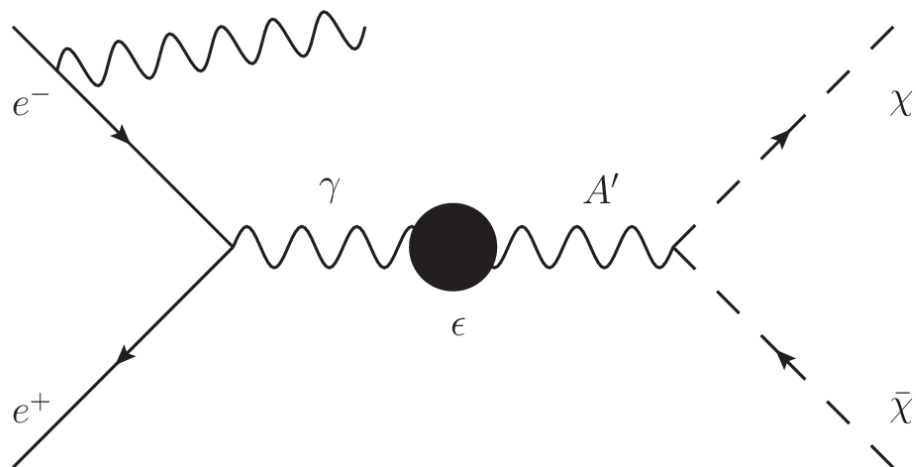


# Dark photon $\rightarrow$ Invisibles & Low multiplicity @ Phase 2



**Giacomo De Pietro**

*Università di Roma Tre*    *INFN Roma Tre*



*8th Belle II Italian Meeting  
20-21 November 2017 @ Pisa*

# Dark photon $\rightarrow$ Invisibles

# Theory

Dark photon = bosonic mediator of a “dark interaction”  
between dark matter (and SM) particles

**A dark photon could explain the muonic g-2 anomaly!**

Minimal model:

$$\mathcal{L}_{A'} = -\frac{1}{4}F'^{\mu\nu}F'_{\mu\nu} + \boxed{\frac{1}{2} \frac{\epsilon}{\cos \theta_W} B^{\mu\nu} F'_{\mu\nu}} - \frac{1}{2}m_{A'}^2 A'^{\mu} A'_{\mu}$$

at low energies:

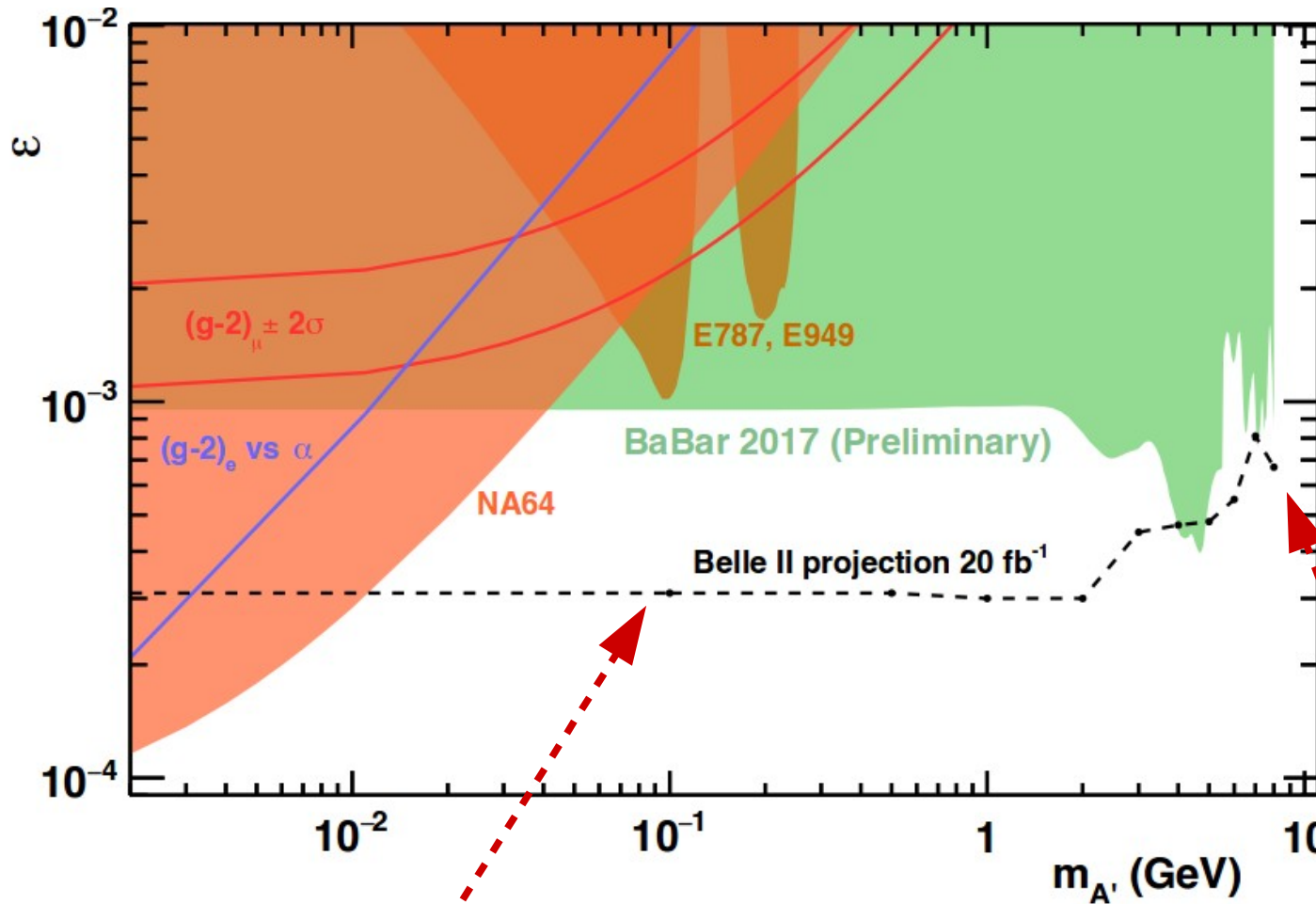
$$\boxed{\mathcal{L}_{\text{kin.mix.}} = \frac{1}{2}\epsilon F^{\mu\nu} F'_{\mu\nu}}$$

**Free parameters (to be measured):**

$\epsilon$  (strength of the mixing)

$m_A$  (mass of the dark photon)

# Belle II perspective



20 fb<sup>-1</sup>  
@ Phase 2

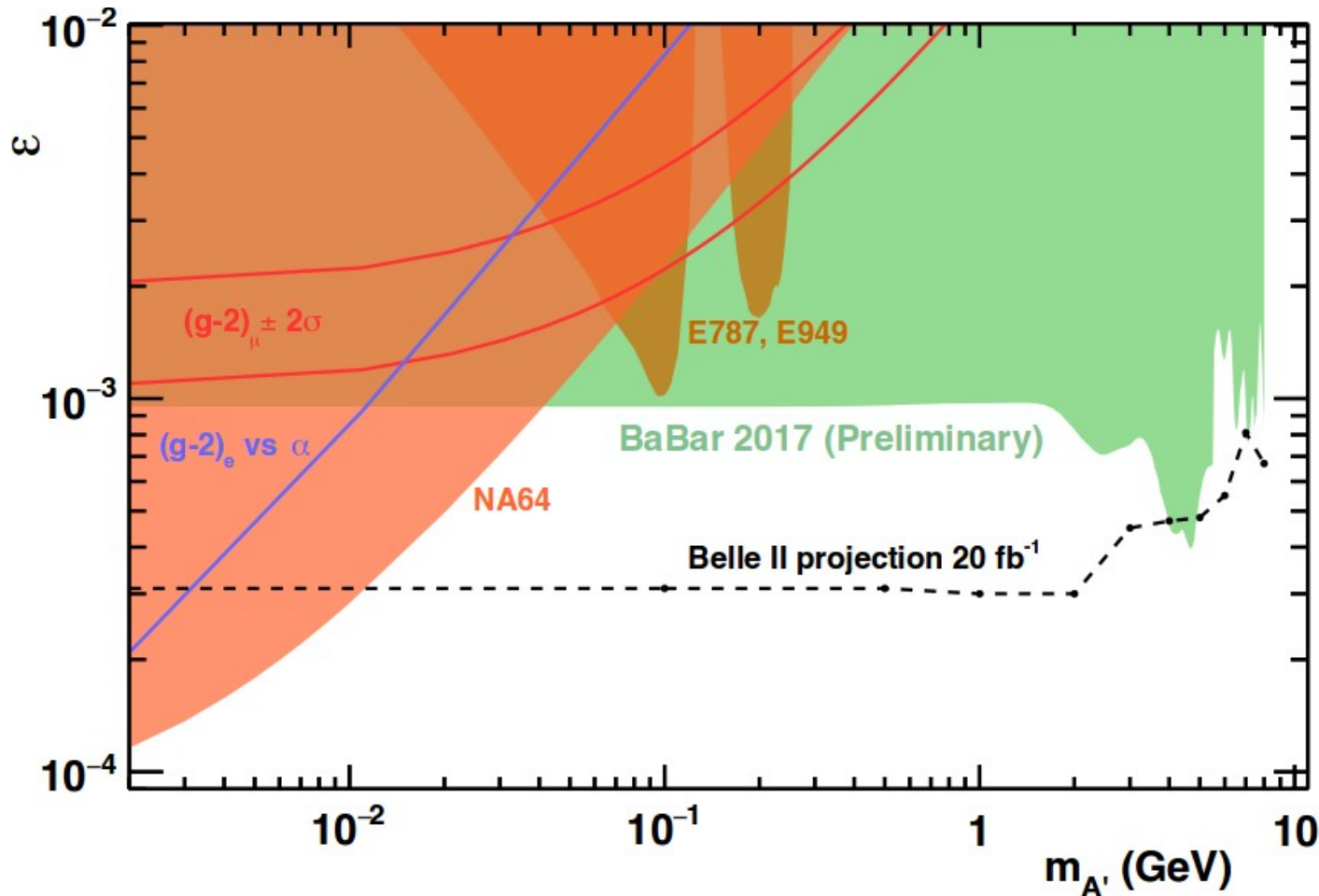
- Barrel ECL without projective cracks in ?

