

# Commissioning of the ATLAS Pixel Detector with Cosmic Ray Data

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on behalf of the ATLAS Collaboration



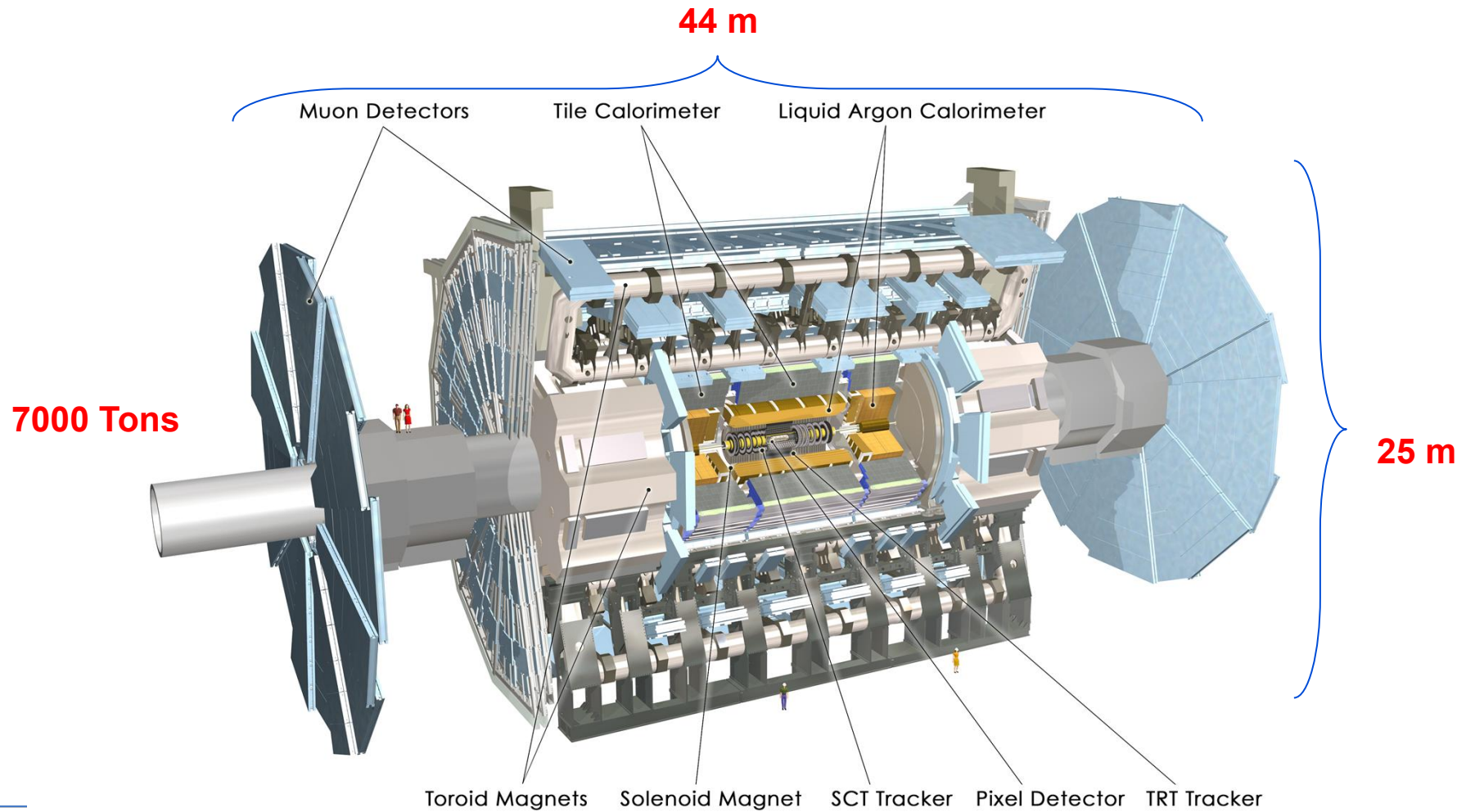
# Outline

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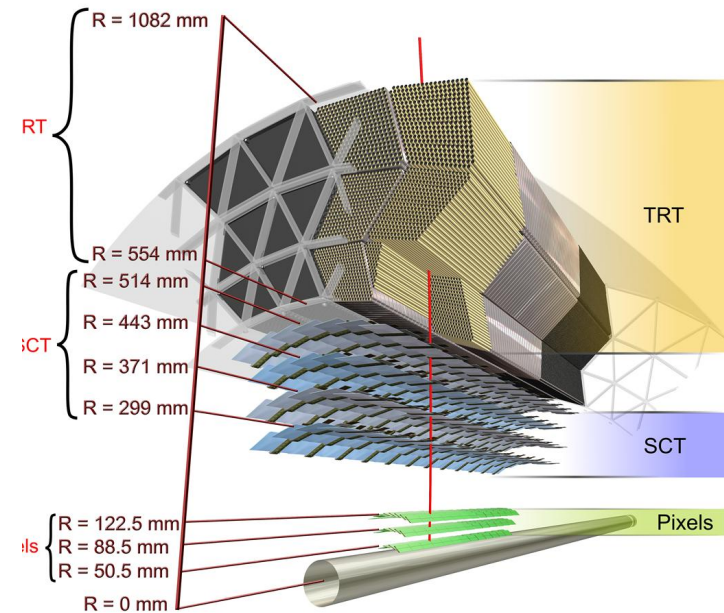
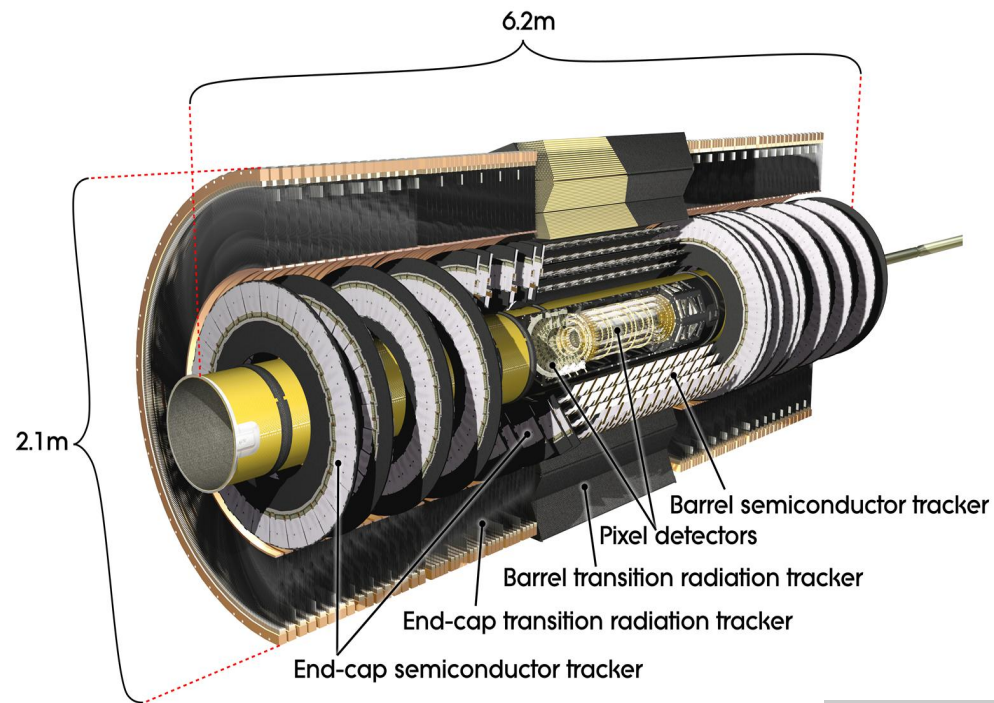
- ATLAS and its Inner Detector
- The ATLAS Pixel Detector and its commissioning
  - ▣ Calibration and Cosmic ray data taking
    - Threshold, noise and masked pixels
    - Time-over-Threshold and timing
    - Resolution, efficiency and noise occupancy
    - Lorentz angle
- Status and expectations
  - ▣ Readiness for collisions and long term operations



# The ATLAS Detector



# The ATLAS Inner Detector



- 2 T solenoidal magnetic field
- Acceptance  $|\eta| < 2.5$   
(transition radiation tracker  $|\eta| < 2$ )

- Momentum resolution  
 $\sigma(p_T)/p_T = 0.05\% p_T [\text{GeV}/c] \oplus 1\%$
- Impact parameter resolution ( $0.25 < |\eta| < 0.5$ )  
 $\sigma(d_0) = 10 \mu\text{m} \oplus 140 \mu\text{m} / p_T [\text{GeV}/c]$





