



Charm quarks to probe Glasma in p-Pb collisions

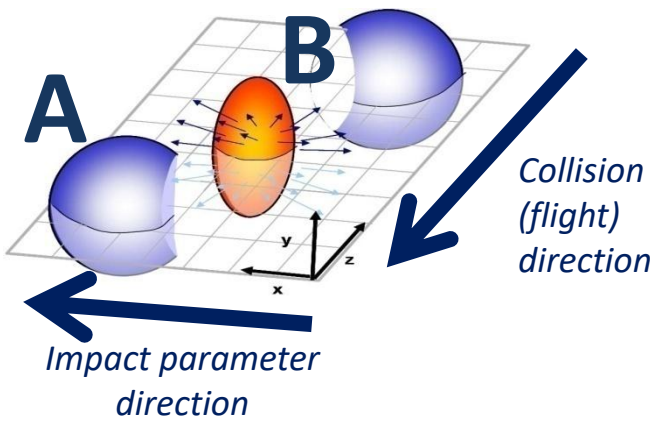
QCD@WORK 2018

The *first* QCD conference ever held in Matera, Italy

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School of Nuclear Science and Technology
Lanzhou University

- High energy nuclear collisions and the Glasma
- Gluon fields in p-Pb collisions
- ***Heavy quarks and the cathode tube effect***
- Conclusions



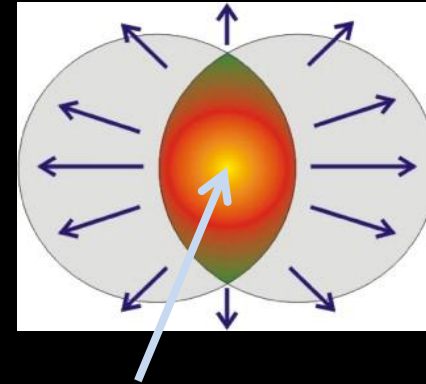


A,B: Cu, Au (RHIC@BNL)
 Pb (LHC@CERN)
 p (LHC@CERN)
 p-Pb collisions (LHC@CERN)
 d-Au collisions (RHIC@BNL)

Au - Au : $\sqrt{s} = 200 \times A$ GeV at RHIC
 Pb - Pb : $\sqrt{s} = 2.76 \times A$ TeV at LHC
 p - Pb : $\sqrt{s} = 5.02$ TeV at LHC
 p - p : $\sqrt{s} = 5, 7$ and 13 TeV at LHC



High energy nuclear collisions



Collision direction

In Au-Au, Pb-Pb, Cu-Cu,.....:
 Hot and dense expanding
QUARK-GLUON-PLASMA (QGP)

QGP formation time

- RHIC ≈ 0.6 fm/c $\approx 10^{-24}$ sec
- LHC ≈ 0.2 fm/c

QGP lifetime

- RHIC ≈ 5 fm/c $\approx 10^{-23}$ sec
- LHC ≈ 10 fm/c





Is it possible to describe collisions within effective field theory?

What kind of system is created after the collision?

Which observable we can look at in order to probe this system?



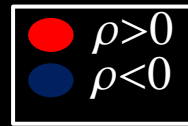
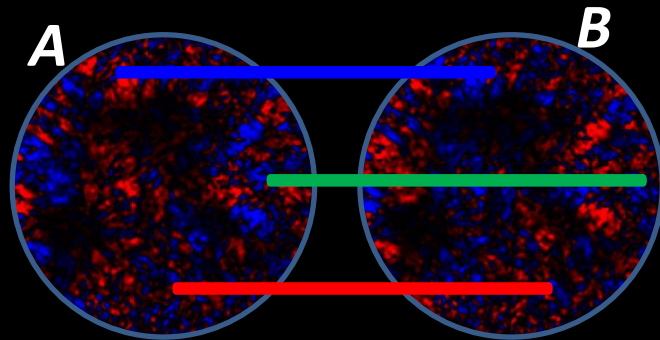
Nucleons/Nuclei at high energy:

Small- x gluons dominate the wave function.

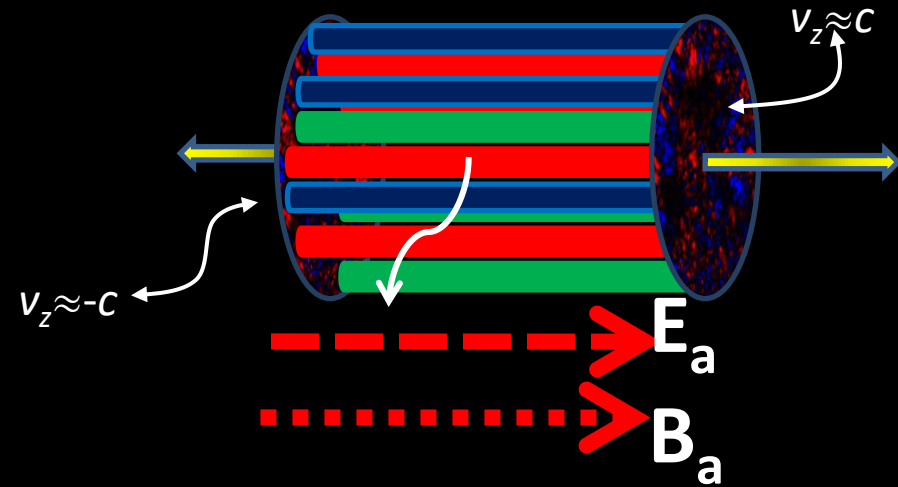
Collision within effective theory (Color-Glass-Condensate, aka CGC):

Interaction of two dense gluon systems.

McLerran and Venugopalan (1996)
and many others



$$\begin{aligned}\nabla \cdot \mathbf{E} &= ig [\mathcal{A}^i, \mathcal{E}^i] \\ \nabla \cdot \mathbf{B} &= ig [\mathcal{A}^i, \mathcal{B}^i]\end{aligned}$$



Glasma

Due to the large density the gluon field behaves like a classical field:

Dynamics is governed by classical EoMs, namely the classical Yang-Mills (CYM) equations.

$$\begin{aligned}\frac{dA_i^a(x)}{dt} &= E_i^a(x), \\ \frac{dE_i^a(x)}{dt} &= \sum_j \partial_j F_{ji}^a(x) + \sum_{b,c,j} f^{abc} A_j^b(x) F_{ji}^c(x)\end{aligned}$$

Here:

$$F_{ij}^a(x) = \partial_i A_j^a(x) - \partial_j A_i^a(x) + \sum_{b,c} f^{abc} A_i^b(x) A_j^c(x)$$

QCD equivalent of the Maxwell Equations in vacuum space.

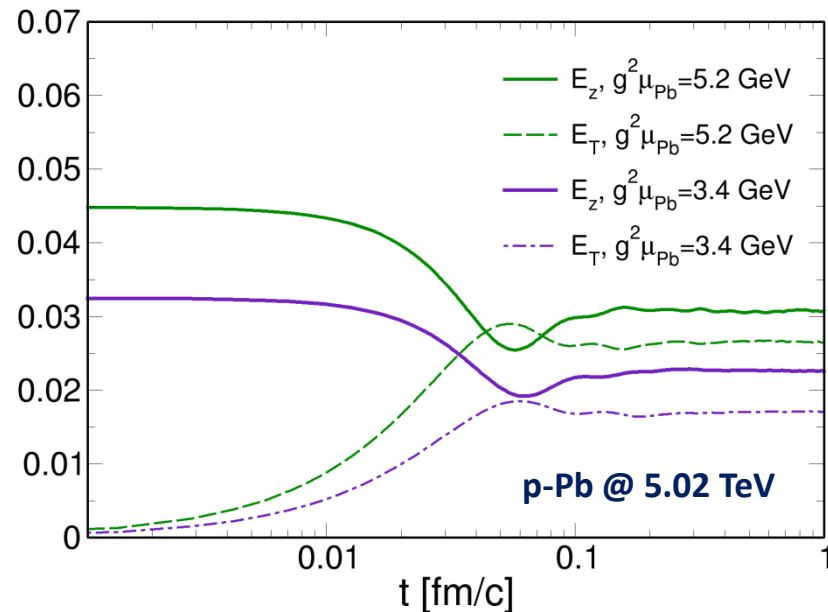
$$A_\mu \rightarrow \frac{A_\mu}{g}$$

Field strength tensor

Evolution of the system is studied assuming the Glasma initial condition, and evolving this condition by virtue of the CYM equations.