- Lower threshold of the trigger  
- Higher CM energy

- Better cover of ECL gaps with KLM

(@BaBar, lower trigger only @ 2S, 3S)

# Belle II perspective



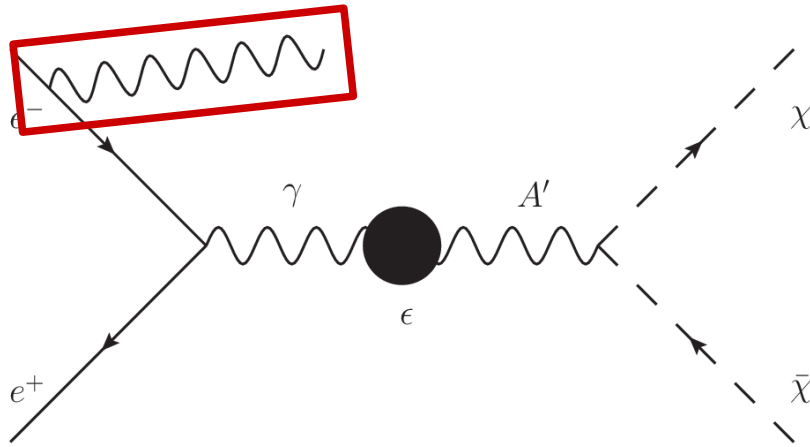
20 fb<sup>-1</sup>  
@ Phase 2

50 ab<sup>-1</sup>  
@ Phase 3 ?

Will it be possible to use the single photon trigger during the Phase 3 with the final luminosity?

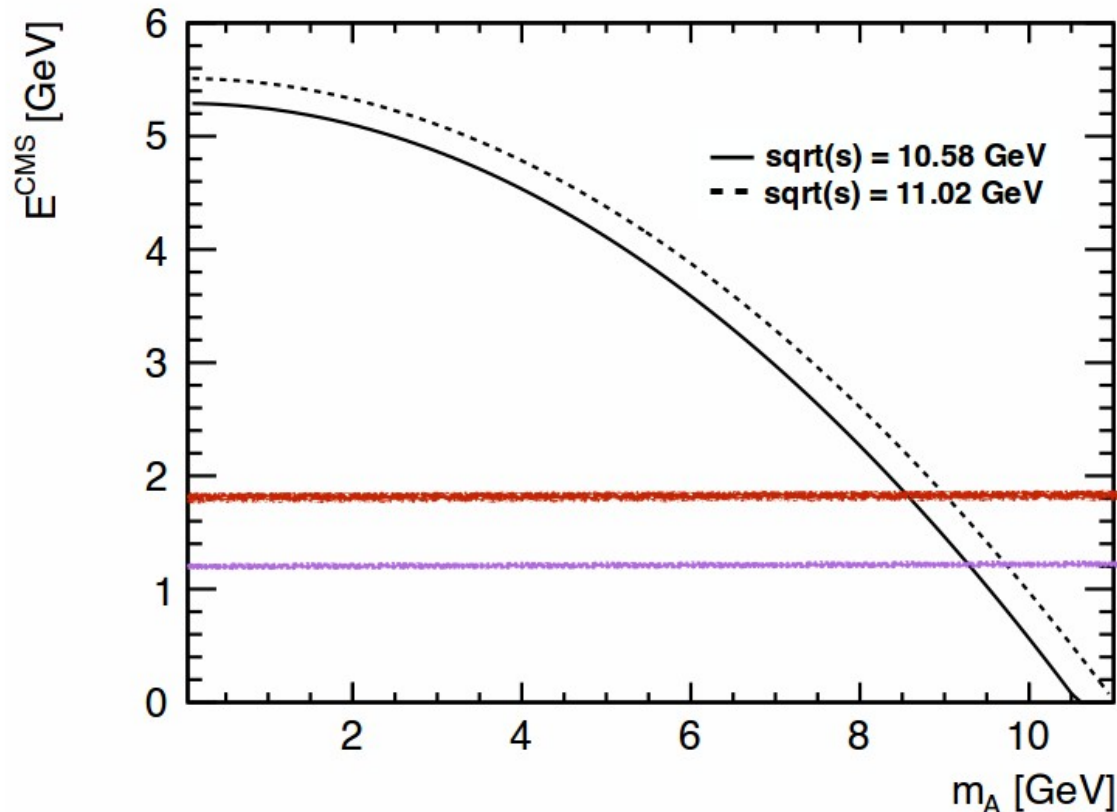
**This could be the only chance for us to do this measurement...**

# Exp. signature



Only 1 photon in the detector  
**Needed a single photon trigger!**

Signal: a “bump” at a  
**given energy of the detected photon**



$$E_{\gamma} = \frac{s - M_{A'}^2}{2\sqrt{s}}$$

$$E_{\text{Trigger}} = 1.8 \text{ GeV}$$

$$E_{\text{Trigger}} = 1.2 \text{ GeV}$$

# Trigger + Evt. selection

Two triggers for single photon:

1)  $E_{CM} > 2 \text{ GeV}$

2)  $E_{CM} > 1 \text{ GeV}$ ,  $E_{CM} \text{ (2nd cluster)} < 0.2 \text{ GeV}$

Common features:  $18.5^\circ < \theta < 139.2^\circ$ , no Bhabba, no  $\gamma\gamma$

Event selection requirements:

- cluster in barrel ECL and  $E > E(\theta)$
- no clusters with  $E_{CM} > 0.1 \text{ GeV}$
- no tracks with  $p_T > 0.2 \text{ GeV}$  in CM
- no KLM clusters outside a  $25^\circ$  cone around the photon or back-to-back

# Backgrounds

## Backgrounds

- ▶  $e^+e^- \rightarrow \gamma\gamma$ , 1 $\gamma$  undetected:  
Peaking, identical to the signal for  $m_{A'} < 1.6 \text{ GeV}/c^2$ . Photons can escape undetected through inefficient detector regions.
- ▶  $e^+e^- \rightarrow \gamma\gamma\gamma$ , 1 $\gamma$  undetected, 2<sup>nd</sup> out of the detector acceptance.
- ▶  $e^+e^- \rightarrow e^+e^-\gamma$ , both electrons out of the detector acceptance ( $\gamma$  energy limited by kinematics).
- ▶ Beam background photons do not fake signal  $\gamma$ , but can be the 2<sup>nd</sup>  $\gamma$  in a signal event.
- ▶ Irreducible SM background  $e^+e^- \rightarrow \nu\nu\gamma$  is negligible.



# After selection

Signal and Background: After cuts.

Release-00-08-00  
Phase2 geometry  
Phase 2 beam backgrounds

## Preselection (reconstructed):

ECL-N1 (gamma) clusters CMS energy sorted: G0, G1

$33^\circ < \text{Theta0Lab} < 127^\circ$

E0CMS dependent cut on Theta0Lab for low E0CMS

**E1CMS < 0.1 GeV\***

All Tracks  $p_t < 0.2$  GeV

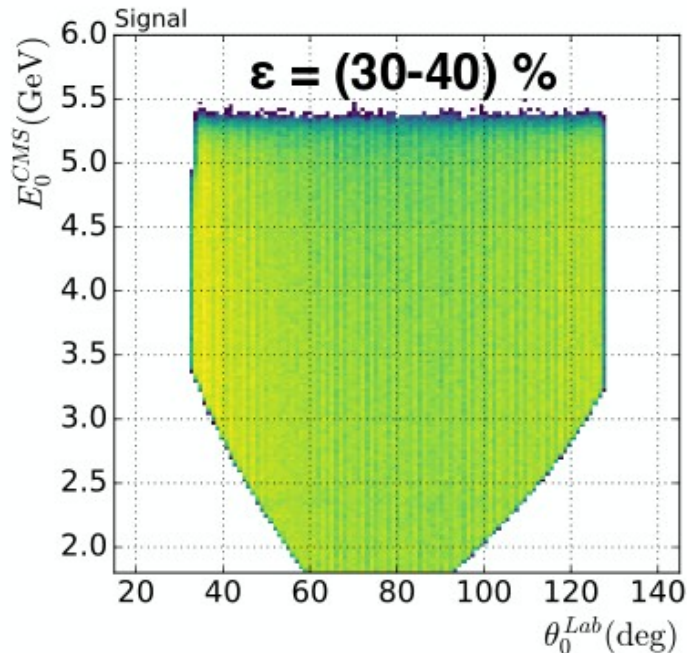
No KLM cluster back to back to G0

**No KLM clusters in KLM veto regions (various gaps)\***

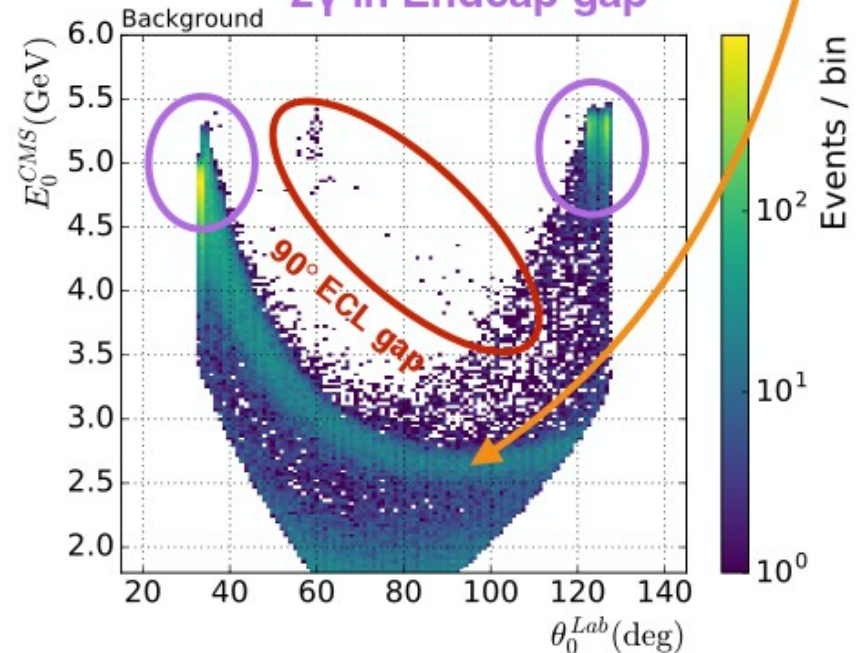
\*Needed after using new MC14 Phase2 background.

