# Exploring for New Physics using Charged Lepton Flavor Violation

Doug Glenzinski Fermilab May 2018

# Outline

- Motivation & Introduction
- Experimental Summary
- Future Expectations
- Summary

# Why Charged Lepton Flavor Violation (CLFV)?

- Quarks mix, v mix... what about I<sup>+</sup>?
   CLFV : neutrino-less transitions of the type μ → e, τ → e, τ → μ
- There is no known Global Symmetry that requires LF conservation
- Many extensions to the Standard Model predict large CLFV effects
- CLFV offers opportunity to probe  $\Lambda_{\rm NP} \sim O(10^3 10^4)$  TeV >> TeV



#### Some CLFV Processes

Process	Current Limit	Next Generation exp
$\tau \rightarrow \mu \eta$	BR < 6.5 E-8	
$\tau \rightarrow \mu \gamma$	BR < 6.8 E-8	10 <sup>-9</sup> - 10 <sup>-10</sup> (Belle II, LHCb)
$\tau  ightarrow \mu \mu \mu$	BR < 3.2 E-8	
$\tau \rightarrow eee$	BR < 3.6 E-8	
$K_L \rightarrow e \mu$	BR < 4.7 E-12	NA62
$K^+ \rightarrow \pi^+ e^- \mu^+$	BR < 1.3 E-11	
$B^0  ightarrow e \mu$	BR < 7.8 E-8	LHCb, Belle II
$B^+  ightarrow K^+ e \mu$	BR < 9.1 E-8	
$\mu^{+} \rightarrow e^{+}\gamma$	BR < 4.2 E-13	10 <sup>-14</sup> (MEG)
$\mu^+ \rightarrow e^+ e^+ e^-$	BR < 1.0 E-12	10 <sup>-16</sup> (PSI)
μ⁻N → e⁻N	R <sub>μe</sub> < 7.0 E-13	10 <sup>-17</sup> (Mu2e, COMET)

(current limits from the PDG)

#### Expect significant progress in near future

#### Some CLFV Processes

Process	Current Limit	Next Generation exp
$\tau \rightarrow \mu$ η	BR < 6.5 E-8	
$\tau  ightarrow \mu\gamma$	BR < 6.8 E-8	10 <sup>-9</sup> - 10 <sup>-10</sup> (Belle II, LHCb)
$τ \rightarrow μμμ$	BR < 3.2 E-8	
$\tau \rightarrow eee$	BR < 3.6 E-8	
$K_L \rightarrow e\mu$	BR < 4.7 E-12	NA62
$K^+ \rightarrow \pi^+ e^- \mu^+$	BR < 1.3 E-11	
B⁰ <b>→</b> eµ	BR < 7.8 E-8	LHCb, Belle II
B⁺ → K⁺eµ	BR < 9.1 E-8	
$\mu^+ \rightarrow e^+ \gamma$	BR < 4.2 E-13	10 <sup>-14</sup> (MEG)
$\mu^+ \rightarrow e^+e^+e^-$	BR < 1.0 E-12	10 <sup>-16</sup> (PSI)
μ⁻N → e⁻N	R <sub>μe</sub> < 7.0 E-13	10 <sup>-17</sup> (Mu2e, COMET)
		(current limits from the PDG)

#### Experiments using muons among the most sensitive

# CLFV in the $\nu$ -Standard Model



- Extremely suppressed in the Standard Model : rate ~  $\Delta m_v^4$  /  $M_w^4$  < 10<sup>-50</sup>
- Many New Physics models predict rates observable at next generation CLFV experiments
- No SM pollution : Observation is unambiguously New Physics

May 2018

#### New Physics Contributions to CLFV



#### A broad array of New Physics models contribute to CLFV

May 2018

#### **CLFV** Predictions



The different channels offer complementary sensitivity. Their comparison is a powerful model discriminant.

#### **CLFV** Predictions



The different channels offer complementary sensitivity. Their comparison is a powerful model discriminant.

#### **CLFV** Predictions



The different channels offer complementary sensitivity. Their comparison is a powerful model discriminant.

#### Using CLFV to Determine New Physics

Modol		$uN \rightarrow eN$	${ m BR}(\mu{ ightarrow}eee)$	$CR(\mu N \rightarrow eN)$
	$\mu \rightarrow eee$	$\mu$ iv $\rightarrow$ erv	$BR(\mu \rightarrow e\gamma)$	${ m BR}(\mu  ightarrow e \gamma)$
MSSM	Loop	Loop	$pprox 6  imes 10^{-3}$	$10^{-3} - 10^{-2}$
Type-I seesaw	$\operatorname{Loop}^*$	Loop*	$3\times 10^{-3}-0.3$	0.1 - 10
Type-II seesaw	Tree	Loop	$(0.1-3) imes 10^3$	$\mathcal{O}(10^{-2})$
Type-III seesaw	Tree	Tree	$pprox 10^3$	${\cal O}(10^3)$
LFV Higgs	$\operatorname{Loop}^\dagger$	$\operatorname{Loop}^{*\dagger}$	$pprox 10^{-2}$	$\mathcal{O}(0.1)$
Composite Higgs	$\operatorname{Loop}^*$	$\operatorname{Loop}^*$	0.05-0.5	2 - 20

from L. Calibbi and G. Signorelli, Riv. Nuovo Cimento, 41 (2018) 71

TABLE VII. – Pattern of the relative predictions for the  $\mu \rightarrow e$  processes as predicted in several models (see the text for details). It is indicated whether the dominant contributions to  $\mu \rightarrow eee$  and  $\mu \rightarrow e$  conversion are at the tree or at the loop level; Loop<sup>\*</sup> indicates that there are contributions that dominate over the dipole one, typically giving an enhancement compared to Eq. (40, 41). <sup>†</sup> A tree-level contribution to this process exists but it is subdominant.

#### The relative rates are model dependent

• Their ratios can be used to probe the underlying theory

May 2018

#### **CLFV** Sensitivity

A A = Discovery Sensitivity

	AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?
$\epsilon_K$	*	***	***	*	*	**	***
$S_{\psi\phi}$	***	***	***	*	*	***	***
$S_{\phi K_S}$	***	**	*	***	***	*	?
$A_{\rm CP} \left( B  o X_s \gamma \right)$	*	*	*	***	***	*	?
$A_{7,8}(B\to K^*\mu^+\mu^-)$	*	*	*	***	***	**	?
$A_9(B  ightarrow K^* \mu^+ \mu^-)$	*	*	*	*	*	*	?
$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$B_s \to \mu^+ \mu^-$	***	***	***	***	***	*	*
$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***
$K_L  ightarrow \pi^0  u ar{ u}$	*	*	*	*	*	***	***
$\mu \to e \gamma$	***	***	***	***	***	***	***
$\tau \to \mu \gamma$	***	***	*	***	***	***	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
$d_n$	***	***	***	**	***	*	***
$d_e$	***	***	**	*	***	*	***
$(g-2)_{\mu}$	***	***	**	***	***	*	?

