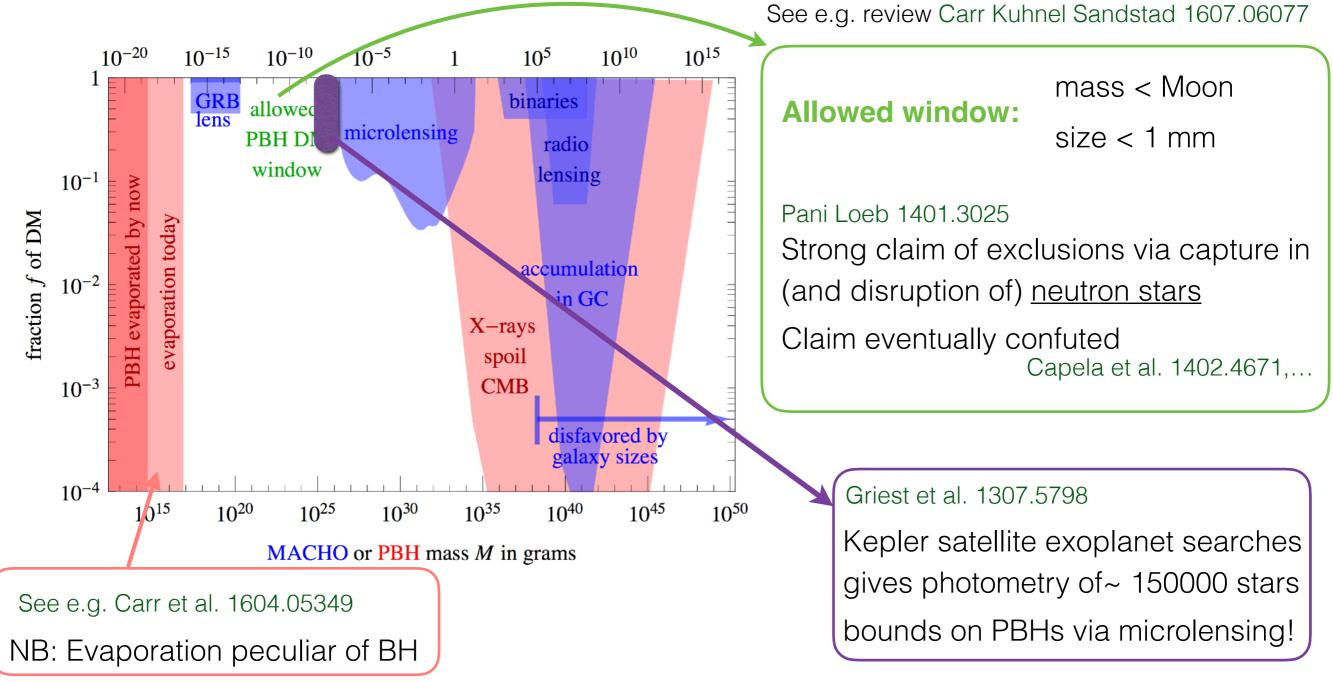


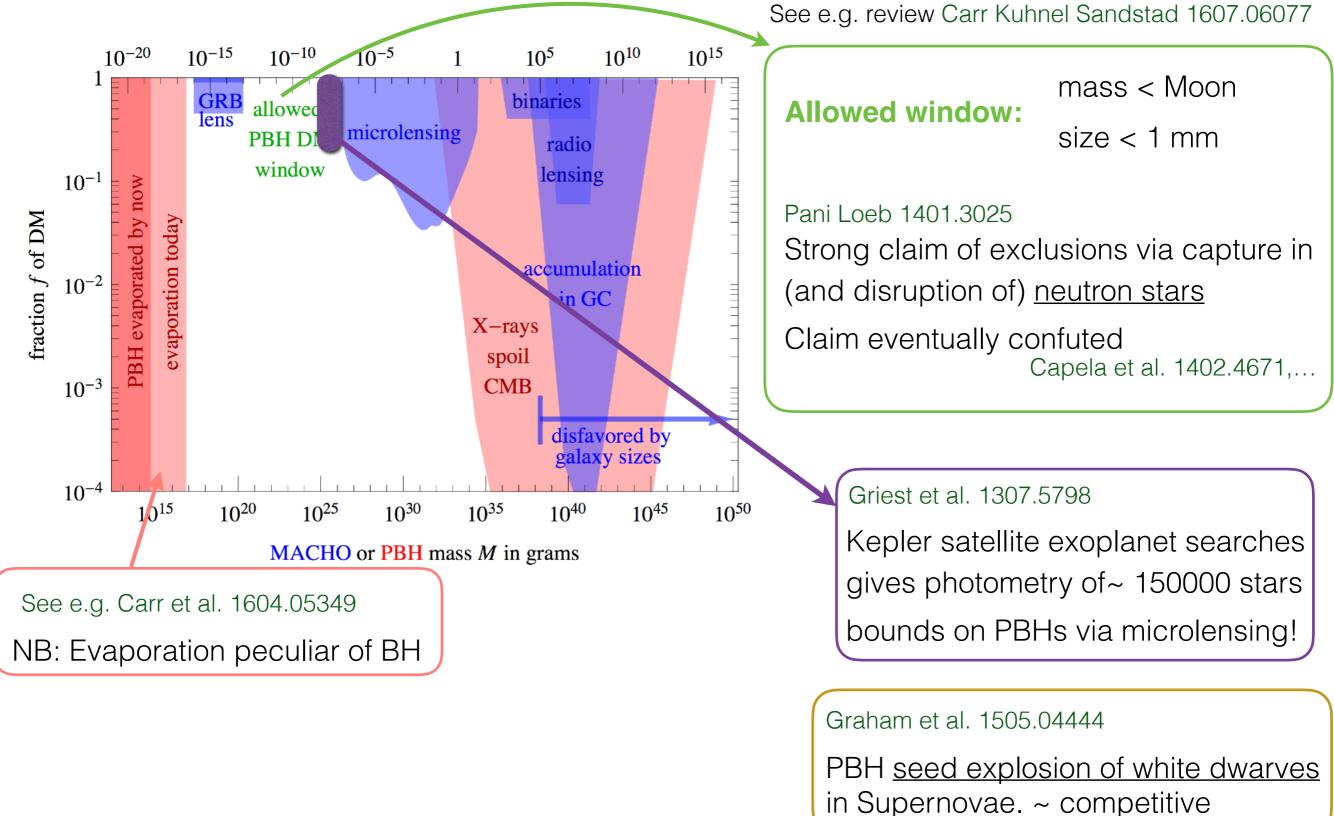
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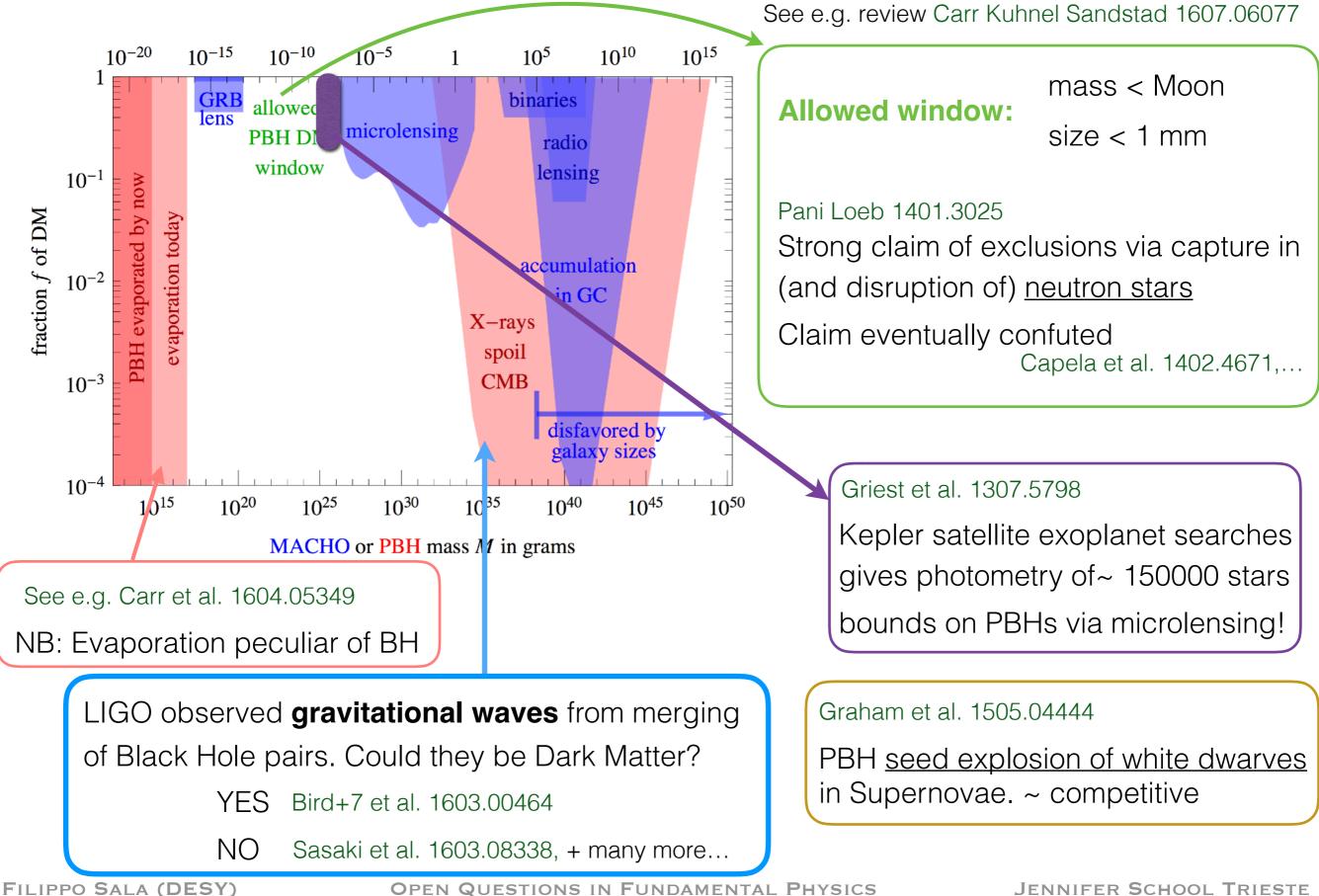


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NB: Dark Energy can be explained by SM vacuum energy (this would pose an enormous theoretical problem, see later)

Open Questions from Data

Dark Matter

Baryon Asymmetry

Neutrino Oscillations

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Open Questions from Data

Baryon Asymmetry

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Baryon Asymmetry

Baryon Particle that is charged under $U(1)_{baryon}$ (NB not same meaning of cosmo baryons) SM particles (and their Baryon charge): quarks (1/3), proton (1), leptons (0), ...

Baryon numberCounts the difference between baryons and anti baryonsConserved in the SM Lagrangian (accidental symmetry, was not imposed)But: not conserved in the SM, at non-perturbative level

Baryon Asymmetry $n_B, n_{\bar{B}}, n_{\gamma}$ number densities of baryons, antibaryons, photons

$$\eta \equiv \frac{n_B - n_{\bar{B}}}{n_{\gamma}} \simeq 6 \times 10^{-10}$$

measured independently from <u>CMB</u> and <u>BBN</u> (so it is a check of theory)

"Thermal" Populations

Thermal freeze-out of SM baryon interactions $n_{B,\bar{B}}\langle\sigma v
angle\simeq H$

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Baryon Asymmetry implies BSM

Sakharov Conditions (1967)

Standard Model

| Baryon Number Viola | ation otherwise no-way to have $n_B eq n_{ar{B}}$ |
