# Experiments with FAZIA from LNS to GANIL

### Giovanni Casini INFN Sezione di Firenze











#### Eurisol- Town Meeting, Pisa 7/2018

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### The physics motivation

### Heavy-Ion Collisions at Fermi energies: explore the Equation of State (EoS) in a range of excitations, densities and n-p asymmetries.

**1) Isovector properties of the EoS,** Isospin diffusion/migration, symmetry energy, iso-yields, vaporization (low density)

**2) Transport properties and connection to the transport models** Stopping and energy dissipation; *NN* cross-section; particles and cluster flows; *NN* and *multi*-body correlations in dense matter

See: Y.Zhang et at PRC 97 2018 output of the **CODE EVALUATION PROJECT** 

### 3) Clustering and decays inside excited nuclear matter

In-medium properties of clusters, links between nuclear dynamics and nuclear structure

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### Investigation of EoS with collisions

$$E(\rho, \delta = \frac{N-Z}{A})/A = E_0(\rho, \delta = 0) + E_{\delta}(\rho, \delta)$$

$$E_{\delta}(\rho, \delta) = E_{sym}(\rho)\delta^2$$

$$E_{sym}(\rho) = \frac{1}{2} \left(\frac{\partial^2 E_{\delta}(\rho, \delta)}{\partial \delta^2}\right)_{\delta=0}$$

$$E_{sym}(\rho) = S + L_{sym}x + \frac{1}{2}K_{sym}x^2$$

$$x = (\rho - \rho_0)/(3\rho_0)$$

e.g.  $L_{sym}$ =78(ave.) +-31( $\sigma$ ) MeV J.Margueron 2018

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### Density and Isospin gradients in HIC

Semiperipheral reactions at Fermi energies produce regions at subsaturation densities (V.Baran et al. N.Phys A 730 2004)



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### The FAZIA detector



 Silicons 300+500micron (special nTD wafer with specific crystalline orientation cut, reverse mounting)
 Csl(TI) 10 cm homogenous TI doping (5%) custom Phodiodes for maximum active area
 Telescope and Block Active area: 20x20mm2 Block (BLK): 16 telescopes, in a 4x4 matrix ΔΩ for a BLK at 100cm: 6.4 msr Granularity at 100cm: 1.2 deg





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### The FAZIA system: flexible and versatile





#### 🧾 FEE

6-ch detectors (3ch x 2Telescopes)
Preamps, pulser; HV bias gen/regulat/control;
12 sampling ADC's, 2 FPGA, FIFO
BLOCK cards and HV generation cards
Optical link to Regional Board

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<u>Easy</u> <u>mounting/operation</u> 3 I/O (used) connections:

48V, optical fiber, water pipes

### Neck dynamics and neutron enrichment

- neck formation and rupture favour IMF at midvelocity
- preferential alignement of the final fragments
- neck region and local neutron enrichment



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### **ISOFAZIA** experiment

PhD Thesis G.Pastore, Firenze 2017



planar geometry

Study of **isospin dynamics** looking at the properties of the Quasi-Projectile (QP) remnant and of its decay products (in practice from Z=1 to Z around 20)

- variation of: neutron content of the target N/Z=1 and 1.4

- variation of: QP decay mode (evaporation, fissionlike)
- comparisons with transport model predictions

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### **FAZIASYM** experiment



Again we study **isospin dynamics** measuring the properties of the QP and of its decay products (<u>from Z=1 to the charge of the QP</u>)

variation of: neutron contents of the proj/targ N/Z=1 and 1.4
 variation of: QP decay mode (evaporation, fissionlike)
 looking at specific N=Z dominated channels (alpha rich neck?)

- comparisons with transport model predictions





#### **FAZIASYM EXPERIMENT** at LNS, Catania

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### neutron abundance of neck (ISOFAZIA result)



Once more confirmed the **neutron enrichment of the midvelocity region**. Effect interpreted as due to isospin drift (see also e.g. De Filippo PRC 86, 2012)

S.Piantelli at IWM-MC 2018 and PhD Thesis G.Pastore, Firenze 2017



AMD+Gemini calculationsshow slight sensitivity to the asy-stiffness ofthe EOS (weak indication for a very stiff behaviour  $L_{sym} \sim 100$ )Giovanni Casini INFN FirenzeEurisol- Town Meeting, Pisa 7/2018

### New approach to fast fission (FaziaSym result)



Can we t a kind of the exam

IMPORTANT: extensive simulations AMD+Gemini to keep systematics LHC con under control and then to further study fission time-scales vs. isospin equilibration and EOS asy-stiffness

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### New approach to fast fission (FaziaSym result)



just a crude idea.....Can we think of a kind of TIER on the example of LHC for best use of computing resources in our Eurisol area?

