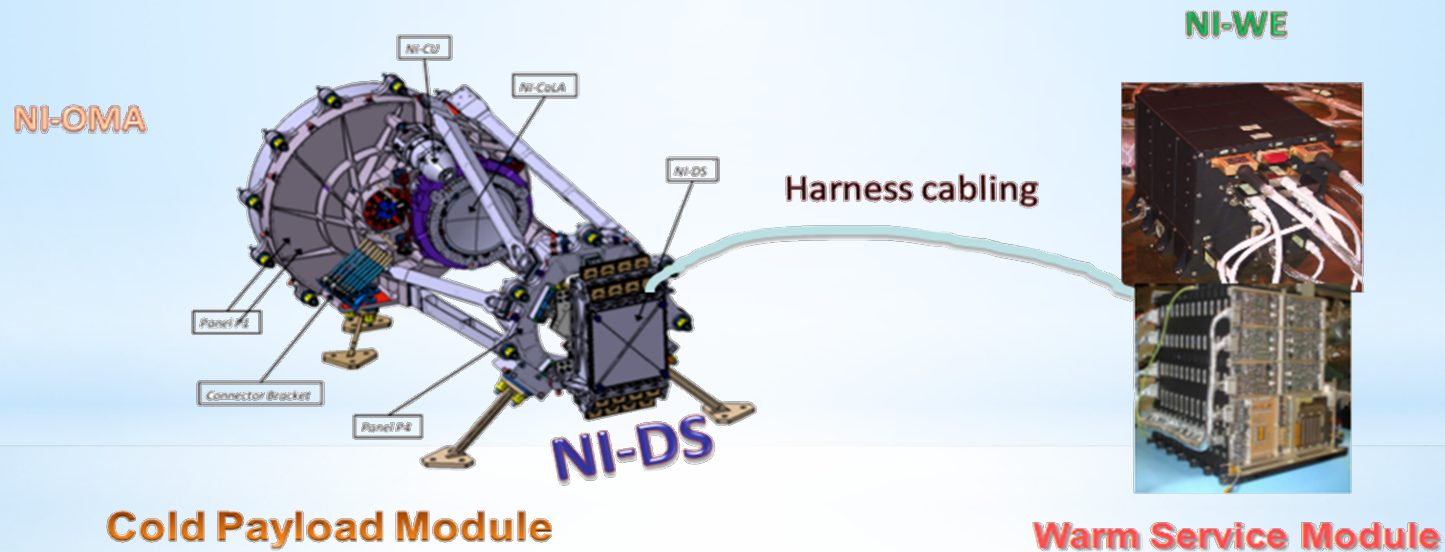


Assembly Integration and Verification (AIV) of the Near Infrared Spectro-Photometer's (NISIP) Warm Electronics (WE) in the EUCLID mission



Fulvio Laudisio



UNIVERSITÀ
DEGLI STUDI
DI PADOVA



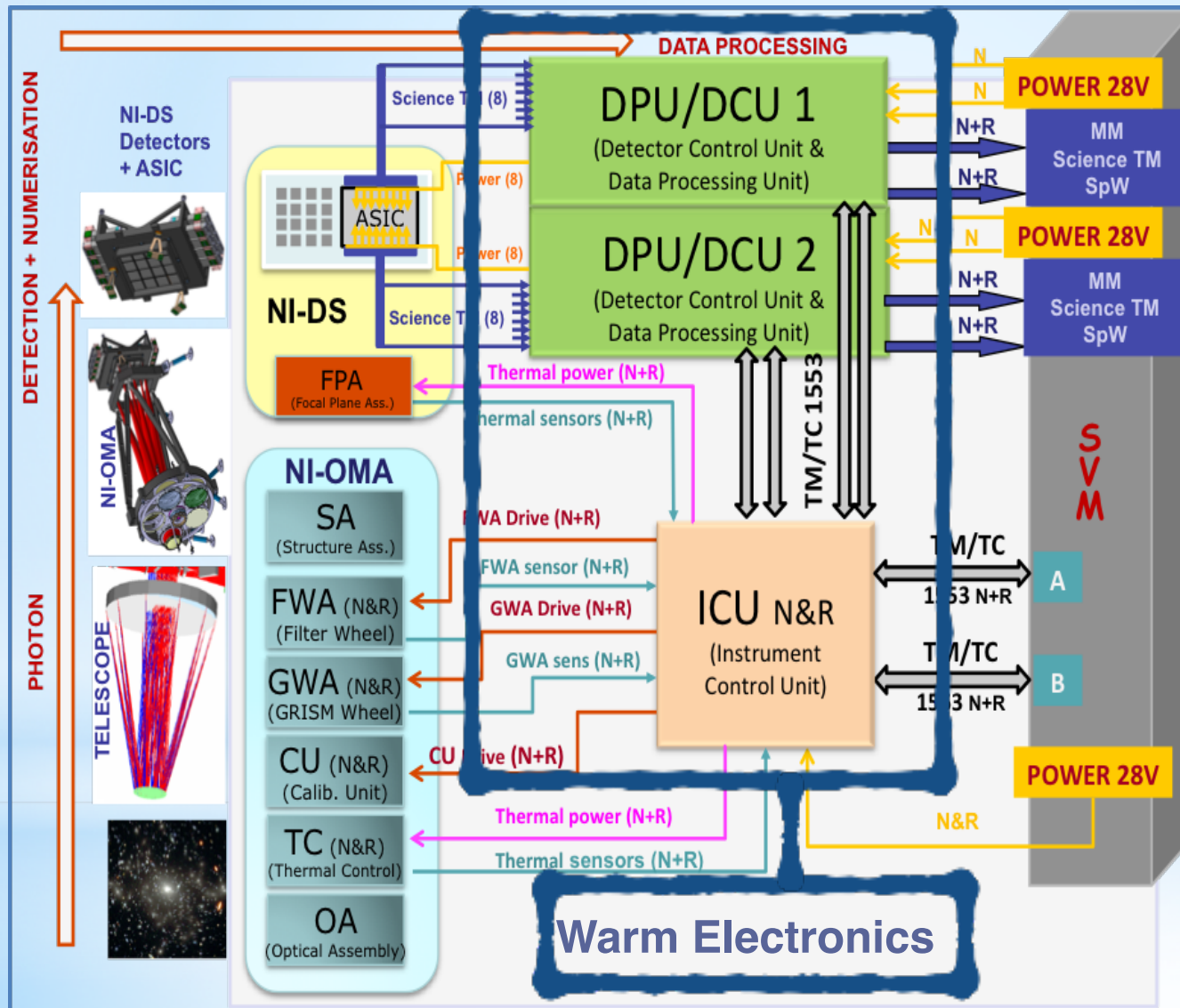
The Euclid Mission



•ESA mission

- Selected in Oct. 2011 - Fully funded
- Partners: ESA, TAS, Airbus DS, Euclid Consortium (EC)
- Overall mass: ~2020 kg, Power : 1920 W (EOL)
- Data rate: 850 Gbit/day
- Telescope (T=125K, passive):
 - 1.2m aperture primary, 3 mirror Korsch anastigmat
- 2 Instruments (VIS, NISP) – T = 100-140 K (passive)
 - Wide field instrument, VIS: 36 e2v 4kx4k CCDs 0.55< λ <0.92 μm , 576 M pixels, 0.11 arcsec/pix, 0.53 deg² FoV
 - Photom. (Y, J, H) +spectrom.: 16 H2GR HgCdTe detectors;
 - 64 Mpixels, 0.30 arcsec/pix, 0.53 deg² FoV (=VIS)
 - Grism slitless spectro (1B + 3R grisms) 0.92< λ <2.05 μm , R>250
- Downlink Rate: X/X + K-band to Ground Station 55 Mbits/s. 850 Gbit/day to transfer 4hr/day.
- Ground Segment: ESA (50%,) EC (50%, EC leads science and external data): 1.5 billion galaxies for WL, 30 million redshifts, 12 billion sources (3sigma)
- L2 orbit
- Launch Vehicle – Soyuz-Fregat
- Launch date 2020, from Kourou space port
- 6.25 years mission + additional surveys (exopl, SN)
- Main surveys: 15,000 deg²+40 deg² 2 mag. deeper
- Science drivers: DE
- Science leads: Euclid Consortium

NISP Warm Electronics



DPU/DCU

- Data acquisition
- Data processing
- Data compression
- Data transfer to satellite memory

ICU

- Filter wheel & grism wheel control
- Telecommands dispatching
- Telemetry acquisition and transfer to SVM

Activities for the AIV

- ✓ Development of simulators for the missing hardware: ad hoc software running on pc implementing the communication interface.
- ✓ Preparation of test sequences for each requirement declared on the starting documents.
- ✓ Running the tests and compile the reports
- ✓ Write the documentation for all the operations that will be repeated on the upcoming hardware for the next phases (Electro Qualified Model and Flight Models)

Hardware Testing

- All the tests will be repeated in a space like environment.
- Thermal, and electrical, response of the hardware and of the detectors will be checked at LAM (Laboratoire d'Astrophysique de Marseille)



45 m³ crio-vacuum chamber
77K and 10⁻⁶ mbar

Thank you for your attention

May I Introduce Myself



Daniela Piccioni Koch daniela.piccioni@kit.edu
KIT – SCC Germany

Ninth  International School on

Architectures, tools and methodologies for developing
efficient large scale scientific computing applications

Ce.U.B. Bertinoro (FC) 23- 28 October 2017



Karlsruhe Institute of Technology (KIT) The Research University in the Helmholtz Association



The Karlsruhe Institute of Technology (KIT), was established by the merger of the ForschungsZentrum Karlsruhe (FZK) GmbH and the Universität Karlsruhe (TH) on 1st October, 2009.



With more than 9,000 employees and an annual budget of about EUR 785 million, KIT is one of the biggest research and education institutions worldwide.

KIT combines the tasks of a university of the state of Baden-Württemberg with those of a research center of the Helmholtz Association in the areas of research, teaching, and innovation.



www.kit.edu

Universität Karlsruhe was founded in 1825 as a Polytechnic and ForschungsZentrum Karlsruhe was founded in 1956 as Nuclear Reactor Facility, whose first reactor (named FR2) started operation in 1962 .

Karlsruhe – The Fan(Fun) City



First Draisine in 1817
by **Karl Freiherr von Drais** (KA, 1785-1851)



“A castle at the center, from which the streets radiate in fan-like form: so 300 years ago, according to legend, the city founder Margrave Karl-Wilhelm von Baden-Durlach, saw his residence in a dream. On 17 June 1715, he had the foundation stone laid that would make his dream a reality.” (Dr. Frank Mentrup, Chief Burgomaster of Karlsruhe).

Karlsruhe Main Train Station Nowadays



„Yes, but with



!“

The Steinbuch Centre for Computing (SCC) is the information technology centre of KIT, which is named after the Karlsruhe computer scientist Karl Steinbuch (1917-2005). Its activities focus on research, development and innovation in the fields of **high performance computing (HPC)**, **data intensive computing (DIC)**, and **secure IT federations**. SCC's research focuses on scientific computing, modelling and simulation, data analysis, grids, cloud and cluster computing, innovative network technologies, IT security, and virtualisation. SCC operates very powerful HPC systems, which are used by regional and federal scientists from academia and industry for computationally intensive projects.

In this context the SCC **Simulation laboratories (SimLabs) 1.Energy, 2.Nanomicro, 3.Climate and Environment** represent the interfaces between users and High Performance Computing providers. Their main task is the software enhancement for an efficient use of supercomputers and distributed systems in interdisciplinary research and development. <https://www.scc.kit.edu/en/research/5960.php>

KIT HPC Systems – bwUniCluster

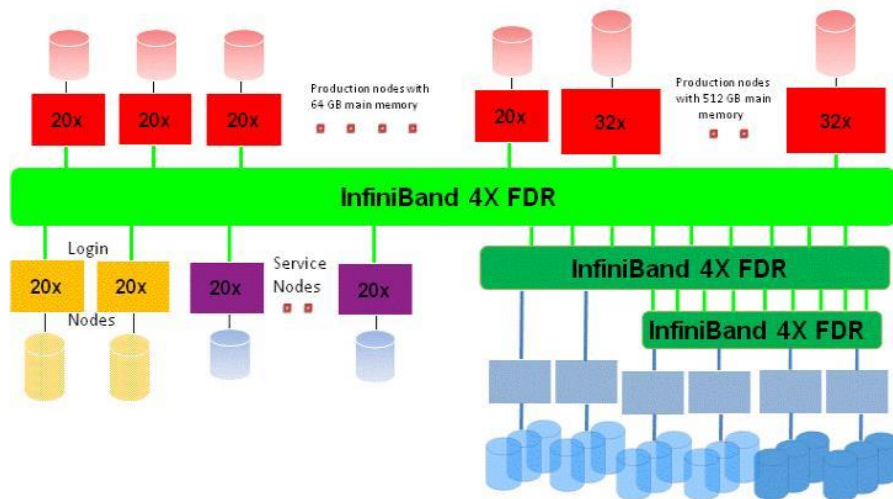
The bwUniCluster is a parallel computer with distributed memory. Each node has sixteen Intel Xeon processors, local memory, disks and network adapters. All nodes are connected by a fast InfiniBand 4X FDR interconnect. In addition the file system Lustre, that is connected by coupling the InfiniBand of the file server with the InfiniBand switch of the compute cluster, is added to bwUniCluster to provide a fast and scalable parallel file system.

	Compute nodes "Thin"	Compute nodes "Fat"	Compute nodes "Broadwell"	Login nodes "Thin / Fat"	Login nodes "Broadwell"	Service nodes
Number of nodes	512	8	352	2	2	10
Processors	Intel Xeon E5-2670 (Sandy Bridge)	Intel Xeon E5-4640 (Sandy Bridge)	Intel Xeon E5-2660 v4 (Broadwell)	Intel Xeon E5-2670 (Sandy Bridge)	Intel Xeon E5-2630 v4 (Broadwell)	Intel Xeon E5-2670 (Sandy Bridge)
Processor frequency (GHz)	2.6	2.4	2.0	2.6	2.2	2.6
Number of sockets	2	4	2	2	2	-
Total number of cores	16	32	28	16	20	16
Main memory	64 GB	1024 GB	128 GB	64 GB	128 GB	64 GB
Local disk	2 TB	7 TB	480 GB	4 TB	480 GB	1 TB
Cache per socket	Level 1: 8x64 KB Level 2: 8x256 KB Level 3: 20 MB	Level 1: 8x64 KB Level 2: 8x256 KB Level 3: 20 MB	Level 1: 8x64 KB Level 2: 14x256 KB Level 3: 35 MB	Level 1: 8x64 KB Level 2: 8x256 KB Level 3: 20 MB	Level 1: 8x64 KB Level 2: 10x256 KB Level 3: 25 MB	Level 1: 8x64 KB Level 2: 8x256 KB Level 3: 20 MB
Interconnect	4xFDR	4xFDR	4xFDR	4xFDR	4xFDR	4xFDR



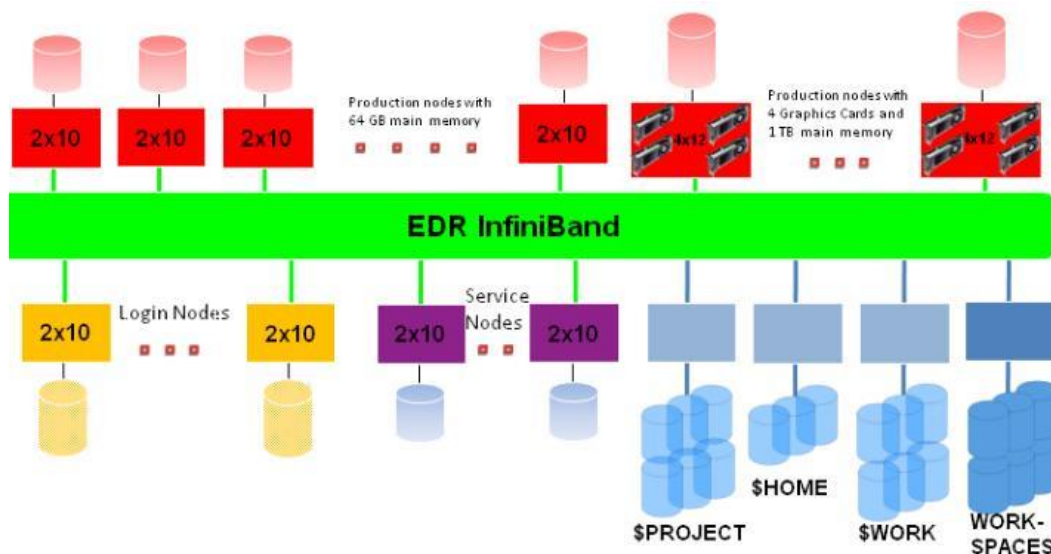
KIT HPC Systems – ForHLR I

- **512** 20-way Intel Xeon compute nodes. 400 GFLOPS peak performance and 64 GB main memory per node.
- **16** 32-way Intel Xeon compute nodes. Each of these nodes has 665,6 GFLOPS peak performance and 64 GB main memory
- **2** 20-way Intel Xeon login nodes. 400 GFLOPS peak performance and 64 GB main memory per node.
- **10** 20-way Intel Xeon service nodes. 400 GFLOPS peak performance and 64 GB main memory per node.