|---------------------|--|
| | starting from $n_B = n_{\bar{B}}$ |
| CP Violation | otherwise conjugate processes generate |
| | the same asymmetry, but with opposite sign |
| Out-of-Equilibrium | otherwise expectation value would remain |
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Standard Model

Yes

The Baryon Asymmetry implies physics Beyond the SM

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No The SM has CP-violation (in Cabibbo-Kobayashi-Maskawa matrix of quarks), but too small

No The SM thermal evolution does not go out enough of thermal equilibrium (e.g. ElectroWeak phase transition is second-order)

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Models for the Baryon Asymmetry

Decays of Heavy Particles

- New particle X that decays out-of-equilibrium: $\Gamma_X/H \ll 1 \ {
m at} \ T \sim M_X$

decay rate is much slower than expansion (Hubble) rate

- with two decay channels with different baryon (or lepton) number $X \to a \quad X \to b$
- and where antiparticles have different Branching Ratio $r = BR(X \to a) \neq BR(\bar{X} \to \bar{a}) = \bar{r}$

$$\Rightarrow$$
 generates net baryon number $\Delta B = (r - \bar{r})(B_a - B_b)$

Electroweak Baryogenesis

SM already breaks baryon number

Let us rely on that, and modify EW phase transition so that it is sufficiently out-of-equilibrium (i.e. strong first-order) and add BSM source of CP violation

can give baryon asymmetry, predicts **New Physics at ~ TeV** and **gravitational waves**!

