

# Rassegna Sperimentale sulla Fisica del Neutrino

**Fabio Bellini**

“Sapienza” Università di Roma “ & INFN Roma

Congresso Prospettive Sezione INFN di Roma *06/05/2009*

# Neutrino Physics

- **How small is the Neutrino mass?** (Fermi, Pauli ~1930)
- **Which Neutrino: Dirac or Majorana?** (Majorana 1937)
- **Do Neutrinos oscillate?** (Pontecorvo ~1960)
- Astrophysical/Cosmological Neutrinos:  
relic neutrinos? Supernova neutrinos? high Energy neutrinos?
- Electromagnetic Neutrino Properties?
- Geo-neutrinos?
- Sterile neutrinos?
- .....

*see Capone &  
De Bernardis' talk*

# Oscillations Results

**Solar**

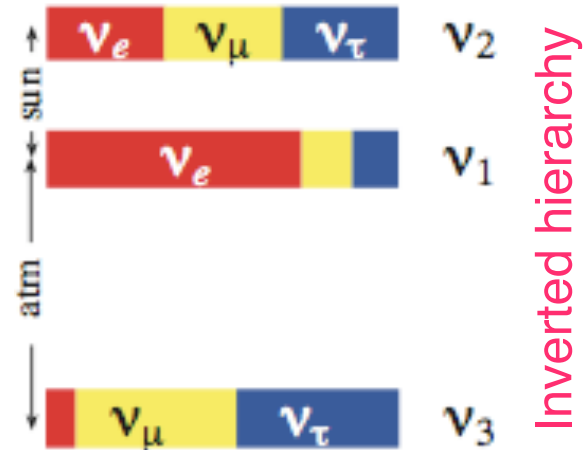
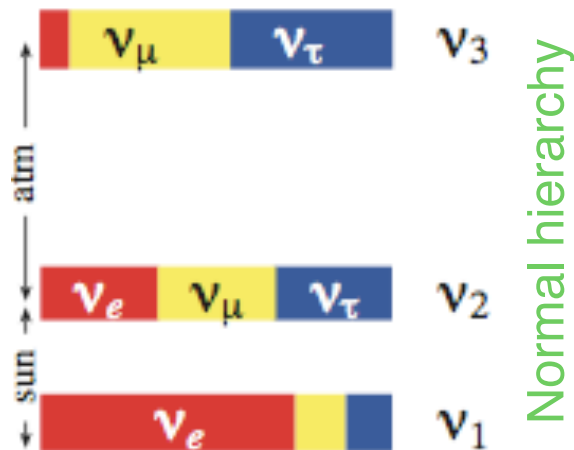
$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{bmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

**Atmospheric + accelerator**      **Mainly Chooz + Minos+solar+atmospheric**      **Solar+ Kamland**

**Atmospheric**       $\theta_{23} \equiv \theta_{\text{atm}} = 45^\circ \pm 8^\circ$        $\theta_{13} < 10^\circ$        $\theta_{12} \equiv \theta_{\text{sol}} = 34^\circ \pm 3^\circ$   
 $\sin^2(2\theta_{23}) = 1.02 \pm 0.04$        $\sin^2(2\theta_{13}) = 0.07 \pm 0.04$        $\tan^2(\theta_{12}) = 0.484 \pm 0.048$

$$\Delta m_{21}^2 \equiv \Delta m_{\text{sol}}^2 = (7.58 \pm 0.21) \cdot 10^{-5} \text{eV}^2 \quad |\Delta m_{32}^2| \equiv \Delta m_{\text{atm}}^2 = (2.40 \pm 0.15) \cdot 10^{-3} \text{eV}^2$$

Strumia, Vissani, hep-ph/0606054



# What's next

- Precision era:  $\Delta m^2_{12}$ ,  $\theta_{12}$ ,  $\Delta m^2_{23}$ ,  $\theta_{23}$
- Appearance paradox: *appearance of new flavors never observed unambiguously*

$$\nu_{\mu} \rightarrow \nu_{\tau}$$

**OPERA**

$$\nu_{\mu} \rightarrow \nu_e$$

**T2K, NOvA, MINOS?**

- Measure  $\theta_{13}$ :

▶ sub-leading transition  $\nu_{\mu} \rightarrow \nu_e$

▶  $\bar{\nu}_e$  disappearance from **reactors**

**DOUBLE CHOOZ, DAYA BAY**

- Depending on  $\theta_{13}$ :

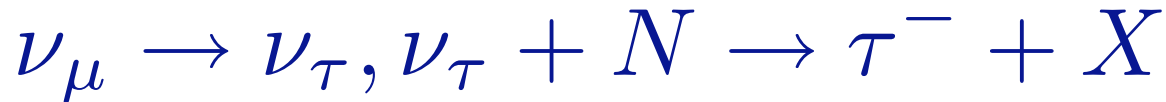
▶ **Mass hierarchy** @ accelerators

▶ **CP violation**:  $P_{CP} \propto 0.23 \sin(\theta_{13}) \sin(\delta) \sin^2(\Delta m^2_{13} L/E) \sin(\Delta m^2_{12} L/E)$

# $\nu_\mu \Rightarrow \nu_\tau$ : OPERA

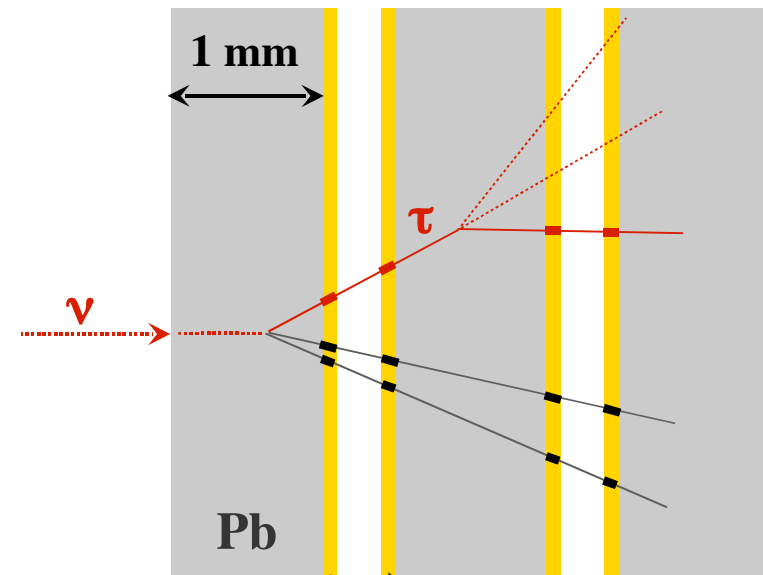
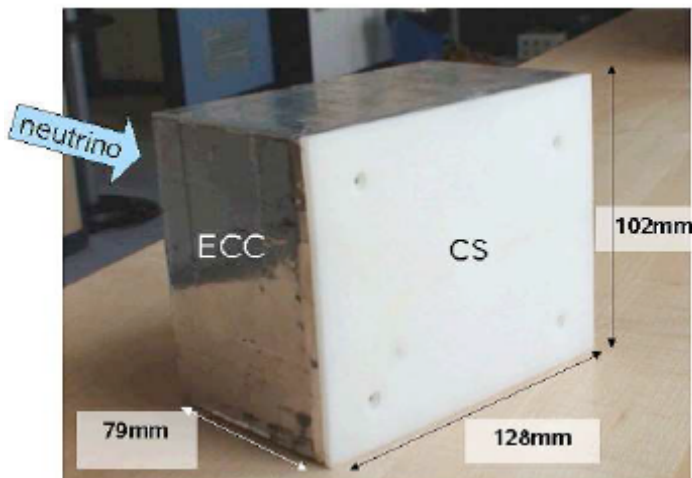
A. Buccheri, M. Capodiferro, P. Pecchi, A. Pelosi, G. Rosa, A. Ruggeri