# The ATLAS Pixel Detector

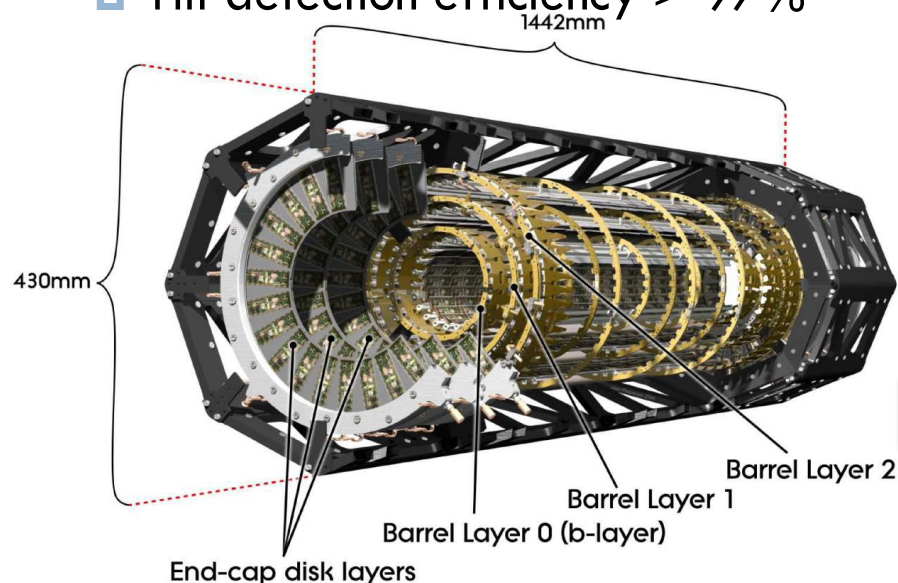
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## □ Requirements:

- Position resolution in  $r\phi$ -direction  $< 15 \mu\text{m}$
- 3 track points for  $|\eta| < 2.5$
- Time resolution  $< 25 \text{ ns}$
- Hit detection efficiency  $> 97\%$

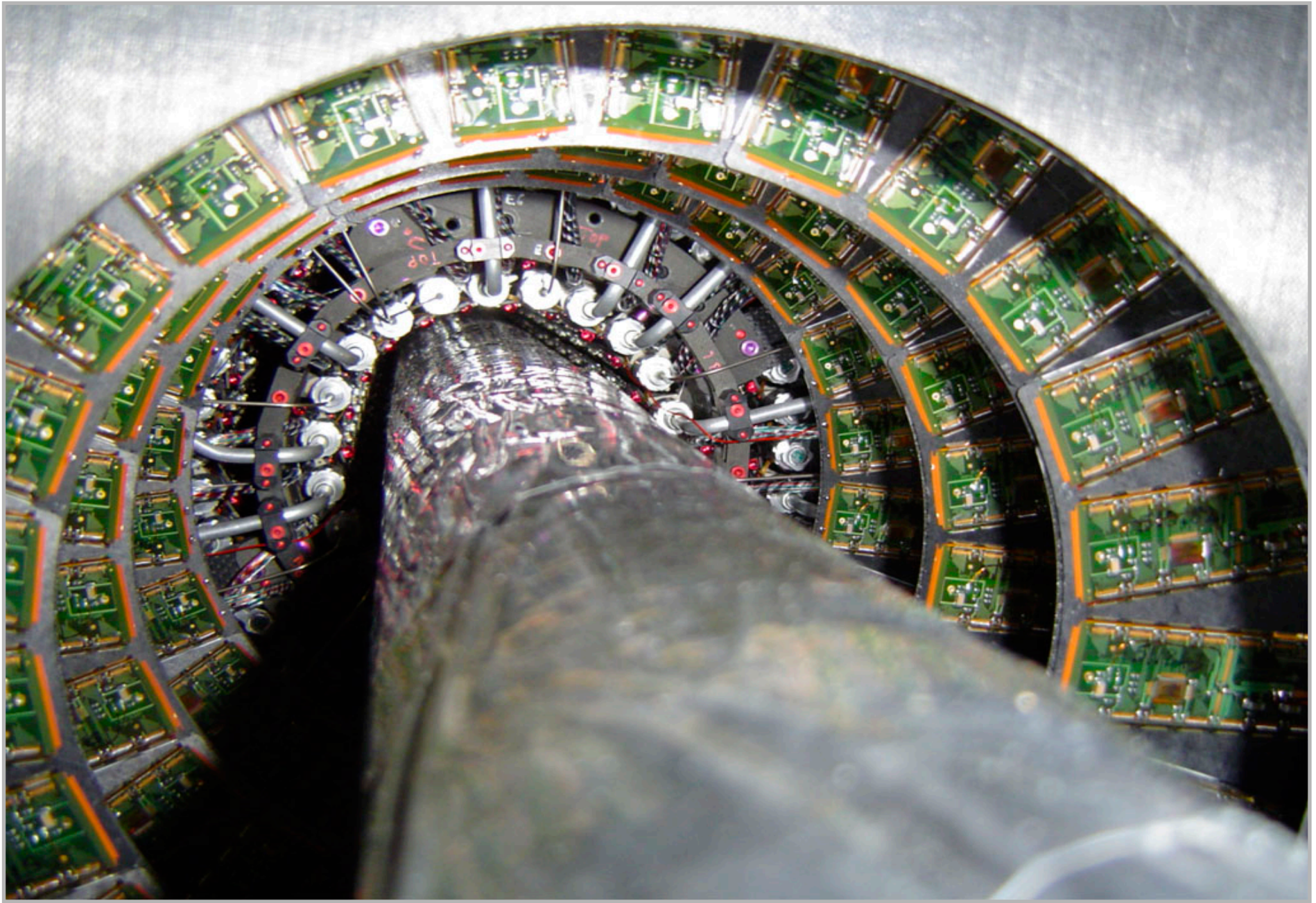
## □ Basic Properties:

- 1744 Pixel Modules on three barrel layers and 2 x 3 disks
- 80M readout channels
- Innermost layer at 5 cm
  - Radiation tolerance  
 $500 \text{ kGy} / 10^{15} \text{ 1 MeV n}_{\text{eq}}\text{cm}^{-2}$
- Evaporative  $\text{C}_3\text{F}_8$  cooling integrated in local support structure  $\rightarrow$  Module temperature below  $0 \text{ }^\circ\text{C}$