KIT HPC Systems – ForHLR II

- 1152 20-way Intel Xeon compute nodes. Each of these nodes has a peak performance of 832 GFLOPS and 64 GB of main memory.
- 21 48-way Intel Xeon rendering nodes. Each of these nodes has 4 NVIDIA GeForce GTX980 Ti graphic cards and 1 TB of main memory.
- 5 20-way Intel Xeon login nodes. Each node has 256 GB main memory.
- 8 20-way Intel Xeon service nodes, with 64 GB of main memory each.





Karlsruhe Institute of Technology



www.bioliq.de

Simulation Lab Energy

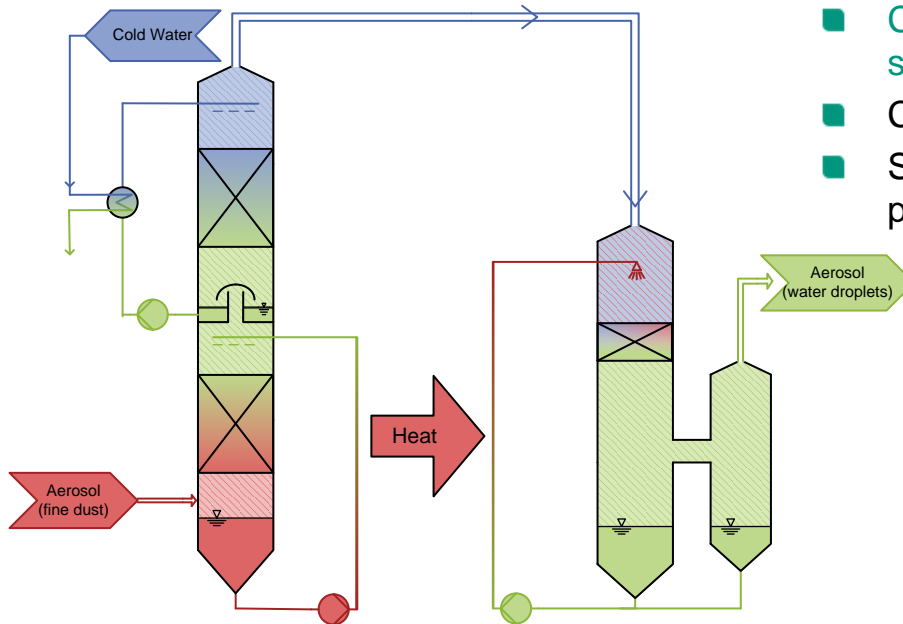
Modelling and simulation of electric power systems for the *Energiewende* in the frame of the Energy Lab. 2.0 Project and in cooperation with the Institute for Applied Computer Science (IAI) of the Karlsruhe Institute of Technology



www.elab2.kit.edu

Possible applications

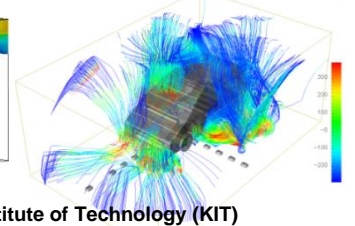
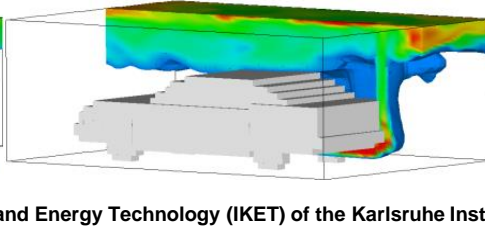
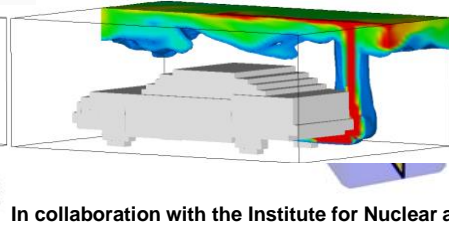
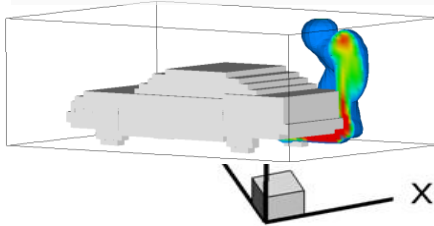
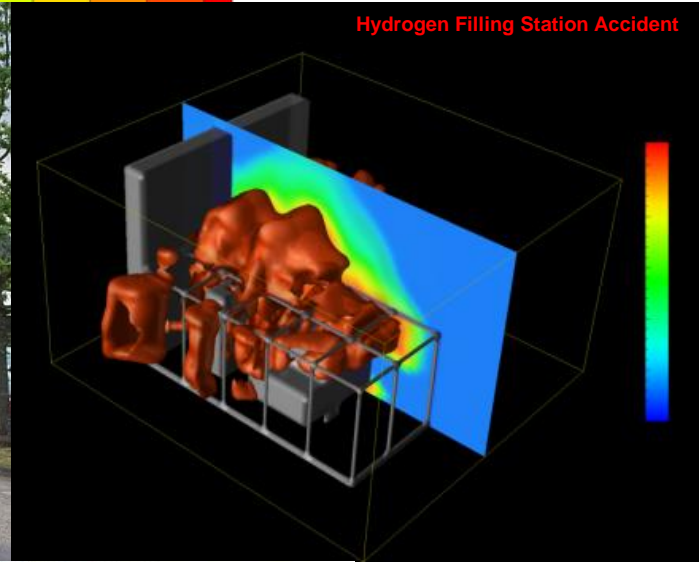
- Column design for condensation assisted fine dust separation
- Cleaning of industrial acidic gases
- Simulation of the behaviour of pyrolysis oil for the production of fuel from biomass (e.g. bioliq)



Concept for a condensation facility with two packed columns and reutilisation of thermal energy developed at ITTK.

Simulation of a contact device to induce heterogeneous condensation

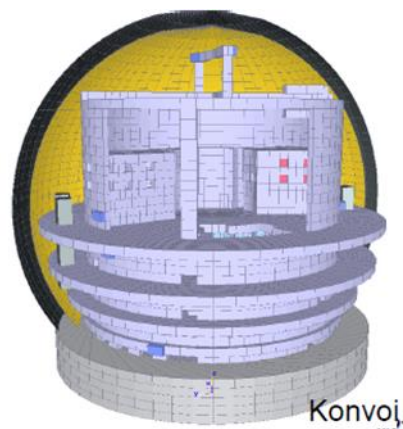
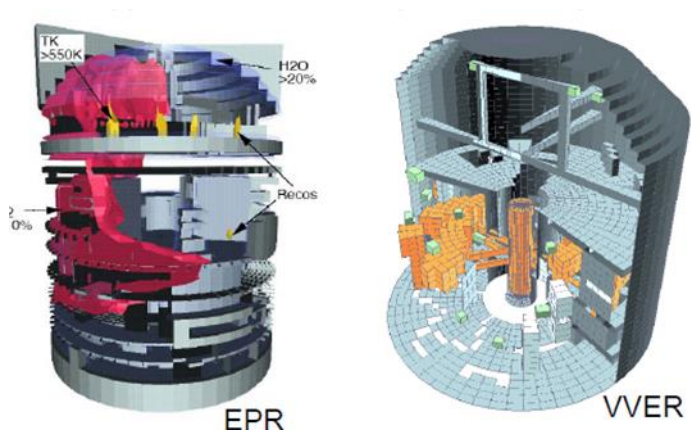
Automotive Hydrogen Safety Simulations with GASFLOW



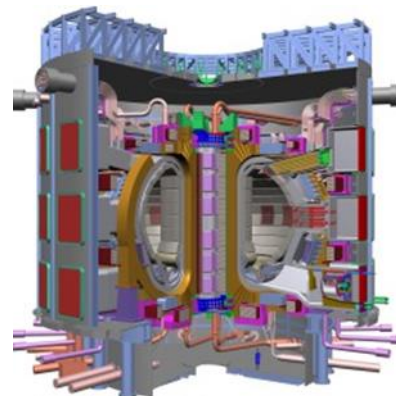
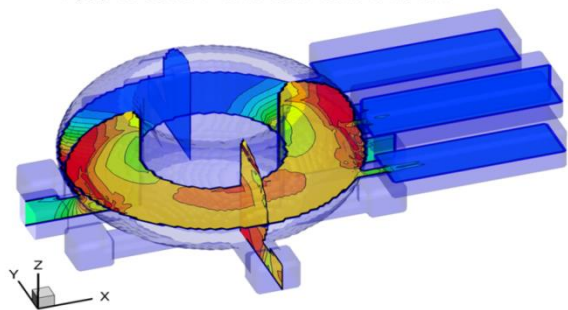
X

In collaboration with the Institute for Nuclear and Energy Technology (IKET) of the Karlsruhe Institute of Technology (KIT)

Hydrogen Safety Simulations for Nuclear and Fusion Research



Simulations for Nuclear Reactor Safety



Simulations for ITER explosion scenarios

Thank you for your attention.

ESC17

Marco Roetta

Software Developer at INFN - LNL

Control systems for PIAVE accelerator

PLC based systems for auxiliary services

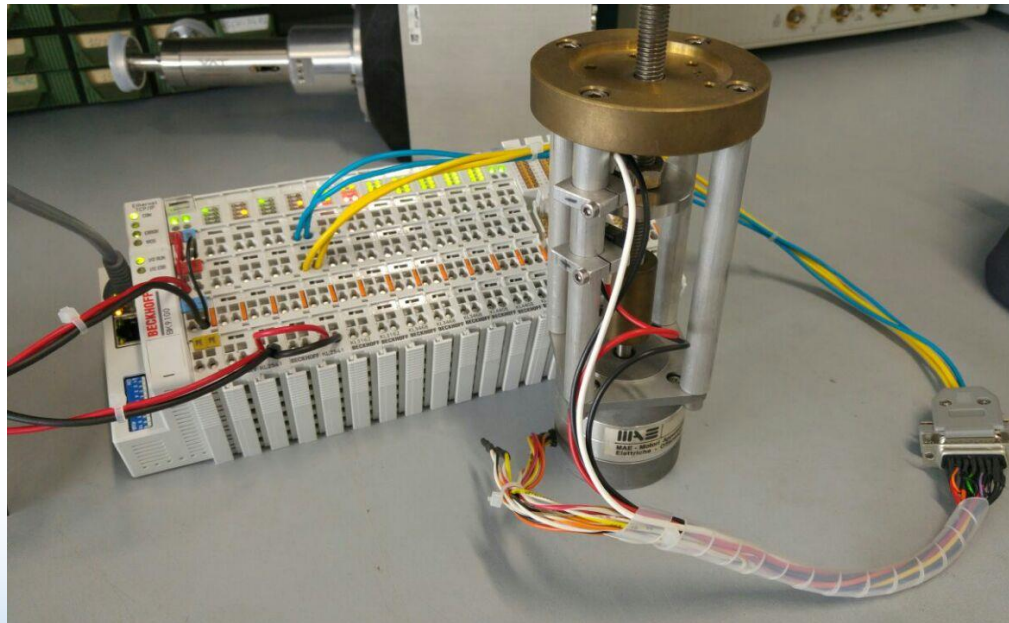
ECR Ion Source - 1

- Software Developer for the ECR Ion Source
- EPICS for logic and hardware communication systems
- CSS and Python for the operator GUI
- Control the hardware inside the HV platform form different remote consoles
- Monitor beam output (Faraday Cup) and machine components status
- Automated spectrum acquisition



ECR Ion Source - 2

- Divided in three components: GUI, PLC, VM (logic)
- Connected using an Ethernet network (copper, fiber)
- The system will be reused to control a Charge Breeder for the SPES project
- Same type of machine (plasma chamber) with a different scope



ESC17

Flavor Physics and Lattice QCD



Giorgio Salerno

Flavor Physics and Lattice QCD

The Standard Model does not:

- include gravity ($M_{\text{Planck}} = (\hbar c/G_N)^{1/2} \approx 10^{19} \text{ GeV}$)

FLAVOR



- explain the origin of flavor, i.e. masses, mixing and CPV ($M_{\text{Flav}} = ??$)

- give a natural explanation of the hierarchy problem ($M_{\text{Weak}} \ll M_{\text{Planck}}$)

- provide (exact) gauge coupling unification ($M_{\text{GUT}} \approx 10^{15}-10^{16} \text{ GeV}$)

- explain the smallness of neutrino masses ($m_\nu \approx (\lambda v)^2/M$, $M \approx 10^{15} \text{ GeV}$)

- produce the observed matter-antimatter asymmetry

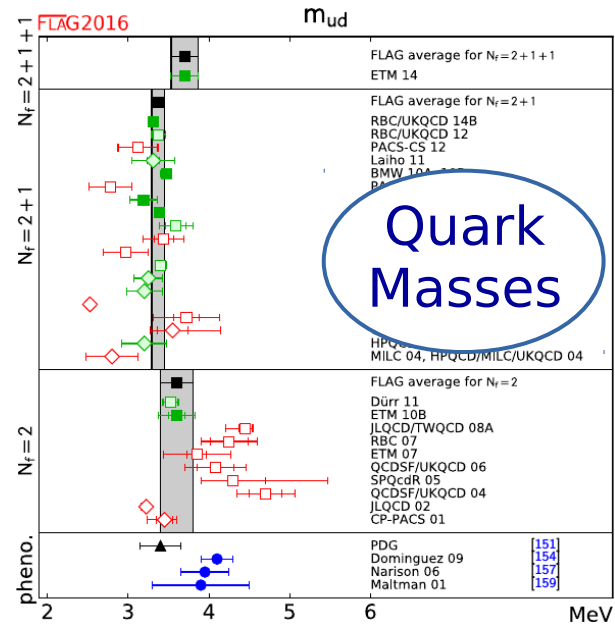
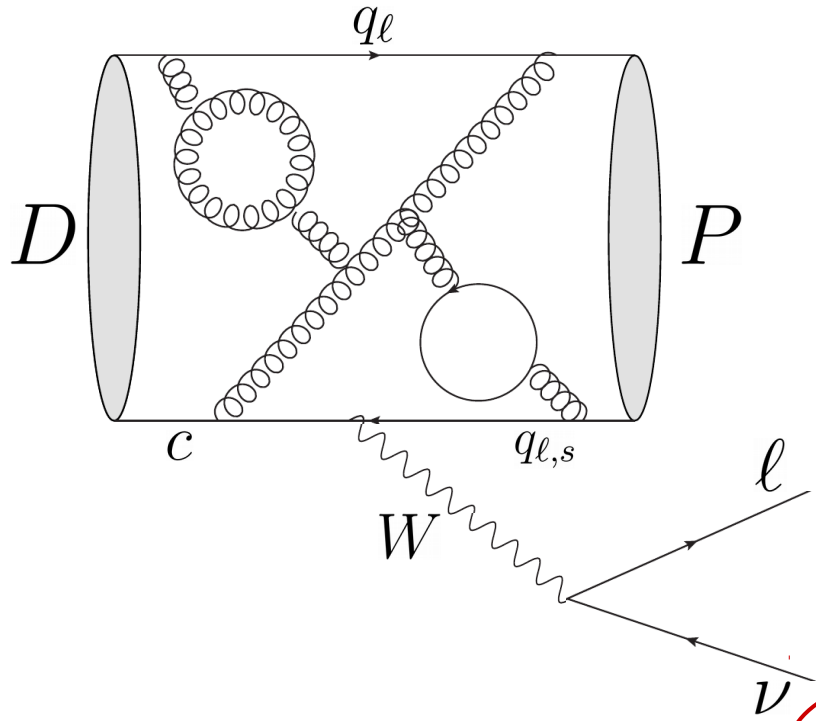
- provide a viable dark matter candidate