- Goal: identify  $\nu_\tau$  appearance through CC interaction

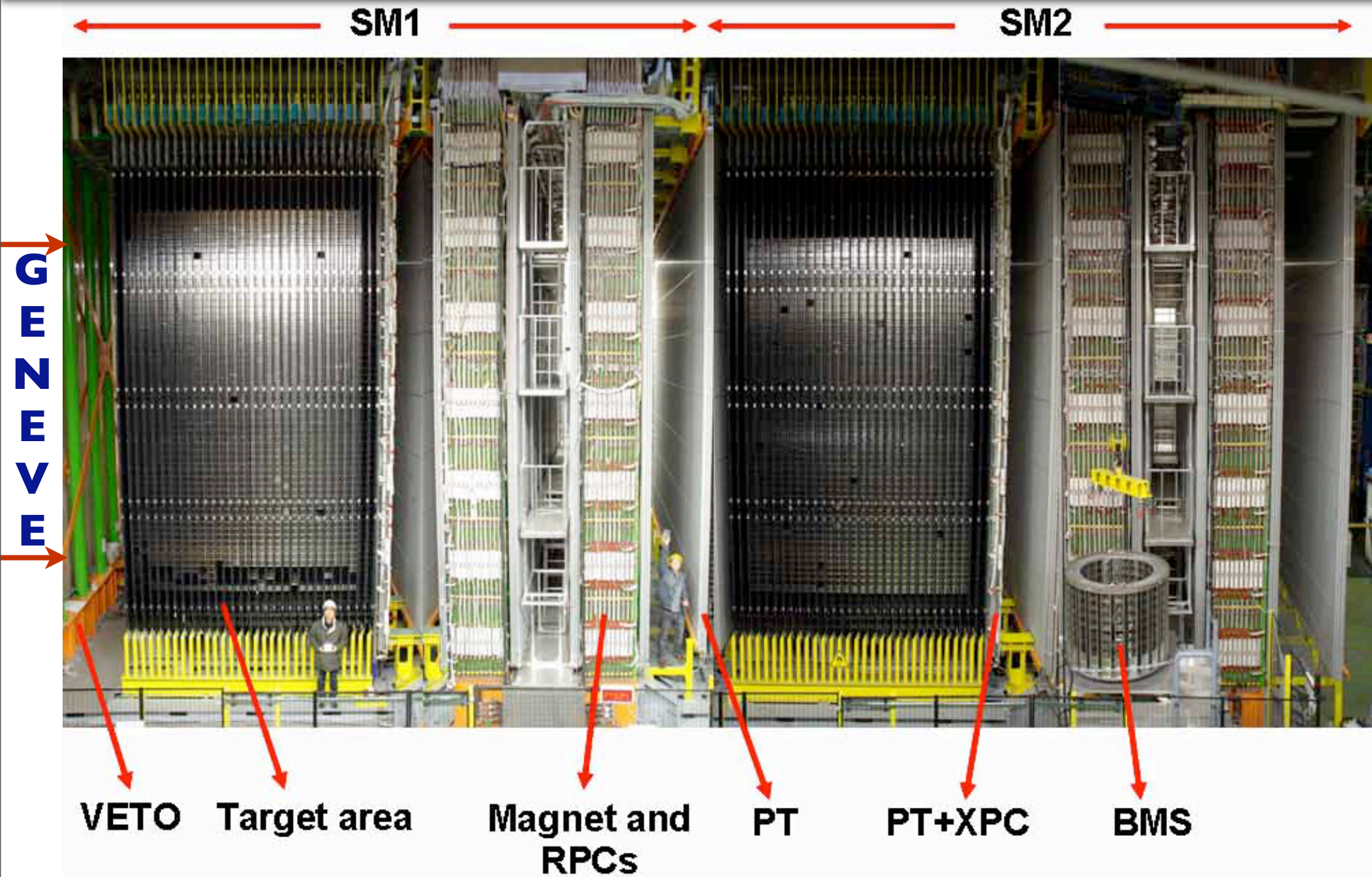


$\mu^- \nu_\tau \nu_\mu$	B. R. $\sim 17\%$
$h^- \nu_\tau n(\pi^0)$	B. R. $\sim 50\%$
$e^- \nu_\tau \nu_e$	B. R. $\sim 18\%$
$\pi^+ \pi^- \pi^- \nu_\tau n(\pi^0)$	B. R. $\sim 14\%$

- Need: **O(kTon)** target mass + **high granularity detector** ( $\sigma \sim \mu\text{m}$ ) for  $\tau$  detection ( $c\tau \sim 87 \mu\text{m}$ ) and bkgd rejection (mainly  $\nu_\mu N \rightarrow c\mu X$ )
- ▶ **Lead-nuclear emulsion** target segmented into basic units (bricks): *57(330 $\mu\text{m}$ ) emulsion films between 1 mm Pb sheets*



# In Picture

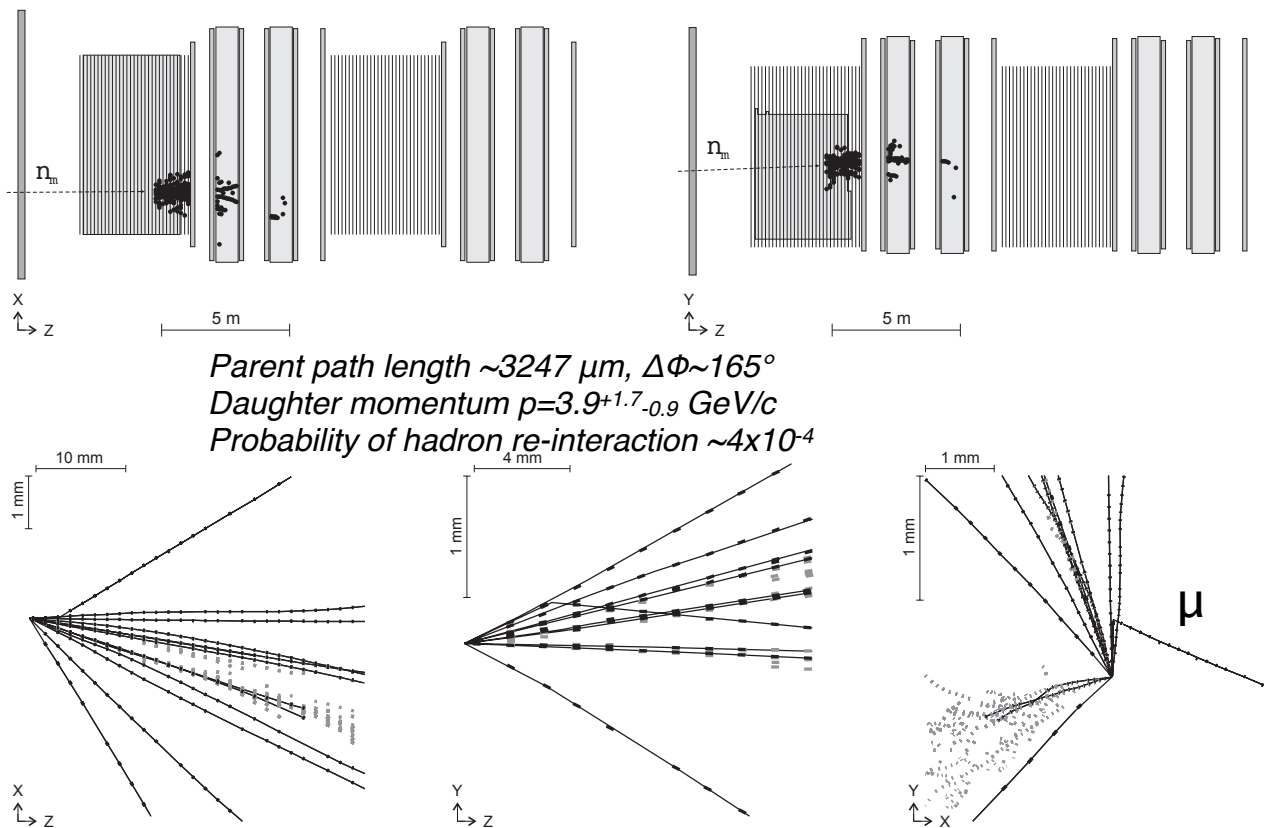


# CNGS Run 2008

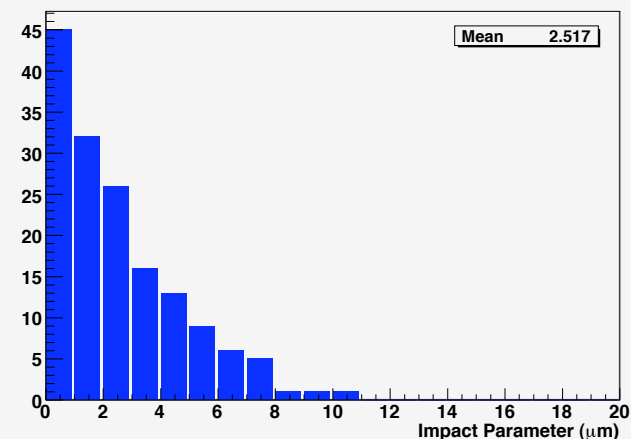
- Total beam intensity:  **$1.8 \cdot 10^{19}$  pot** (exp.  $4.5 \cdot 10^{19}$  pot/y)
- **1663 v candidate** interactions in the target
- **850 bricks** scanned: 650 interactions (8 charm event)

$\langle E (v_\mu) \rangle$	17 GeV
L	730 km
L/E	43 Km/GeV
$(v_e + \bar{v}_e)/v_\mu$	0.87%
$\bar{v}_\mu / v_\mu$	4%
$v_\tau$ prompt	negligible

## first charm event



[hep-ex 0903.2973v1](#)  
**live time >99%**  
**first brick finding eff.  $\sim 70\%$**   
**vertex location eff.  $\sim 80-90\%$**



# Opera perspectives

- Expectation for 2009:  $3.6 \cdot 10^{19}$  pot  $\Rightarrow$  3500  $\nu$  events on detector
  - ▶ Looking forward to see the first  $\tau$  event
- Total “Nominal” pot in 5 years:  $22 \cdot 10^{19}$  pot

$\tau$ decay channels	$\epsilon(\%)$	BR (%)	Signal $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$	Background:
$\tau \rightarrow \mu^-$	17.5	17.7	2.9	0.17
$\tau \rightarrow e^-$	20.8	17.8	3.5	0.17
$\tau \rightarrow h^-$	5.8	49.5	3.1	0.24
$\tau \rightarrow 3h$	6.3	15	0.9	0.17
ALL	$\epsilon \times \text{BR} = 10.6\%$		10.4	0.75

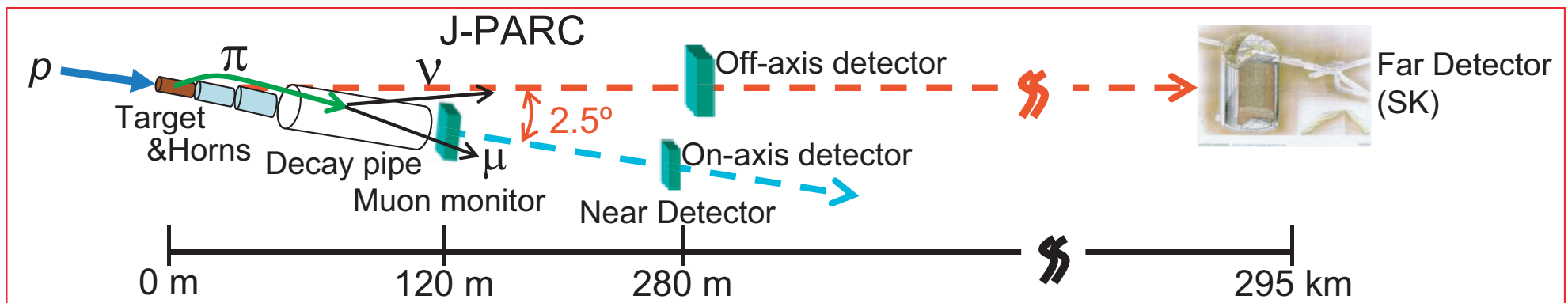
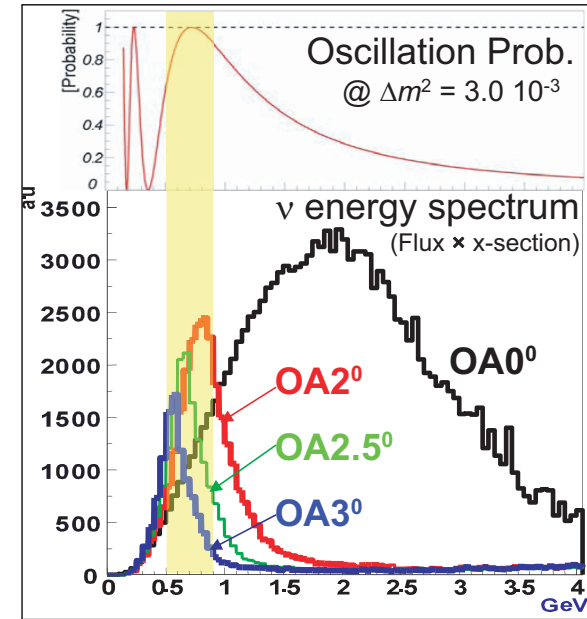
- Activity in Rome: *emulsion scanning station for data analysis*