How close to the beam pipe? See next slide





B. Di Girolamo - Commissioning of the ATLAS Pixel Detector - RD09 30 September 2009

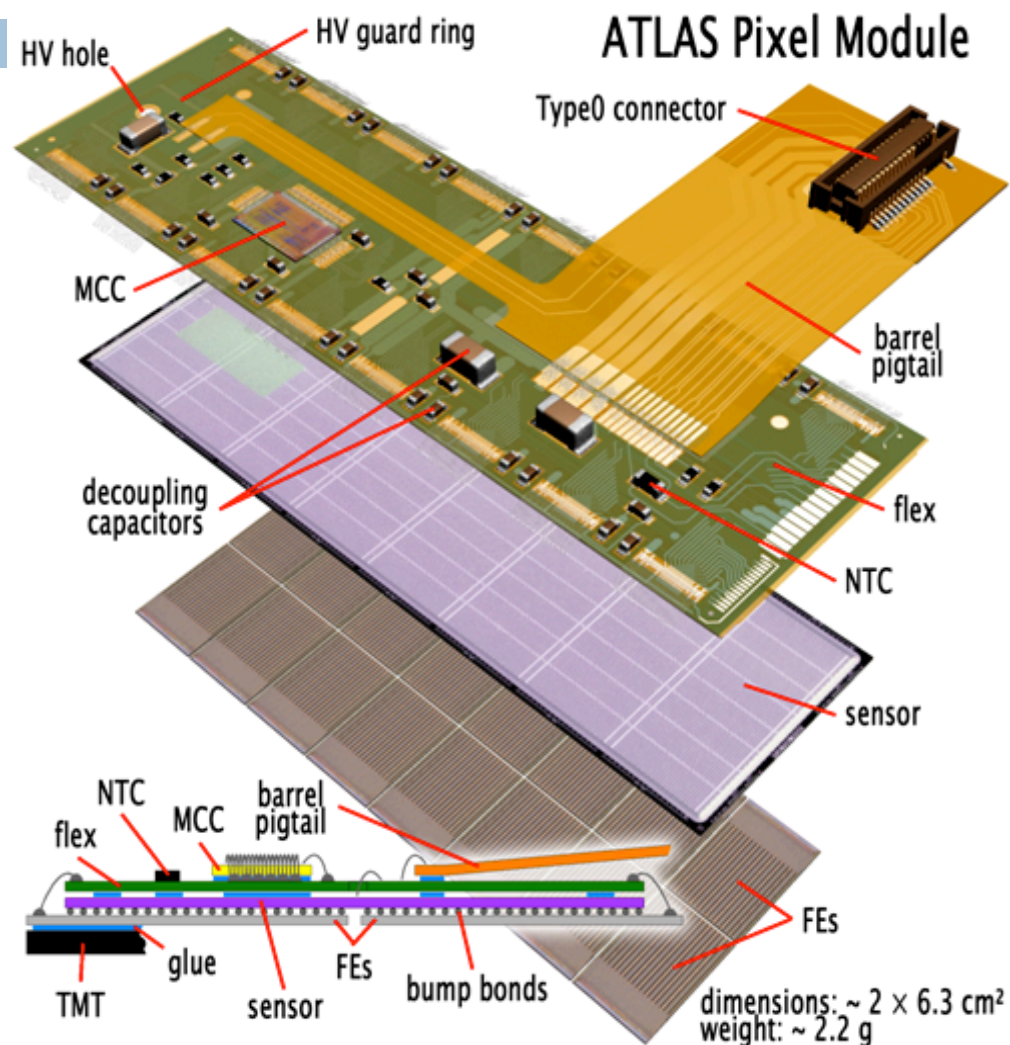




# The ATLAS Pixel Module

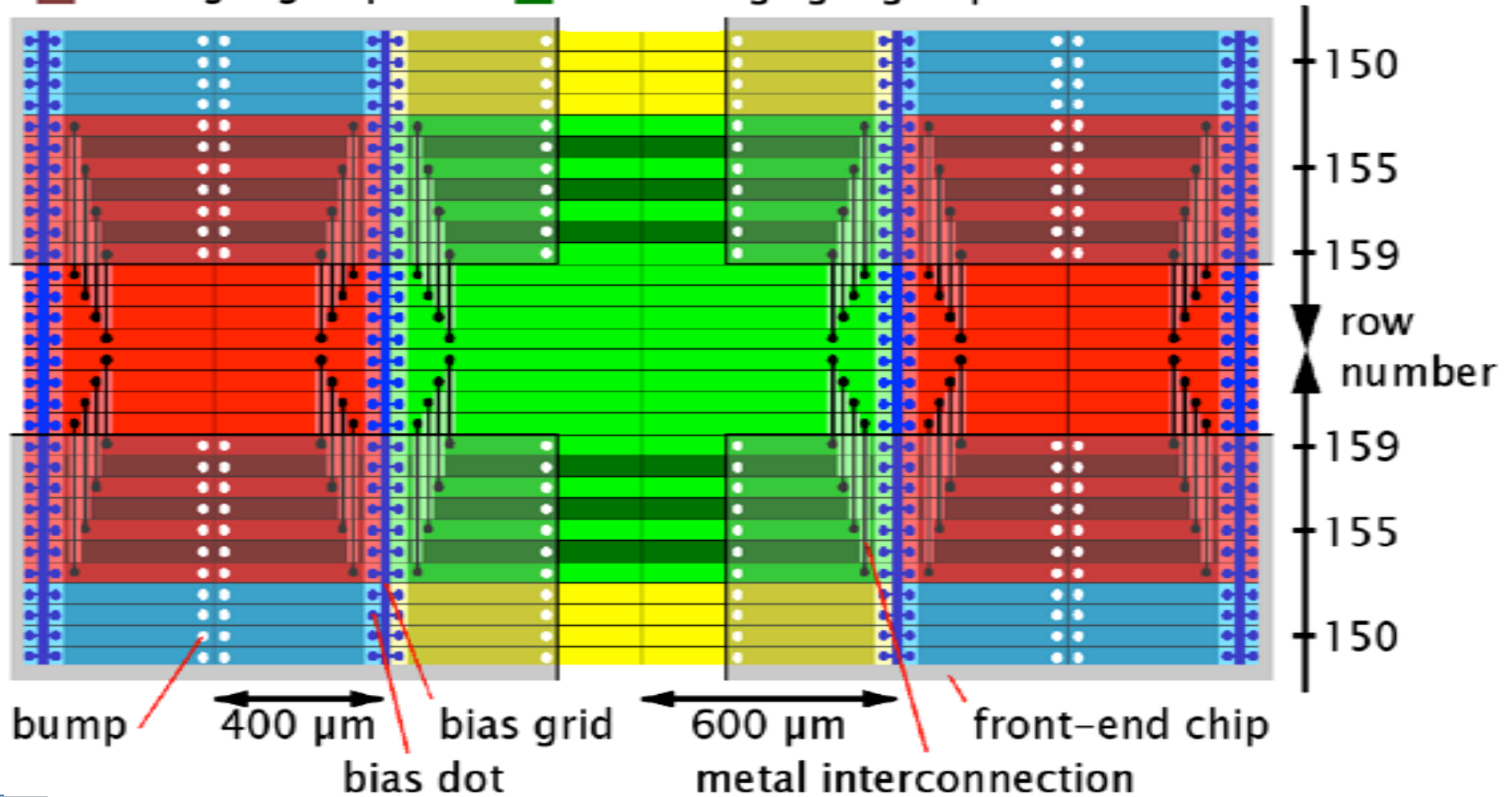
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- Sensor
  - ▣ 250  $\mu\text{m}$  thick n-on-n Si sensor
  - ▣ 47232 (328 x 144) Pixels (46080 read channels)
  - ▣ Typical pixel size 50 x 400  $\mu\text{m}$  (50 x 600  $\mu\text{m}$  pixels in gaps between FE chips)
  - ▣ Bias voltage 150 – 600 V
  
- Readout
  - ▣ 16 FE Chips with 2880 pixels each
  - ▣ Pulse height measured by means of Time over Threshold
  - ▣ Zero suppression in the FE chips, MCC chip builds module event
  - ▣ Data transfer 40 to 160 MHz depending on layer (occupancy)



# Inter-chip Region

- pixel
- ganged pixel
- inter-ganged pixel
- long pixel
- long+ganged pixel
- inter-long+ganged pixel





