

## On the "unreasonable" effectiveness of Homogeneous Transport of Intensity imaging



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- Common knowledge tells us that any low-pass filtering can increase signalto-noise ratio (SNR) or contrast-to-noise ratio (CNR) in an image, but it (almost) inevitably comes together with image blurring (unless some *a priori* information is used in the process, either explicitly or implicitly).
- The trade-off between SNR/CNR and spatial resolution has been recently expressed mathematically in the form of the noise-resolution uncertainty (NRU) principle, which is inherent to optics, scattering and even QED.
- In spite of this, the reconstructive imaging with Homogeneous Transport of Intensity equation is capable of "magically" increasing SNR (or CNR) while preserving the spatial resolution. This capability has been conclusively demonstrated both theoretically and experimentally.

<u>Question</u>: what is the solution of this apparent paradox?

Answer: will be given in this talk!



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## **Denoising: better SNR => less sharpness**





Image with 20% Poisson noise SNR=5  $n_{\rm phot}$ =25 pp 9-pixel wide median filter SNR=32

 $n_{\rm phot}$ =25 pp





5-pixel wide Gaussian filter SNR=40  $n_{\rm phot}$ =25 pp

9-pixel wide average filter SNR=43  $n_{\rm phot}$ =25 pp

Increase in SNR comes at the expense of image sharpness



## **Sharpening: better sharpness => lower SNR**



10-pixel blur 1% noise SNR=60  $\Delta X$ =10 pix





Wiener deconv. reg.=0.52 (optim.) SNR=47  $\Delta X$ =7 pix ( $n_{\text{phot}}$  unchanged)



Richardson-Lucy iter.=214 SNR=54  $\Delta X$ =8 pix  $(n_{phot}$  unchanged)



Iterative Wiener reg.=1, iter.=2 SNR=47  $\Delta X$ =7 pix ( $n_{phot}$  unchanged) Increase in image sharpness always comes at the expense of SNR (increased noise)



# Duality between signal-to-noise and spatial resolution under the fixed dose condition

### Near field imaging (geometrical optics)



The ratio of SNR<sup>2</sup> to spatial resolution depends only on the photon fluence (if the photon statistics is fixed).



If the total number of quanta (e.g. photons) forming an image is fixed, then:

1) increase in SNR comes at the expense of image sharpness (i.e. blurring increases = number of resolvable units decreases = spatial resolution deteriorates)

2) increase in image sharpness (increase in the number of resolvable units = improvement in the spatial resolution) comes at the expense of SNR

Can this "noise-resolution" trade-off in imaging be formulated quantitatively in a general form?



# Intrinsic quality characteristic of imaging systems

We have introduced a dimensionless "intrinsic imaging quality characteristic"  $Q_S$  which incorporates both the average signal-to-noise ratio and the spatial resolution of an imaging system:

$$Q_{S}^{2} = \frac{SNR^{2}(\mathbf{x})}{\overline{S}_{in}(\mathbf{x})(\Delta x)^{d}} = < \dots \times \frac{A^{d}}{A^{d}} > = \frac{M_{pix} < SNR^{2} >}{N_{phot}}$$

where

- $N_{phot} = \langle \bar{S}_{in} \rangle A^d$  is the total number of incident photons or other imaging quanta
- $M_{pix} = (A / \Delta x)^d$  is the effective number of pixels (spatial resolution units)
- *d* is the dimensionality of the data (*d*=2 corresponds to planar images)
- *A* is the linear size of the image (illuminated aperture)

Compare this with the information capacity per imaging particle (a quantity introduced following the approach used by Shannon, Gabor and others):

$$C_s \cong 0.5(M_{pix} / N_{phot}) \log_2 < SNR^2 >$$



Noise – resolution uncertainty principle

## Mathematical result proved:

$$Q_{S}[T] \leq C_{d}^{-1/2}$$

Unlike the case of the Heisenberg uncertainty principle, where the minimal phase-space volume is achieved for Gaussian distributions, the minimum of the new uncertainty functional is achieved for the Epanechnikov kernel (which is well known in statistics):  $T_{Ep}(\mathbf{x}) = (1 - |\mathbf{x}|^2)_+$ 