# $\nu_\mu \Rightarrow \nu_\mu, \nu_\mu \Rightarrow \nu_e$ : T2K

*U.Dore, C. Gargiulo, P. Loverre, L.Ludovici, C.Mariani*

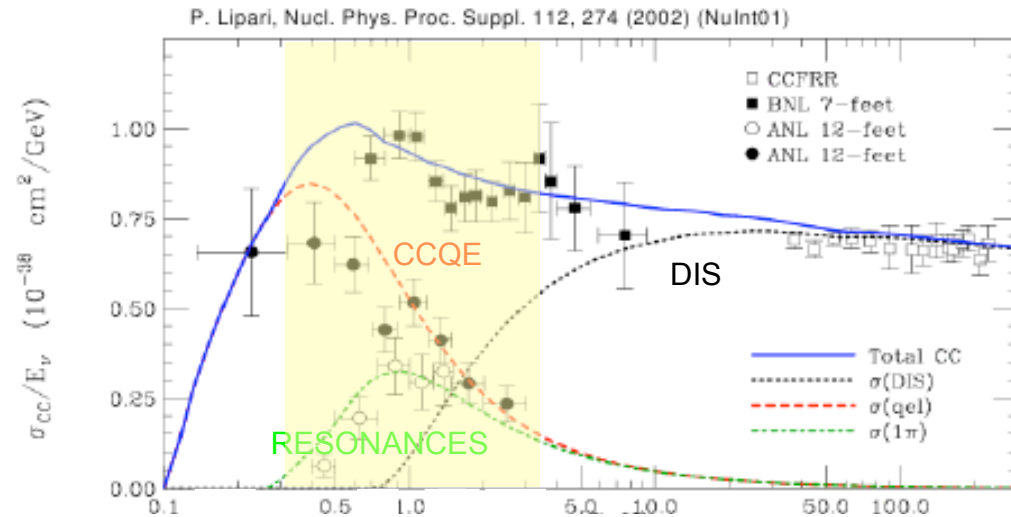
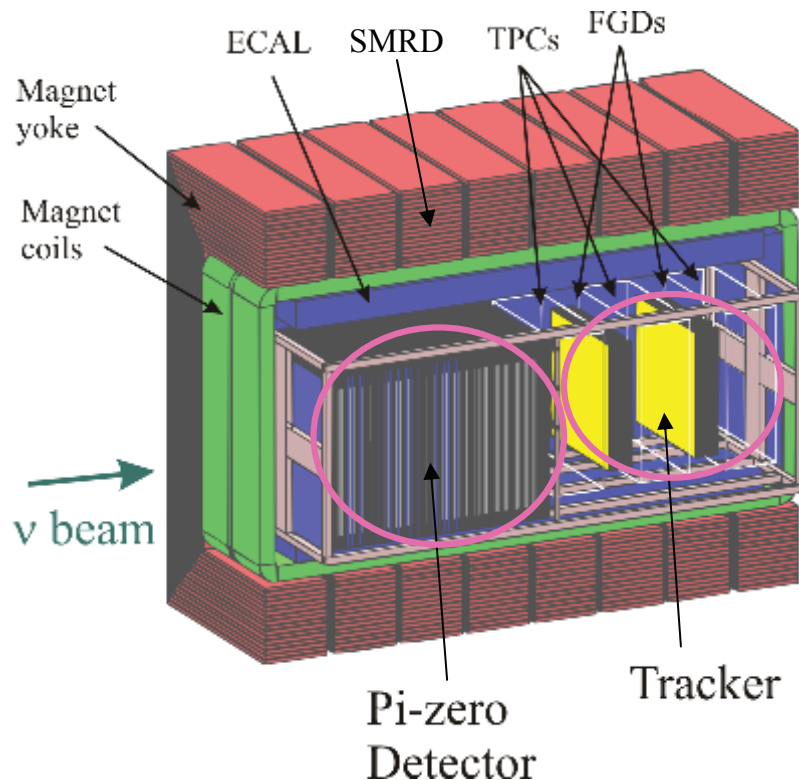
- Precise measurement of  $\Delta m^2_{23}, \theta_{23}$  from  $\nu_\mu$  disappearance
  - ▶ signal:  $\nu_\mu n \rightarrow p \mu^-$       bkgd: CC-1 $\pi$  events
- Search for  $\theta_{13}$ :  $\nu_\mu \Rightarrow \nu_e$  appearance
  - ▶ signal:  $\nu_e n \rightarrow p e^-$       bkgd: Intrinsic beam  $\nu_e$ , NC-1 $\pi^0$
- **Off-axis** beam technique:  $\nu_e$  contamination @ peak  $\sim 0.4\%$



# The Close detector

Key for good  $E_\nu$  spectrum and background estimate

Installation foreseen in this summer



Rome activity:

- data taking/electron final state analysis (beam  $\nu_e$  contamination)
- feasibility study for LAr near detector for T2K upgrade

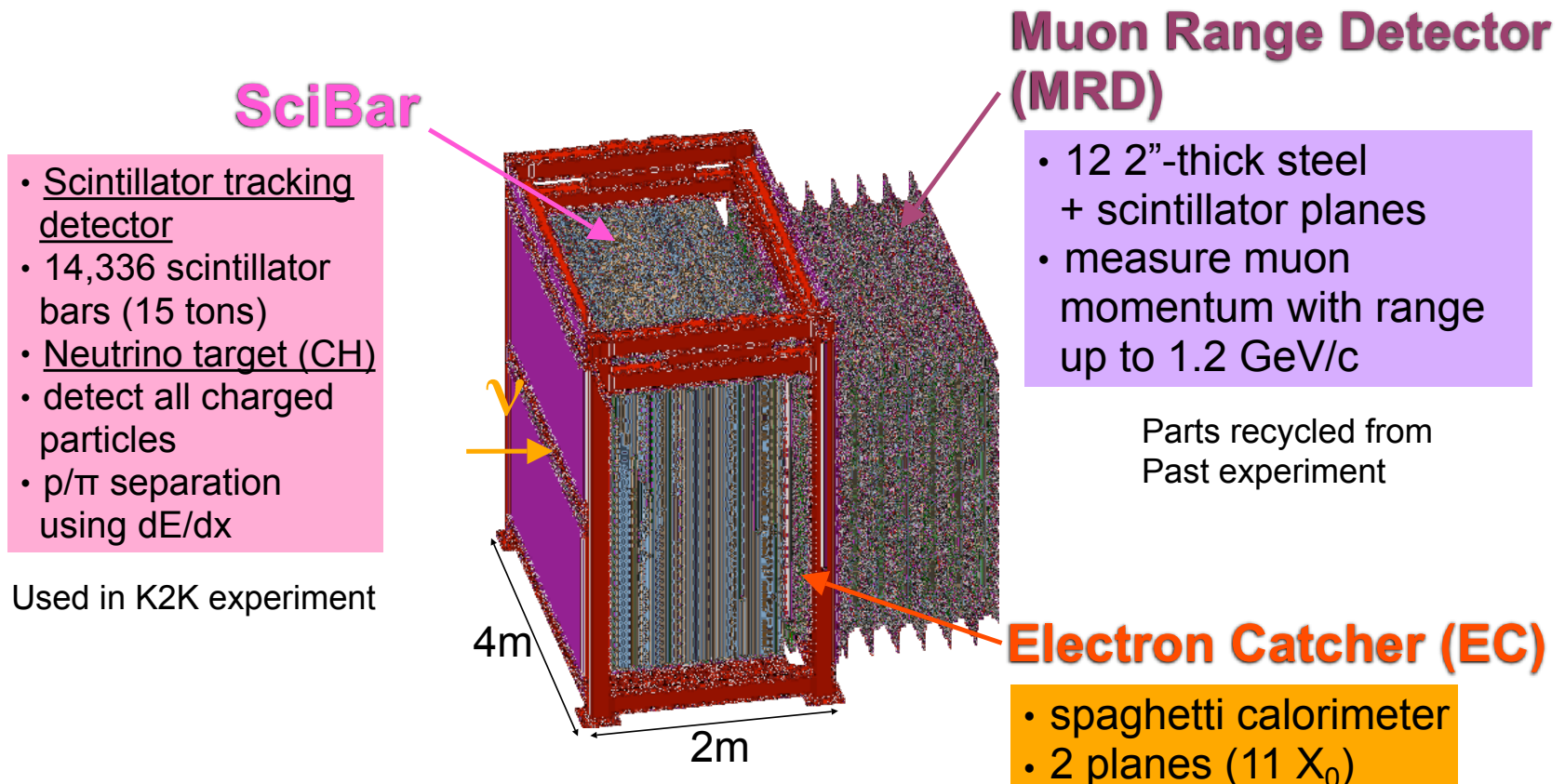
- **April 23<sup>rd</sup> 2009: first protons on target**

<http://www.kek.jp/intra-e/press/2009/J-PARCT2K.html>

- **Start Data taking** at **December 2009** and accumulate  **$\sim 10^{20}$  POT** by **summer 2010**

# $(\bar{\nu})\nu$ cross section: SciBoone

- **K2K-SciBar detector at FNAL Booster Neutrino Beam**



- Data taking is over. Collected statistics:  $1(1.5) \cdot 10^{20}$  POT di  $\nu$  ( $\bar{\nu}$ )
- **No evidence** of CC coherent  $\pi$  production *Phys.Rev.D 78:112004, 2008*
- **CC-QE, CC1 $\pi$ , NC1 $\pi$ , NC-elastic** measurement cross section on going