#### W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models  $\star \star \star$  signals large effects,  $\star \star$  visible but small effects and  $\star$  implies that the given model does not predict sizable effects in that observable.

#### **CLFV** Sensitivity

r = Discovery Sensitivity

$d_n$	***	***	***	**	***	*	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
$\tau \rightarrow \mu \gamma$	***	***	*	***	***	***	***
$\mu \rightarrow e \gamma$	***	***	***	***	***	***	***
$K_* \rightarrow \pi^0 \mu \bar{\nu}$	+	+	+	+	+	***	***
$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***
$B_s \to \mu^+ \mu^-$	***	***	***	***	***	*	*
$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$A_9(B  ightarrow K^* \mu^+ \mu^-)$	*	*	*	*	*	*	?
$A_{7,8}(B\to K^*\mu^+\mu^-)$	*	*	*	***	***	**	?
$A_{\rm CP}\left(B \to X_s \gamma\right)$	*	*	*	***	***	*	?
$S_{\phi K_S}$	***	**	*	***	***	*	?
$S_{\psi\phi}$	***	***	***	*	*	***	***
$\epsilon_K$	*	***	***	*	*	**	***
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?

#### W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models  $\star \star \star$  signals large effects,  $\star \star$  visible but small effects and  $\star$  implies that the given model does not predict sizable effects in that observable.

#### The sensitivity of these experiments is exciting and compelling.

May 2018



#### All 3 experiments use the same basic principles



- High intensity, high purity μ source
- High  $\mu$  stopping rate
- Detector to precisely measure particles consistent with having originated from stopping target

May 2018



NB. In these experiments, 100 MeV/c is "high momentum"

	P (MeV/c)	E (MeV)	KE (MeV)		
e :	100	100	100		
μ:	100	145	40		
π:	100	170	30		
D.Glenzinski   Fermilab					

May 2018

# CLFV Experiments using $\mu^-$ : Muon-to-Electron Conversion ( $\mu^-N \rightarrow e^-N$ )



Current State-of-the-art (@ 90% CL):

$$R_{\mu e} = \frac{\Gamma(\mu^{-} Au \rightarrow e^{-} Au)}{\Gamma(\mu^{-} Au \text{ capture})} < 7 \times 10^{-13}$$

W. Bertl, et al. (SINDRUM-II) Eur. Phys. J. C47 (2006) 337.

May 2018



#### Next generation experiments:

– DeeMee (J-PARC, 3 GeV)	x10	
– Mu2e (Fermilab)	x10,000	Expected improvement
– COMET (J-PARC, 8 GeV)		(relative to current state-of-the-art)
<ul> <li>Phase-I</li> </ul>	x10-100	
Phase-II	x10,000	



Mono-energetic electron  $E_{\mu e} = m_{\mu} - B(A,Z) - R(A,Z) \sim 105 \text{ MeV}$ 

Coherent interaction with nucleus

**Background** 

Decay in Orbit (DIO) ( $\mu^-N \rightarrow e^-\nu\nu N$ )

Radiative Pion Capture (RPC)  $(\pi^- N \rightarrow \gamma N' \rightarrow e^- e^+ N')$ 

Cosmogenic

May 2018

# Aside : muonic atoms

- Stopped μ<sup>-</sup> is captured in atomic orbit –Quickly (~fs) cascades to 1s state
- Bohr radius ~20 fm (for aluminum)
  - –Significant overlap of  $\mu^{\scriptscriptstyle \text{-}}$  and N wavefunctions
  - –Lifetime of the  $\mu\text{-}atom$  ~few 100 ns for stopping targets of interest
- Once in orbit, 3 things can happen
  - -Decay :  $\mu^- N(A,Z) \rightarrow e^- \nu \nu N(A,Z)$  (background)
  - -Capture :  $\mu^- N(A,Z) \rightarrow \nu N^*(A, Z-1)$  (normalization)
  - -Conversion :  $\mu^- N(A,Z) \rightarrow e^- N(A,Z)$  (signal)

May 2018

# Aside : muonic atoms

- Stopped μ<sup>-</sup> is captured in atomic orbit –Quickly (~fs) cascades to 1s state
- Bohr radius ~20 fm (for aluminum)
  - –Significant overlap of  $\mu^{-}$  and N wavefunctions
  - –Lifetime of the  $\mu\text{-}atom$  ~few 100 ns for stopping targets of interest
- Once in orbit, 3 things can happen
  - $-\text{Decay}: \mu^- N(A,Z) \rightarrow e^- \nu \nu N(A,Z) (39\%)$
  - -Capture :  $\mu^- N(A,Z) \rightarrow \nu N^*(A, Z-1)$  (61%)

for an aluminum stopping target

# Aside : muonic atoms

- Stopped μ<sup>-</sup> is captured in atomic orbit –Quickly (~fs) cascades to 1s state
- Bohr radius ~20 fm (for aluminum)
  - –Significant overlap of  $\mu^{-}$  and N wavefunctions
  - –Lifetime of the  $\mu\text{-}atom$  ~few 100 ns for stopping targets of interest
- Once in orbit, 3 things can happen – Decay :  $\mu^- N(A,Z) \rightarrow e^- v v N(A,Z)$  (39%)
  - -Capture :  $\mu^- N(A,Z) \rightarrow \nu N^*(A, Z-1)$  (61%)

Produces 1n, 2γ, 0.1p per capture

for an aluminum stopping target

#### One Mu2e Event (500-1695 ns after proton pulse)



#### Timing information helps mitigate this

May 2018

#### Decay in Orbit Background for $(\mu^-N \rightarrow e^-N)$



- $E_e$  follows the Michel spectrum... but with a long tail from nuclear recoil  $E_{max} = E_{\mu e}$ 
  - Requires excellent  $\sigma_p$  (<200 keV/c) & FWHM < 1 MeV/c to suppress

May 2018

### Radiative Pion Capture Background for $(\mu - N \rightarrow e - N)$



- Pions that survive to the stopping target are promptly captured on the nucleus
  - few% of the time, radiate  $\gamma$  with  $E_{\gamma} \sim m_{\mu}$
  - Suppressed by 10<sup>9</sup>-10<sup>10</sup> with pulsed proton beam and utilizing a delayed search window while maintaining a high efficiency for signal (~50%)