The absolute constants  $C_d$  are smaller than one, which means, that  $Q_S^{(d)}$  can be slightly larger than 1, for example for d=1,2,3 we have:

$$C_1^{-1/2} \cong 1.025$$
  $C_2^{-1/2} \cong 1.061$   $C_3^{-1/2} \cong 1.104$   
 $Q_S[T_{Gaussian}] = 1 < Q_S[T_{Ep}] = C_d^{-1/2}$ 



$$\Delta x = \left(\frac{4\pi}{d} \frac{\int |\mathbf{x}|^2 P(\mathbf{x}) d\mathbf{x}}{\int P(\mathbf{x}) d\mathbf{x}}\right)^{1/2}$$

- conventional measure of the width of a function; normalization  $(4\pi / d)$  is consistent with the width of a rectangular function

d = 1, 2, 3... is the dimensionality of the image



- has the meaning of spatial resolution (dimensionality of length) as a measure of the width of the PSF  $P(\mathbf{x})$  (which can depend on the source, the detector, the geometrical parameters, ...)

Compare this with the definition of the spectral width (Mandel-Wolf):

 $\Delta_2 v = \frac{(\int s(v) dv)^2}{\int s^2(v) dv}$  and the corresponding definition of coherence time:  $T_c = 1/\Delta v$ 



## **Two definitions of spatial resolution**

PSF	$P(\mathbf{x}; \sigma)$	$\hat{P}(\mathbf{u};\sigma)$	$(\Delta x)^2$	$(\Delta_2 x)^2$	$Q_{\scriptscriptstyle S}^{\scriptscriptstyle 4/d}$
<i>d</i> -dim Gauss	$\frac{1}{\left(2\pi\right)^{d/2}\sigma^d}e^{-\frac{ \mathbf{x} ^2}{2\sigma^2}}$	$e^{-2\pi^2\sigma^2 \mathbf{u} ^2}$	$4\pi\sigma^2$	$4\pi\sigma^2$	1
<i>d</i> -dim Rect	$\frac{1}{\left(2\sigma\right)^{d}}\chi_{\left[-\sigma,\sigma\right]^{d}}(\mathbf{x})$	$\prod_{k=1}^d \frac{\sin(2\pi\sigma u_k)}{2\pi\sigma u_k}$	$\frac{4\pi}{3}\sigma^2$	$4\sigma^2$	$\frac{3}{\pi}$
Exp	$(\pi  /  \sigma) e^{-2\pi  x /\sigma}$	$\frac{1}{1+\sigma^2 u^2}$	$\frac{2}{\pi}\sigma^2$	$\frac{4}{\pi^2}\sigma^2$	$\frac{2}{\pi}$
Lorentz	$\frac{\pi^{-1}\sigma}{\sigma^2 + x^2}$	$e^{-2\pi\sigma u }$	<mark>∞</mark>	$4\pi^2\sigma^2$	0
Sech	$\frac{\sigma^{-1}}{e^{\pi x/2\sigma} + e^{-\pi x/2\sigma}}$	$\frac{2}{e^{2\pi\sigma u}+e^{-2\pi\sigma u}}$	$4\pi\sigma^2$	$\pi^2 \sigma^2$	$\frac{\pi}{4}$
Epanechnikov <sup>a)</sup>	$\frac{B_d}{\pi^{d/2}\sigma^d} \left(1 - \frac{ \mathbf{x} ^2}{\sigma^2}\right)_+$	$B_{d} \frac{J_{d/2+1}(2\pi\sigma \mid \mathbf{u} \mid)}{(\pi\sigma \mid \mathbf{u} \mid)^{d/2+1}}$	$\frac{4\pi}{d+4}\sigma^2$	$\frac{4\pi\sigma^2}{(d+4)C_d^{2/d}}$	$rac{1}{C_d^{2/d}}$

<sup>a)</sup>  $B_d = (d/2+1)\Gamma(d/2+1)$   $C_d = 2^d \Gamma(d/2)d(d+2)/(d+4)^{d/2+1} < 1$ 



## **Noise / resolution invariance**

$$S(\mathbf{x}) = \int P(\mathbf{x} - \mathbf{y}) S_{in}(\mathbf{y}) d\mathbf{y}$$