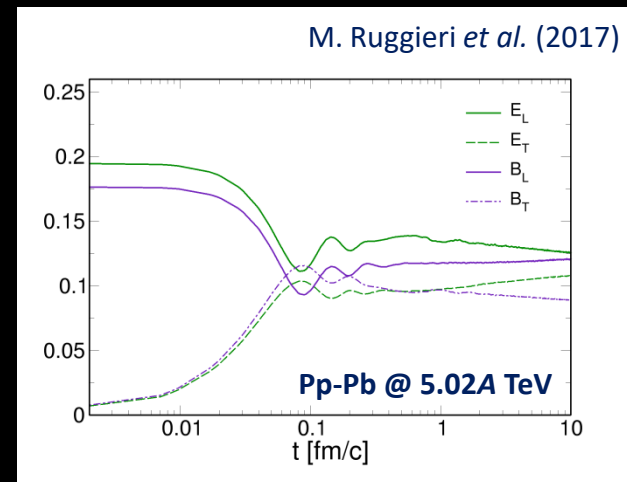
No longitudinal expansion

M. R. and S. K. Das, 1805.09617



The evolving Glasma: fields in p-Pb collisions

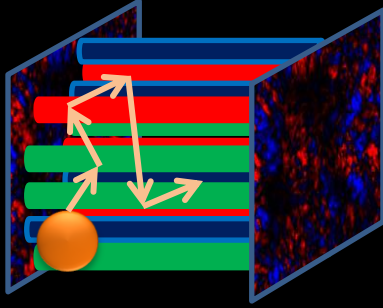
Fields



$$\left. \frac{dE_a^x}{dt} \right|_{t=0+} = \partial_y B_z^a + f_{abc} A_y^b B_z^c$$

Formation time of transverse fields:
 $\Delta t \approx 0.1$ fm/c

Heavy quarks as probes of the evolving Glasma



$$t_{\text{formation}} \approx \frac{1}{2m_c} \approx 0.06 \text{ fm}/c$$



HQs can probe the *very early evolution* of the Glasma fields

Hamilton equations of motion of c-quarks:

$$\frac{dx_i}{dt} = \frac{p_i}{E} \quad E = \sqrt{\mathbf{p}^2 + m^2}$$

$$E \frac{dp_i}{dt} = gQ_a F_{i\nu}^a p^\nu$$

$$E \frac{dQ_a}{dt} = -gQ_c \varepsilon^{cba} \mathbf{A}_b \cdot \mathbf{p}$$

Wong (1979)

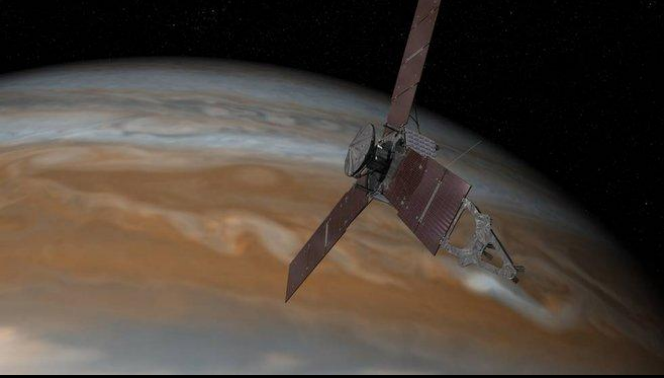
$$\mathbf{v} \equiv \frac{\mathbf{p}}{E} \quad (\text{Relativistic}) \text{ Velocity}$$

$$\frac{d\mathbf{p}}{dt} = q\mathbf{E} + q(\mathbf{v} \times \mathbf{B}) \quad \text{Lorentz force}$$

$$D_\mu J_a^\mu = 0 \quad \text{Gauge-invariant conservation of the color current carried by charm quarks + gluons}$$

$$J_a^\mu = \bar{c} \gamma^\mu T_a c$$

Equations of motion of heavy quarks are solved in the background given by the evolving Glasma fields



Heavy quarks as **probes** of the evolving Glasma

S. K. Das *et al.* (2017)
Rapp *et al.* (2014)
Mrorcinsky (2017)
Scardina *et al.* (2016)
Greiner (2018)
Goussiaux *et al.* (2015)

Heavy quarks are:

- Quite massive
- Quite diluted

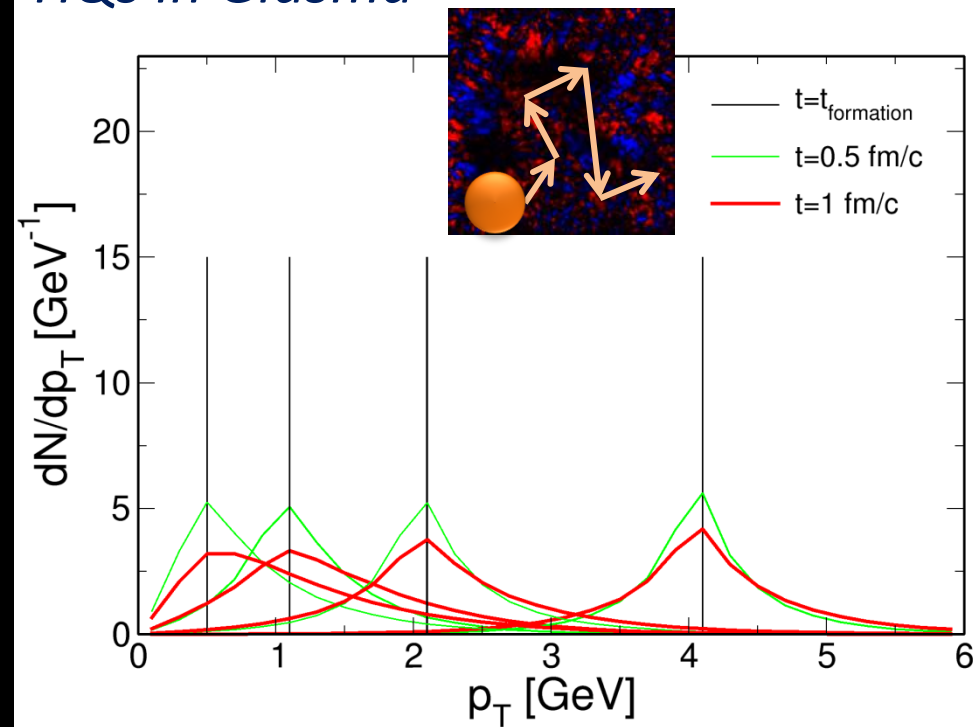
- *Carry negligible color current*
- *Self-interactions occure rarely*