3 $\gamma$  in ECL BWD gap  
and KLM BWD gap

2 $\gamma$  in Endcap gap



7



# ECL inefficiencies

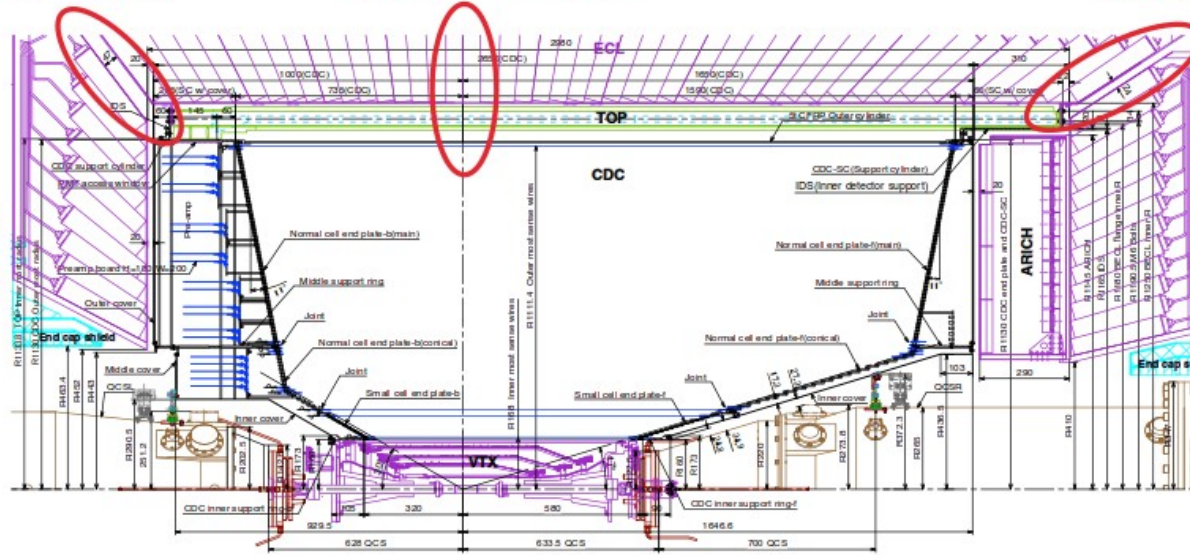
1) BWD gap

4) 1.5mm structure at 90°

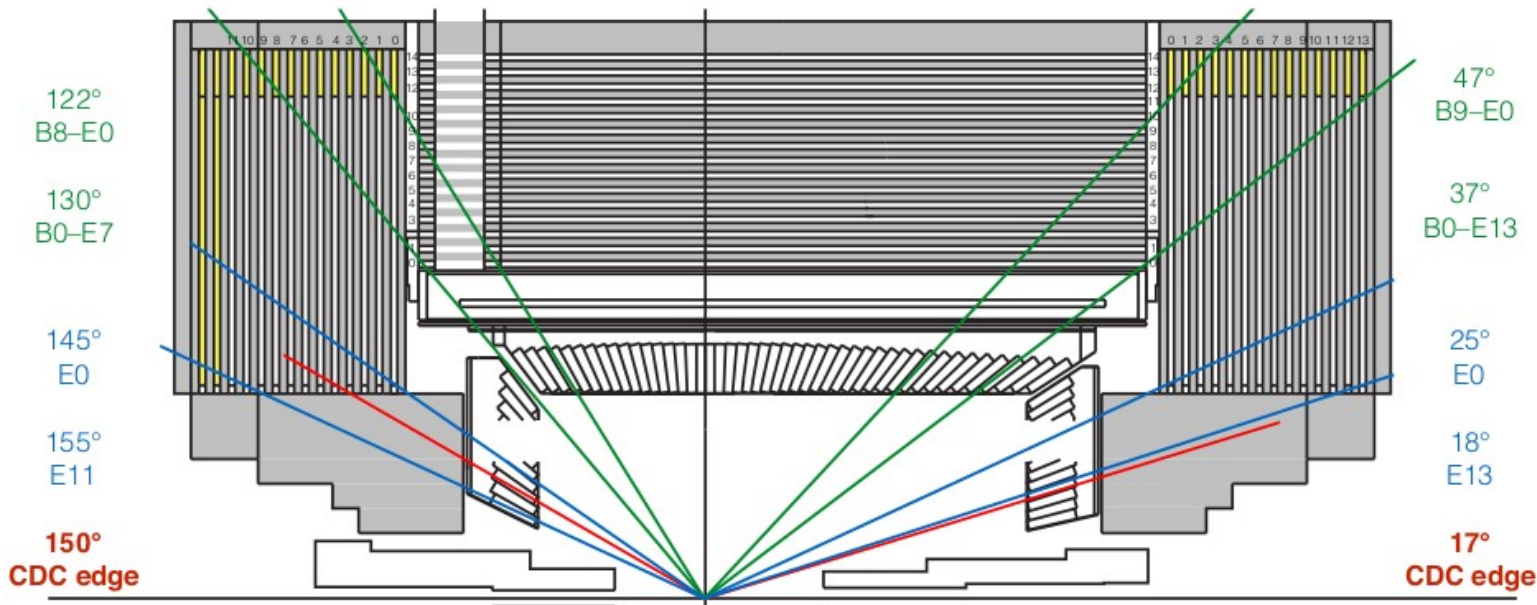
2) FWD gap

3) Projective cracks in endcaps

3) Projective cracks in endcaps



5)  $\gamma$  non-conversion probability  $\sim 3 \times 10^{-6}$



KLM can cover the ECL inefficiencies!

# KLM studies @ Roma Tre

Main goal(s):

- map of the KLM efficiency for the photon detection
  - comparison data/MC: how much reliable are the geometry and simulation of the KLM?

Info contained in the KLMClusters data objects:

- time
- vector with cluster position
  - momentum
- position and momentum errors
  - # of layers with a cluster hit
- innermost layer with a cluster hit

# KLM studies @ Roma Tre

Main goal(s):

- map of the KLM efficiency for the photon detection
  - comparison data/MC: how much reliable are the geometry and simulation of the KLM?

Info contained in the KLMClusters data objects:

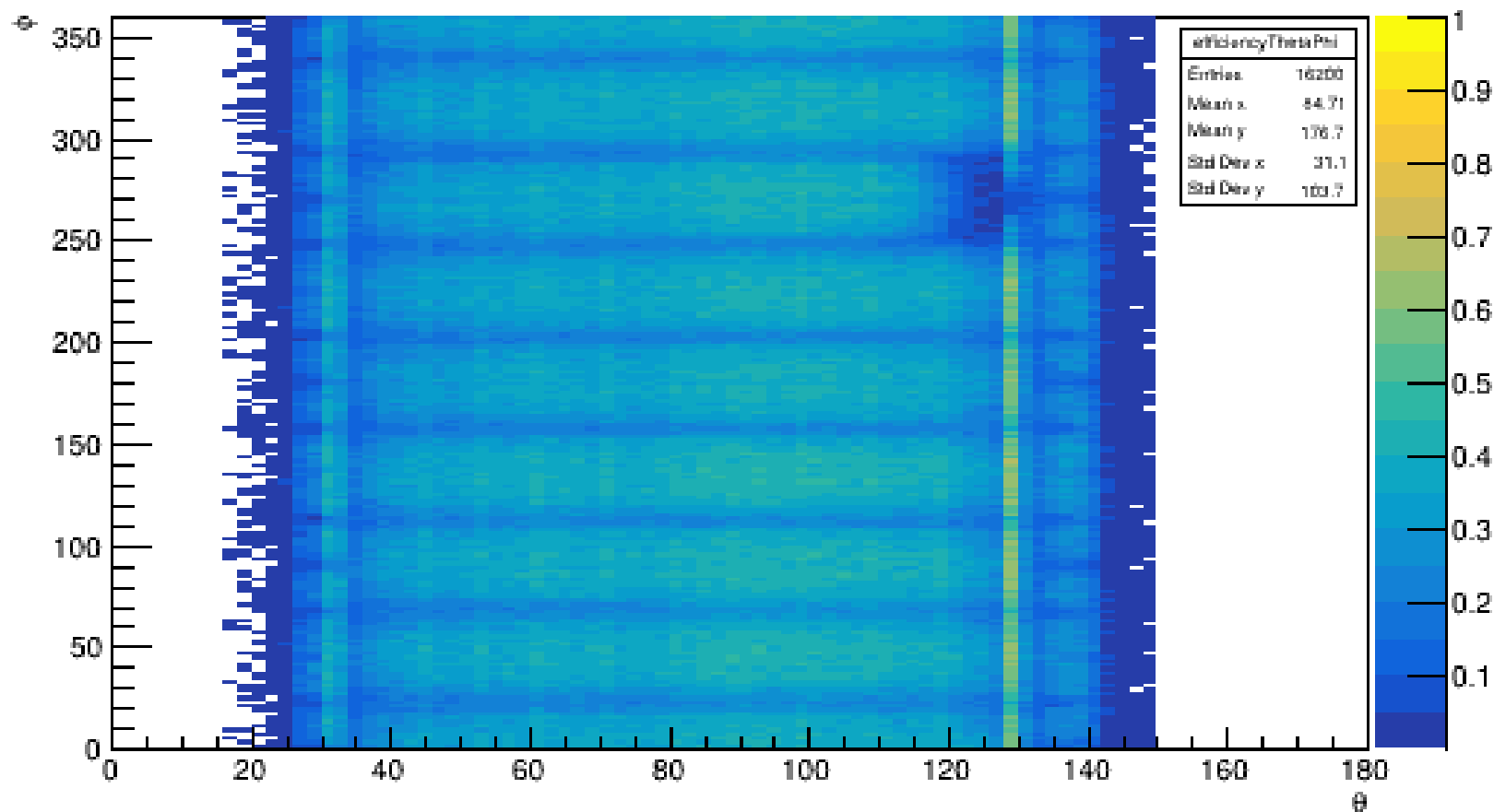
- time
- vector with cluster position
  - momentum
- position and momentum errors
  - # of layers with a cluster hit
- innermost layer with a cluster hit
- requested this morning to Leo additional info about the shape of the cluster (number of 2D hits per layer)

