



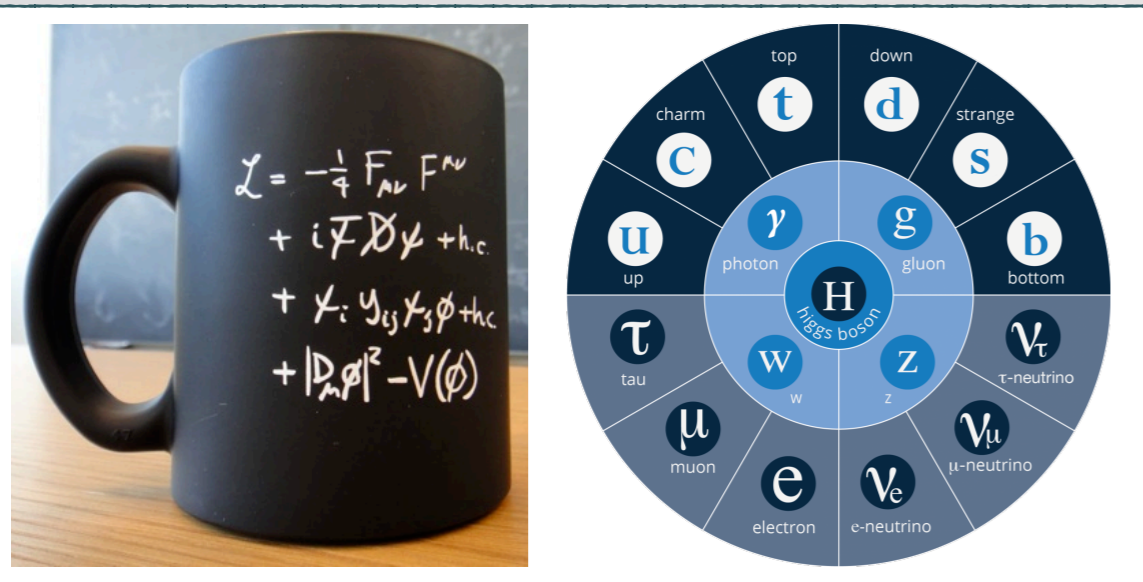
A personal short review on
**Experimental
Neutrino Physics**

Marco Pallavicini
Università di Genova and INFN

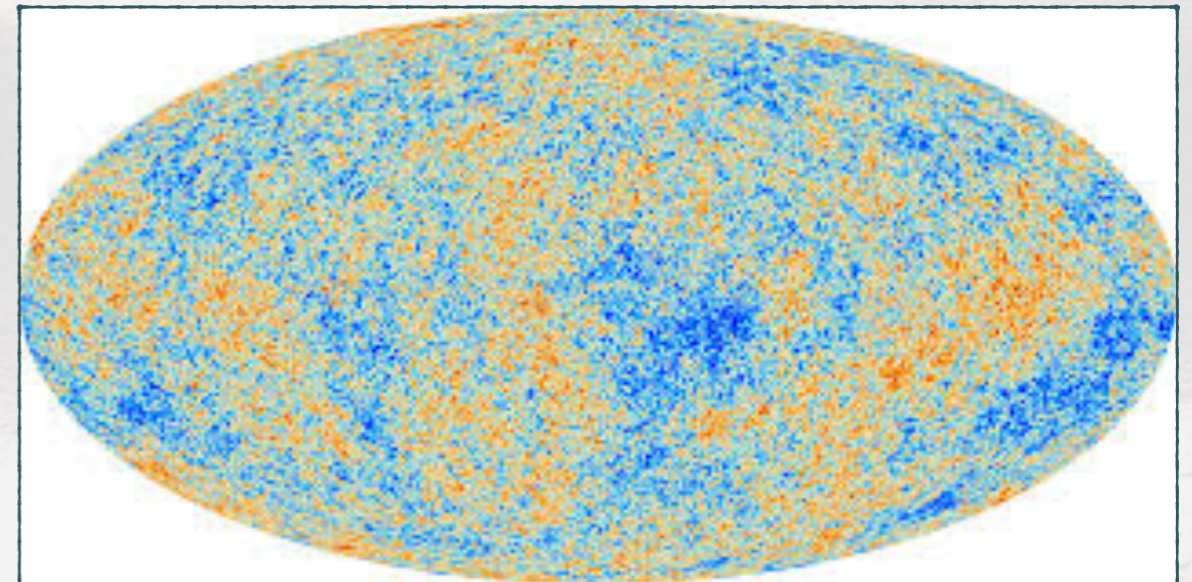
FPCapri2018
Capri, June 8-10, 2018

- Neutrino physics (as all astroparticle physics) is a bridge between ***two standard models***

Standard Model (of particles and fields)



Standard Model (Λ CDM Cosmology)



- Beautiful, but certainly **incomplete**

- **Neutrinos:** m , $\nu = \bar{\nu}$?, δ_D ($\delta_{M1,2}$), hierarchy
- No **dark matter candidates**
 - θ_{QCD} ? Super-symmetry? **Steriles**? BSM
- **No inflation** (connected with Higgs?)
- **Gravity is classical only**
 - Quantum gravity? Different paradigm?
- **CPV insufficient**
 - Can **neutrinos** help?

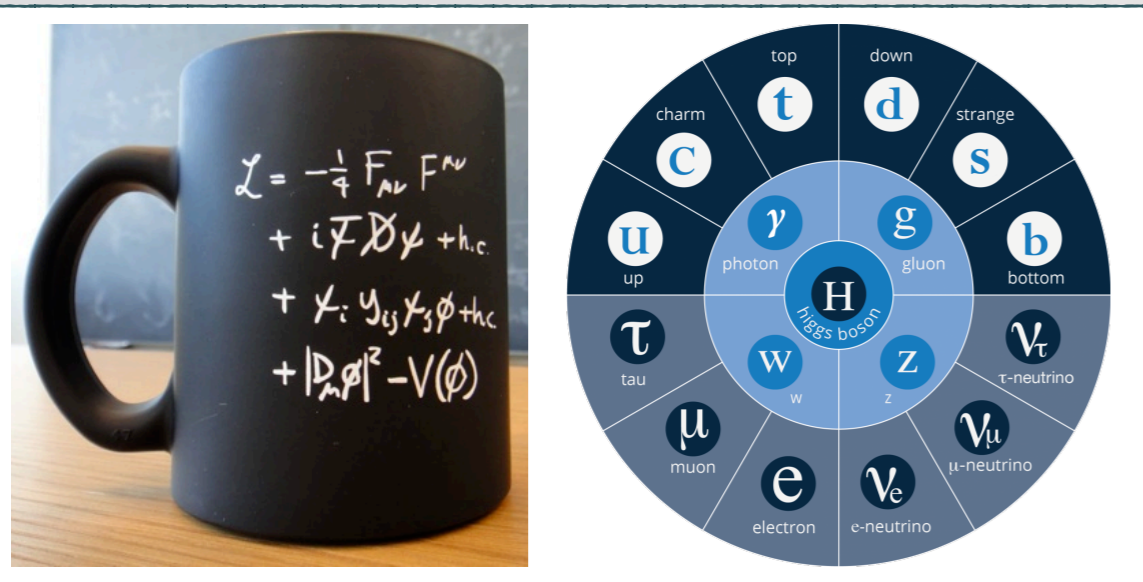
- Beautiful, but somewhat **ad hoc**

- $\sim 95\%$ of energy density is “dark”
 - **“Matter”**
 - Which role for neutrinos?
 - **“Energy”**
 - A constant or a field? Why so small?
- Inflation is postulated
- Sakharov conditions imposed but not explained. **Which role for leptons?**

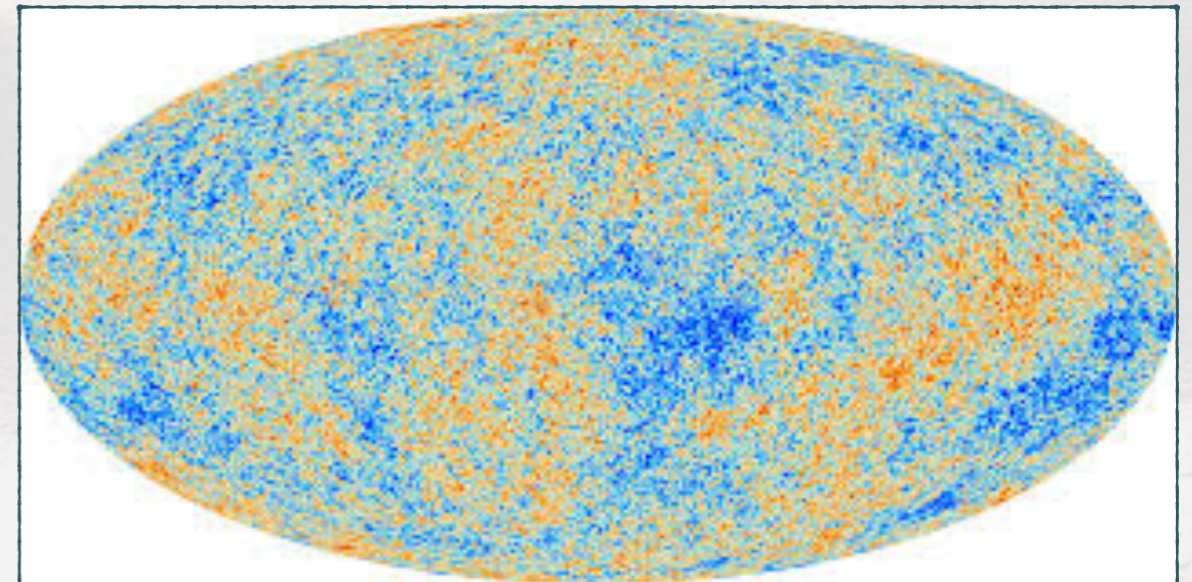


- Neutrino physics (as all astroparticle physics) is a bridge between ***two standard models***

Standard Model (of particles and fields)



Standard Model (Λ CDM Cosmology)



- Beautiful, but certainly **incomplete**

• **Neutrinos:** m , $\nu = \bar{\nu}$?, δ_D ($\delta_{M1,2}$), hierarchy

• No **dark matter candidates**

• θ_{QCD} ? Super-symmetry? **Steriles?** BSM

• **No inflation** (connected with Higgs?)

• **Gravity is classical only**

• Quantum gravity? Different paradigm?

• **CPV insufficient**

• Can **neutrinos** help?

- Beautiful, but somewhat **ad hoc**

• $\sim 95\%$ of energy density is “dark”

• **“Matter”**

• Which role for **neutrinos**?

• **“Energy”**

• A constant or a field? Why so small?

• Inflation is postulated

• Sakharov conditions imposed but not explained. **Which role for leptons?**

Astrophysics & Cosmology



SN1987A

SuperNovae
G.W. and ν
CvB
High Energy ν
Leptogenesis

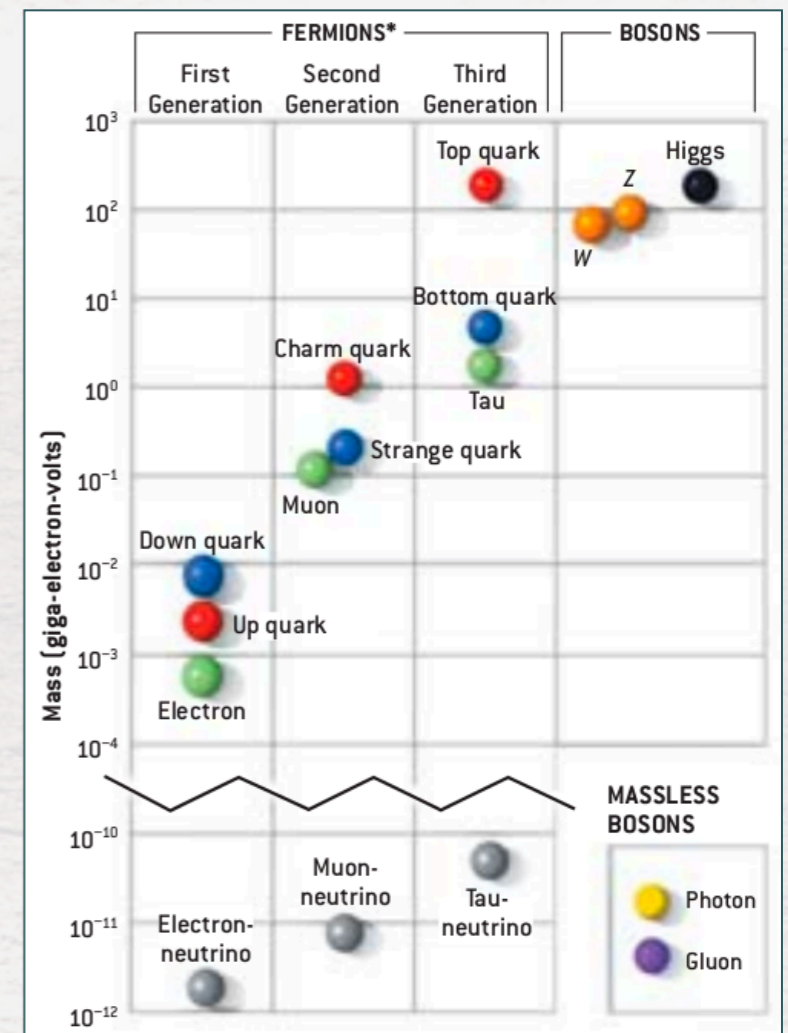


New Physics

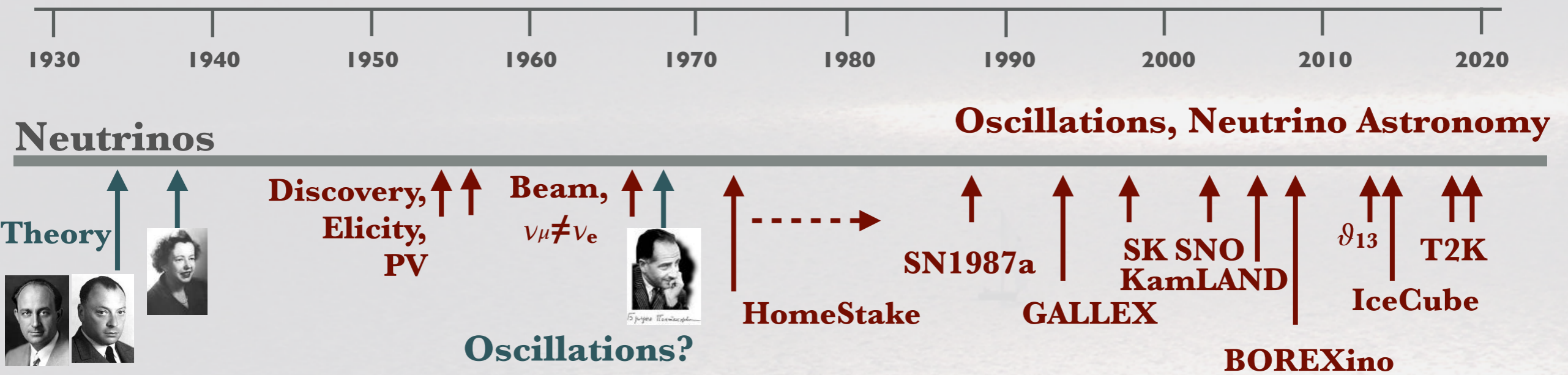
$0\nu\beta\beta$
Mass scale
Sterile ν
Dark Matter

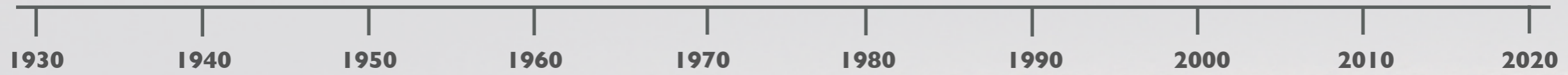
Standard Model

CP Violation
Majorana vs Dirac
Cross sections
Coherent scattering



(image: G. Kane, Sci. Am.)





Neutrinos

Oscillations, Neutrino Astronomy



$$|\Delta m^2| = 2.47 \pm 0.04 \cdot 10^{-3} \text{ eV}^2$$

$$\theta_{23} = 47.5 \pm 3.2^\circ$$

$$\delta_D = ? \text{ } (-\pi/2 ?)$$

$$\theta_{13} = 8.56 \pm 0.15^\circ$$

$$\delta m^2 = 7.40 \pm 0.21 \cdot 10^{-5} \text{ eV}^2$$

$$\theta_{12} = 33.6 \pm 0.77^\circ$$

$$\mathbf{U} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta_D} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta_D} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_1} & 0 \\ 0 & 0 & e^{i\alpha_2} \end{pmatrix}$$

**Atmospheric
Accelerators LBL
L ~ 700 km**

**Reactors L ~ 1 km
LBL L ~ 200 km**

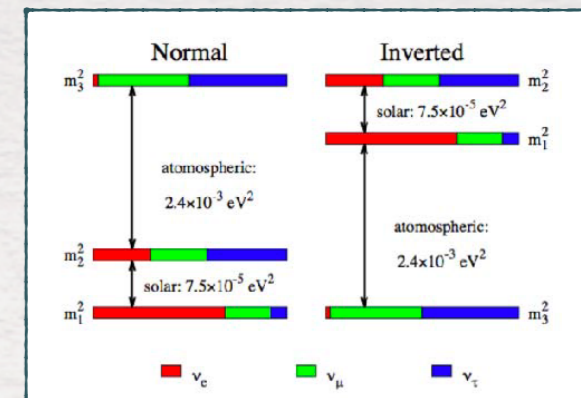
**Solar
Reactors
L ~ 200 km**

$0\nu\beta\beta$

Next generation (JUNO, T2HK, DUNE) has sufficient precision for global fits to almost all parameters

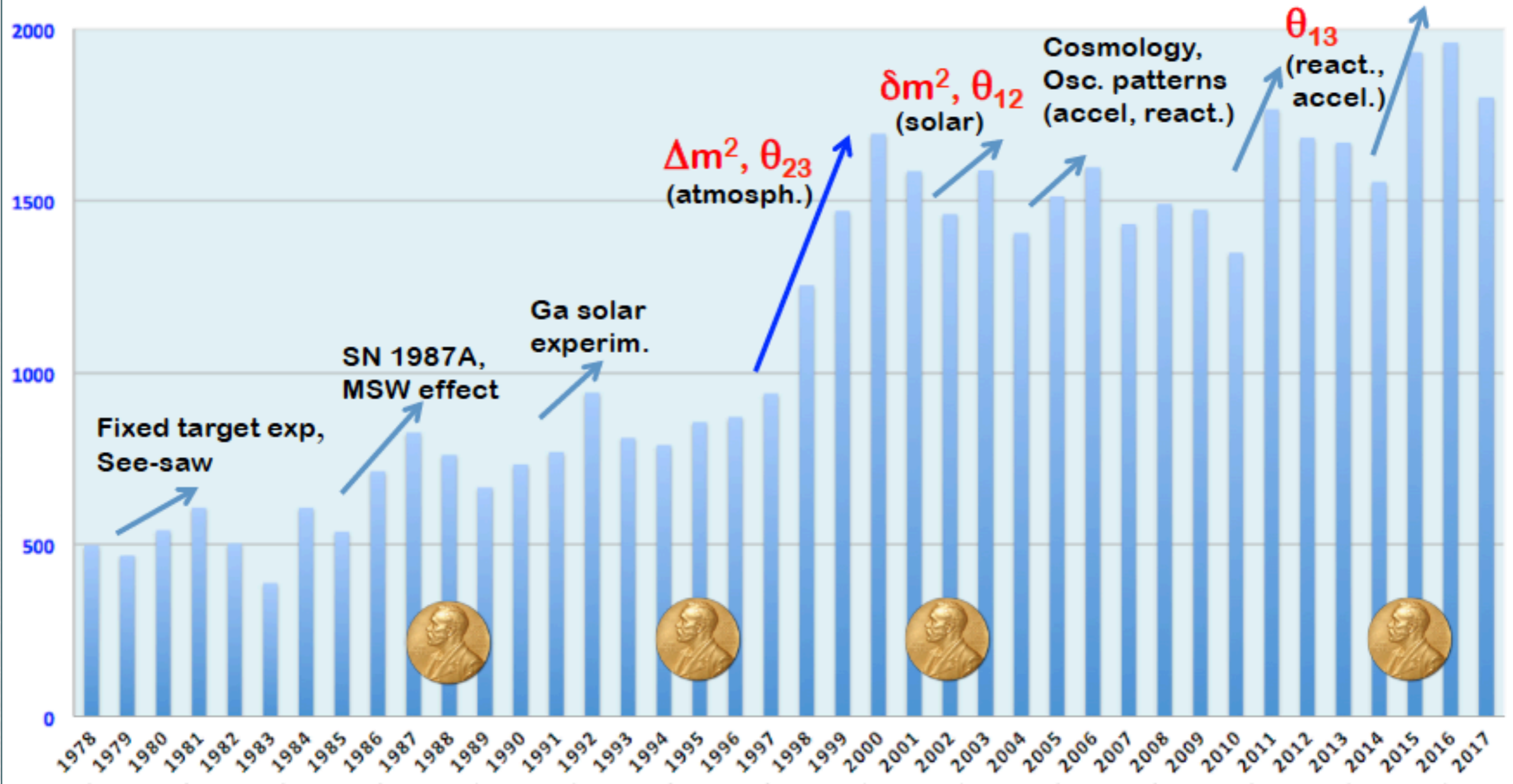
Combined T2K, Nova, etc analysis may yield an early “detection” of CP violation phase δ_D

$$\mathbf{V} \neq \overline{\mathbf{V}} ?$$



N. of #neutrino# preprints per year (1978-2018) from **INSPIRE**

Pattern reflects breakthroughs and peaks of interest...



From: E. Lisi Neutrino 2018

- Example: **NuFIT 3.2** (www.nu-fit.org) updated to **2018**

					NuFIT 3.2 (2018)
	Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 4.14$)		Any Ordering
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
$\sin^2 \theta_{12}$	$0.307^{+0.013}_{-0.012}$	0.272 \rightarrow 0.346	$0.307^{+0.013}_{-0.012}$	0.272 \rightarrow 0.346	0.272 \rightarrow 0.346
$\theta_{12}/^\circ$	$33.62^{+0.78}_{-0.76}$	31.42 \rightarrow 36.05	$33.62^{+0.78}_{-0.76}$	31.43 \rightarrow 36.06	31.42 \rightarrow 36.05
$\sin^2 \theta_{23}$	$0.538^{+0.033}_{-0.069}$	0.418 \rightarrow 0.613	$0.554^{+0.023}_{-0.033}$	0.435 \rightarrow 0.616	0.418 \rightarrow 0.613
$\theta_{23}/^\circ$	$47.2^{+1.9}_{-3.9}$	40.3 \rightarrow 51.5	$48.1^{+1.4}_{-1.9}$	41.3 \rightarrow 51.7	40.3 \rightarrow 51.5
$\sin^2 \theta_{13}$	$0.02206^{+0.00075}_{-0.00075}$	0.01981 \rightarrow 0.02436	$0.02227^{+0.00074}_{-0.00074}$	0.02006 \rightarrow 0.02452	0.01981 \rightarrow 0.02436
$\theta_{13}/^\circ$	$8.54^{+0.15}_{-0.15}$	8.09 \rightarrow 8.98	$8.58^{+0.14}_{-0.14}$	8.14 \rightarrow 9.01	8.09 \rightarrow 8.98
$\delta_{CP}/^\circ$	234^{+43}_{-31}	144 \rightarrow 374	278^{+26}_{-29}	192 \rightarrow 354	144 \rightarrow 374
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.40^{+0.21}_{-0.20}$	6.80 \rightarrow 8.02	$7.40^{+0.21}_{-0.20}$	6.80 \rightarrow 8.02	6.80 \rightarrow 8.02
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.494^{+0.033}_{-0.031}$	$+2.399 \rightarrow +2.593$	$-2.465^{+0.032}_{-0.031}$	$-2.562 \rightarrow -2.369$	$[+2.399 \rightarrow +2.593]$ $[-2.536 \rightarrow -2.395]$

Phantom of the Opera: sterile neutrinos

Dirac vs Majorana ($\nu \neq \bar{\nu}$?) **$0\nu\beta\beta$**

U_{PMNS} unitary?

$\delta_{CP} \neq 0$?

$\Delta m^2 > 0$?

ϑ_{23} maximal? Octant?