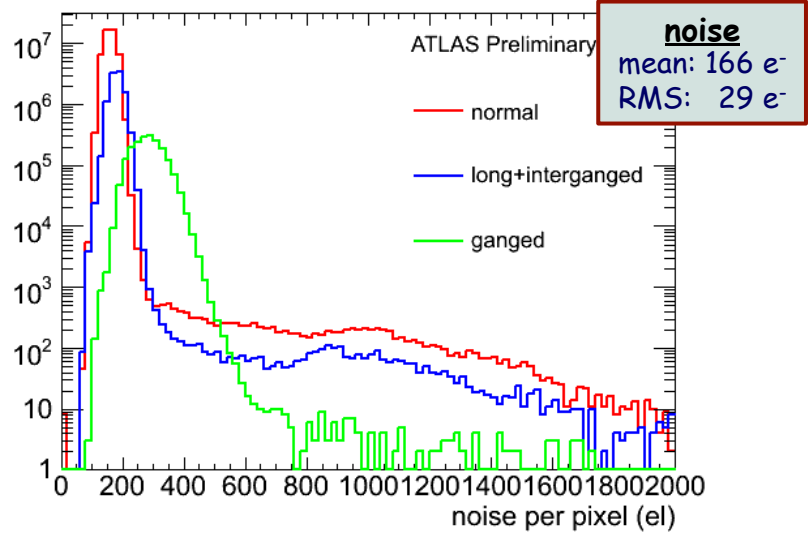
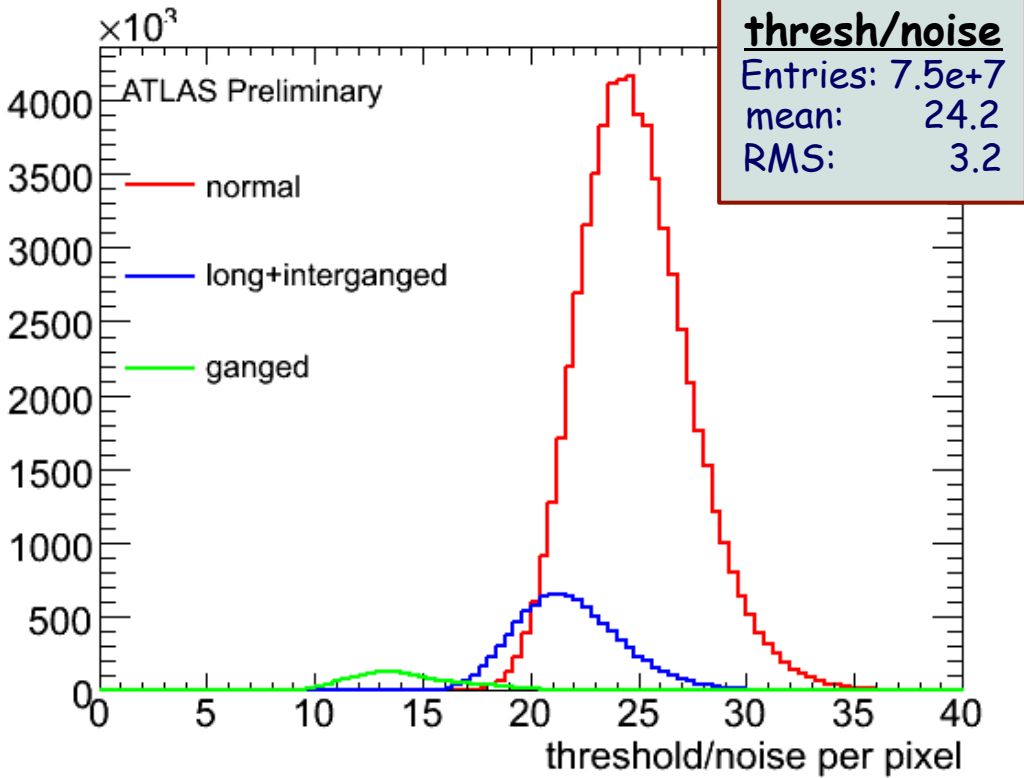
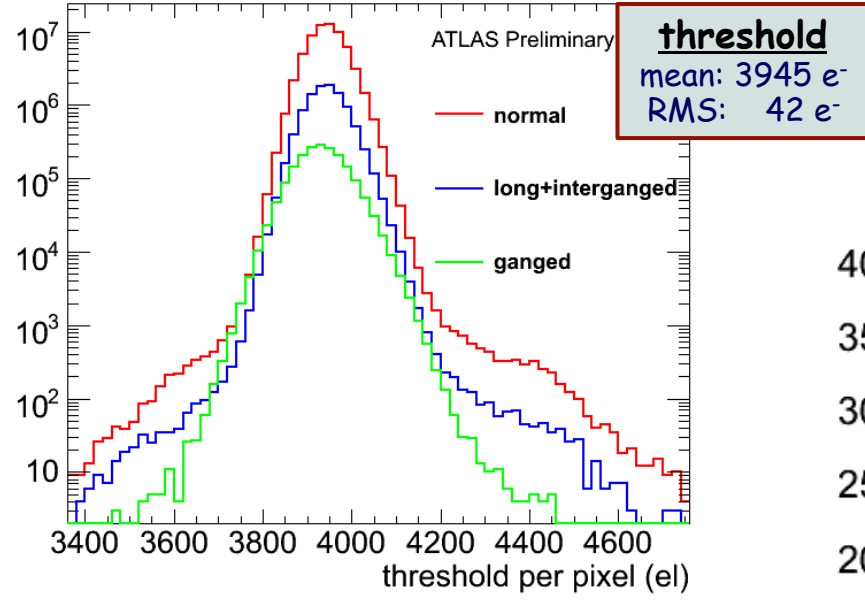
# The Pixel Detector commissioning

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- August-December 2008
  - ▣ Functionality checks, calibrations and cosmic ray data
  - ▣ 240 k tracks with field off, 190 k tracks with field on
- May-July 2009
  - ▣ Short calibration period and cosmic ray data
  - ▣ 90 k tracks with field off, 180 k tracks with field on
- Restarted mid-August 2009
  - ▣ 5 weeks for calibration and very soon in continuous cosmic ray data taking until beam comes



# Threshold and Noise

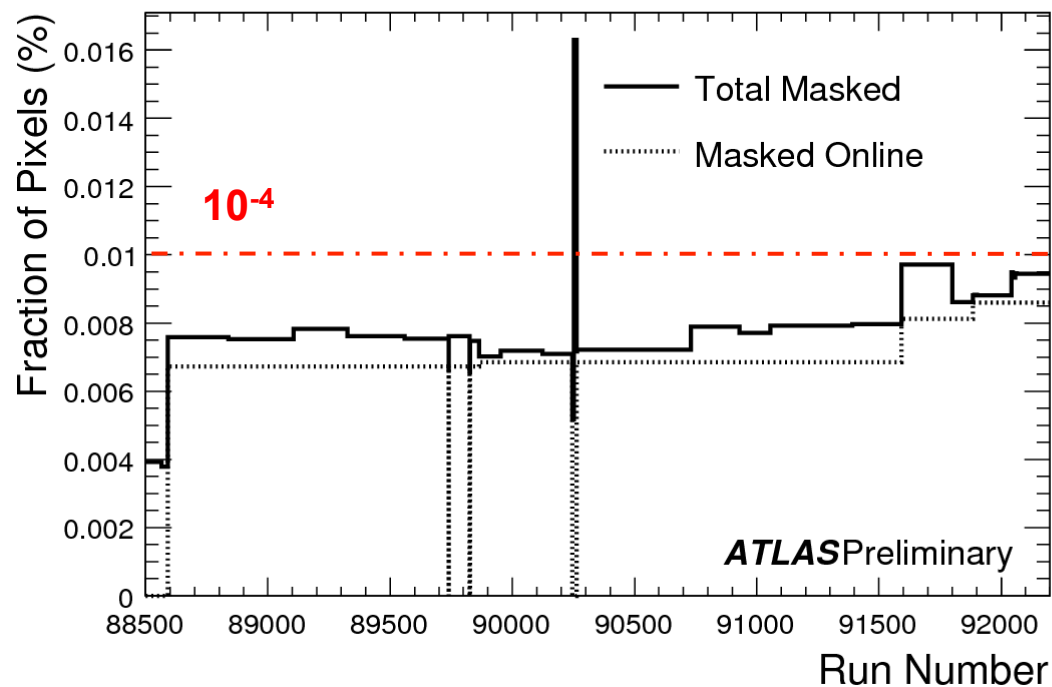
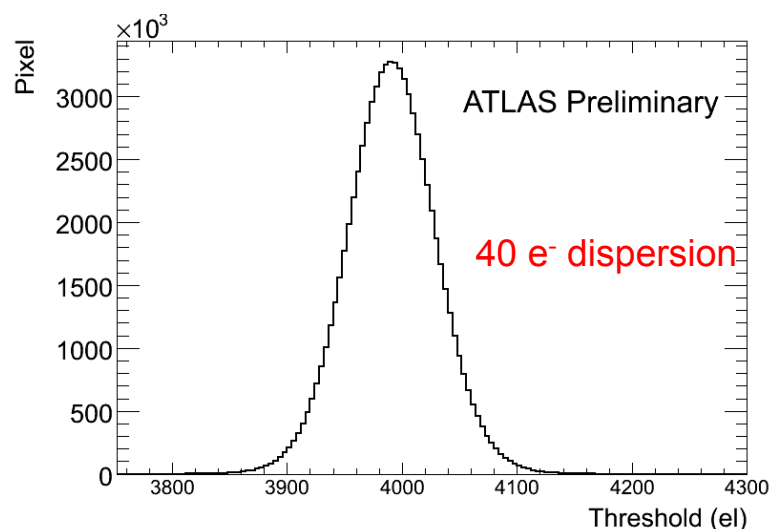


**75M pixels in the plots!**  
 Noise approx. 200 e<sup>-</sup> for most pixels,  
 slightly higher for special pixels  
**Threshold/Noise approx. 25 for normal pixels**