May 2018



Background

Decay in Orbit (DIO)  $(\mu^-N \rightarrow e^-\nu\nu N)$ 

Radiative Pion Capture (RPC)  $(\pi^- N \rightarrow \gamma N')$ Cosmogenic

Keys to success: excellent spectrometer resolution, pulsed proton beam, high efficiency cosmic veto

May 2018

#### Proton Beams for Mu2e and COMET





# Proton Beams for Mu2e and COMET





May 2018



about 25 meters end-to-end

#### • Consists of 3 solenoid systems



**Production Solenoid:** 

8 GeV protons interact with a tungsten target to produce  $\mu$ - (from  $\pi$ - decay)

• Consists of 3 solenoid systems

May 2018



Transport Solenoid:

Captures  $\pi$ - and subsequent  $\mu$ -; momentum- and sign-selects beam

#### • Consists of 3 solenoid systems



**Detector Solenoid:** 

Upstream – Al. stopping target, Downstream – tracker, calorimeter (not shown – cosmic ray veto system, proton-beam monitor, stopped-muon monitor)

#### • Consists of 3 solenoid systems

May 2018



Graded fields important to suppress backgrounds, to increase muon yield, and to improve geometric acceptance for signal electrons

#### • Consists of 3 solenoid systems

May 2018



Graded fields important to suppress backgrounds, to increase muon yield, and to improve geometric acceptance for signal electrons

#### • Derived from MELC concept originated by Lobashev and Djilkibaev in 1989

May 2018
### **COMET-II** Apparatus



 Also inspired by Lobashev and Djilkibaev

May 2018

### **COMET Evolution**







### Mu2e Expected Background Yield

(COMET Phase-II very similar)

Category	Source	Events
	$\mu$ Decay in Orbit	0.14
Intrinsic	Radiative $\mu$ Capture	<0.01
	Radiative $\pi$ Capture	0.02
	Beam electrons	<0.01
	$\mu$ Decay in Flight	<0.01
Late Arriving Beam	$\pi$ Decay in Flight	<0.01
	Anti-proton induced	0.04
Miscellaneous	Cosmic Ray induced	0.21
Total Background		0.41

(assuming 6.7E17 stopped muons in 6E7 s of beam time)

#### • Designed to be nearly background free

### Detector development and prototypes



• Experiment designs finalized

Required performance demonstrated in test beams

### **Construction well underway** (Phase-I COMET, Mu2e)



Commissioning begins 2020-2021

# CLFV Experiments using $\mu^+$ : Muon to e + gamma ( $\mu^+ \rightarrow e^+ \gamma$ ) Muon to 3 electrons ( $\mu^+ \rightarrow e^+ e^+ e^-$ )

# Muon to electron+gamma ( $\mu^+ \rightarrow e^+ \gamma$ )



Current State-of-the-art (@ 90% CL):

 $BF(\mu^+ \rightarrow e^+\gamma) < 4.2 \times 10^{-13}$ 

A. M. Baldini, et al. (MEG) Eur. Phys. J. C76, 8 (2016) 434.

May 2018

# Muon to electron+gamma ( $\mu^+ \rightarrow e^+ \gamma$ )



Next generation experiments:

– MEG-II (PSI)

x10 Expected improvement (relative to current state-of-the-art)

# MEG Experiment ( $\mu^+ \rightarrow e^+ \gamma$ )



#### **Background**

#### $\mu^+ \rightarrow e^+ \nu \nu \gamma$ Radiative Muon Decay (RMD)

Accidentals (ACC)

Back-to-back  $e\gamma$ E<sub>e</sub> = E<sub> $\gamma$ </sub> = m<sub> $\mu$ </sub>/2





$$\mathsf{BF}_{\mathsf{ACC}} \propto \left(\frac{R_{\mu}}{D}\right) (\Delta t_{e\gamma}) \frac{\Delta E_e}{m_{\mu}/2} \left(\frac{\Delta E_{\gamma}}{15m_{\mu}/2}\right)^2 \left(\frac{\Delta \theta_{e\gamma}}{2}\right)^2$$





Keys to success: excellent energy, timing, angular resolutions, particularly  $\Delta E_{\gamma}$  and  $\Delta \theta_{e\gamma}$ 

May 2018

#### **MEG Proton Beam**



- 1.3 MW of 0.6 GeV protons
- DC muon beam using "surface" muons,  $p_{\mu}$  ~ 28 MeV/c
- MEG uses few  $10^7 \,\mu^+/s$

### **MEG Detector**



- Liquid Xe calorimeter
  - PMT readout
  - 11% of solid angle
- Drift Chamber (DC)

- Radius : 19 - 28 cm

- Scintillator timing counters (TC)
- DC and TC inside graded solenoid field
- 205 μm polyethelene target

### **MEG Solenoid**



Sweeps e+ out of central region

Bending radius independent of pitch angle

#### **COBRA = COnstant Bending Radius**

- 1.3 T in central region
- 0.5 T in outer regions

### **MEG** Analysis



**Utilizes 5 variables** 

• E<sub>e</sub>, E<sub>γ</sub>

• 
$$t_{e\gamma} = t_e - t_{\gamma}$$

• φ<sub>eγ</sub>

Blind Analysis Full Likelihood fit to data

• Published results uses full data set (2009-2013)  $-7.5 \times 10^{14}$  stopped  $\mu^+$ 

May 2018

#### **MEG** Analysis



### **MEG** Calibrations

	Process	Energy	Main Purpose	Frequency
Cosmic rays	$\mu^{\pm}$ from atmospheric showers	Wide spectrum O(GeV)	LXe-DCH relative position	annually
			DCH alignment	
			TC energy and time offset calibration	
			LXe purity	on demand
Charge exchange	$\pi^- \mathrm{p}  o \pi^0 \mathrm{n} \ \pi^0  o \gamma \gamma$	55, 83, 129 MeV photons	LXe energy scale/resolution	annually
Radiative $\mu$ -decay	$\mu^+  ightarrow { m e}^+ \gamma  u ar{ u}$	photons > 40 MeV,	LXe-TC relative timing	continuousl
		positrons > 45 MeV	Normalisation	
Normal $\mu$ -decay	$\mu^+  ightarrow { m e}^+  u ar{ u}$	52.83 MeV end-point positrons	DCH energy scale/resolution	continuousl
			DCH and target alignment	
			Normalisation	
Mott positrons	$e^+$ target $\rightarrow e^+$ target	$\approx 50$ MeV positrons	DCH energy scale/resolution	annually
-			DCH alignment	
Proton accelerator	$^{7}\text{Li}(p,\gamma)^{8}\text{Be}$	14.8, 17.6 MeV photons	LXe uniformity/purity	weekly
	$^{11}$ B(p, $\gamma$ ) $^{12}$ C	4.4, 11.6, 16.1 MeV photons	TC interbar/ LXe-TC timing	weekly
Neutron generator	<sup>58</sup> Ni(n, <i>γ</i> ) <sup>59</sup> Ni	9 MeV photons	LXe energy scale	weekly
Radioactive source	$^{241}\mathrm{Am}(\alpha,\gamma)^{237}\mathrm{Np}$	5.5 MeV $\alpha$ 's, 56 keV photons	LXe PMT calibration/purity	weekly
Radioactive source	${}^{9}\text{Be}(\alpha_{241}\text{Am}, n){}^{12}\text{C}^{\star}$ ${}^{12}\text{C}^{\star}(\gamma){}^{12}\text{C}$	4.4 MeV photons	LXe energy scale	on demand
LED			LXe PMT calibration	continuousl