# KLM studies @ Roma Tre

Map of the KLM efficiency for the photon detection:  
in principle, already available from MC photons generated with the ParticleGun

$0 < E < 10 \text{ GeV}$

Efficiency:  $\theta$  vs.  $\phi$



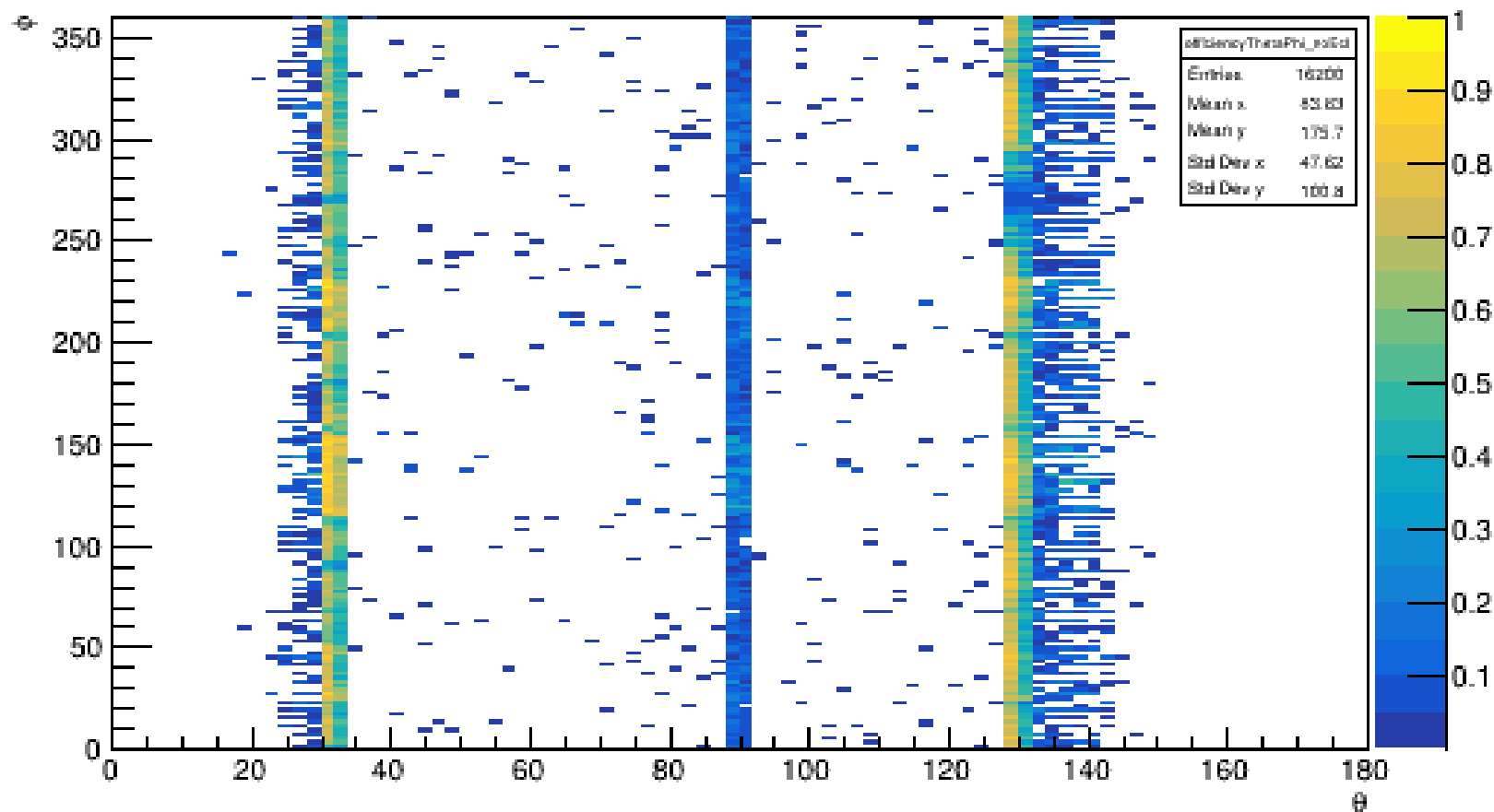
$\frac{\text{\#MCPhotons with } \geq 1 \text{ KLMCluster}}{\text{\#MCPhotons generated}}$

# KLM studies @ Roma Tre

Map of the KLM efficiency for the photon detection:  
in principle, already available from MC photons generated with the ParticleGun

$0 < E < 10 \text{ GeV}$

Efficiency:  $\theta$  vs.  $\phi$  - No ECLClusters associated



$$\frac{\text{\#MCPhotons with } \geq 1 \text{ KLMCluster \&\& } 0 \text{ ECLClusters}}{\text{\#MCPhotons generated with } 0 \text{ ECLClusters}}$$

# KLM studies @ Roma Tre

But there is a “problem” with the simulation:  
there are no informations at the mDST level  
that tell us if a photon “reached” the KLM or not,  
in order to apply the correct normalization.

NB: the function `hasSeenInDetector(Const::KLM)`  
returns always 0 for photons,  
even if they have a `KLMCluster` associated

# KLM studies @ Roma Tre

But there is a “problem” with the simulation:  
there are no informations at the mDST level  
that tell us if a photon “reached” the KLM or not,  
in order to apply the correct normalization.

*It's impossible to apply  
the correct normalization  
at the MC level :(*

NB: the `SeenInDetector(Const::KLM)`  
returns always 0 for photons,  
even if they have a `KLMCluster` associated



# KLM studies @ Roma Tre

Datasets available during Phase 2:

Sample	Note	Generated sigma nb	Percentage selected	Accepted sigma nb	Rate Hz 40 nb-1/sec	Fiducial efficiency %	Barrel efficiency %
Bhabha	0.5 & 5 deg	122760	0.150	184	7358	92.2	100
gamma gamma		25.2	12.4	3.1	125	96.9	100
e e e e		1693	0.28	4.7	188		
e e mu mu		67.8	3.1	2.1	84		
tau tau		0.919	91.9	0.8	34	94.6	97.6
mu mu		1.115	70.8	0.8	32	92.5	100
BB		1.05	100.0	1.1	42		
u u-bar		1.61	90.7	1.5	58		
d d-bar		0.4	90.4	0.4	14		
s s-bar		0.38	95.9	0.4	15		
c c-bar		1.3	100.0	1.3	52		
2gamma production of ALP	0.2 GeV					12.1	
	0.5 GeV					85.9	
	2 GeV					97.6	
	10 GeV					99.0	100
2gamma production of pi0	no tag					2.1	0.2
	1 tag						
ALP--> invisible	9.3 GeV					82.7	93.1
ALP --> gamma gamma	0.2 GeV					99.1	100
	0.5 GeV					99.3	
	3 GeV					99.6	
	9.3 GeV					99.7	
a' --> e e	0.5 GeV					97.8	100
a' --> invisible	0.5 GeV					83.6	100.0
	9.3 GeV					74.4	94.0
gamma pi+ pi-	0.5 GeV					96.3	99.9
tau --> e gamma						99.4	100.0
tau --> mu gamma						98.8	99.8
Y3S --> pi pi Y1S						44.0	49.8
TOTAL				200	8003		

# KLM studies @ Roma Tre

Datasets available during Phase 2:

Sample	Note	Generated sigma nb	Percentage selected	Accepted sigma nb	Rate Hz 40 nb-1/sec	Fiducial efficiency %	Barrel efficiency %
Bhabha	0.5 & 5 deg	122760	0.150	184	7358	92.2	100
gamma gamma		25.2	12.4	3.1	125	96.9	100
e e e e		1693	0.28	4.7	188		
e e mu mu		67.8	3.1	2.1	84		
tau tau		0.919	91.9	0.8	34	94.6	97.6
mu mu		1.115	70.8	0.8	32	92.5	100

For systematic studies:

? (mu mu) accepted: 0.8 nb

? (mu mu gamma) accepted: ~ 0.008 nb

mu mu gamma events: ~ 150000 (Phase 2)

e e gamma events are "less clean",  
but more abundant!

	9.3 GeV					99.7	
a' --> e e	0.5 GeV					97.8	100
a' --> invisible	0.5 GeV					83.6	100.0
	9.3 GeV					74.4	94.0
gamma pi+pi-	0.5 GeV					96.3	99.9
tau --> e gamma						99.4	100.0
tau --> mu gamma						98.8	99.8
Y3S --> pi pi Y1S						44.0	49.8
TOTAL				200	8003		

# KLM studies @ Roma Tre

I learned during the last days

*how to use correctly gbasf2,*

so I will start as soon as possible to look at the mu mu and mu mu gamma events from the MC production, in order to measure the KLM efficiency with the Phase 2 data.