# $\theta_{13}$ Experiments

- **T2K:**  $\sin^2(2\theta_{13}) < 0.01$  in 5 years (no sensitivity for CP violation and matter effects)

- **Double Chooz:**

- ▶ End of **2009**: start with only far detector:  $\sin^2(2\theta_{13}) < 0.06$  90%CL in 1.5 y

- ▶ Far+Near from **2011**:  $\sin^2(2\theta_{13}) < 0.03$  90%CL in 3 y

- **Daya Bay:**

- ▶ Start in **2012**  $\sin^2(2\theta_{13}) < 0.01$  90%CL

- **NOvA:**

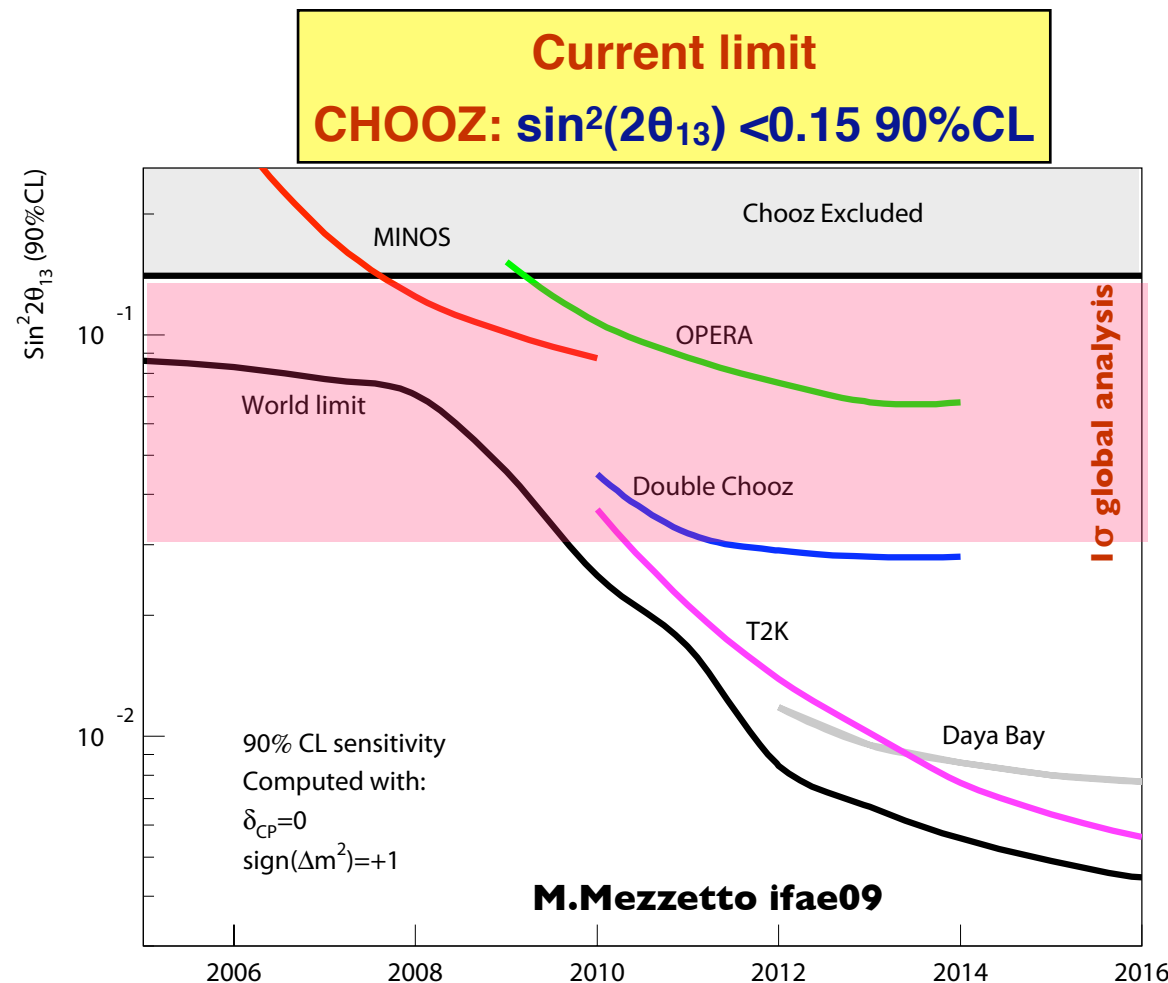
- ▶ 2<sup>nd</sup> generation LBL experiment

(15kTon tracking liquid scintillator at  
810 km off-axis from NUMI(700kW))

- ▶ Full detector start:**2013**

Groundbreaking started May 1st 2009

- ▶  $\sin^2(2\theta_{13}) < 0.01$  + matter and CP effect

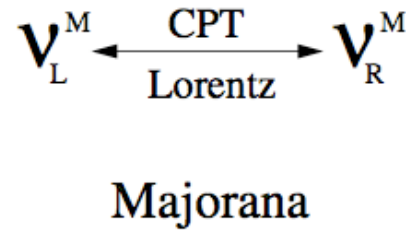
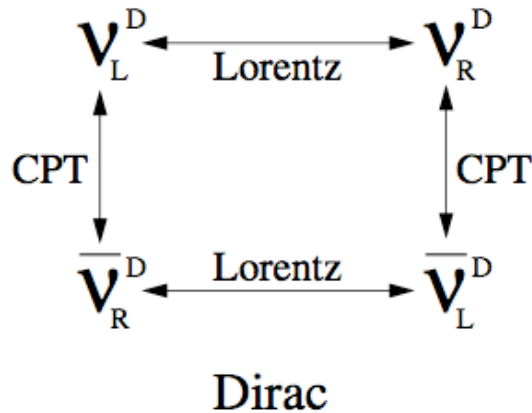


# $\theta_{13}$ dilemma

- All mixing matrix and mass hierarchy could be explored with accelerator experiment
- If  $\sin^2(2\theta_{13}) > 0.01$  ( $\theta_{13} > 3^\circ$ )
  - ▶ **Superbeam** (2-4 MW) + **new detectors**  
Electronic final state  $\rightarrow$  *low density detector*: Mton WC, 0.1Mton LAr, Liquid scintillator  
Systematic limited: flux composition and low energy nuclear cross section
- If  $\sin^2(2\theta_{13}) < 0.01$  ( $\theta_{13} < 3^\circ$ )
  - ▶ **Neutrino factories**: initial beam  $\nu_\mu + \bar{\nu}_e$  or  $\bar{\nu}_\mu + \nu_e$   
Golden channel  $\nu_e \Rightarrow \nu_\mu$  **wrong sign muons**: Minos like detector (*wrong charge rejection*  $< 10^{-3}$ )  
Silver channel  $\nu_e \Rightarrow \nu_\tau$ : 4xOPERA or 10 kTon LAr detector  
High density detector:  $E > 10\text{GeV}$ ,  $L \sim 3000\text{ km}$ , consistent matter effect
  - ▶ **Beta beams**:  $\nu_\mu$  appearance but only one flavor at  $t=0$ :  $\nu_e$  or  $\bar{\nu}_e$   
Drawback low energy neutrinos sub GeV: single  $\pi$  NC production and e mis-id

# Dirac or Majorana?

- Majorana conjecture (1937)  $\nu = \nu^c$

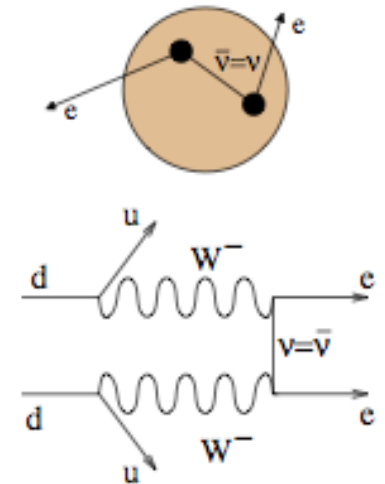


Majorana

- Massless neutrino **DOES NOT** allow testing of the Majorana nature
- Implication: Lepton number violation
- Neutrinoless Double Beta Decay:  $(A,Z) \rightarrow (A,Z+2) + 2 e^-$

$$\Gamma_{1/2}^{0\nu} \propto Q^5 |M_{nucl}|^2 |m_{\beta\beta}|^2$$

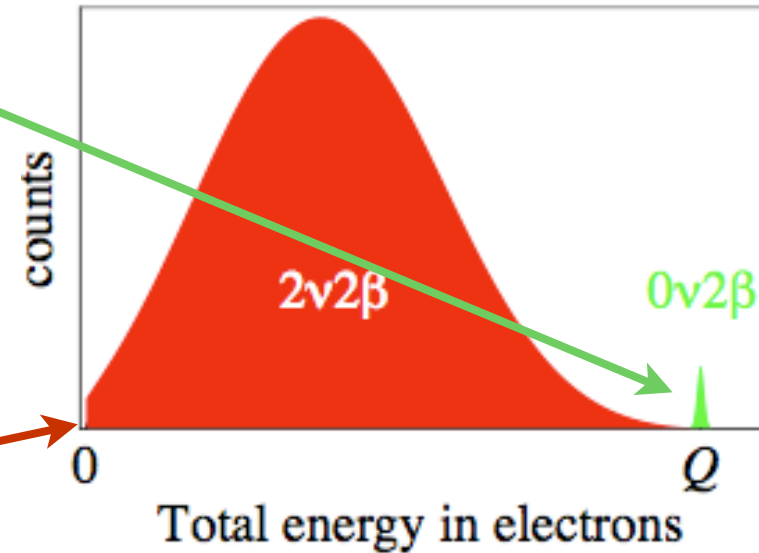
$$m_{\beta\beta} = \cos^2 \theta_{13} (m_1 \cos^2 \theta_{12} + m_2 e^{2i\alpha} \sin^2 \theta_{12}) + m_3 e^{2i\beta} \sin^2 \theta_{13}$$