#### Table 1 The calibration tools of the MEG experiment.

#### Scale and resolutions determined with high degree of confidence

May 2018

#### **MEG Final Result**



May 2018

# MEG-II Upgrade – another x10 better



PDF paramete	ers	MEG	MEG II
$E_{e^+}$ (keV)		380	130
$\theta_{e^+}$ (mrad)		9.4	5.3
$\phi_{e^+}$ (mrad)		8.7	3.7
$z_{e^+}/y_{e^+}$ (mm) co	ore	2.4/1.2	1.6/0.7
$E_{\gamma}(\%) \ (w > 2 \ {\rm cm})$	(w < 2  cm)	2.4/1.7	1.1/1.0
$u_{\gamma}, v_{\gamma}, w_{\gamma} \text{ (mm)}$		5/5/6	2.6/2.2/5
$t_{e^+\gamma}$ (ps)		122	84
Efficiency (%)			
Trigger		≈ 99	≈ 99
Photon		63	69
$e^+$ (tracking $\times$ n	natching)	30	70

- Commissioning with beam has begun!
- Physics data taking will begin late 2018 – early 2019

### Muon to three electrons ( $\mu^+ \rightarrow e^+e^+e^-$ )



#### Current State-of-the-art (@ 90% CL):

#### $BF(\mu^+ \rightarrow e^+e^-) < 1 \times 10^{-12}$

U. Bellgardt, et al. (SINDRUM) Nucl. Phys. B299 (1988) 1.

May 2018

### Muon to three electrons ( $\mu^+ \rightarrow e^+e^+e^-$ )



#### Next generation experiments:

- Mu3e (PSI)
  - Phase la
  - Phase Ib
  - Phase II

x20 x400 (rela x10,000

Expected improvement (relative to current state-of-the-art)

May 2018

# Mu3e Experiment ( $\mu^+ \rightarrow e^+e^-$ )



**Signal** 

#### **Background**

#### $\mu^+ \rightarrow e^+ \nu \nu \gamma \rightarrow e^+ \nu \nu e^+ e^-$ Radiative Muon Decay (RMD)

Accidentals



# Mu3e Experiment ( $\mu^+ \rightarrow e^+e^+e^-$ )



**Background** 

#### $\mu^+ \rightarrow e^+ \nu \nu \gamma \rightarrow e^+ \nu \nu e^+ e^-$ Radiative Muon Decay (RMD)

Accidentals

• Keys to success: excellent momentum, timing, and vertex resolutions

# Mu3e Experiment ( $\mu^+ \rightarrow e^+e^+e^-$ )



### Mu3e (Phase-I) beam



- 1.3 MW of 0.6 GeV protons
- DC muon beam using "surface" muons,  $p_{\mu}$  ~ 28 MeV/c
- Mu3e will use  $10^7 10^8 \,\mu^+/s$
- Utilizes same beam line as MEG

## Mu3e (Phase-II) beam



To achieve Phase-II sensitivity requires an upgraded facility at PSI: High Intensity Muon Beam

- Currently under development
- Not (yet) approved



#### May 2018

# Mu3e Detector development (for Phase-Ib, II)



• Finalizing experiment design & prototyping. Aim to have Phase 1a detector ready for data taking by end of MEG-II running.

May 2018

### **Future Expectations**

### Expected data taking schedule



#### • Significant progress in all three $\mu$ channels in next 5-7 years

May 2018

### Comparing sensitivities

# Constraints from $\mu \rightarrow e$ Experiments



\* S. Nussinov, R.D. Peccei, and X.M. Zhang, Phys. Rev. D63 (2001) 016003; (arXiv:0004153 [hep-ph]).

# Constraints from $\mu \rightarrow e$ Experiments



Constraints on LFV Yukav	va couplings
$\frac{\text{From CLFV experiments}^*}{\mu \rightarrow e\gamma : B(h \rightarrow \mu e) < 10^{-8}}$ $\mu N \rightarrow eN : B(h \rightarrow \mu e) < 10^{-10}$	(today) (future)
Collider experiments LHC : $B(h \rightarrow \mu e) < 10^{-2} - 10^{-3}$	(today – future)

\* R. Harnik, J. Kopp, and J. Zupan, JHEP 03 (2013) 026; (arXiv:1209.1397 [hep-ph]).

## Examples of CLFV and LHC sensitivity



#### • CLFV experiments probe parts of NP parameter space LHC does not

May 2018

#### Constraints from $\tau \rightarrow e, \mu$ Experiments



- Results using CLFV  $\tau$  decays correspond to B(h $\rightarrow$  $\tau e$ ,  $\tau \mu$ ) ~ 10%
  - CMS and ATLAS already exploring  $B(h \rightarrow \tau \mu) \approx 1\%$

May 2018

### Summary

- CLFV experiments provide deep, broad probes of New Physics parameter space
  - Will probe  $\Lambda_{\rm NP}$  ~ O(10^3- 10^4) TeV >> LHC
- Near future experiments have compelling discovery sensitivity over a broad range of New Physics models (SUSY, GUT, ED, LHT, 2HDM,...)
- Combining information from >1 CLFV channel can allow a determination of underlying New Physics mechanism
- The next ~5 years promise to be very exciting for CLFV searches!

May 2018
# For more information

#### • Useful reviews

- L. Calibbi and G. Signorelli, Riv. Nuovo Cimento, 41 (2018) 71
- T. Gorringe & D. Hertzog, Prog.Part.Nucl. Phys. 84 (2015) 73.
- S. Mihara, J.P. Miller, P. Paradisi, G. Piredda, Annu.Rev.Nucl.Part.Sci. 63 (2013) 552.
- R.H. Bernstein & P.S. Cooper, Phys. Rept. 532 (2013) 27.
- Y. Kuno & Y. Okada, Rev.Mod.Phys. 73 (2001) 151.

