



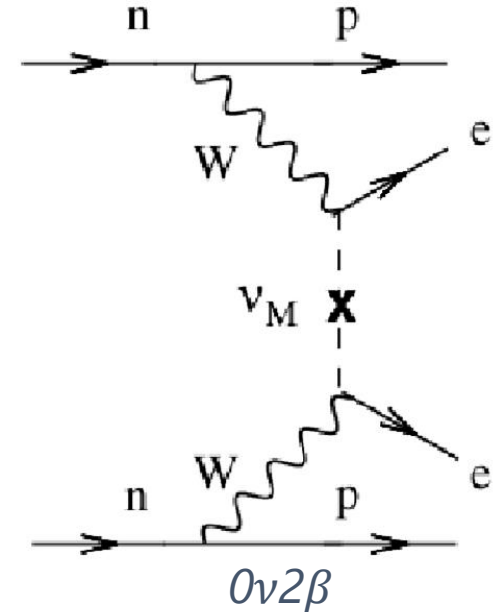
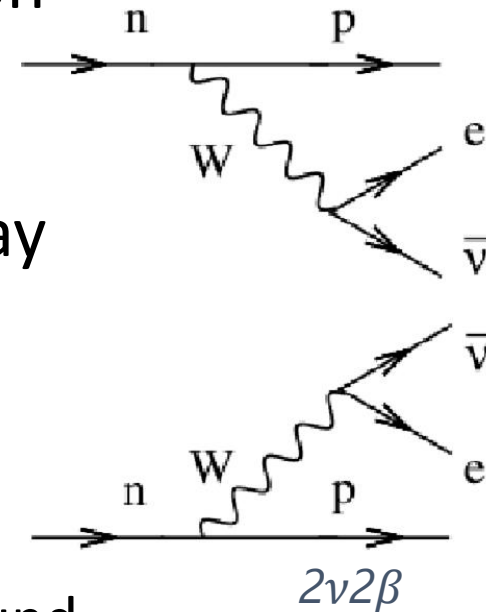
SNO+ from water to scintillator