# Threshold and Noise: Masked Pixels

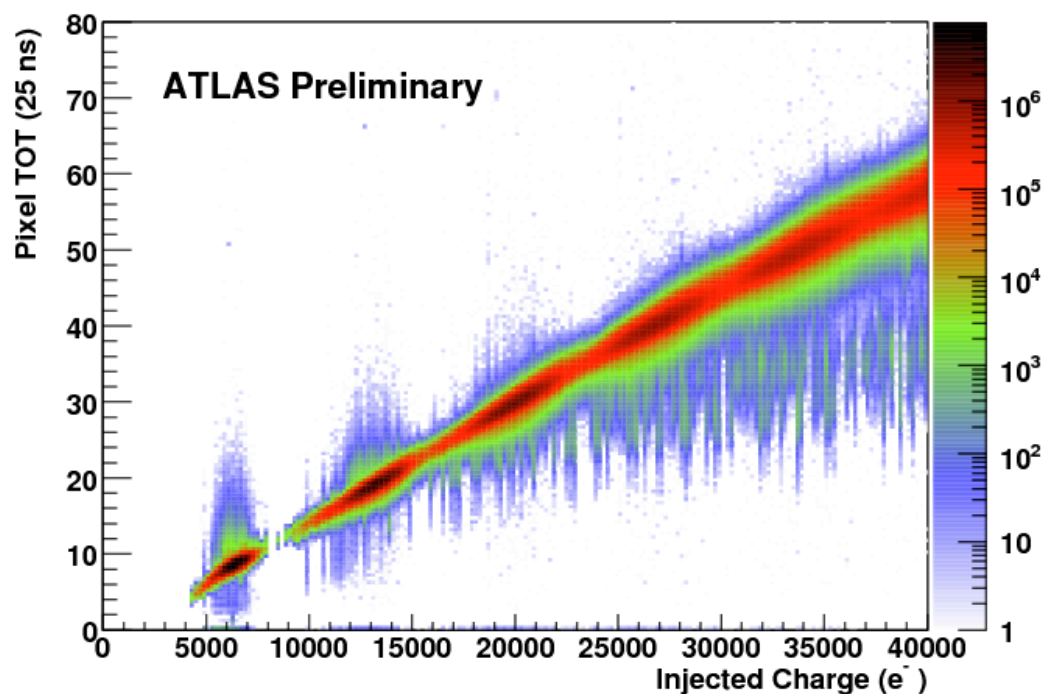
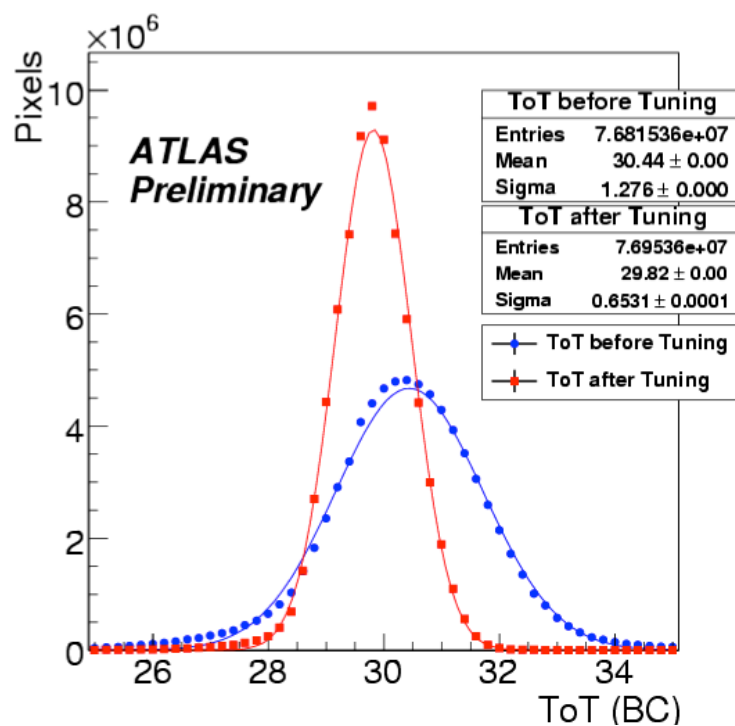
- Threshold setting: 4000  $e^-$
- Threshold tuned pixel by pixel, threshold dispersion  $\sim 40 e^-$
- Fraction of masked pixel  $\sim 0.01\%$





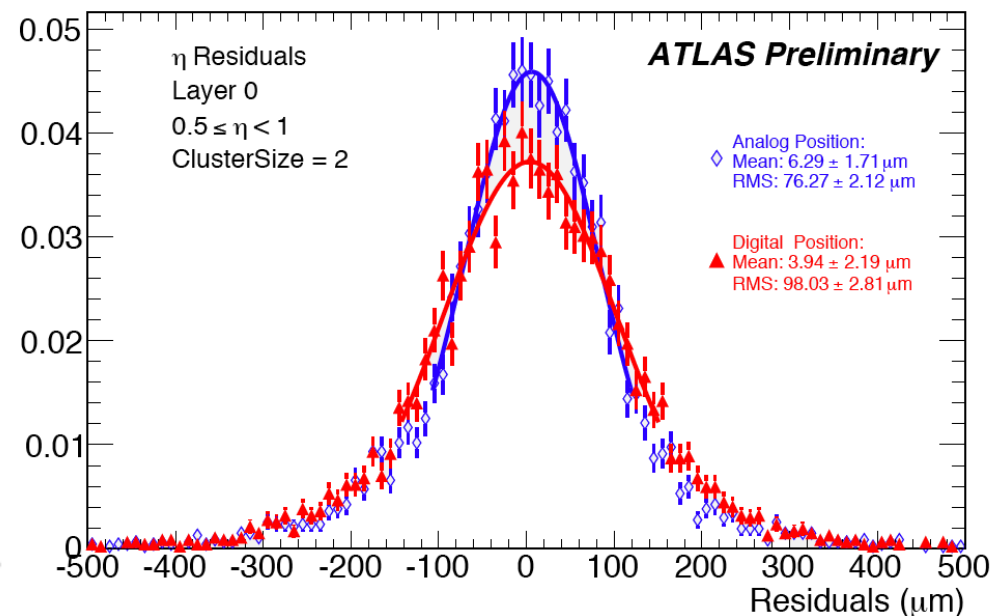
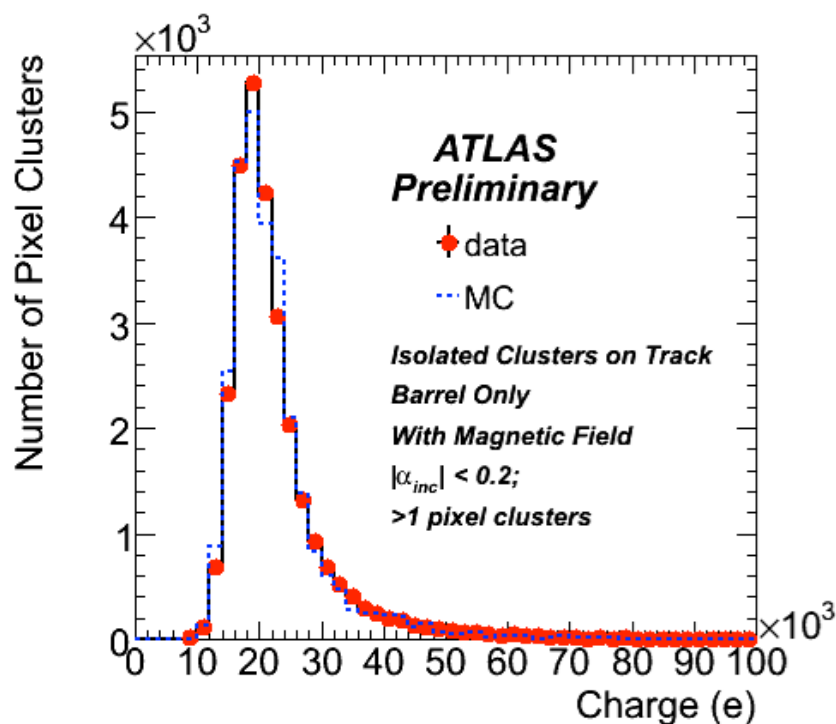
# Time over Threshold

- FE chips provide Time over Threshold information for each hit
  - ▣ Nearly linear dependence on deposited charge
- Pixel-by-pixel tuning; chosen tuning: 30 BC for 20 ke<sup>-</sup>
- Calibration by means of test charge injection