#### • About the experiments

- MEG: <u>http://meg.icepp.s.u-tokyo.ac.jp</u> (MEG-II TDR: arXiv:1801.04688)
- Mu2e: <u>http://mu2e.fnal.gov</u> (TDR: arXiv:1501.05241)
- COMET: <u>http://comet.kek.jp/Introduction.html</u> (Proposal: http://comet.kek.jp/Documents\_files/Phase-I-Proposal-v1.2.pdf)
- DeeMee: <u>http://deeme.hep.sci.osaka-u.ac.jp</u> (Proposal: http://deeme.hep.sci.osaka-u.ac.jp/documents/deeme-proposal-r28.pdf/view)
- Mu3e: <u>https://www.psi.ch/mu3e/mu3e</u> (Proposal: https://www.psi.ch/mu3e/documents)

# **Backup Slides**

# Using CLFV to Determine New Physics

# Can use ratio of rates to determine dominant operator contribution

- multiple ratios can determine multiple operators and the ratio of their couplings
- e.g.  $(\mu \rightarrow e\gamma) / (\mu N \rightarrow eN)$
- e.g.  $\mu N \rightarrow eN$  with different nuclei
- Also information in angular distributions

 $-\mu \rightarrow eee, \tau chnls$ 



D.Glenzinski | Fermilab





### **CLFV** Sensitivity

W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

~
<u> </u>
>
- <del>1</del>
<u>.</u>
.0
~
<u> </u>
Ð
S
1.1
~
<u> </u>
(1)
~
~
0
U U
Ś
·=
_
- 11
11
1
-

	AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?
$\epsilon_K$	*	***	***	*	*	**	***
$S_{\psi\phi}$	***	***	***	*	*	***	***
$S_{\phi K_S}$	***	**	*	***	***	*	?
$A_{\rm CP}\left(B \to X_s \gamma\right)$	*	*	*	***	***	*	?
$A_{7,8}(B ightarrow K^*\mu^+\mu^-)$	*	*	*	***	***	**	?
$A_9(B  ightarrow K^* \mu^+ \mu^-)$	*	*	*	*	*	*	?
$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$B_s  ightarrow \mu^+ \mu^-$	***	***	***	***	***	*	*
$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***
$K_L  ightarrow \pi^0  u ar u$	*	*	*	*	*	***	***
$\mu \rightarrow e \gamma$	***	***	***	***	***	***	***
$\tau \rightarrow \mu \gamma$	***	***	*	***	***	***	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
$d_n$	***	***	***	**	***	*	***
$d_e$	***	***	**	*	***	*	***
$(g-2)_{\mu}$	***	***	**	***	***	*	?

AC	U(1) flavor symmetry
RVV2	Non-abelian SU(3)- flavored MSSM
AKM	SU(3)-flavored SUSY
δμ	LH CKM-like currents
FBMSSM	Flavor-blind MSSM
LHT	Little Higgs w/T parity
RS	Randall-Sundrum

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models  $\bigstar \bigstar \bigstar$  signals large effects,  $\bigstar \bigstar$  visible but small effects and  $\bigstar$  implies that the given model does not predict sizable effects in that observable.

# **MEG** Calibration

- E<sub>e</sub> calibrated using Michel edge
- $E_{\gamma}$  calibrated during special runs using a liquid hydrogen target
  - $-\pi^{-}p \rightarrow \pi^{0} n \rightarrow \gamma\gamma n$  (pion charge exchange)
  - Tag opposite side  $\gamma$  using a small movable BGO array dedicated to this calibration
  - The angle and energy of the opposite side tag can be used to define a monoenergetic source of  $\gamma$  in the LXe (~55 MeV)

### **MEG** Calibration



- Photon energy scale determined to <1% using special runs with a pion beam, a custom target, and exploiting charge exchange processes to identify monoenergetic photons with energies ~55 MeV
- Monitor stability of Eγ with short special runs that use a <1 MeV proton beam and a Li<sub>2</sub>B<sub>4</sub>O<sub>7</sub> target to produce photons at 18, 12, & 4 MeV

### **MEG** Calibration



• Relative  $e-\gamma$  timing calibrated using radiative muon decays

May 2018

### MEG Backgrounds



#### • Accidental backgrounds dominate.

# **MEG Accidental Background**



- Both curves are steeply falling in region of interest... sensitivity strong function of  $\sigma_{\text{E}}$ 

### MEG-II Upgrade : aim for x10 improvement



### • Expect to begin 3y data taking in 2019 —Aim to improve sensitivity by x10

# Mu3e acceptance

- Inner bore of detector solenoid large enough to fully contain tracks with momenta up to ~50 MeV/c
- Central pixels fully fiducial for track momenta down to ~12 MeV/c
- Forward pixels provide additional hits from "recurl"



May 2018

# DeeMee ( $\mu$ -N $\rightarrow$ e-N)







Signal Region: 102.0 -- 105.6 MeV/c

- New concept at JPARC – 3 GeV from RCS H-Line
- Use thick target as production, decay, and stopping volumes (graphite, SiC)
- Customize beam line to select momentum bite near  $E_{\mu e} \sim m_{\mu}$ so that you're sensitive to  $\mu N \rightarrow eN$  that occurs near the target surface
- Goal: R<sub>μe</sub> < 2 x 10<sup>-14</sup> @ 90%CL
  - 2-3y of running at 1 MW
  - Currently operating at ~400 kW

### COMET Phase I & II

28

#### Summary of COMET phase-I/II

	COMET-Phase-I	COMET-Phase-II
experiment starts (*)	2020	~2022
beam intensity	3.2kW (8GeV)	56kW (8GeV)
running time	1.5 x 10 <sup>6</sup> (sec)	2.0 x 10 <sup>7</sup> (sec)
# of protons	$3.8 \times 10^{18}$	$8.5 \times 10^{20}$
# of muon stops	8.7 × 10 <sup>15</sup>	$2.0 \times 10^{18}$
muon rate	$5.8 \times 10^{9}$	$1.0 \times 10^{11}$
# of muon stops / proton	0.0023	0.0023
# of BG	0.03	0.3
S.E.S.	3.1 x 10 <sup>-15</sup>	2.6 x 10 <sup>-17</sup>
U.L. (90%CL.)	7.0 × 10 <sup>-15</sup>	6.0 × 10 <sup>-17</sup>

(\*) Engineering runs and Physics runs

H.Nishiguchi(KEK)

Project of Muon LFV at J-PARC

Tau2012, Nagoya

#### • Will employ straw technology

- Low mass
- Can reliably operate in vacuum
- Robust against single-wire failures