J. P. Yáñez for the SNO+ Collaboration
Weak Interactions and Neutrinos 2019
j.p.yanez@ualberta.ca



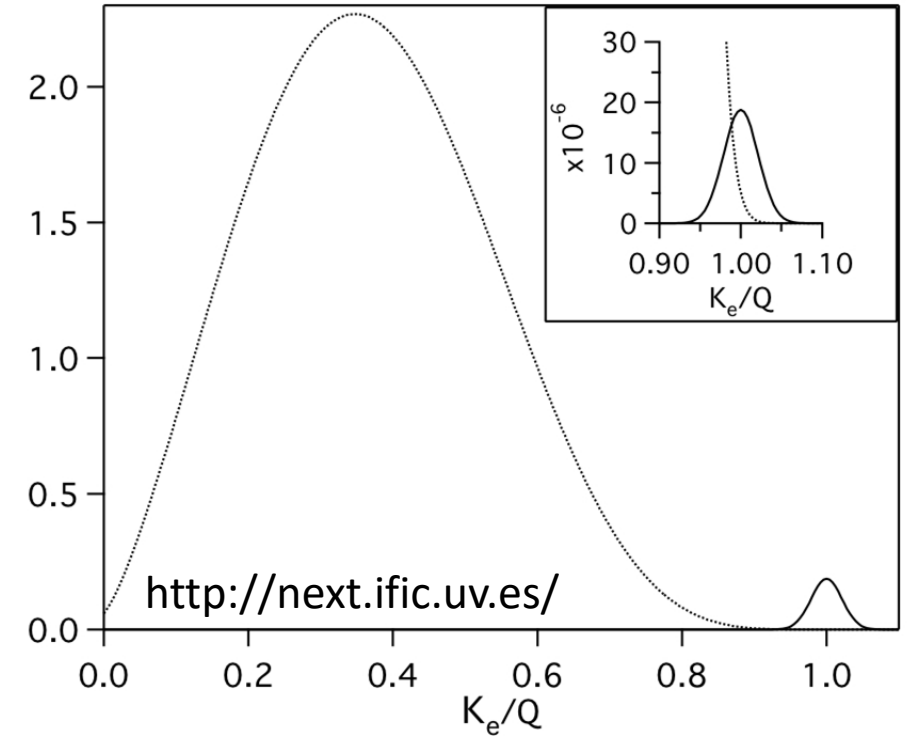
A matter/antimatter asymmetry search

- The neutrino could be a Majorana fermion
 - Only possible for neutral particles
 - Would be its own antiparticle
- Some isotopes undergo double beta decay
 - If the neutrino is Majorana, the decay can produce zero neutrinos
- $2\nu 2\beta$ is rare; $0\nu 2\beta$ would be even rarer.
Detection requires to:
 - Achieve (and understand) very low background
 - Accurately determine detector response
 - Consider scalability of the technique

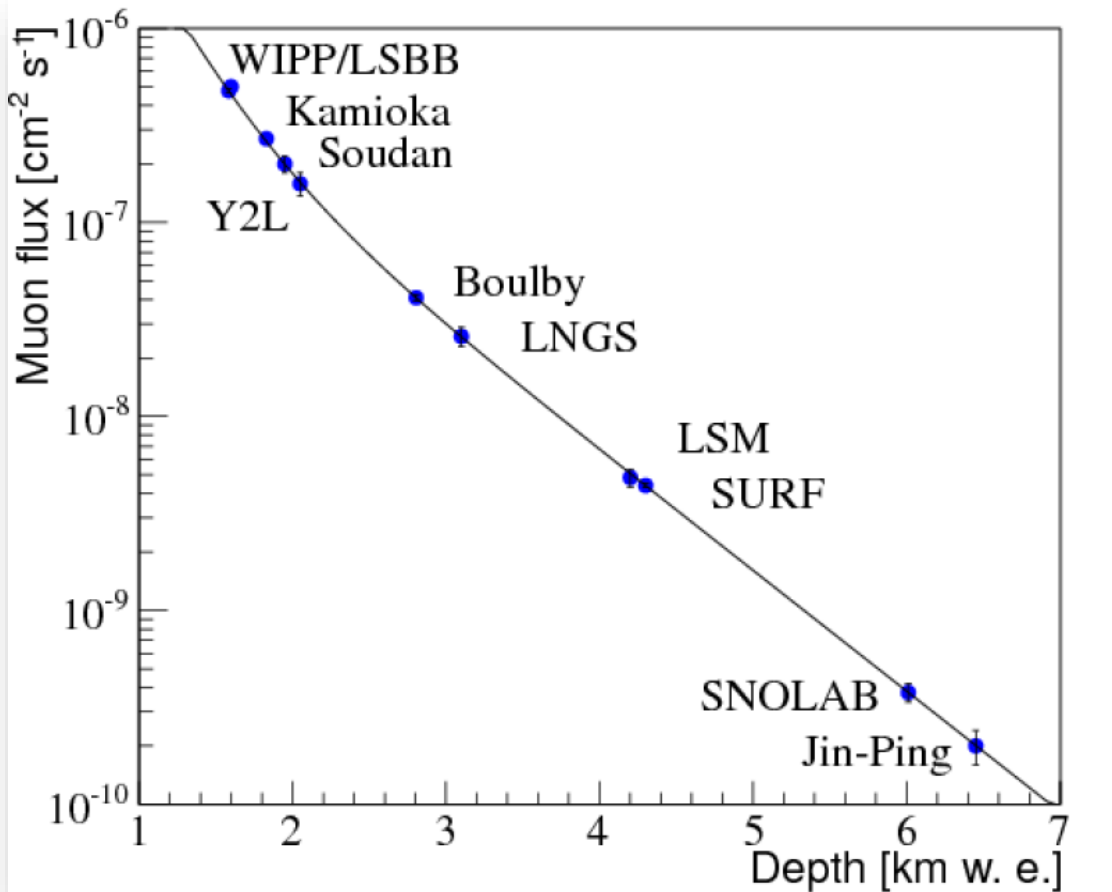


A matter/antimatter asymmetry search

- The neutrino could be a Majorana fermion
 - Only possible for neutral particles
 - Would be its own antiparticle
- Some isotopes undergo double beta decay
 - If the neutrino is Majorana, the decay can produce zero neutrinos
- $2\nu 2\beta$ is rare; $0\nu 2\beta$ would be even rarer.
Detection requires to:
 - Achieve (and understand) very low background
 - Accurately determine detector response
 - Consider scalability of the technique



The SNO+ detector



- Very low background neutrino detector
- Located in SNOLAB
 - Sudbury, ON, Canada
- At a depth of 2km (rock, 5900 mwe)
 - About 63 cosmic muons/day

This Is What Life In The Happiest City In Canada Is Like

Welcome to Sudbury, Ontario.