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X could have any mass, also very heavy

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Like DM, the Baryon Asymmetry does not tell us the energy scale of New Physics

Open Questions from Data

Neutrino Oscillations

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Neutrino Oscillations



In the SM, neutrinos do not change flavour while they propagate (= "oscillate")

1960's-2000's Solar Neutrinos Deficit of ν_e from the Sun wrt to prediction w/ massless neutrinos

1990's Atmospheric Neutrinos $u_{\mu}
ightarrow
u_{e}$ deficit of muon neutrinos wrt to electron ones

2012 Reactor Neutrino yet another oscillation angle

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Neutrino Masses imply Neutrino Oscillations

Degree of freedom that propagates = by def. eigenstate of p^2 , so of mass

In the flavor basis $\,
u_e,\,
u_\mu,\,
u_ au$ the mass matrix is non-diagonal

$$\left|\nu_{e}\right\rangle = \sum_{i=1}^{3} U_{ei} \left|\nu_{i}\right\rangle$$

(where mass basis = $u_1, \
u_2, \
u_3$)

$$ig| P_{lpha
ightarrow eta} = |\langle
u_eta(t) |
u_lpha
angle|^2 = \left| \sum_i U^*_{lpha i} U_{eta i} e^{-im_i^2 L/2E}
ight|^2$$

Neutrino Masses beyond the SM

Measured oscillations \Rightarrow we know two mass differences, we do not know mass scale

 \Rightarrow at least one neutrino should have mass $m_{
u}\gtrsim 0.05~{
m eV}$

$$\Omega_{\nu} \simeq \frac{\sum m_{\nu}}{50 \text{ eV}} + \text{CMB} + \text{other cosmology} \implies \sum m_{\nu} \lesssim 0.2 \text{ eV}$$

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Only operator of <u>dimension 5</u> that you can write in the SM, others start at dimension 6

$$\frac{1}{\Lambda_{\rm NP}} L^2 H^2$$

$$L = \begin{pmatrix} \nu_{\ell} \\ \ell^- \end{pmatrix} \quad H = \begin{pmatrix} \frac{G^+}{\sqrt{2}} \end{pmatrix} \quad m_{\nu} \simeq \frac{v^2}{\Lambda_{\rm NP}} \quad \Rightarrow \Lambda_{\rm NP} \simeq 10^{15} \text{ GeV}$$

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Dimension 5 violates our recipe (non-renormalizable): needs "UV completion"

$$\Lambda_{\rm NP} = \frac{M_N}{y_{\nu}^2} \qquad \begin{array}{c} y_{\nu} \approx 1 \quad M_{\rm N} \approx 10^{15} \text{ GeV} \\ y_{\nu} \ll 1 \quad M_{\rm N} \ll 10^{15} \text{ GeV} \end{array} \qquad \begin{array}{c} \text{Again, not an indication} \\ \text{of a New Physics scale!} \end{array}$$

Open Questions: 1900 to Today

Open Questions from Data

Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

| | SM + GR = ? | |
|-----------|--|-----------------|
| | EW vacuum stable? | |
| | Why Qe = -Qp?? | |
| | Unification of Interactions | |
| | Unification of Matter and Interactions | (Supersymmetry) |
| | Hierarchy Problems | |
| Not here: | Why 3 generations? | |
| | Why hierarchical flavour couplings? | |
| | Why Universe flat? (Cosmic Inflation) | |
| | | |

Open Questions from Theory

 $\cdot \cdot \cdot$

SM + GR = ?

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Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

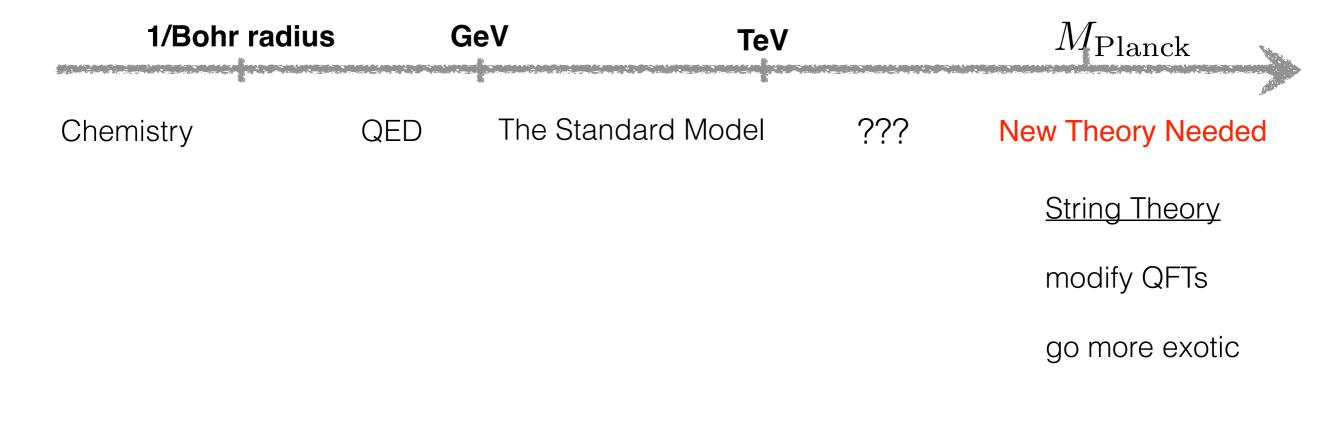
NP scale from consistency of the theory

Gravitational Interaction contains a new fundamental scale

$$M_{\rm Planck} \equiv 1/G_{\rm Newton} \simeq 10^{19} {\rm GeV}$$

At those scales one needs to go beyond either **SM + GR** or **QFT**, possibly **both** (GR is non-renormalizable)