- action of an LSI system on the input fluence

$$SNR^{2}(\mathbf{x}) = \frac{\overline{S}^{2}(\mathbf{x})}{\sigma^{2}(\mathbf{x})} = SNR_{in}^{2}(\mathbf{x})\frac{(\Delta_{2}x)^{d}}{(\Delta_{2}x)_{in}^{d}}$$

- effect of LSI system on the input  $SNR^2$ 



- the ratio of  $SNR^2$  and the spatial resolution volume is invariant with respect to the action of LSI transformations, if  $P(\mathbf{x})$  is much narrower than  $SNR^2(\mathbf{x})$ , but much wider than the noise correlation, i.e. if  $P(\mathbf{x})$  effectively filters the noise, but does not smear the input signal (best case scenario).

$$M_{pix} SNR^{2}(\mathbf{x}) = M_{pix,in} SNR_{in}^{2}(\mathbf{x}) \qquad M_{pix} = A^{d} / (\Delta_{2}x)^{d} \qquad \text{- the product of } SNR^{2} \text{ and the number of resolution units } M_{pix} \text{ is a } (\text{dimensionless!}) \text{ invariant.}$$



## Application to idealized generic "black-box" scattering (imaging) experiment



## Assumed:

- Uniform coherent illumination (incident photons are uncorrelated)
- Weak elastic single-scattering interaction,  $\gamma(\mathbf{x}) = \sigma \rho(\mathbf{x})$ .
- Ideal position-sensitive dark-field detection (only scattered photons enter the detector)

The image noise and the spatial resolution are determined primarily by the number and correlation length of detected photons scattered by each elementary "voxel" of the sample!



## Information capacity of a scattering experiment (source + sample + detector)

$$\left\langle SNR^{2} \right\rangle = \overline{S}(x)(\Delta_{2}x)^{d} = \overline{\gamma}A \overline{S_{0}^{(d-1)}A^{d-1}} \underbrace{\left(\Delta_{2}x\right)^{d}}_{A^{d}} = \frac{\Omega N_{phot}}{M_{vox}} \qquad \begin{array}{l} \text{Registered SNR}^{2} \text{ is proportional to the scattering power } \Omega = \gamma A \text{ of the sample } (\text{dark field}) \text{ or its square } (\text{bright field}) \end{array}$$

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The information capacity of the considered scattering setup is equal to the number of spatial resolution units  $M_{vox}$  times the log<sub>2</sub> of the number of scattered photons  $\Omega N_{phot}$  per unit. In other words, the total number of distinguishable samples is approximately  $(\Omega N_{phot} / M_{vox})^{M_{vox}/2}$ . This number is maximized, when  $M_{vox}$  is large, i.e. for low SNR.

$$< SNR_{CT}^2 > \cong \frac{2\Omega^2 N_{phot}}{M_{vox}^{4/3}}$$

Unlike the simple generic scheme considered here, the registered SNR<sup>2</sup> in CT is proportional to the inverse 4<sup>th</sup>(!) power of the spatial resolution (noise increases stronger because of the singularity of the inverse Radon transform).

## **Relationship to Shannon's information capacity**

- C. Shannon (1949): Maximum number of bits,  $N_{bits}$ , that can be transmitted within a time interval  $A_t$  over a communication channel with bandwidth  $1/\Delta T$  is limited by  $N_{bits} \le \frac{A_t}{\Delta T} \log SNR$
- Our result for 1D case (noise resolution uncertainty principle) is complementary to the above one:

$$\frac{A_x}{\Delta X}SNR^2 \le C_1^{-1} N_{photons}$$

• Combining these results, we can obtain in particular that the information capacity of a communication channel or an imaging system is ultimately limited by the number of signal quanta used:

$$N_{bits} \leq M_{res.units} \log SNR \leq M_{res.units} SNR^2 \leq C_d^{-1} N_{photons}$$

• This relationship provides a link between the information capacity, spatial resolution and SNR (incl. radiation dose) in imaging systems. Note, however, that in principle the right-hand side contains also a quantity  $O(M_{res.units})$ !



- A duality exists between noise and spatial resolution in LSI imaging systems.
- This duality ("noise-resolution uncertainty") applies also to (some) deconvolution operations and system transformations described by PDEs with constant coefficients. The ratio of SNR and spatial resolution  $(\Delta_2 x)^{d/2}$  is invariant with respect to such transformations.
- Noise-resolution duality is closely related to (Shannon's) information capacity of imaging and communication systems.
- Noise-resolution uncertainty also exists in quantum electrodynamics, where the ultimate lower limit of the product of the spatial width of a mode of an electromagnetic field and the variance of its intensity is determined by vacuum fluctuations.