- explain "dark" (vacuum) energy

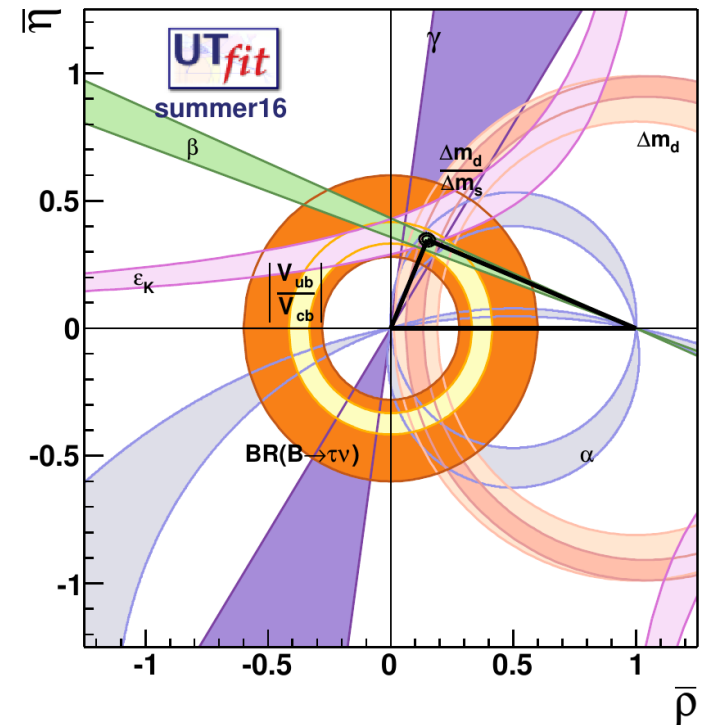
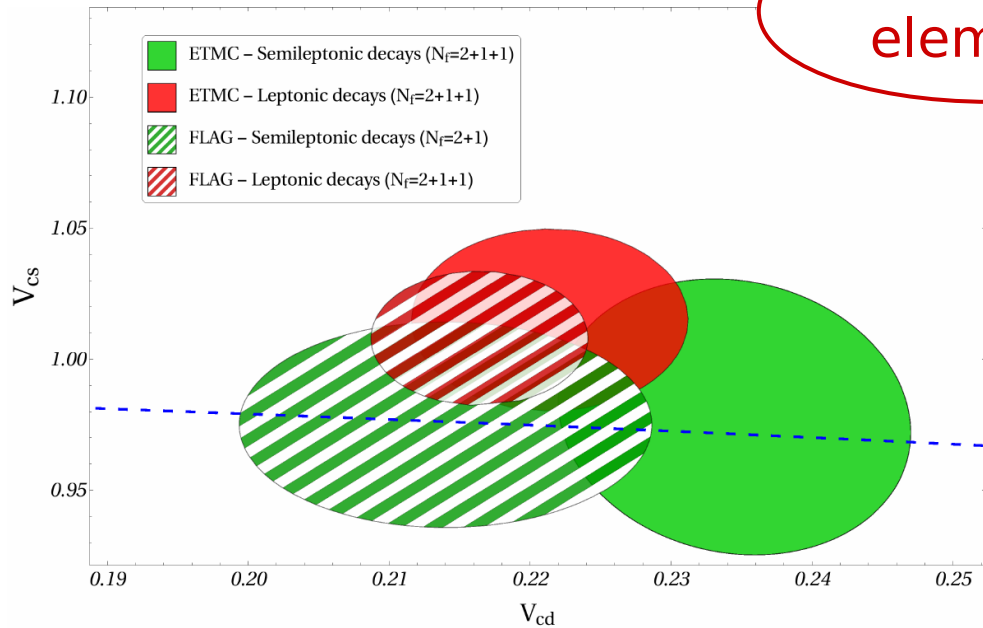
Cosmology



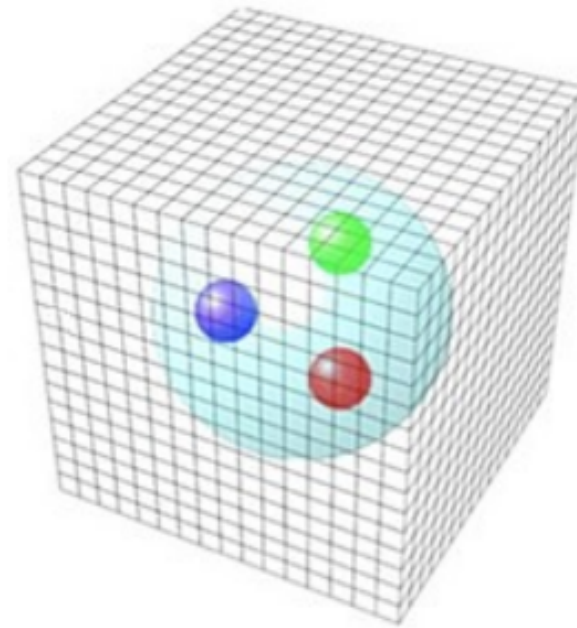
Flavor Physics and Lattice QCD



CKM matrix elements



Lattice QCD



Non-perturbative long-distance QCD
contributions from first principles

QCD simulated on a discrete
space-time and finite volume

Flavor Physics and Lattice QCD

$$C(t) \equiv \langle O_\pi(t) O_\pi^\dagger(0) \rangle \approx \sum_n \frac{|\langle 0 | O_\pi(0) | n \rangle|^2}{2E_n} \exp[-E_n t] \xrightarrow{t \rightarrow \infty} \frac{|\langle 0 | O_\pi(0) | \pi \rangle|^2}{2M_\pi} \exp[-M_\pi t]$$



Calculated on the lattice (Importance sampling Monte Carlo)

$$\frac{1}{Z} \int \mathcal{D}[\phi] O_1(\phi) O_2(\phi) e^{-S_E[\phi]} \simeq \frac{1}{N_c} \sum_{c=1}^{N_c} O_1(\phi_c) O_2(\phi_c)$$

$$(\Delta O)^2 = \frac{1}{N_c} \sum_{c=1}^{N_c} (O(\phi_c) - \langle O \rangle)^2 \quad \longrightarrow \quad \text{Statistical error} \sim \frac{1}{\sqrt{N_c}}$$

Thank you for the attention

Performance Optimizations & HEP

Servesh Muralidharan

Senior Fellow

IT-DI-LCG, UP Team

24 Oct 2017

Understanding Performance

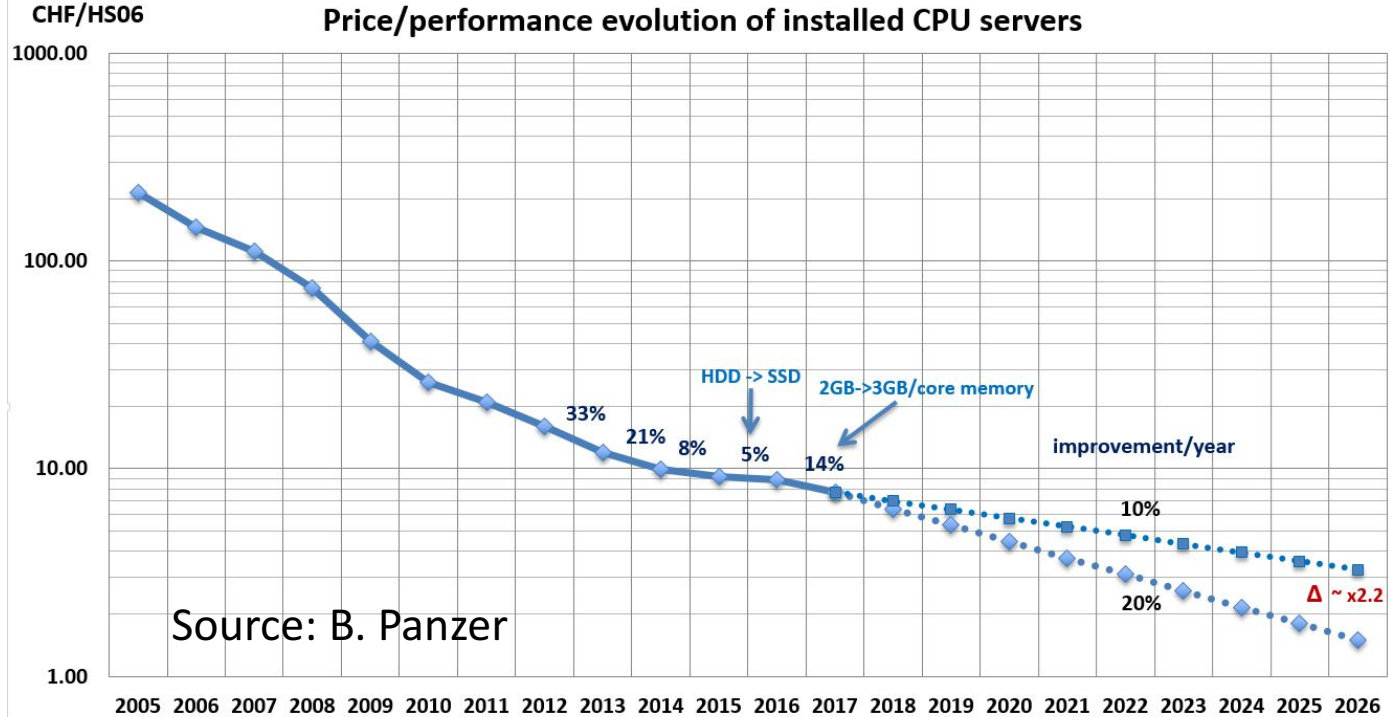
- **Team in IT**
 - Part of the WLCG team (IT-DI-LCG)
 - Closely working with the HSF and Software Technology Forum
 - Started January 2016
 - Regular meetings with some experiments : ATLAS, CMS, LHCb
- **Long term activity**
 - Focus on linking activities in the community
 - Software and infrastructure
 - Aggregation of knowledge about existing tools and data
 - Providing tools to understand, measure and improve performance
 - Analyze software and workflows
- **Members**
 - Markus Schulz
 - Andrea Sciaba
 - David Smith
 - Andrea Valassi
 - Servesh Muralidharan

Motivation - Gap between estimated compute resources for Run3 (~2021) and Run4 (~2024, HL-LHC)

Gap:

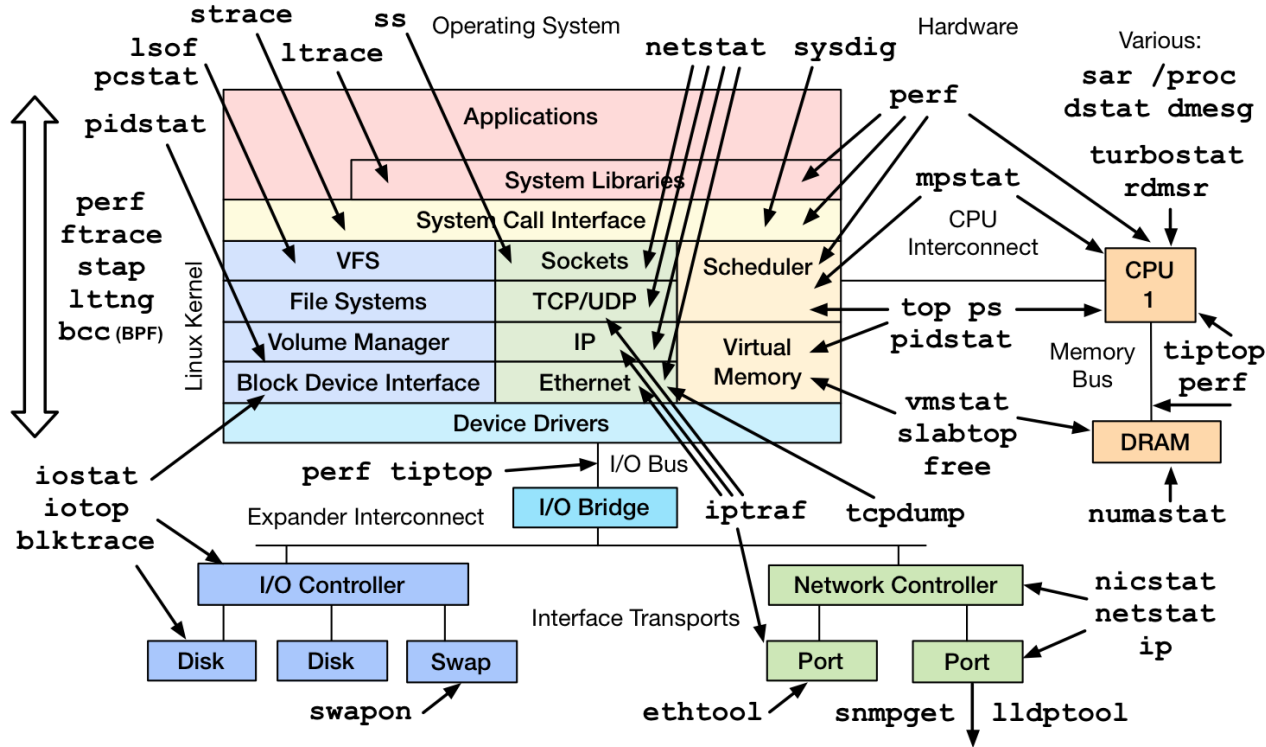
8 – 25x
(optimistic)

30 – 85x
(conservative)



Understanding Performance – Sounds Easy???

Linux Performance Observability Tools






Vectorization Advisor

Vectorization Advisor is a vectorization analysis toolset that lets you identify loops that will benefit most from vector parallelism, discover performance issues preventing from effective vectorization and characterize your memory vs. vectorization bottlenecks with Advisor Roofline model automation.