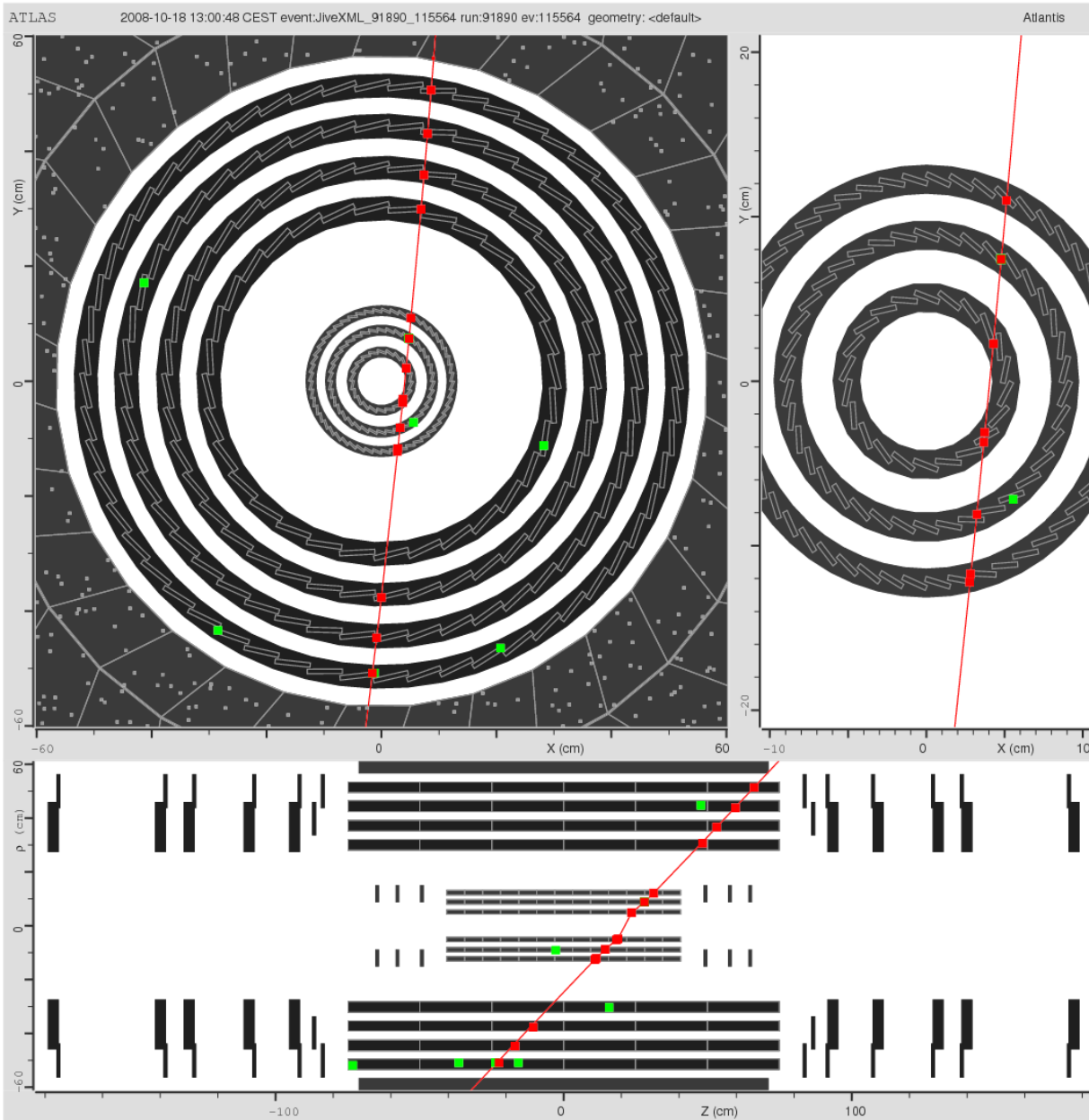
# Time over Threshold and Resolution

- Charge measurement with ToT in cosmic ray data taking
  - “Landau” peak at 18300 e<sup>-</sup> (Simulation 19000 e<sup>-</sup>): Confirms ToT Calibration
- Impact on resolution: still limited by statistics, but noticeable



$\sigma : 98 \mu\text{m} \rightarrow 76 \mu\text{m}$





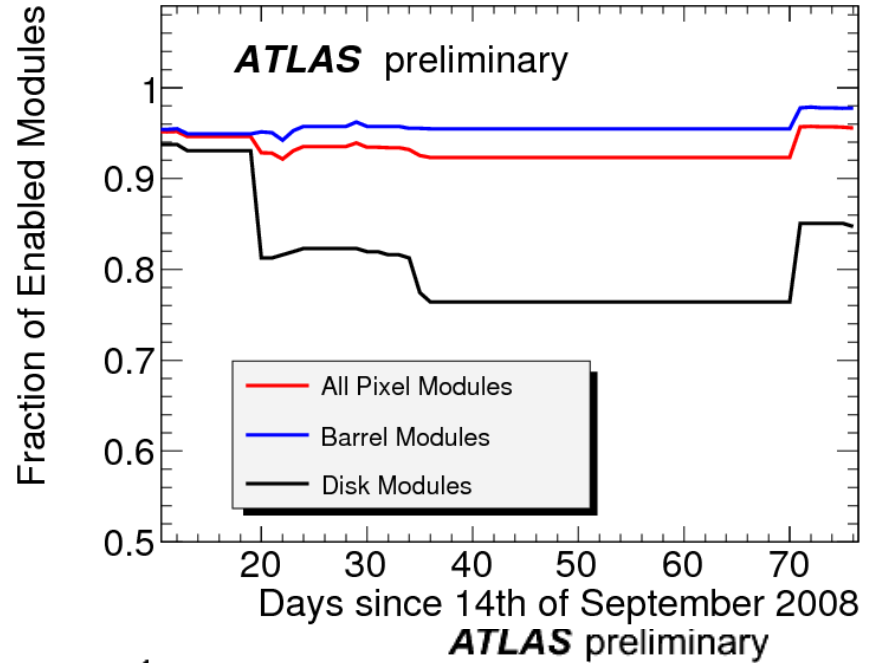
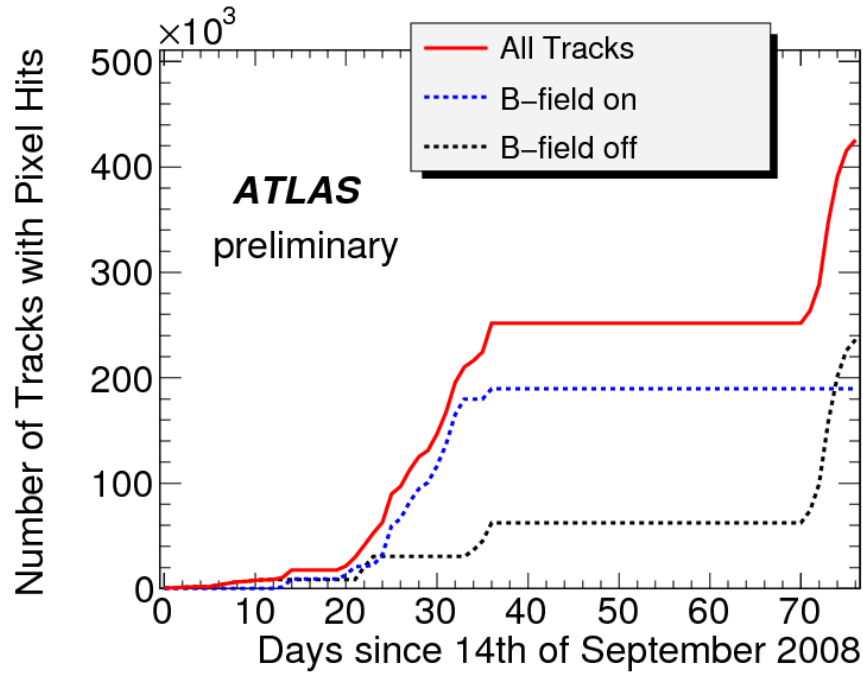
# Tracks

- Track with 8 pixel hits on track (2 x 2hits in module overlap regions)
- **Red**: hits on track
- **Green**: isolated hits (noise)
- Noise occupancy:  
  