Posted on April 28, 2015, at 10:10 a.m.



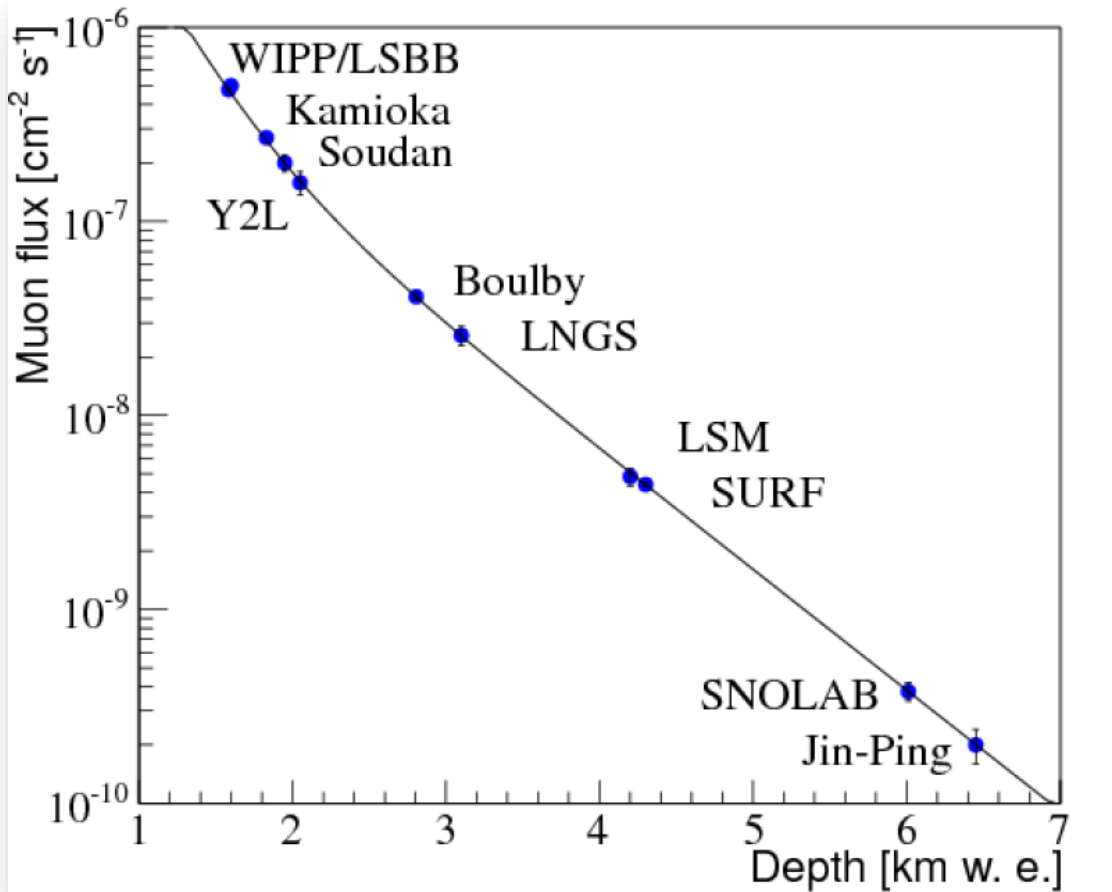
Amber Dowling
BuzzFeed Contributor

[View 34 comments ↓](#)