Program metrics

Elapsed Time	170.37s		
Vector Instruction Set	AVX, AVX2	Number of CPU Threads	20
Total GFLOP Count	936.64	Total GFLOPS	5.50
Total Arithmetic Intensity [®]	0.04		

Loop metrics

Total CPU time	2088.55s		100.0%
Time in 19 vectorized loops	7.02s		
Time in scalar code including time in 29 vectorized completely unrolled loops [®]	2081.53s		99.7%

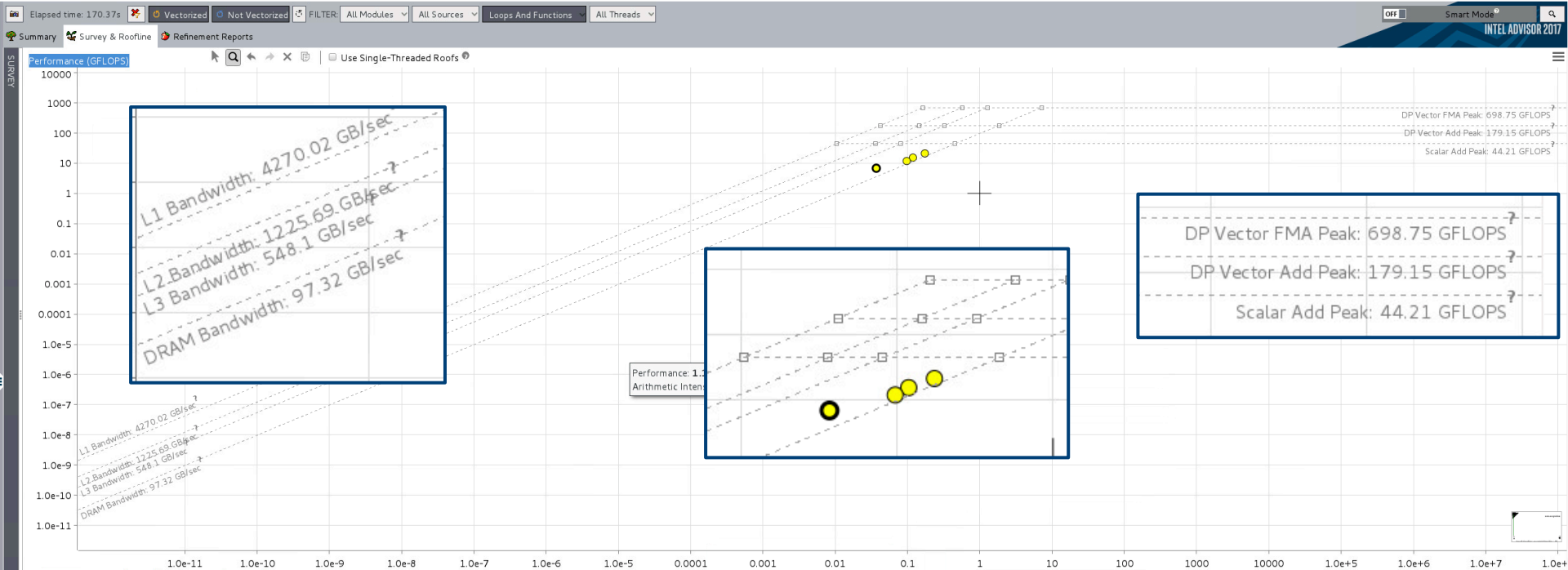
Vectorization Gain/Efficiency

Vectorized Loops Gain/Efficiency [®]	2.14x		
Program Approximate Gain [®]	1.00x		

Top time-consuming loops[®]

Loop	Self Time [®]	Total Time [®]	Trip Counts [®]
[loop in G4PolyPhiFace::InsideEdges at G4PolyPhiFace.cc:853]	30.015s	30.015s	15
[loop in G4ElasticHadrNucleusHE::HadrNucDifferCrSec at G4ElasticHadrNucleusHE.cc:892]	24.898s	44.284s	1
[loop in G4SteppingManager::DefinePhysicalStepLength at G4SteppingManager2.cc:165]	18.295s	270.912s	9
[loop in G4VoxelNavigation::VoxelLocate at G4VoxelNavigation.icc:52]	9.641s	24.850s	3
[loop in G4VCSGfaceted::Inside at G4VCSGfaceted.cc:227]	8.490s	143.584s	8

Intel Advisor – Roofline



Hotspot 1

File: MTWistEngine.cc:115 CLHEP::MTWistEngine::flat

Line	Source	Total Time	%	Loop/Function Time	%
114		7.688s			
115	double MTWistEngine::flat() {	7.688s		29.807s	
116	unsigned int y;	0.259s			
117		0.080s			
118	if(count624 >= N) {	1.549s			
119	int i;				
120					
121	for(i=0; i < NminusM; ++i) {	0.110s			
122	y = (mt[i] & 0x80000000) (mt[i+1] & 0x7fffffff);	0.060s			
123	mt[i] = mt[i+M] ^ (y >> 1) ^ ((y & 0x1) ? 0x9908b0df : 0x0);	0.100s			Shifts
124	}				
125					
126	for(; i < N-1 ; ++i) {	28.616s			
127	y = (mt[i] & 0x80000000) (mt[i+1] & 0x7fffffff);				
128	mt[i] = mt[i-NminusM] ^ (y >> 1) ^ ((y & 0x1) ? 0x9908b0df : 0x0);	0.030s			Shifts
129	}				
130		0.010s			
131	y = (mt[i] & 0x80000000) (mt[0] & 0x7fffffff);				
132	mt[i] = mt[M-1] ^ (y >> 1) ^ ((y & 0x1) ? 0x9908b0df : 0x0);	1.039s			
133					
134	count624 = 0;				
Selected (Total Time):		7.688s			

Hotspot 1

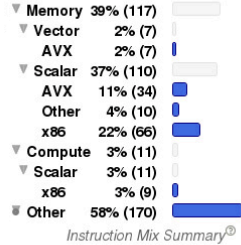
CLHEP::MTwistEngine::flat function at MTwistEngine.cc:115

f 36.324s

Function Total time

AVX; AVX2 28.556s

Instruction Set Self time



Code Optimizations

Compiler: Intel(R) C++ Intel(R) 64 Compiler for applications running on Intel(R) 64,
Version: 17.0.2.174 Build 20170213

Statistics for FLOPS And Data Transfers

GFLOPS 6.85449

AI 0.037

GFLOP 14.531

Giga Floating-point Operations Per Second
Per-loop GFLOPS = Total GFLOP / Elapsed Time. Elapsed time is the exclusive (self-time-based) wall time from the beginning to the end of loop/function execution. For single-threaded applications Elapsed time is equal to Self-Time.

AI - Arithmetic Intensity - Ratio of Floating-point Operations to L1 Transferred Bytes

Giga Floating-point Operations

Data transfers between CPU and memory sub-system (total traffic, including L1, L2, LLC and DRAM traffic)

in Giga Bytes 393.484

in Giga Bytes Per Second 185.609

Traits

Shifts

Hotspot 2

File: G4ElasticHadrNucleusHE.cc:892 G4ElasticHadrNucleusHE::HadrNucDifferCrSec

Line	Source	Total Time	%	Loop/Function Time	%	
890	Din2 = 0;					
891	DmedTot = 0;					
892	for(G4int l = 0; l<=1; l++)			44.284s		
893	{					
894	if(l == 0) BinCoeff = 1;					
895	else if(l != 0) BinCoeff = BinCoeff*(i-1+1)/1;	4.688s [Divisions; Type Conversions
896						
897	exp1 = 1/R22B+(i-1)/R12B;	0.200s [Divisions; Type Conversions
898	exp1p = exp1+R12Apd;	0.380s [
899	exp2p = exp1+R12ApdR22Ap;					
900	exp3p = exp1+R22Apd;	0.020s [
901						
902	Din2 = Din2 + N2p*BinCoeff*	1.350s [FMA
903	(C1/exp1p*std::exp(-theQ2/4/exp1p)-	4.672s [Divisions
904	C2/exp2p*std::exp(-theQ2/4/exp2p)+	6.400s [Divisions
905	C3/exp3p*std::exp(-theQ2/4/exp3p));	1.260s [Divisions
906						
907	DmedTot = DmedTot + N2p*BinCoeff*	5.929s [FMA
908	(C1/exp1p-C2/exp2p+C3/exp3p);					
909						
910	N2p = -N2p*R23dR13;					
911	} // 1					

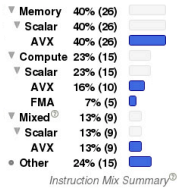


Hotspot 2

Loop in G4ElasticHadrNucleusHE::HadrNucDifferCrSec at G4ElasticHadrNucleusHE.cc:892

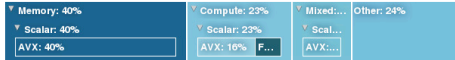
 **44.284s**
Scalar Total time

AVX 24.898s
Instruction Set Self time



Instruction Mix

Memory: 26 Compute: 15 Mixed: 9 Other: 15 Number of Vector Registers: 16



Trip Counts

Maximum: 40
Average: 1
Minimum: 1
Call Count: 63118400
Iteration Duration: < 0.001s

Code Optimizations

Compiler: Intel(R) C++ Intel(R) 64 Compiler for applications running on Intel(R) 64,
Version: 17.0.2.174 Build 20170213

Statistics for FLOPS And Data Transfers

GFLOPS 11.8651

AI 0.097

GFLOP 20.640

FLOP Per Iteration 29

Giga Floating-point Operations Per Second
Per-loop GFLOPS = Total GFLOP / Elapsed Time. Elapsed time is the exclusive (self-time-based) wall time from the beginning to the end of loop/function execution. For single-threaded applications Elapsed time is equal to Self-Time.

AI - Arithmetic Intensity - Ratio of Floating-point Operations to L1 Transferred Bytes

Giga Floating-point Operations

Floating-point Operations Per Loop Iteration

Data transfers between CPU and memory sub-system (total traffic, including L1, L2, LLC and DRAM traffic)

in Giga Bytes 212.110

in Giga Bytes Per Second 121.934

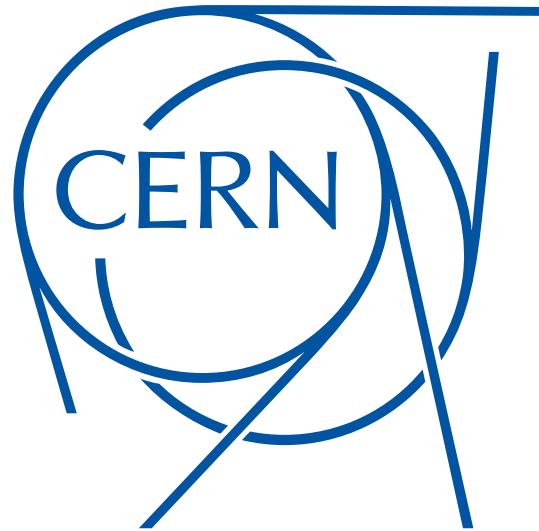
in Bytes Per Loop Iteration 298.023

Traits

Divisions
Type Conversions
FMA

Conclusions

- Good compilers and tools are available
- Architecture's peak performance
- Compute kernels
- Algorithm complexity
- Compiler readable code
- Good loops – Bounded, Linear and Minimal Intra Dependencies
- Iterative optimization – Design, Analyze, Optimize, Repeat



Nicholas Terranova

Self-Introduction

- 31 years old
- Nuclear Engineer
- From Bologna (Italy)
- Research fellow at CNAF-INFN of Bologna, since 16-Oct-2017



Nicholas Terranova (cont'd)

- Master degree in Nuclear Engineering in 2012;
- Master thesis at the University of Arizona, title: *Numerical benchmarks for the multi-group diffusion equation*;
- **PhD in Reactor Physics** in 2016;
- 2,5 year at the CEA (**French Atomic Energy Commission**) of Cadarache (south of France);
- Title: *Covariance matrix generation for nuclear data of interest to the reactivity loss uncertainty estimation of the Jules Horowitz Material Testing Reactor.*



- In the last year: **post-doc fellow at CEA-Saclay** (close to Paris);
- Development of reactivity perturbation and eigenvalue sensitivity features in **Tripoli-4**;
- **Tripoli-4 is a Monte Carlo transport code** for instrumentation, radiation protection and reactor criticality calculations developed by the Reactor Physics and Applied Mathematics Section of the CEA-Saclay.

**Thank you for your
attention**
nicholas.terranoa@cnaif.infn.it



Andres Tiko



INFN Padova

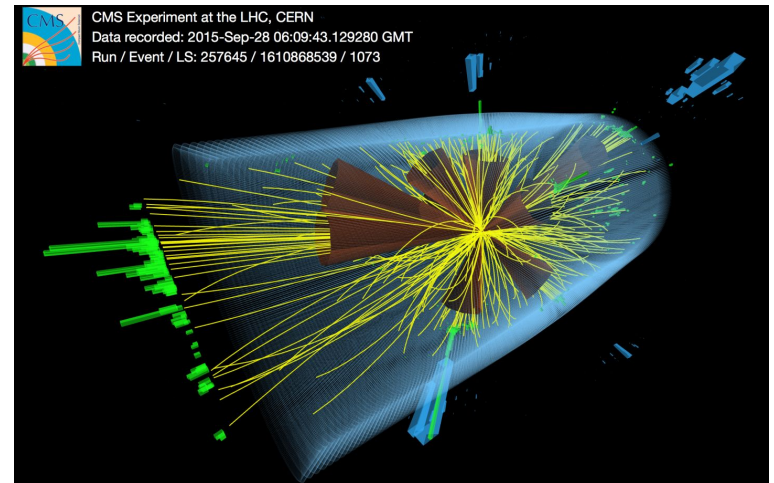
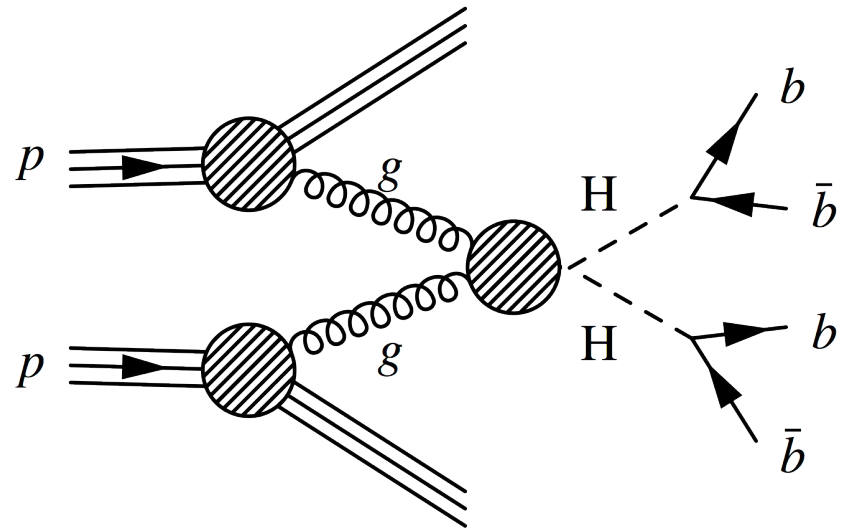
Experimental particle
physics postdoc

Member of CMS
— experiment

Current work

Trying to measure the production of Higgs boson pairs decaying to b-quarks

Rare process, but gives important insight into Higgs boson properties and further understanding of the Standard Model



Challenges

- Dealing with huge amounts of data - and increasing all the time
- Need to be able to process it efficiently within the framework of LHC computing grid - large, but not unlimited resources
- Overwhelming amounts of background processes mimicking the signal signature - need to distinguish between the two

Future interests

So far, high energy physics has mostly focused on the knowledge of physical processes to model and analyze data

New approaches are being introduced - using deep learning, we can treat the tracks of a collision in a detector as an image and use it directly for separating background and signal

Signal



Background



*Average fractional
polarization of
extragalactic radio
sources at
Planck frequencies*



© Tiziana Trombetti

ESC17 23-28/10/2017

Work in collaboration with of Carlo Burigana & Gianfranco de Zotti

The detection of CMB primordial B-mode polarization is the current challenge in cosmology & particle physics ... They are crucial since they carry a clean signature of primordial inflation

This signal is very faint \rightsquigarrow provides a measure of the energy scale of inflation \rightsquigarrow

its amplitude is related to the tensor to scalar ratio $r = T/S$ & is generated by weak tensor perturbations ($r < 0.07$)

But

Even in the cleanest 70% sky, Galactic & extragalactic POLARIZED FOREGROUND emissions DOMINATE over B-MODES & instrumental noise of MOST SENSITIVE forthcoming or proposed CMB EXPERIMENTS \rightsquigarrow CORE, CMB Probe, PIXIE, LiteBIRD & CMB-S4

Foreground cleaning must reach at least 99.9% level @ $l \simeq 10$...

99% level @ $l \simeq 100$... 90% level @ $l \simeq 1000$

B-mode APS due to gravitational lensing exceeds primordial B- for $l \geq 10$ if $r \approx 10^{-2}$ \rightsquigarrow
primordial B- requires very accurate control of lensing effects
(contaminated by fluctuations of unresolved extragalactic sources...)

Crucial \rightsquigarrow precise understanding of extragalactic sources polarization properties

Different observations and analysis of radio
sources polarimetric properties evidence
broad variety of spectral shapes

*The key is to characterize source contamination
@ ν range [60-120] GHz ... where Planck data have
min in T brightness spectra of diffuse polarized foregrounds*

Although # extraG pol sources detected by Planck quite limited ...
<P fraction> of fainter sources can be estimate \rightarrow stacking techniques
coadding the P signal from many objects detected in total intensity
(not in polarization) ... $\hat{=}$ signal-to-noise (S/N) ratio!