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### The last experiments at LNS



configuration of SiX blocks at 100cm distance from target

FAZIACOR: 32S and 20Ne+12C @ 25-45MeV/u (completed march 2017) FAZIAPRE: 40,48Ca+12C @ 25-40MeV/u (completed may 2018)

alpha-cluster decay modes, sequential vs. direct (in primis 12C Hoyle state).

resonance decay modes in the hot nuclear environment fast particle emissions vs isospin dynamics with increasing bombarding energy

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### Nuclear resonances in hot environment

Spectroscopy of hot Fragments created during the nuclear reactions



### PRELIMINARY apparently almost 100% decay occur through 8Be

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Various cases of hot fragments above particle emission threshold



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# Transmission-type Experiment

### FAZIA ZERO experiment (B.Sun Behang Univ) 12C+12C at 60-80 MeV/u, this july

- Total reaction cross-section TCS and charge-changing cross-section CCCS allow to determine the nuclear and charge radii. Their difference gives access to nuclear skin (sensitive to Lsym next to the saturation density)
- Neutron skin and neutron star radius related both to EOS (Fattoyev ArXiv:171106615)
- Single experiment with FAZIA (at zero deg) gives TCS and CCCS!



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### Coupling FAZIA and INDRA at GANIL





FAZIA will cover the forward region 2-12deg; beyond we will use the INDRA telescopes

Memomandum of Understandings on FAZIA\_INDRA just now approved by parties

### 2018

 ★ Two FAZIA blocks now mounted in the INDRA vacuum chamber
 ★ Commissioning with parasitic beams is on going (24 june 2018 till end july)
 ★ Proposal to the PAC sent just now 2019
 ★ Start measurements with FAZIA+ INDRA

### Scientific reach with FAZIA\_INDRA

#### first proposed experiment : 58,64Ni+58,64Ni @ 32 and 52MeV/u



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# Where we are with thresholds



# **Possible improvements**

#### Use of thin first layers to lower the DE-E threshold





**INFN** funded

Si

SiCilia project S.Tudisco et al.

First results from a DE-E telescope where the first layer is a very thin (10µm) **SiC** epitaxial detector



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## **Possible improvements**

# Reduce E-thresolds for mass determination usinig **Time of Flight**

S.Valdrè INFN Fi IWM\_MC 2018

 Efforts to improve the signal synchronization (the common clock to ADC allows around 1ns)

Validated a method using fast infrared LED to synchronize Si1 of FAZIA during the experiments (accuracy 10ps)

Already verified E-th for proton discrimination down to 2MeV



### Conclusions: FAZIA towards EURISOL

A new detection system based on solid state telescopes has been built with <u>excellent ion identification capability</u>, by optimizing the DE-E and PSA techinques

Various experiments done at LNS demonstrated the <u>versatility and</u> <u>flexibility of FAZIA</u>

The commissioning of the <u>FAZIA-INDRA setup is ongoing</u> and this is the ouverture of the activity for next 4-5 years at GANIL with Fermi energies stable (and LISE ?) beams

The obtained perfomances are optimal above 20MeV/u. At the lower energies of ISOL facilities, the <u>threshold for isotopic separation</u> of IMF Z>4 is an issue and future efforts will be done on that

The great FAZIA versatility permits further <u>experimental collaborations</u> and to exchange our (hardware and software) expertise along the road to EURISOL

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### The accessible (Z,A) regions



### charge identification from 2 MeV/u (for Z = 6) isotopic discrimination from 5 MeV/u (for Z = 6)

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### neck emissions



preferential aligned configuration where a small fragment is in between QP and QT



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6	sampling ADCs per telescope			
	Si 1	14 bit, 100 MHz 14 bit, 250 MHz 14 bit, 250 MHz	4 GeV full-scale charge signal 250 MeV full-scale charge signal current signal	QH1 QL1 I1
	Si 2	14 bit, 100 MHz 14 bit, 250 MHz	4 GeV full-scale charge signal current signal	Q2  2
	CsI(TI)	14 bit, 100 MHz	300 MeV Si-eq. f.s. charge signal	Q3

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### Neck dynamics and neutron enrichment



Semiperipheral reactions a neck formation and rupture favour IMF at midvelocity preferential alignement of the final fragments neck region at subsaturation densities isospin drift can be responsible of neutron enrichment of neck matter