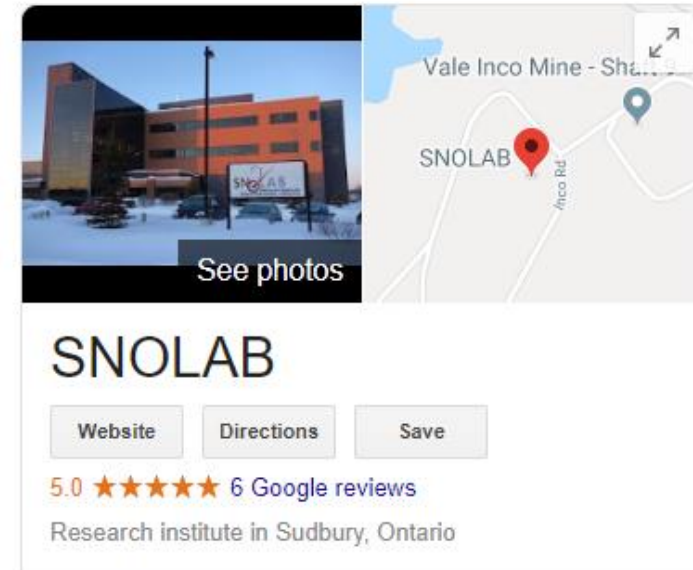


Welcome to the happiest city in Canada: Sudbury, Ontario.