OSCILLATIONS

Absolute Mass scale

CNO from the Sun

Astrophysics

Spectrometers, μ Bolometers, EUCLID

BOREXino

IceCUBE, KM3Net

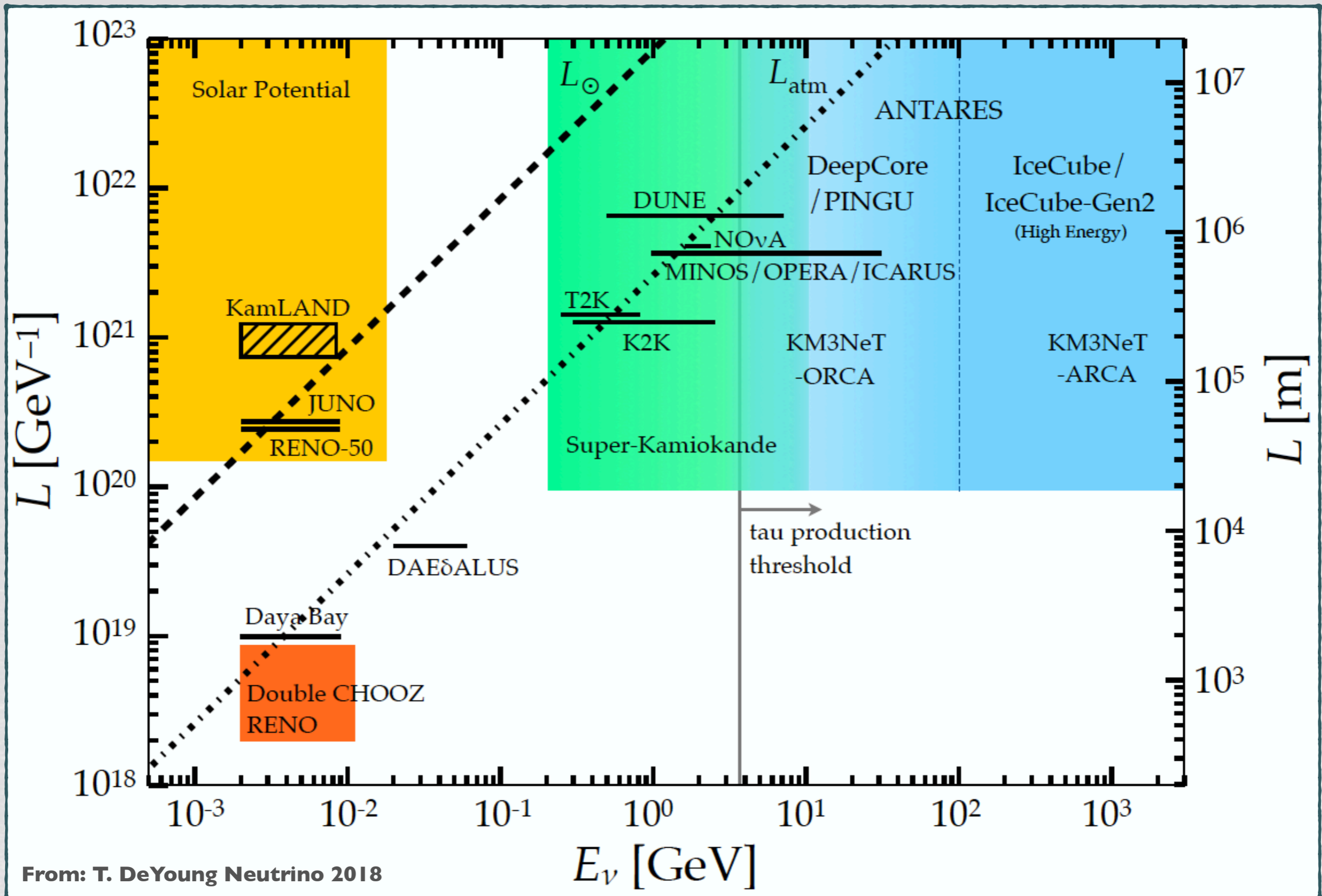
Multi-messenger (GW, photons) **VIRGO-LIGO + Astronomy**

$C\nu B$

R&D for PTolemy, Euclid, CMB fits

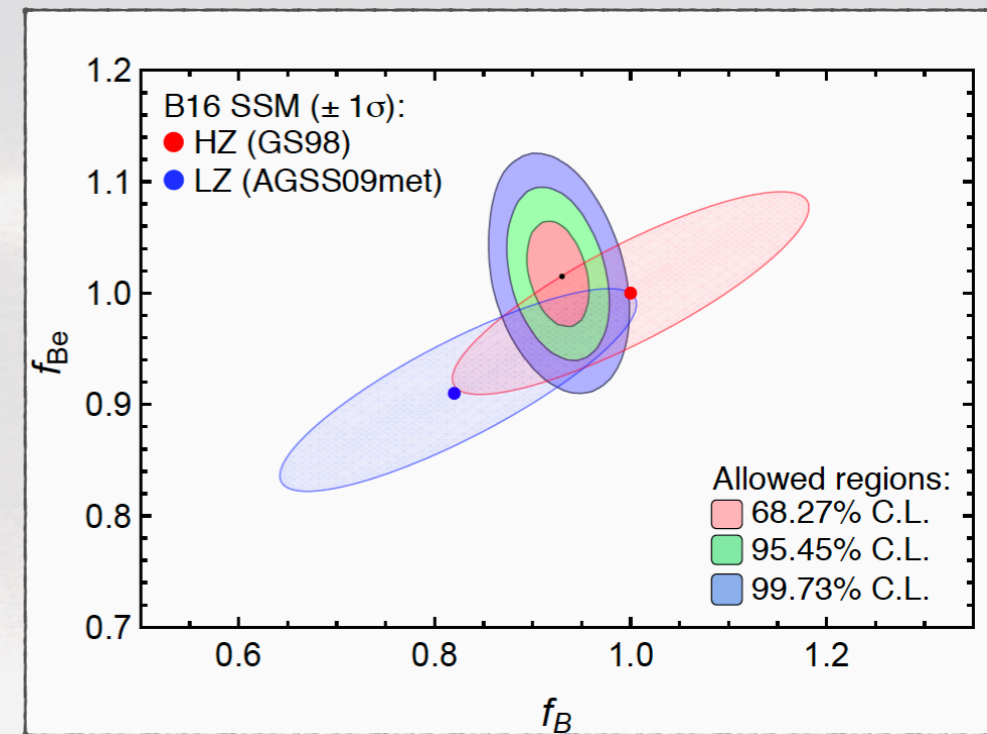
SN (pulse and relics)

Borexino, LVD, JUNO, SK, HK, DUNE



From: T. DeYoung Neutrino 2018

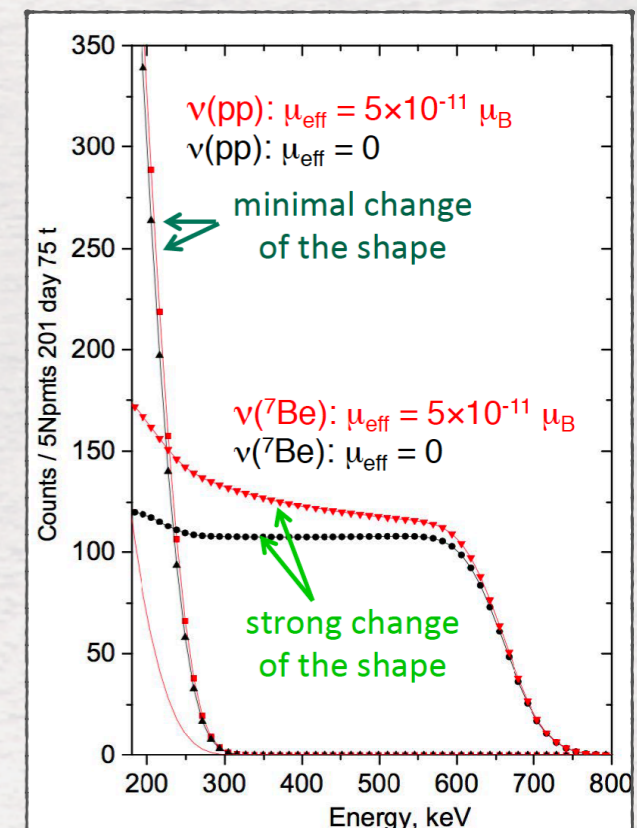
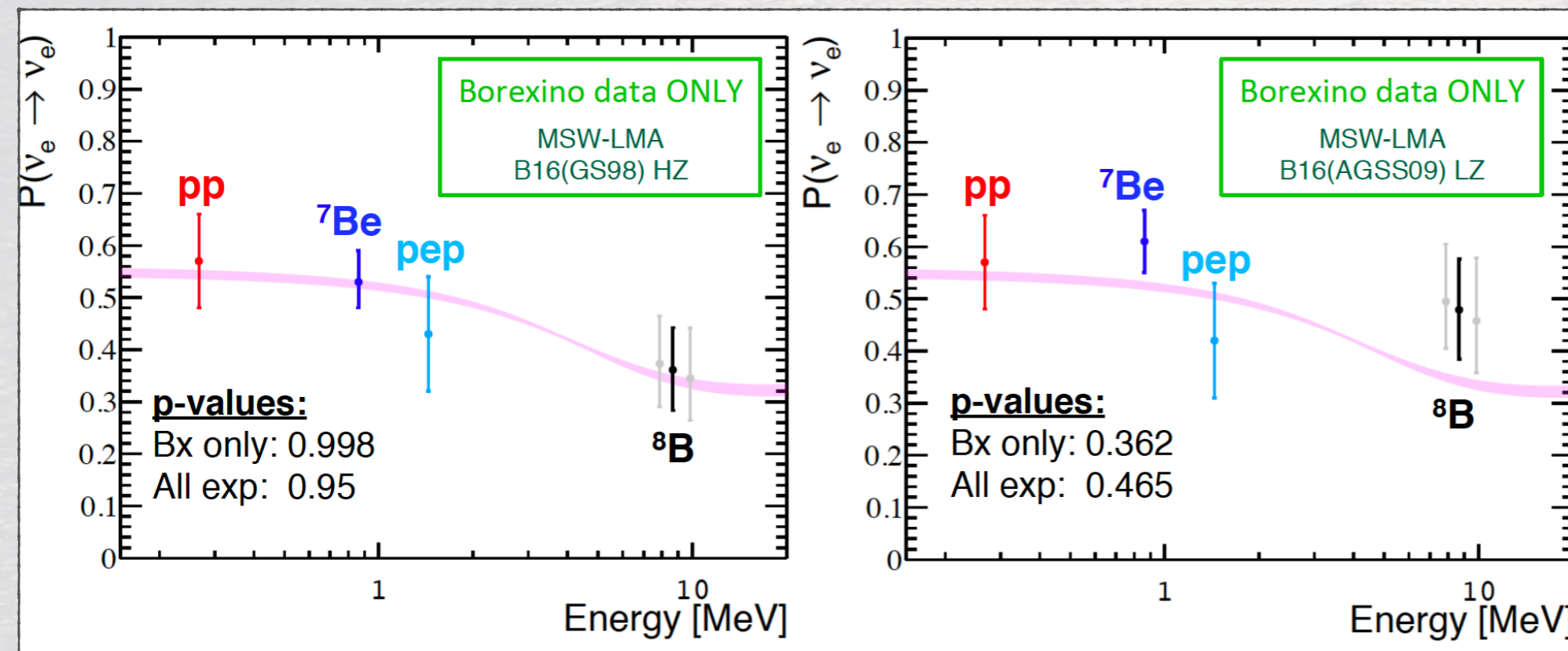
- **2017:** final **Borexino** results on **pp cycle**
 - Complete precision measurement of *all components of the chain*
 - Precision Test of Standard Solar Model
 - Evidence of MSW on P_{ee}
 - **First hint on solar metallicity**
 - HZ favoured at 2σ



HZ

LOW Z

Magnetic Moment



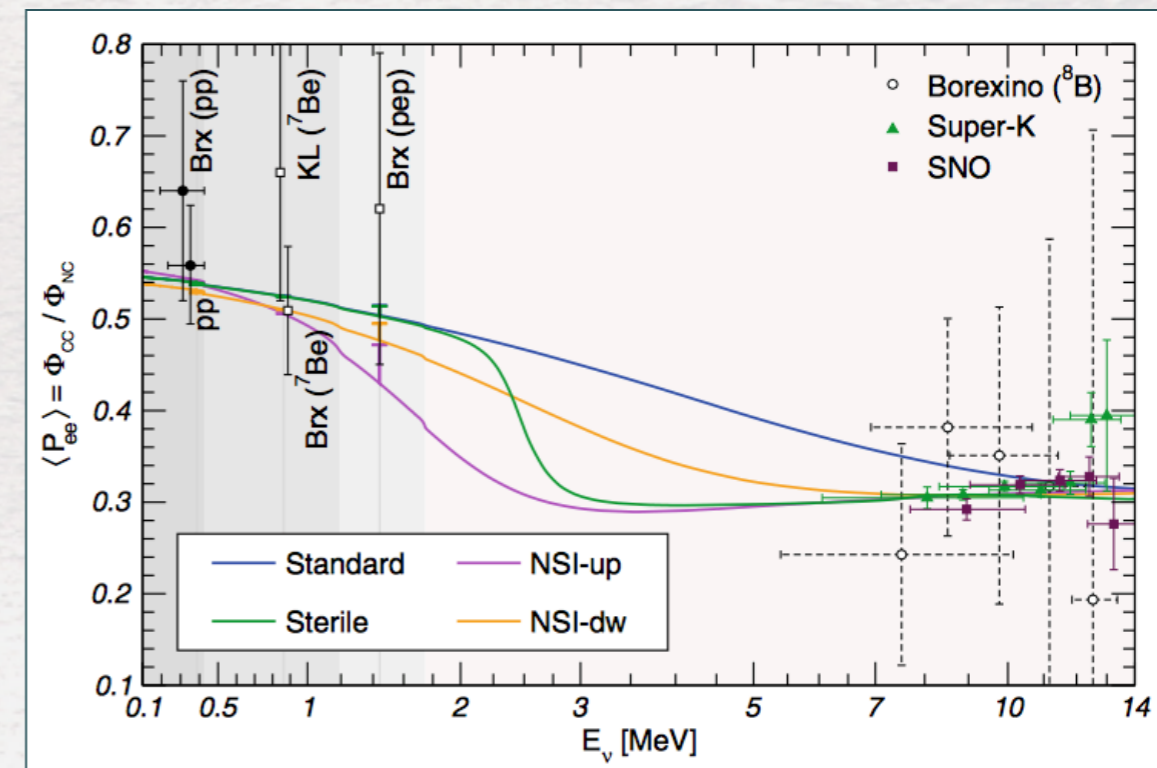
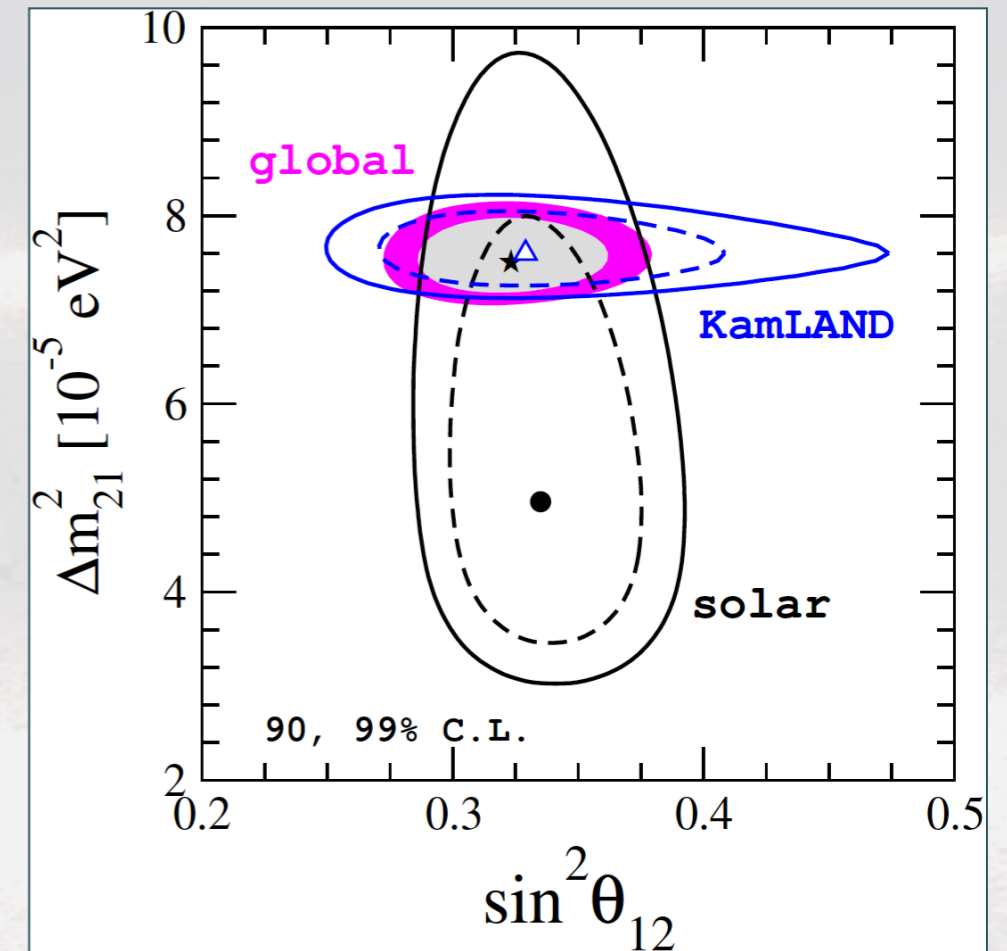
- Fit parameters measured by **KamLAND** at reactors and by global fit to **solar neutrino experiments** are in **small tension**

- 2σ tension on δm^2

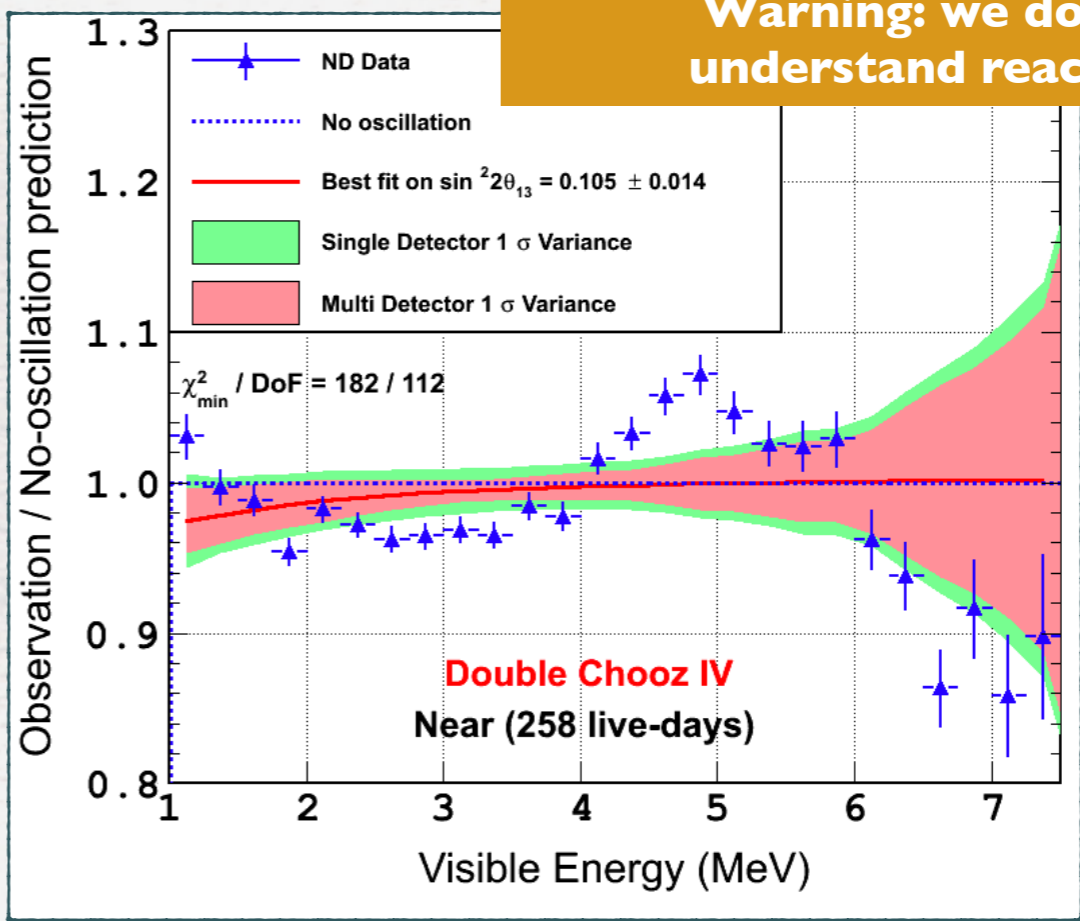
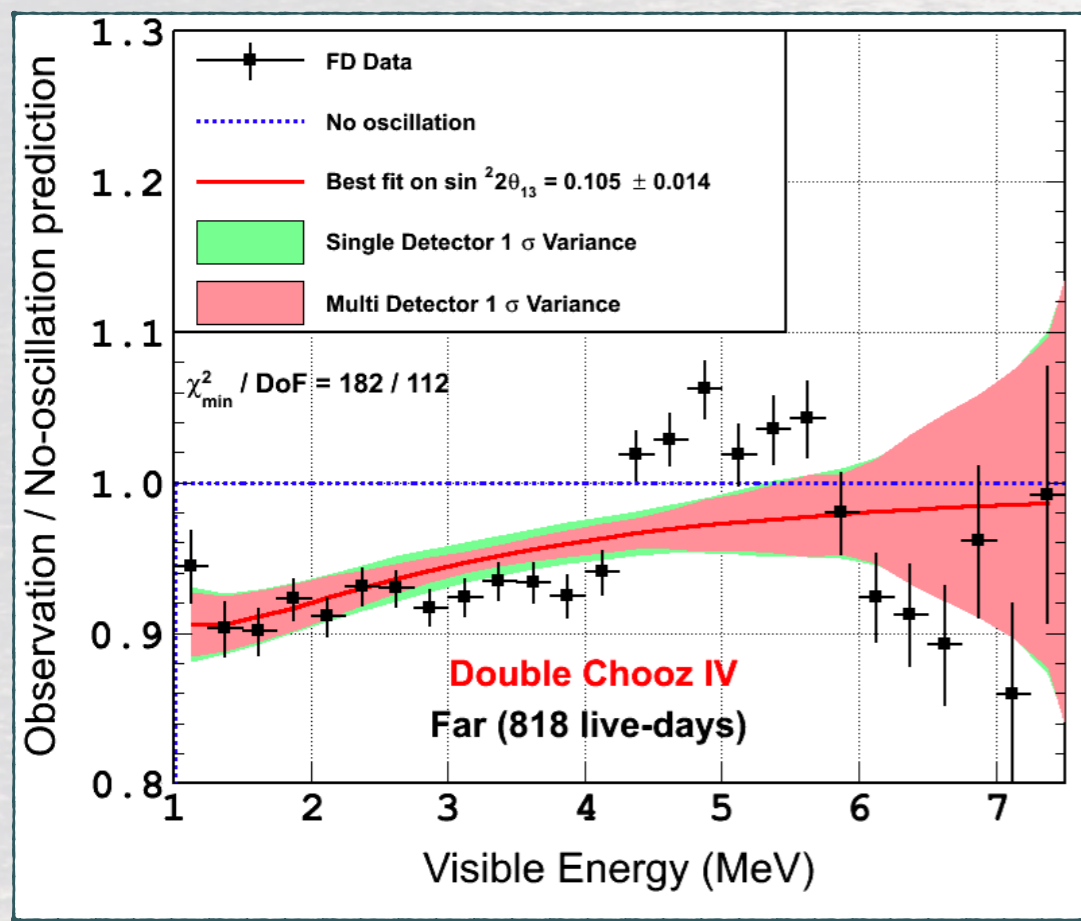
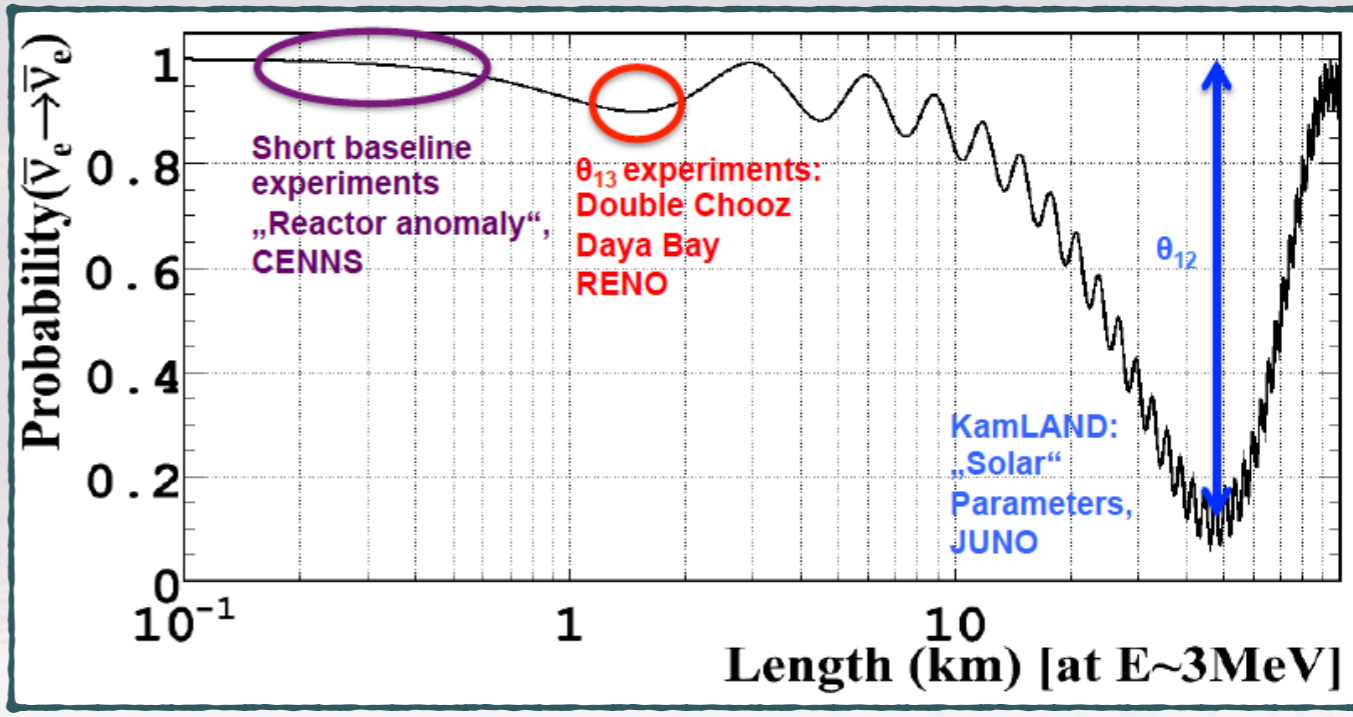
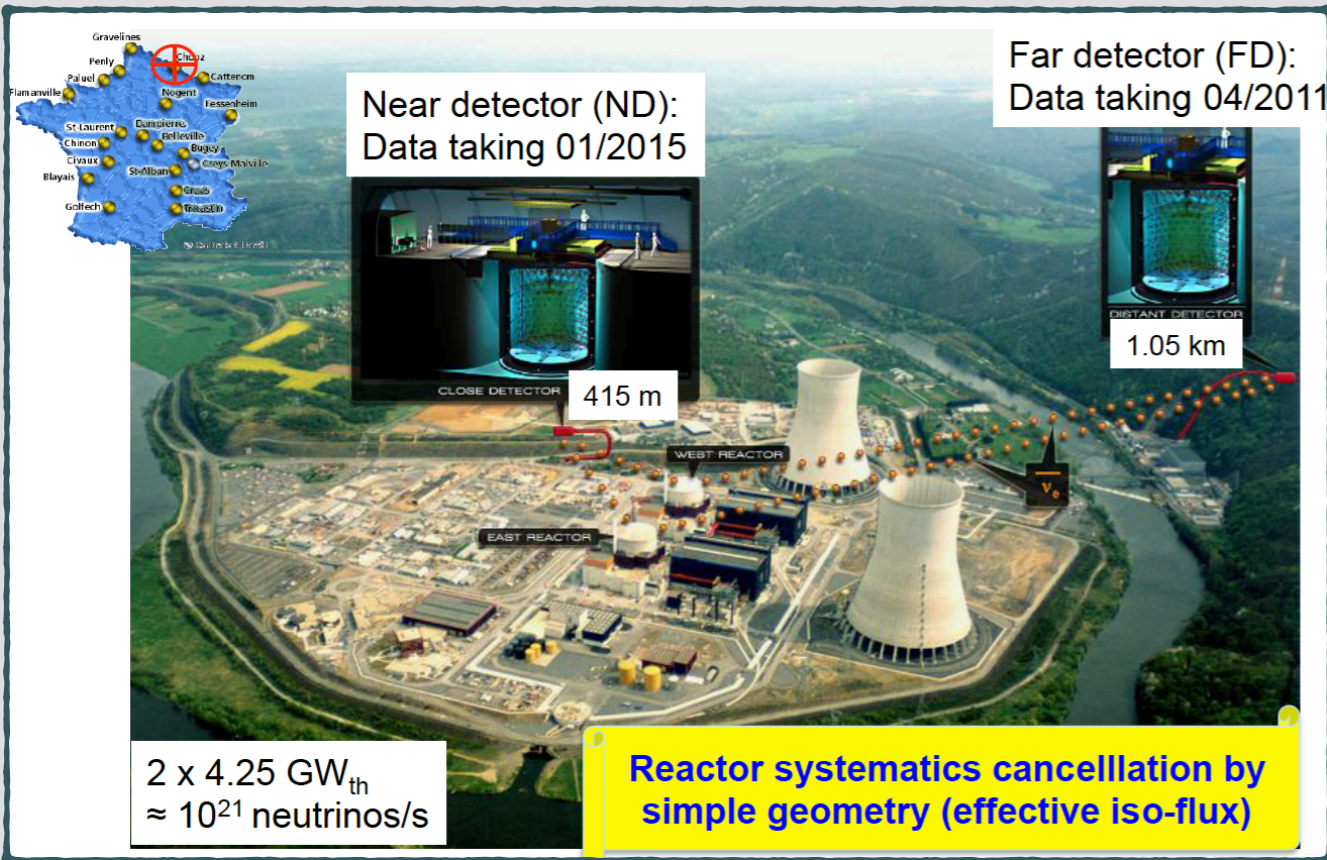
- KamLAND value would give a **steep upturn** and smaller **day-night effect** in ^8B