Same area of an axial CT slice of an excised breast tissue sample obtained using conventional CT (left) and in-line phase-contrast CT with TIE-Hom phase retrieval (right) at the same X-ray dose. Factor of x20 improvement in SNR demonstrated, equivalent to x400 reduction in X-ray dose at the same SNR and spatial resolution as in the equivalent absorption image.



# How can we explain the "unreasonable" effectiveness of the TIE-Hom imaging

The problem:

1) On one hand, the ratio of SNR to spatial resolution cannot change in any intensity-linear transformation (including TIE-Hom propagation, and TIE-Hom phase retrieval).

2) On the other hand, it has been repeatedly demonstrated in experiments, that a combination of near-field free-space propagation between the object and the detector, with TIE-Hom phase retrieval is capable of increasing SNR by a factor of up to x100s, while not worsening (or only moderately worsening) the spatial resolution.



"Contact" image



Propagated image with TIE-Hom phase retrieval



Variance of exposure-averaged intensity ("self" noise) 1D case expressions considered for simplicity

$$V_T(x,R) \equiv < I_T^2(x,R) > - < I_T(x,R) >^2$$

- definition of (noise) variance

exposure time average ensemble average  

$$< I_T(x,R) > = T^{-1} \int_{-T/2}^{T/2} < |U(x,R;t)|^2 > dt = I(x,R)$$

- the photon flux is assumed to be stationary

$$< I_T^2(x,R) > = < \left(T^{-1} \int_{-T/2}^{T/2} |U(x,R;t)|^2 dt\right)^2 > \neq I^2(x,R)$$

$$V_T^{Poisson}(x, R) = < I_T^2(x, R) > - < I_T(x, R) >^2 = I(x, R)$$

$$V_T^{Gauss}(x,R) = \langle I_T^2(x,R) \rangle - \langle I_T(x,R) \rangle^2 = I^2(x,R)$$

- Gaussian statistics (e.g. "intrinsic" or "self" variance of light emitted by thermal light sources)



Variance of detected(!) time-averaged intensity in a plane z = R:

$$V_{P,T}(x,R) = V_{P,T}^{self}(x,R) + V_{P,T}^{det}(x,R) \cong T^{-1} \int S^2(x,R;\nu) d\nu + (\Delta_2 x)^{-1} \int \eta^{-1}(\nu) S(x,R;\nu) d\nu$$

 $\Delta_2 x = 1/\int P^2(x)dx$  - width of the PSF (~1/2 of the correlation length)  $\eta(v)$  - detector (quantum) efficiency constant (photons/Joules)

S(x, R; v) = I(x; R) s(v) - cross-spectrally pure case (for simplicity)

 $\int s(v)dv = 1$ ,  $\int s^2(v)dv = T_c$  - coherence time (inverse of the spectral width)

$$V_{P,T}(x,R) \cong (T_c / T)I^2(x,R) + T^{-1}\eta^{-1}(\Delta_2 x)^{-1}I(x,R)$$
  
Gaussian self-  
noise term  
Shot-noise term



The source (self) noise term is typically much (much) smaller than the detector (shot) noise term

$$V_{P,T}(x,R) \cong (T_c/T)I^2(x,R) + T^{-1}\eta^{-1}(\Delta_2 x)^{-1}I(x,R) - J^2/m^2/s^2$$

$$Gaussian \ self-$$

$$noise \ term$$

$$noise \ term$$

$$I^2\eta^2(\Delta_2 x)^2 \times V_{P,T}(x,R) \cong (T_c/T)[\overline{n}(x)]^2 + \overline{n}(x)$$
- dimensionless version

 $\overline{n}(x) = \eta(\Delta_2 x)T I_R(x)$  is the average number of photons registered within the "pixel" with area  $\Delta_2 x$  centred at point *x*, during the exposure time *T*.