The SNO+ detector

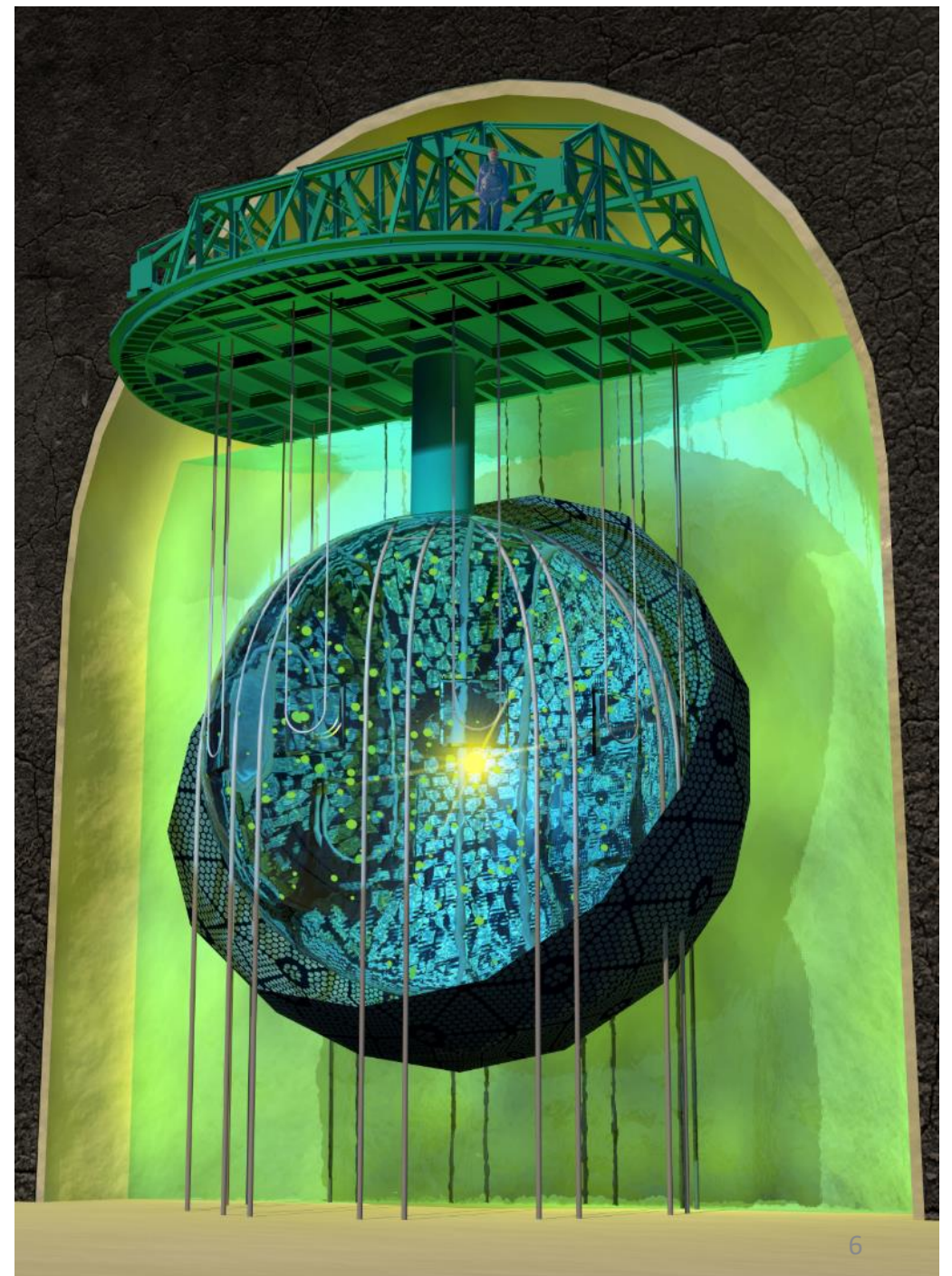


- Very low background neutrino detector
- Located in SNOLAB
 - Sudbury, ON, Canada
- At a depth of 2km (rock, 5900 mwe)
 - About 63 cosmic muons/day



The SNO+ detector

- Detector itself is over 9000 PMTs monitoring an acrylic vessel
 - Mounted at 8.5m radius
 - 54% of photocoverage
 - Cavity is flooded with ultra-pure water
- Fiducial volume: spherical acrylic vessel
 - Radius of 6m
 - Held in place by tensylon rope systems
 - Access via neck at the top



Fiducial volume material

Material in vessel sets the physics goal

1. Ultra-pure water phase

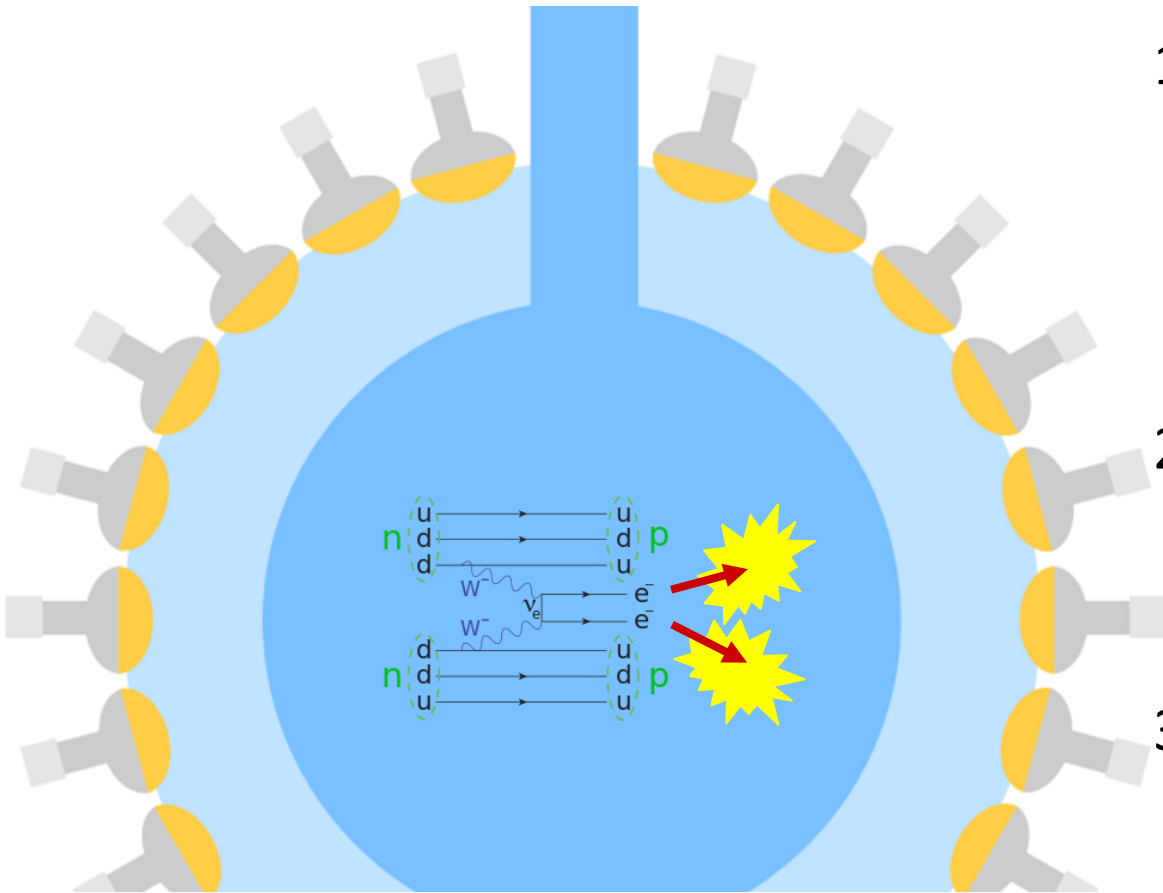
- Recording 7 hits/MeV deposited energy (Cherenkov)
- Physics goals are nucleon decay, solar and reactor neutrinos
- Understanding of detector optics and external backgrounds