Domains of validity



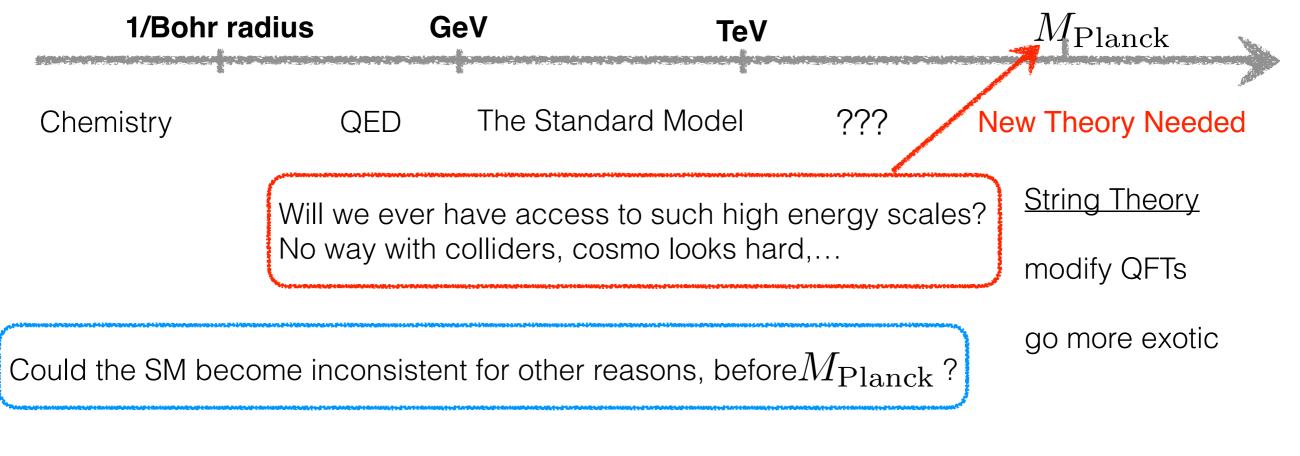
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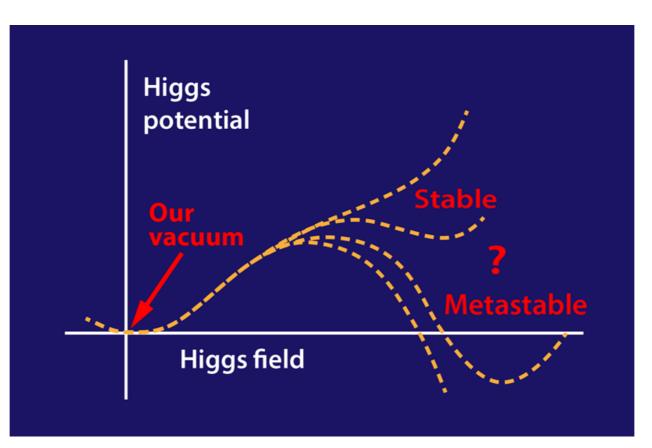
OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Is the EW vacuum stable?

If not, then need NP at the scale where unstable, because we are still here!

$$V(h) = -m^2h^2 + \frac{\lambda(\mu)}{4}h^4$$

 $\lambda(\mu = m_h) \simeq 0.13$ and SM + QFT predicts its running at higher energies



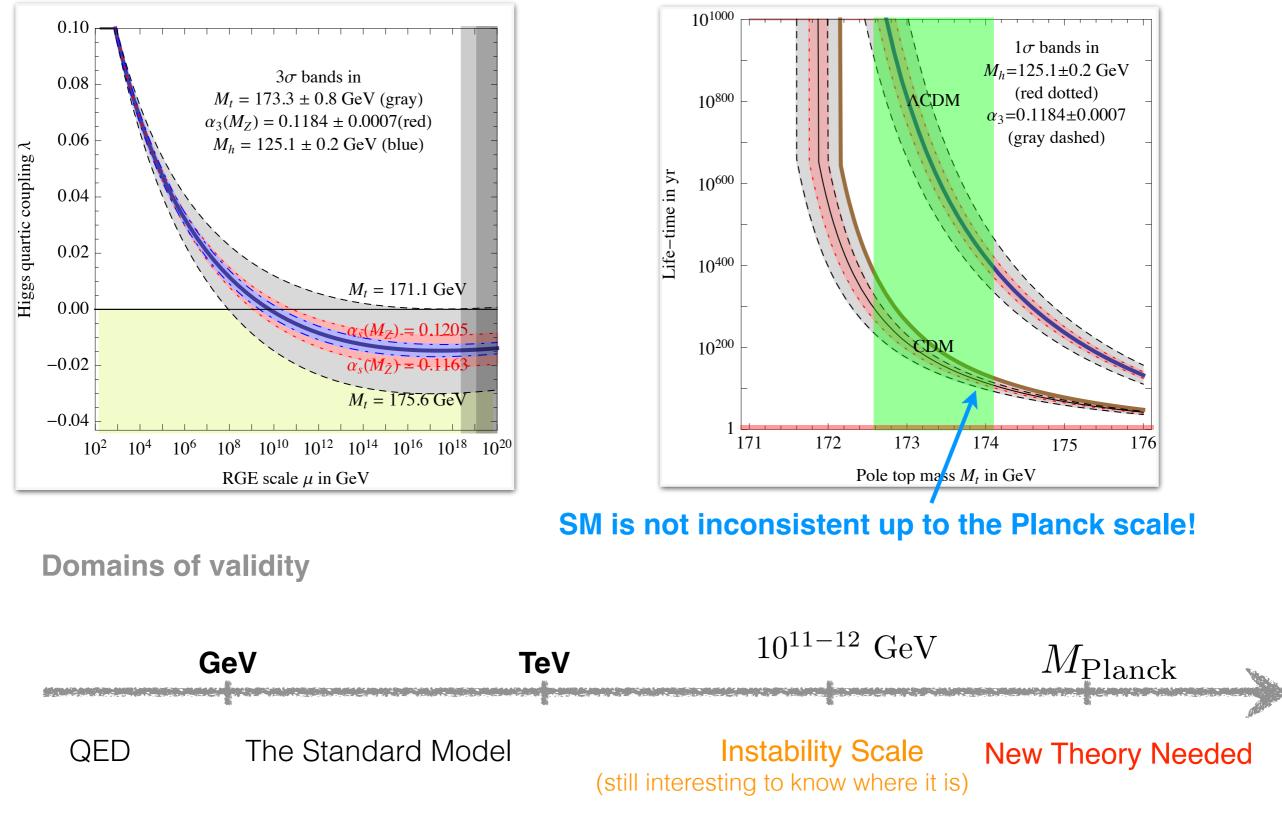
If $\lambda(\mu) < 0$ then our **vacuum unstable** & decays via <u>quantum tunneling</u> to the true vacuum

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Our Fate (according to the SM)

Most precise computation to date Buttazzo+ 1307.3536



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Why Qe = -Qp??