≈ No disturbance to the evolving gluon fields

HQs are real probes of the evolving Glasma

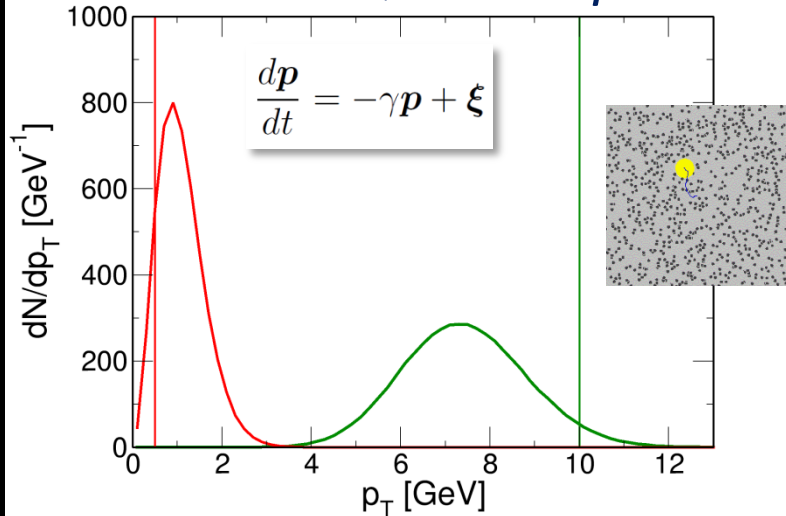
HQs in Glasma

M. Ruggieri and S. K. Das, *in preparation*

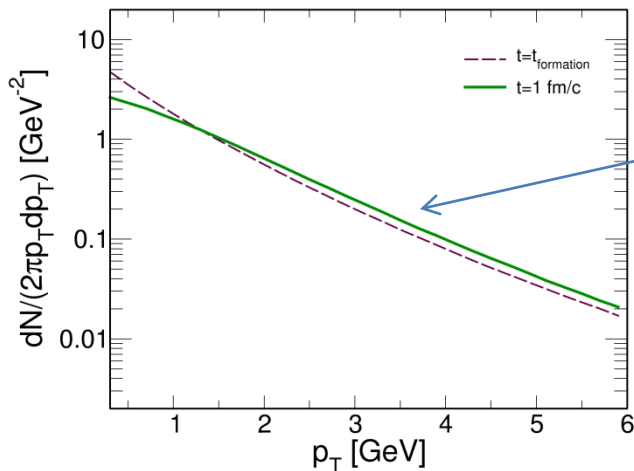


(Anomalous) Diffusion
in momentum space

T=500 MeV HQs in hot plasma



p-Pb @ 5.02 TeV



Nuclear modification factor (R_{pPb}) for p-Pb collisions

Initial distribution: from perturbative QCD, aka **prompt**

Evolution: **interaction with the Glasma**

$$R_{pPb} = \frac{(dN/d^2p_T)_{\text{final}}}{(dN/d^2p_T)_{pQCD}}$$

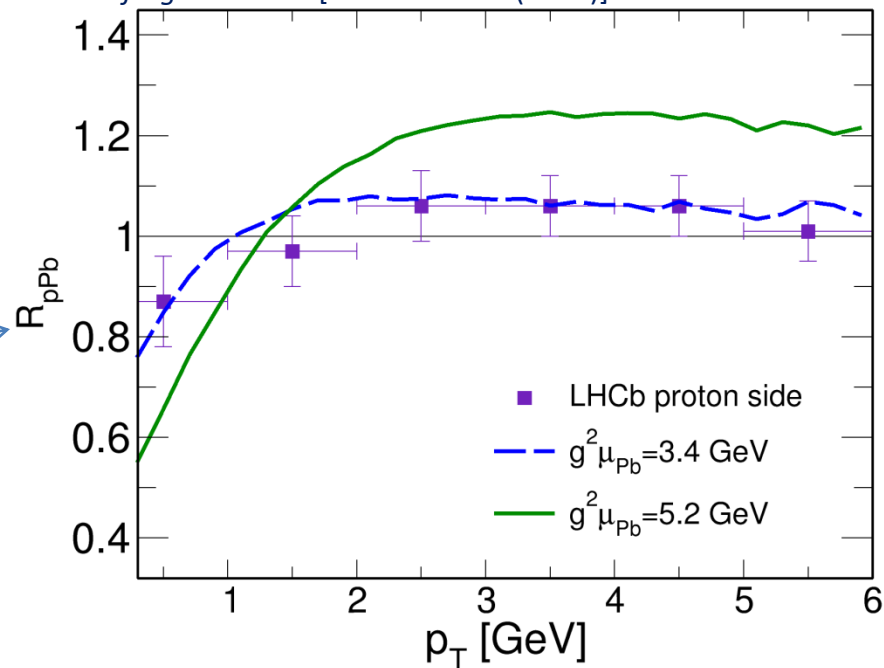
$R_{pPb} \neq 1$

Interaction with the fields created
by the collision

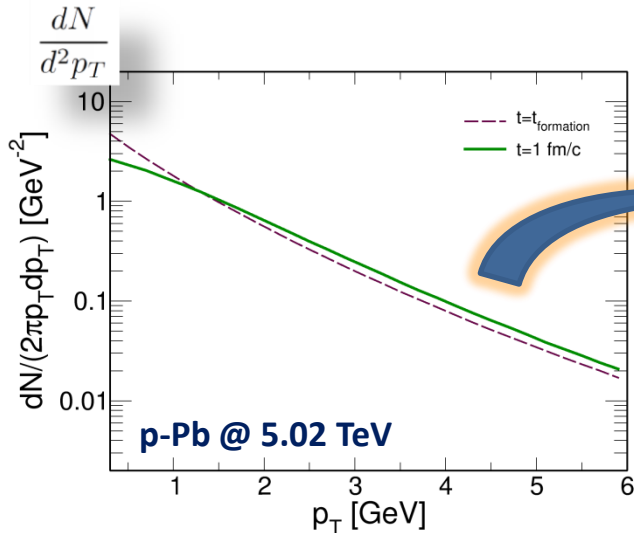
D-mesons R_{pPb}

M. Ruggieri and S. K. Das, 1805.09617

Standard fragmentation [Peterson et al.(1983)]

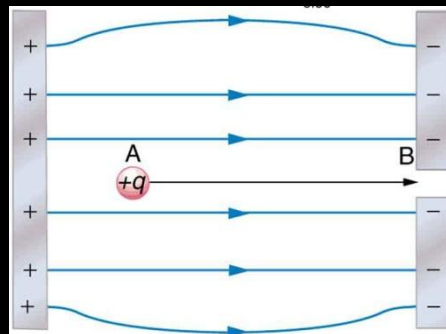


Diffusion results in acceleration

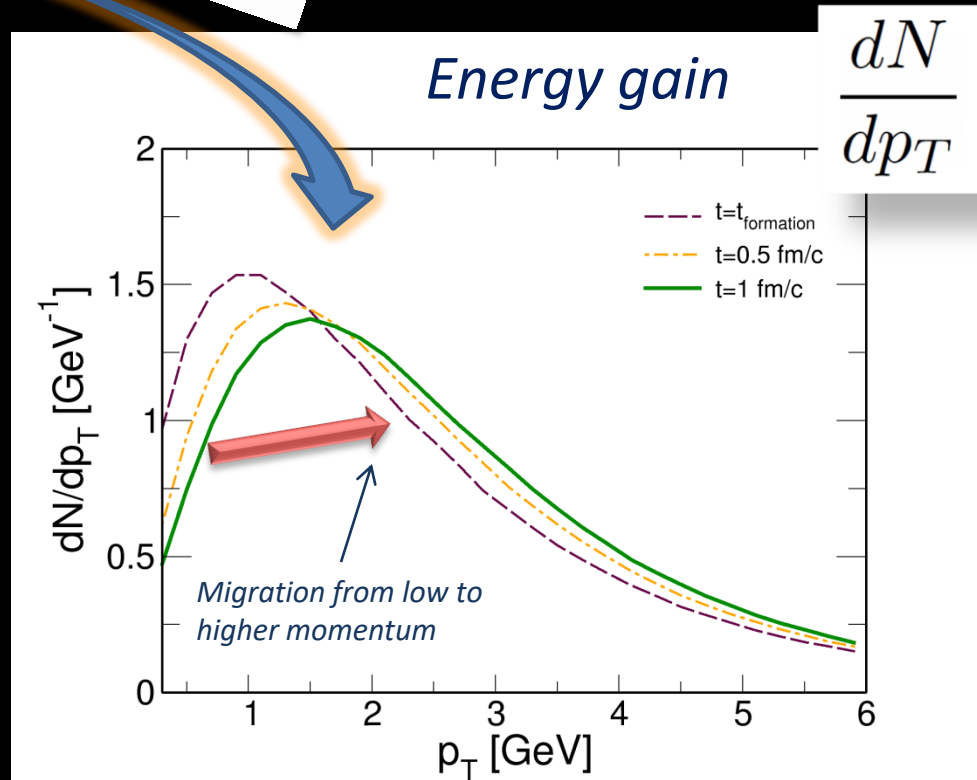


$$\times p_T (\times 2\pi)$$

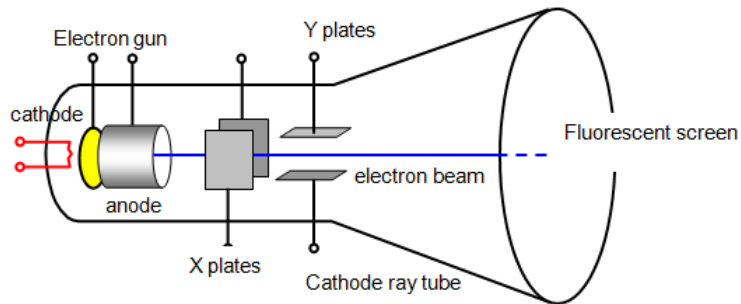
Heavy quarks seem to be **accelerated** by the **transverse** (color-)electric field



$$\frac{\Delta p}{\Delta t} = qE$$



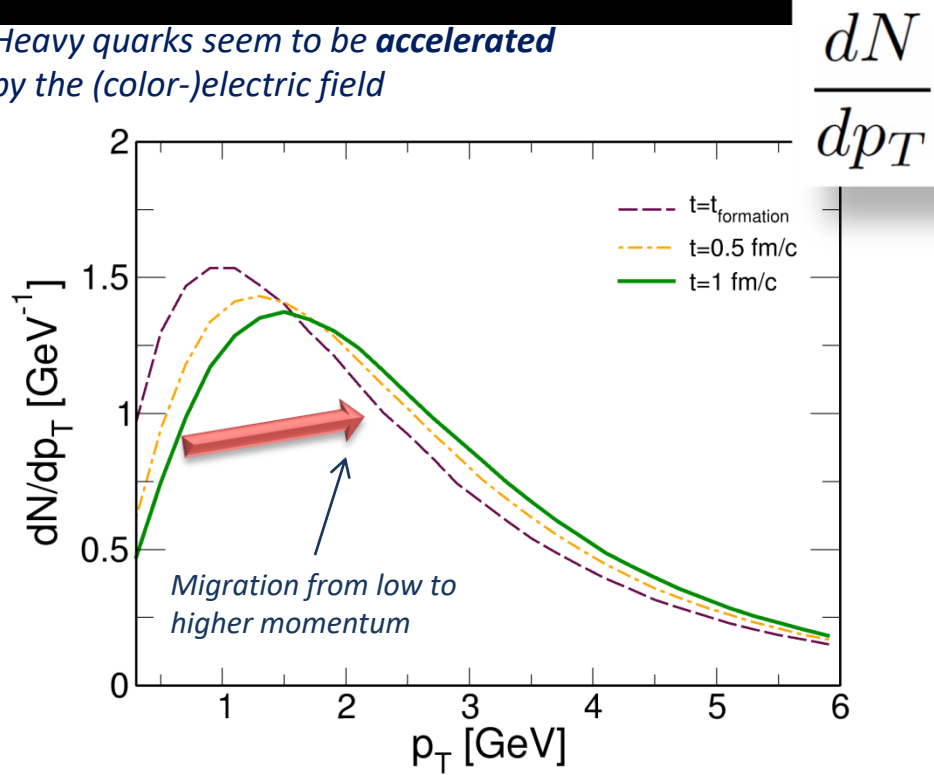
Why cathode tube?