2. Intermediate stage: scintillator

- Estimated 500 hits/MeV deposited energy
- Studies of solar, geo and reactor neutrinos
- Understanding of scintillator backgrounds

3. Ultimate goal: Tellurium-loaded scintillator

- Search for $0\nu 2\beta$ in ^{130}Te (Q value 2.5 MeV)
- Expect 400 hits/MeV deposited energy



Fiducial volume material

Material in vessel sets the physics goal

1. Ultra-pure water phase

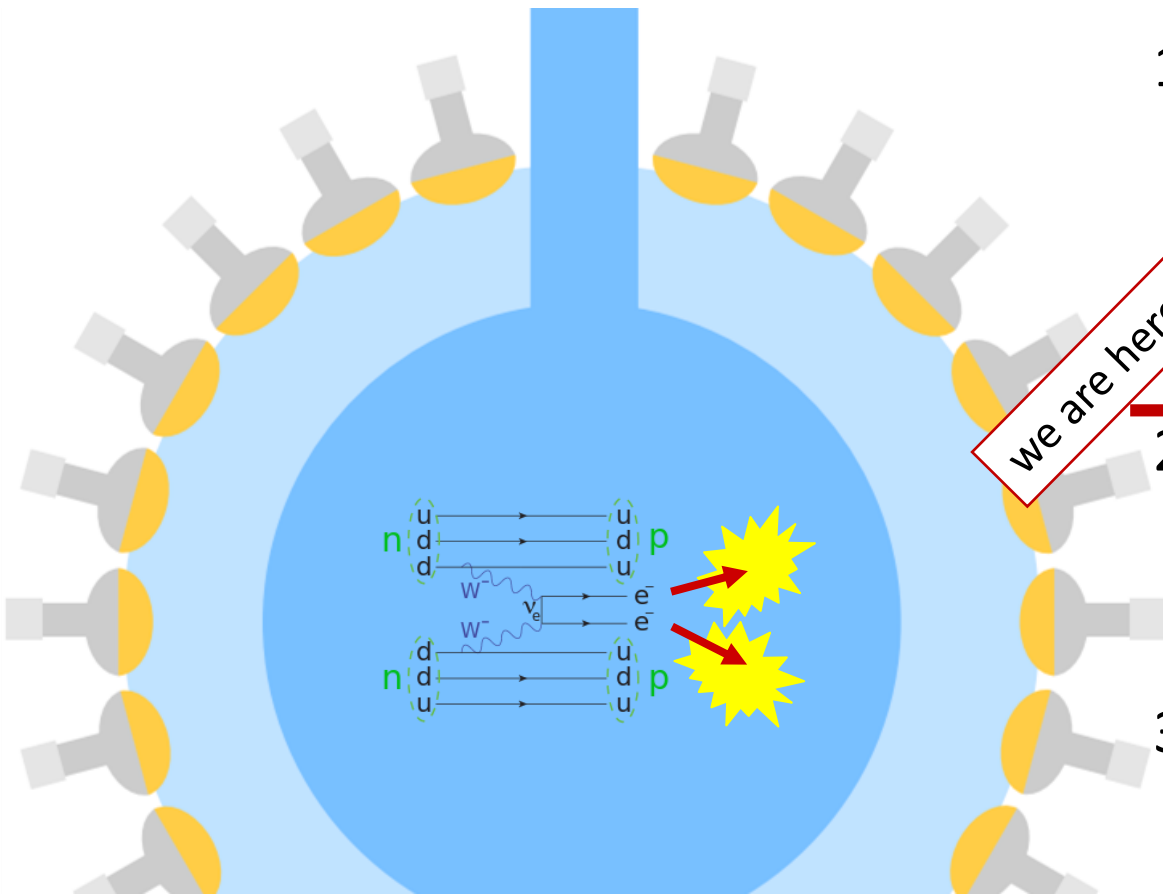
- Recording 7 hits/MeV deposited energy (Cherenkov)
- Physics goals are **nucleon decay, solar** and reactor neutrinos
- **Understanding of detector optics and external backgrounds**