- Better measurements are needed both at reactors (**JUNO**) and on solar neutrinos (**HK ?**)

- N.S.I. with $\epsilon \sim 0.3$ can improve the fit



Maltoni & Smirnov - Eur. Phys. J. A (2016) 52: 87

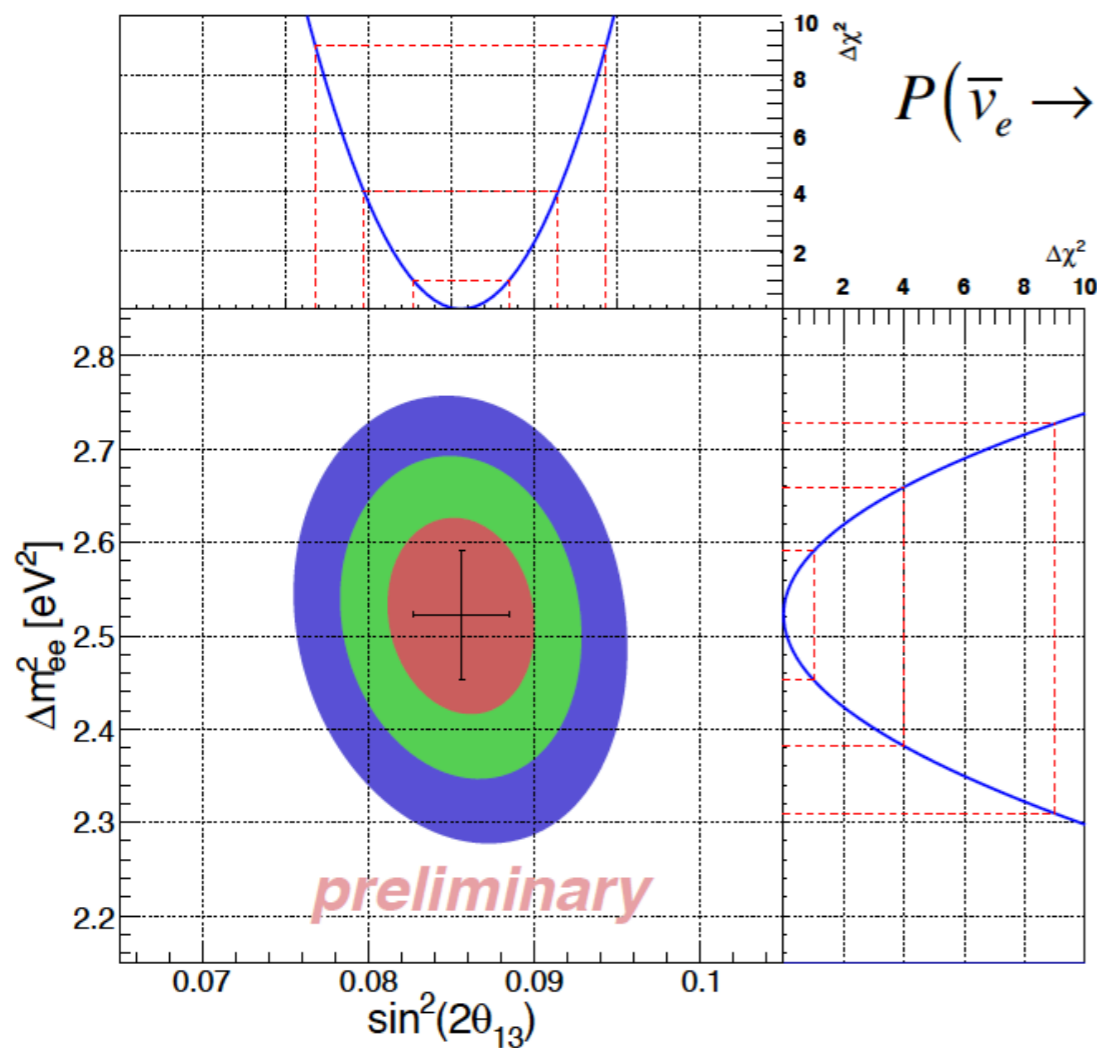


Warning: we do not understand reactors!

Neutrino 2018 Results with 1958 Days

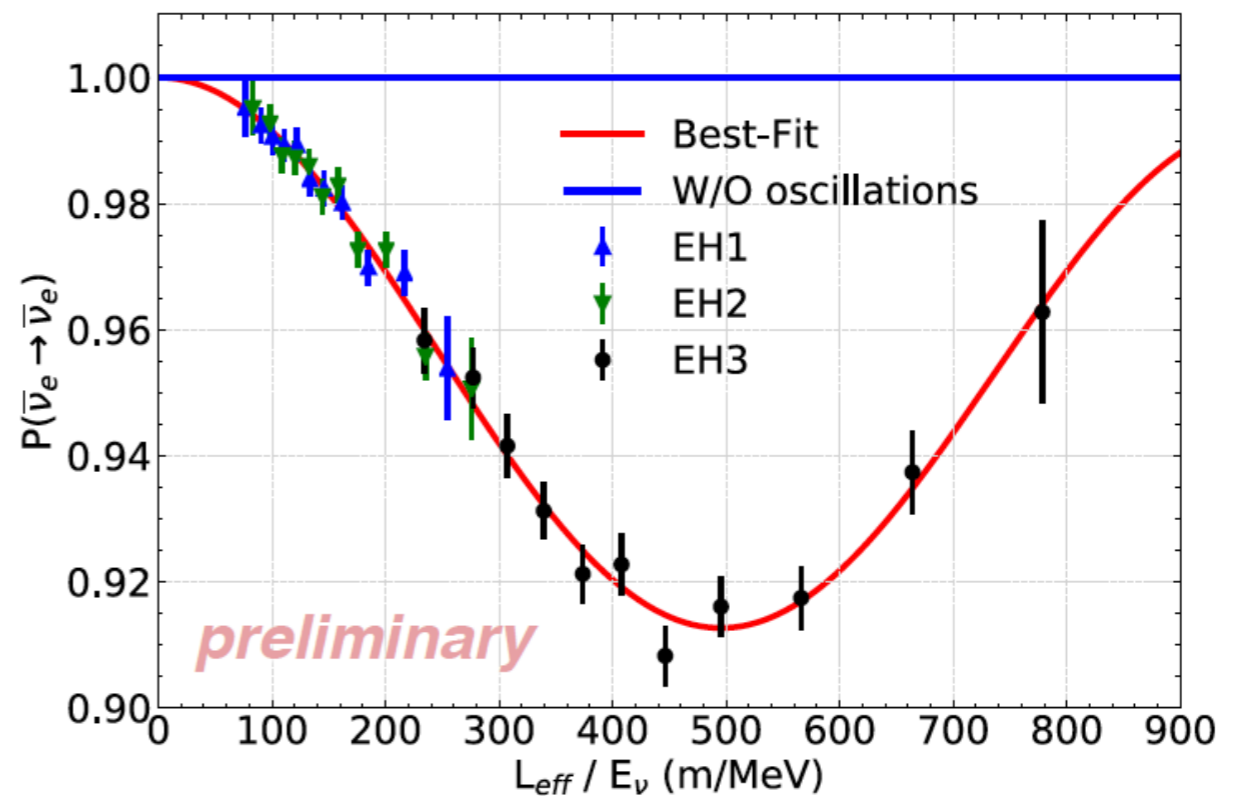


- Measure $\sin^2 2\theta_{13}$ and $|\Delta m_{ee}^2|$ to **3.4%** and **2.8%** respectively



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \frac{1.267 \Delta m_{ee}^2 L}{E} - \text{solar term}$$

effective mass splitting



results with 1958 days →

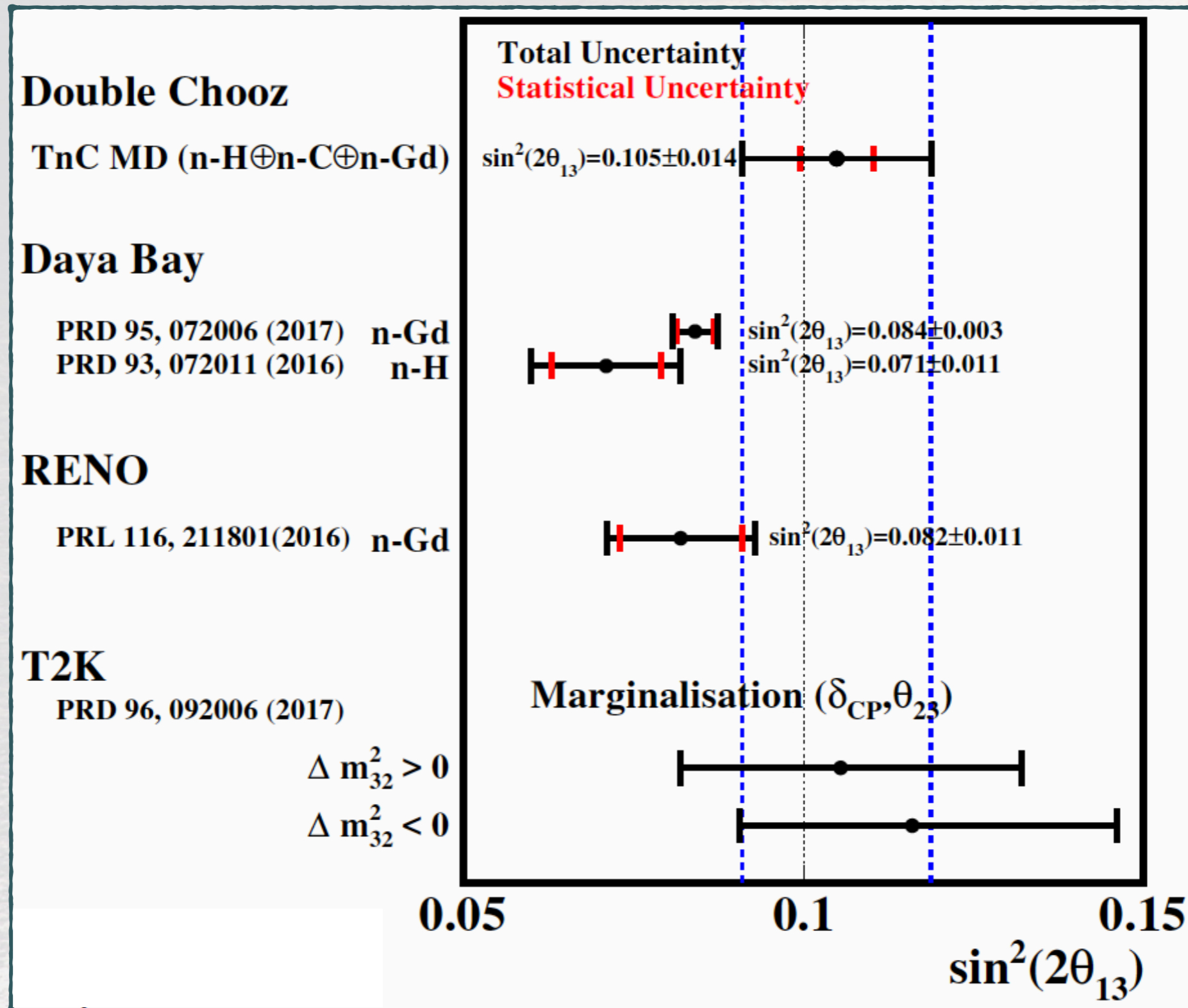
$$\sin^2 2\theta_{13} = 0.0856 \pm 0.0029$$

$$|\Delta m_{ee}^2| = (2.52 \pm 0.07) \times 10^{-3} \text{ eV}^2$$

The statistical uncertainty contributes about 60% (50%) of the total θ_{13} (Δm_{ee}^2) uncertainty.

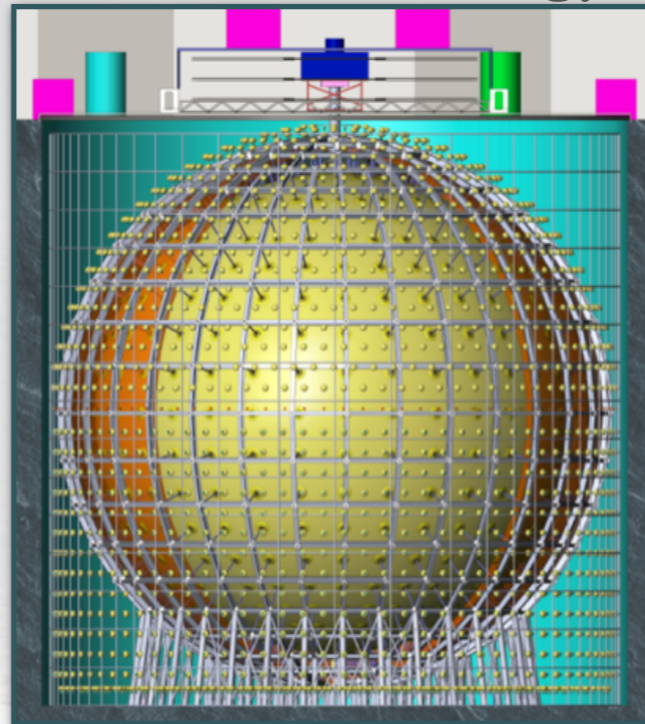
- Reactor experiments (disappearance) or T2K (appearance)

Neutrino 2018

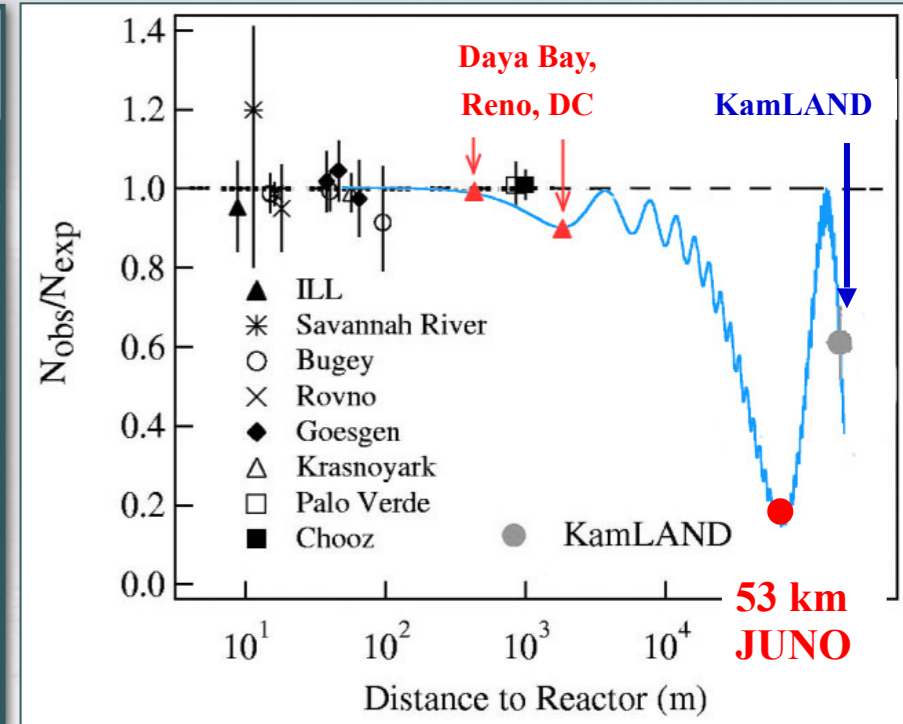


- **JUNO**: coming soon
 - Precision measurement of angles
 - **Unitarity, Hierarchy**
- Relevant INFN investment
 - **Purification** (Borexino legacy)
 - **Electronics PMTs and Tracker**

20 kt LS, 3% energy



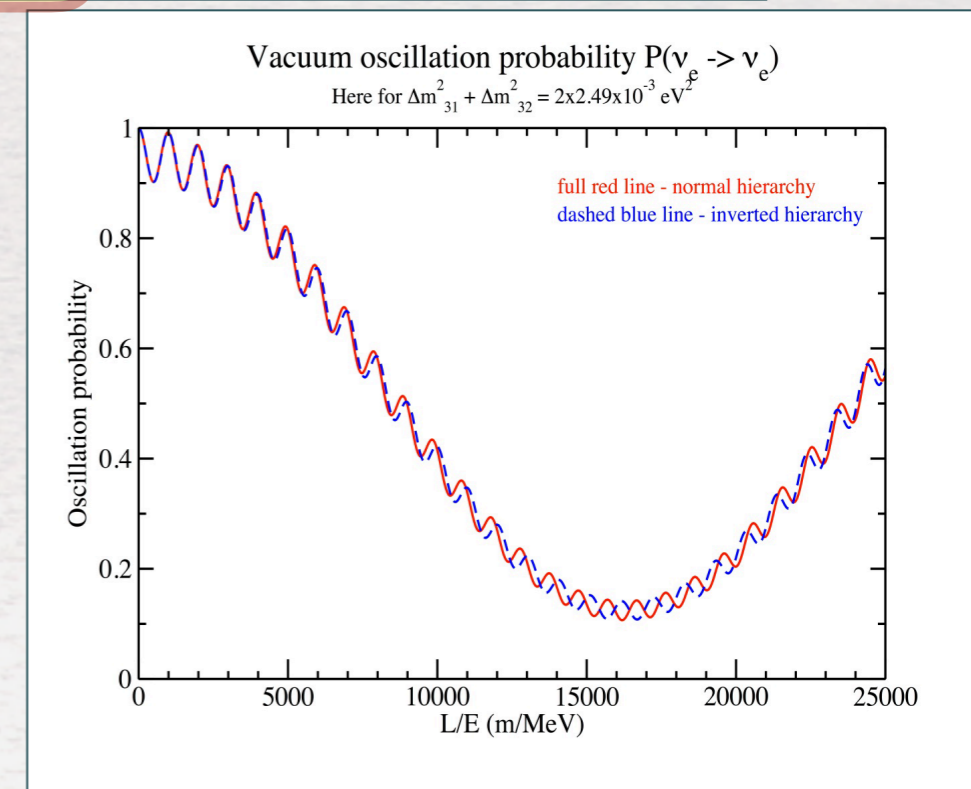
“Medium” Baseline



$$\begin{aligned}
 P(\bar{\nu}_e \rightarrow \bar{\nu}_e) &= 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta m_{21}^2 \frac{L}{4E} - \sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \Delta m_{31}^2 \frac{L}{4E} + \sin^2 \theta_{12} \sin^2 \Delta m_{32}^2 \frac{L}{4E} \right) \\
 &\approx 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta m_{21}^2 \frac{L}{4E} - \sin^2 2\theta_{13} \sin^2 \Delta m_{ee}^2 \frac{L}{4E}, \text{ for } \Delta m_{12}^2 \ll \Delta m_{32}^2
 \end{aligned}$$

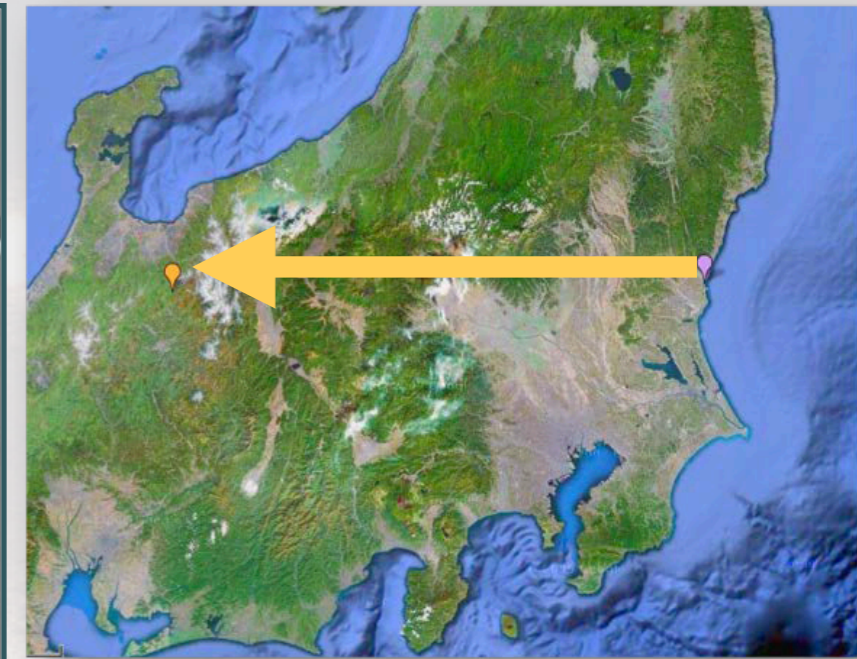
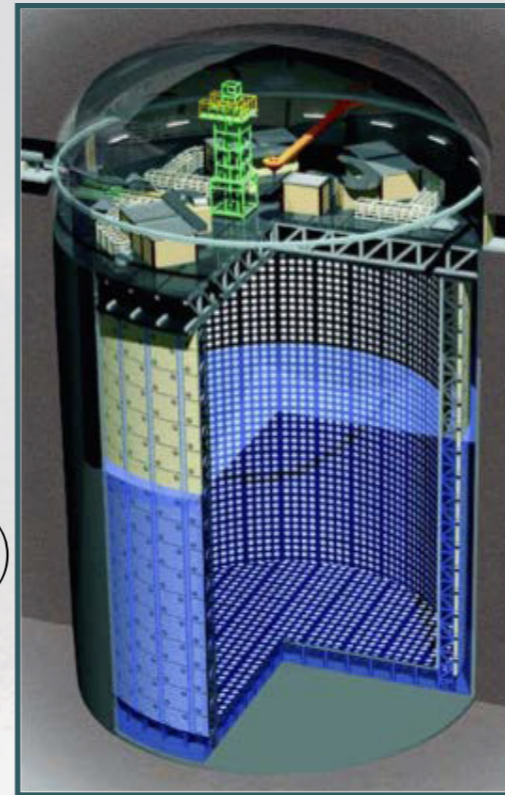
- Very challenging detector resolution, accuracy and stability

$$\begin{aligned}
 \Delta m_{31}^2 &= \Delta m_{32}^2 + \Delta m_{21}^2 \\
 \text{NH: } |\Delta m_{31}^2| &= |\Delta m_{32}^2| + |\Delta m_{21}^2| \\
 \text{IH: } |\Delta m_{31}^2| &= |\Delta m_{32}^2| - |\Delta m_{21}^2|
 \end{aligned}$$



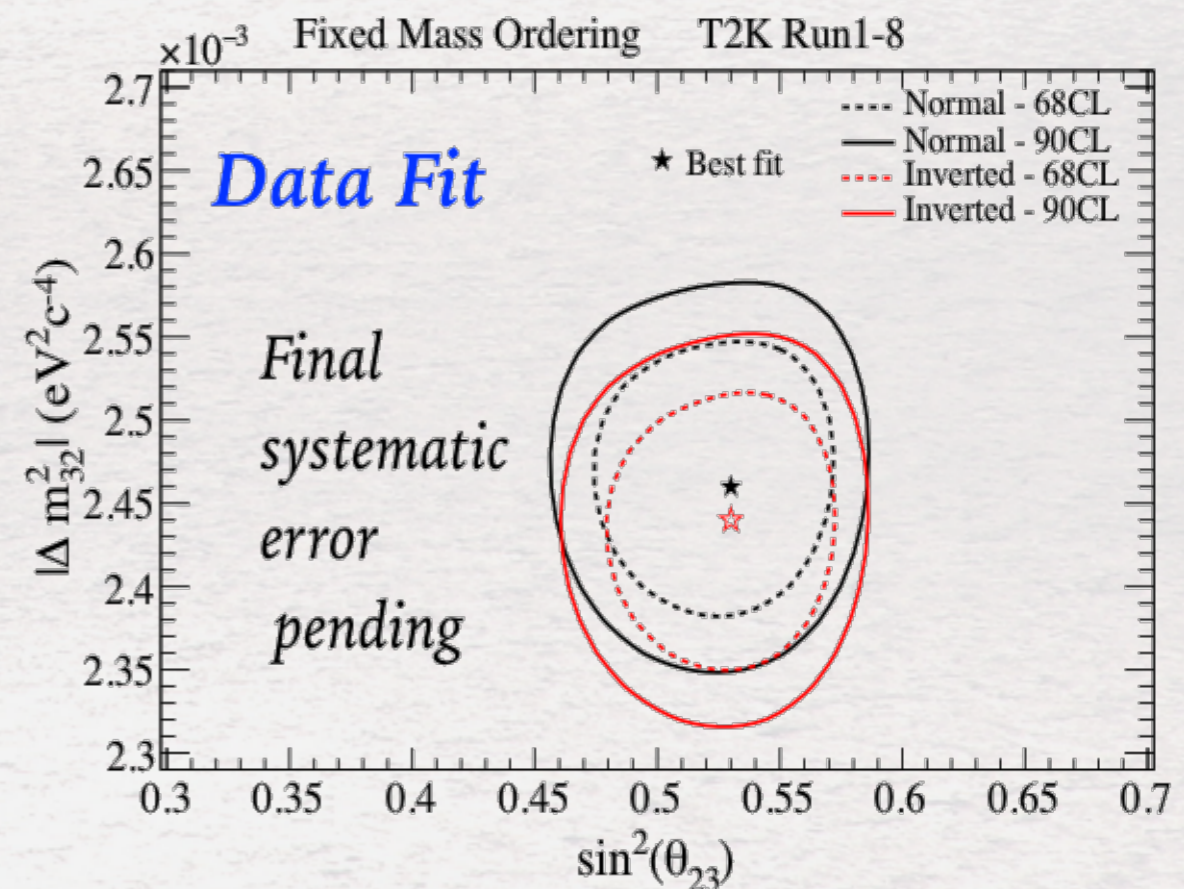
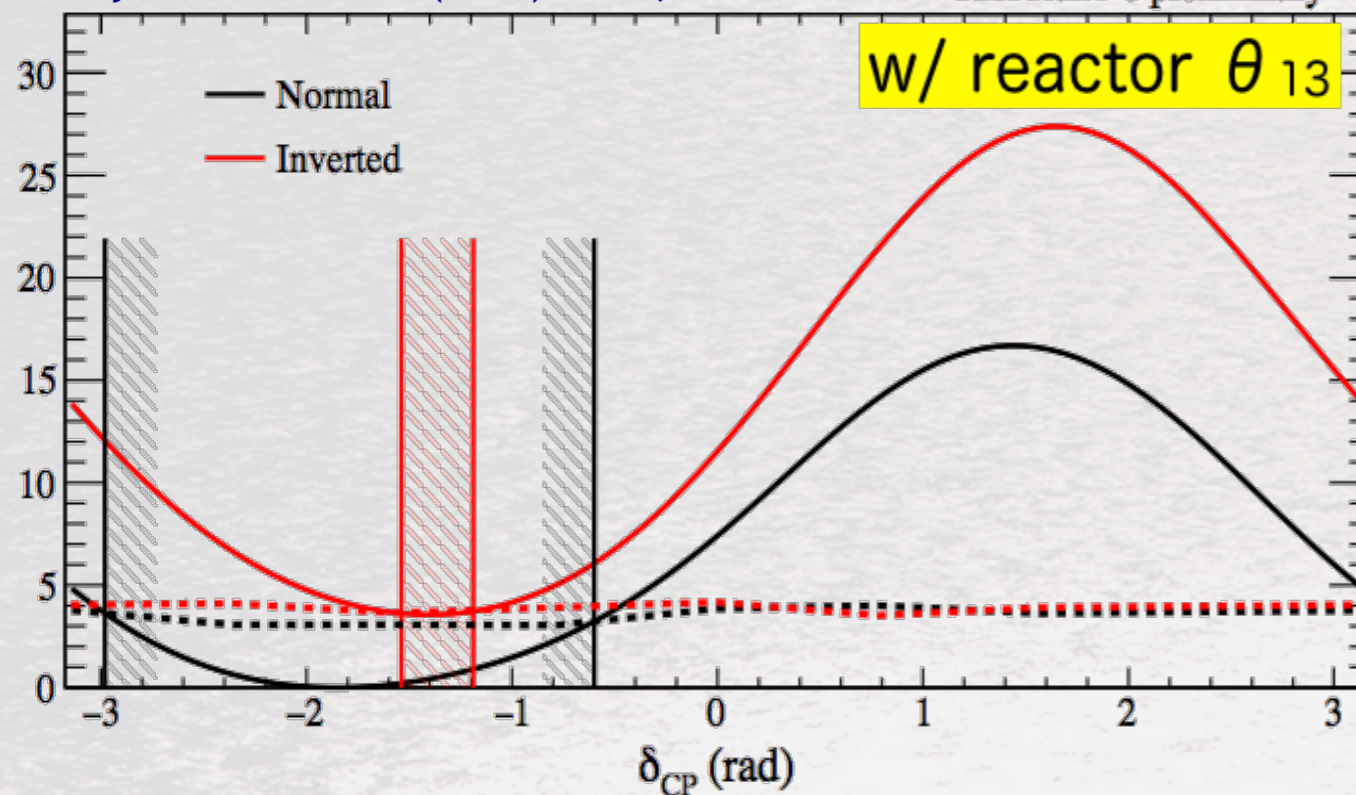
See **Ciro Riccio's** talk today