Note that for X-rays with  $\lambda = 1$ Å and monochromaticity  $\Delta\lambda/\lambda = 10^{-4}$ , the coherence time  $T_c \sim 1/\Delta v = \lambda^2/(c\Delta\lambda) \sim 0.3 \times 10^{-15}$ s! Therefore, if the exposure time is a few ms, then, unless there are  $10^{12}(!)$  or more photons registered per pixel(!) per exposure(!), the self noise term is much smaller than the shot noise term. Therefore,

$$V_{P,T}(x,R) \cong T^{-1}\eta^{-1}(\Delta_2 x)^{-1}I(x,R)$$

Image noise is well approximated by the shot noise alone!



$$V_{P,T}(x,R) \cong T^{-1}\eta^{-1}(\Delta_2 x)^{-1} \int \overline{S}(x,R;\nu) d\nu$$

- photodetection shot noise (assuming spatial ergodicity)

 $S(x, R; v) \cong [1 - \sigma^2 \nabla_x^2] S(x, 0; v)$  - homogeneous TIE (TIE-Hom)

 $\sigma^2 = \gamma R' \lambda \,/\, (4\pi)$ 

$$\left|\frac{\sigma^2 \nabla^2 S(x,0;\nu)}{S(x,0;\nu)}\right| = \left|1 - \frac{S(x,R;\nu)}{S(x,0;\nu)}\right| <<1 \quad \text{-TIE validity condition}$$
  
$$\overline{S}(x,R;\nu) \cong \overline{S}(x,0;\nu) \implies V_{P,T}(x,R) \cong V_{P,T}(x,0) \quad SNR(x,R) = \frac{\overline{I}(x,R)}{V_{P,T}^{1/2}(x,R)} \cong \frac{\overline{I}(x,0)}{V_{P,T}^{1/2}(x,0)}$$

Variance of image noise and SNR do not change upon forward free-space propagation in the near-Fresnel region!



## Image noise decreases and SNR increases upon TIE-Hom retrieval

 $S_{ret}(x,0;\nu) = \int P_{\sigma}(x-x')S(x',R;\nu)dx'$  - TIE-Hom "phase" retrieval

 $P_{\sigma}(x) = \int \frac{\exp(-i2\pi x\xi)}{1 + 4\pi^2 \sigma^2 \xi^2} d\xi = \frac{1}{2\sigma} e^{-|x|/\sigma} - 1\text{D case (2D and 3D - qualitatively similar)}$ 

$$V_{P,T}^{ret}(x,0) = V_{P_{\sigma},P,T}(x,R) = \frac{(\Delta_2 x)}{4\sigma} V_{P,T}(x,R)$$

$$\frac{\Delta_2 x}{4\sigma} = \left(\frac{\pi h^2}{4\gamma\lambda R'}\right)^{1/2} = \left(\frac{\pi}{4}\frac{N_F^{\min}}{\gamma}\right)^{1/2} \equiv G^{-1/2} \ll 1 \quad \text{- provided } N_F \ll \gamma$$

$$SNR_{ret} = \frac{\overline{I}_{ret}(x,0)}{(V_{P,T}^{ret})^{1/2}(x,0)} \cong G^{1/4} \frac{\overline{I}(x,0)}{V_{P,T}^{1/2}(x,0)} \equiv G^{1/4} SNR_0$$

Variance of image noise decreases and SNR increases upon TIE-Hom (phase) retrieval! The effect is stronger in 2D ( $\sim G^{1/2}$ ), and still stronger in 3D ( $\sim G$ ).



Spatial resolution improves upon forward propagation, and then deteriorates upon TIE-Hom retrieval

 $(\Delta x)[P] = (4\pi \int x^2 P(x) dx)^{1/2}$  - spatial resolution (defined as *std* of the PSF)

$$4\pi \int x^2 (1 - \sigma^2 \nabla_x^2) P(x) dx = (\Delta x)_0^2 [P] - 4\pi \sigma^2 \int (\nabla_x^2 x^2) P(x) dx$$

$$(\Delta x)_{R}^{2}[P] = (\Delta x)_{0}^{2}[(1 - \sigma^{2} \nabla_{x}^{2})P] = (\Delta x)_{0}^{2}[P] - 8\pi\sigma^{2}$$

Spatial resolution improves upon "forward" free-space propagation

$$(\Delta x)_{0,ret}^2[P] = (\Delta x)_R^2[(1 - \sigma^2 \nabla_x^2)^{-1}P] = (\Delta x)_R^2[P] + 8\pi\sigma^2 = (\Delta x)_0^2[P]$$

Spatial resolution deteriorates upon TIE-Hom retrieval, back to the level it had in the contact plane z=0.