# The challenge

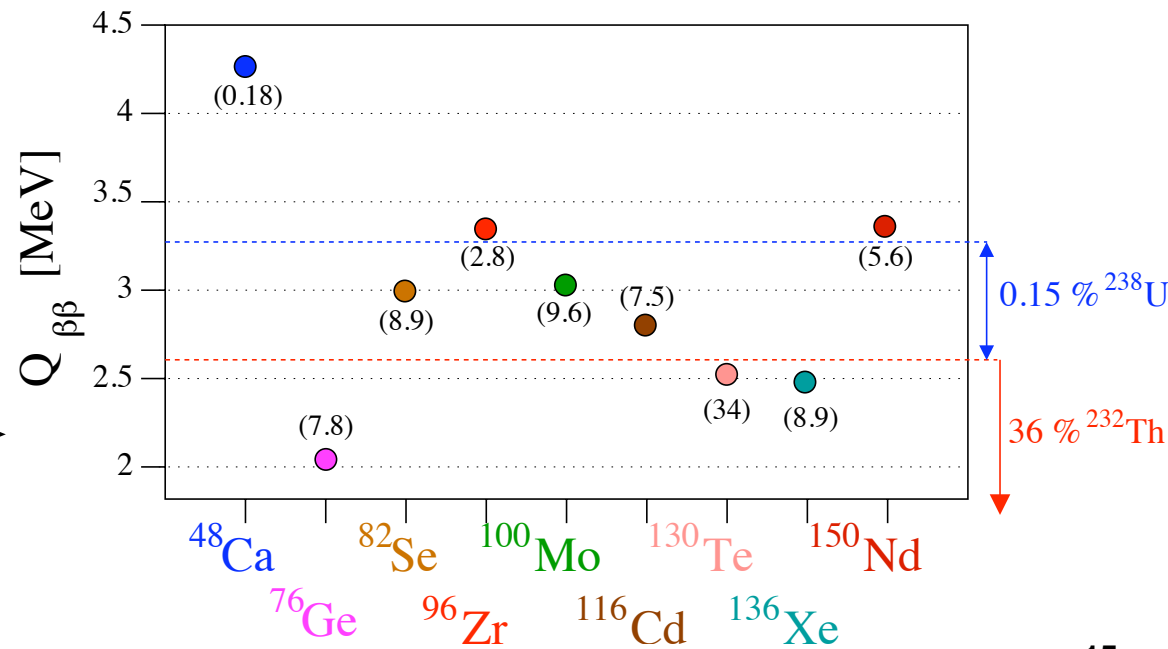
- **Signature:** monochromatic line at the **Q-value**

- $\tau_{0\nu\beta\beta} \sim 10^{25\sim 26} \text{ y}$  ( $m_{\nu\text{light}} \sim 50 \text{ meV}$ )
  - ▶ 1 event/year  $\rightarrow 10^3$  moles  $\sim O(100\text{kg})$



- **Background:**

- ▶  $\tau_{2\nu\beta\beta} \sim 10^{20} \text{ y}$  ( $N_{2\nu\beta\beta}/N_{0\nu\beta\beta} \sim (\delta E/Q)^6$ )
- ▶ Primordials: **U, Th** ( $\tau \sim 10^9 \text{ y}$ )
- ▶ neutrons & cosmic induced
- ▶  $\beta$  decay forbidden
- $\Rightarrow$  **even-even nuclei**

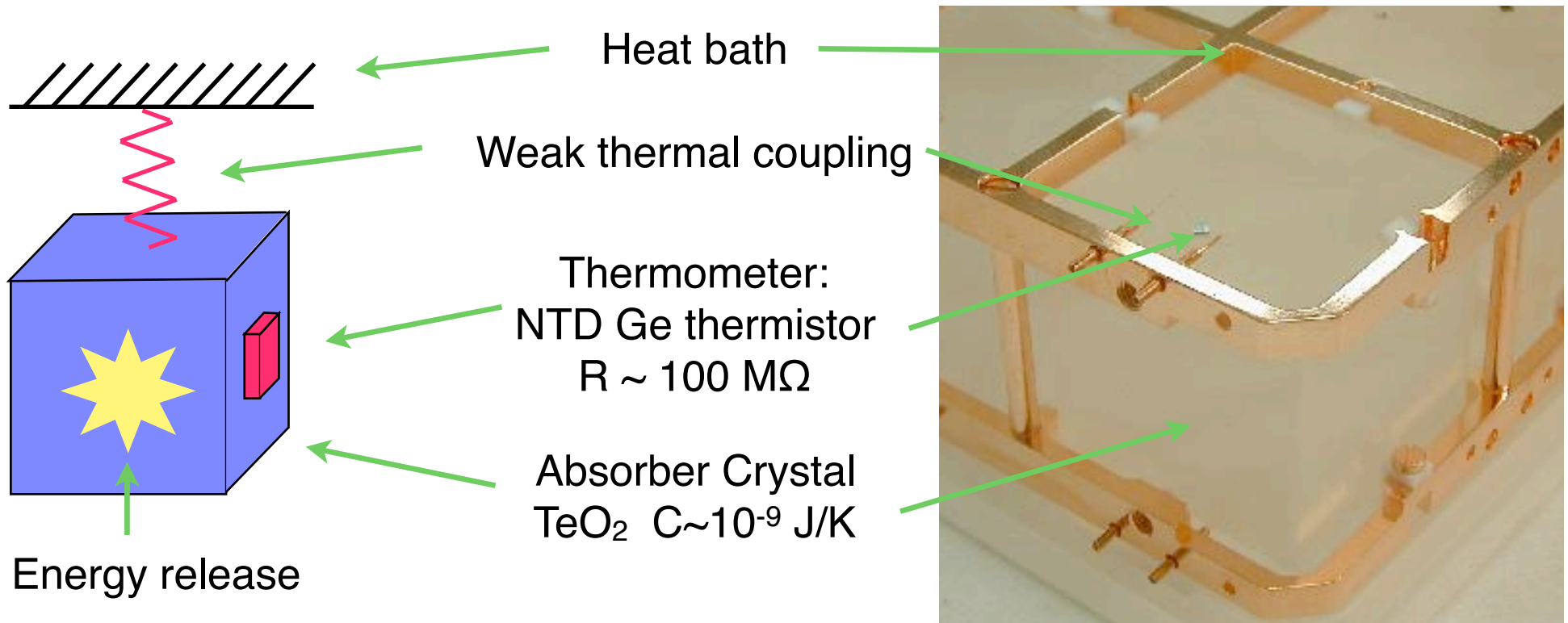


# CUORICINO

*F.Bellini, F.Orio. C.Tomei, M.Vignati*

- Particle energy converted into phonons  $\rightarrow$  temperature variation  $\Delta T = E/C$
- Need very low heat capacity: crystals (dielectric, diamagnetic) @  $\sim 10\text{mK}$

$\text{TeO}_2$   
detector = source



- Detector response in this configuration:  $\sim 0.2\text{ mK} / \text{MeV} \sim 0.2\text{ }\mu\text{V}/\text{MeV}$
- FWHM Resolution @  $0\nu\text{DBD} \sim 0.3\%$

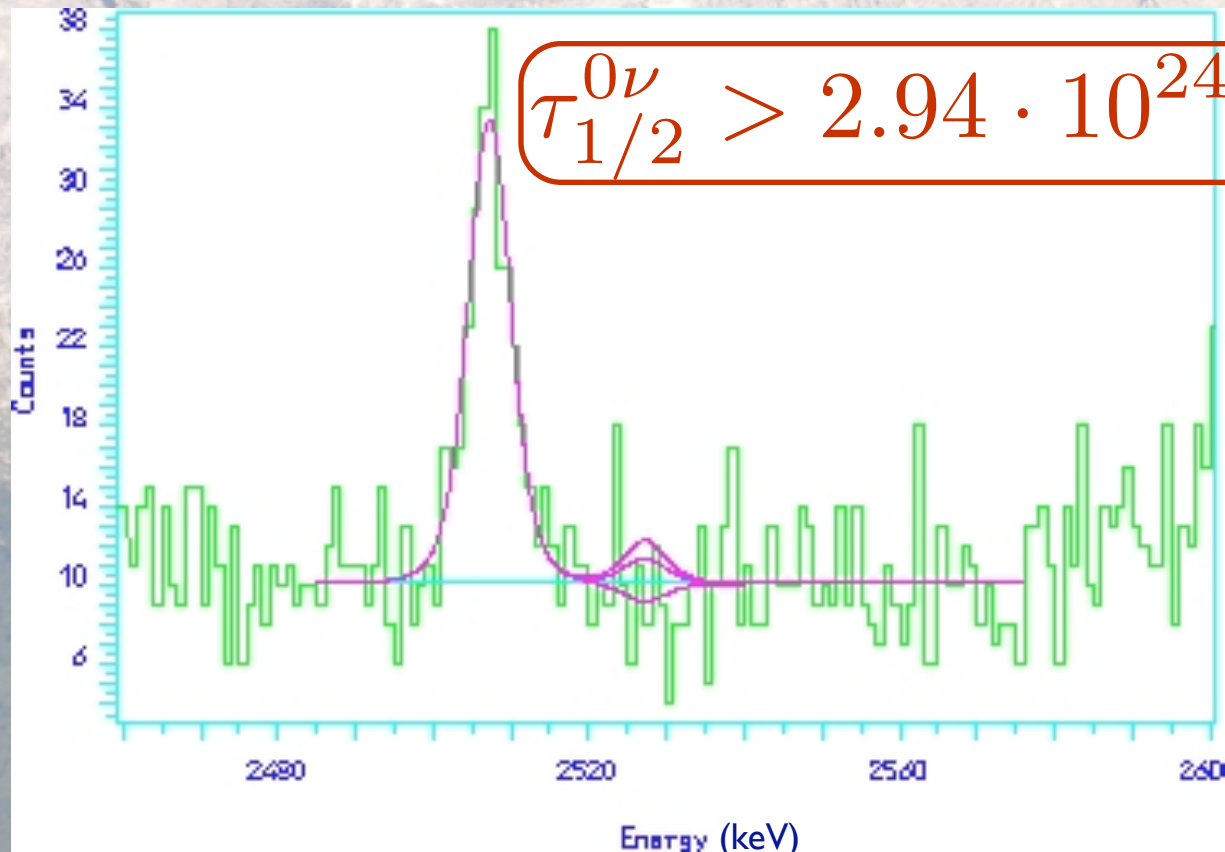
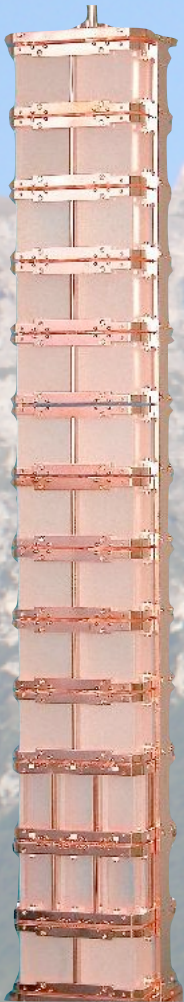