Unification of Interactions

Unification of Matter and Interactions (Supersymmetry)

Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Unifying Interactions - the past



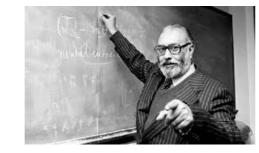
Unified celestial bodies and apples



Unified Electricity and Magnetism







Related Electromagnetism and Weak force

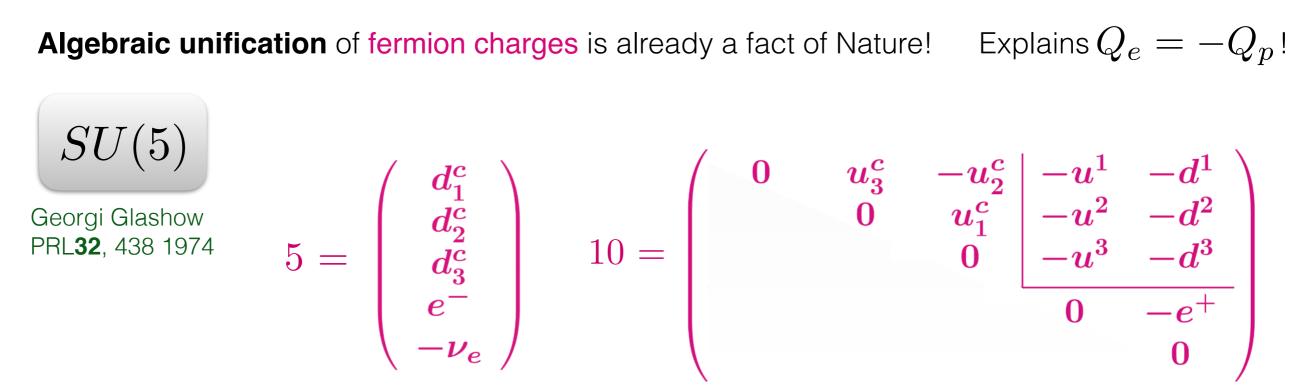
Why should Nature stop unifying?

Why should Nature keep unifying?

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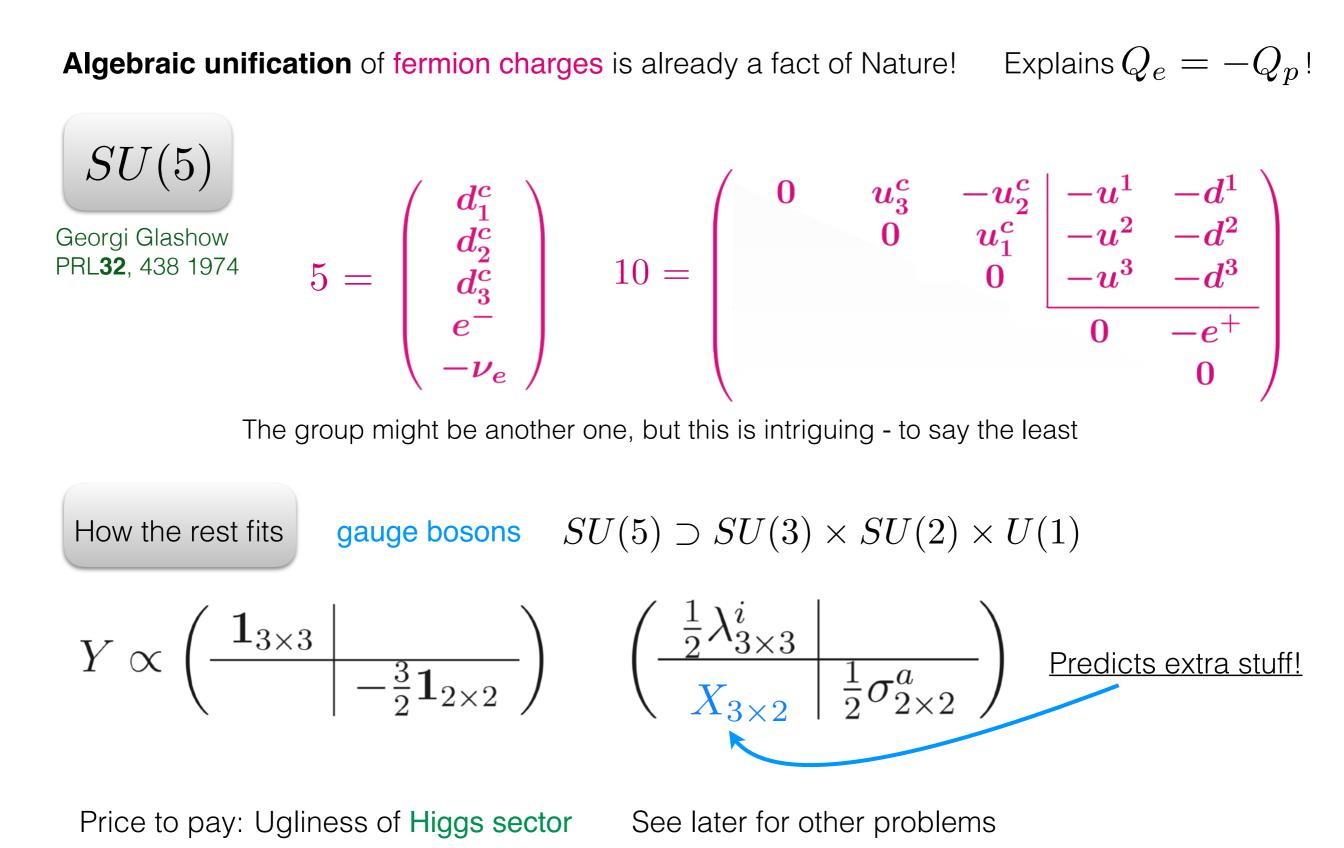
OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Unifying Interactions - the present