Low-res Planck data is complicated \rightarrow each resolution element contains a source &
other polarized signals: CMB itself & diffuse polarized emissions from the Galaxy (non Gaussian!)

BIAS!!!!

MeTHOD carry out a new investigation
adopting an independent simpler & analytical approach...

Signals measurements in
map @ given source catalogue...

IDA \Rightarrow intensity distribution
analysis

The distribution of signals is compared with
random positions away from sources \rightarrow control fields

The mean polarized flux density of sources is then estimated as

$$\langle P \rangle = \left(\langle P_{sources}^2 \rangle - \langle P_{CFields}^2 \rangle \right)^{1/2}$$

to statistically remove the noise

contributions to P, to the CMB & to polarized Galactic emissions ...

the method directly corrects for the 'noise bias' without the need of simulations!!!



UNIVERSITÀ
DEGLI STUDI
DI TORINO



INDIGO - DataCloud
Better Software for Better Science

Batch System as a Service on HPC resources

Speaker: Vallero Sara - INFN Torino

ESC17 School - Bertinoro (FC) October 23-28 2017

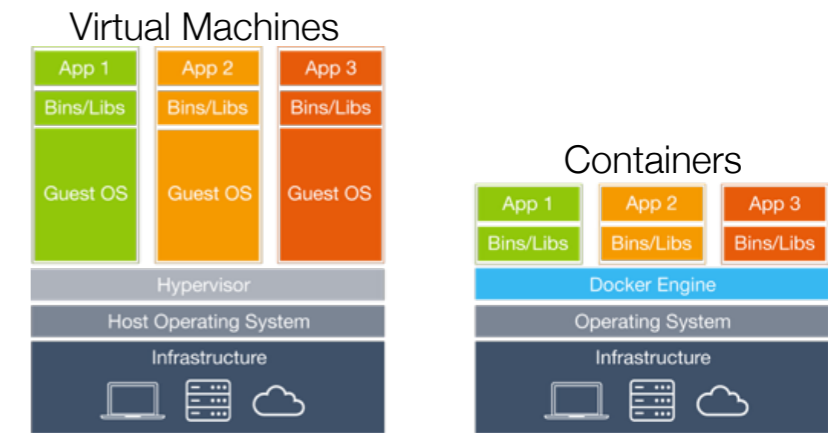
Cloud-like cluster management

OCCAM cluster @ C3S (Torino)
<https://c3s.unito.it>

➔ light-weight virtualization (Linux containers)

According to GARTNER:

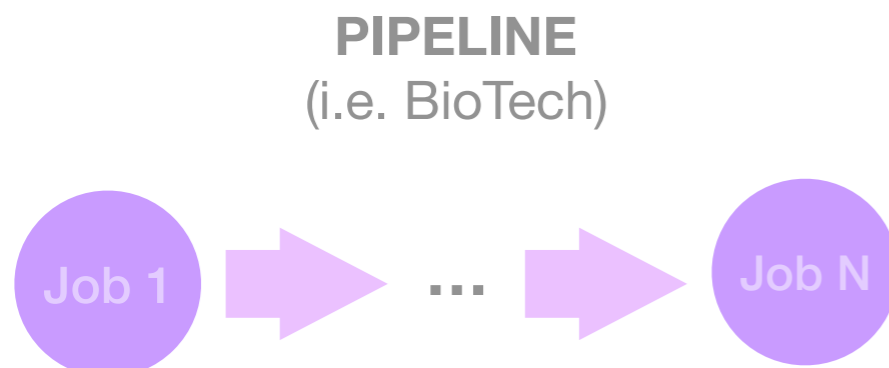
Cloud Computing is a style of computing in which **scalable and elastic** IT-enabled capabilities are delivered **as a service** using Internet technologies.



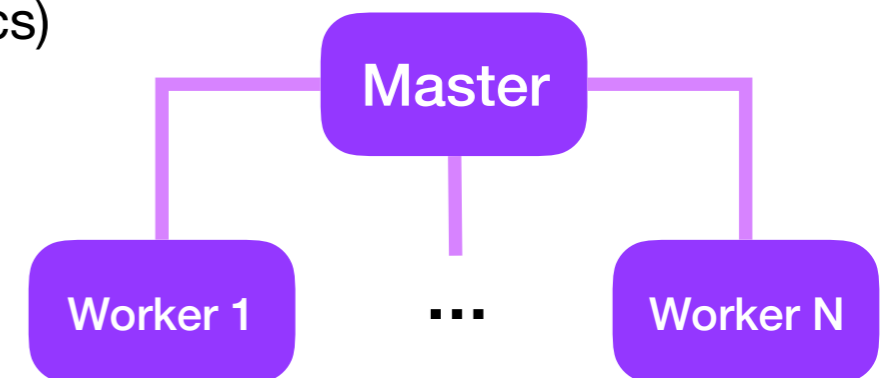
From VMs to Containers:

- **package**, ship and run distributed application components with guaranteed **platform parity** across different environments
- *democratizing* virtualization by providing it to developers in a usable, **application-focused** form
- full virtualization not suited for HPC applications

➔ enable different applications (computing models) to run concurrently



BATCH SYSTEM
(i.e. Physics)



➔ dynamic re-allocation of resources, autoscaling

Batch System as a Service

Automatically deploy a batch system cluster in highly-available and scalable configurations.

Schedule tasks on distributed resources



Apache
MESOS

A DISTRIBUTED KERNEL

1) Ad hoc Mesos framework

(<https://github.com/svallero/HTMframe>)



- scheduler: implements custom policies on Mesos resource offers
 - instantiates *static roles* in ordered sequence: master, submitter and some executors
 - re-instantiate static roles on failure
- health-checks on static roles
- **auto-scaling according to HTCondor metrics:**
 - scale-out: in case of idle jobs (submitter publishes queue and nodes status)
 - scale-in: idle nodes publish unhealthy state
- HTTP API
- GUI (coming soon)

THE SCHEDULERS

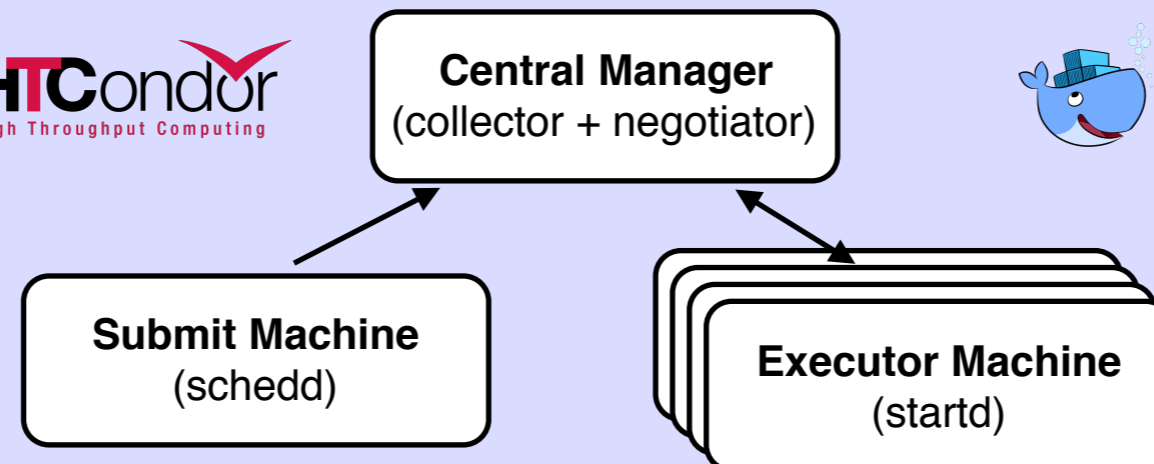
What we used so far...

2) Marathon framework



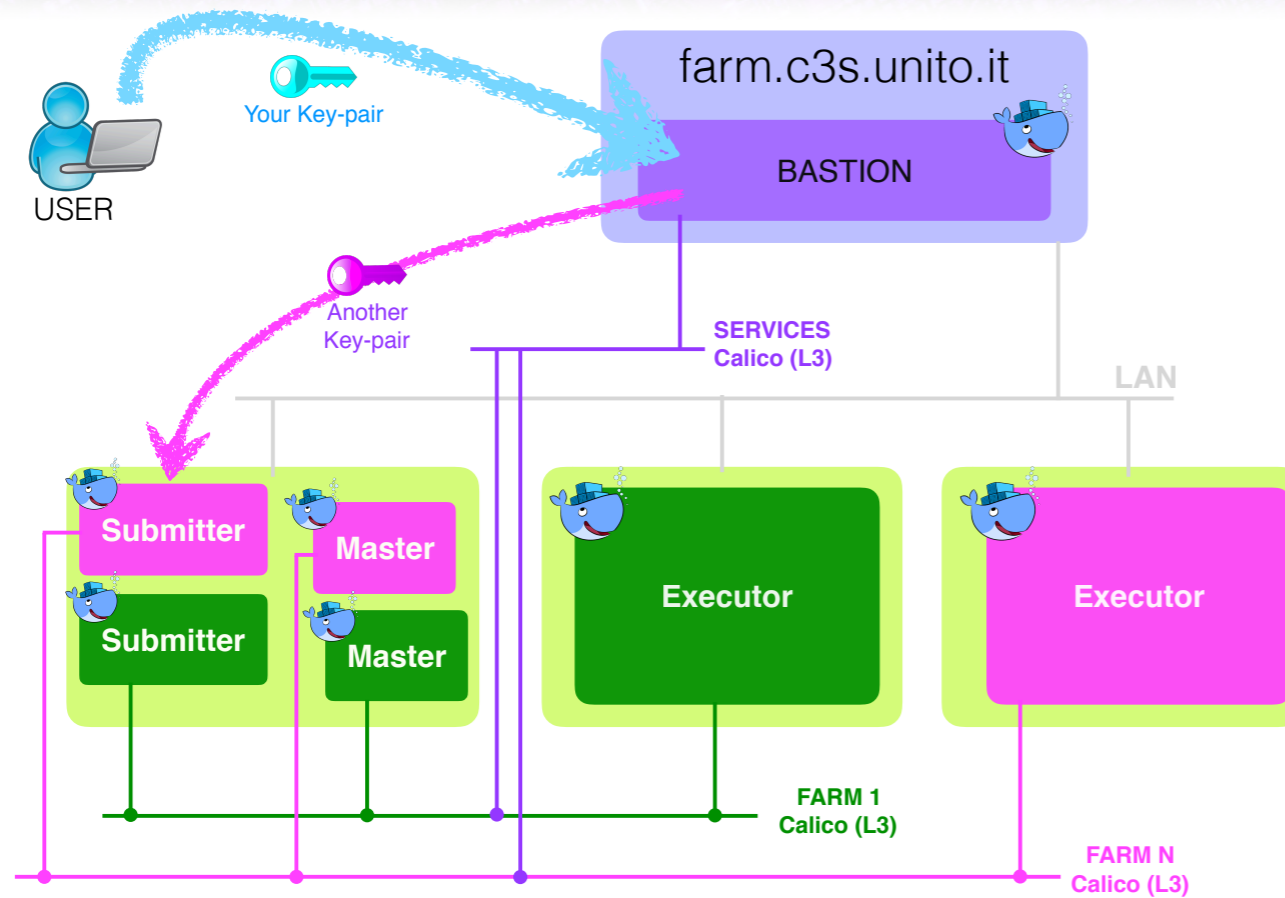
- containers are deployed as Long Running services
- health checks and failover
- no application specific auto-scaling

All farm components are packaged in separate Docker containers



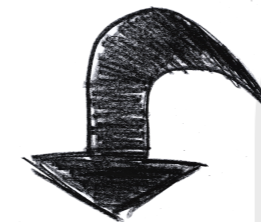
THE APPLICATION

Application isolation



Calico functionalities:

- **Orchestrator Plugin**: to manage the Calico network as with the orchestrator (Docker) tools
- **Felix agent**: programs routing rules and ACLs into the Linux kernel (FIB and iptables)
- **etcd**: communication between components and datastore
- **BGP client**: distributes the routing tables around the datacenter



PoC of geographical deployment

- 3 autonomous systems: Torino, Bologna and Bari
- VPN connection among the 3 sites (IPoIP)
- each site runs an instance of: configuration store (etcd and Zookeeper), Calico, Mesos and Marathon
- one instance of each of the above services is elected as leader (fault tolerance)

Mesos Frameworks Slaves Offers

Master c21131d6-7f80-40dc-a9e6-c159c604313f

Cluster: IndigoCluster
 Server: 90.147.102.106:5050
 Version: 0.28.0
 Built: 7 months ago by root
 Started: a week ago
 Elected: 7 days ago

LOG

Slaves
 Activated 4
 Deactivated 0

Active Tasks

ID	Name	State	Started	Host
htcondor-submitter-debian.91889e13-8a13-11e6-828c-567d9b6063b0	htcondor-submitter-debian	STAGING		90.147.169.190
htcondor-submitter-debian.8e88f2a2-8a13-11e6-828c-567d9b6063b0	htcondor-submitter-debian	STAGING		90.147.169.190
htcondor-e-workload-to-1.6af12bc4-864e-11e6-828c-567d9b6063b0	htcondor-e-workload-to-1	RUNNING	5 days ago	193.205.66.172
htcondor-e-workload-to-1.2462310a-864c-11e6-828c-567d9b6063b0	htcondor-e-workload-to-1	RUNNING	5 days ago	193.205.66.172
htcondor-e-workload-cnaf.3cec9c33-864a-11e6-828c-567d9b6063b0	htcondor-e-workload-cnaf	RUNNING	5 days ago	131.154.98.2
htcondor-s-workload.f1e88231-8649-11e6-828c-567d9b6063b0	htcondor-s-workload	RUNNING	5 days ago	131.154.98.2
htcondor-m-workload.f7105a3f-8648-11e6-828c-567d9b6063b0	htcondor-m-workload	RUNNING	5 days ago	131.154.98.2
htcondor-master-debian.5284d039-7a65-11e6-aef0-aab8a9362413	htcondor-master-debian	RUNNING	3 weeks ago	131.154.98.2