Electrons are produced by the electron gun,
then accelerated by the electric field



Heavy quarks seem to be **accelerated**
by the (color-)electric field





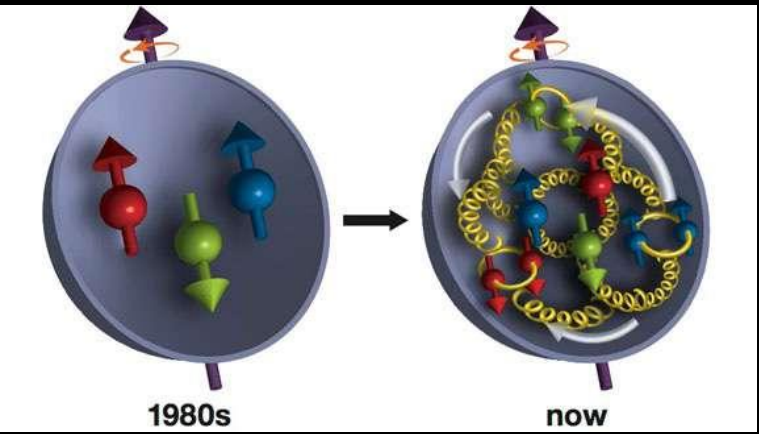
- Borrowing the Glasma picture, the **evolution** of the system after the collision can be probed by **heavy quarks observables**.
- (At least part of) the measured nuclear modification factor in p-Pb can be understood as the (anomalous) diffusion of heavy quarks in the evolving Glasma: **Cathode Tube Effect**.
- *J/ψ survival in the Glasma*
- *Diffusion coefficient of c and b quarks*
- *Anomalous vs standard diffusion*
- *$v_2, v_3, \dots$*
- *p-Pb, p-p, Pb-Pb*



Appendix

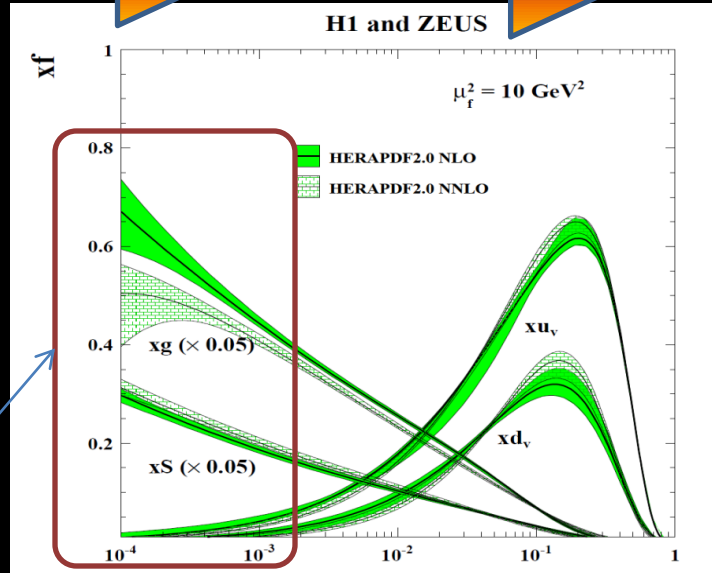
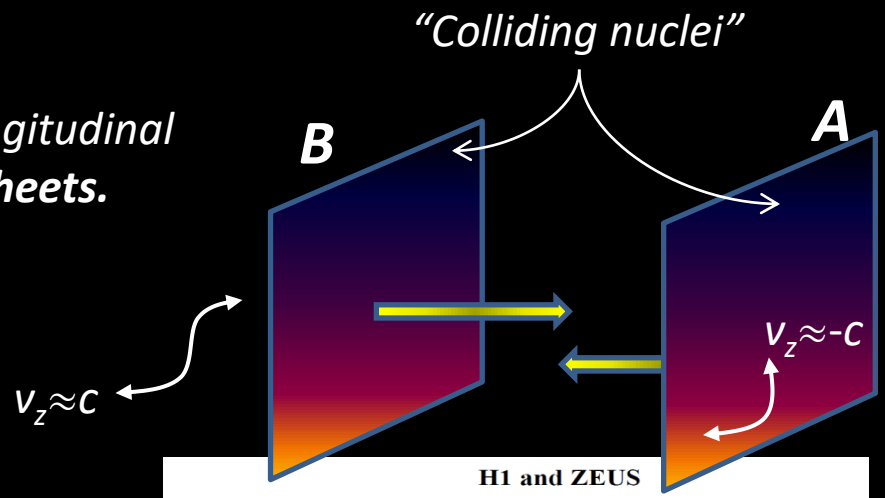
Cold nuclear matter effects: gluon saturation

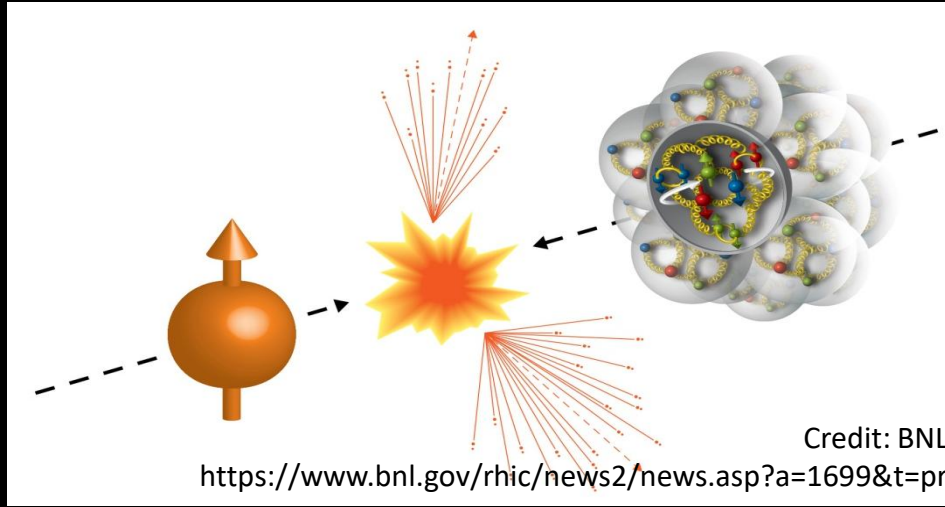
The two nuclei are Lorentz contracted along the longitudinal direction: in Lab frame they appear like two thin sheets.



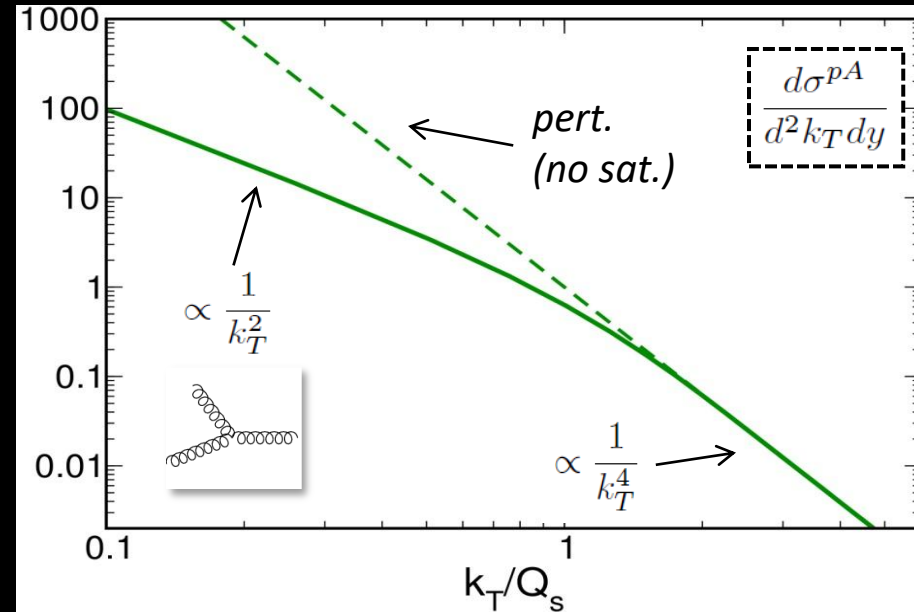
x : parton momentum/nucleon momentum

The small- x proton wave function is dominated by the sea of virtual gluons.





*High gluon density:
Gluon recombination*



Saturation

Gluon production is suppressed due to the abundance of the $2 \rightarrow 1$ processes.

Saturation scale, Q_s

Momentum scale at which saturation becomes important

McLerran and Venugopalan (1994)
and many others

Proton: GBW fit

Golec-Biernat and Wusthoff (1999)

$$Q_s^2 = Q_0^2 \left(\frac{x_0}{x} \right)^\lambda$$

$$x_0 = 4.1 \times 10^{-5}$$

$$\lambda = 0.277$$

$$Q_0 = 1 \text{ GeV}$$

$x \approx 10^{-2}$ RHIC energy: $Q_s \approx 0.46 \text{ GeV}$ [In agreement with Dumitru *et al.* (2013)]

$x \approx 10^{-4}$ LHC energy: **$Q_s \approx 0.80 \text{ GeV}$**

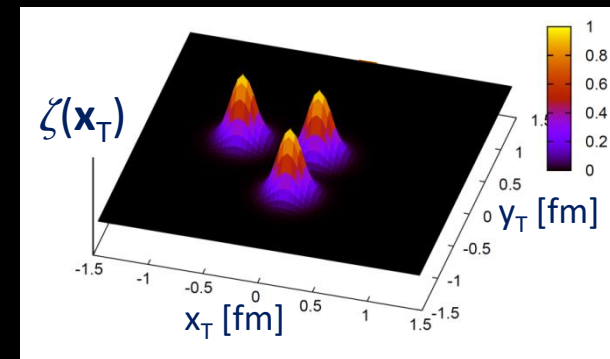
valence quarks

assumption

$$\langle g^2 \mu(\mathbf{x}_T) \rangle \stackrel{\text{valence quarks}}{=} g^2 \mu \cdot \langle \zeta(\mathbf{x}_T) \rangle \stackrel{\text{assumption}}{=} Q_s$$