~  $10^{-10}$  hits/pixel/BC

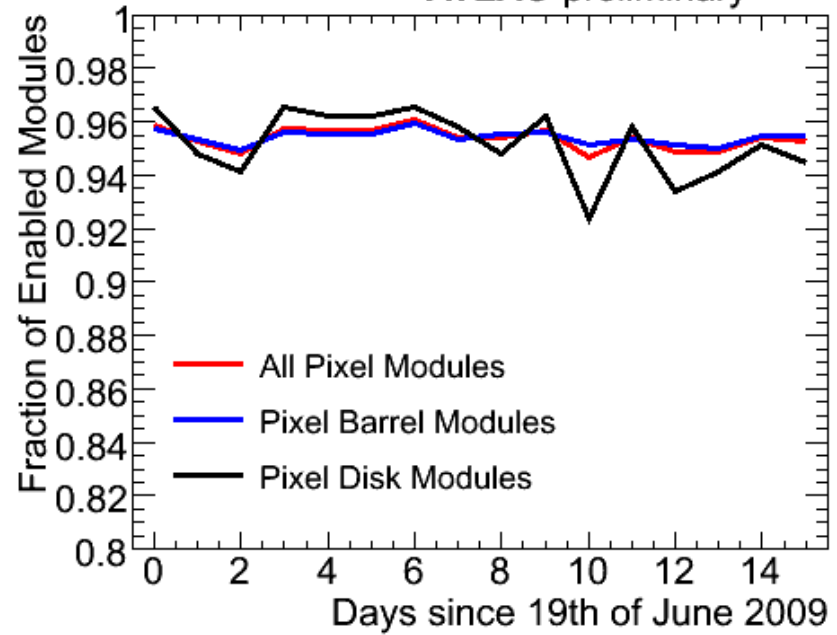
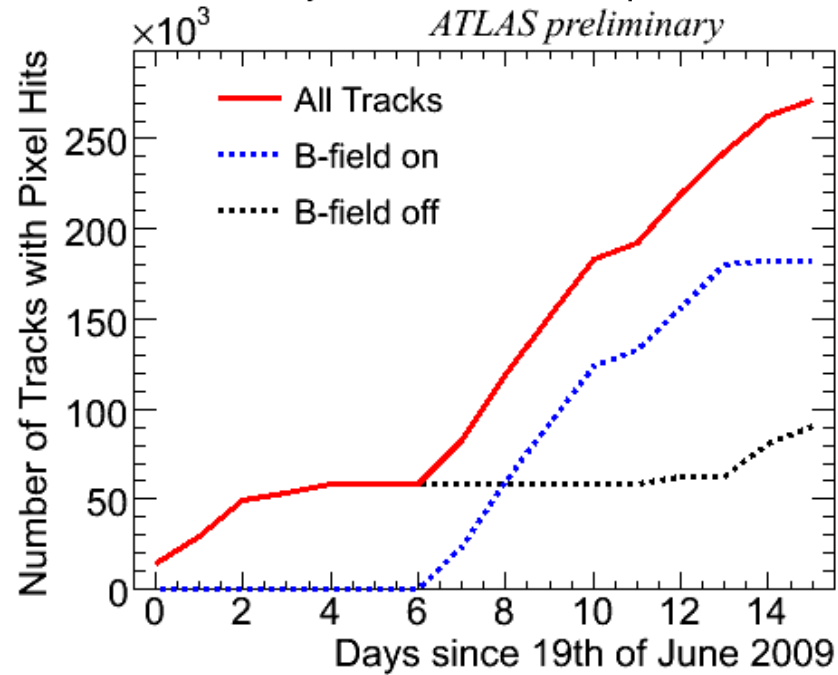




2008



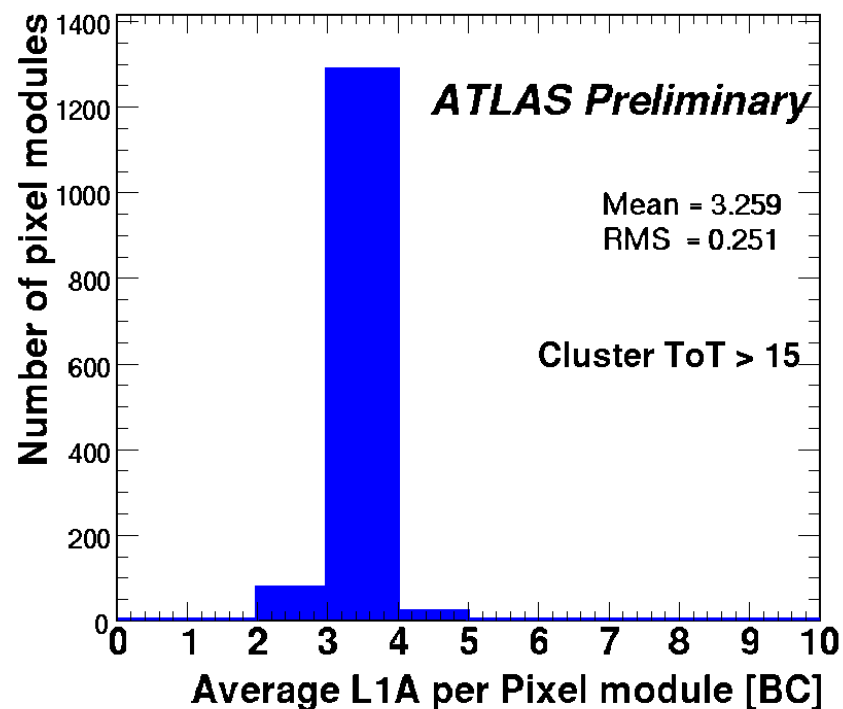
2009



# Timing

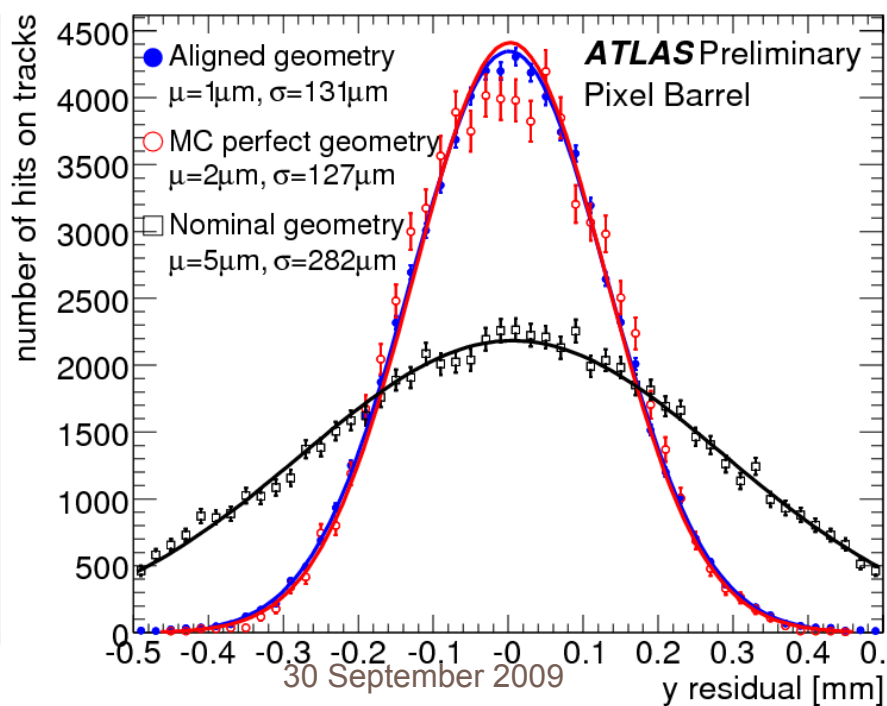
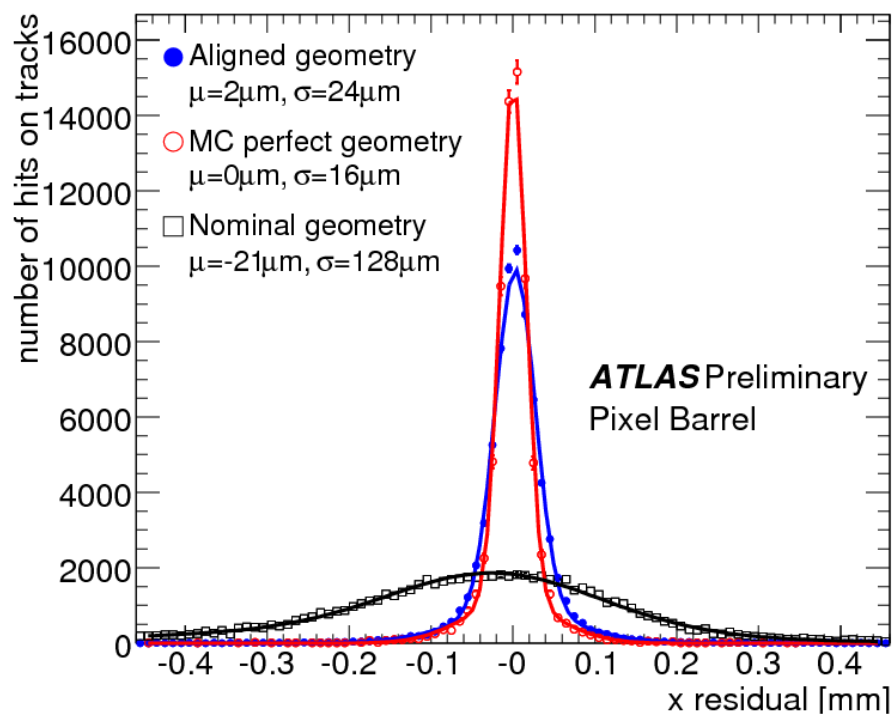
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- Each hit has to be assigned to the correct bunch crossing (25 ns)
- Module clocks have to be precisely aligned with the bunch crossing clock
- In cosmic ray data taking: readout of 8 consecutive BCs (plot shows hit time w.r.t. beginning of readout window)
- Correction of propagation delays:
  - ▣ First step: time alignment of readout crates (oscilloscope measurements)
  - ▣ Second step: time alignment of modules (using cable length data)
- Remaining effects:
  - ▣ Trigger jitter
  - ▣ Random phase of cosmics
  - ▣ Timewalk; “in-time” for less than  $\sim 5000 e^-$
- Plan to start data taking with 5 BC, later reduce readout window to 3 BC and 1 BC