- **T2K**: working well, **470 kW**
 - $1.5 \cdot 10^{21}$ p.o.t (ν) e $7.6 \cdot 10^{20}$ ($\bar{\nu}$)
 - **2017**: new release δ_{CP} e θ_{12}
 - Upgrade T2K-II in progress (ND280)
 - **CPV 3σ ? Possible.**

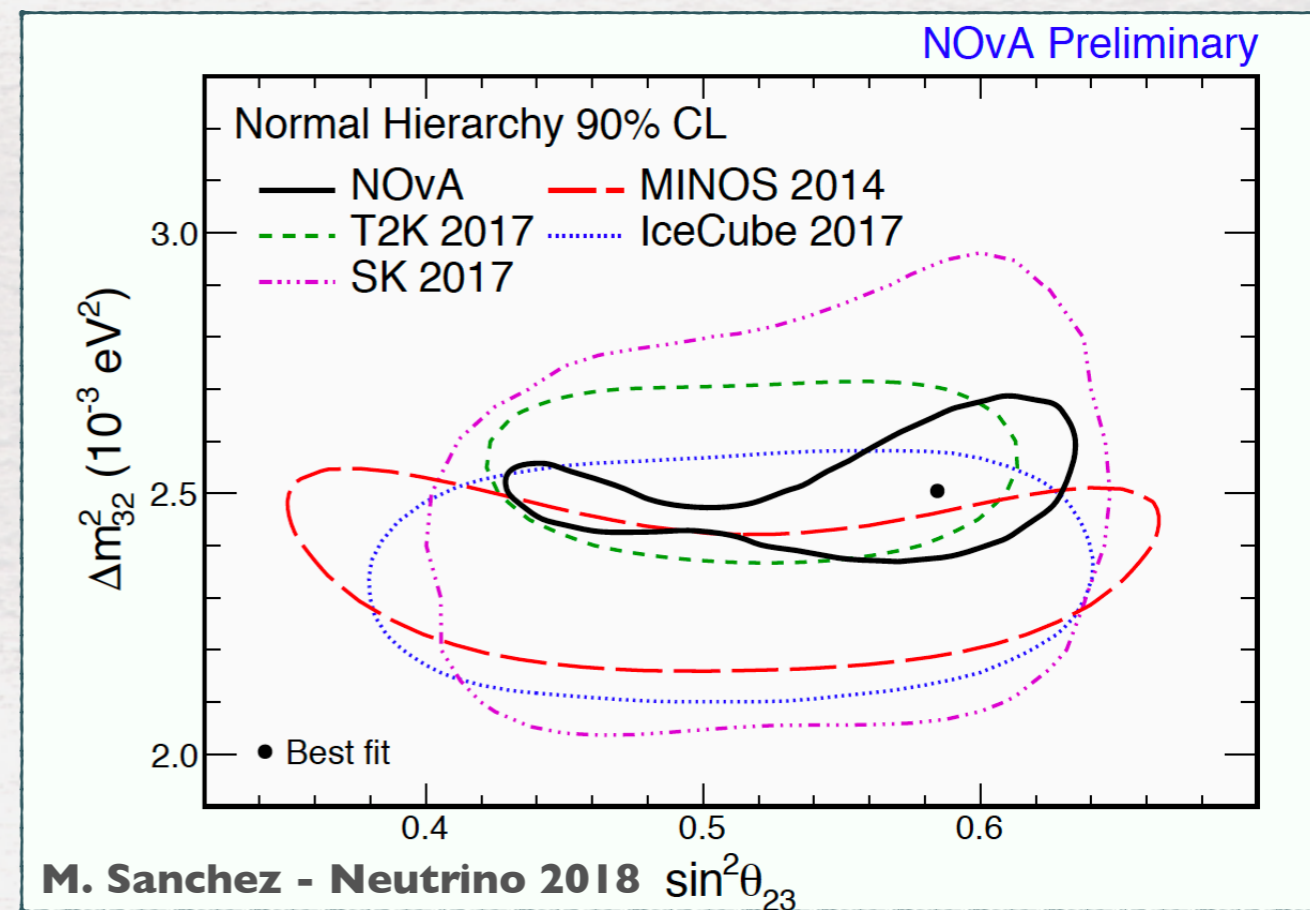
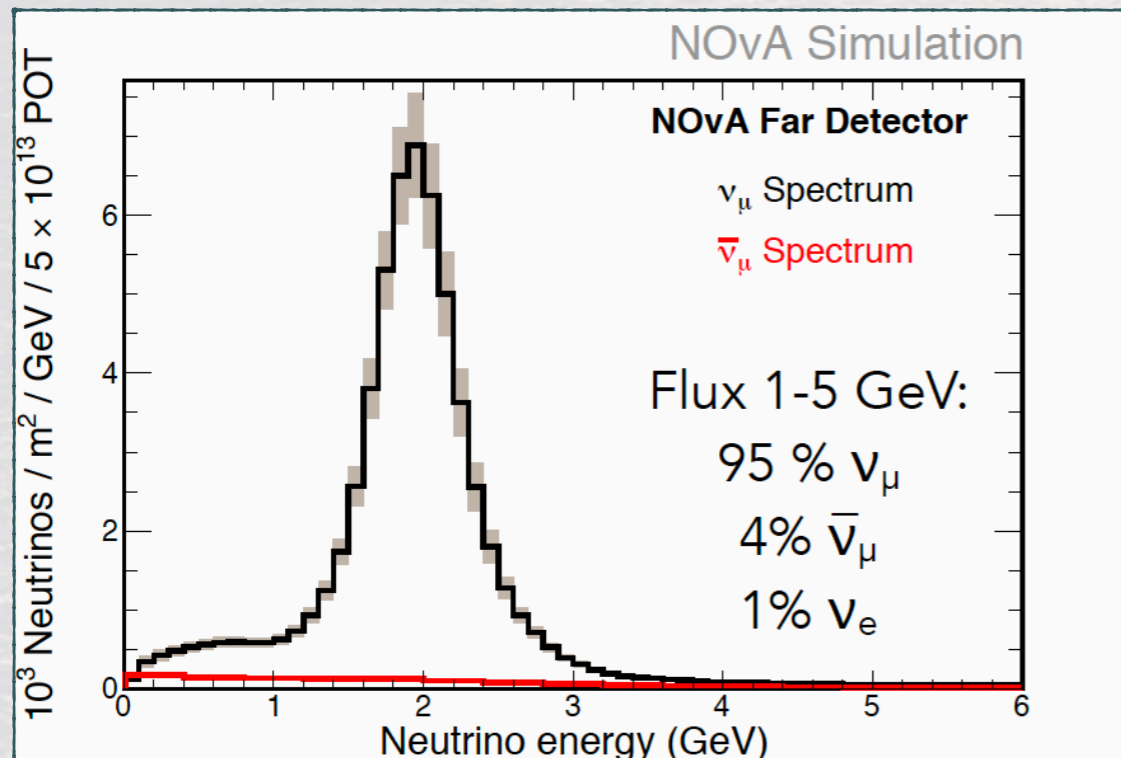
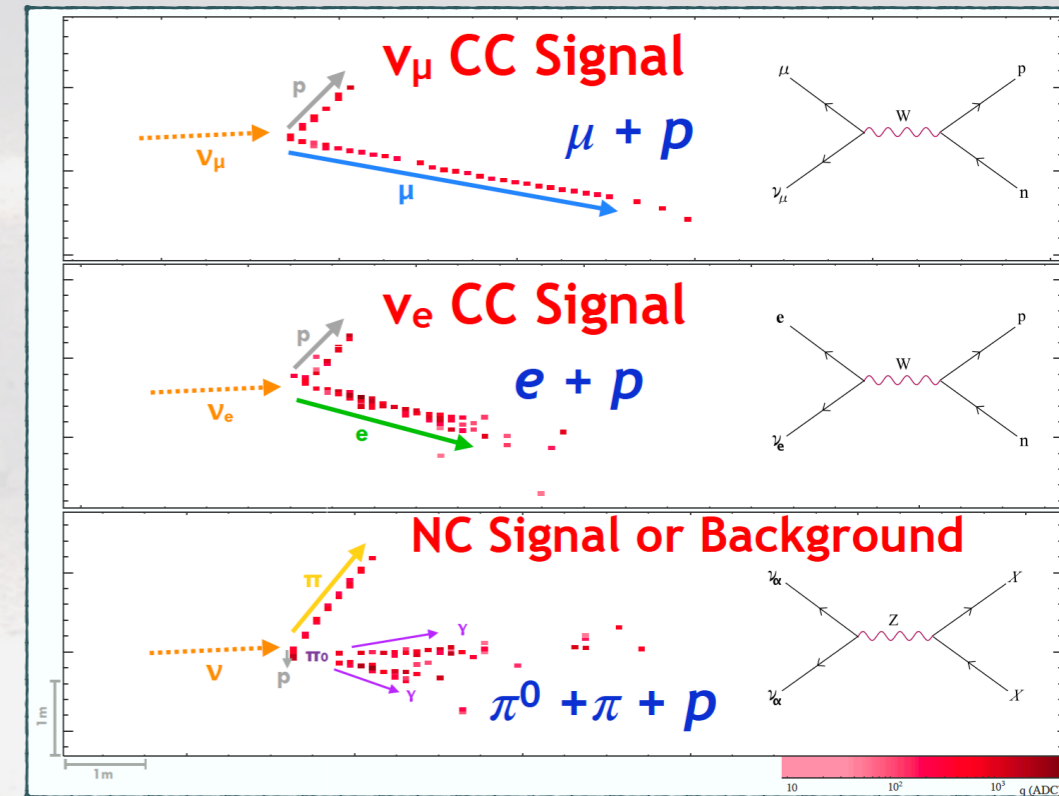
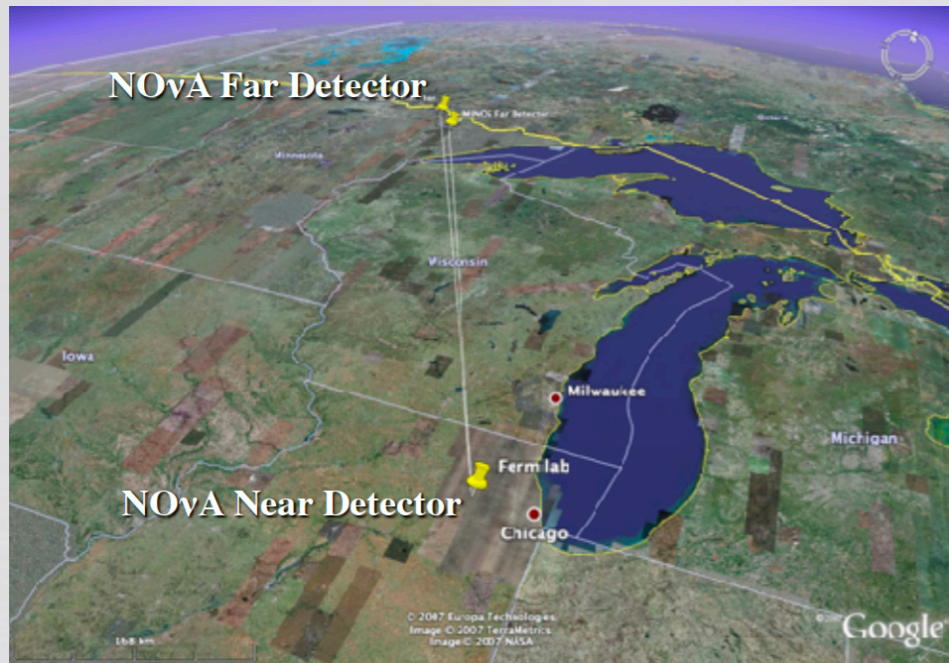


Phys. Rev. Lett. 118 (2017) no.15, 151801

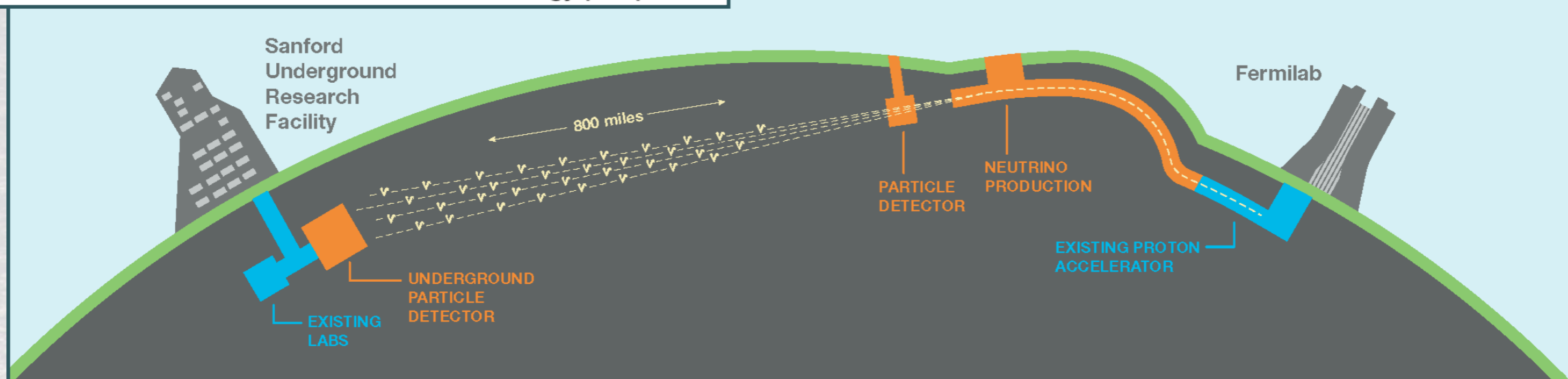
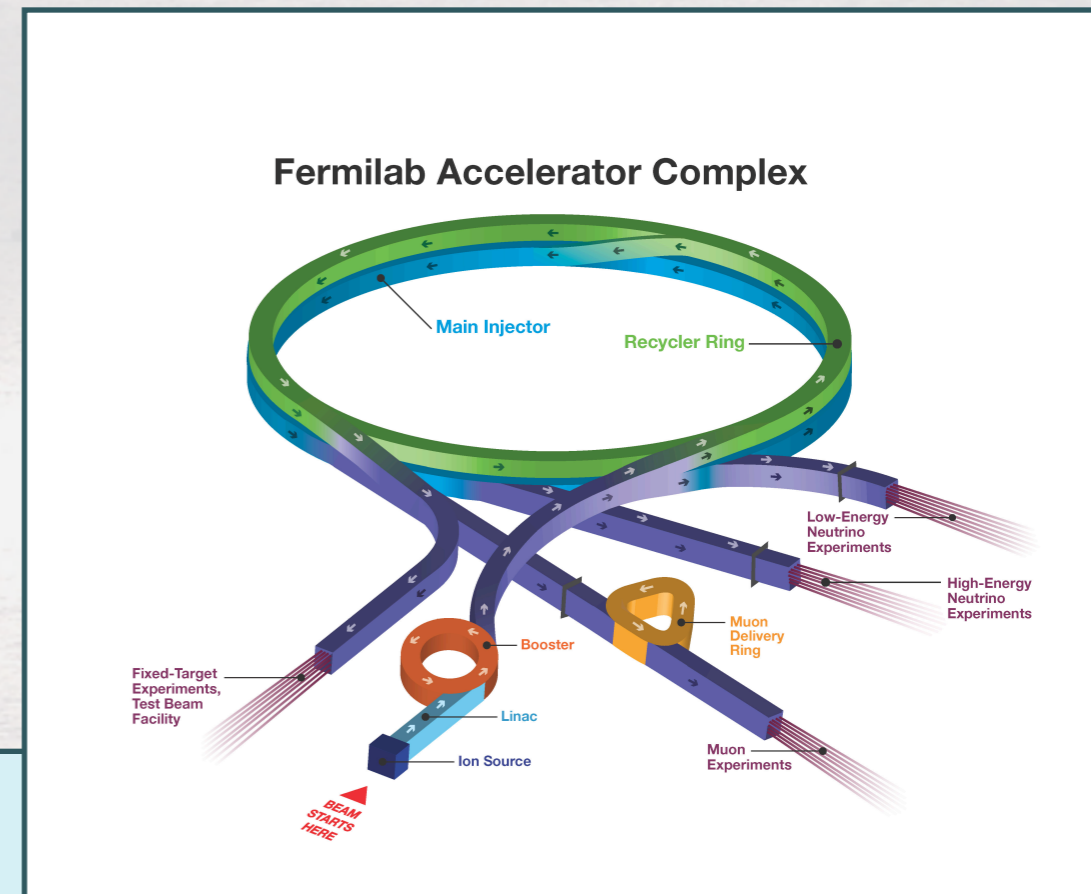
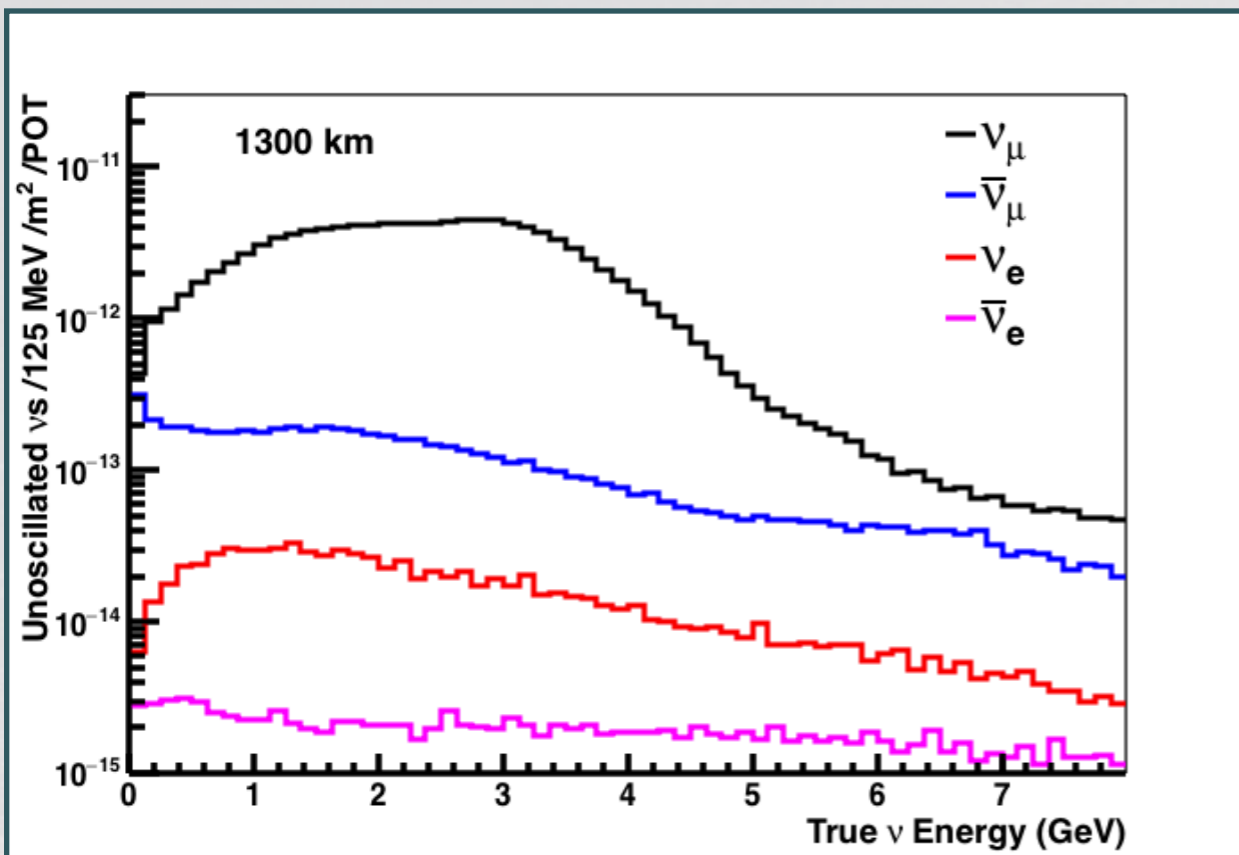
T2K Run1-8 preliminary



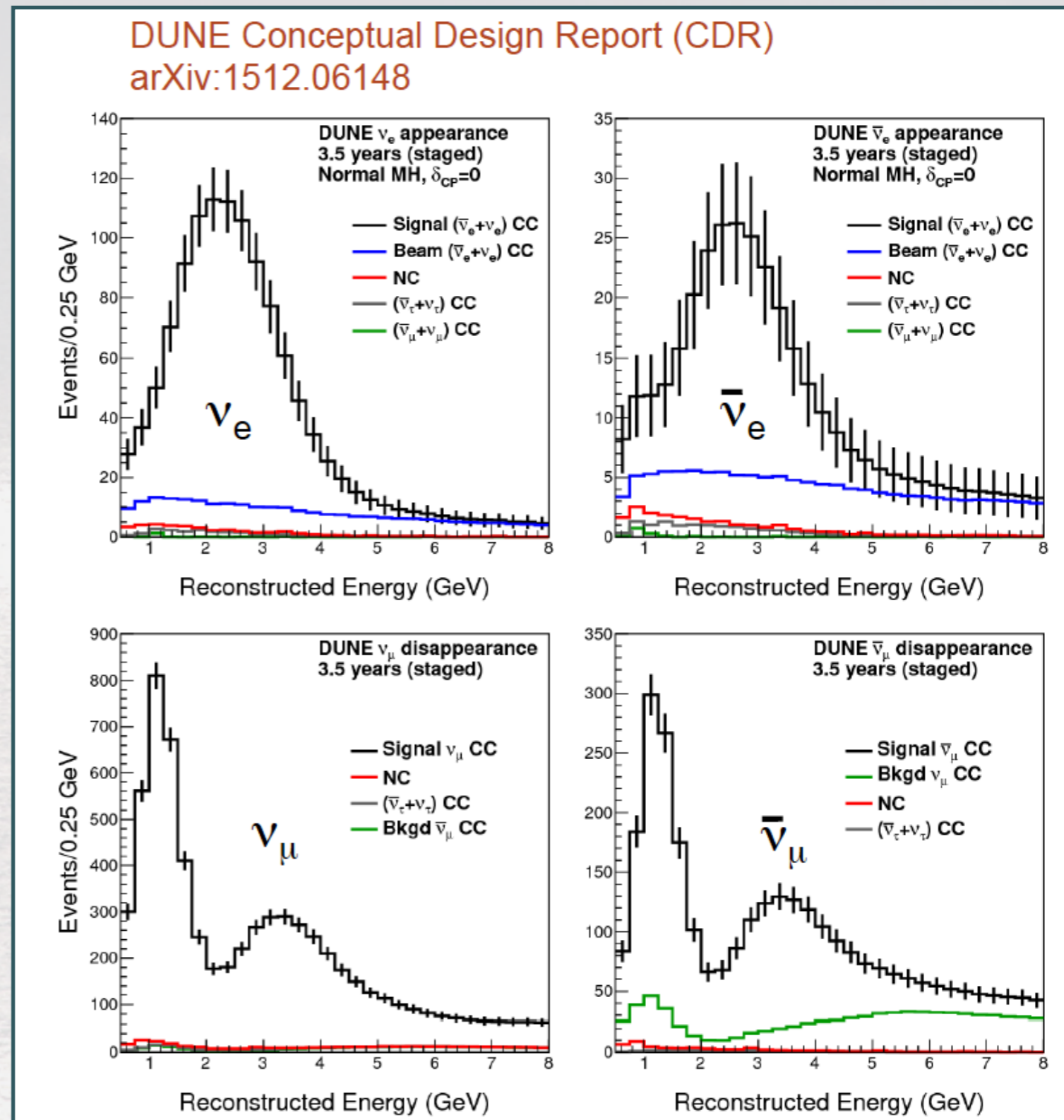
- NuMI: longest baseline, from FNAL to Minnesota
 - 700 kW, 810 km, off axis (narrow) neutrino beam
 - Disappearance, Appearance