# KLM studies @ Italy

The Napoli group (with Francesco Di Capua) will join us for these analysis and studies (the activities need to be defined)



# Brief overview of the other low multiplicity analyses during Phase 2

# Overview

$ee \rightarrow \gamma + \text{ALP}$ , $\text{ALP} \rightarrow \gamma\gamma$ (single, di, and tri photons) and $ee \rightarrow ee + \text{ALP}$ (two photon production)	Searches	<a href="#">@ Christopher Hearty</a> <a href="#">@ Torben Ferber</a>	in progress, <a href="https://arxiv.org/abs/1409.4792">https://arxiv.org/abs/1409.4792</a>  <a href="https://confluence.desy.de/download/attachments/53769418/20170526_ferber.pdf?version=1&amp;modificationDate=1495742716842&amp;api=v2">https://confluence.desy.de/download/attachments/53769418/20170526_ferber.pdf?version=1&amp;modificationDate=1495742716842&amp;api=v2</a>  <a href="https://kds.kek.jp/indico/event/24563/session/21/contribution/120/material/slides/0.pdf">https://kds.kek.jp/indico/event/24563/session/21/contribution/120/material/slides/0.pdf</a>	yes* (tri photon final state only)
$e^+ e^- \rightarrow \pi \pi$ $\gamma_{\text{ISR}} (g-2)$	Low Multiplicity	<a href="#">@ Boris Shwartz</a> <a href="#">@ Torben Ferber</a> <a href="#">@ Yosuke Maeda</a> <a href="#">@ John Michael Roney</a>	in progress, <a href="https://arxiv.org/pdf/1205.2228v1.pdf">https://arxiv.org/pdf/1205.2228v1.pdf</a>	no* (repeat BaBars QED study as a first step)
$\text{AFB } e^+ e^- \rightarrow l^+ l^-$ , $l = \mu, \tau$	Low Multiplicity	<a href="#">@ Torben Ferber</a>	in progress, help needed! <a href="http://www.actaphys.uj.edu.pl/fulltext?series=Reg&amp;vol=46&amp;page=2285">http://www.actaphys.uj.edu.pl/fulltext?series=Reg&amp;vol=46&amp;page=2285</a>	no* (repeat BaBars QED study as a first step)

# Overview

Luminosity Measurement	Low Multiplicity	@ Robert Seddon	in progress, help needed!	yes
$ee \rightarrow \tau\tau\Phi$ , $\Phi \rightarrow l^+l^-$ , $l = e, \mu, \tau$ (dark skalar)	Searches	-	<a href="https://arxiv.org/abs/1606.04943">https://arxiv.org/abs/1606.04943</a>	yes
$e^+e^- \rightarrow \gamma_{ISR} + A', A' \rightarrow W'W'$ (4l)	Searches	-	help needed!	?
$e^+e^- \rightarrow \gamma_{ISR} + A', A' \rightarrow A'H', H' \rightarrow A'A'$ (6l)	Searches	-	help needed!	?
$e^+e^- \rightarrow \gamma_{ISR} + A', A' \rightarrow A'H', H' \rightarrow \text{invisible}$	Searches	-	help needed!	?
$e^+e^- \rightarrow \gamma_{ISR}$ MM (magnetic monopoles)	Searches	@ Dmitrii Neverov	<a href="http://pdg.lbl.gov/2010/reviews/rpp2010-rev-mag-monopole-searches.pdf">http://pdg.lbl.gov/2010/reviews/rpp2010-rev-mag-monopole-searches.pdf</a> <a href="https://arxiv.org/abs/1707.05295">https://arxiv.org/abs/1707.05295</a>	yes

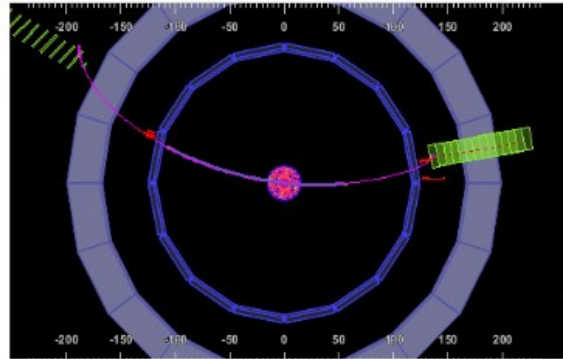
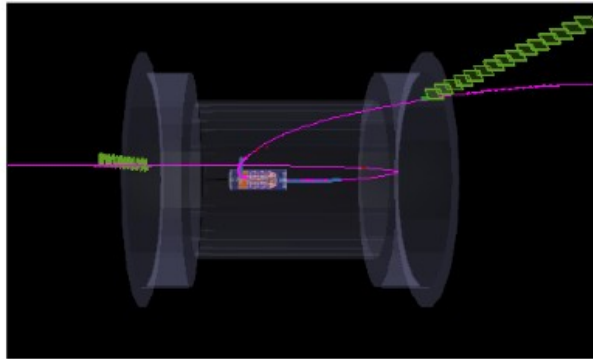
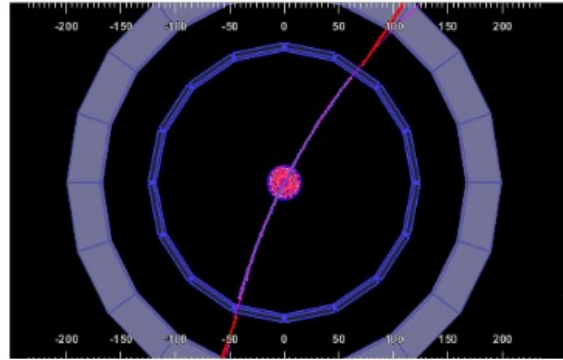
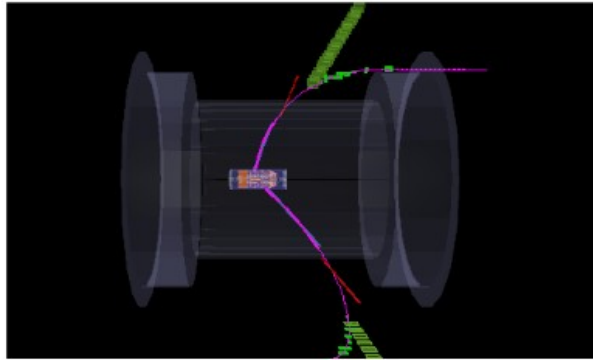
# Overview

Luminosity Measurement	Low Multiplicity	@ Robert Seddon	in progress, help needed!	yes
$ee \rightarrow \tau\tau\Phi$ , $\Phi \rightarrow l^+l^-$ , $l = e, \mu, \tau$ (dark skalar)	Searches	-	<a href="https://arxiv.org/abs/1606.04943">https://arxiv.org/abs/1606.04943</a>	yes
$\tau \rightarrow e/\mu$ invisible	Tau	-	Not published. *One can probably re-interpret a result from an existing Belle or BaBar 1 prong decay analysis but looking at the muon momentum spectrum and not detecting any peaks.  <a href="http://pdglive.lbl.gov/BranchingRatio.action?parCode=S035&amp;desig=102">http://pdglive.lbl.gov/BranchingRatio.action?parCode=S035&amp;desig=102</a> ,  <a href="http://pdglive.lbl.gov/BranchingRatio.action?parCode=S035&amp;desig=103">http://pdglive.lbl.gov/BranchingRatio.action?parCode=S035&amp;desig=103</a>	yes*
$e^+e^- \rightarrow \gamma_{ISR} + A'$ , $A' \rightarrow \text{hadrons}$	Searches	-	Best limits from KLOE2 up to $\sim 1$ GeV	
$e^+e^- \rightarrow \gamma_{ISR} + A'$ , $A' \rightarrow \text{Pseudo Dirac DM with large mass splitting}$	Searches	-	Pseudo-Dirac DM with a large mass splitting. The higher mass DM particle will decay into the lower mass DM and an off-shell $A'$ , $A'$ must decay to SM. Mass splitting $< 0.1$ mass, final state will be missing energy and $e^+e^- \gamma_{ISR}$ .	yes, but very difficult final state



# Search for monopoles

Example events,  $q=1, g=1e$



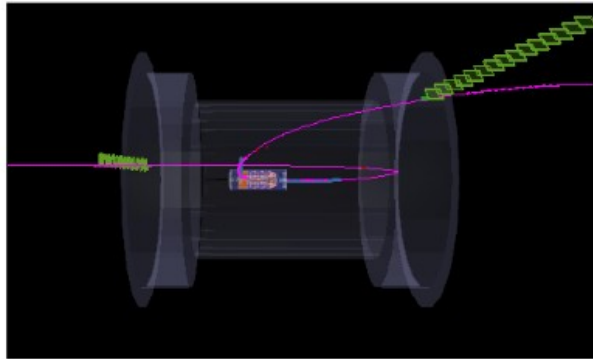
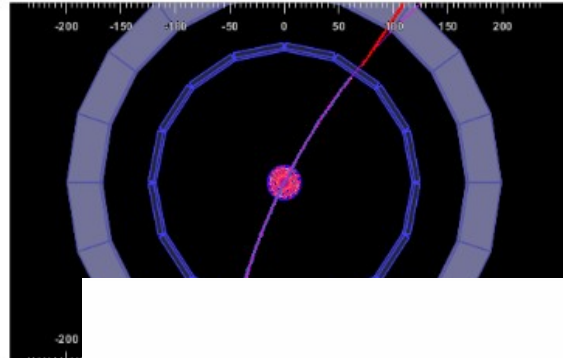
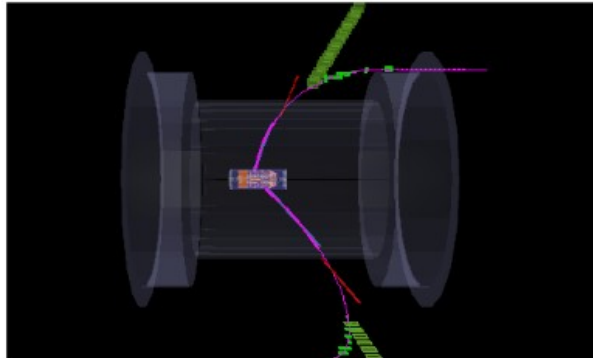
11 Oct 2017