## "Unreasonable" effectiveness of TIE-Hom imaging explained at last

 $h^2$ 



Total effect: 
$$\frac{SNR_{ret}(x,0)}{(\Delta x)_{0,ret}} = G^{1/4} \frac{SNR(x,0)}{(\Delta x)_0} \qquad \qquad G \cong \frac{4}{\pi}$$

<u>Total effect</u>: the ratio of SNR and spatial resolution increases in TIE-Hom imaging. Note that all the "magic" happens during free-space propagation, phase retrieval does not do anything really special! This "magic" is possible only because the image noise is dominated by the photodection noise.



#### SNR gains for ideal detector with a single-pixel PSF.

$h \bigcup R \Longrightarrow$	1.5m	2m	6m	9m
30µm	24	31	80	114
50µm	11	<u>14</u>	33	47
70µm	6	8	19	26
100µm	4	<u>6</u>	11	15

#### E=32keV

 $\gamma(E)$  corresponds to glandular relative to adipose tissue:

$$\gamma = \frac{\delta_{gland} - \delta_{fat}}{\beta_{gland} - \beta_{fat}}$$

#### SNR gains for detector with a Gaussian PSF, std $\sigma$ = pixel size.

$h \downarrow   R \Longrightarrow$	1.5m	2m	6m	9m	SNR(x,0)
30µm	4	8	15	20	$Gain = \frac{Gain = Sint(x, 0)}{SNR(x, 0)} \sim$
50µm	2	4	7	10	$C = \frac{4 \gamma \lambda R'}{2}$
70µm	1.8	2	4	6	$G = \frac{1}{\pi} \frac{1}{h^2}$
100µm	1.6	1.8	3	4	



### DRFs for an ideal detector with a single-pixel PSF.

$h \bigcup R \Longrightarrow$	1.5m	2m	6m	9m
30µm	576	961	6400	13000(!)
50µm	121	<u>196</u>	1089	2209
70µm	36	64	361	676
100µm	16	<u>36</u>	121	225

#### E=32keV

 $\gamma(E)$  corresponds to glandular relative to adipose tissue:

$$\gamma = \frac{\delta_{gland} - \delta_{fat}}{\beta_{gland} - \beta_{fat}}$$

#### DRFs for a detector with a Gaussian PSF, std $\sigma$ = pixel size.

$h \bigcup   R \Longrightarrow$	1.5m	2m	<b>6</b> m	9m	Dose(x,0)
30µm	16	64	225	400	$DRF = \frac{\langle y \rangle}{Dose_{ret}(x,0)} \sim G^2$
50µm	4	16	49	100	$C = 4 \gamma \lambda R'$
70µm	3.2	4	16	36	$G = \frac{\pi}{\pi} \frac{h^2}{h^2}$
100µm	2.6	3.2	9	16	



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Advantage of short exposures and large number of projections (from *Kitchen et al, arXiv:1704.03556, 2017*)



TIE-Hom imaging of mouse lungs as a function of exposure time (dose) and propagation distance (phase retrieval).

X-ray energy E = 24 keV Detector pixel size h = 15.3 µm  $\gamma = 1773$ 

Dose Reduction Factor of 27600 demonstrated at a fixed SNR and spatial resolution.

At these *E*,*h*,*R* and  $\gamma$ , theoretical SNR Gain Factor for an ideal detector is 243, corresponding to DRF = 59000!



- TIE-Hom imaging is a unique method which allows one to increase SNR in images at a fixed radiation dose, without a loss of spatial resolution.
- The "magic" of TIE-Hom imaging happens during the forward freespace propagation, which improves the spatial resolution (reduces the photon correlation length) without increasing the noise variance.
- This "magic" is only possible because the image noise is dominated by photodetection shot noise, which is "added" only at the image detection stage and does not depend on the propagation distance.
- With good photon-counting detectors (high DQE, relatively small pixels, single-pixel PSF), SNR gain factors of the order of 10<sup>2</sup> may be possible, which translates into dose reduction factors of the order of 10<sup>4</sup>, in PB-CT of breast tissue.



## The end