# Alignment

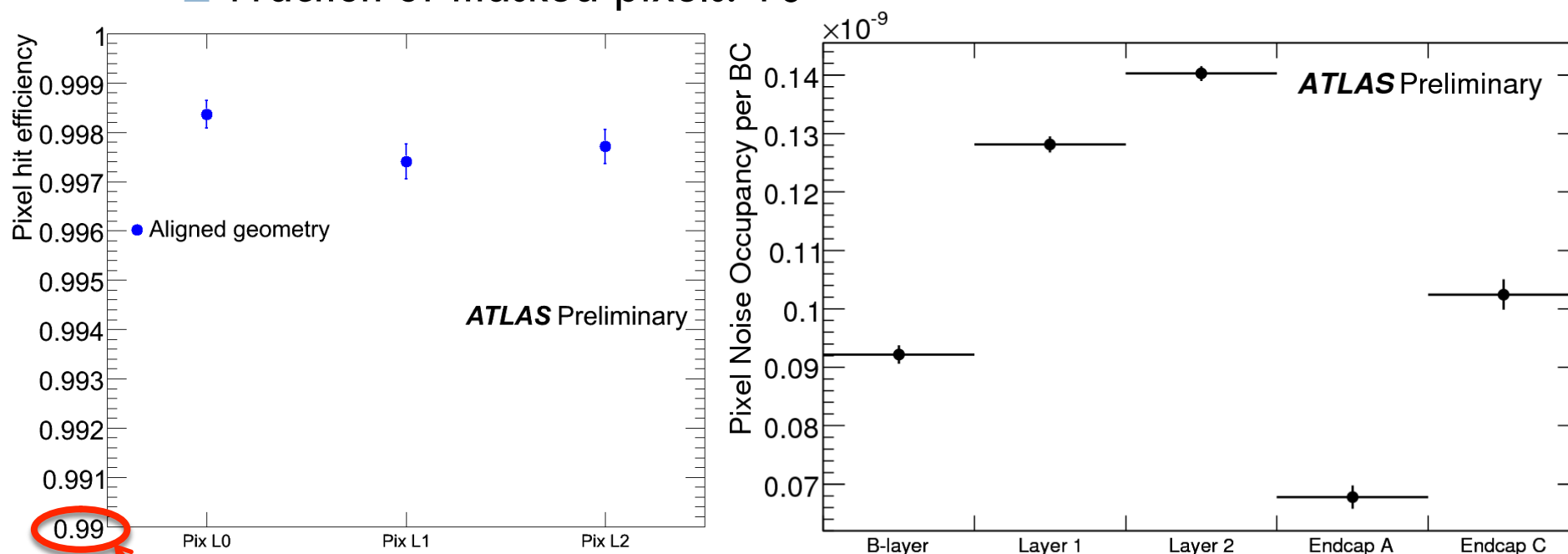
- Alignment of pixel barrel modules from cosmic data
  - ▣ Beam data needed for end-cap alignment
- Alignment not yet perfect due to limited statistics, but large improvement w.r.t. nominal geometry and good starting point for alignment with beam:
  - ▣ Precision direction:  $128\ \mu\text{m} \rightarrow 24\ \mu\text{m}$                       beam direction:  $282\ \mu\text{m} \rightarrow 131\ \mu\text{m}$





# Efficiency and Noise Occupancy

- After alignment measured efficiency is  $> 99.7\%$  for active modules
- Noise occupancy after masking of noisy pixels:  $\sim 10^{-10}$ 
  - Fraction of masked pixels:  $10^{-4}$



0.99

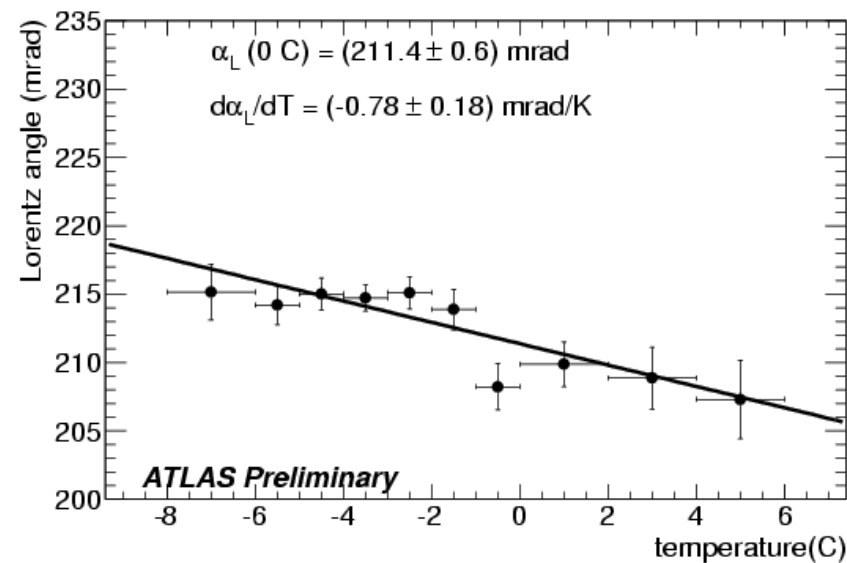
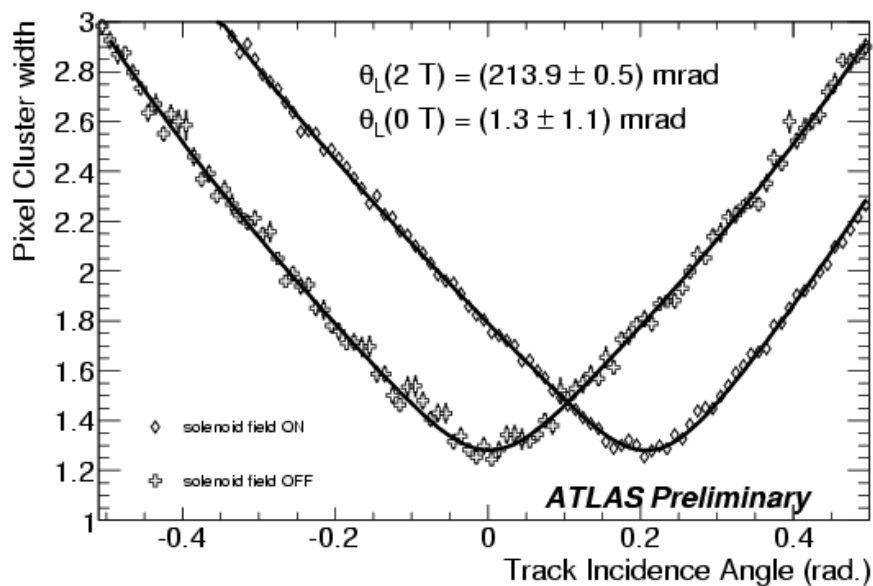
Zoom!



# Lorentz Angle Measurement

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- Cluster size vs. track angle with and without magnetic field → Measurement of the Lorentz angle
- Measured value close to expected value (225 mrad)
- Theoretically expected dependence on mobility can be nicely seen when including modules of different temperature
  - ▣ Measured:  $(-0.78 \pm 0.18)$  mrad/K, expected:  $-0.74$  mrad/K



# Status and expectations

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- The ATLAS Pixel Detector has been commissioned in a relatively short time
- Noise  $\sim 200 e^-$ , threshold tuned at  $4000 e^-$  with a dispersion  $\sim 40 e^-$ , m.i.p. signal at  $\sim 19000 e^-$
- The cosmic data taking has been extremely useful
  - ▣ Timing already in good shape: plan to reduce the readout window rapidly
  - ▣ Resolution  $\sim 24 \mu\text{m}$  in the precision direction, efficiency  $> 99.7 \%$
  - ▣ Noise occupancy  $\sim 10^{-10}$ ,  $10^{-4}$  fraction of masked pixels
- Many Pixel notes and a summary article in preparation
- Starting from this year all cooling loops are operated
  - ▣ Modest amount of coolant leak: studying the effects under irradiation
- 1.6 % of the detector is not functional due to on-detector failures
- Tuning at lower thresholds for beam related studies

**The Pixel Detector with  $\geq 98\%$  working modules is ready for LHC**