# Final (preliminary) Results

*Hosted @ Laboratori Nazionali del Gran Sasso, Italy.*

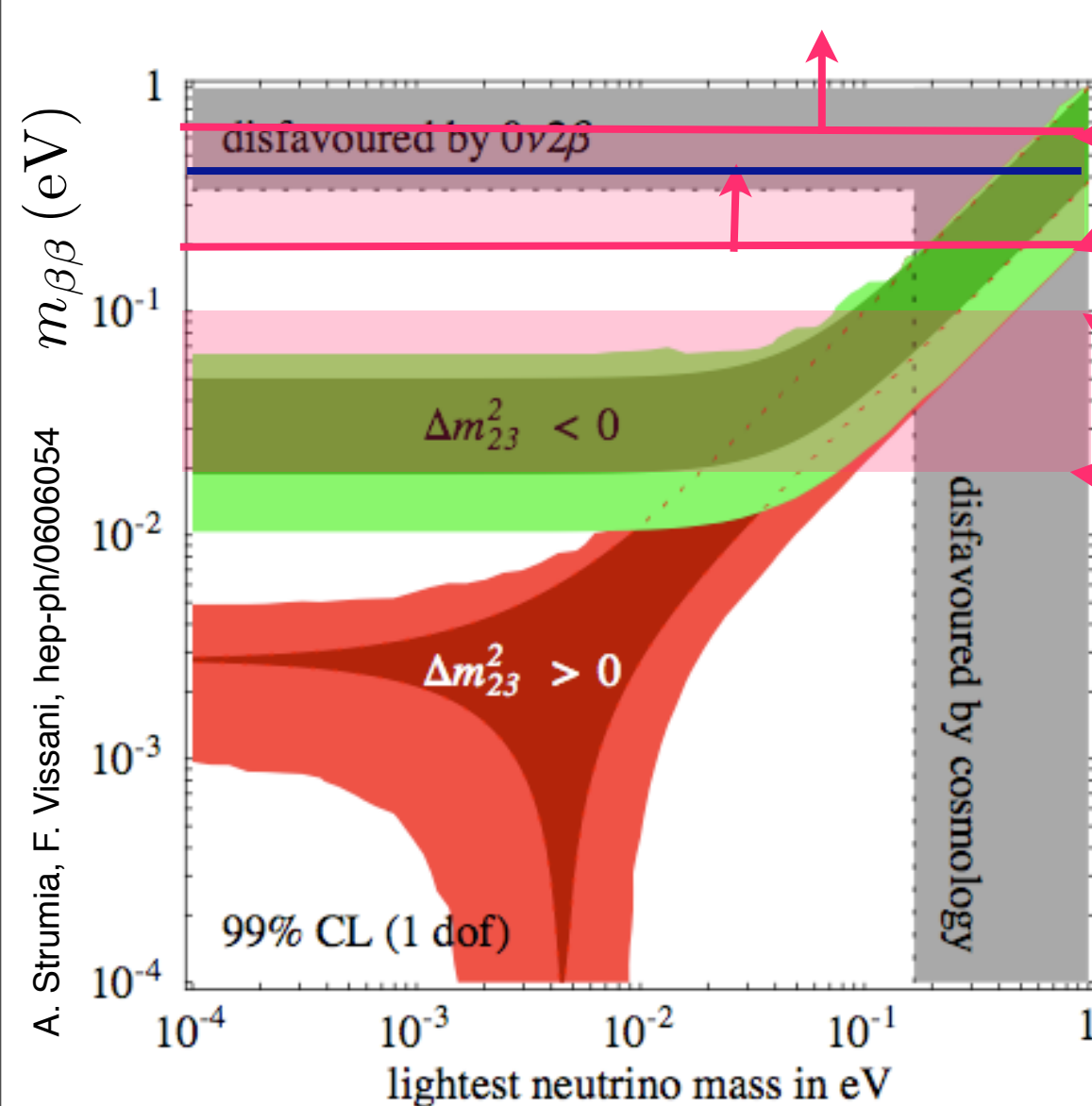
Final statistics:  $M \cdot t = 18 \text{ kg } ^{130}\text{Te} \cdot \text{y}$

Background level:  $b = (0.18 \pm 0.01) \text{ counts/keV/kg/y}$



# $0\nu\beta\beta$ : where we are & we will be

- Express  $m_{\beta\beta}$  as function of measured oscillation parameters and the unknown lightest neutrino mass:  $m_{\beta\beta} = f(U_{ek}, m_{lightest}, \Delta m_{12}, \Delta m_{13})$



## CUORICINO

$$m_{\beta\beta} < 0.20 \div 0.68 \text{ eV}$$

## CUORE goal:

test inverted hierarchy

$$\tau_{1/2}^{0\nu} > 2.1 \cdot 10^{26} \text{ y @90\%CL}$$

$$m_{\beta\beta} < 20 \div 100 \text{ meV}$$

Scaling CUORICINO to CUORE

$$M = 20 \times m$$

$$\delta E = 0.7 \times \delta E$$

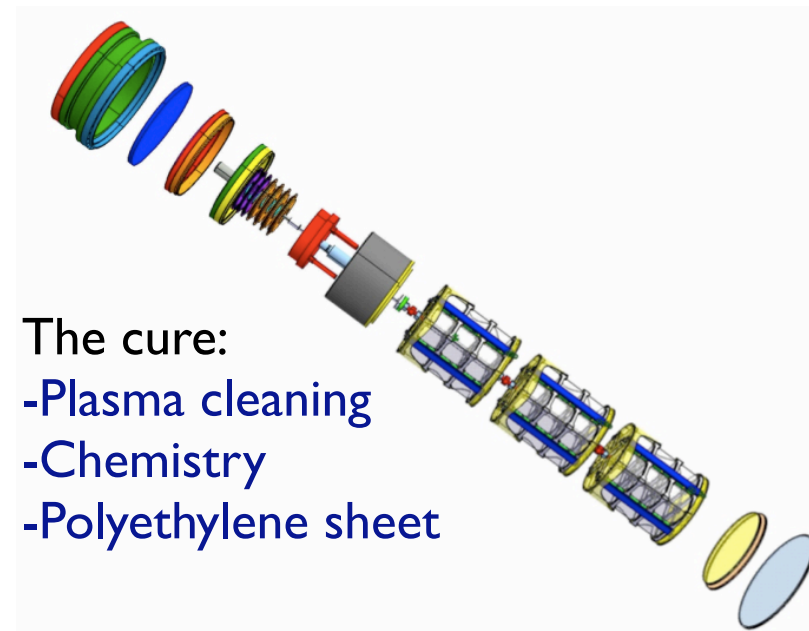
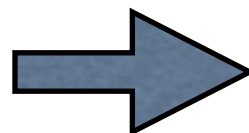
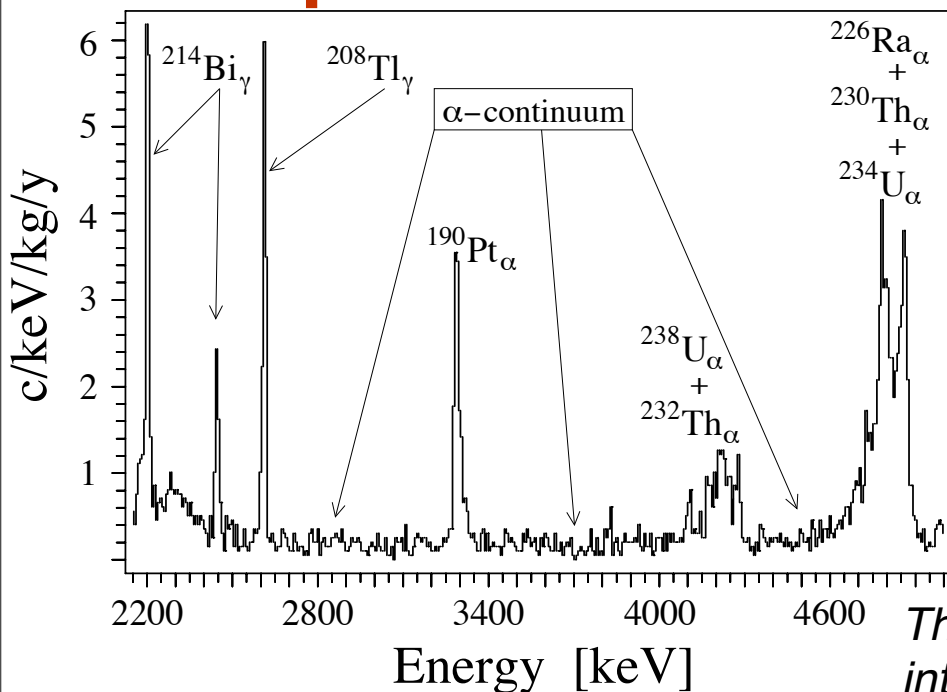
$$B = b / 20$$

$$\Rightarrow 0.01 \text{ counts/keV/kg/y}$$

# The way to CUORE

F.Bellini, A.Buccheri, C.Cosmelli, I.Dafinei, R.Faccini, F.Ferroni, C.Gargiulo  
E.Longo, S.Morganti, F.Orio, V.Pettinacci, C.Tomei, M.Vignati