Averaging over events:

$$g^2 \mu \approx 1.6 \text{ GeV @ } 5.02 \text{ TeV}$$



Pb: GBW fit again

Freund *et al.* (2002)
 Albacete *et al.* (2004)
 Armesto *et al.* (2005)
 Kowalski *et al.* (2006)
 Kowalski *et al.* (2008)
 Lappi (2008)

$$Q_s^2 = f(A) Q_0^2 \left(\frac{x_0}{x} \right)^\lambda$$

$$f(A) = A^{1/3}, \text{ naive}$$

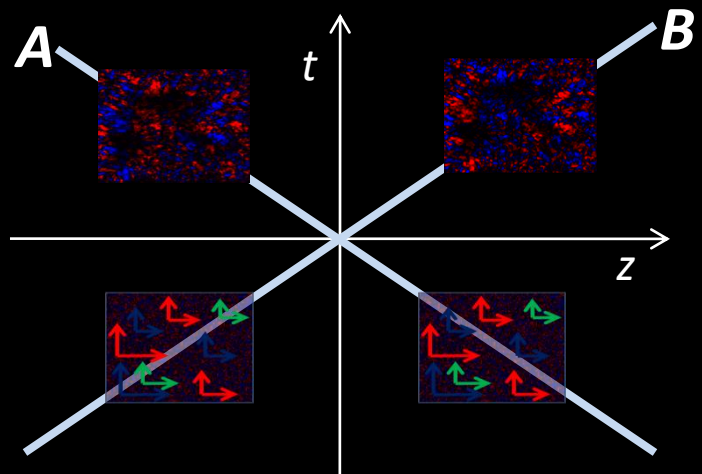
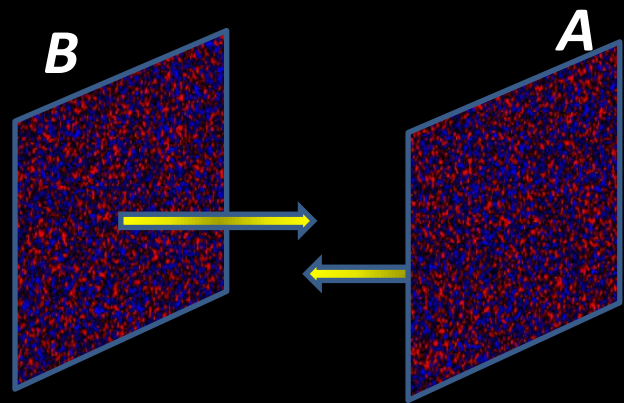
$$f(A) = cA^{1/3} \log(A), \text{ IP-Sat}$$

<i>naive</i>	{	$x \approx 10^{-2}$ RHIC energy: $Q_s \approx 1.7 \text{ GeV}$, $g^2\mu \approx 3 \text{ GeV}$
$x \approx 10^{-4}$ LHC energy: $Q_s \approx 3 \text{ GeV}$, $g^2\mu \approx 5.3 \text{ GeV}$		
<i>IP-Sat</i>	{	$x \approx 10^{-2}$ RHIC energy: $Q_s \approx 1.1 \text{ GeV}$, $g^2\mu \approx 2 \text{ GeV}$
$x \approx 10^{-4}$ LHC energy: $Q_s \approx 1.90 \text{ GeV}$, $g^2\mu \approx 3.3 \text{ GeV}$		

We consider

$g^2\mu_{\text{Pb}} \approx$ in the range $3.3 \text{ GeV} - 5.3 \text{ GeV}$ @ 5.02 TeV

Building up the Glasma fields



$$-\partial_{\perp}^2 \Lambda^{(A)}(\mathbf{x}_T) = \rho^{(A)}(\mathbf{x}_T)$$

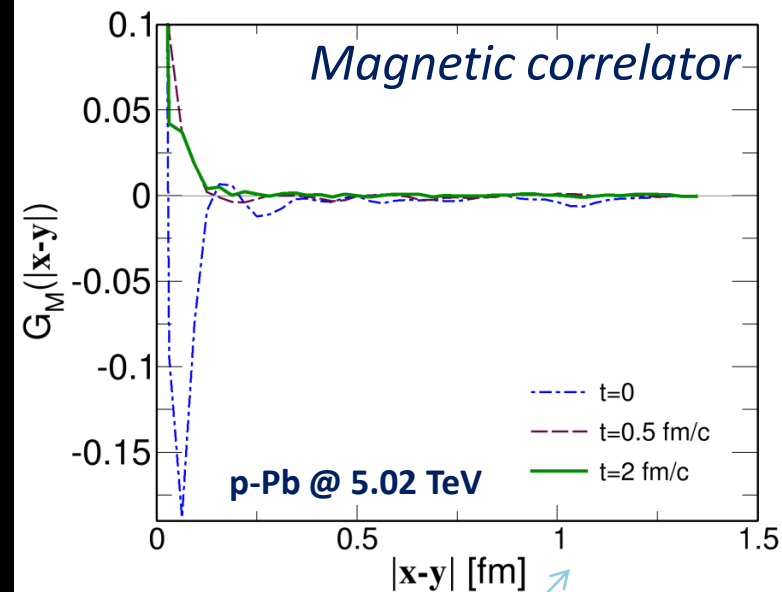
$$V^{\dagger}(\mathbf{x}_T) = e^{i\Lambda^{(A)}(\mathbf{x}_T)}, \quad W^{\dagger}(\mathbf{x}_T) = e^{i\Lambda^{(B)}(\mathbf{x}_T)}$$

$$\alpha_i^{(A)} = iV\partial_i V^{\dagger}, \quad \alpha_i^{(B)} = iW\partial_i W^{\dagger}$$

Glasma color fields

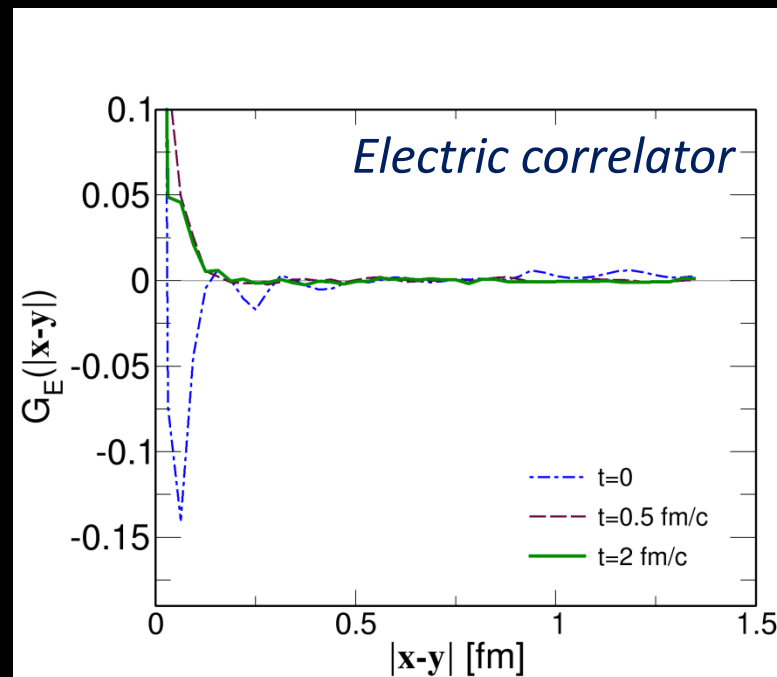
$$E^z = i \sum_{i=x,y} \left[\alpha_i^{(B)}, \alpha_i^{(A)} \right],$$

$$B^z = i \left(\left[\alpha_x^{(B)}, \alpha_y^{(A)} \right] + \left[\alpha_x^{(A)}, \alpha_y^{(B)} \right] \right)$$