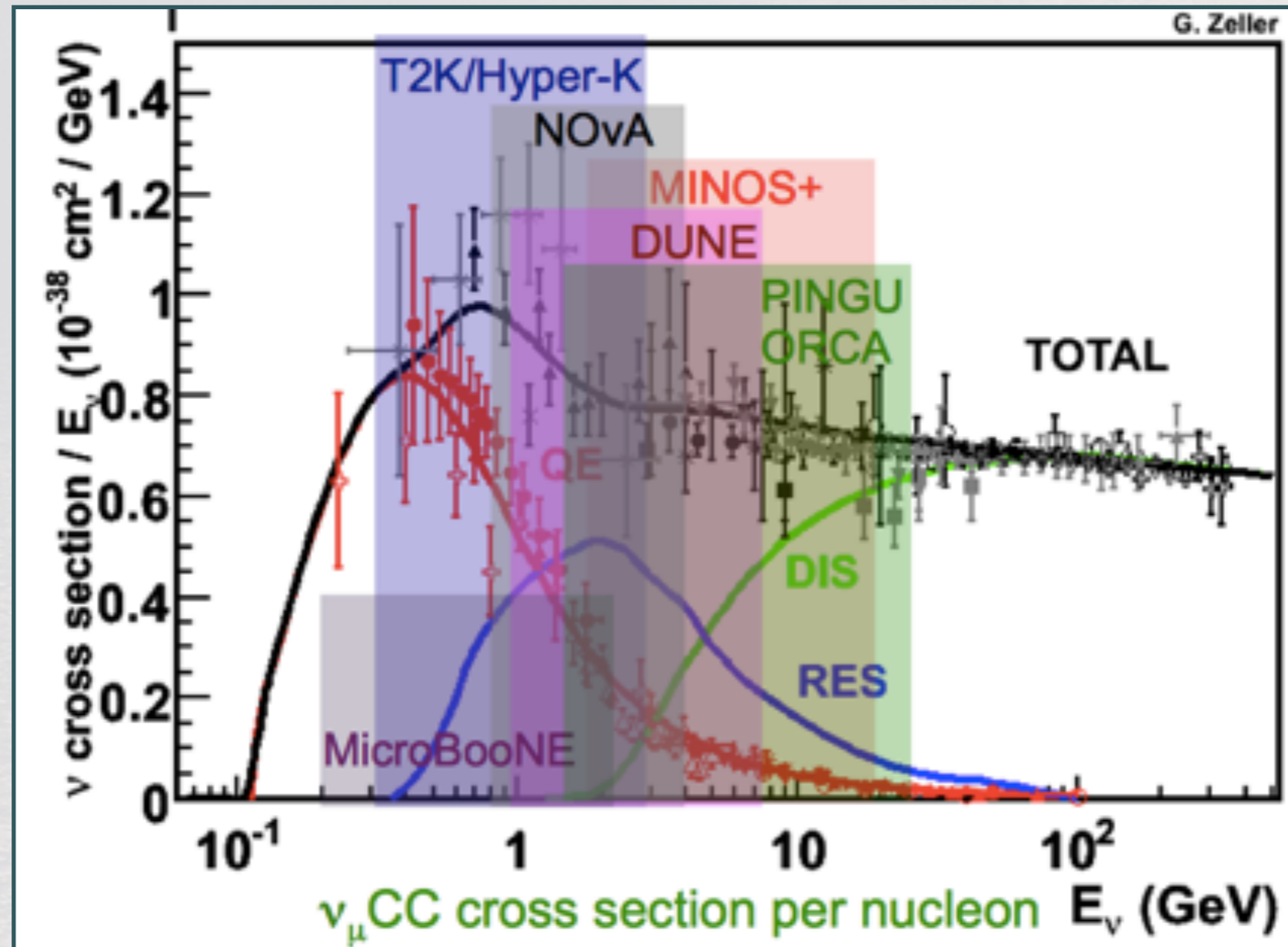
- Long base line experiment from Fermilab to SURF at Sanford (S.D.)
 - 60-120 GeV proton beam, 1.2 (2.4) MW, 1300 km
 - Neutrino and Anti-Neutrino mode



- GENIE event generation, Geant-4 beam, fast MC detector response
 - About **1000** ν_e and $\bar{\nu}_e$ in 7 years
 - Full fit to four spectra to extract parameters
 - GLoBES configurations [arXiv: 1606.09550](https://arxiv.org/abs/1606.09550)
- In addition:
 - SN, proton decay, ν cross sections, sterile ν , new physics



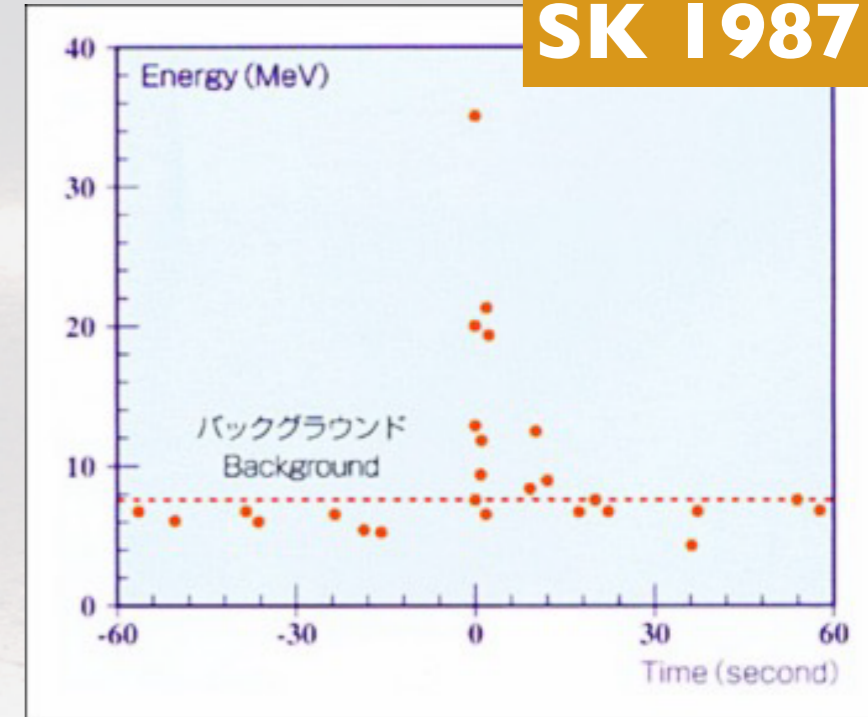
- The current understanding of neutrino cross sections with nuclei is not sufficient to extract all potential science from next generation neutrino experiments



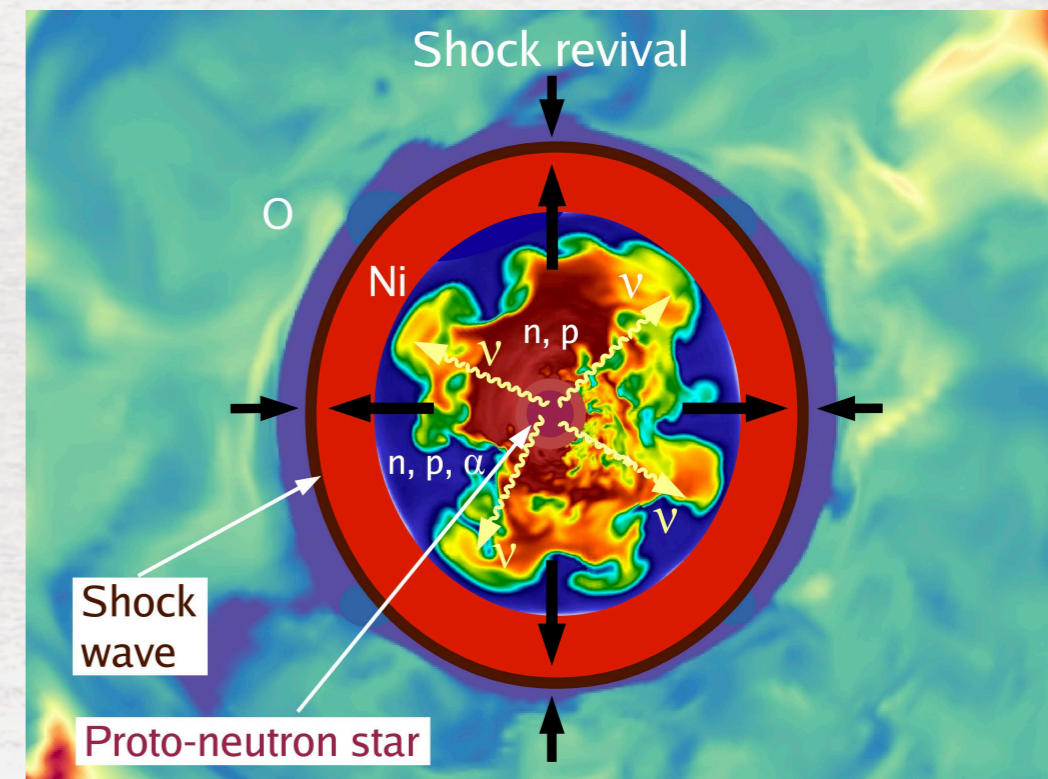
- Theoretical progress is crucial. Another example of the importance of strong interactions corrections to “weak physics” (ϵ'/ϵ , g_μ , $0\nu\beta\beta$ matrix elements,

SK 1987

- One of the most **challenging theoretical problems**, all **at the extreme**
 - Strong coupling effects of EW interactions
 - Highly non-linear and poorly understood **flavour evolution** problem
 - A unique lab for
 - **Particle physics, Nuclear Physics, General Relativity**
 - Strong feedback from neutrino-matter dynamics in truly unique conditions

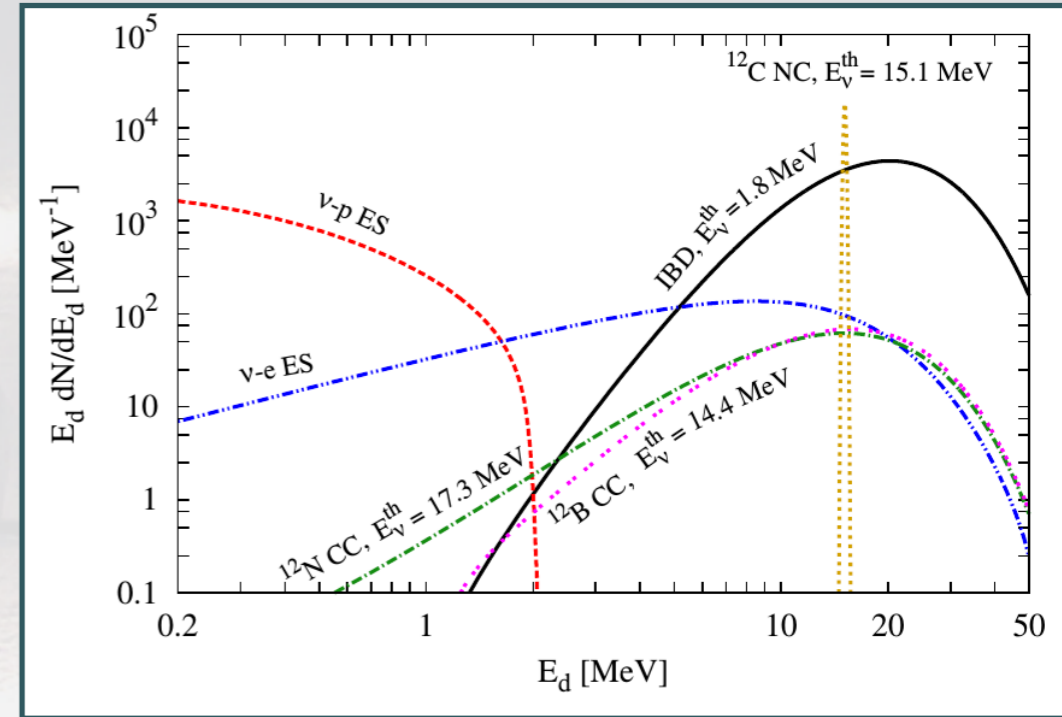


+ ν ?



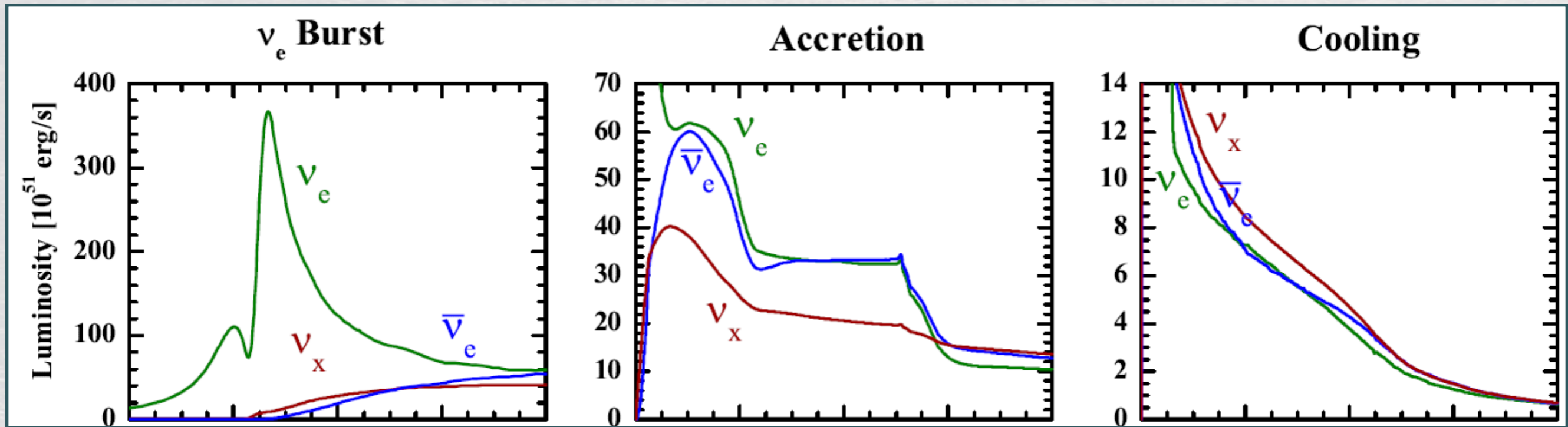
Janka 2018

- Potential breakthrough because of the very large number of events and resolution



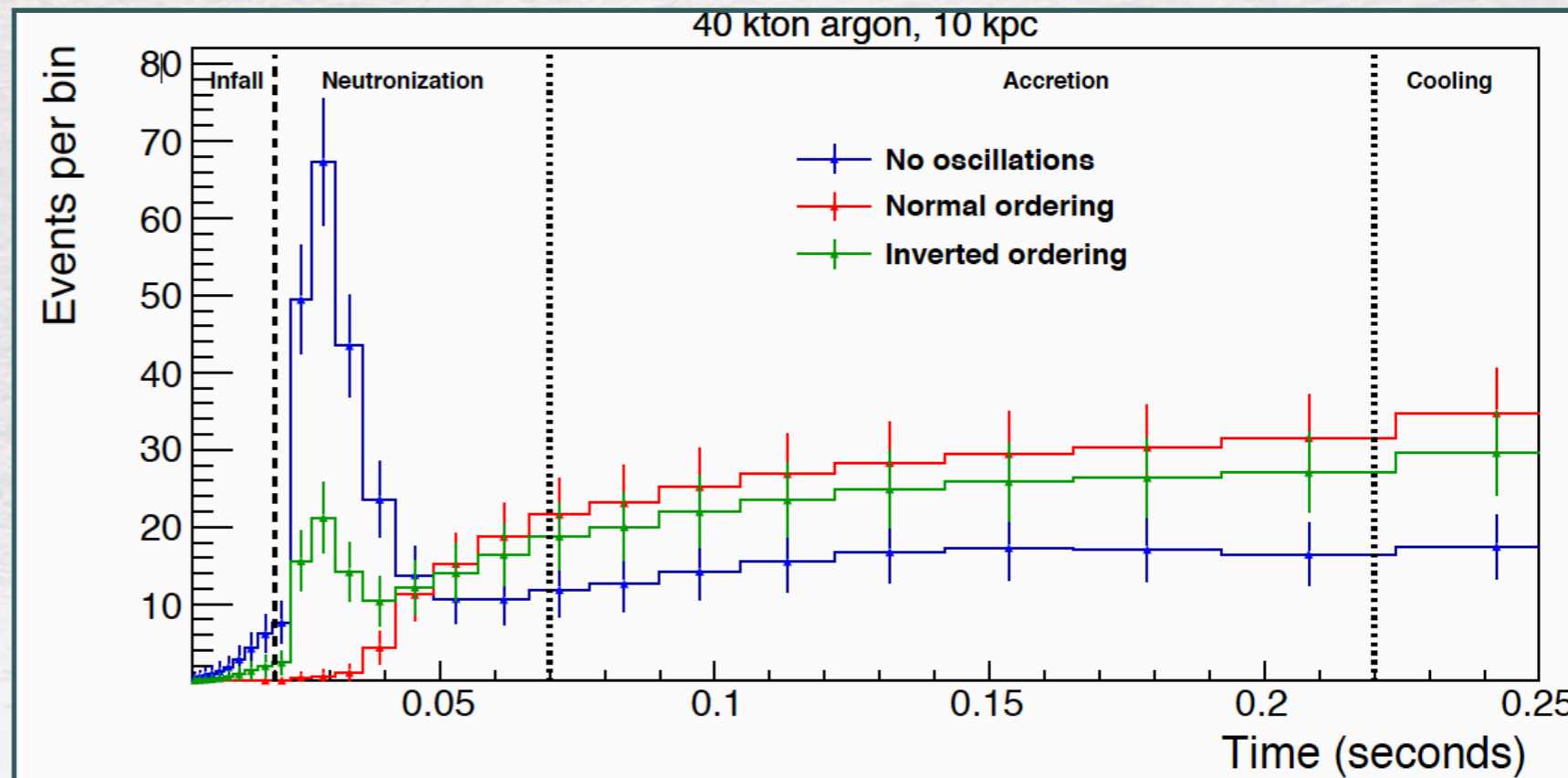
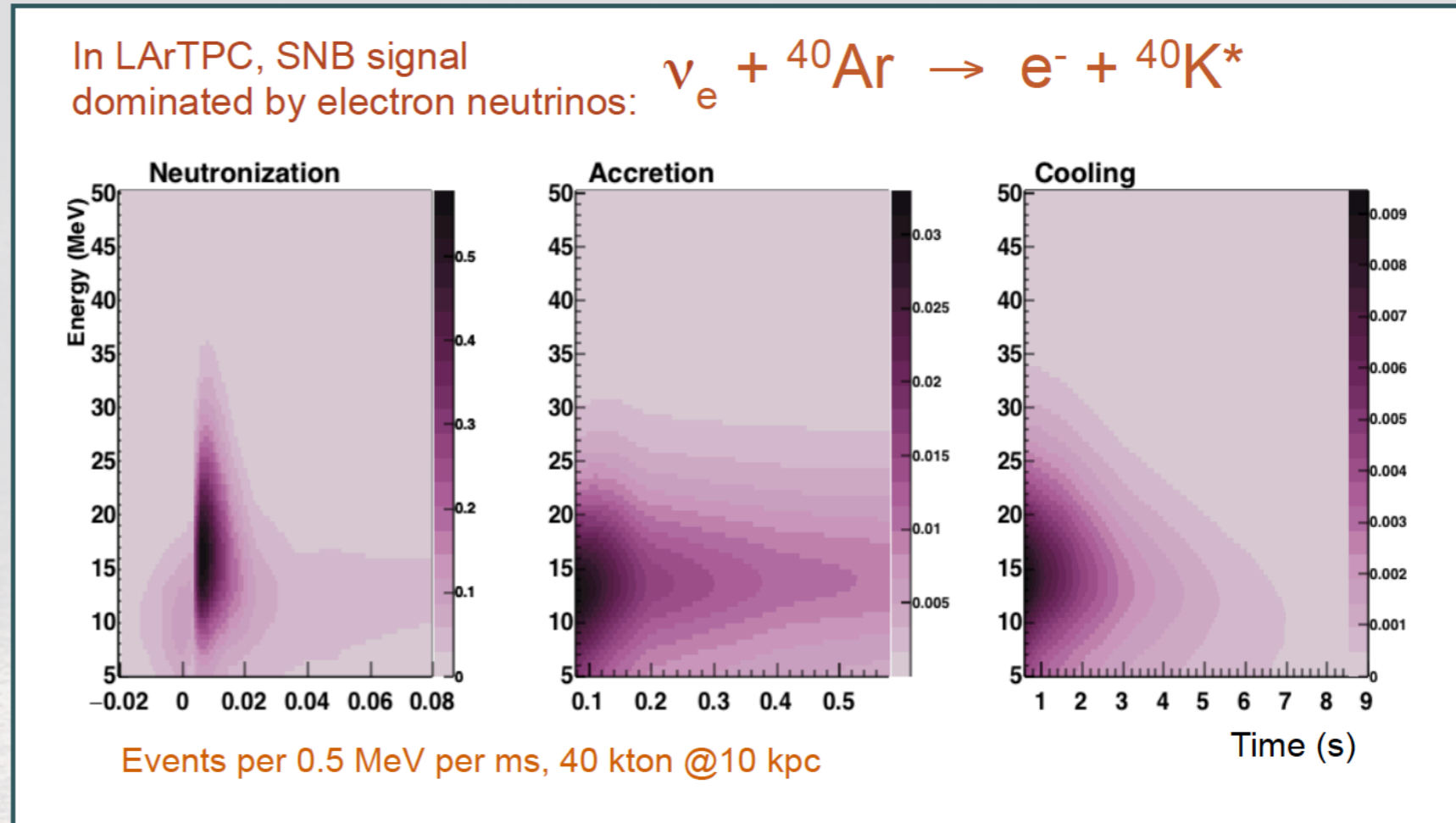
JUNO collab.,
arXiv:1507.05613

Channel	Type	Events for different $\langle E_\nu \rangle$ values		
		12 MeV	14 MeV	16 MeV
$\bar{\nu}_e + p \rightarrow e^+ + n$	CC	4.3×10^3	5.0×10^3	5.7×10^3
$\nu + p \rightarrow \nu + p$	NC	0.6×10^3	1.2×10^3	2.0×10^3
$\nu + e \rightarrow \nu + e$	ES	3.6×10^2	3.6×10^2	3.6×10^2
$\nu + {}^{12}\text{C} \rightarrow \nu + {}^{12}\text{C}^*$	NC	1.7×10^2	3.2×10^2	5.2×10^2
$\nu_e + {}^{12}\text{C} \rightarrow e^- + {}^{12}\text{N}$	CC	0.5×10^2	0.9×10^2	1.6×10^2
$\bar{\nu}_e + {}^{12}\text{C} \rightarrow e^+ + {}^{12}\text{B}$	CC	0.6×10^2	1.1×10^2	1.6×10^2

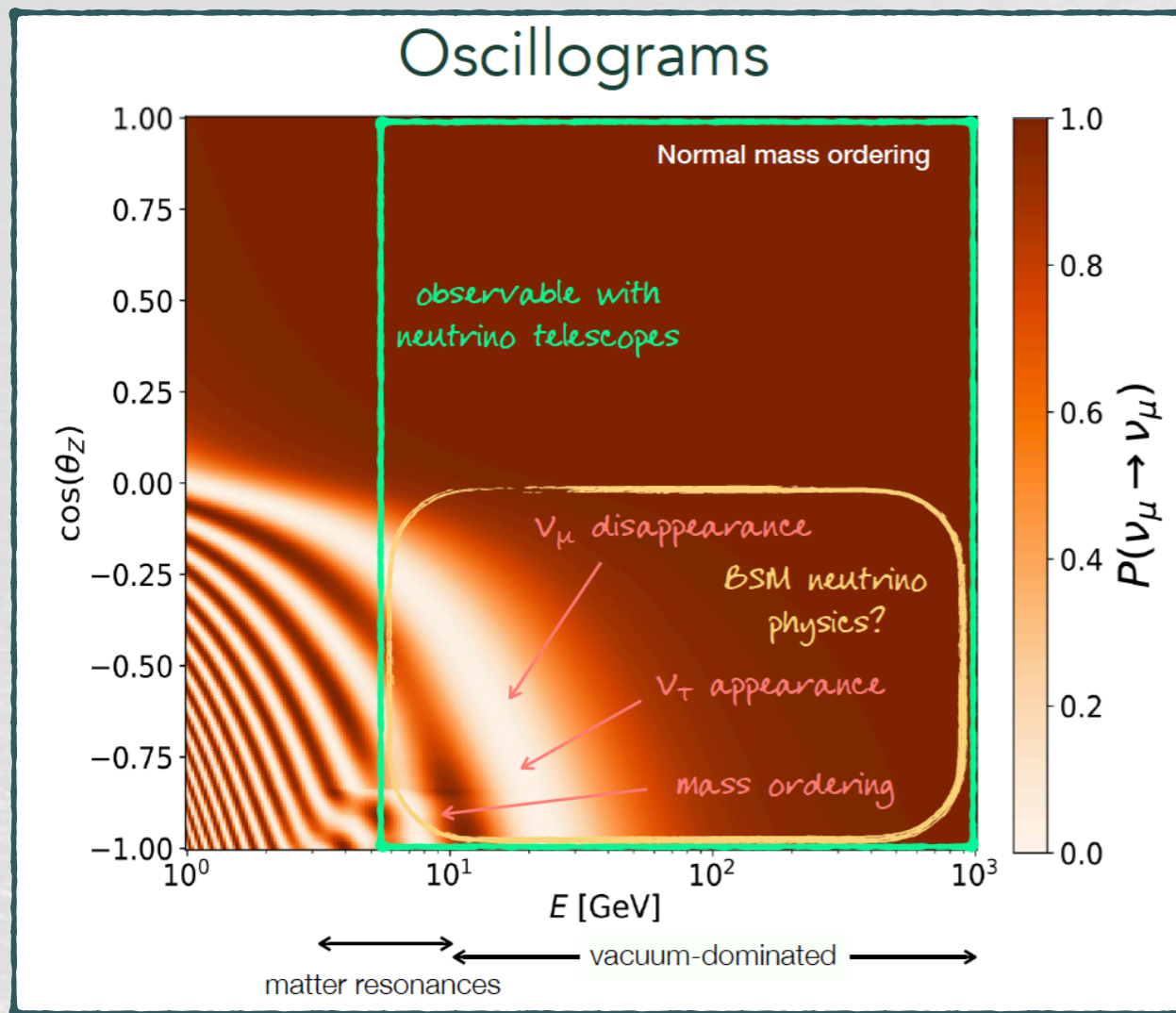


A. Mirizzi et. al., arXiv:1508.00785

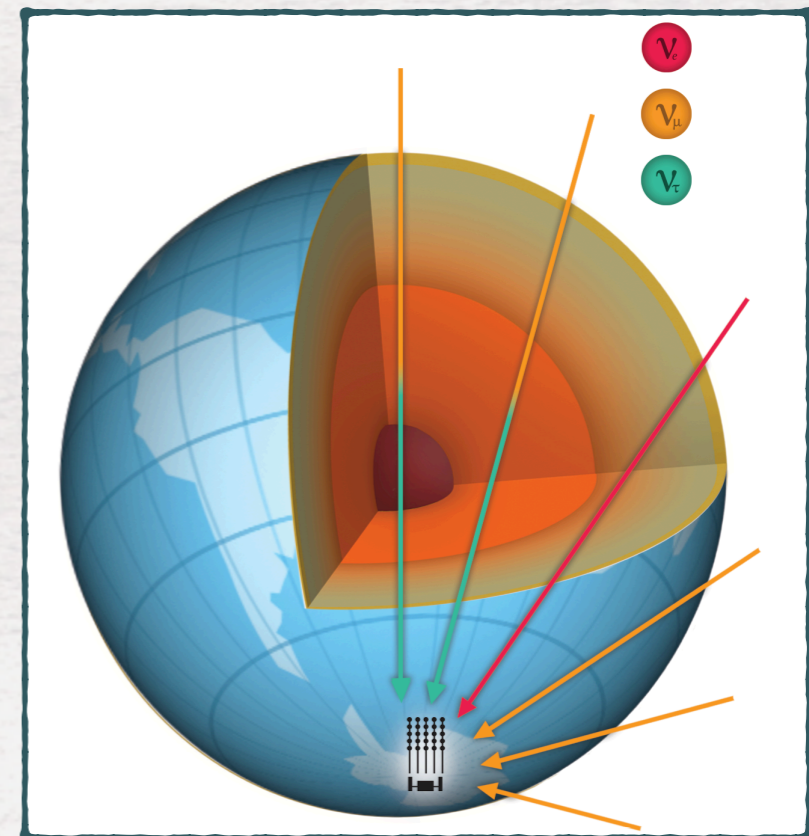
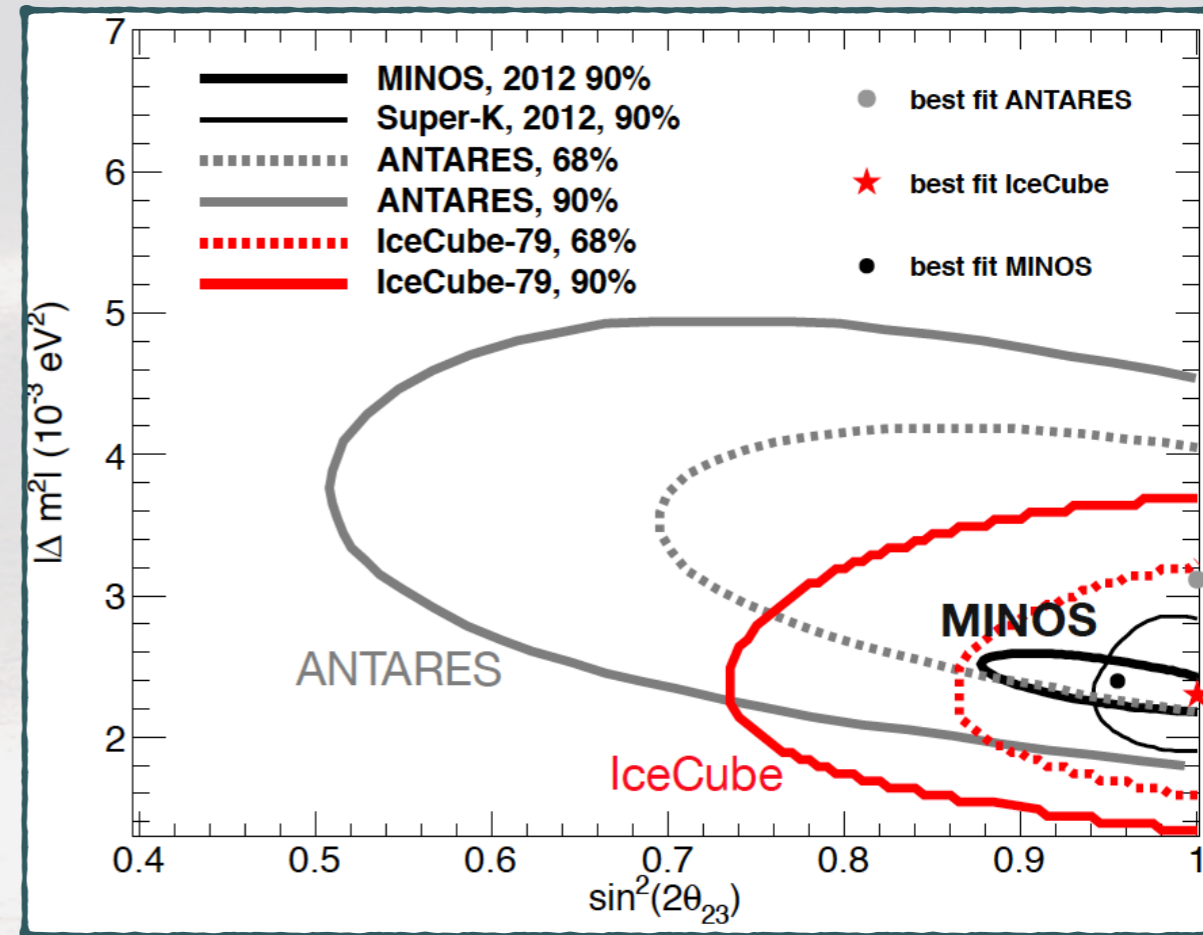
- No single event analysis, total charge (or light) signal as a function of time
- Sensitivity to mass hierarchy and SN model with early time detection