MPI jobs and InfiniBand

Executor role configuration (Marathon)

```

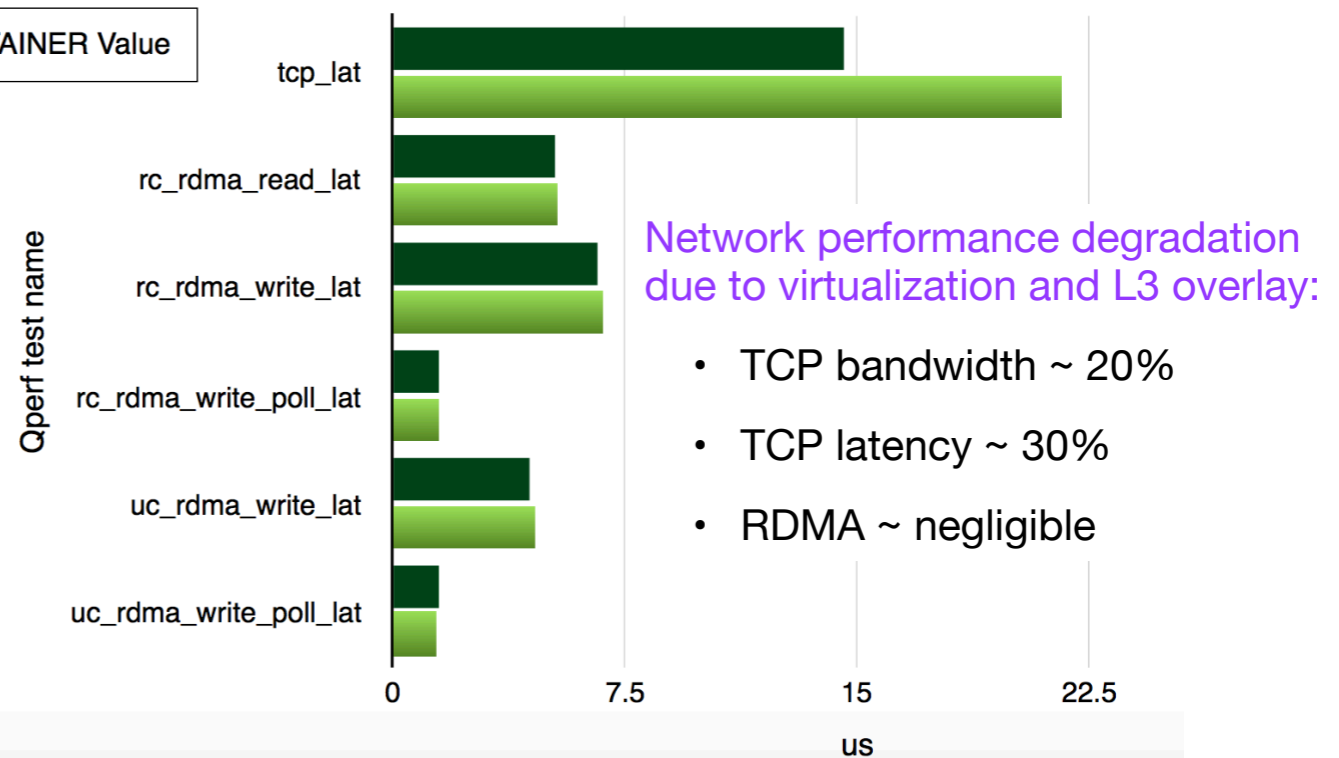
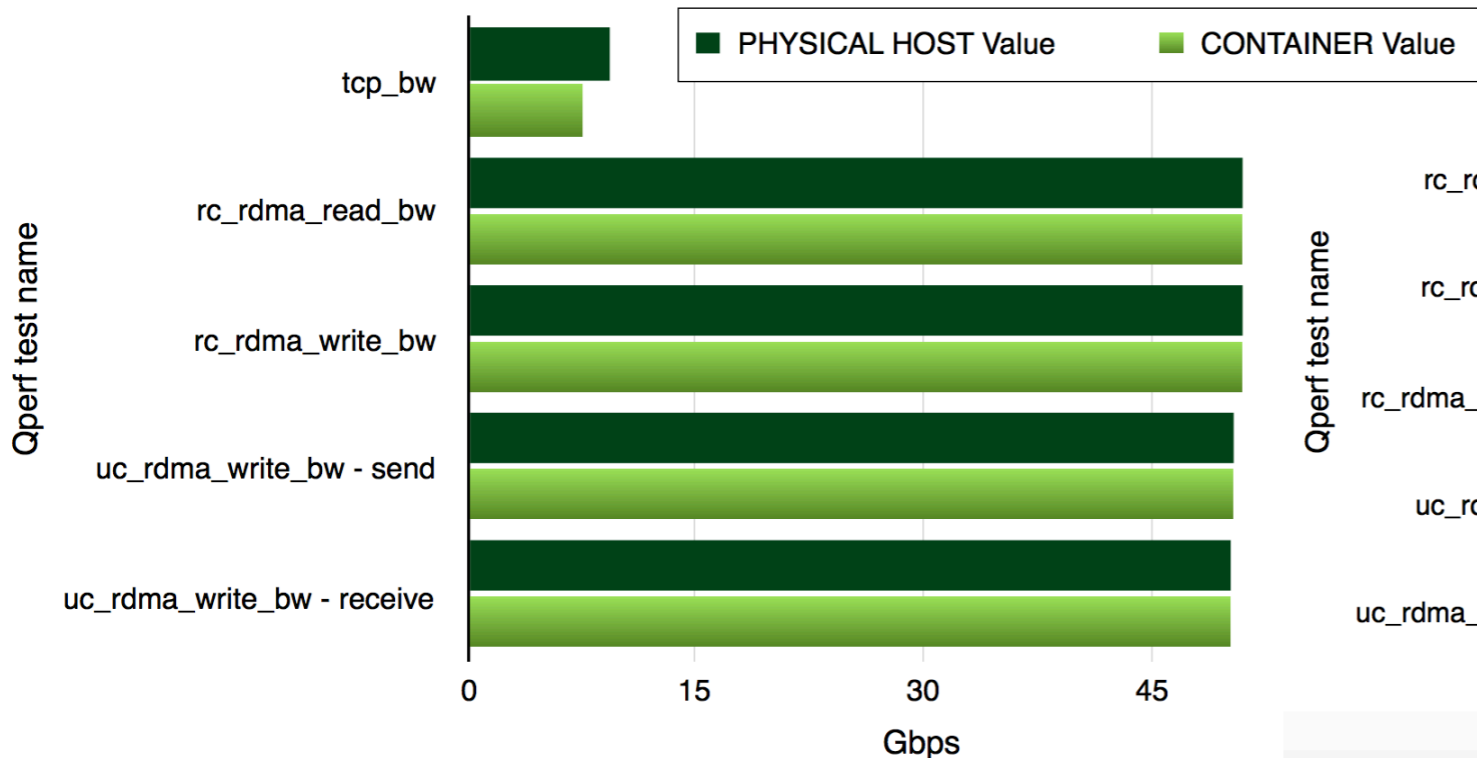
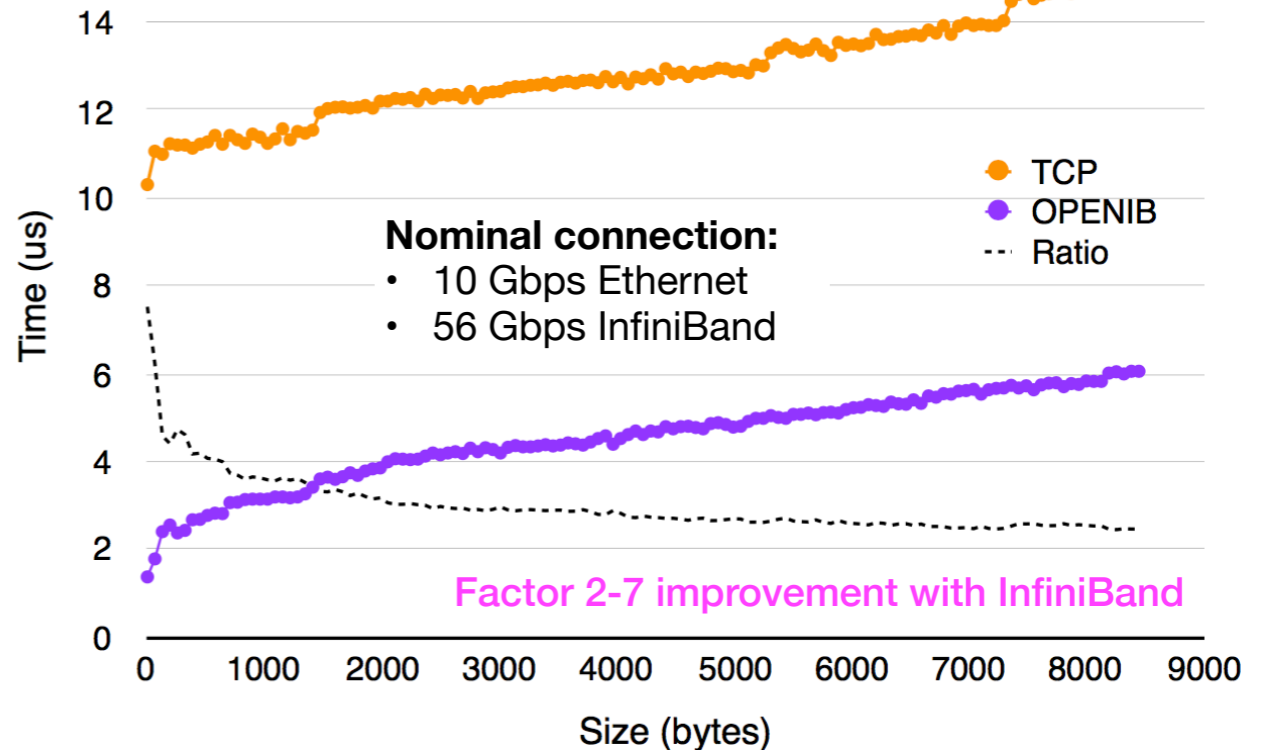
{
  "id": "/omkaar-executor-mpi",
  "cpus": 1,
  "mem": 2000,
  "disk": 0,
  "instances": 1,
  "container": {
    "type": "DOCKER",
    "volumes": [],
    "docker": {
      "image": "gitlab.c3s.unito.it:5000/htadmin/createdockers",
      "parameters": [
        {
          "key": "net",
          "value": "omkaarnet"
        },
        {
          "key": "device",
          "value": "/dev/infiniband:/dev/infiniband"
        },
        {
          "key": "ulimit",
          "value": "memlock=-1:-1"
        }
      ]
    }
  }
}

```

Calico net

InfiniBand

Roundtrip blocking communication between 2 farm nodes (mpptest suite on OpenMPI)



Hi!

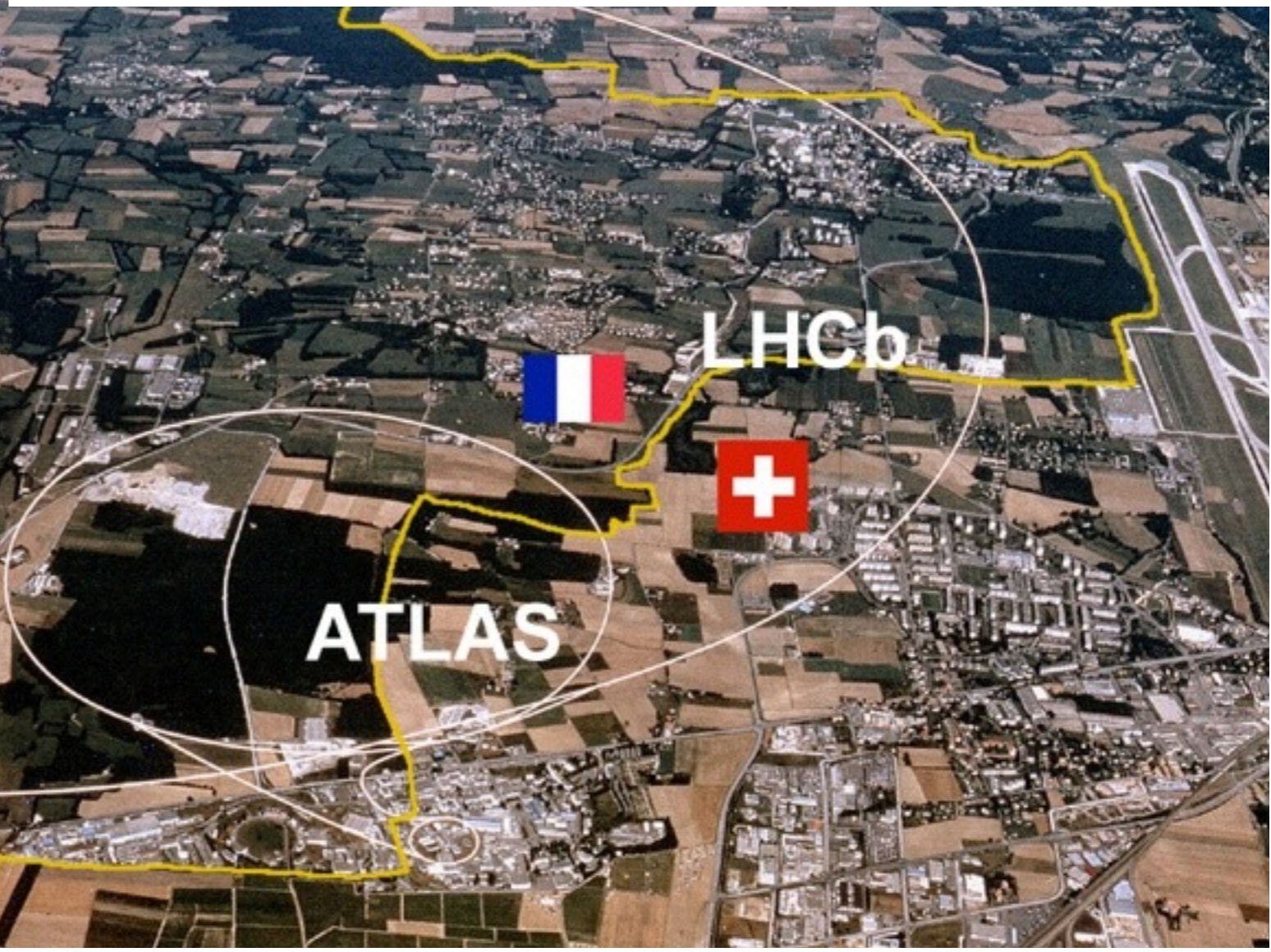
Mauro Verzetti



The LHC



- proton-proton collider
- length: 27 km
- \sqrt{s} : 7 TeV (2011), 8 TeV (2012), 13 TeV (2015)