The group might be another one, but this is intriguing - to say the least

Unifying Interactions - the present



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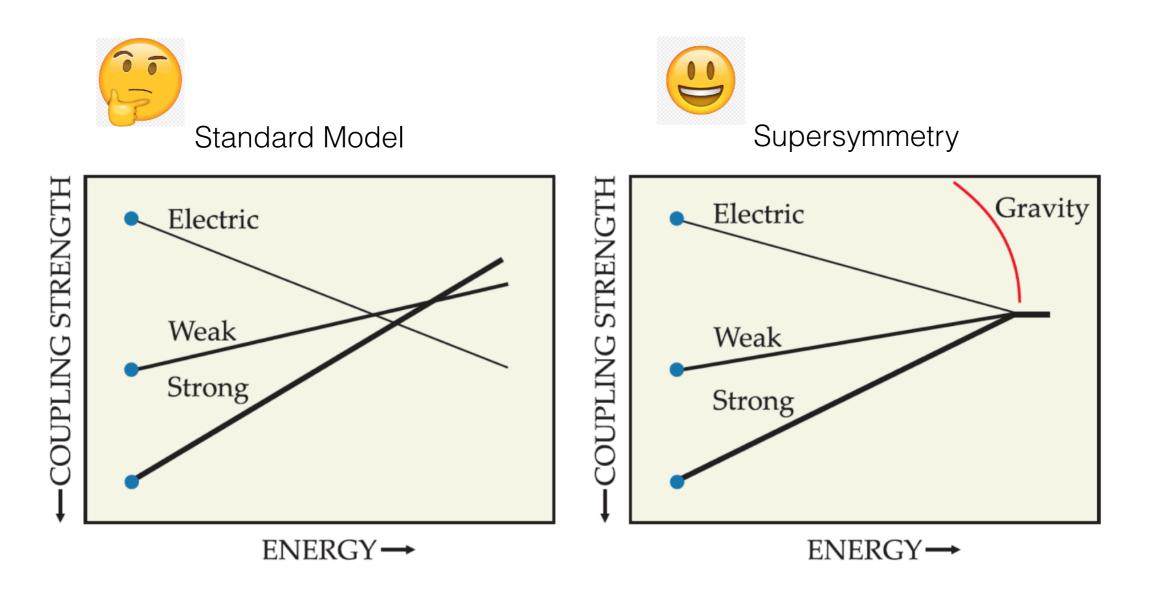
OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Unifying Interactions - the present

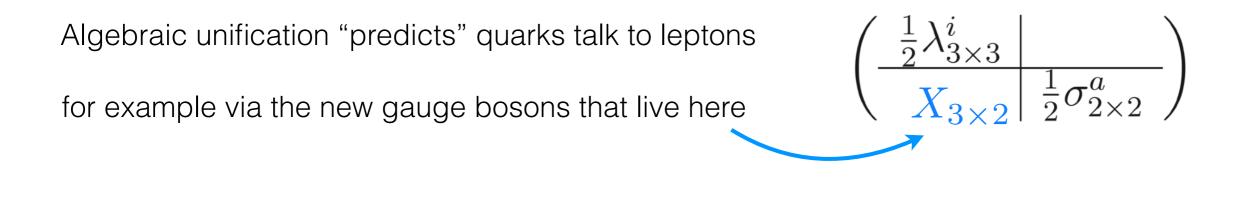
Dynamical unification of **gauge couplings** is to be checked - can be computed!

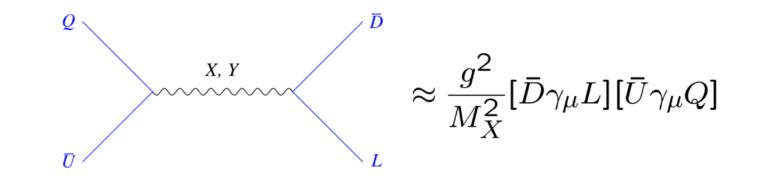
- 3 input parameters: electromagnetic, weak and strong coupling
- 2 output parameters: "the" coupling, the scale of unification

1 highly non-trivial consistency condition!