## The problem: $\alpha$



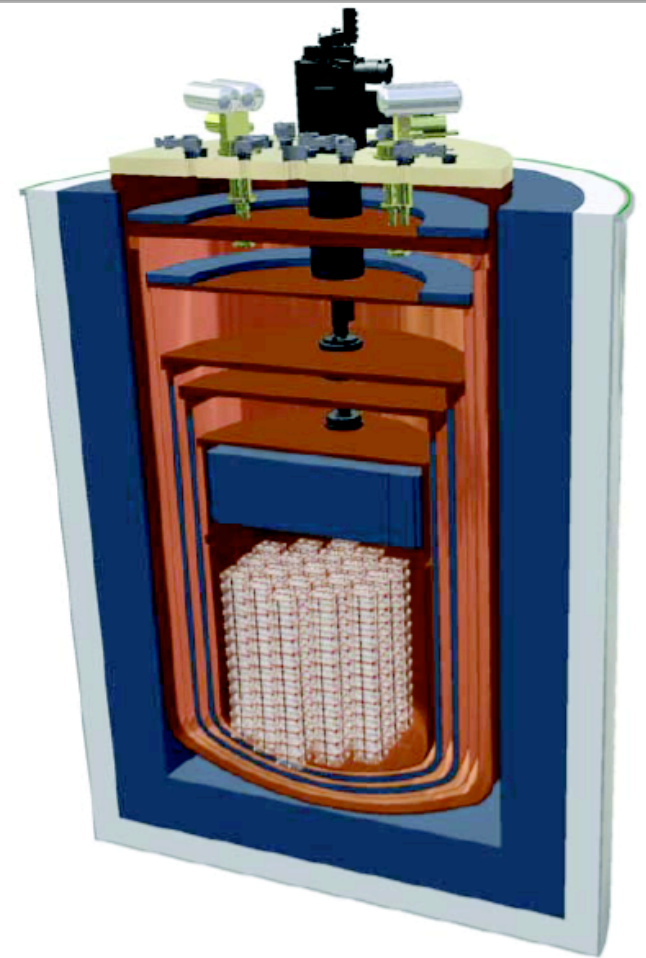
The cure:  
-Plasma cleaning  
-Chemistry  
-Polyethylene sheet

*The chosen one will hold the new(56) CUORE crystals into the CUORICINO cryostat and CUORE-0 will start(2010)*

**Rome activities:**  
Engineering of CUORE(0) detector construction  
Crystal production  
Construction/installation/data taking/data analysis  
Bolometer response function development  
Alternative CUORE calibration system study  
Dark matter studies

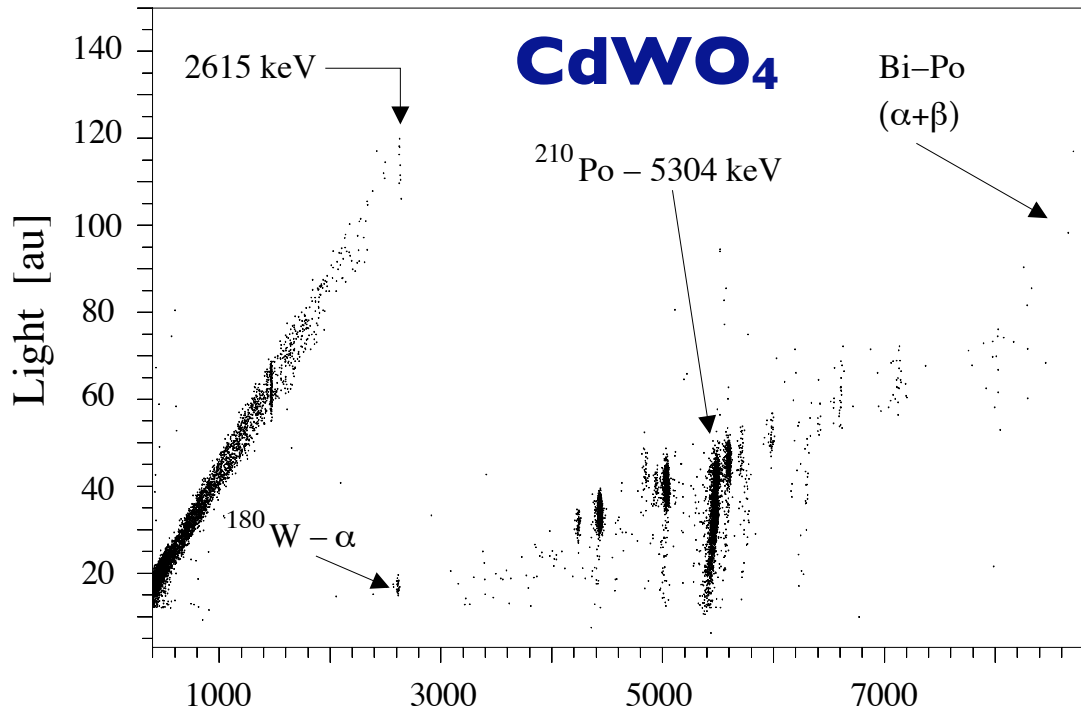
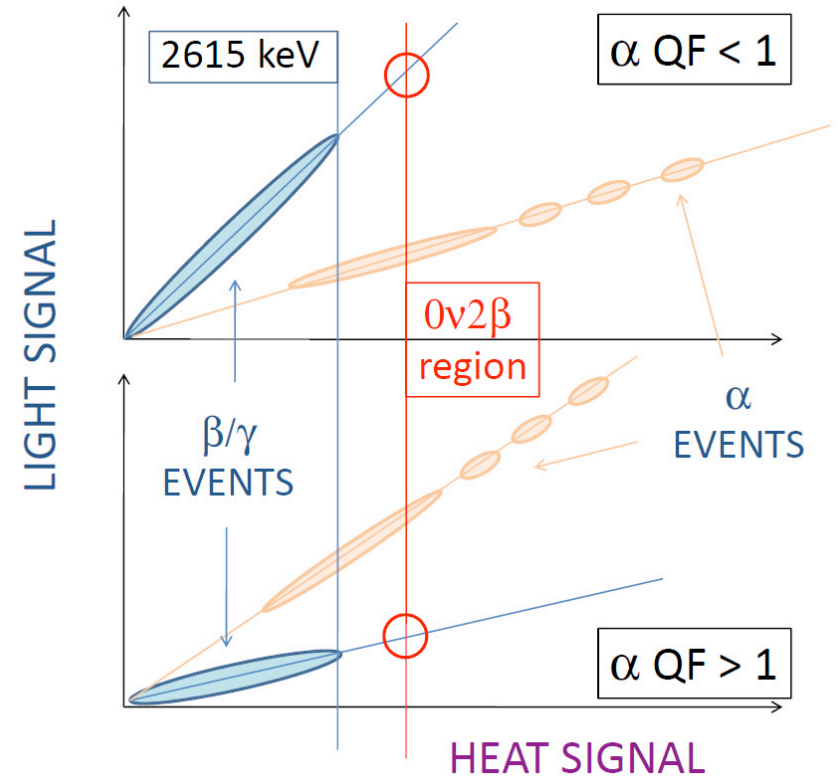
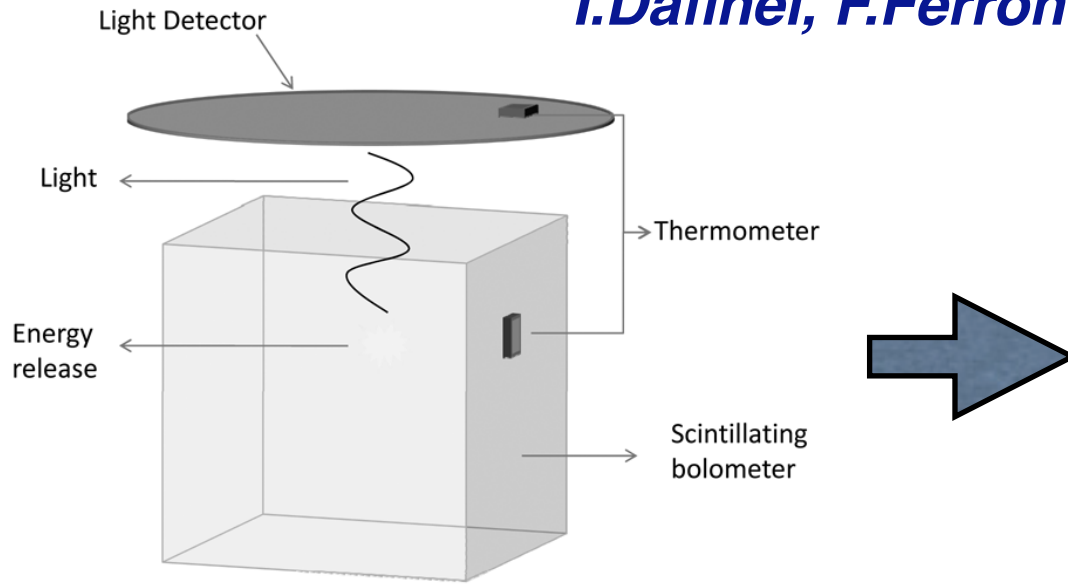
# CUORE and competitors

- **CUORE**: 200 kg  $^{130}\text{Te}$  (2012)
  - ▶  $\tau > 2.1 \cdot 10^{26} \Leftrightarrow 0.02 < m_\nu < 0.1 \text{ eV}$
- **GERDA @ LNGS**: naked  $^{76}\text{Ge}$  in high purified Ar
  - ▶ Phase I: 15 kg  $\tau > 2 \cdot 10^{25}$  in 1 y (end 2009)
  - ▶ Phase II(III): 40(1000) kg enriched  
 $\tau > 2 \cdot 10^{26} \Leftrightarrow 0.07 < m_\nu < 0.3 \text{ eV}$  in 3 y
- **SUPERNEMO(R&D)**: improved tracking detectors(2013)
  - ▶ 100-20 kg  $^{82}\text{Se}$   $\tau > 1-2 \cdot 10^{26} \Leftrightarrow 0.04 < m_\nu < 0.1 \text{ eV}$
- **EXO**: Liquid enriched  $^{136}\text{Xe}$  TPC
  - ▶ phase I: 200 kg  $\tau > 2 \cdot 10^{25} \Leftrightarrow 0.1 < m_\nu < 0.2 \text{ eV}$
  - ▶ phase II: 1 Ton (Ba<sup>++</sup> tag)  $\tau > 8 \cdot 10^{26} \Leftrightarrow 0.03 < m_\nu < 0.06 \text{ eV}$