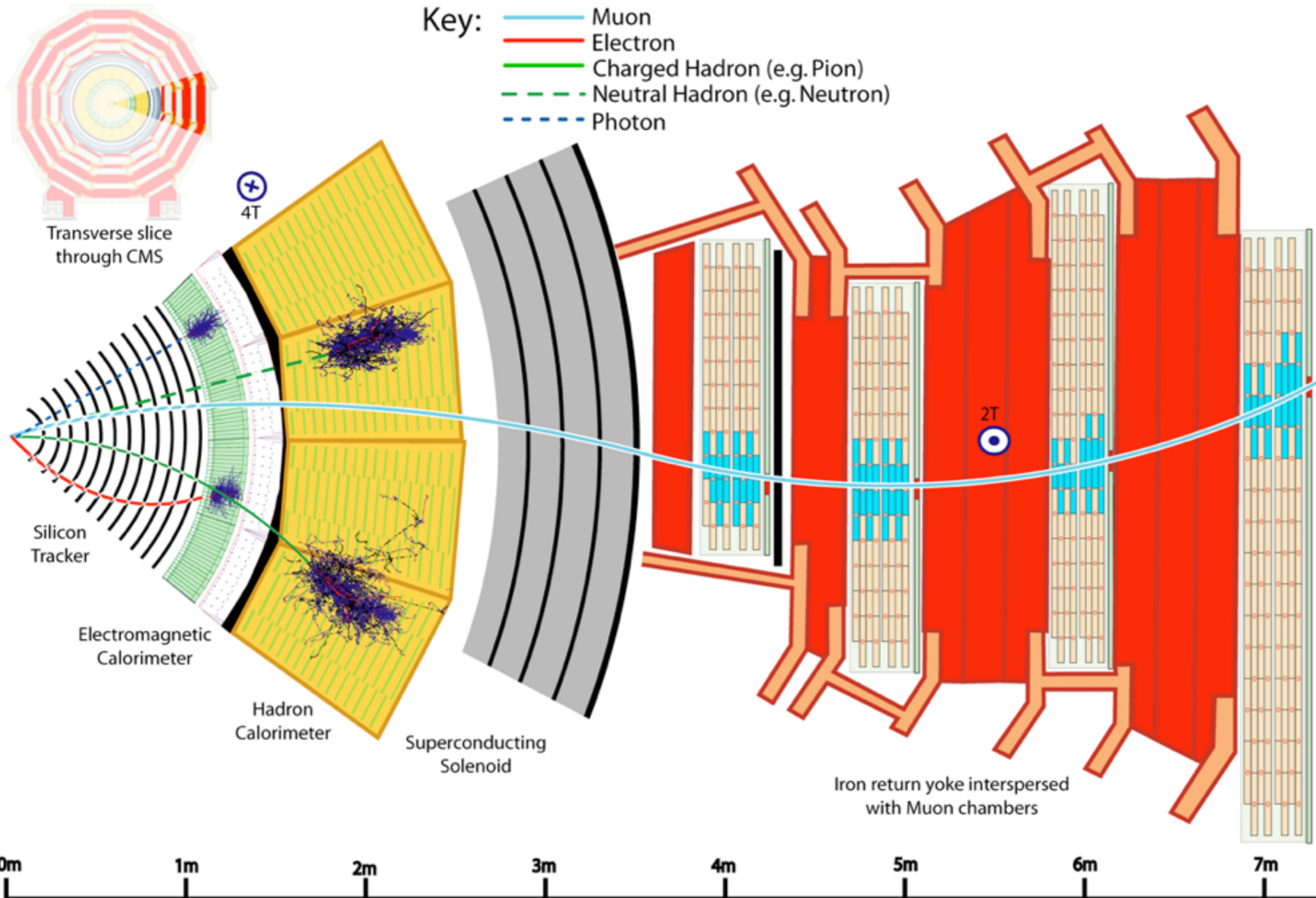


ALICE

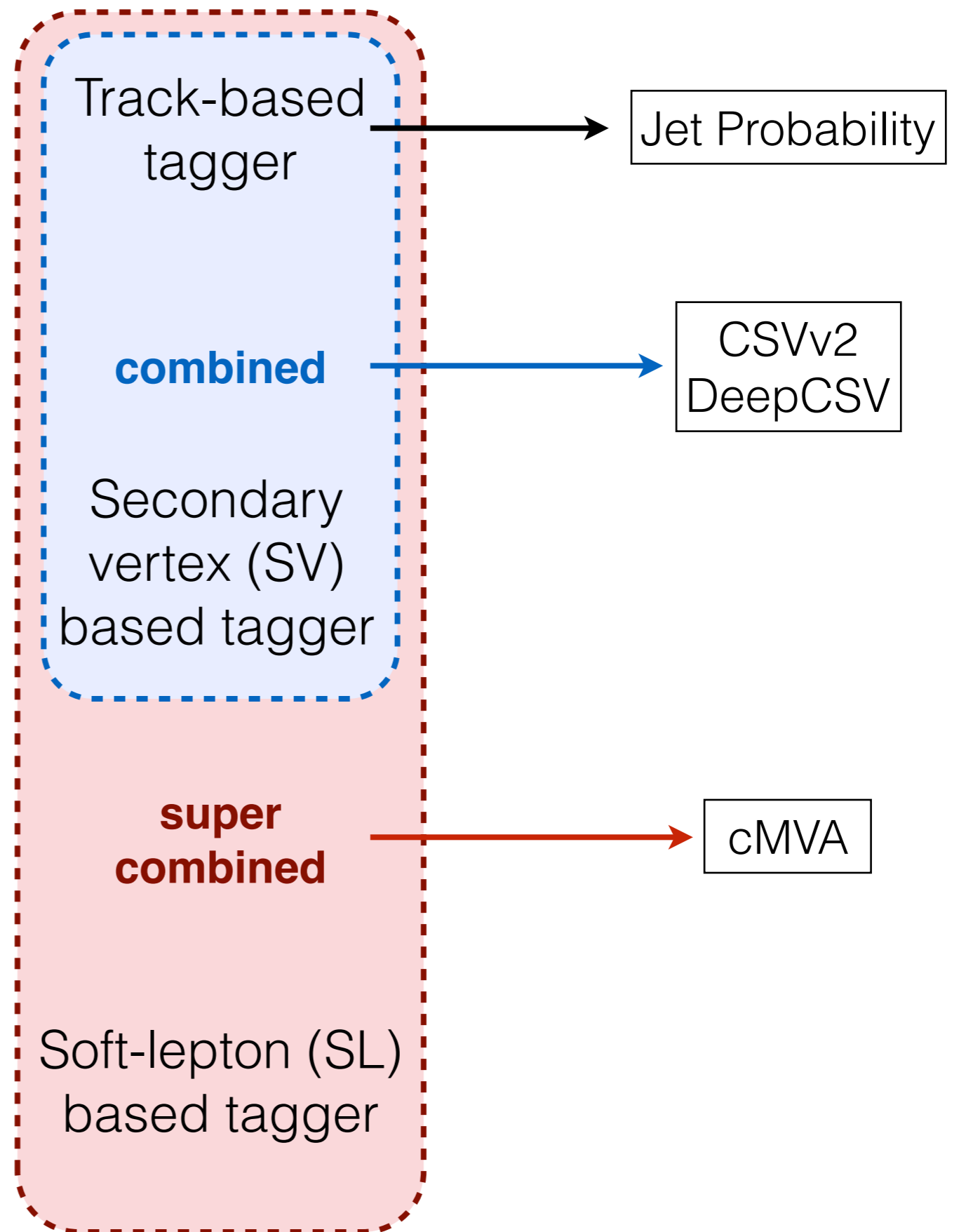
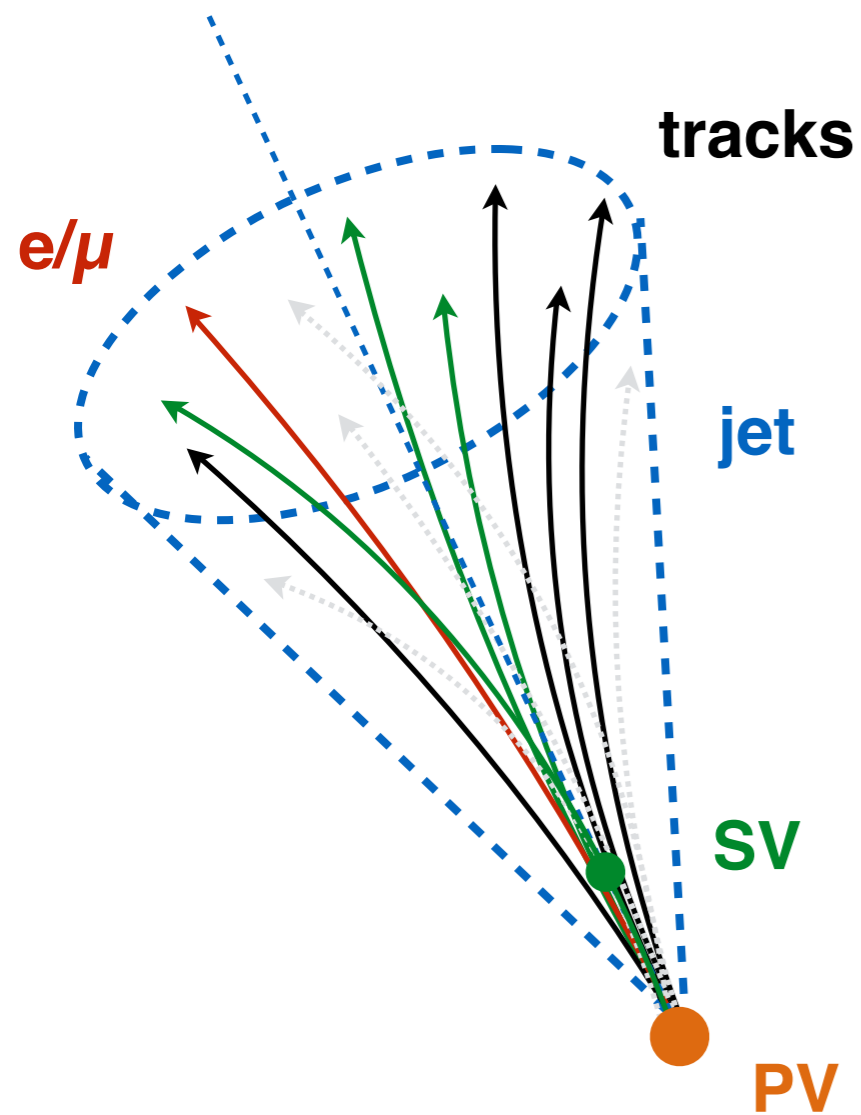
ATLAS

LHCb

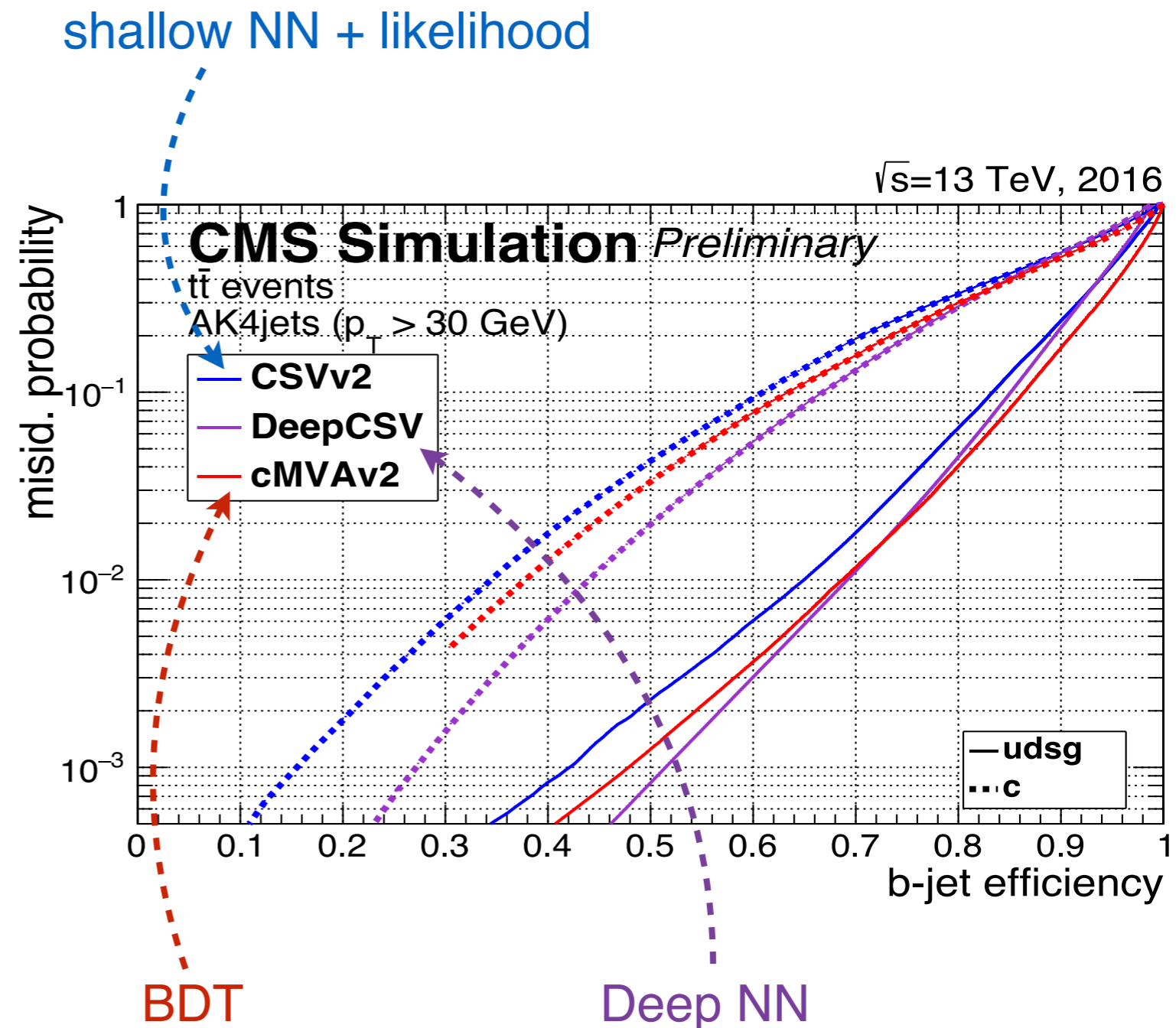
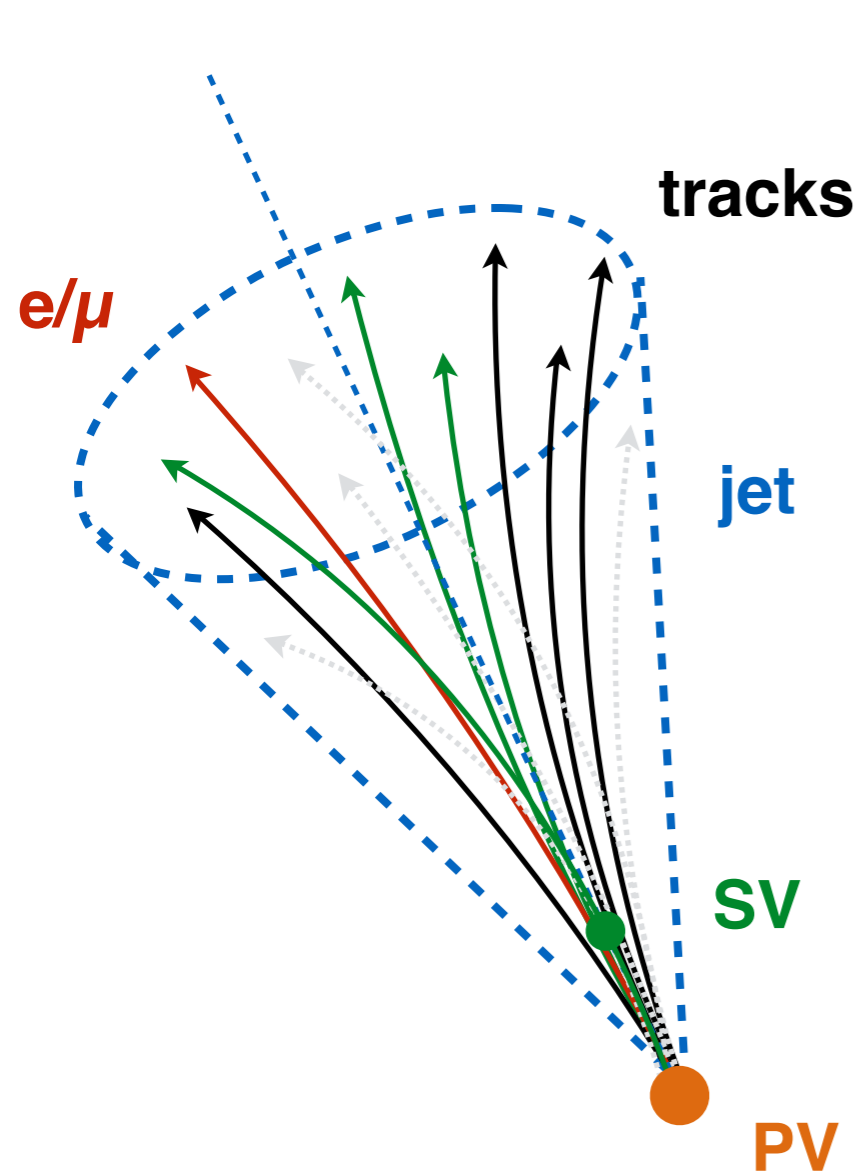
The CMS Experiment



Heavy flavor jets



Heavy flavor jets



DeepNN shows best performance

ESC17: Architectures, tools and methodologies for
developing efficient large scale scientific computing
applications