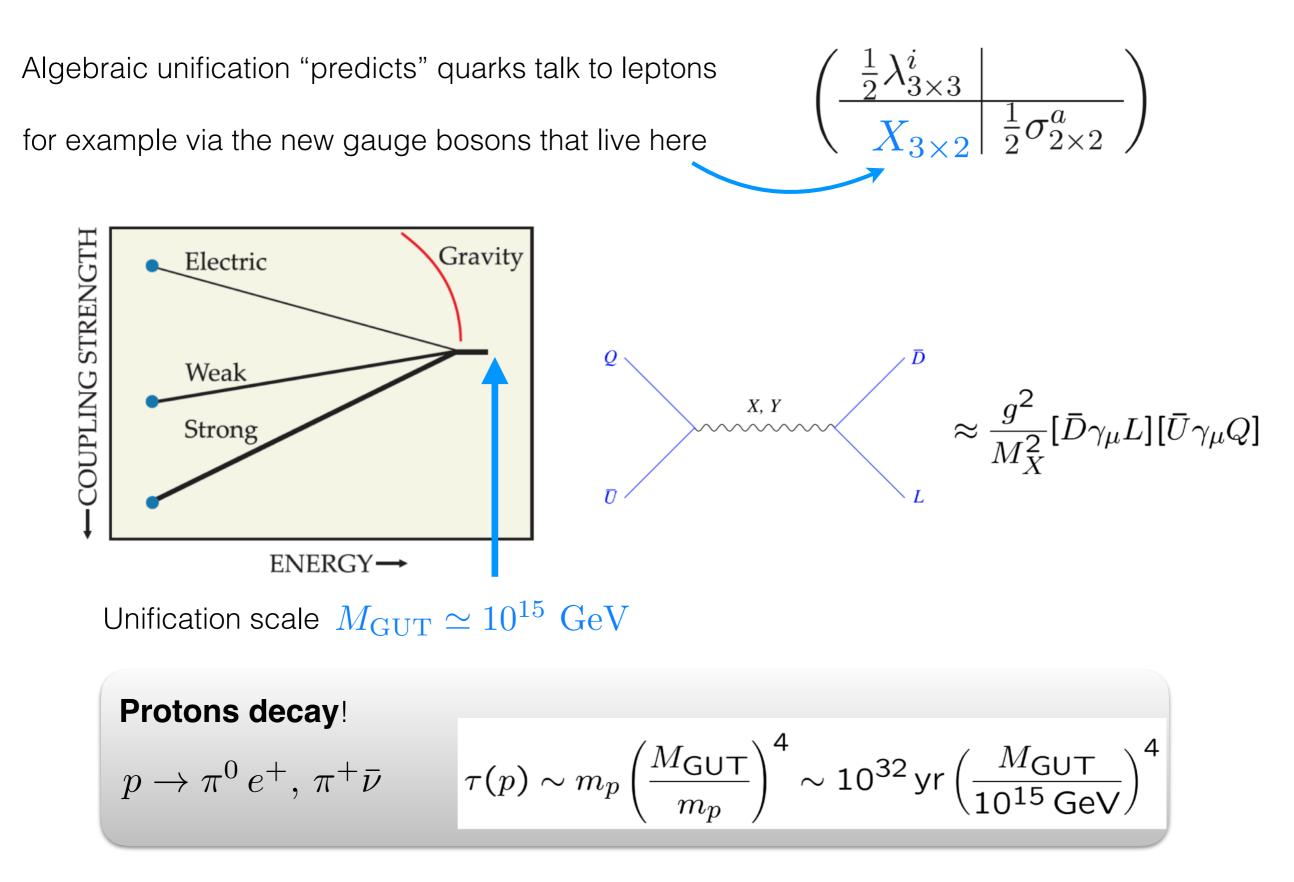


Coupling Unification - Predictions



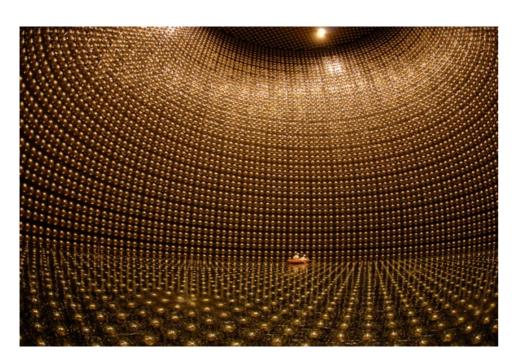


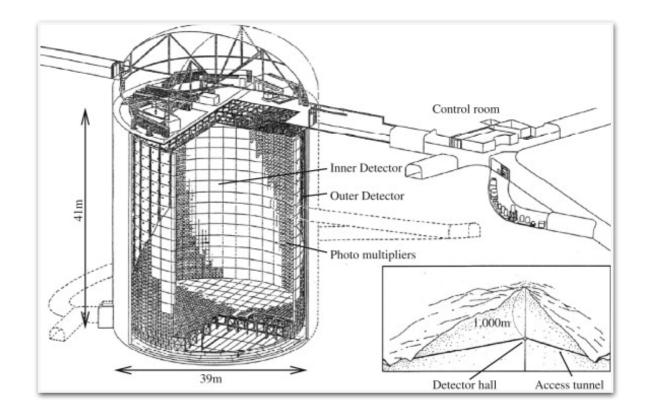
Coupling Unification - Predictions



Coupling Unification - experimental tests

SuperKamiokande

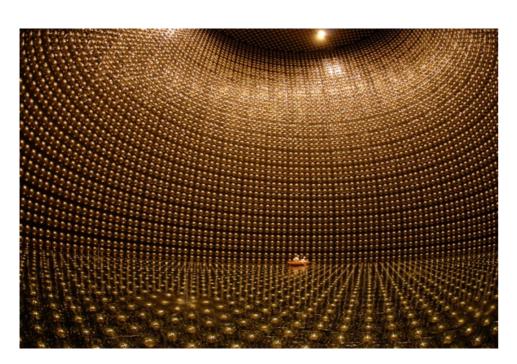


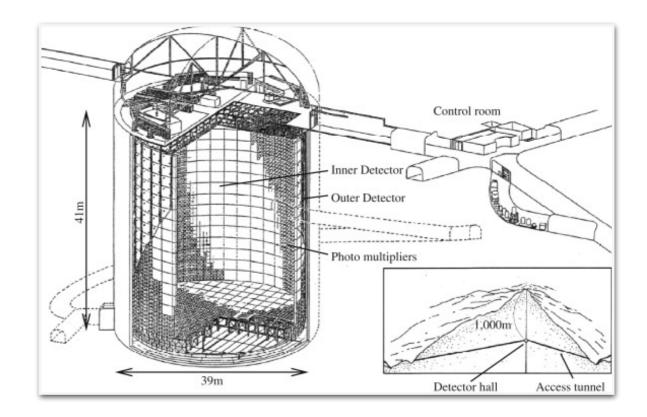


 $pprox 20 imes 10^3$ tons of water, surrounded by detectors

Coupling Unification - experimental tests

SuperKamiokande





 $\approx 20 \times 10^3~$ tons of water, surrounded by detectors

 $pprox 10^{10}\,$ moles of protons

$$au(p) > 10^{10} N_{\text{A}} ext{yr} \sim 10^{33} ext{yr} \quad M_{\text{GUT}} \gtrsim 10^{15 \div 16} \, ext{GeV}$$

Observe for some years:

Limit "touches" current predictions, but not enough to disproof unification

Another lesson: precision measurements can tell us about extremely large scales!

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Deeper Unification: Supersymmetry

Unifies Matter with Interactions

SUSY does so by relating fermions and bosons

 $Q |\text{fermion}\rangle = \text{boson}$ $Q |\text{boson}\rangle = \text{fermion}$

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Roots: **1967** Coleman & Mandula

Most general symmetry structure of QFT can be Poincaré x internal symmetries

Assumptions: causality, locality, ..., bosonic symmetry generators

What's the physical motivation?

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Assumptions: causality, locality, ..., bosonic symmetry generators

What's the physical motivation?

Drop this assumption ——> Supersymmetry!

Every particle has a partner with spin differing by 1/2 same mass same quantum numbers Partners of SM particles not seen \longrightarrow Break SUSY!

OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

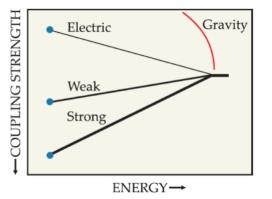
SUSY: Many Virtues and one Vice



By itself a deeper Unification of Laws of Nature (matter and interactions)



Predicts New Particles that automatically induce **gauge couplings unification**





Necessary ingredient of String Theory ("leading" candidate for quantum gravity)



Accidental Symmetries provide many **Dark Matter** candidates (WIMP, Gravitino...)



