

Quantum enhanced measurements

from ultra-high sensitivity in absorption measurements to detection of quantum gravity effects

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- Optical measurements are widespread in many branch of science.
- They are simple to be experimentally implemented and they can be used for a lot of different scopes, from development of new technologies to investigation of more fundamental questions.



Using classical light, said N the mean number of photons in the beam, it holds:



N.B. Light power cannot be increased arbitrarly!





If classical light is used, we are limited by the shot noise.

- It comes from the quantization of the field and is an unovoidable limit in sensitivity.
- It scales as $1/\sqrt{N}$ but often we cannot increase N arbitrarly.
- In high precision experiments it can be the principal source of uncertainty, in these cases going beyond the shot noise limit would be extremely important.

But...

Quantum light can improve the information gained /photon !

Therefore, using opportune "quantum states of light" it is possible to go below the shot noise limit.



Quantum enhanced measurements







Quantum states of light can offer important advantages in imaging protocols, in particular in presence of low illumination level.

PHYSICAL REVIEW A 77, 053807 (2008)

High-sensitivity imaging with multi-mode twin beams

E. Brambilla, L. Caspani, O. Jedrkiewicz, L. A. Lugiato, and A. Gatti INFM-CNR-CNISM, Dipartimento di Scienze Fisiche e Matematiche, Università dell'Insubria, Via Valleggio 11, 22100 Como, Italy (Received 28 September 2007; revised manuscript received 20 December 2007; published 8 May 2008)

Sub-shot noise imaging has been realized in our labs in 2010.

From a proof of principle experiment we've recently moved toward real applications implementing a sub shot-noise microsope.











nature notonics

72 pixels at full resolution Resolution $L = 480 \,\mu m$

[G. Brida et al, Nat. Phot. 4, 227 (2010) PRA 83, 033811 (2011)]

8000 pixels at full resolution Resolution $L = 5\mu m$

[Samantaray N, Ruo-Berchera I, Meda A, Genovese M, Light: Science and Appl. 6, e17005(2017)]



• Advantages of quantum optical states in intereferometric schemes have been studied theoretically in the '80 by Caves.



• Recently their use has been experimentally implemented in the field of gravitational wave detection, where high sensitivity is required.

nature photonics

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Enhanced sensitivity of the LIGO gravitational wave detector by using squeezed states of light

The LIGO Scientific Collaboration*





They are optical states which cannot be described by the semi-classical theory of light.

- They can have properties very far from the coherent states
- Not every quantum state is easly experimentally realizable!

For example, Fock states have fantastic potential properties but their realization for high *n* is extremely challenging.

We focus on two states, easily experimentally implementable with current technologies and used in our labs.

Squeezed vacuum Twin-Beam state (TWB)

Experimentally these states are obtained by non linear optical wave mixing processes: pairs of photons are emitted into degenerate (single-mode squeezing, e.g. squeezed vacuum) or non degenerate (TWB) modes.



Parametric Down Conversion (PDC)



• Photon number statistics:

$$\langle n \rangle = \sinh^2 r$$
 $(\Delta n)^2 = 2\cosh^2 r \sinh^2 r$

• Quadrature operators have uncertainty:

$$X_{1} = \frac{1}{2}(a + a^{+}) \qquad X_{2} = \frac{1}{2i}(a - a^{+})$$
$$[X_{1}, X_{2}] = \frac{i}{2} \qquad \Delta X_{1} \Delta X_{2} \ge \frac{1}{4}$$
$$(AX_{2})^{2} = \frac{1}{4} = \frac{1}{2} = \frac{1}{2} = \frac{1}{4} = \frac{1}{2} = \frac{1}{4} = \frac{1}{4}$$

$$(\Delta X_1)_{sq}^2 = \frac{1}{4}e^{2r} < (\Delta X_1)_{coh}^2$$

$$(\Delta X_2)_{sq}^2 = \frac{1}{4}e^{2r} > (\Delta X_1)_{coh}^2$$



saturates Unc. Relation $\Delta X_1 \Delta X_2 = \frac{1}{4}$

Workshop Quantum Foundations, 29 Nov-1 Dec, Frascati



$$|\Psi^{TWB}\rangle_{12} = S_{12}(\zeta)|0\rangle_{a_1}|0\rangle_{a_2} \quad \text{where} \quad S_{12}(\zeta) = \exp(\zeta a_1^+ a_2^+ - \zeta^* a_1 a_2)$$

• Twin-beam in the Fock basis:

$$|\Psi^{TWB}\rangle_{12} = S_{12}(\zeta)|0\rangle_{a_1}|0\rangle_{a_2} = \sum_{m=0}^{\infty} c_m |m,m\rangle_{a_1a_2}$$

• The two modes are maximally entangled in the photon number:

$$\left\langle \Psi^{TWB} \left| \left(m_{a_1} - m_{a_2} \right)^M \right| \Psi^{TWB} \right\rangle_{a_1 a_2} = 0$$



$$\sigma = \frac{Var(N_1 - N_2)}{\langle N_1 + N_2 \rangle} = 1 - \eta < 1$$

Ideal TWB $\sigma = 0$; Coherent state $\sigma = 1$

 η : system losses

• The two-modes are also correlated in quadratures :

$$\Delta^2(X_1 - X_2) = \frac{e^{-2r}}{2} \qquad \Delta^2(Y_1 + Y_2) = \frac{e^{-2r}}{2}$$









The experimental approach must take into account the importance of **unbiased strategy**, therefore a twomode approach is the most appropriate.