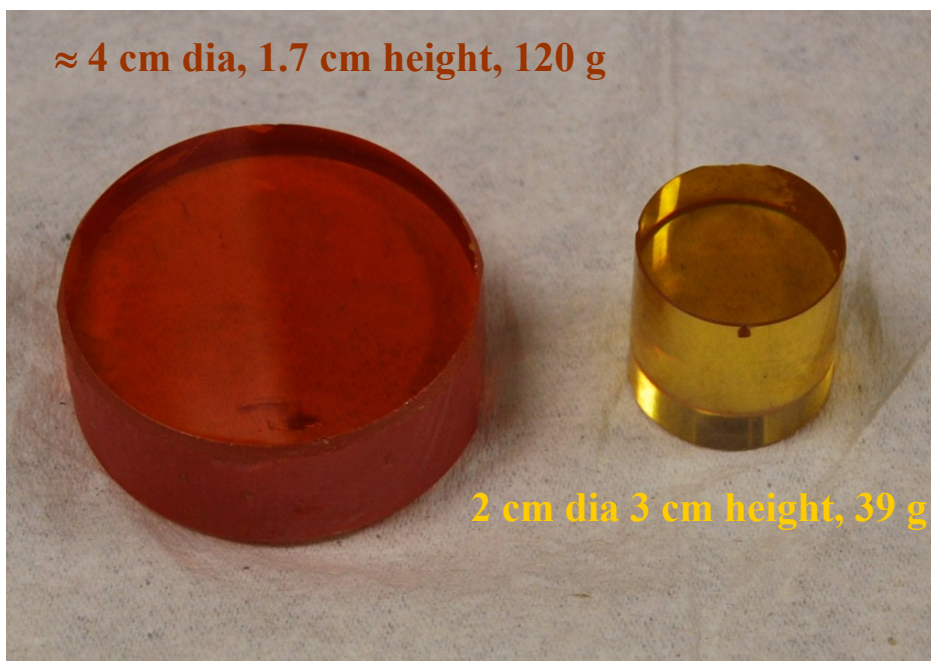
# Bolux: double read-out

*I.Dafinei, F.Ferroni, F.Orio*



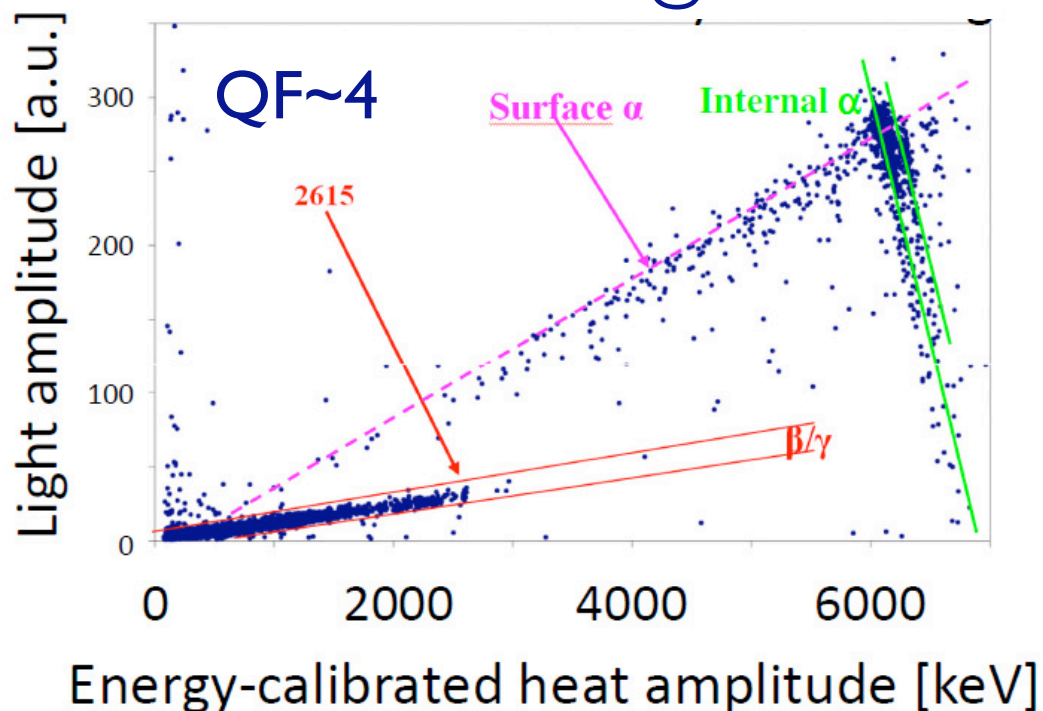
**Bolux @ CSN5**  
**Stefano Pirro @ L.N.G.S.**  
 $\delta E(\text{light detector}) \sim 250 \text{ eV} @ 6 \text{ keV}$

# A promising candidate: ZnSe



Preliminary:  
α rejection efficiency ~99%  
exp. bkgd ~  $10^{-3/4}$  counts/keV/kg/y

Stefano Pirro @ LNGS



## **Rome activities:**

- development and characterization of ZnSe scintillating crystals
- construction and data taking/analysis of scintillating bolometers @LNGS

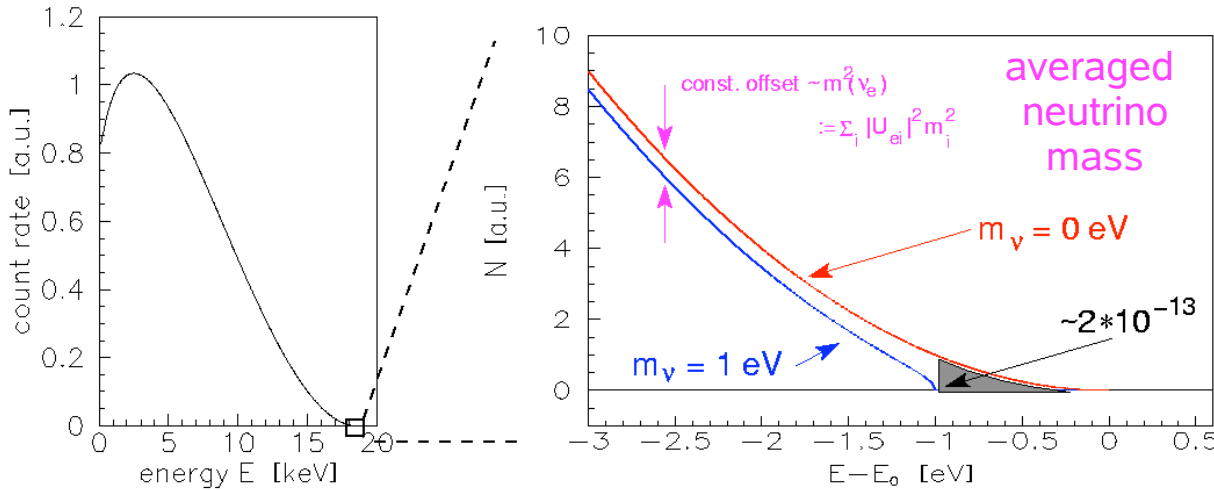
# Conclusions

- **Neutrino physics is one of the leading field in HEP today**
- **Oscillation:**
  - ▶ **2009:** the generation of experiment optimize to detect  $\theta_{13}$  is starting:  
**Double Chooz, T2K**, (soon after **Daya Bay**, then **NOvA**)
  - ▶ **Neutrino oscillation future** dictated by  $\theta_{13}$  value
  - ▶ **Accelerator neutrino beam** might have a bright future
- **Dirac or Majorana** nature is a fundamental question that needs to be answered
  - ▶ **The second generation experiment** could win or might show the path to victory
  - ▶ It could represent the only chance to give a measure of the **neutrino mass**

Thanks to F.Ferroni, P.Loverre, L.Ludovici, S.Pirro, G.Rosa

# $m_\nu$ from beta decay: MARE

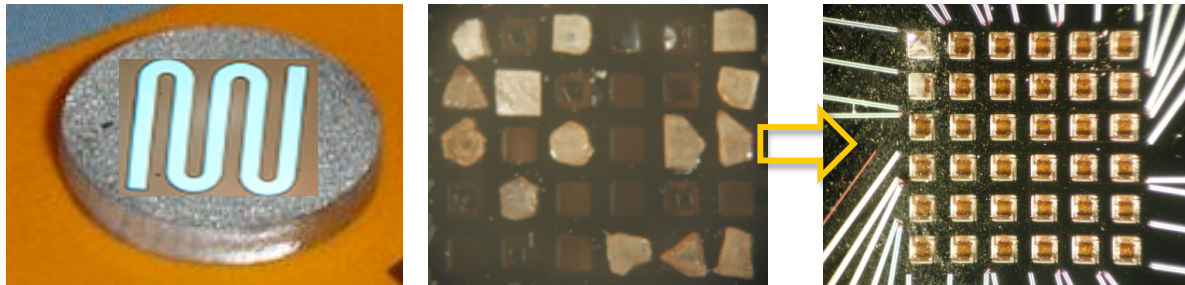
P. De Bernardis, M. Calvo, S. Masi



$$m(\nu_e) = \sqrt{\sum_i |U_{ei}^2| \cdot m_i^2}$$

Previous results:  $m_\nu < 15$  eV

- **Rhenium Bolometer array:**  $^{187}\text{Re} \rightarrow ^{187}\text{Os} e^- \bar{\nu}$  ( $Q=2.47$  keV,  $\tau_{1/2} \sim 43.2$  Gy)
- Phase I:  $\Delta E \sim 5$  eV,  $\Delta t \sim 50 \mu\text{s} \Rightarrow m_\nu \sim 0.2$  eV



Rome R&D:  
Kinetic  
Inductance  
detector

- Phase II: magnetic micro-calorimeters R&D:  $\Delta E \sim 5$  eV  $\Delta t \sim 1 \mu\text{s} \Rightarrow m_\nu \sim 0.2$  eV
- **KATRIN:** the ultimate spectrometer  $m_\nu \sim 0.2$  eV