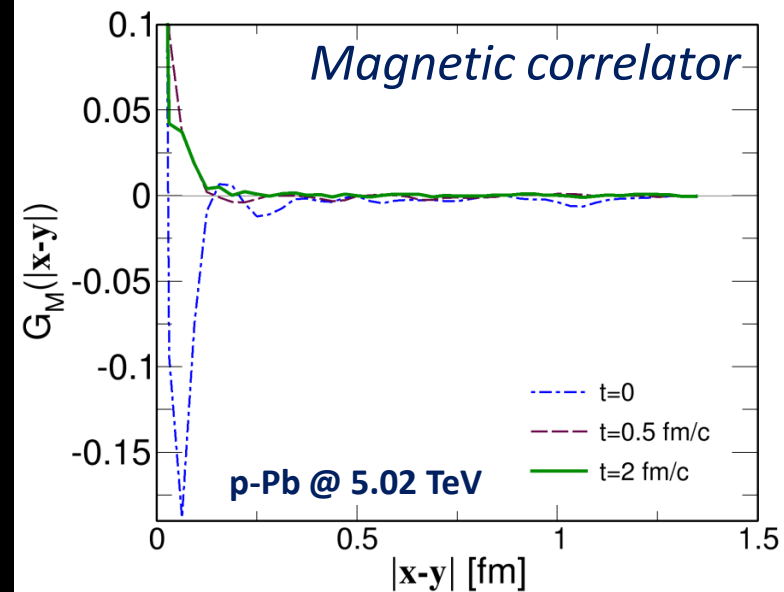


Initial time:
Anticorrelation

$t \approx 0.5$ fm/c:
Correlation length ≈ 0.2 fm
Correlation domains

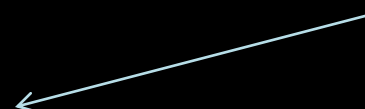


M. Ruggieri, *in preparation*



Correlation domains: p-Pb versus Pb-Pb

p-Pb



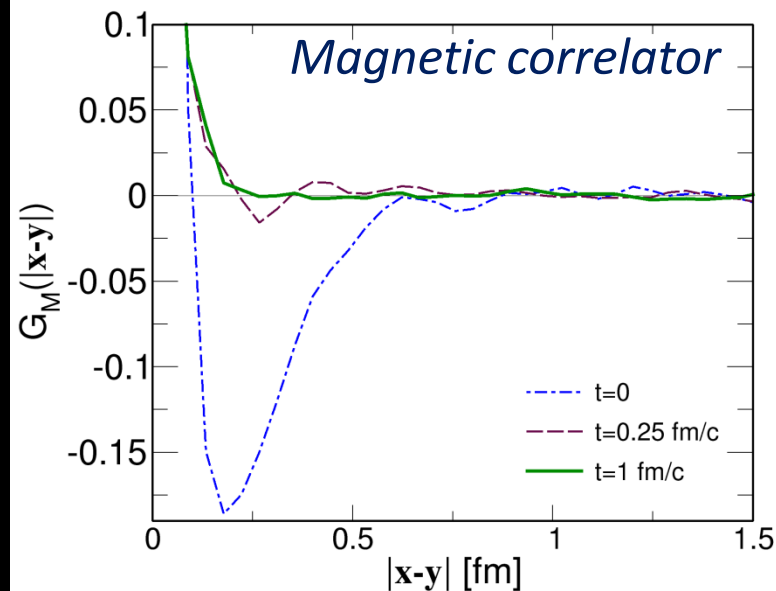
Pb-Pb



In agreement with Dumitru *et al.* (2014)
Dumitru *et al.* (2013)

Pb-Pb @ 5.02 TeV

M. Ruggieri *et al.* (2017)



Fluctuations appear at the *next-to-leading order in the QCD coupling*:

- *Longitudinal electric fields* (E_z): break longitudinal invariance.
- *Transverse electric fields* (E_x, E_y): added on the top of the longitudinal Glasma fields.

$$\langle \xi_i^a(\mathbf{x}_T) \xi_j^b(\mathbf{y}_T) \rangle = \delta_{ab} \delta_{ij} \delta^{(2)}(\mathbf{x}_T - \mathbf{y}_T);$$

$$g^2 \mu \langle F(z) F(z') \rangle = \Delta^2 \delta(z - z').$$

No correlation in:

- *Transverse plane*
- *Longitudinal direction*

Fluctuations of the electric field:

$$\begin{aligned} \delta E_i^a(\mathbf{x}_T, z) &= \partial_z F(z) \xi_i^a(\mathbf{x}_T), \\ \delta E_z^a(\mathbf{x}_T, z) &= -F(z) D_i \xi_i^a(\mathbf{x}_T). \end{aligned}$$

White noise

See:

Romatschke and Venugopalan (2006)

Ohnishi *et al.* (2014)

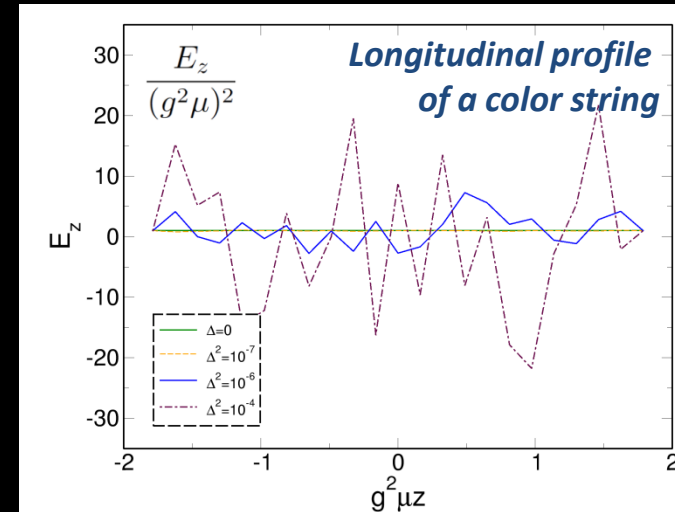
Fukushima and Gelis (2012)

Fukushima (2013)

Berges and Schlichting (2013)

A more rigorous treatment is also possible:

Epelbaum and Gelis (2013)



Fluctuations appear at the *next-to-leading order in the QCD coupling*:

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No correlation in:

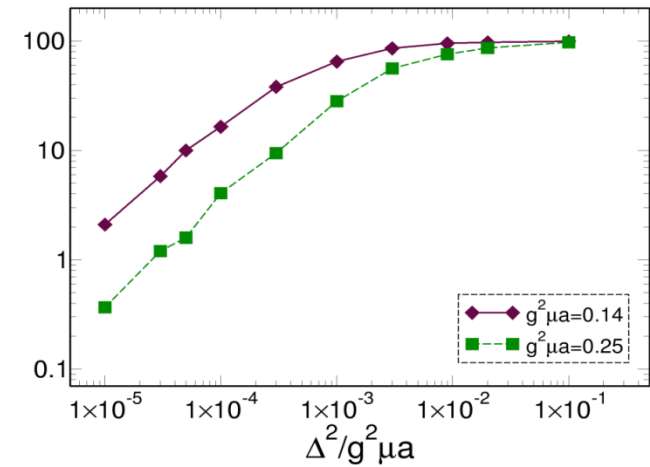
- *Transverse plane*
- *Longitudinal direction*

With these the fluctuations of the electric field are computed as:

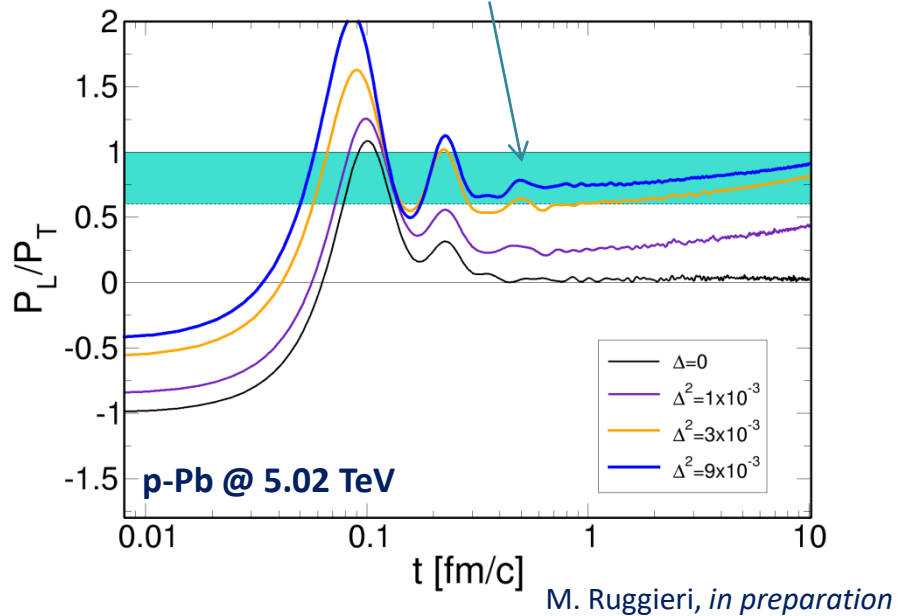
$$\begin{aligned} \delta E_i^a(\mathbf{x}_T, z) &= \partial_z F(z) \xi_i^a(\mathbf{x}_T), \\ \delta E_z^a(\mathbf{x}_T, z) &= -F(z) D_i \xi_i^a(\mathbf{x}_T). \end{aligned}$$

$$\Delta \sim g$$

% of energy carried by fluctuations



Fair isotropy band: $P_L/P_T \geq 0.6$



The evolving Glasma: pressures

$$P_T = \frac{E_z^a E_z^a + B_z^a B_z^a}{2},$$

$$P_L = -P_T + \frac{E_T^a E_T^a + B_T^a B_T^a}{2},$$

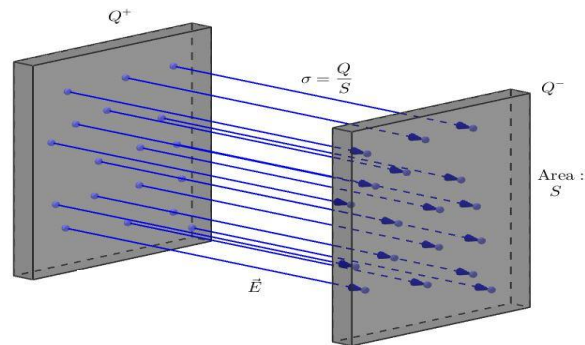
Romatschke and Venugopalan (2006)
 Berges (2012)
 Epelbaum and Gelis (2012)
 Fukushima (2014)
 M. R. et al. (2017)

Negative pressure

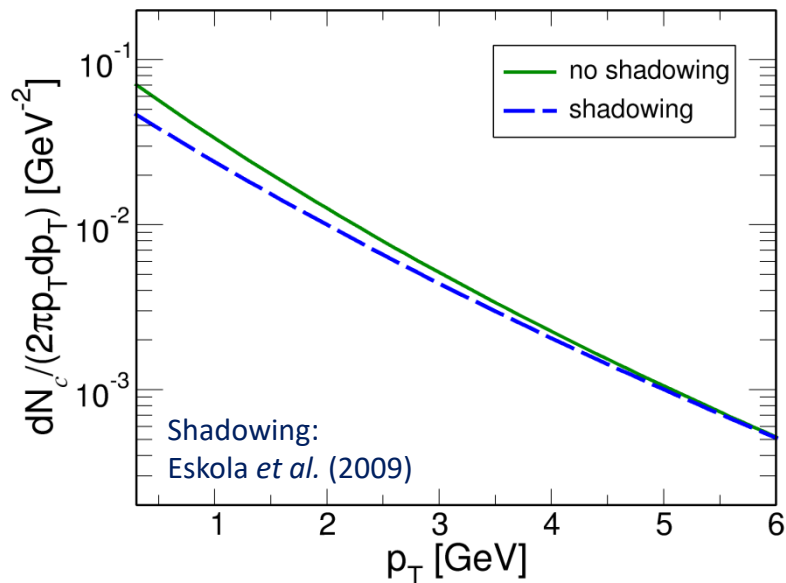
Attraction among the two nuclei.