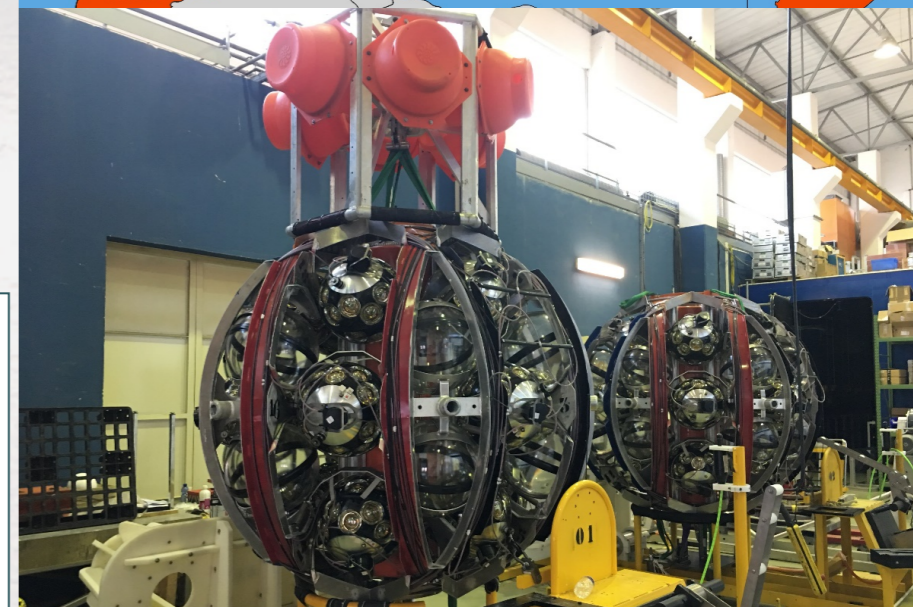
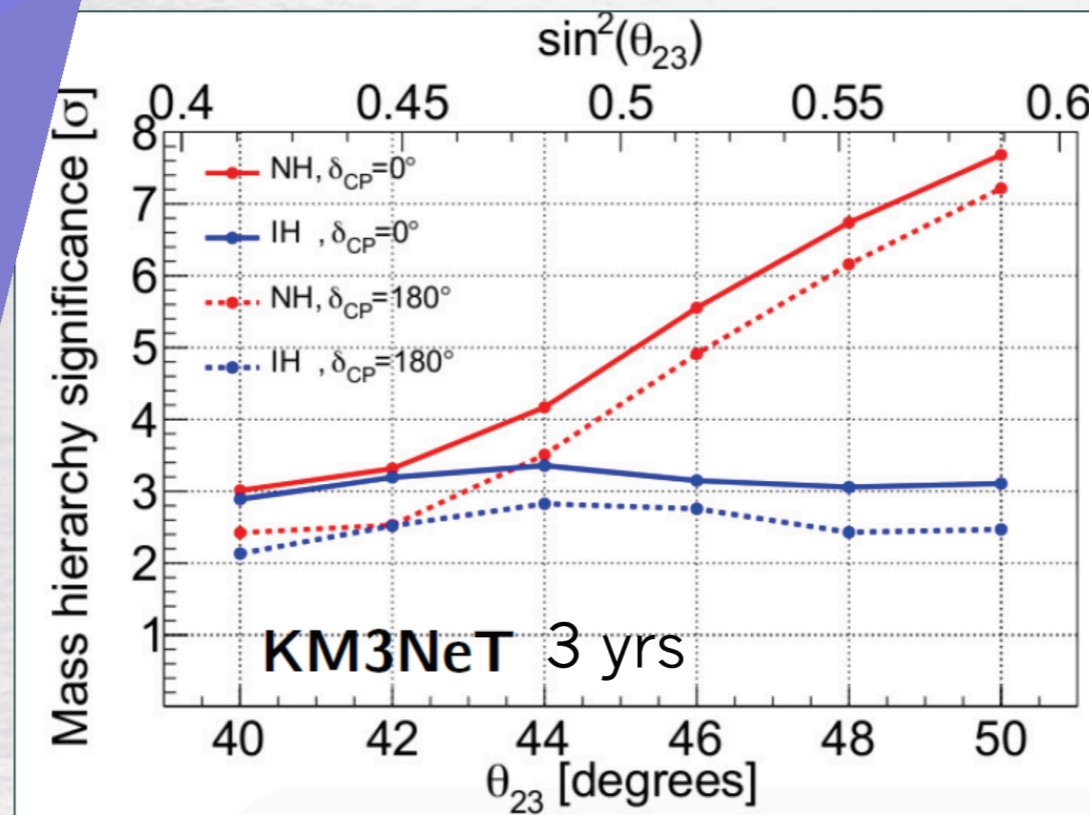
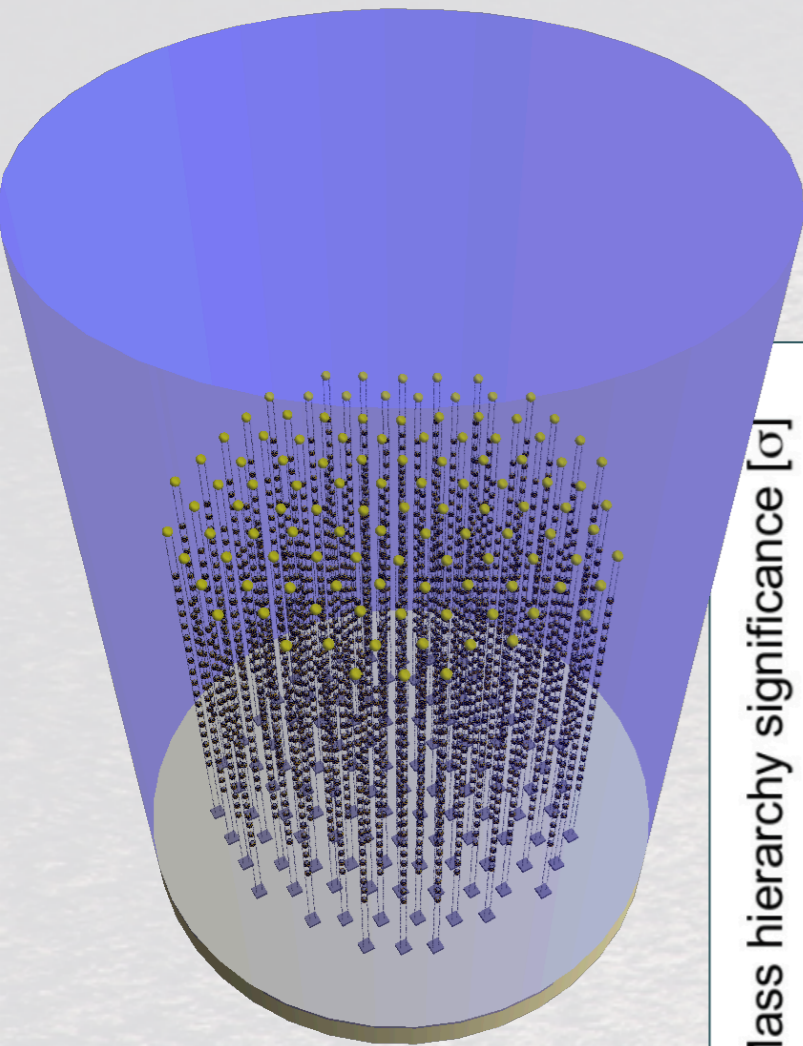
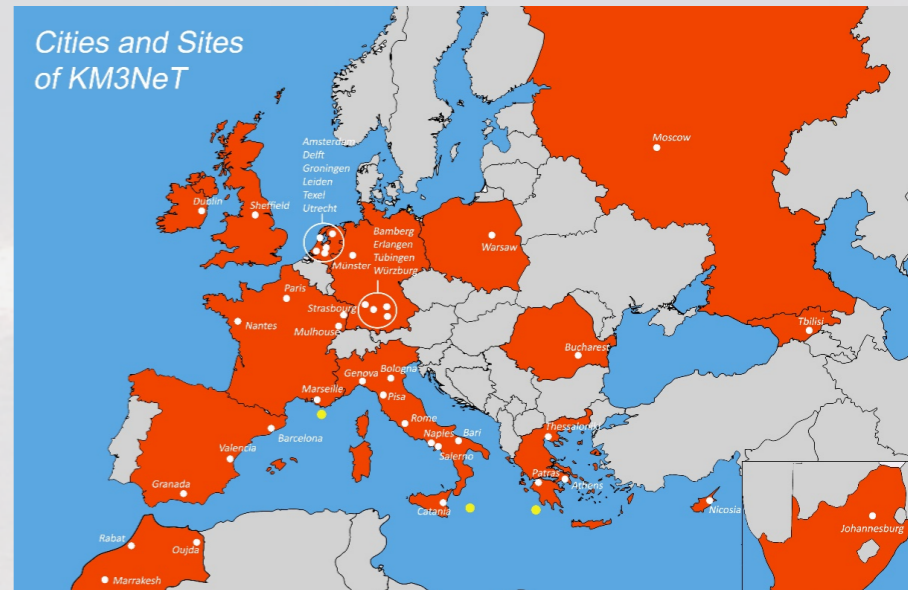
- Smaller precision, but completely different energy range (up to 100 TeV)
 - IceCube and Antares data so far
 - Sensitivity to new physics and mass hierarchy



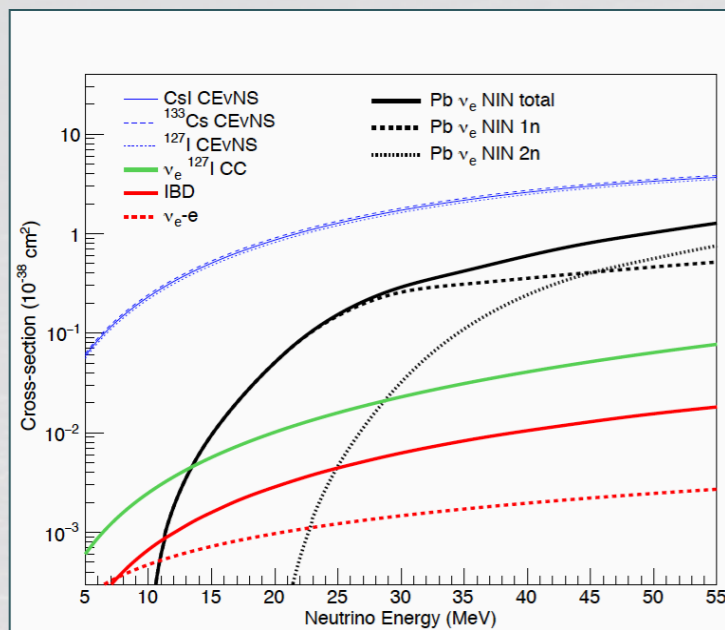
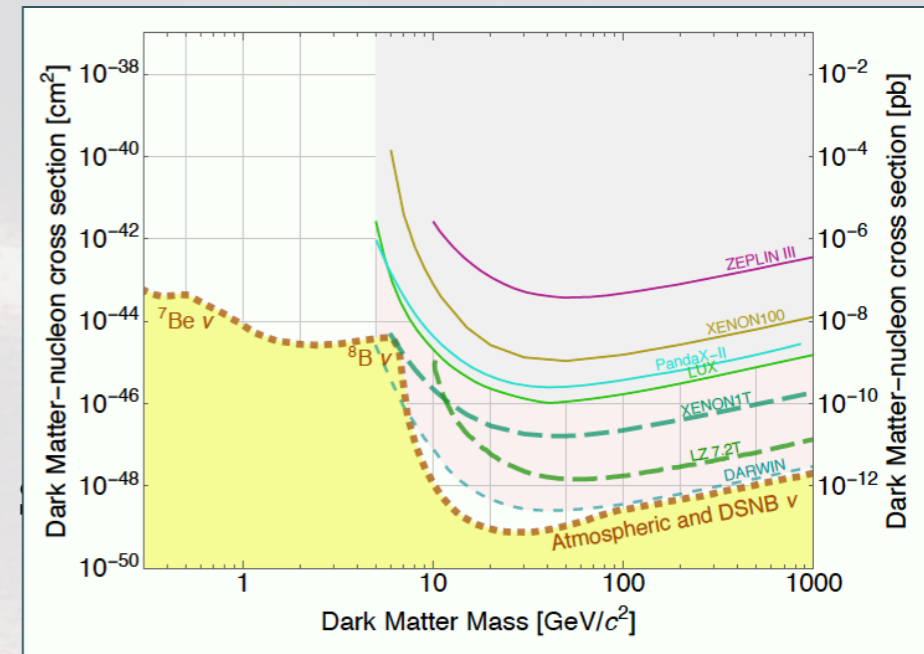
From: T. DeYoung Neutrino 2018



- **KM3Net**: key European project for **multi-messenger ν**
 - ARCA: Neutrino Astrophysics
 - Sources, Acceleration mechanisms
 - ORCA (Mass hierarchy)



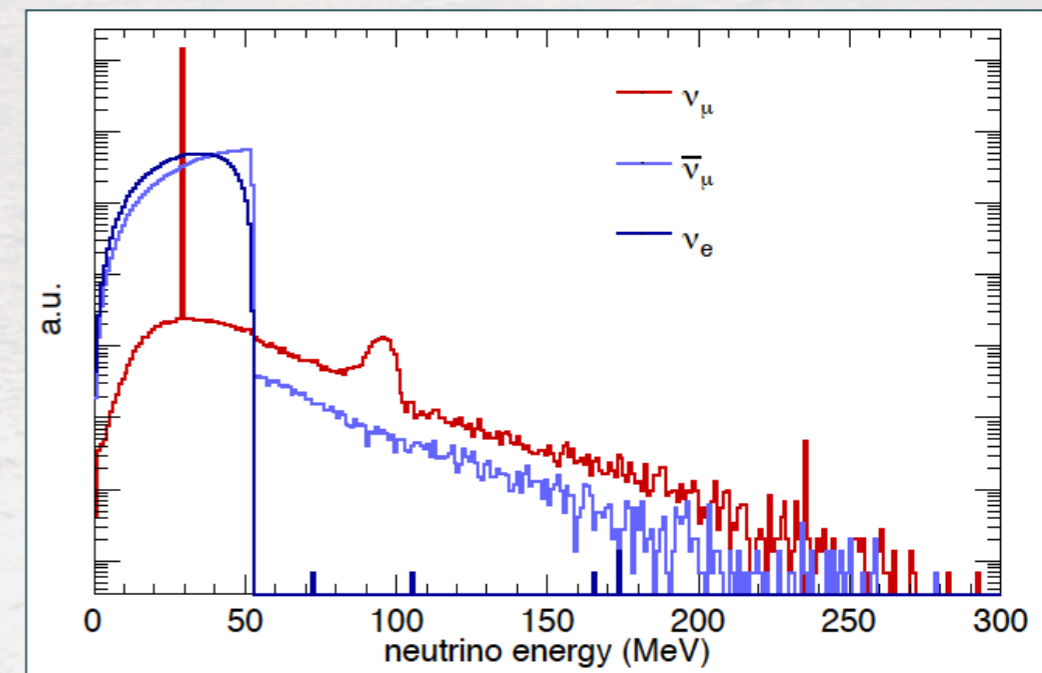
- Coherent neutrino scattering finally observed
 - Predicted in 1974, very difficult measurement
 - Cross section is large, energy recoil very small
 - Fundamental physics, SuperNovae, Dark Matter



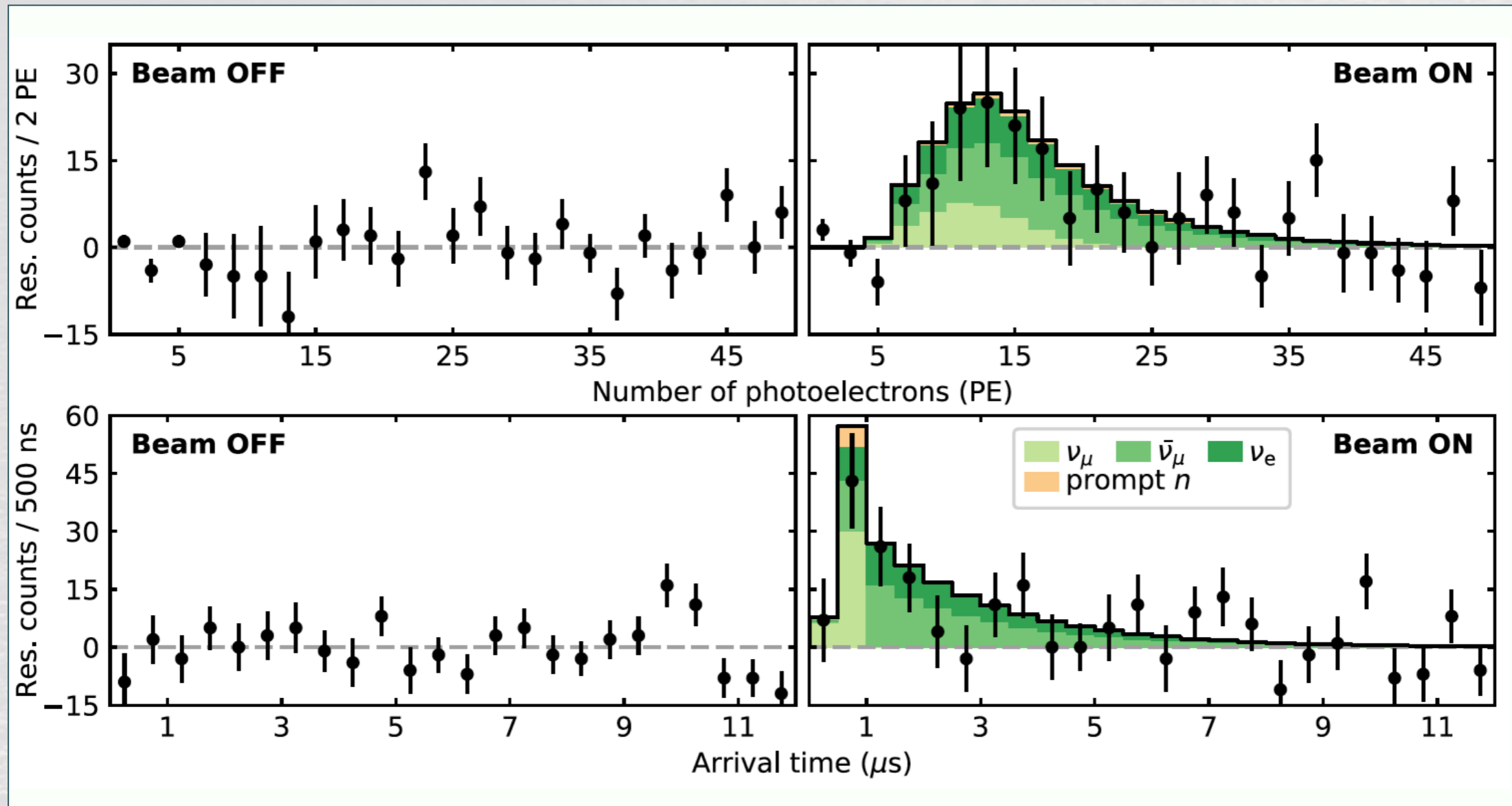
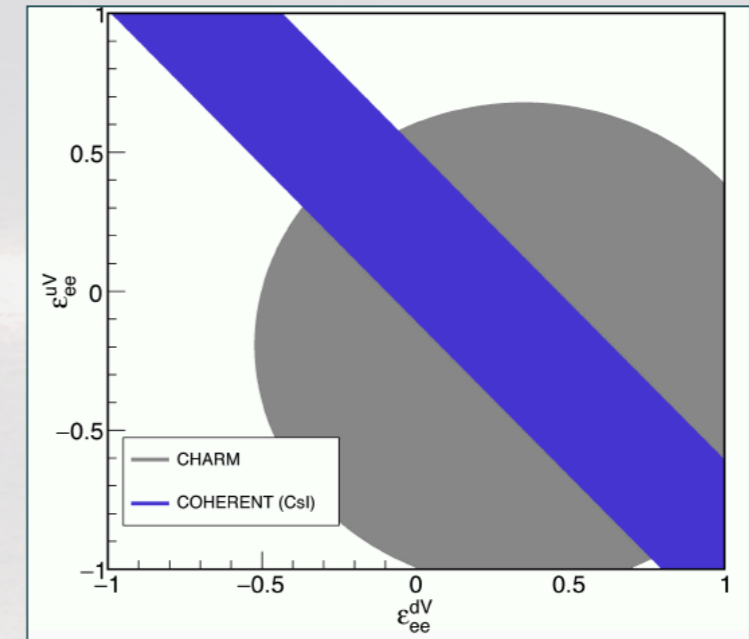
$$\sigma_{tot} = \frac{G_F^2 E_\nu^2}{4\pi} [Z(1 - 4\sin^2\theta_W) - N]^2 F^2(Q^2) \quad E_\nu \leq 50\text{MeV}$$

$$E_{nr} = \frac{2E^2}{M} = 15 - 160\text{keV} \quad @E_\nu = 30\text{MeV}$$

- Experiment at the SNS, Spallation Neutron source at Oak Ridge, Tennessee
 - Most intense pulsed neutron source
 - Very low threshold CsI detectors