- 5 mm diameter straw
- Spiral wound
- Walls: 12 μm Mylar + 3 μm epoxy + 200 Å Au + 500 Å Al
- $\bullet$  25  $\mu m$  Au-plated W sense wire
- 33 117 cm in length
- 80/20 Ar/CO2 with HV < 1500 V

May 2018



- Self-supporting "panel" consists of 100 straws
- 6 panels assembled to make a "plane"
- 2 planes assembled to make a "station"
- Rotation of panels and planes improves stereo information
- >20k straws total

- 18 "stations" with straws transverse to the beam
- Naturally moves readout and support to large radii, out of the active volume





- Inner 38 cm is purposefully un-instrumented
  - Blind to beam flash
  - Blind to >99% of DIO spectrum

# Mu2e Spectrometer Performance



#### • Performance well within physics requirements

May 2018

### Mu2e Track Reconstruction and Selection





### Mu2e Performance



#### • Robust against increases in rate

# Mu2e Calorimeter

- Baseline design : Cesium Iodide (CsI)
  - Radiation hard, fast, compact

	Csl
Density (g/cm3)	4.51
Radiation length (cm)	1.86
Moliere Radius (cm)	3.57
Interaction length (cm)	39.3
dE/dX (MeV/cm)	5.56
Refractive index	1.95
Peak luminescence (nm)	310
Decay time (ns)	26
Light yield (rel. to Nal)	3.6%
Variation with temperature	-1.4% / deg-C

# Mu2e Calorimeter



May 2018

### Mu2e Calorimeter



- With 60 ns integration, expect to achieve an energy resolution ~5% for 105 MeV electrons
  - Performance a weak function of rate in relevant range

May 2018

### Mu2e Cosmic-Ray Veto



- Cosmic  $\mu$  can generate background events via decay, scattering, or material interactions

### Mu2e Cosmic-Ray Veto



#### • Veto system covers entire DS and half TS

# Mu2e Cosmic-Ray Veto



- Will use 4 overlapping layers of scintillator
  - Each bar is  $5 \times 2 \times 450 \text{ cm}^3$
  - 2 WLS fibers / bar
  - Read-out both ends of each fiber with SiPM
  - Have achieved  $\varepsilon$  > 99.4% (per layer) in test beam

# Cosmic Ray Veto



• Test beam data to vet design/performance

# Mu2e Selection Requirements

Parameter	Requirement		
Track quality and backgrou	and rejection criteria		
alman Fit Status Successful Fit			
Number of active hits	$N_{active} \ge 25$		
Fit consistency	$\chi^2$ consistency > 2x10 <sup>-3</sup>		
Estimated reconstructed momentum uncertainty	$\sigma_p < 250 \text{ keV/c}$		
Estimated track t <sub>0</sub> uncertainty	$\sigma_t < 0.9$ nsec		
Γrack t <sub>0</sub> (livegate)	700 ns $< t_0 < 1695$ ns		
Polar angle range (pitch)	$45^{\circ} < \theta < 60^{\circ}$		
Ainimum track transverse radius	$-80 \text{ mm} < d_0 < 105 \text{ mm}$		
Aaximum track transverse radius	$450 \text{ mm} < d_0 + 2/\omega < 680 \text{ mm}$		
rack momentum	103.75 < p < 105.0 MeV/c		
Calorimeter matching and partic	cle identification criteria		
rack match to a calorimeter cluster	$E_{cluster} > 10 \text{ MeV}$		
	$\chi^2$ (track-calo match) < 100		
latio of cluster energy to track momentum	E/P < 1.15		
Difference in track $t_0$ to calorimeter $t_0$	$\Delta t =  t_{track} - t_{calo}  < 3$ ns from peak		
Particle identification	$\log(L(e)/L(\mu)) < 1.5$		

Full set of selection criteria employed to estimate backgrounds and sensitivity reported in TDR (Summer 2014)

# Mu2e Systematic Uncertainties

Effect	Uncertainty in DIO background yield	Uncertainty in CE single- event-sensitivity (×10 <sup>-17</sup> )
MC Statistics	±0.02	±0.07
Theoretical Uncertainty	±0.04	-
Tracker Acceptance	±0.002	±0.03
Reconstruction Efficiency	±0.01	±0.15
Momentum Scale	+0.09, -0.06	±0.07
µ-bunch Intensity Variation	±0.007	±0.1
Beam Flash Uncertainty	±0.011	±0.17
µ-capture Proton Uncertainty	±0.01	±0.016
µ-capture Neutron Uncertainty	±0.006	±0.093
µ-capture Photon Uncertainty	±0.002	±0.028
Out-Of-Target µ Stops	±0.004	±0.055
Degraded Tracker	-0.013	+0.191
Total (in quadrature)	+0.10, -0.08	+0.35, -0.29

#### • Evaluated for all background sources

# Mu2e Tracker Occupancy



Reco Hit Time by Generator Particle

- Accidental occupancy from beam flash,  $\mu$  capture products, out-of-target  $\mu$  stops, etc.

May 2018

# Mu2e Signal Momentum Spectrum



 Smearing dominated by interactions in stopping target and in (neutron/proton) absorbers upstream of tracker

May 2018



- We need to understand contributions from accidentals and correlated-accidentals
  - For neutrons and photons as a function of time, energy, timing resolution, and read-out threshold

### Muon momentum distribution



#### • The muons that stop are low momentum

# Improving the Previous Experiment

- Current world's best limit on  $\mu N \rightarrow eN$  is from SINDRUM-II:
  - W. Bertl, et al. (SINDRUM II Collaboration), Euro. Phys. C47 (2006) 337.
  - $-R_{\mu e}(\mu N_{Au} \rightarrow eN_{Au}) < 7 \times 10^{-13} @ 90\% CL (2006)$
  - Limited by
    - Backgrounds from prompt pions
    - Stopped- $\mu$  rate (~10<sup>7</sup>  $\mu$ /s using ~1 MW beam)
- Any improvement to SINDRUM-II needs to address these limitations

# Improving the Previous Experiment

- In 1989 Lobashev and Djilkabaev published a paper proposing an experiment that solved these two problems by
  - 1. Utilizing a pulsed proton beam
  - 2. Employing solenoids to collect muons
- Mu2e is the realization of their proposed technique
  - Pulsed beam from the Fermilab accelerator complex
  - Solenoid system capable of delivering high intensity stopped-muon beam
## Using Solenoids to Collect Muons



#### Fermilab's Muon Campus



- New facilities to host muon experiments
  - Two new experimental halls and the associated beam lines
  - Will produce the world's highest intensity muon beams
  - Physics data taking has begun for Muon (g-2) experiment