Energy has to be given *by the environment* to allow for the expansion of the system.

In realistic collisions *this energy is given by the kinetic energy of the colliding nuclei.*

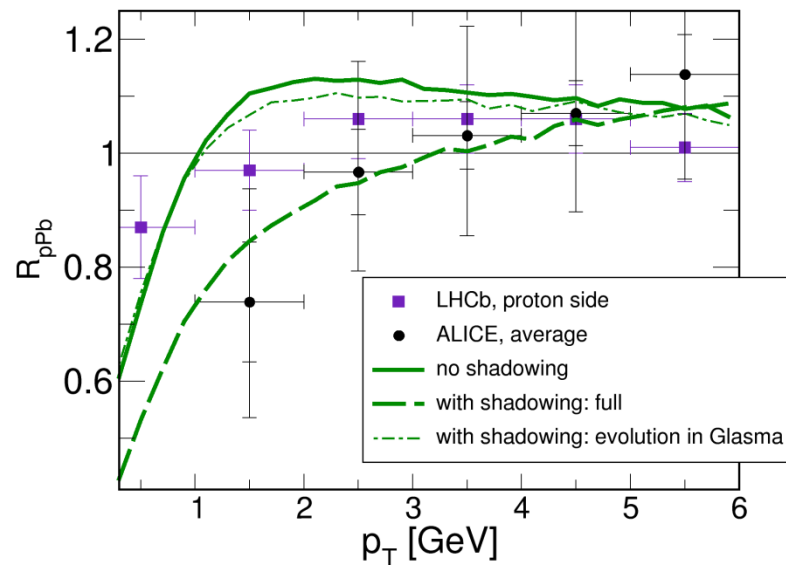


Nuclear shadowing does not affect the cathod tube

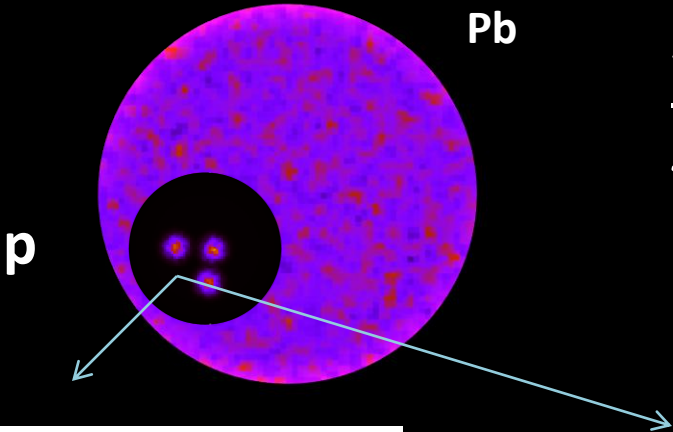
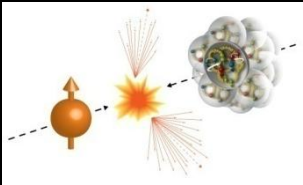


$$R_{pPb} = \frac{(dN/d^2p_T)_{\text{final}}}{(dN/d^2p_T)_{pQCD}}$$

p-Pb @ 5.02 TeV

M. Ruggieri and S. K. Das, *in preparation*

Implementation of p-Pb



Valence quarks
Sources of the $x \approx 1$ gluon fields

Mantysaari *et al.* (2017)
 Mantysaari and Schenke (2016)
 Schlichting and Schenke (2014)

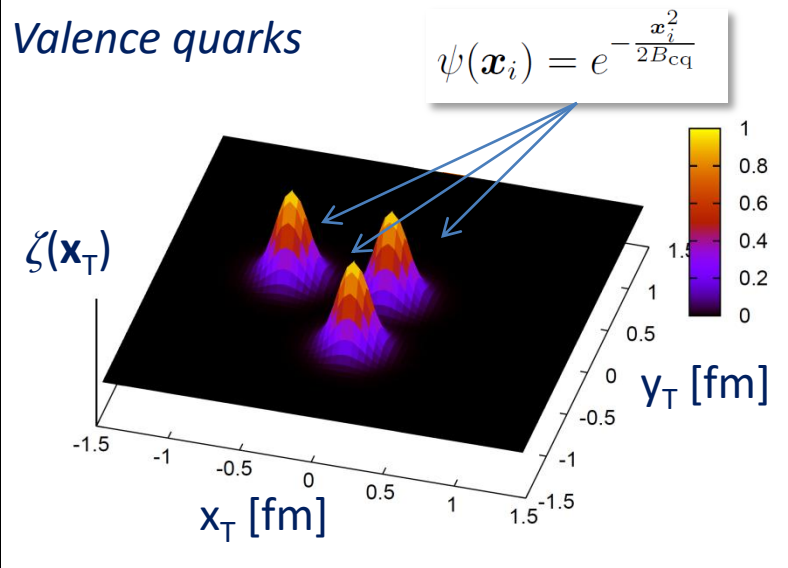
$$\langle \rho_a(\mathbf{x}_T) \rho_b(\mathbf{y}_T) \rangle = (g^2 \mu(\mathbf{x}_T))^2 \delta_{ab} \delta^2(\mathbf{x}_T - \mathbf{y}_T)$$

$$g^2 \mu(\mathbf{x}_T) = g^2 \mu \cdot \zeta(\mathbf{x}_T)$$

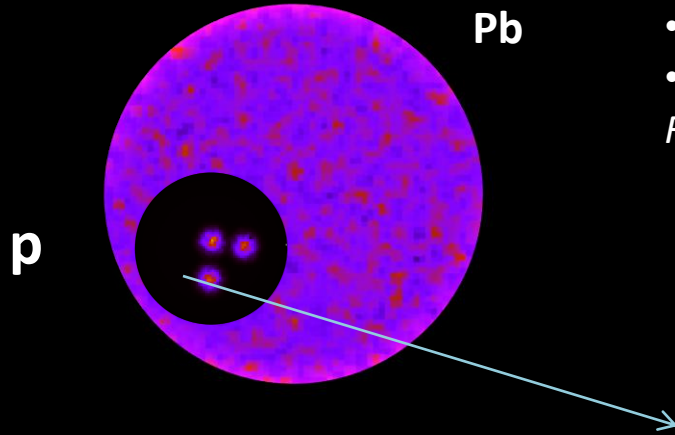
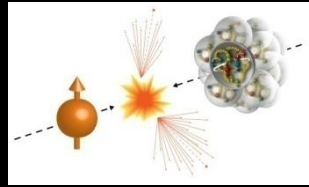
$$\psi(\mathbf{x}_i) = e^{-\frac{\mathbf{x}_i^2}{2B_{cq}}}$$

$$\zeta(\mathbf{x}_T) = \frac{1}{3} \sum_{i=1}^3 e^{-\frac{(\mathbf{x}_T - \mathbf{x}_i)^2}{2B}}$$

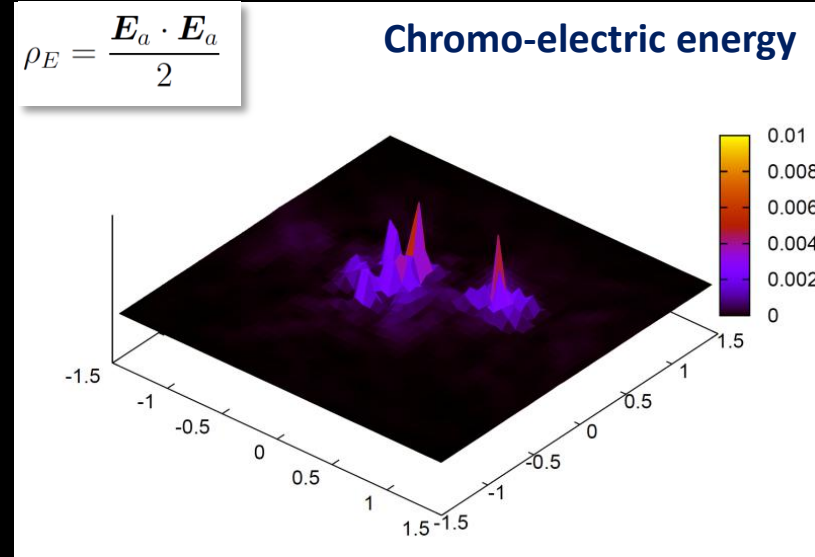
$B_{cq} = 3 \text{ GeV}^{-2} \approx 0.12 \text{ fm}^2$
 $B = 0.3 \text{ GeV}^{-2} \approx 0.01 \text{ fm}^2$