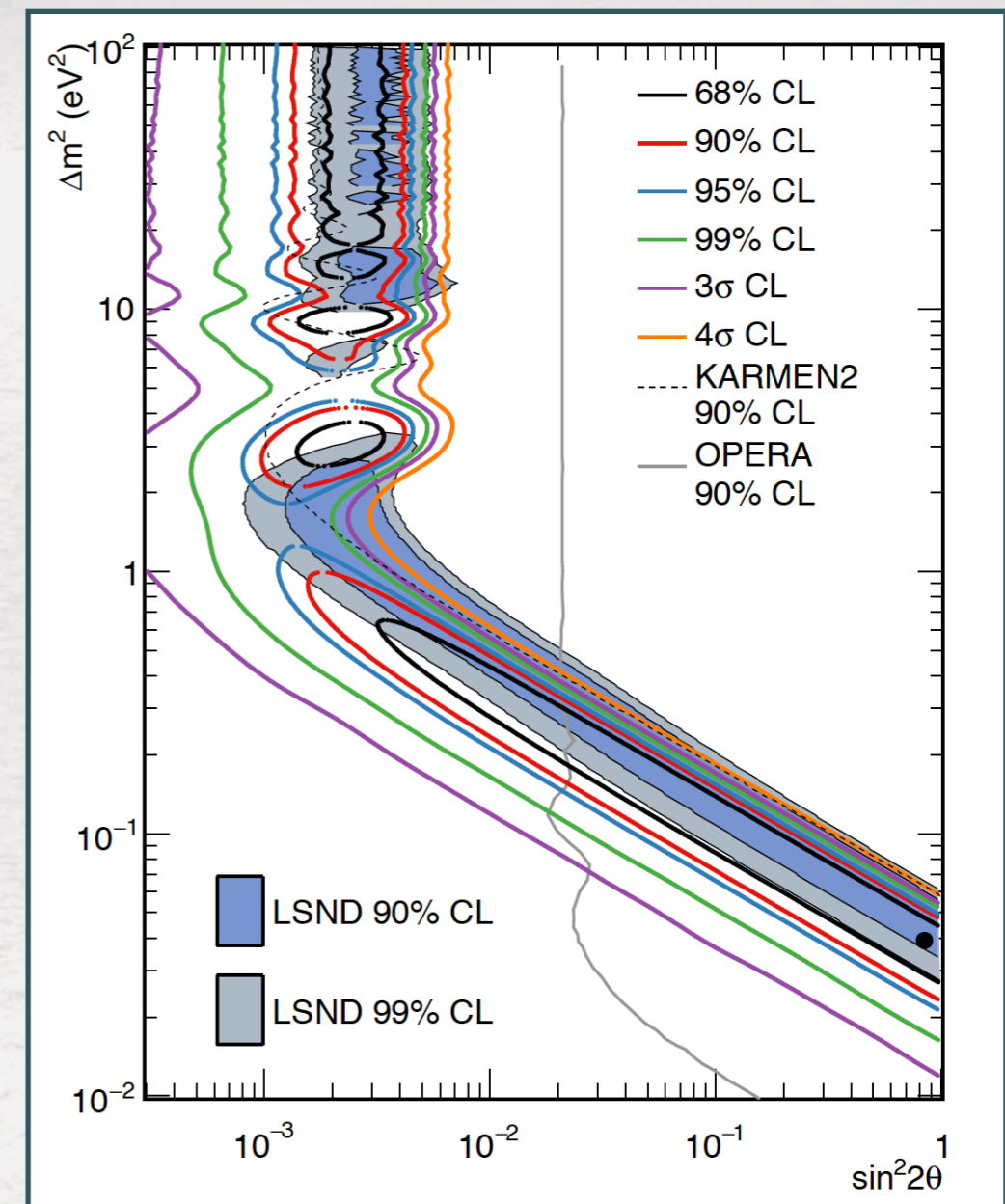
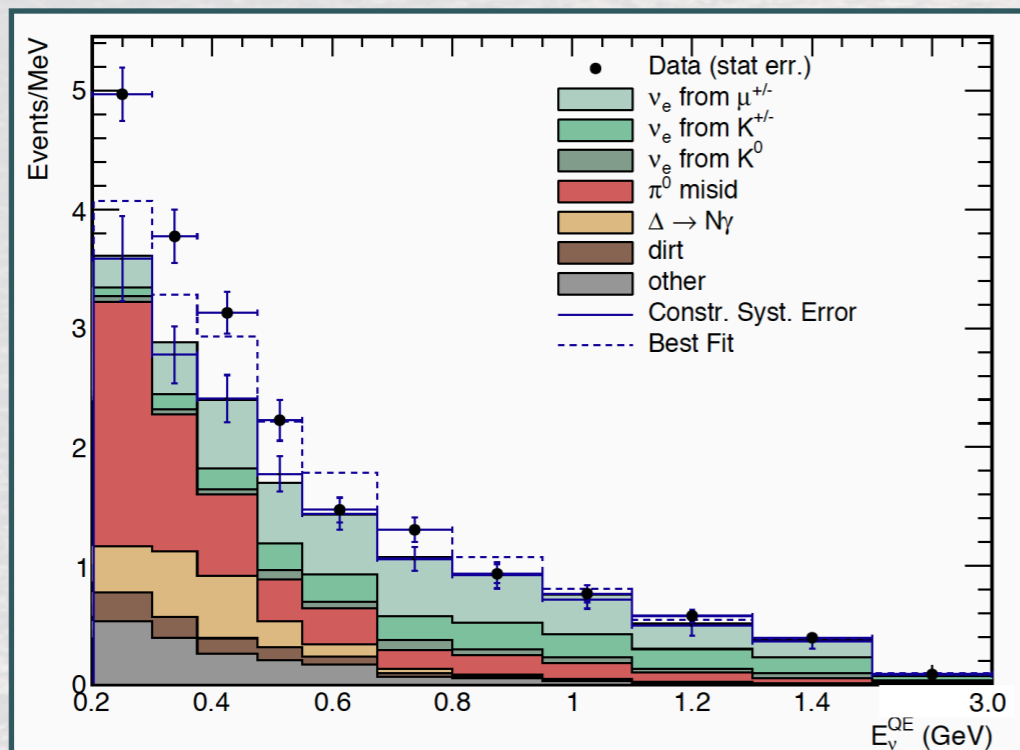
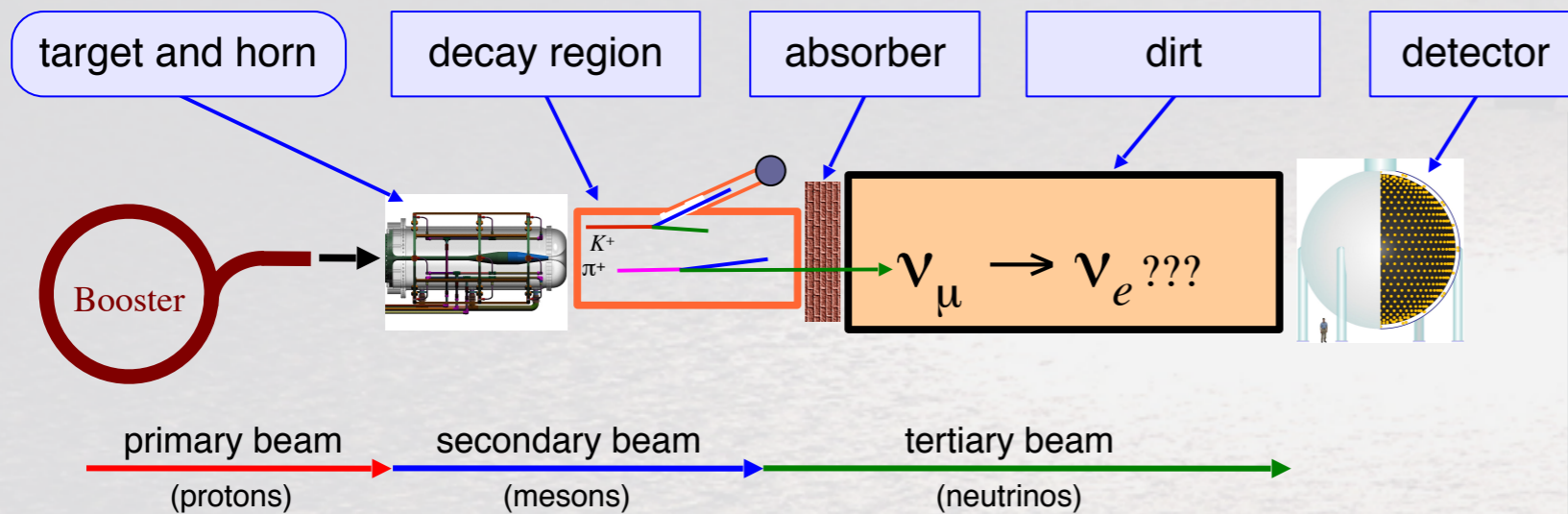
- **$1.4 \cdot 10^{23}$ p.o.t (6 GWhr)**
- **Observed: 136 ± 31 ev, expected 134 ± 22**
- No signal excluded at **6.7σ**
- Already *constraining NSI*



arXiv:1805.12028v1 May 30th, 2018

- Breaking news (well, maybe....)
 - MiniBoone excess of low energy electron appearance confirmed!
 - Very very confusing.....

500 MeV, 500 m

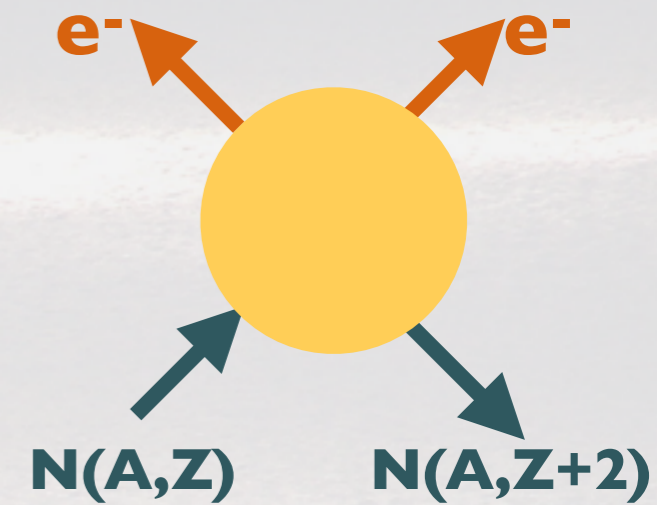


• Neutrino-less double beta decay is a *portal to new physics*

• Its observation requires **probes scales beyond SM**

- **Lepton number** violation
- Majorana vs Dirac mass terms
- New **mass generation** mechanisms
- **Baryogenesis and/or Leptogenesis** mechanisms

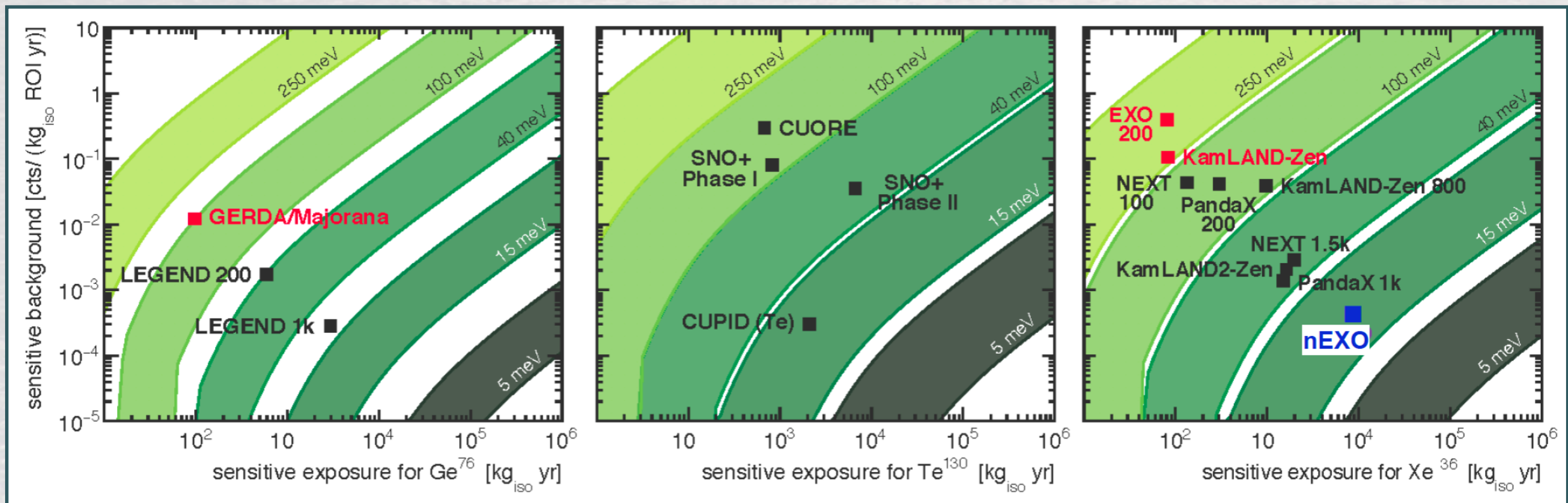
J. Schechter and J.W.F. Valle,
PRD25 (1982) 2951



• NOT necessarily linked to Majorana ν , although it implies non-zero Majorana mass

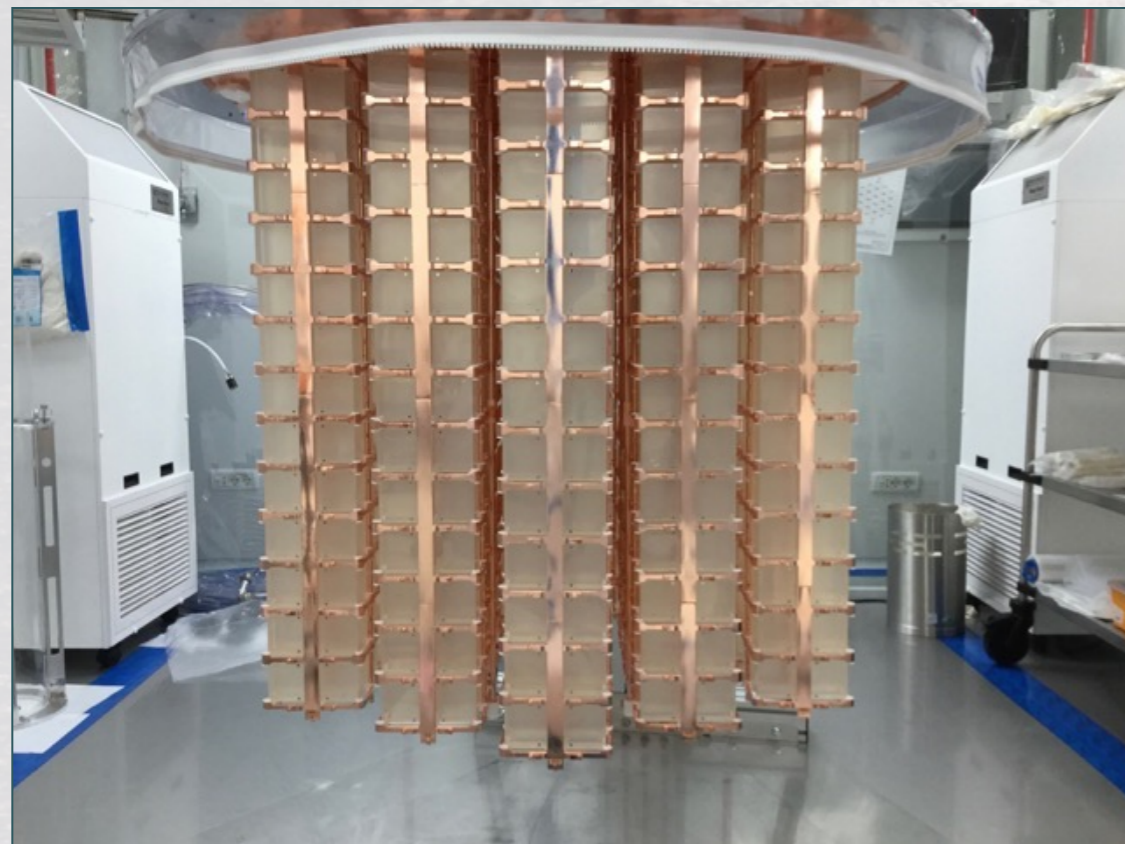
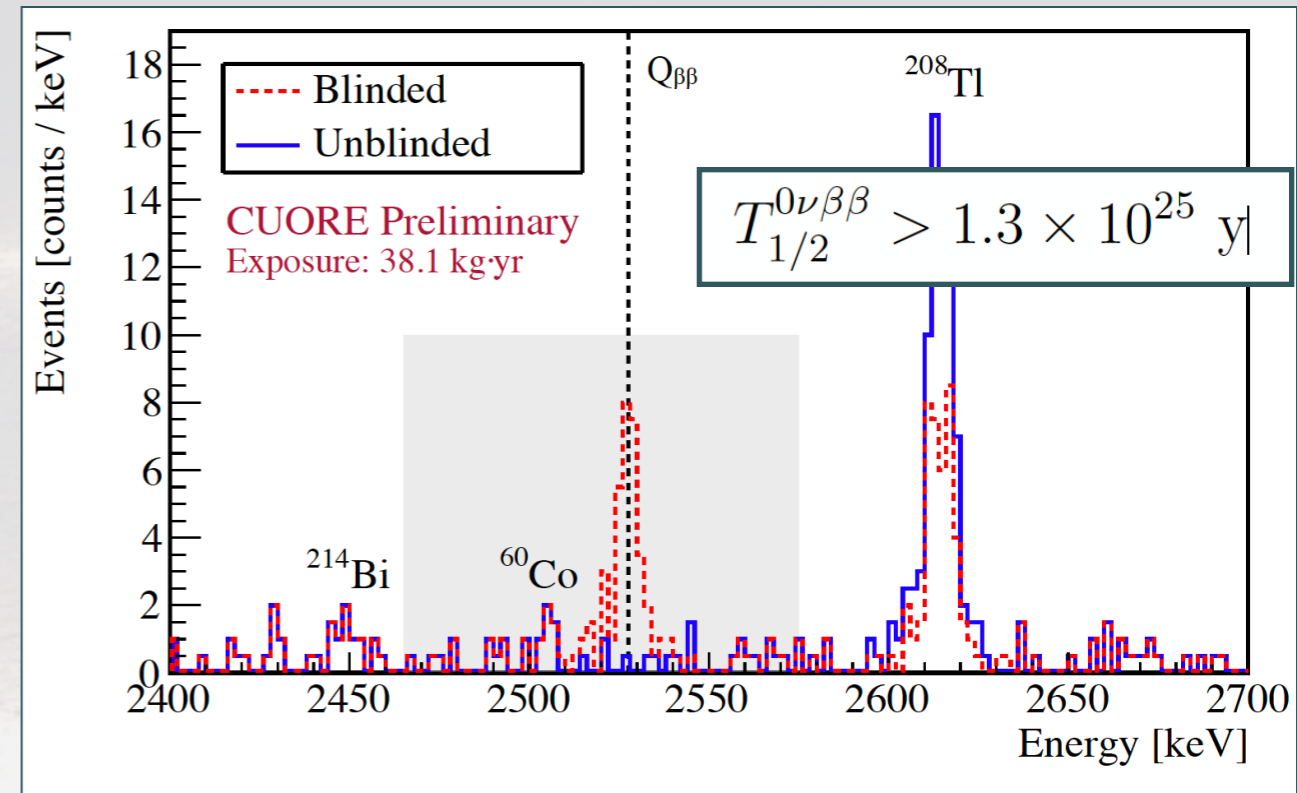
• Roughly, sensitivity from **10^{26} y to 10^{28} y**, corresponds to collider energy increase by a factor 20 (e.g. from Tevatron to LHC)

• Well....., with many less channels for exploration



- **CUORE:** Cryogenic bolometers ^{130}Te

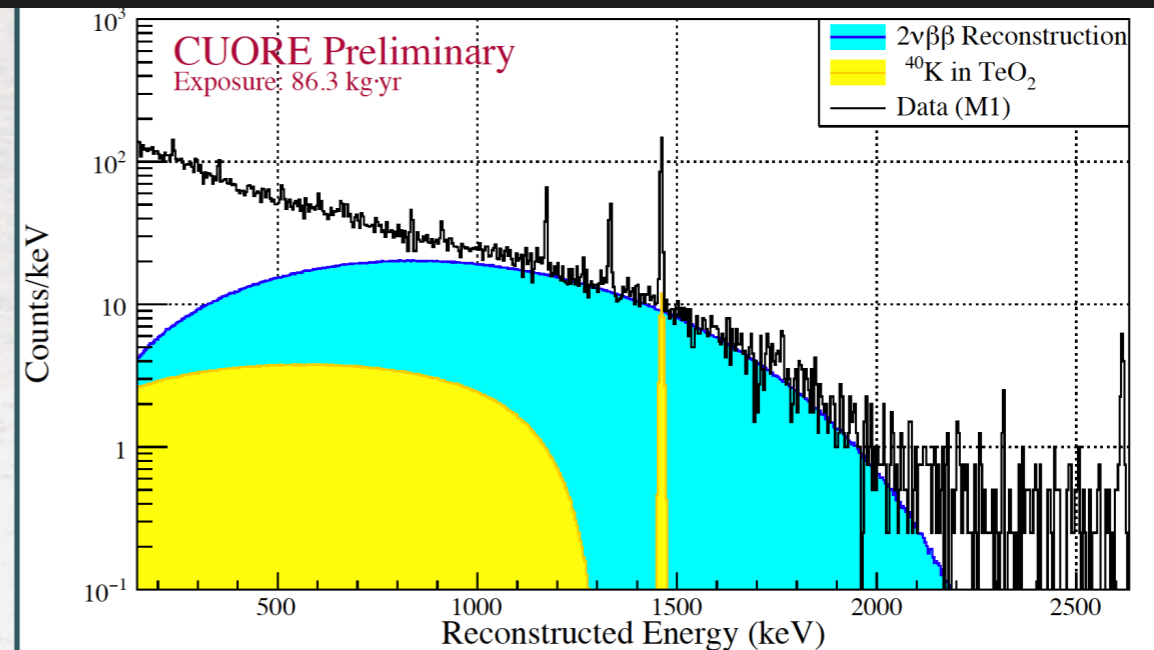
- The coldest m³ of the Universe!
- First release: Oct. 23rd, 2017
- Neutrino 2018: 2 neutrino signal



Measurement of the $2\nu\beta\beta$ Decay Half-life

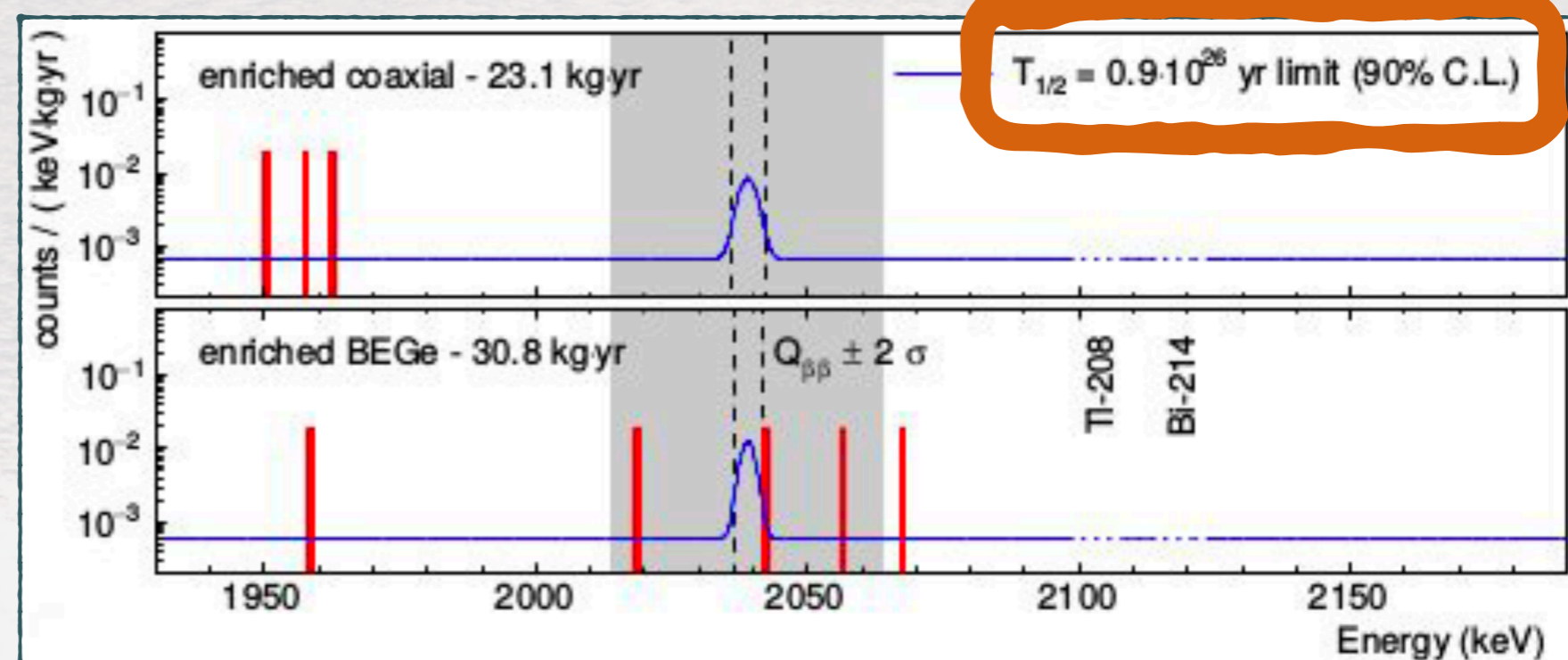
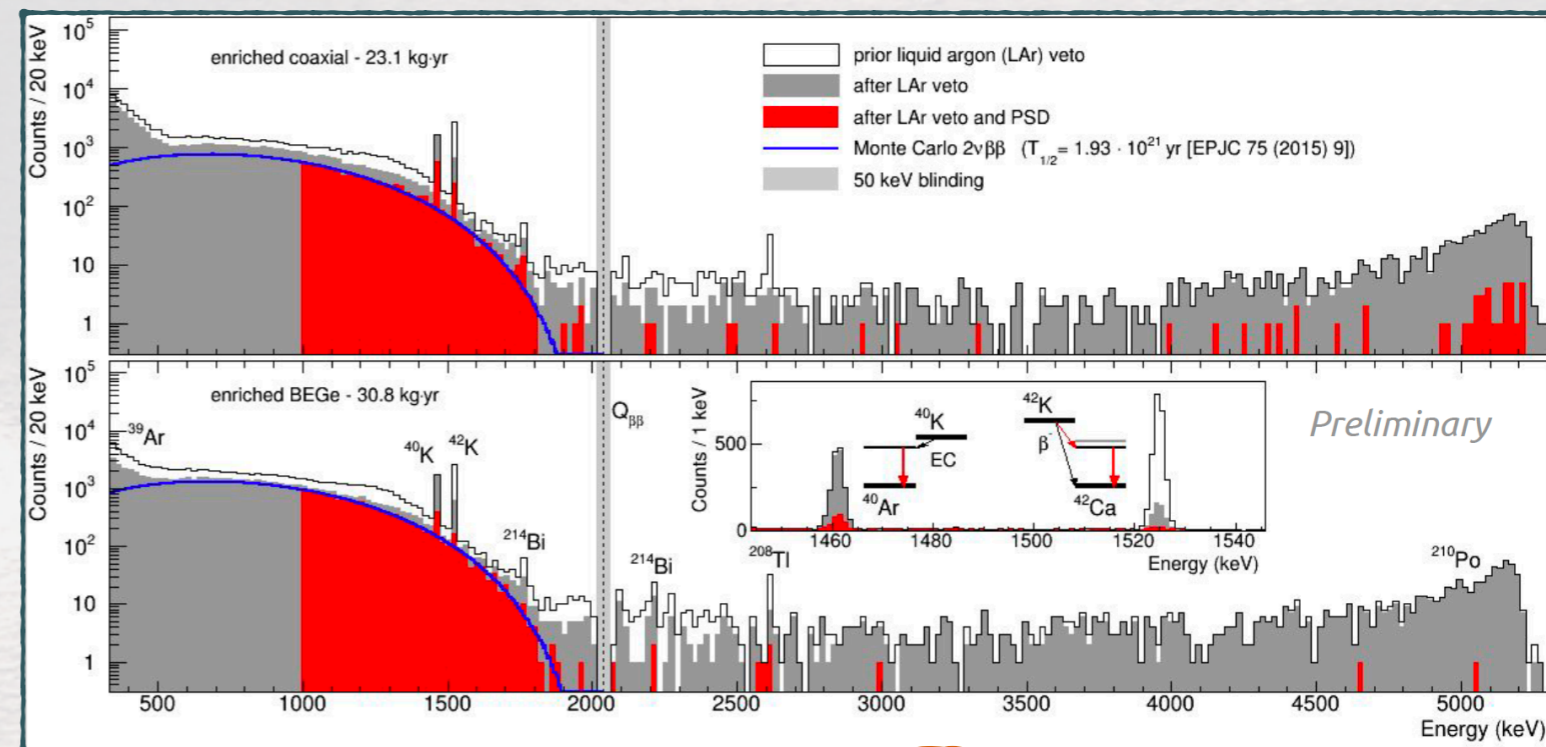
$$\Gamma_{1/2}^{2\nu} = [8.7 \pm 0.1 \text{ (stat.)} \pm 0.2 \text{ (syst.)}] \times 10^{-22} \text{ yr}^{-1} \quad (\text{Preliminary})$$

$$T_{1/2}^{2\nu} = [7.9 \pm 0.1 \text{ (stat.)} \pm 0.2 \text{ (syst.)}] \times 10^{20} \text{ yr}$$



- ^{76}Ge enriched, ~ 30 kg
 - First $0\nu\beta\beta$ experiment with zero background
 - $6 \cdot 10^{-4}$ c keV $^{-1}$ kg $^{-1}$
 - Upgrade at 200 kg approved
 - LEGEND 200