Monopole searches; Dmitrii Neverov; 28th B2GM

5

# Search for monopoles

Example events,  $q=1, g=1e$

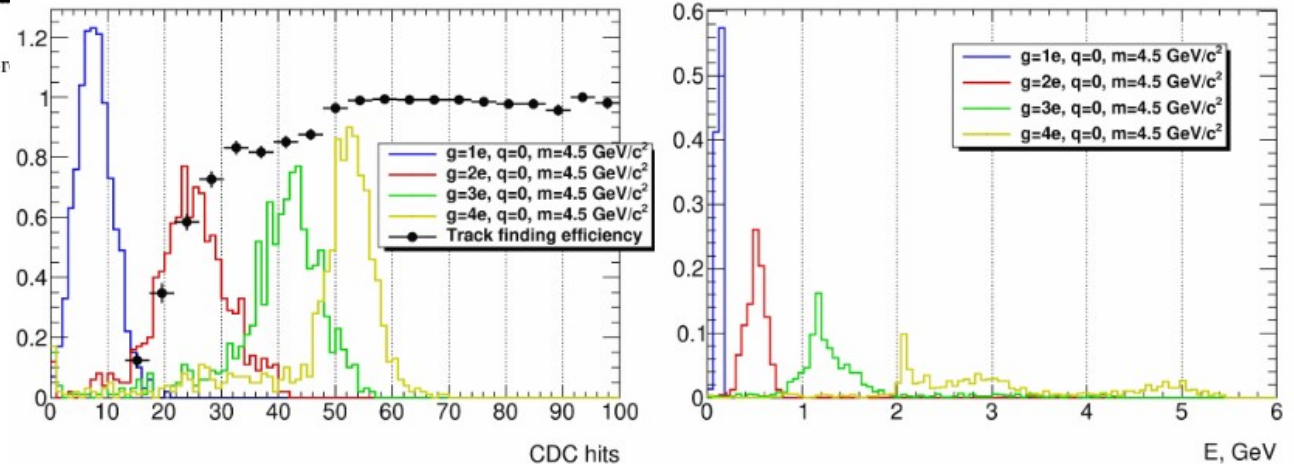


## Single monopole hits

- Leaves CDC, ECL, and KLM hits
- Because of extra  $\beta^2$  factor in ionisation, heavy ( $4.5 \text{ GeV}/c^2$ ) monopoles of low charge give a faint signal
- Higher charge monopoles and dyons produce clear signal

11 Oct 2017

Monopole searches; Dmtirii Never



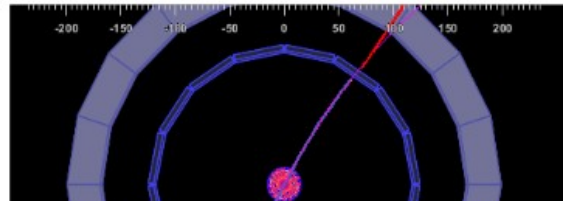
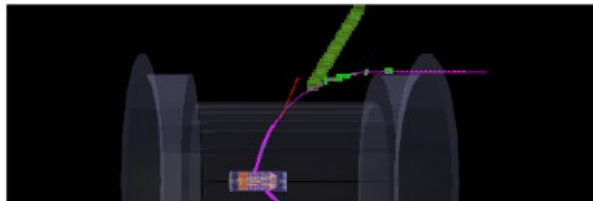
11 Oct 2017

Monopole searches; Dmtirii Neverov; 28th B2GM

6

# Search for monopoles

Example events,  $q=1, g=1e$



Few results for low-charge monopoles:  
during Phase 2 Belle II will be  
competitive with present results!

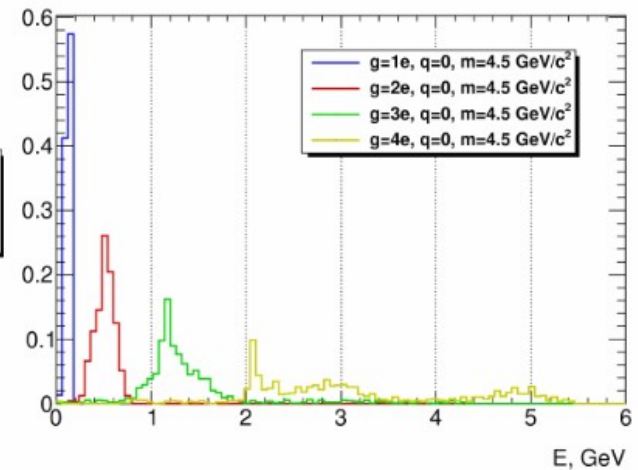
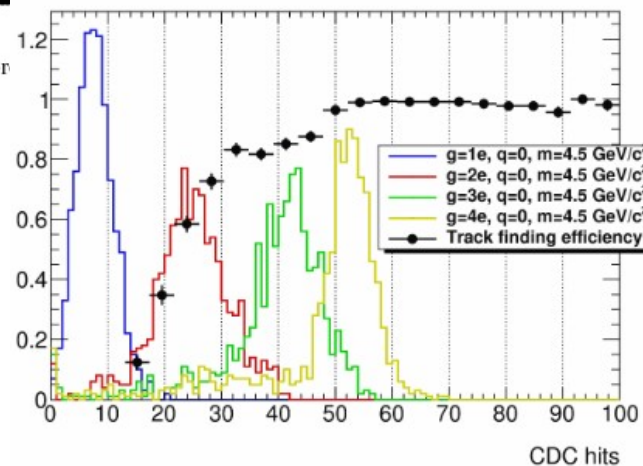
We have monopole simulation in basf2  
from release-01-00-00 :)

e hits

poles of

11 Oct 2017

Monopole searches; Dmitrii Never



11 Oct 2017

Monopole searches; Dmitrii Neverov; 28th B2GM

6

# Backup slides

# Dark photon

# Search for dark photon



**Dedicated experiments + Multipurpose experiments**

# Theory

Dark photon = bosonic mediator of a “dark interaction”  
between dark matter (and SM) particles

**A dark photon could explain the muonic g-2 anomaly!**

Minimal model:

$$\mathcal{L}_{A'} = -\frac{1}{4}F'^{\mu\nu}F'_{\mu\nu} + \boxed{\frac{1}{2} \frac{\epsilon}{\cos \theta_W} B^{\mu\nu} F'_{\mu\nu}} - \frac{1}{2}m_{A'}^2 A'^{\mu} A'_{\mu}$$

at low energies:

$$\boxed{\mathcal{L}_{\text{kin.mix.}} = \frac{1}{2}\epsilon F^{\mu\nu} F'_{\mu\nu}}$$

**Free parameters (to be measured):**

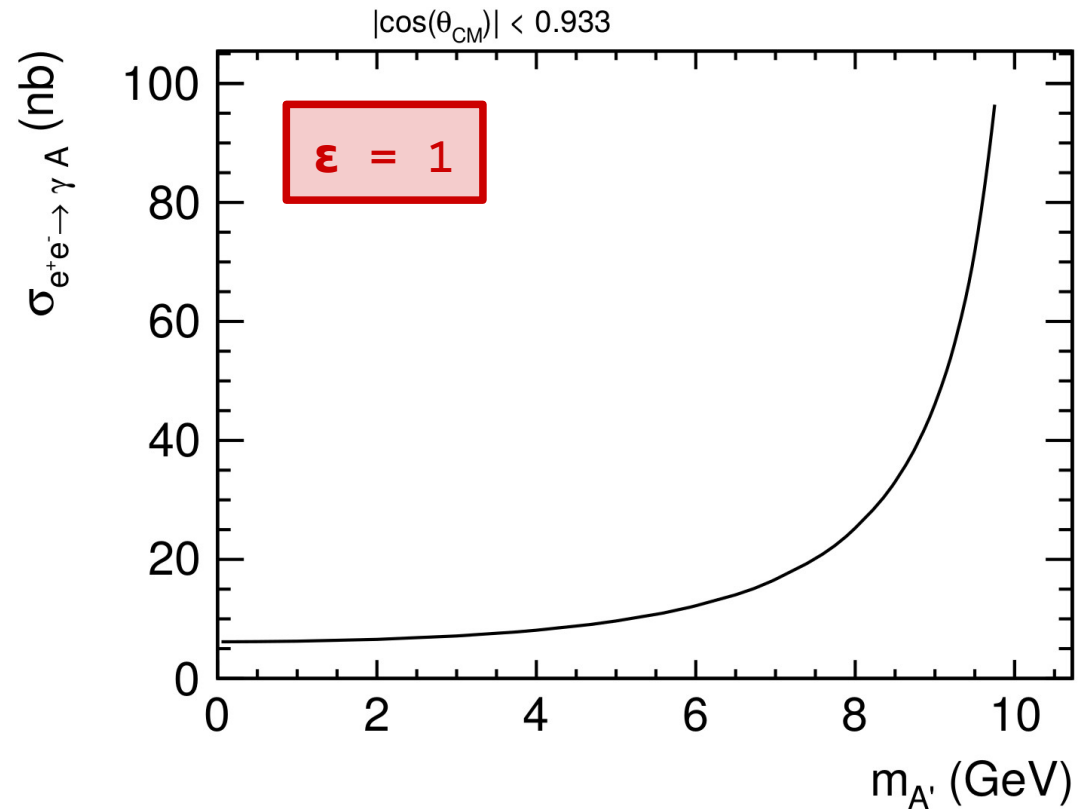
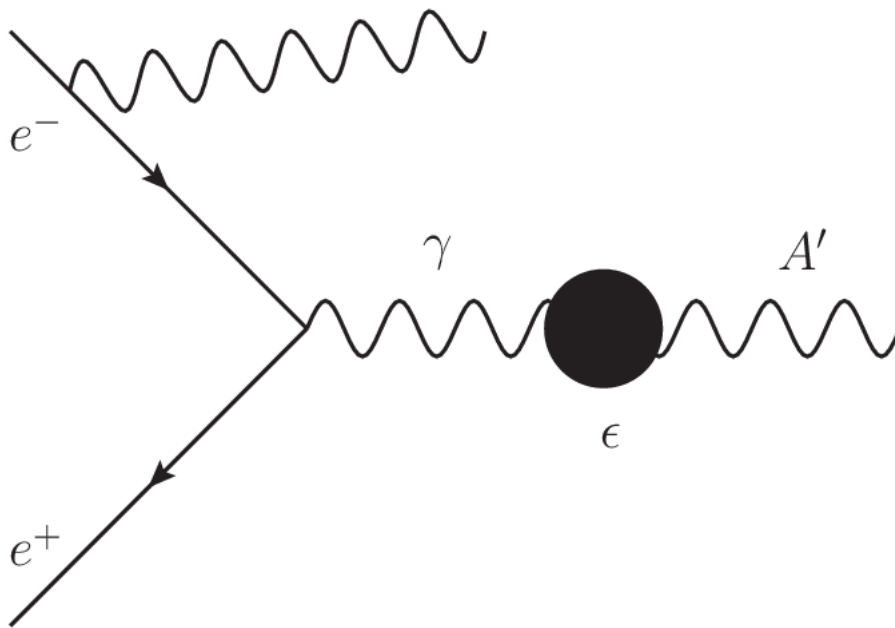
$\epsilon$  (strength of the mixing)