May 2018

## Mu2e Solenoid Summary

	PS	TS	DS
Length (m)	4	13	11
Diameter (m)	1.7	0.4	1.9
Field @ start (T)	4.6	2.5	2.0
Field @ end (T)	2.5	2.0	1.0
Number of coils	3	52	11
Conductor (km)	14	44	17
Operating current (kA)	10	3	6
Stored energy (MJ)	80	20	30
Cold mass (tons)	11	26	8

- PS, DS is being built by General Atomics
  - TS is being built by ASG + Fermilab

May 2018

## Mu2e Schedule

- Full scale solenoid construction has started
- Full scale detector construction ramping-up in 2018
- Solenoid and detector installation in 2019-2020
- Initial commissioning in 2021
- First physics running in 2022
- At full intensity
  - Reach Sindrum-II sensitivity in 100 min
  - x10 in 17 hours running
  - x100 in 7 days running
  - x10000 in 700 days running

## Mu2e Collaboration

Spokespersons: Jim Miller (Boston), DG (Fermilab)

#### **Over 200 Scientists from 37 Institutions**

Argonne National Laboratory, Boston University, University of California Berkeley, University of California Davis, University of California Irvine, California Institute of Technology, City University of New York, Joint Institute of Nuclear Research Dubna, Duke University, Fermi National Accelerator Laboratory, Laboratori Nazionale di Frascati, University of Houston, Helmholtz-Zentrum Dresden-Rossendorf. INFN Genova, Institute for High Energy Physics, Protvino, Kansas State University, Lawrence Berkeley National Laboratory, INFN Lecce, University Marconi Rome, Lewis University, University of Liverpool, University College London, University of Louisville, University of Manchester, University of Minnesota, Muon Inc., Northwestern University, Institute for Nuclear Research Moscow, INFN Pisa, Northern Illinois University, Purdue University, Rice University, Sun Yat-Sen University, University of South Alabama, Novosibirsk State University/Budker Institute of Nuclear Physics, University of Virginia, University of Washington, Yale University













May 2018

# $h \rightarrow \tau$ constraints from $\mu \rightarrow e$ CLFV

 $\tau$ μ-τe couplings can contribute to μ $\rightarrow$ e transitions. As an example:



- $\mu \rightarrow e\gamma$  constrains dipole contributions
- $\mu$ -N  $\rightarrow$  e-N constrains vector contributions
- Future improvements in  $\mu$ -N  $\rightarrow$  e-N will probe B(h $\rightarrow \tau\mu$ )B(h $\rightarrow \tau e$ ) < 10<sup>-7</sup>

cf. I.Dorsner, S. Fajfer, A. Greljo, J. Kamenik, N, Kosnik, I. Nisandzic, JHEP 1506 (2015) 108; (1502.07784). R. Harnik, J. Kopp, J. Supan (1209.1397).

May 2018

#### **Direct Searches for CLFV h decays**



<u>CMS</u> B(h $\rightarrow \tau\mu$ ) < 1.51 x 10<sup>-2</sup> Best fit : (0.84 +/- 0.40)%

<u>ATLAS</u> B(h $\rightarrow \tau\mu$ ) < 1.43 x 10<sup>-2</sup> Best fit : (0.53 +/- 0.51)%

#### • Looking forward to the updated analysis...

#### NP Contributions to CLFV



• As has been discussed, LFV higgs couplings will contribute

### CLFV using taus

#### cLFV Experiments using taus



#### Sensitivity dominated by Belle & BaBar

#### cLFV Experiments using taus

E>	Physics process	Cross section (nb)
	$\Upsilon(4S) \to B\bar{B}$	1.2
	$e^+e^- \rightarrow {\rm continuum}$	2.8
	$\mu^+\mu^-$	0.8
	$\tau^+ \tau^-$	0.8
	Bhabha ( $\theta_{\rm lab} \ge 17^{\circ}$ )	44
	$\gamma\gamma~(\theta_{\rm lab} \ge 17^\circ)$	2.4
	$2\gamma$ processes $^{b}$	$\sim 80$
	Total	$\sim 130$

#### $\tau^+\tau^-$ production cross section ~ 1 nb Latest results use 500-800 fb<sup>-1</sup> ~ 500-800M $\tau^+\tau^-$ pairs

#### cLFV Experiments using taus



- Exploit (E,p) constraints available at e<sup>+</sup>e<sup>-</sup> collider
- On Search side:  $m_{reco} = m_{\tau}$ ,  $E_{reco} = E_{beam}$  (in CM)
- Employ excellent particle identification algorithms
- Additional requirements to suppress qq,  $\mu\mu$ , ee,  $\gamma\gamma$

May 2018

#### Some cLFV Results



•  $\tau \rightarrow e\gamma$ ,  $\mu\gamma$  already observe background

May 2018

#### Some cLFV Results







Ellipse = Signal Region Solid lines : m<sub>reco</sub> sidebands

- $\tau \rightarrow 3$  leptons very clean
- No events in signal region



#### cLFV tau decays at LHCb



- Lots of  $\tau$  from b $\rightarrow \tau v X$  and c $\rightarrow \tau v X$
- Effective cross section of 85  $\mu b$
- Competitive  $\tau \rightarrow \mu \mu \mu$  sensitivity

#### cLFV Result



• Normalized to  $D_s \rightarrow \phi \pi \rightarrow (\mu \mu) \pi$ 



#### Most Recent HFAG Results



#### Future Sensitivity



## LHCb $\tau \rightarrow \mu \mu \mu$ Results

	7 TeV	8 TeV	
$\mathcal{B}\left(D^s\to\phi\left(\mu^+\mu^-\right)\pi^-\right)$	$(1.32 \pm 0.10) \times 10^{-5}$		
$\mathcal{B}\left(D_s^- \to \tau^- \bar{\nu}_\tau\right)$	$(5.61 \pm 0.24) \times 10^{-2}$		
$f_{ au}^{D_s}$	$0.78\pm0.04$	$0.80\pm0.03$	
$\epsilon_{ m cal}{}^{ m R}/\epsilon_{ m sig}{}^{ m R}$	$0.898 \pm 0.060$	$0.912\pm0.054$	
${\epsilon_{\mathrm{cal}}}^{\mathrm{T}}/{\epsilon_{\mathrm{sig}}}^{\mathrm{T}}$	$0.659 \pm 0.006$	$0.525\pm0.040$	
$N_{ m cal}$	$28200\pm440$	$52130\pm700$	
lpha	$(7.20 \pm 0.98) \times 10^{-5}$	$^{-9}$ (3.37 ± 0.50) × 10 <sup>-9</sup>	