Neutrino 2018

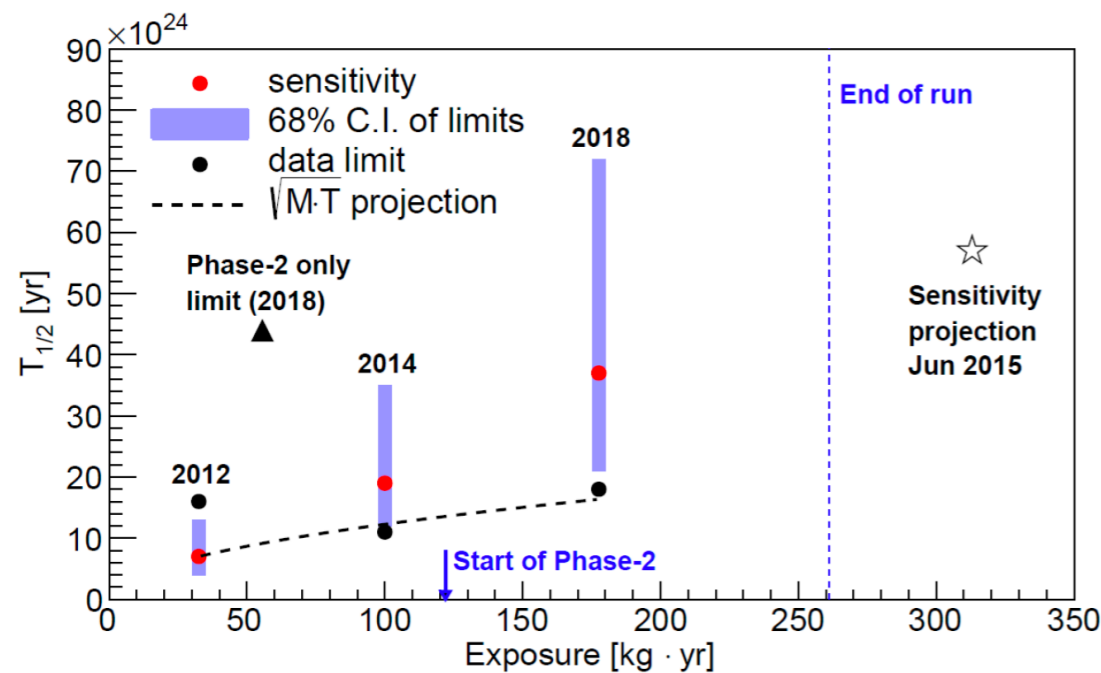


- EXO (liquid) and KamLAND-ZEN (dissolved in liquid scintillator) search for $0\nu\beta\beta$ of ^{136}Xe
 - Larger backgrounds, lower E resolution, but much larger masses (ton scale)

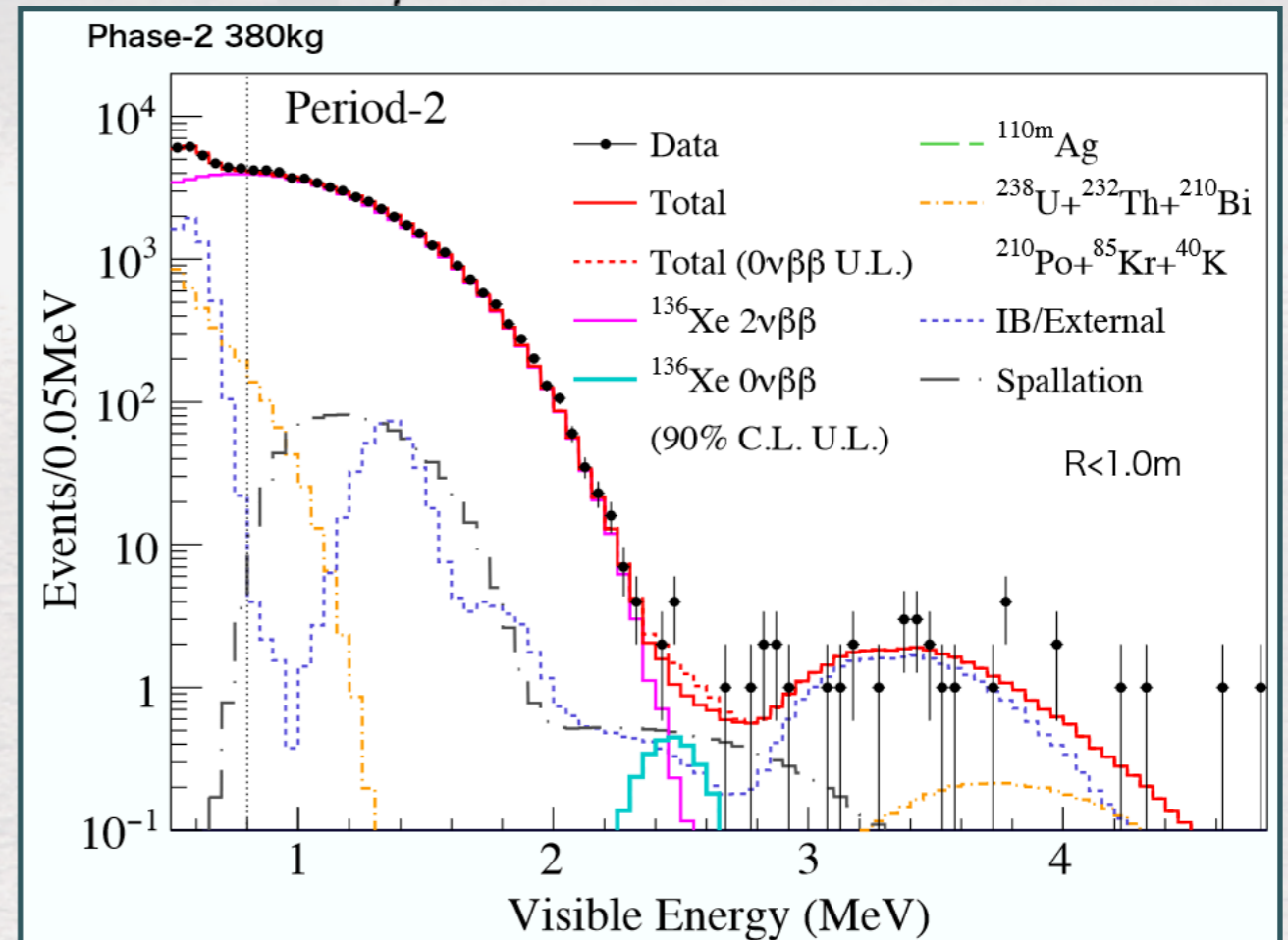
$$T_{1/2}^{0\nu} > 1.07 \times 10^{26} \text{ yr}$$

A brief history of EXO-200 results

	Sensitivity (yr)	90% CL Limit (yr)	$\langle m_{\beta\beta} \rangle$ (meV)
PRL 109, 032505 (2012)	0.7×10^{25}	1.6×10^{25}	
Nature 510, 229 (2014)	1.9×10^{25}	1.1×10^{25}	
PRL 120 072701 (2018)	3.8×10^{25}	1.8×10^{25}	147-398



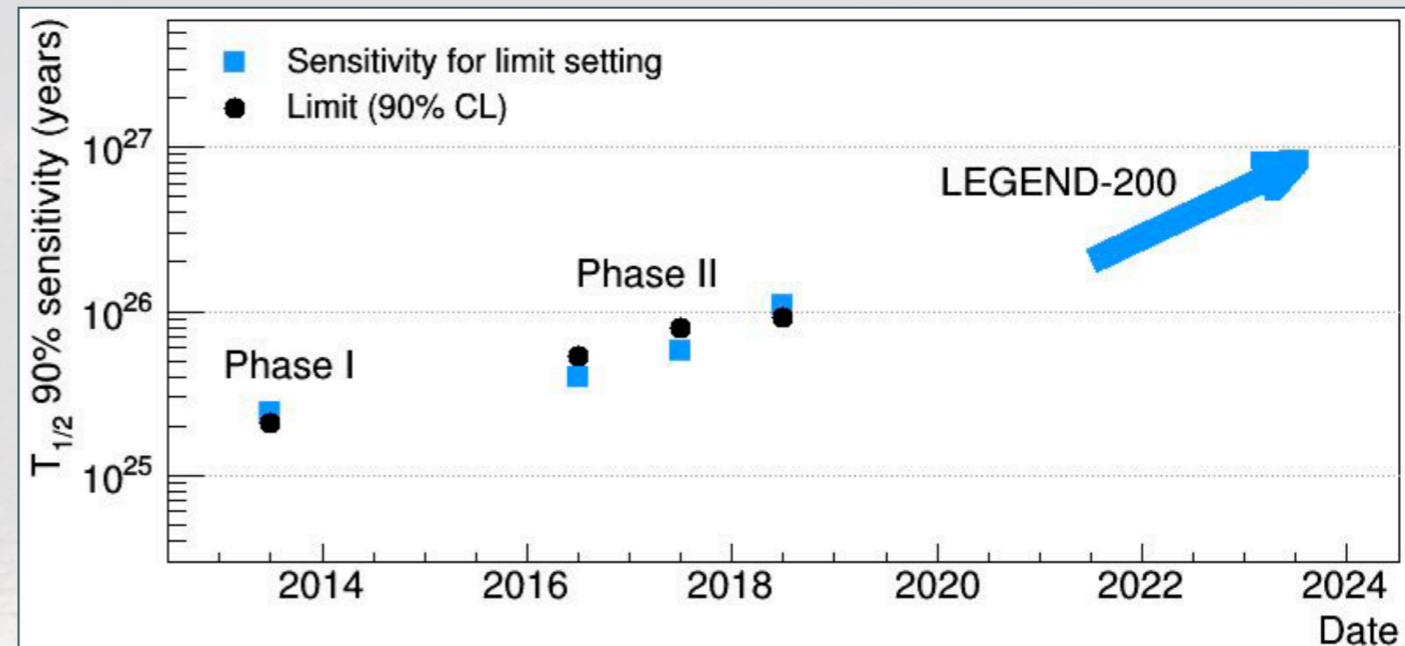
Gratta - Neutrino 2018



London - Dec. 2017

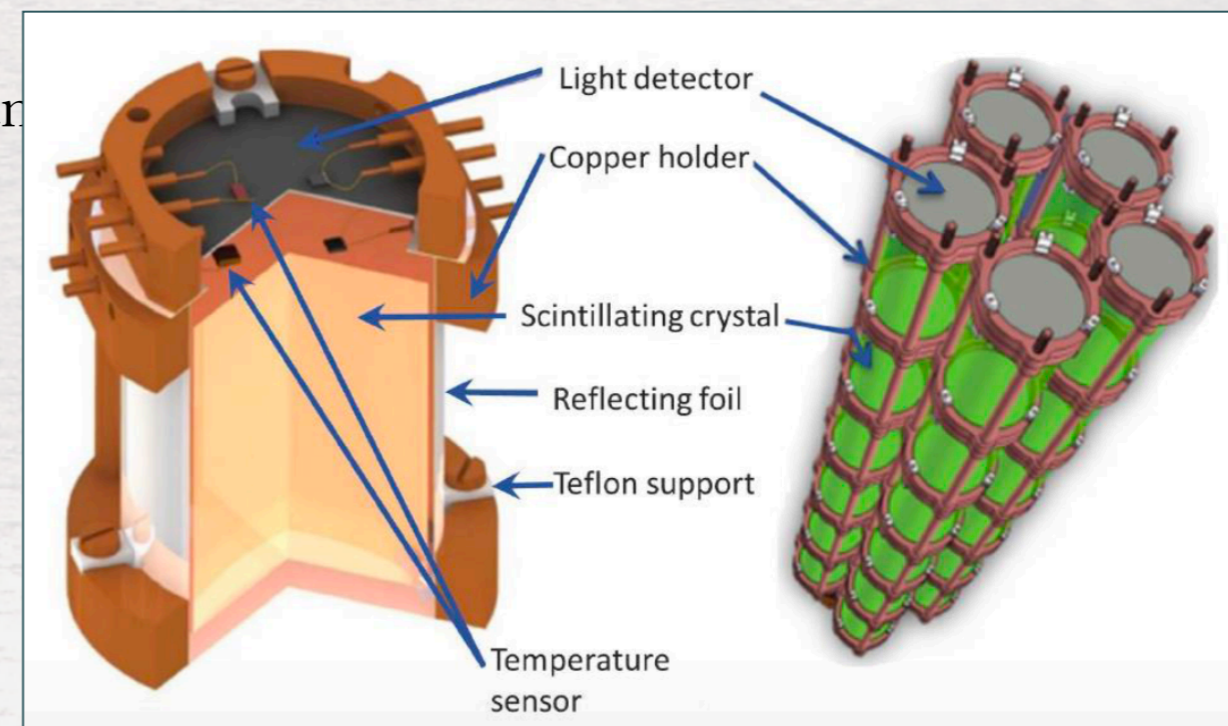
$$\langle m_{\beta\beta} \rangle < (61 - 165) \text{ meV}$$

- **^{76}Ge** : GERDA II-2 and LEGEND 200 kg
 - LNGS 200 kg
 - Background free experiment possible
 - Agreement with US Majorana group



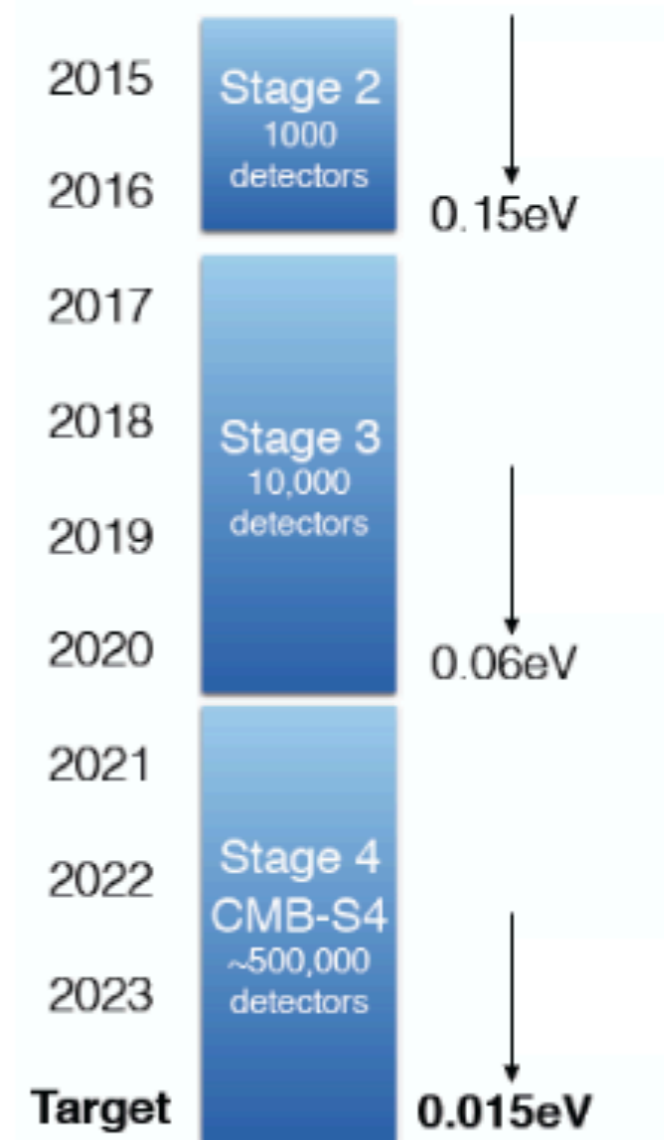
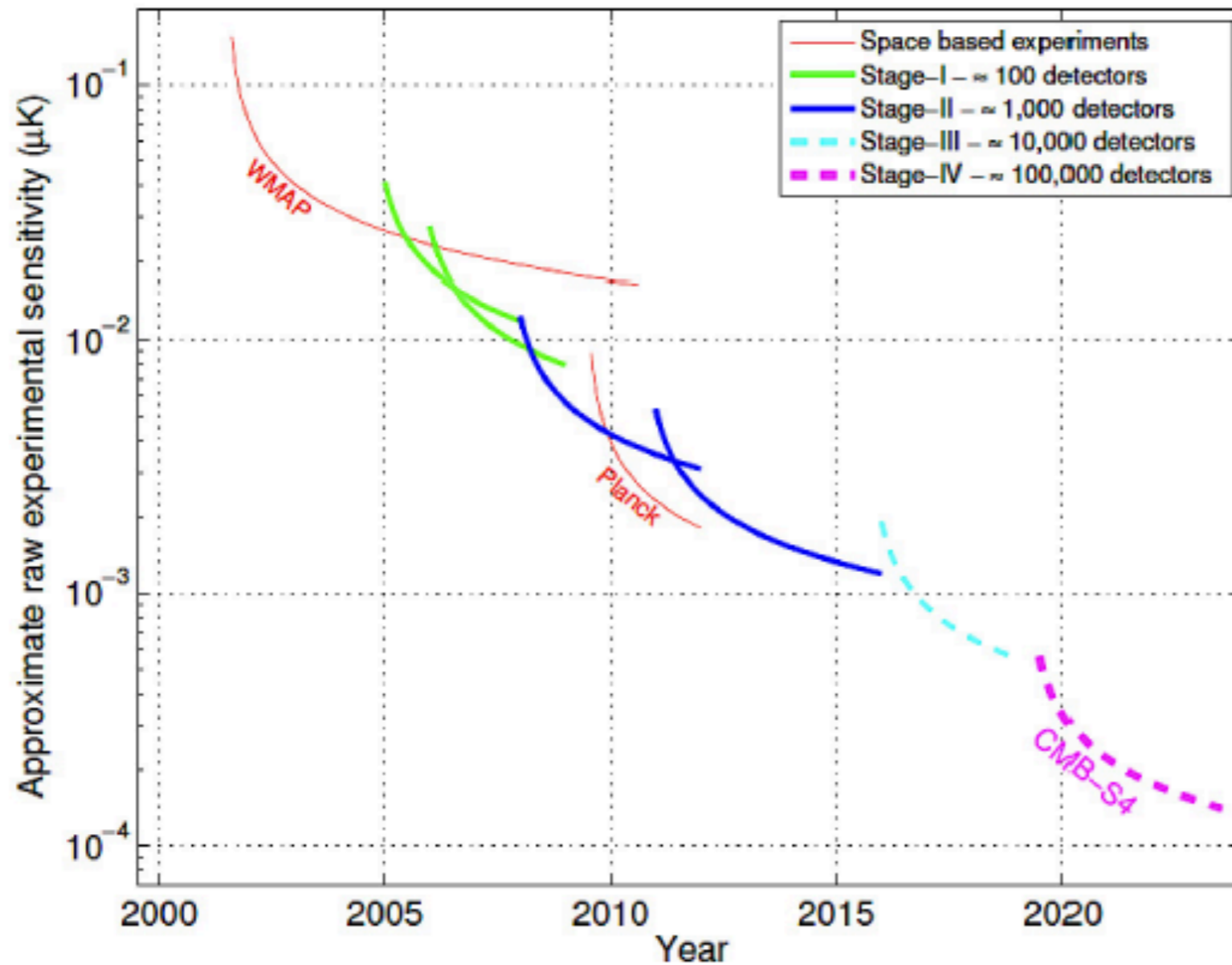
• Bolometers

- **CUPID** program to identify the best technology and materials
 - **Scintillating Bolometers ?**
 - **Cherenkov from Te ?**
 - **^{130}Te , ^{100}Mo , ^{82}Se ?**
- Zn^{82}Se run in progress and first results
- ^{100}Mo at Modane



Moore's law of CMB sensitivity in μK

...and expected error on Σm_ν



- $\Lambda \neq 0$? **RG wrong on large scale ?** **Field ?**

- **Mission:** Measure the evolution of gravitational potential and matter over 10 Gy ($z < 2$) using:

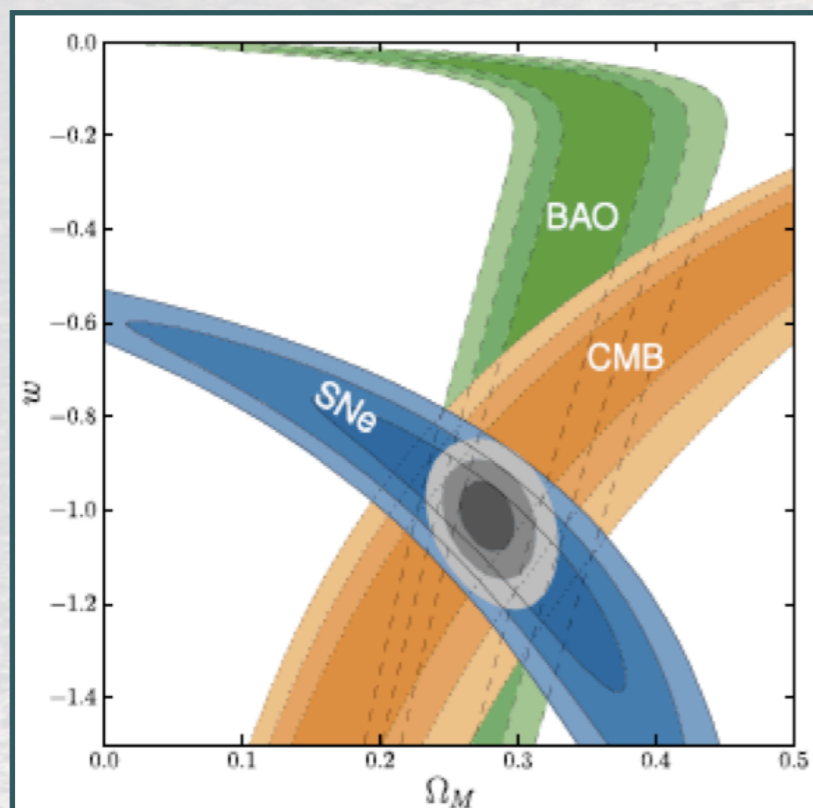
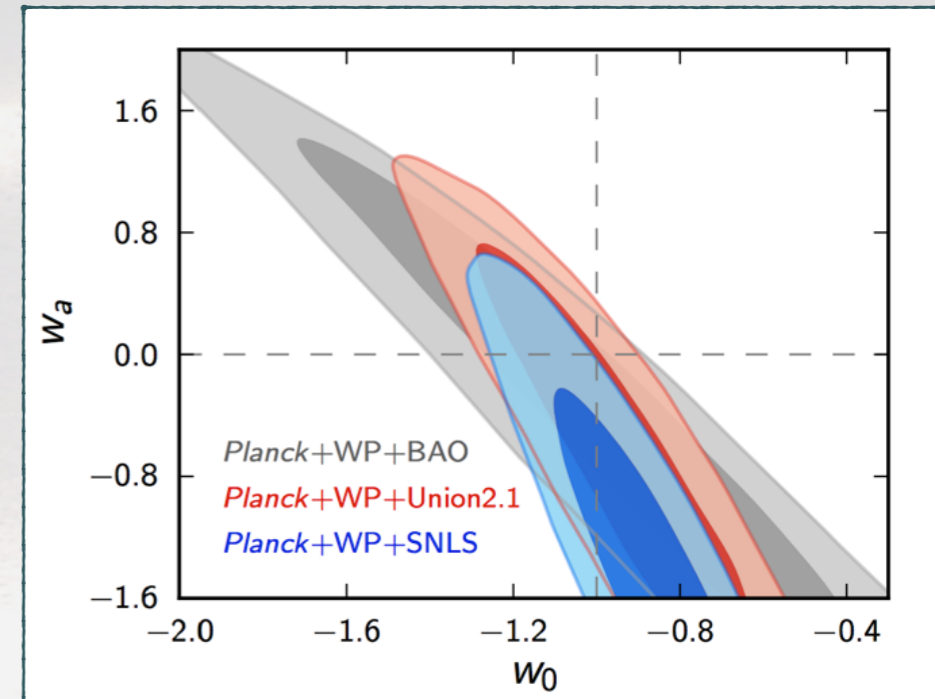
- Cluster formation (BAO)
- Weak Lensing

- **Neutrino mass ! Sensitivity down to 0.3 eV in the sum of all neutrinos**

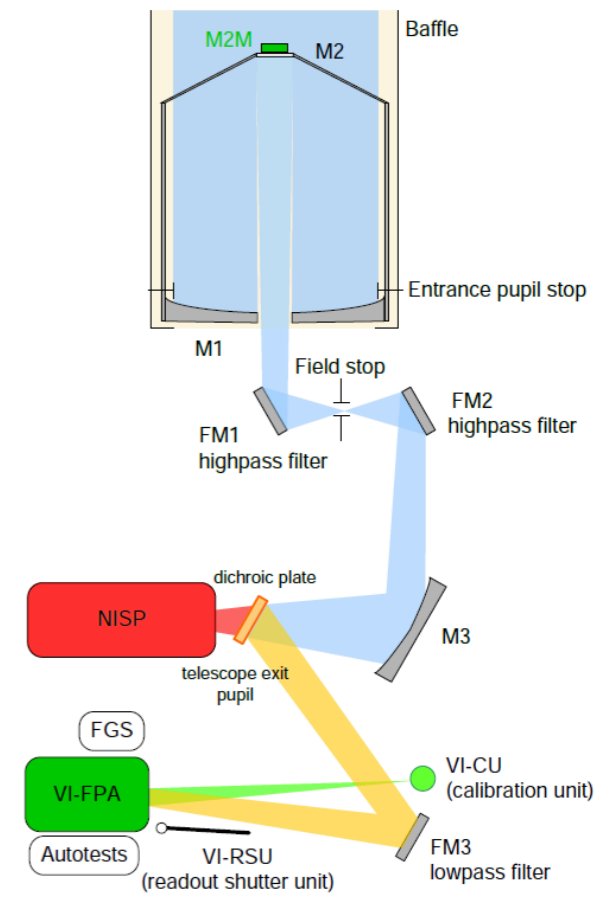
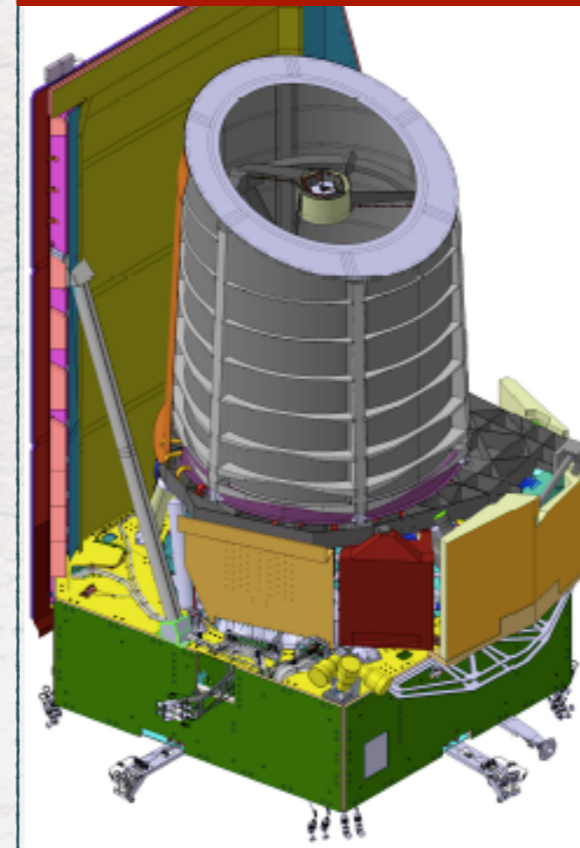
- **Two instruments:**

- Near **Infrared** Spectro-Photometer (NISP)
- **Visible CCD (VIS)**

$$w = w_0 + w_a(1-a)$$



Launch 2022 - 6.5 y



- Not covered for lack of time
 - Direct neutrino mass with **Katrin, Project 8**, microBolometers
 - Many **sterile neutrino searches** at reactors
 - Many **DBD experiments around the world**
 - Many future projects (e.g. ESS)
- **Neutrino physics is a very active field of research**
- Next decade might bring:
 - Hierarchy and CPV in lepton sector
 - Neutrino mass scale (likely from Cosmology first)
- **Hard to tell**
 - DBD ?
 - Sterile ?

**Neutrinos are always
Surprising !**