$m_A$  (mass of the dark photon)

# Production

Several production mechanisms at  $e^+ e^-$  colliders:

I will focus on  $e^+ e^-$  annihilation



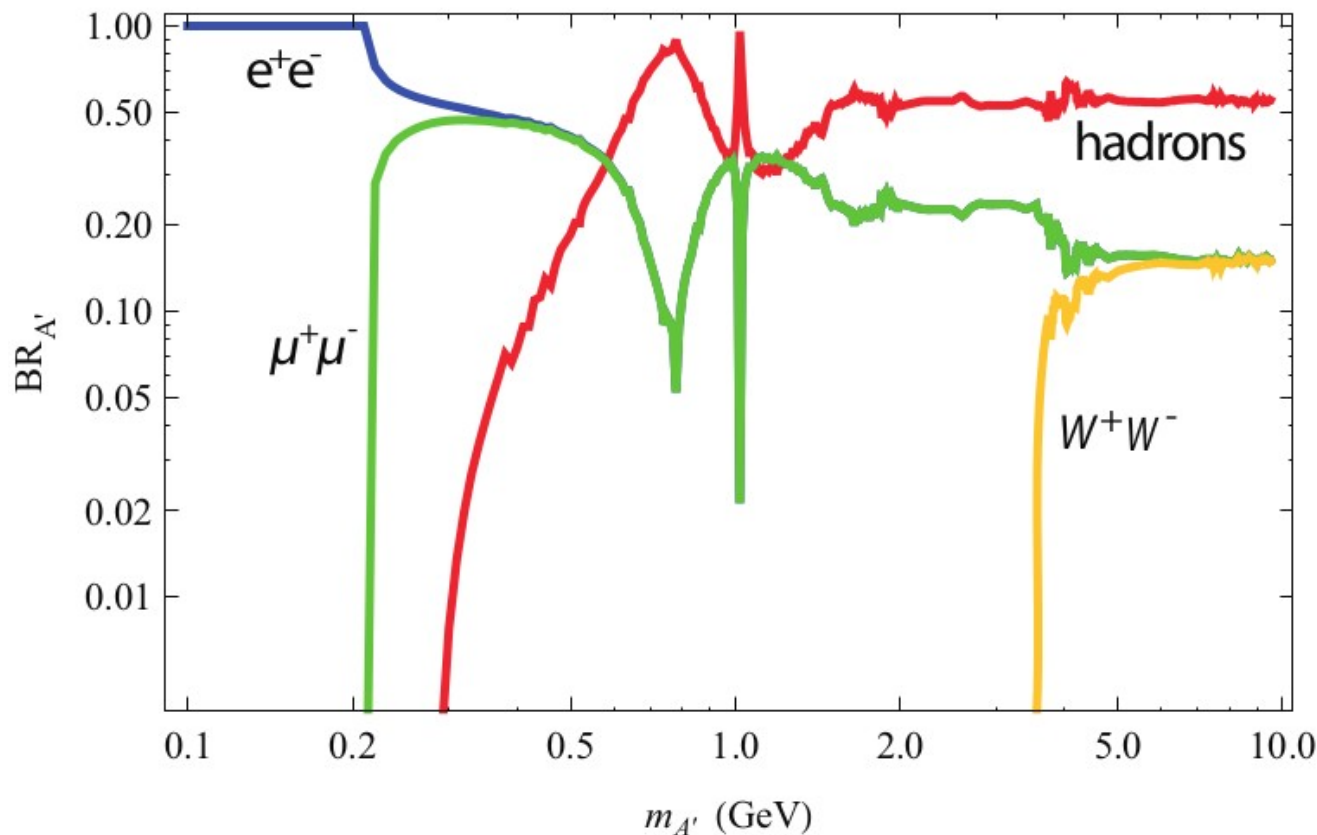
$$\frac{d\sigma(e^+e^- \rightarrow \gamma A')}{d\cos\theta} = \frac{\alpha\epsilon^2}{2s^2(s - m_{A'}^2)} \left( \frac{s^2 + m_{A'}^4}{\sin^2\theta} - \frac{(s - m_{A'}^2)^2}{2} \right)$$



# Visible decay

Two different scenarios: visible vs. invisible decay

If  $m_x > \frac{1}{2} m_{A'}$ : dark photon decays into SM particles

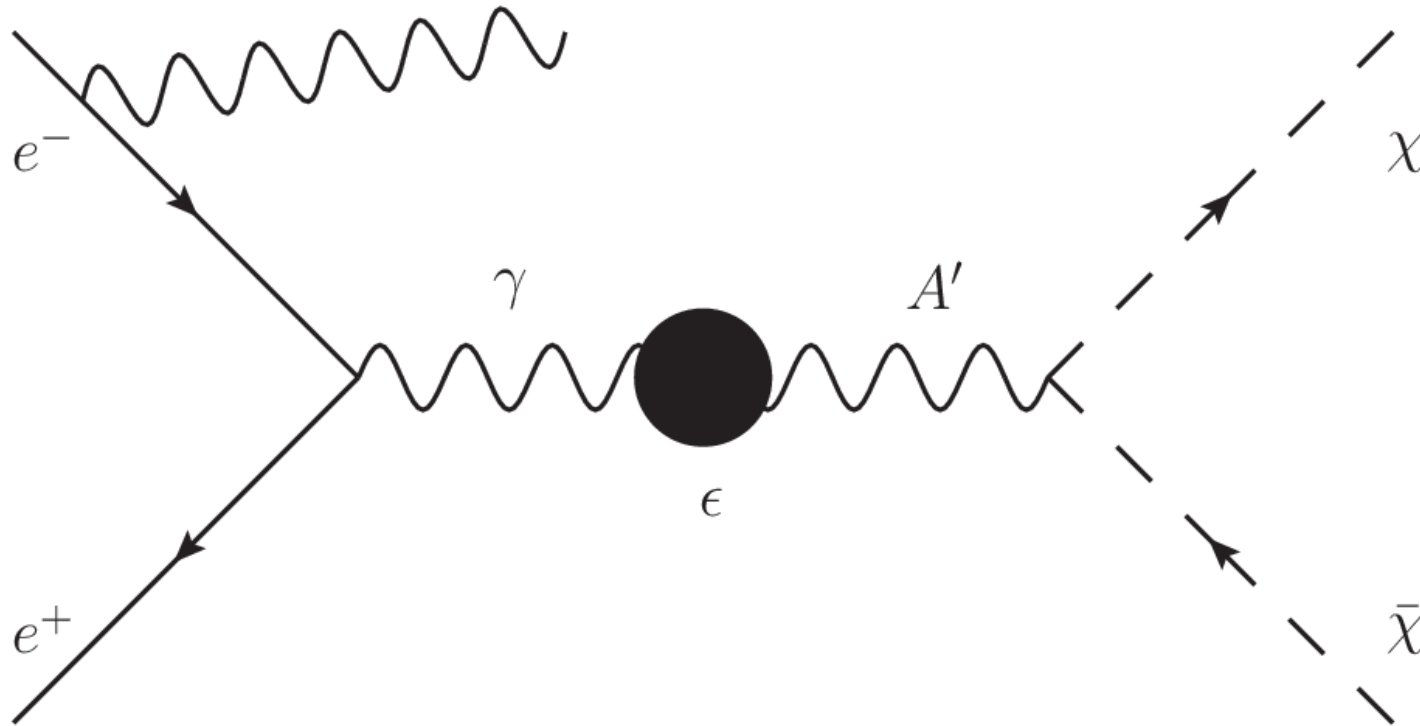


$$\Gamma_{A' \rightarrow \text{had}} = \frac{1}{3} \alpha \epsilon^2 M_{A'} \sqrt{1 - \frac{4m_\mu^2}{M_{A'}^2}} \left( 1 + \frac{2m_\mu^2}{M_{A'}^2} \right) \times \frac{\Gamma(e^+e^- \rightarrow \text{hadrons})}{\Gamma(e^+e^- \rightarrow \mu^+\mu^-)} (E = M_{A'})$$

# Invisible decay

If  $m_\chi < \frac{1}{2} m_{A'}$ : dark photon can decay into DM particles

**Main decay if the coupling with DM isn't suppressed**



$$\Gamma_{A' \rightarrow \chi\bar{\chi}} = \frac{1}{3} \alpha_D M_{A'} \sqrt{1 - \frac{4m_\chi^2}{M_{A'}^2}} \left( 1 + \frac{2m_\chi^2}{M_{A'}^2} \right)$$