Glasma in p-Pb collisions



- Transverse area $\approx (4 \text{ fm})^2$
 - Overlap of color charges of proton and nucleus
- Finite size system*

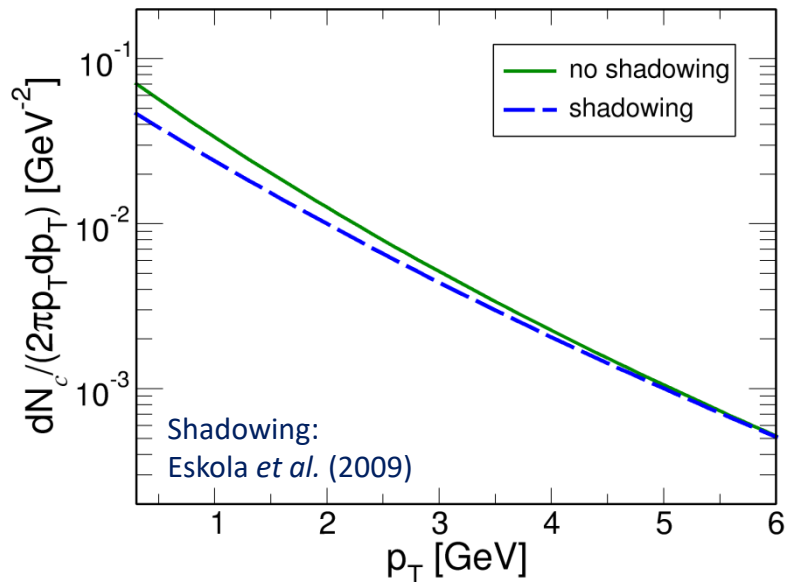


Energy distribution depends on the model.

It concentrates around the peaks of the color charges (valence + sea) in the proton.



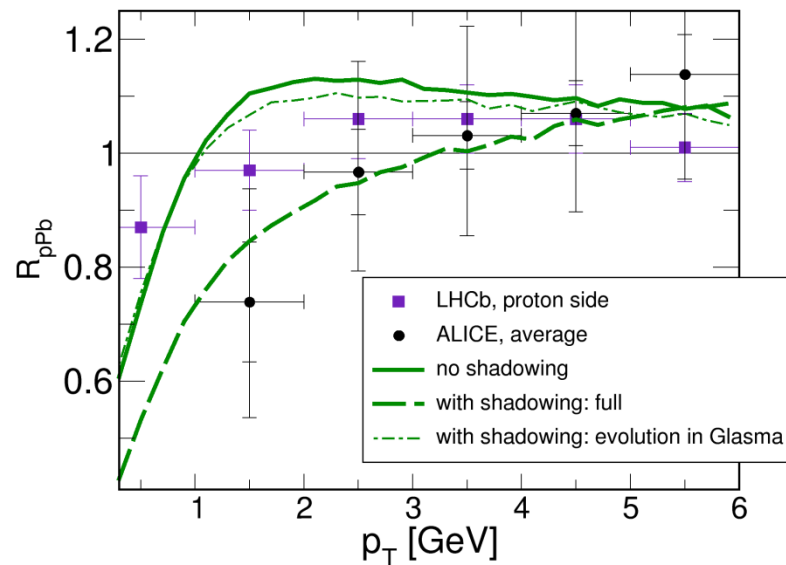
Nuclear shadowing does not affect the cathode tube



$$R_{\text{pPb}} = \frac{(dN/d^2p_T)_{\text{final}}}{(dN/d^2p_T)_{\text{pQCD}}}$$

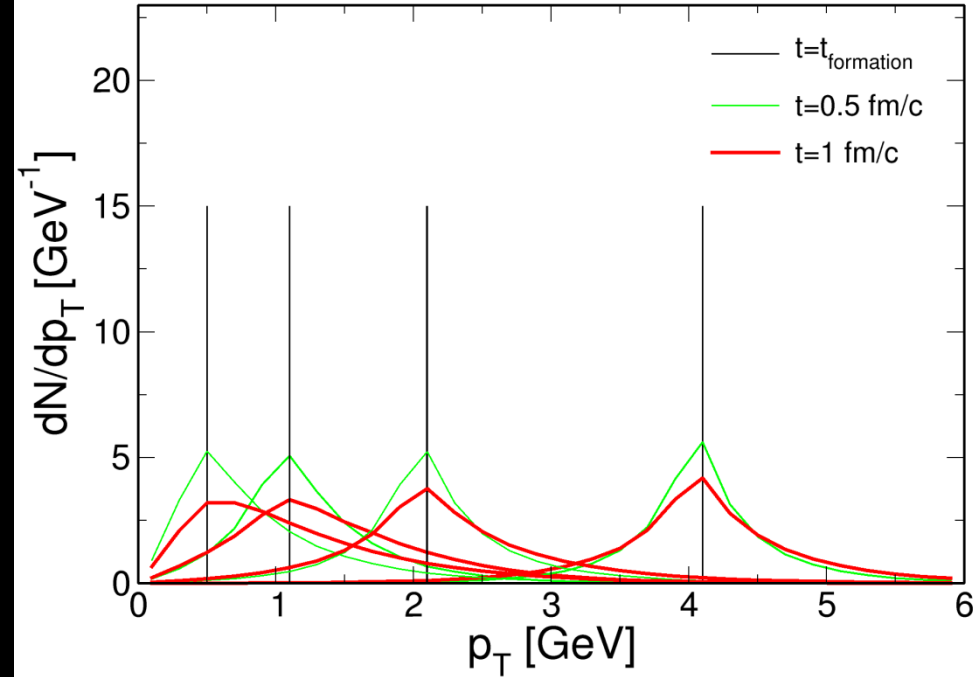
p-Pb @ 5.02 TeV

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HQs in Glasma

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Standard Brownian motion:
Random walk with fixed probability

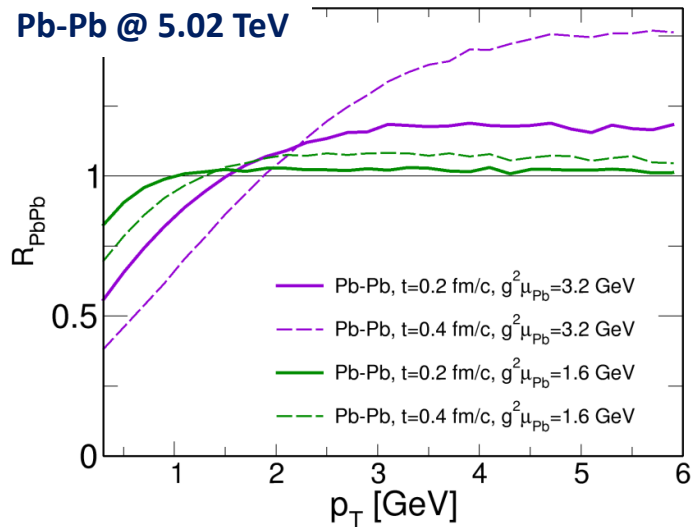
$$\langle p^2 - \langle p \rangle^2 \rangle = 2Dt$$

Anomalous diffusion:
Random walk in disordered structures

$$\langle p^2 - \langle p \rangle^2 \rangle = 2Dt^{1/d_w}$$

For HQs in Glasma:

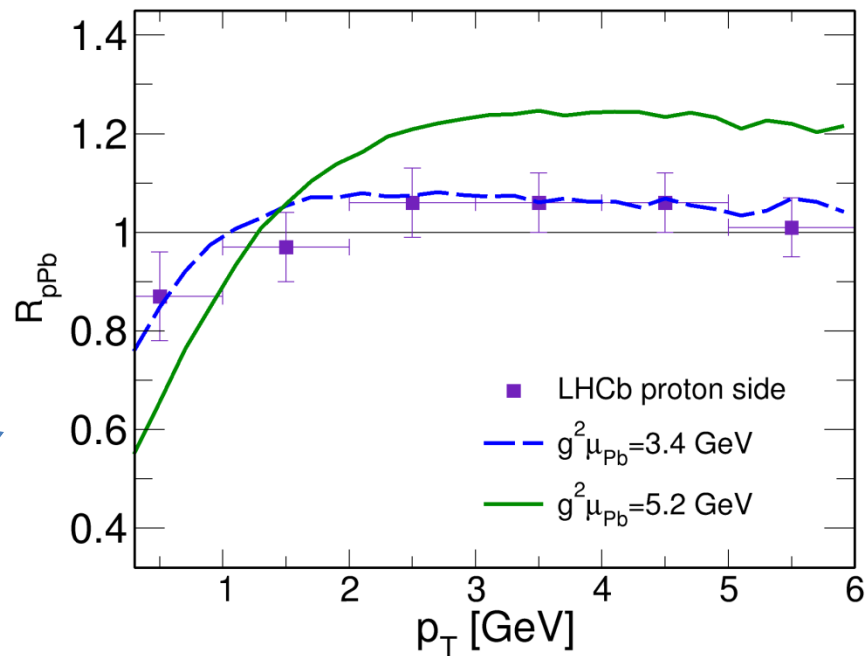
$$d_w \approx 1.66$$

Pb-Pb @ 5.02 TeV

Interaction with Glasma affects the spectrum of c in Pb-Pb

p-Pb @ 5.02 TeV

M. Ruggieri and S. K. Das, 1805.09617



$$R_{\text{pPb}} = \frac{(dN/d^2 p_T)_{\text{final}}}{(dN/d^2 p_T)_{\text{pQCD}}}$$