Gravitational-wave luminosity of Binary Neutron Star Mergers

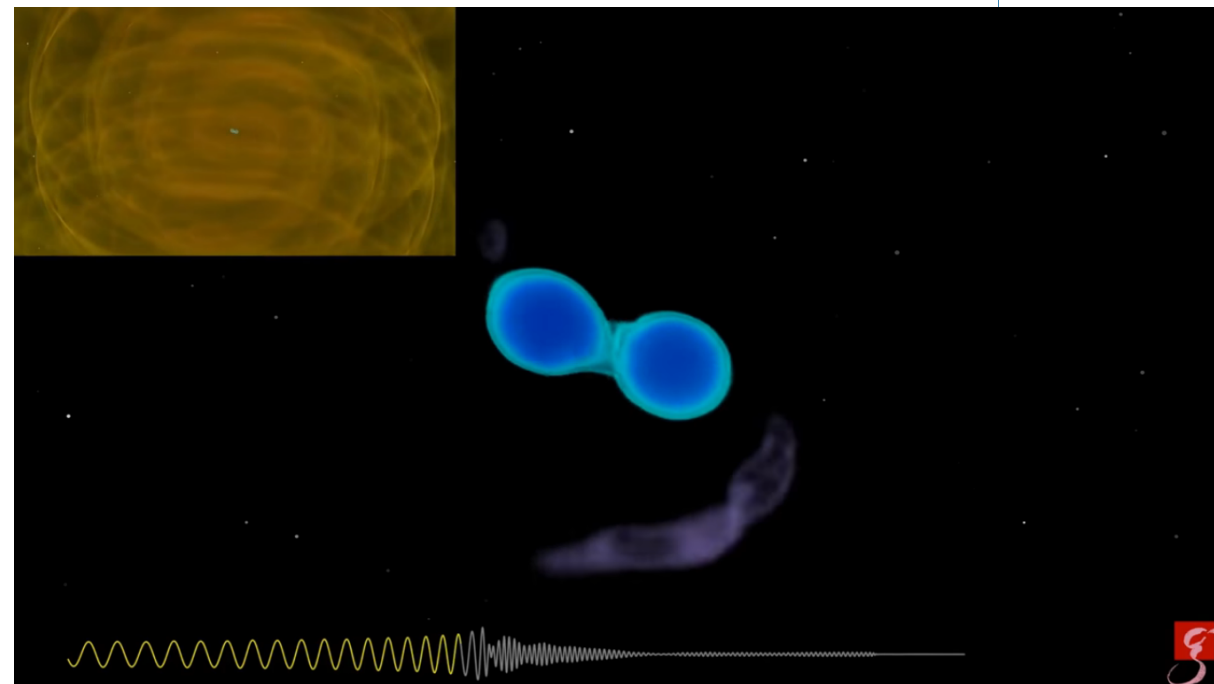
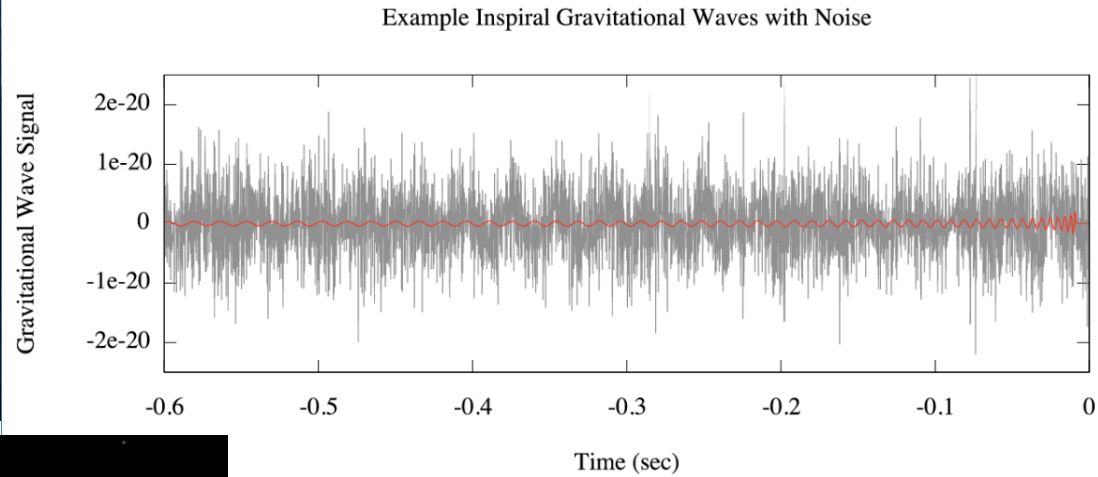
Numerical relativity for gravitational-wave modeling

Francesco Zappa



Ce.U.B. Bertinoro(FC), 23-28.10.2017

Gravitational-wave observations require waveform models



Simulations in general relativity (Numerical relativity)

GW170817, simulation of the merger (T. Dietrich)

Gravitational waves

Ricci curvature,
proportional to second
derivatives of $g_{\mu\nu}$

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

Four-tensor metric,
the unknown.

Stress-energy tensor
(source term).

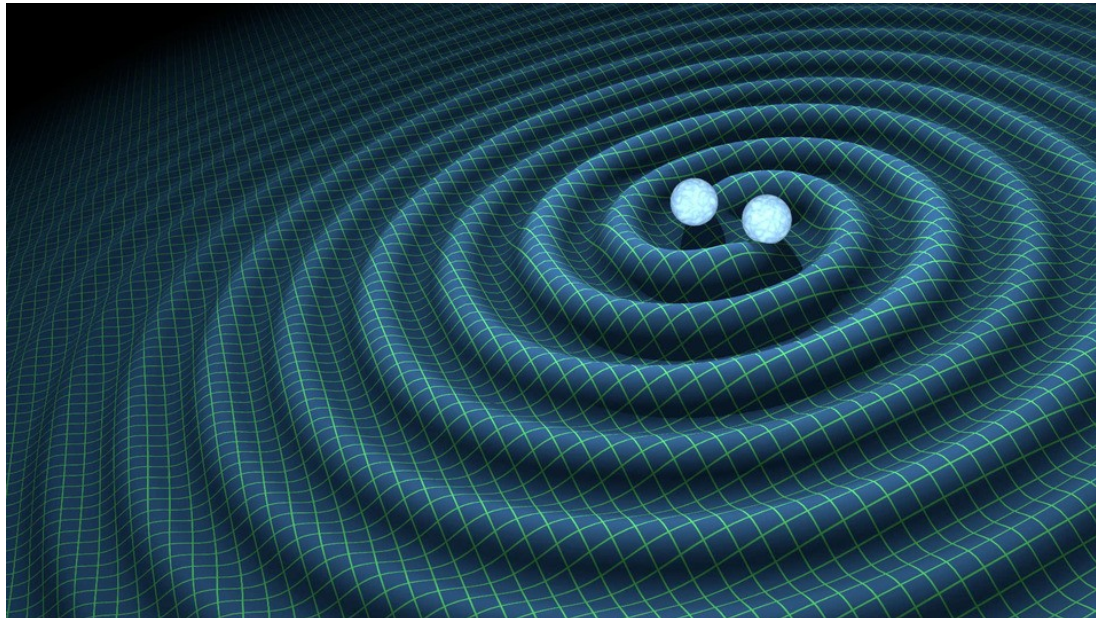
Ansatz GW:

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

Small perturbation over the
flat background

Linearized Einstein equations are tensor wave equations:

$$\partial_t^2 h_{ij} = \eta^{kl} \partial_k \partial_l h_{ij}$$



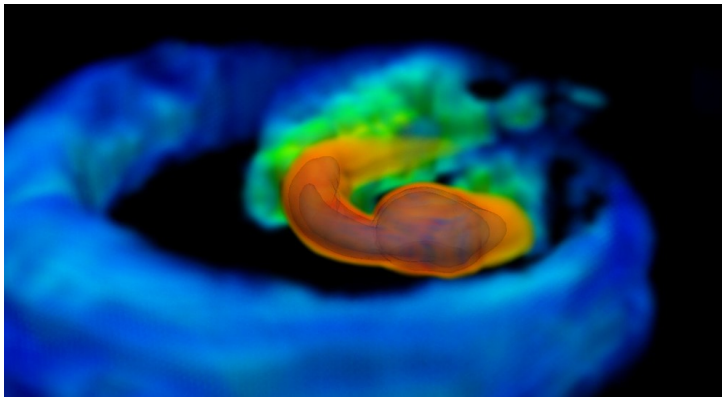
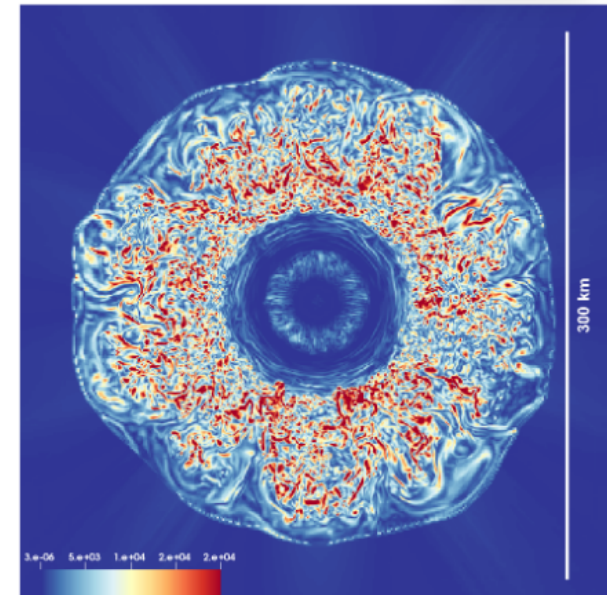
Numerical relativity

Challenges:

- GR formulation and Cauchy problem
- Coordinates
- GR hydrodynamics (Neutron star mergers, supernova bursts..)
- Strong and dynamical gravitational field
- Multi-physics and multiscales

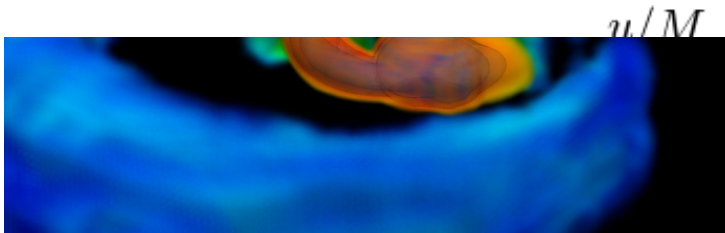
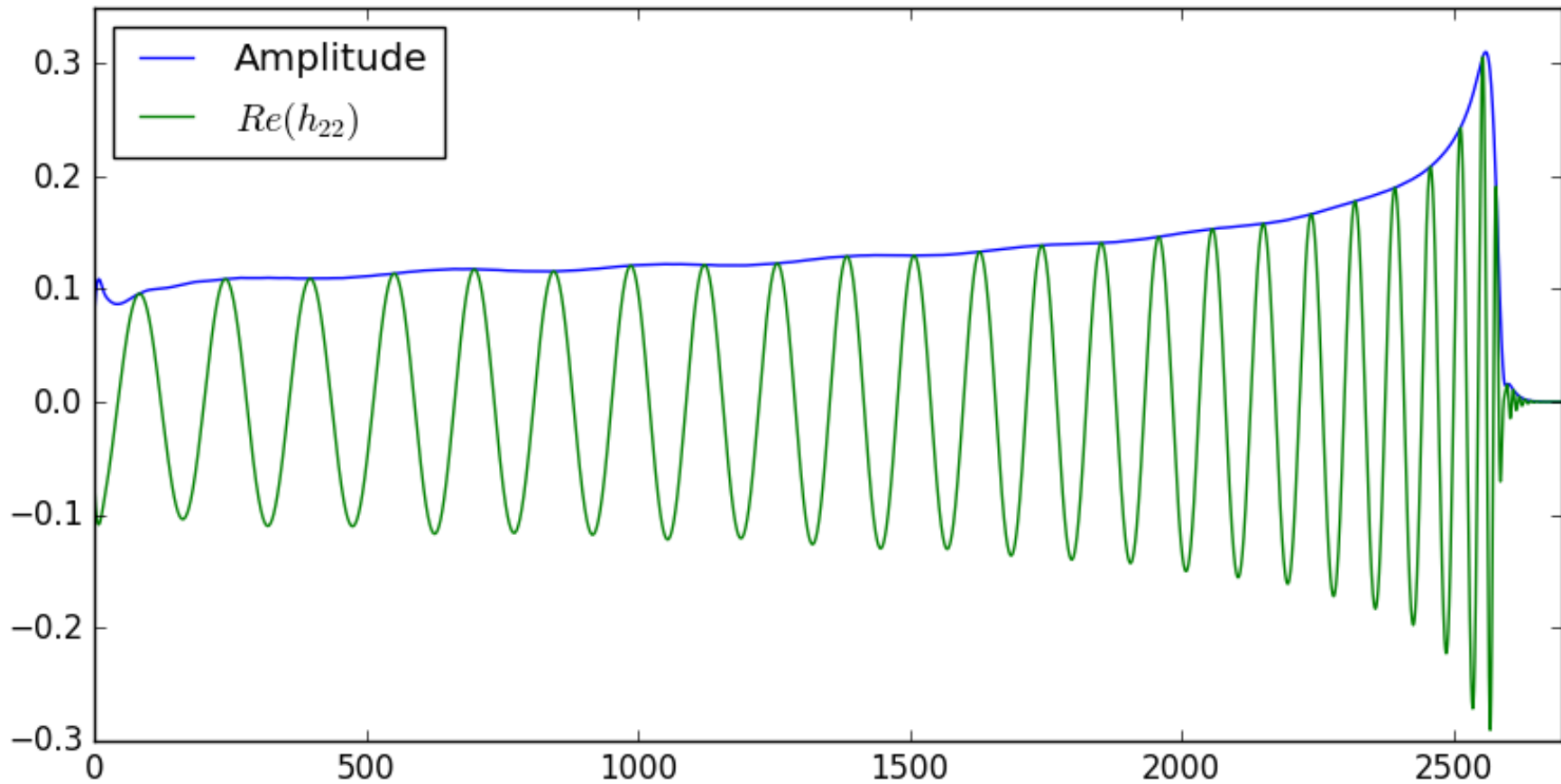
Application examples:

- Compact binary mergers [S. Bernuzzi 2012, D. Radice 2015]
- Supernova core collapse [Winteler 2012, D. Radice 2016]
- Gravitational wave modeling [S. Bernuzzi 2016, T. Dietrich 2017]

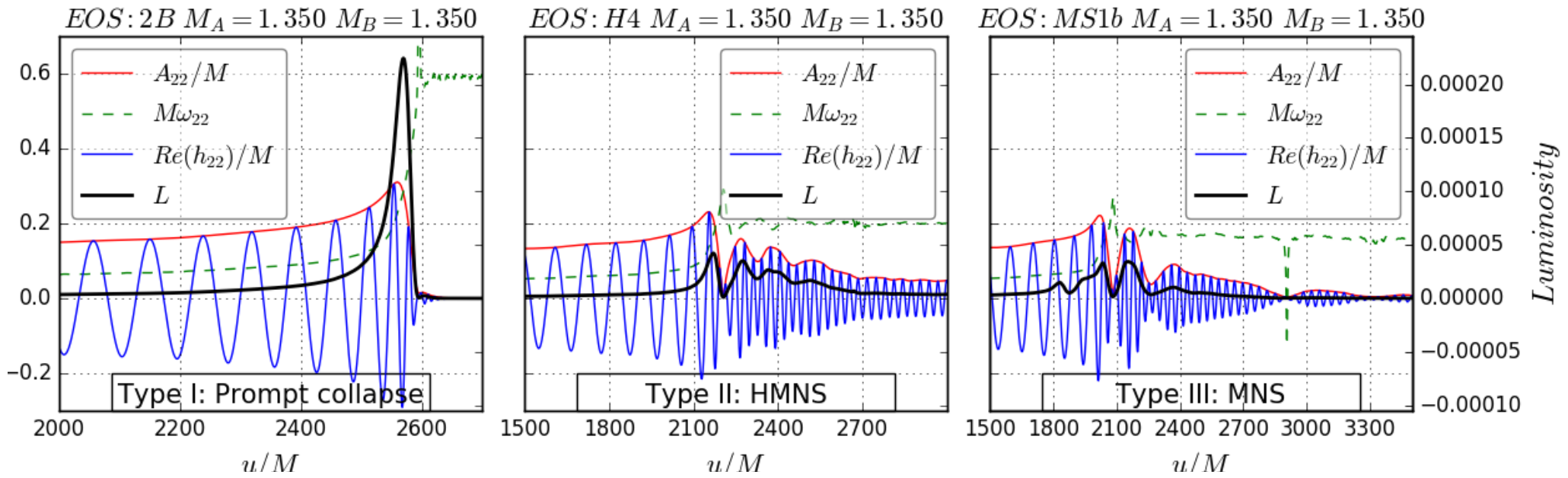


Numerical relativity

Challenges:



My work: luminosity of GWs



Tidal effects are quantified by one parameter κ_2^T that strongly affects the merger's dynamic:

$$L_{\text{peak}} / \kappa_2^T \text{ for many simulations} \rightarrow \text{fit}$$

My work: luminosity of GWs