Role of losses for different estimators is deeply discussed in [E. Losero et al. arXiv:1710.09307]



Experimental implementation at INRiM





Toward an experimental quantum gravity

• Several QG theories (string theories, holographic theory, heuristic arguments from black holes,...) predict non-commutativity of position variables at Planck scale.

In particular, according to C. Hogan, phenomenological theory:

 $l_P = 1.6 \cdot 10^{-35} m$ $t_P = 5.4 \cdot 10^{-44} s$ $E_P = 1.2 \cdot 10^{19} GeV$



$$[x_i, x_j] = x_k \epsilon_{ijk} ict_P / \sqrt{4\pi}$$

$$\Delta x_1 \Delta x_2 \ge l_P^2/2$$

holographic uncertainty principle

C. Hogan

- This new quantum uncertainty of space-time induces a slight random wandering of transverse position (called "holographic noise")
- Due to the presence of l_p this noise is negligible at normal scales, therefore an experimental confirmation of this model is extremely challenging.

[C. Hogan, Arxiv: 1204.5948; C. Hogan, Phys. Rev. D 85, 064007 (2012)]



The presence of holographic noise in one interferometers manifests as a noise which affects the position of the beam-splitter.

Its order of magnitude can be estimated as:

$$\Delta X \sim \sqrt{L \ l_P}$$

- an interferometer is able to «accumulate» the hologrphic noise, $\Delta X \propto \sqrt{L}$
- Due to *l_P* the holographic noise is covered by other noises!

To overpass this problem Hogan proposed the so called «holometer»

The holometer is a system of two Michelson interferometers placed close to each other, whose arms can be placed in two different configurations.

Experimentally realized at Fermilab, $L \sim 40m$





Holographic noise in the holometer



@Fermilab HOLOMETER : principle of operation

- holographic noise \rightarrow correlated (in II configuration)
- other sources of noise → uncorrelated (in [⊥] and II both configuration); their contribution vanish over a sufficiently long integration time.



Performing a measurement of cross correlation between the two interferometers output we can detect, with current technology, the holographic noise! ($L \sim 40 m$)

Control measurement can be performed «turned off» HN correlation just by separating the space-time volumes (⊥ conf.)





Can holometer performances improve using quantum light?







Squeezed vacuum states in input



- Quantum state: $|\xi\rangle_{a_1}\otimes|\xi\rangle_{a_2}=S_{a_1}(\xi)S_{a_2}(\xi)|0\rangle_{a_1}\otimes|0\rangle_{a_2}$
- Measured observable: $\widehat{C} = \{Y_1(\phi_1) \mathscr{E}[Y_1]\}\{Y_2(\phi_2) \mathscr{E}[Y_2]\}$
- Phase covariance est.: $\mathscr{E}_{\parallel}[\delta\phi_1\delta\phi_2] \approx \frac{\mathscr{E}_{\parallel}[Y_1Y_2] \mathscr{E}_{\perp}[Y_1Y_2]}{\left\langle \partial_{\phi_1,\phi_2}^2 Y_1(\phi_0)Y_2(\phi_0) \right\rangle}$
- **Principle:** each squeezed state reduces the photon noise in each interferometer (still uncorrelated!), then the covariance performs statistical cancellation (like the Fermilab Holometer)





Twin-Beam state in input



• Quantum state:
$$|\Psi(\lambda)\rangle_{a_1,a_2} = \frac{1}{\sqrt{1+\lambda}} \sum_{m=0}^{\infty} \left(e^{i\theta} \sqrt{\frac{\lambda}{1+\lambda}} \right)^m |m,m\rangle_{a_1,a_2}$$

• Measured observable:
$$\widehat{C}(\phi_1,\phi_2) = [N_1(\phi_1) - N_2(\phi_2)]^2 = N_1^2 + N_2^2 - 2N_1N_2$$

• Phase covariance est.:
$$\mathscr{E}_{\parallel}[\delta\phi_1\delta\phi_2] \approx \frac{\mathscr{E}_{\parallel}[N_1N_2] - \mathscr{E}_{\perp}[N_1N_2]}{\left\langle \partial_{\phi_1,\phi_2}^2 N_1(\phi_0) N_2(\phi_0) \right\rangle}$$

• **Principle:** noise in each interferometer is higher than shot-noise but they are perfectly correlated an can be subtracted (different from a statistical cancellation)



The performance of quantum light in the experimental situation strictly depends on the losses

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VRiM



Uncertainty of phase correlation for TWB can be arbitrarily small (not depends on the intensity) thanks to entanglement! But it is very sensitive to losses.



Scheme of one of the two Michelson interferometers, developed in strict collaboration with the Technical University of Denmark (DTU)







Preliminary measurements





- When using classical light the shot-noise limits the sensitivity of the measurement.
- Quantum light, in particular TWB and squeezed vacuum, permits to overcome this limit.
- In absorption measurement we demonstrate how TWB can lead in the ideal case to the UQL. Moreover the best sensitivity per photon never achieved has been experimentally demonstrated.
- In the holometer the use of quantum light can increase the sensitivity. This could offer the possibility of experimentally testing phenomenological models of quantum gravity.

THANK YOU!