Table 1: Terms entering into the normalisation factors,  $\alpha$ , and their combined statistical and systematic uncertainties.

$$\mathcal{B}\left(\tau^{-} \to \mu^{-} \mu^{+} \mu^{-}\right) = \frac{\mathcal{B}\left(D_{s}^{-} \to \phi\left(\mu^{+} \mu^{-}\right) \pi^{-}\right)}{\mathcal{B}\left(D_{s}^{-} \to \tau^{-} \bar{\nu}_{\tau}\right)} \times f_{\tau}^{D_{s}} \times \frac{\epsilon_{\mathrm{cal}}^{\mathrm{R}}}{\epsilon_{\mathrm{sig}}^{\mathrm{R}}} \times \frac{\epsilon_{\mathrm{cal}}^{\mathrm{T}}}{\epsilon_{\mathrm{sig}}^{\mathrm{T}}} \times \frac{N_{\mathrm{sig}}}{N_{\mathrm{cal}}} \equiv \alpha N_{\mathrm{sig}},$$

May 2018

	$\mathcal{M}_{\mathrm{PID}}$ response	$\mathcal{M}_{3body}$ response	Expected	Observed
		0.26 - 0.34	$39.6\pm2.3$	39
		0.34 - 0.45	$32.2\pm2.1$	34
		0.45 - 0.61	$28.7\pm2.0$	28
	0.40 - 0.54	0.61 - 0.70	$9.7 \pm 1.2$	5
		0.70 - 0.83	$11.4 \pm 1.3$	7
		0.83 - 0.94	$7.3 \pm 1.1$	6
		0.94 - 1.00	$6.0\pm1.0$	0
		0.26 - 0.34	$13.6\pm1.4$	8
		0.34 - 0.45	$12.1 \pm 1.3$	12
		0.45 - 0.61	$8.3 \pm 1.0$	13
	0.54 - 0.61	0.61 - 0.70	$2.60\pm0.62$	1
		0.70 - 0.83	$1.83\pm0.60$	5
		0.83 - 0.94	$2.93\pm0.72$	6
		0.94 - 1.00	$2.69\pm0.63$	3
		0.26 - 0.34	$13.5\pm1.4$	7
		0.34 - 0.45	$10.9 \pm 1.2$	11
		0.45 - 0.61	$9.7 \pm 1.2$	12
	0.61 - 0.71	0.61 - 0.70	$3.35\pm0.69$	2
		0.70 - 0.83	$4.60\pm0.89$	5
		0.83 - 0.94	$4.09\pm0.81$	4
		0.94 - 1.00	$2.78\pm0.68$	1
		0.26 - 0.34	$7.8 \pm 1.1$	6
		0.34 - 0.45	$7.00\pm0.99$	8
		0.45 - 0.61	$6.17\pm0.95$	6
	0.71 - 0.80	0.61 - 0.70	$1.57\pm0.56$	2
		0.70 - 0.83	$2.99\pm0.72$	0
		0.83 - 0.94	$3.93\pm0.81$	0
		0.94 - 1.00	$3.22\pm0.68$	1
		0.26 - 0.34	$5.12\pm0.86$	3
		0.34 - 0.45	$4.44\pm0.79$	6
		0.45 - 0.61	$3.80\pm0.78$	5
	0.80 - 1.00	0.61 - 0.70	$2.65\pm0.68$	2
		0.70 - 0.83	$3.05\pm0.67$	2
	0.83 - 0.94	$1.74\pm0.54$	2	
		0.94 - 1.00	$3.36\pm0.70$	3

Table 3: Expected background candidate yields in the 8 TeV data set, with their uncertainties, and observed candidate yields within the  $\tau^-$  signal window in the different bins of classifier response. The classifier responses range from 0 (most background-like) to +1 (most signal-like). The first bin in each classifier response is excluded from the analysis.

## LHCb $\tau \rightarrow \mu \mu \mu$ Results (2fb<sup>-1</sup> @ 8TeV)

#### Data are consistent with Background expectations

Table 2: Expected background candidate yields in the 7 TeV data set, with their uncertainties, and observed candidate yields within the  $\tau^-$  signal window in the different bins of classifier response. The classifier responses range from 0 (most background-like) to +1 (most signal-like). HCb  $\tau \rightarrow \mu \mu \mu$ The first bin in each classifier response is excluded from the analysis.

$\mathcal{M}_{\mathrm{PID}}$ response	$\mathcal{M}_{3\mathrm{body}}$ response	Expected	Observed
	0.28 - 0.32	$3.17 \pm 0.66$	4
	0.32 - 0.46	$9.2 \pm 1.1$	6
0.40 - 0.45	0.46 - 0.54	$2.89 \pm 0.63$	6
	0.54 - 0.65	$3.17\pm0.66$	4
	0.65 - 0.80	$3.64\pm0.72$	2
	0.80 - 1.00	$3.79\pm0.80$	3
	0.28 - 0.32	$4.22\pm0.78$	6
	0.32 - 0.46	$8.3 \pm 1.1$	10
0.45 - 0.54	0.46 - 0.54	$2.3\pm0.57$	4
	0.54 - 0.65	$2.83\pm0.63$	8
	0.65 - 0.80	$2.72\pm0.69$	5
	0.80 - 1.00	$4.83\pm0.90$	7
	0.28 - 0.32	$2.33\pm0.58$	6
	0.32 - 0.46	$8.3 \pm 1.1$	8
0.54 - 0.63	0.46 - 0.54	$2.07\pm0.53$	1
	0.54 - 0.65	$3.29\pm0.68$	1
	0.65 - 0.80	$2.96\pm0.65$	4
	0.80 - 1.00	$3.11\pm0.69$	3
	0.28 - 0.32	$2.69\pm0.62$	1
	0.32 - 0.46	$7.5\pm1.0$	5
0.63-0.75	0.46 - 0.54	$2.06\pm0.53$	3
	0.54 - 0.65	$2.00\pm0.55$	5
	0.65 - 0.80	$3.16\pm0.66$	2
	0.80 - 1.00	$4.67\pm0.84$	2
	0.28 - 0.32	$2.19\pm0.55$	2
	0.32 - 0.46	$3.38\pm0.76$	5
0.75 - 1.00	0.46 - 0.54	$1.52\pm0.46$	3
	0.54 - 0.65	$1.28\pm0.47$	1
	0.65 - 0.80	$2.78\pm0.65$	1
	0.80 - 1.00	$4.42\pm0.83$	7

#### **Results** (1fb<sup>-1</sup> @ 7TeV)

#### Data are consistent with Background expectations