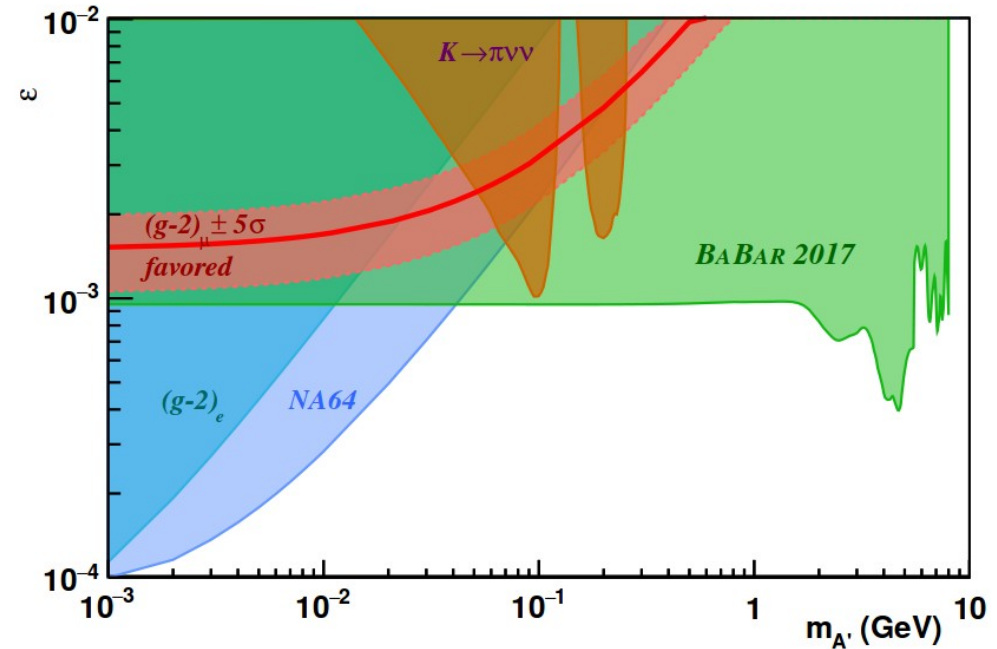
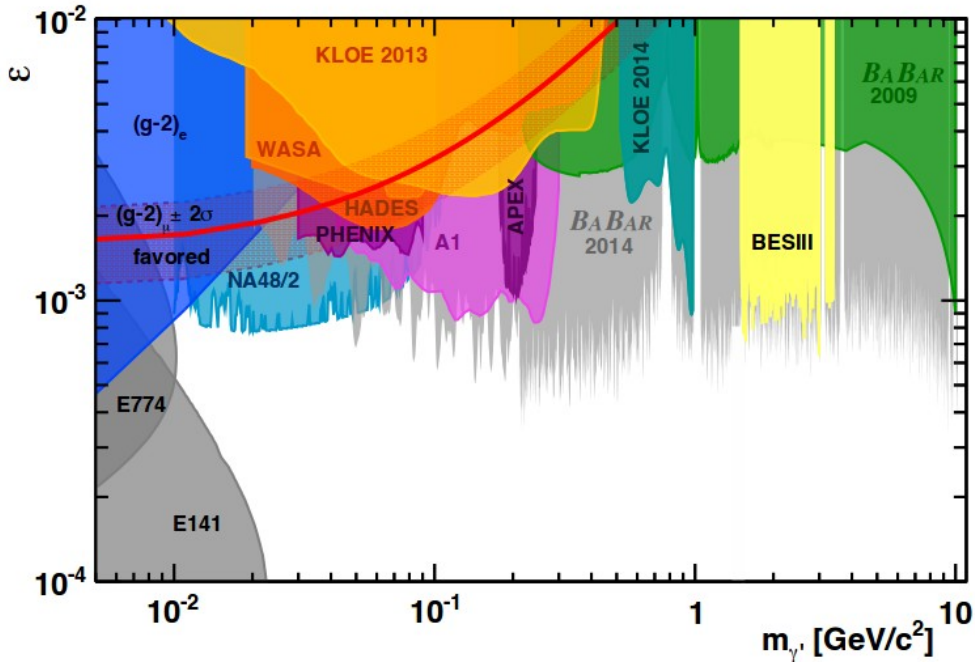
No suppression:

$$\alpha_D \gg \alpha \epsilon^2$$

# Experimental status

Visible decay ( $\rightarrow l^+ l^-$ )

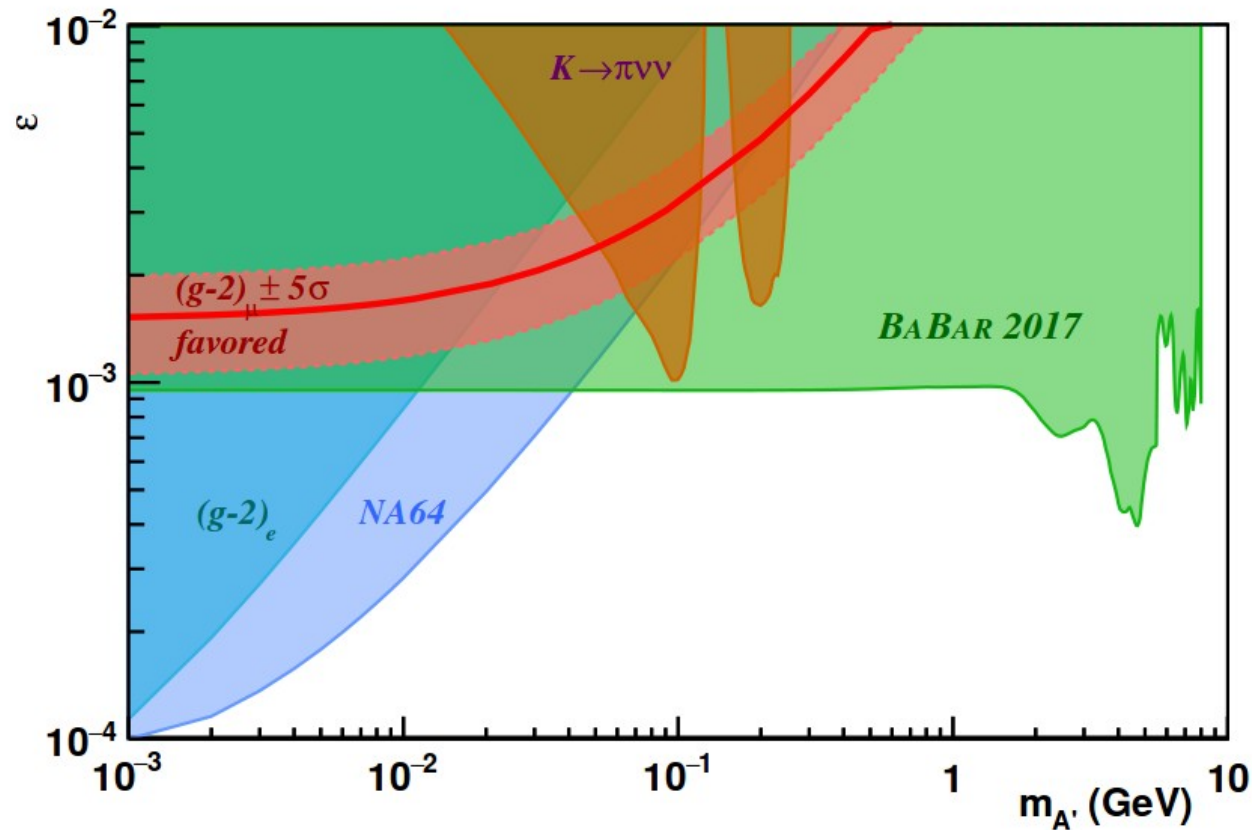
Invisible decay



**No direct comparison between results  
regarding visible and invisible decay**

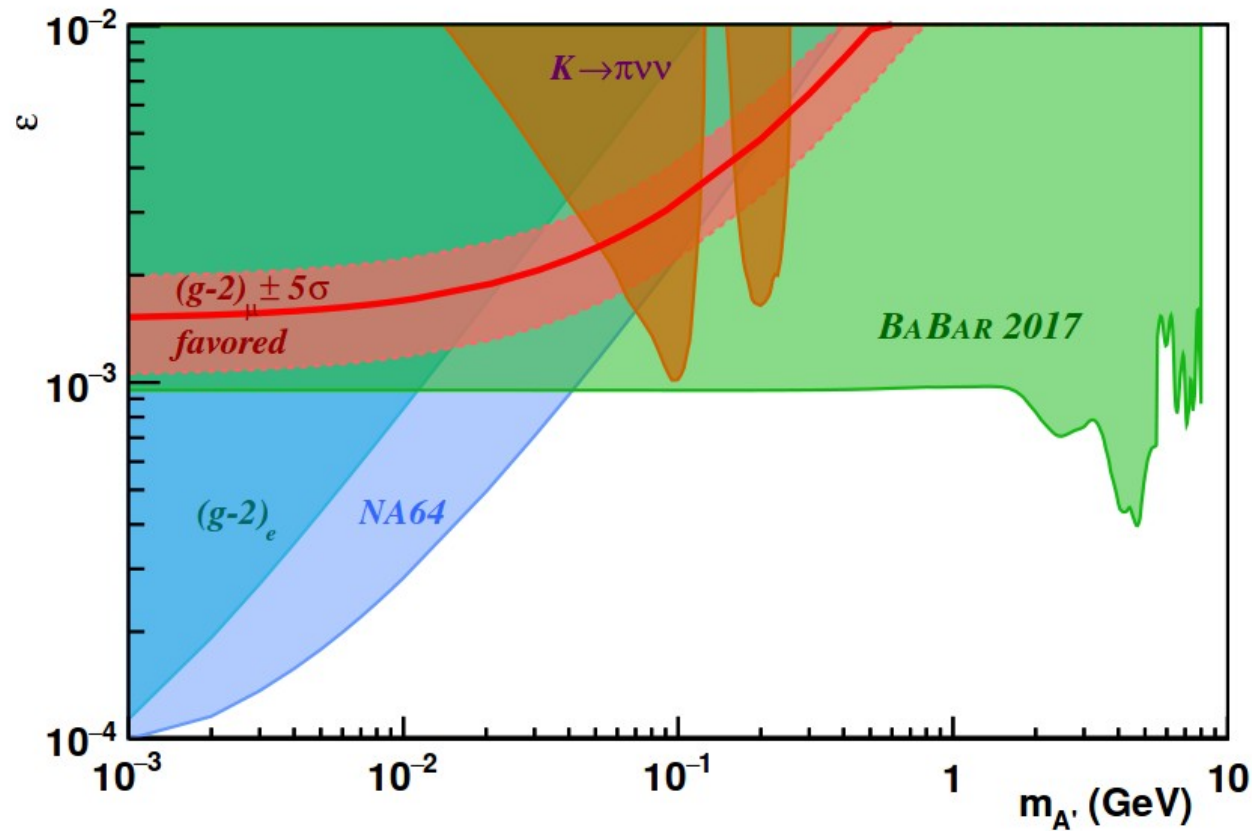
Many other results related to hadronic decays  
and different models/mechanisms

# Dark photon $\rightarrow$ Invisibles



BaBar and NA64 **ruled out** the possibility to explain completely the  $g-2$  anomaly introducing a dark photon  
(**assuming light DM:  $m_x < \frac{1}{2} m_{A'}$** )

# Dark photon $\rightarrow$ Invisibles



But:

- it can still **partially explain** the  $g-2$  anomaly (+ other NP)
- it's still an important **portal to light DM**