Solves the Hierarchy Problem of the Fermi Scale (see next slides)

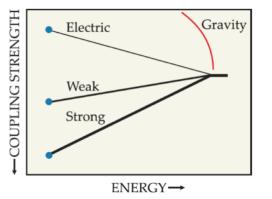
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Solves the Hierarchy Problem of the Fermi Scale (see next slides)



Where are the superpartners??

| | SM + GR = ? | |
|-----------|--|-----------------|
| | EW vacuum stable? | |
| | Why Qe = -Qp?? | |
| | Unification of Interactions | |
| | Unification of Matter and Interactions | (Supersymmetry) |
| | Hierarchy Problems | |
| Not here: | Why 3 generations? | |
| | Why hierarchical flavour couplings? | |
| | Why Universe flat? (Cosmic Inflation) | |
| | | |

Open Questions from Theory

 $\cdot \cdot \cdot$

Hierarchy Problems

Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Beauty and Ugliness in the SM



Gauge sector

Scalar sector

- Higgs mass and couplings
- Cosmological constant
 - 2 hierarchy problems

flavour problem

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

The Hierarchy (or Naturalness) Problems

Physical system (SM + GR) with 3 fundamental scales They should be of the same order

Why
$$\frac{\Lambda}{M_{\rm Planck}^4} \approx 10^{-120}$$
, $\frac{m_h^2}{M_{\rm Planck}^2} \approx 10^{-34}$???

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E.g. NP energy scale responsible for: DM, Neutrino Masses, Quantum Gravity, GUT,...

So why not
$$m_h^2 \approx M_{\rm DM}^2 + M_{\rm GUT}^2 + M_N^2 + \dots$$
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Needs huge fine-tuning of coefficients of each contribution

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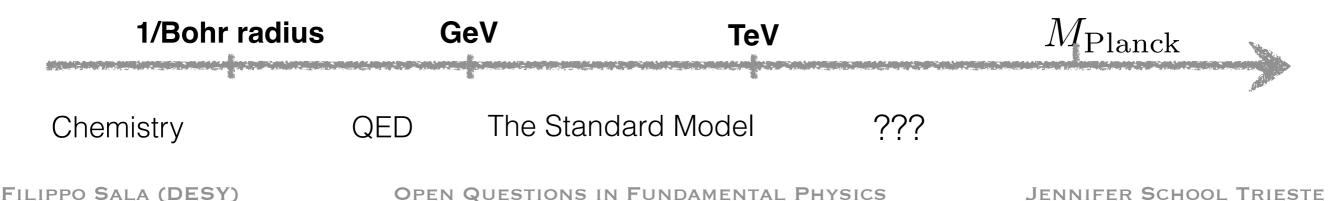
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SM depends on detail of much larger scales: a challenge to reductionism?



Hierarchy Problem and New Physics

In the past, naturalness problems signaled new physics:ProblemNatural solution?New physics $\delta m_e = \alpha \Lambda$ Yes: chiral symmetrypositron $\delta(m_{\pi^+}^2 - m_{\pi^0}^2) = \alpha \Lambda^2$ Yes: π are compositeQCD

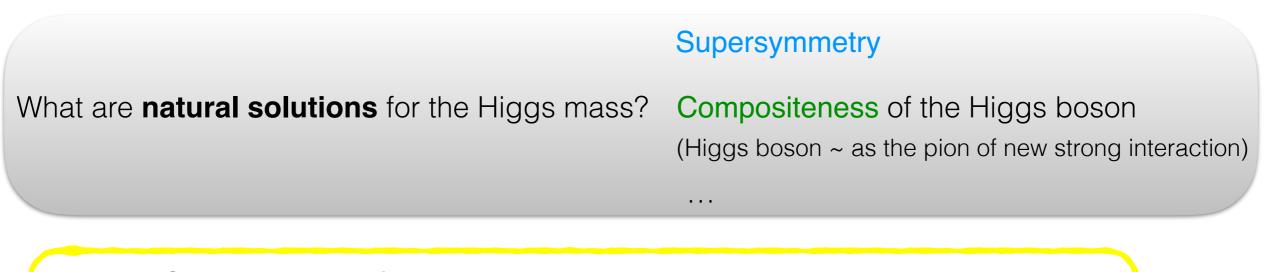
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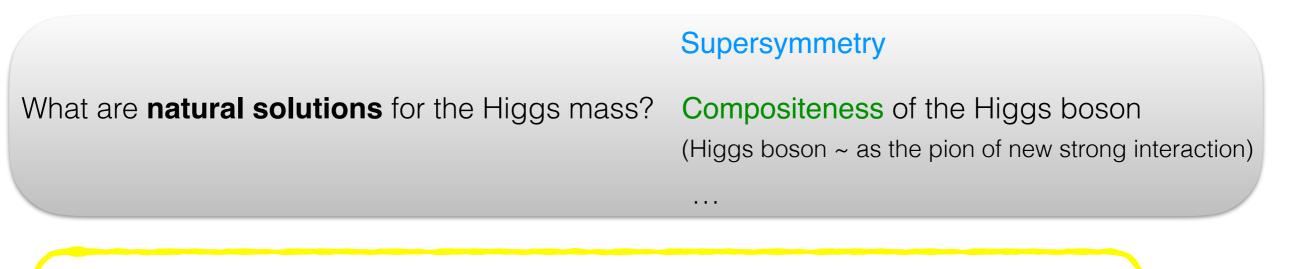
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"Natural Solutions" by definition predict New Physics close to the Higgs mass

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But LEP and LHC found no BSM!

Also, they do not work at all for the cosmological constant!

Other Solutions?

Warning: my opinion

Current experimental exclusions not enough to discard Naturalness **

However, they definitely motivate to explore alternatives (+ we still have to explain Λ)

**Quantitatively, they imply tuning at % level and we observe similar tunings in Nature (e.g. Solar Eclipses!)

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Anthropic Selection on a Multiverse

Some parameters, if slightly different, would not allow for life to develop

But: how to test?



Weinberg obtained this way an upper bound on Λ , very close to value measured much later!

Same reasoning can work for Higgs mass (and also light Yukawa couplings)...

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

Open Questions: 1900 to Today

Open Questions from Data

Open Questions from Theory

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OPEN QUESTIONS IN FUNDAMENTAL PHYSICS

What Next?

Like in 1900: we understand almost everything, but for some "clouds"

Like in 1900: paradigms that worked in the past century are suffering!

Unlike 1900: current model passed loads of experiment and theory tests

The optimist Dream-like situation! On the verge of a revolution!

The pessimist The needed NP could not, and so will not, show up at experiments corollary: go to math/finance/agriculture/...

The pragmatic Let's make more tests

Good News is that data are coming!

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