2. Intermediate stage: scintillator

- Estimated 500 hits/MeV deposited energy
- Studies of solar, geo and reactor neutrinos
- Understanding of scintillator backgrounds

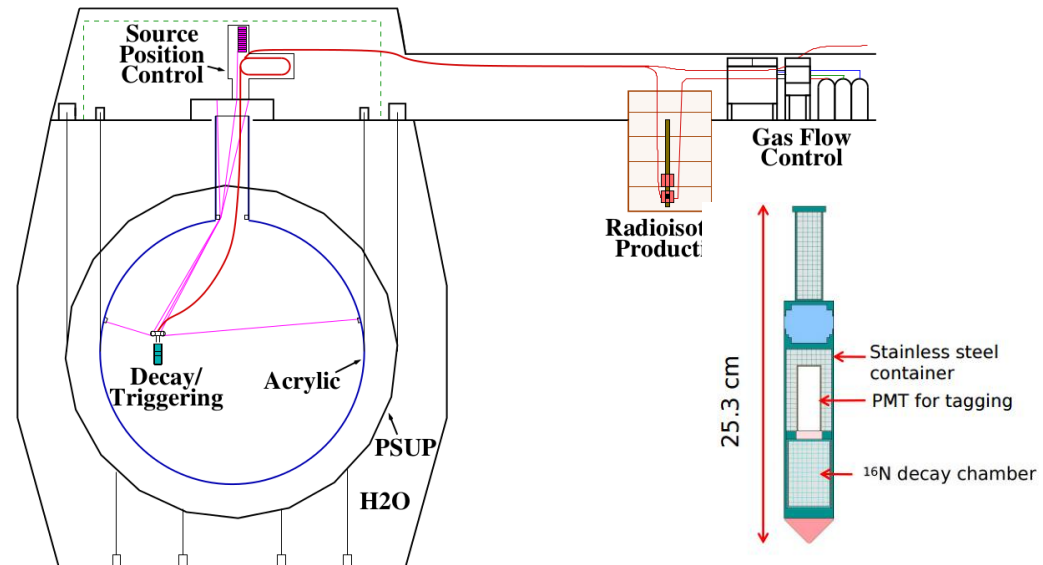
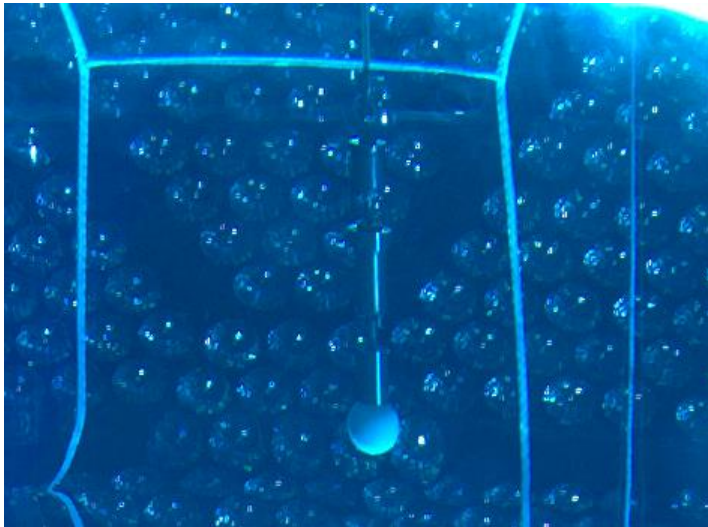
3. Ultimate goal: Tellurium-loaded scintillator

- Search for $0\nu 2\beta$ in ^{130}Te (Q value 2.5 MeV)
- Expect 400 hits/MeV deposited energy



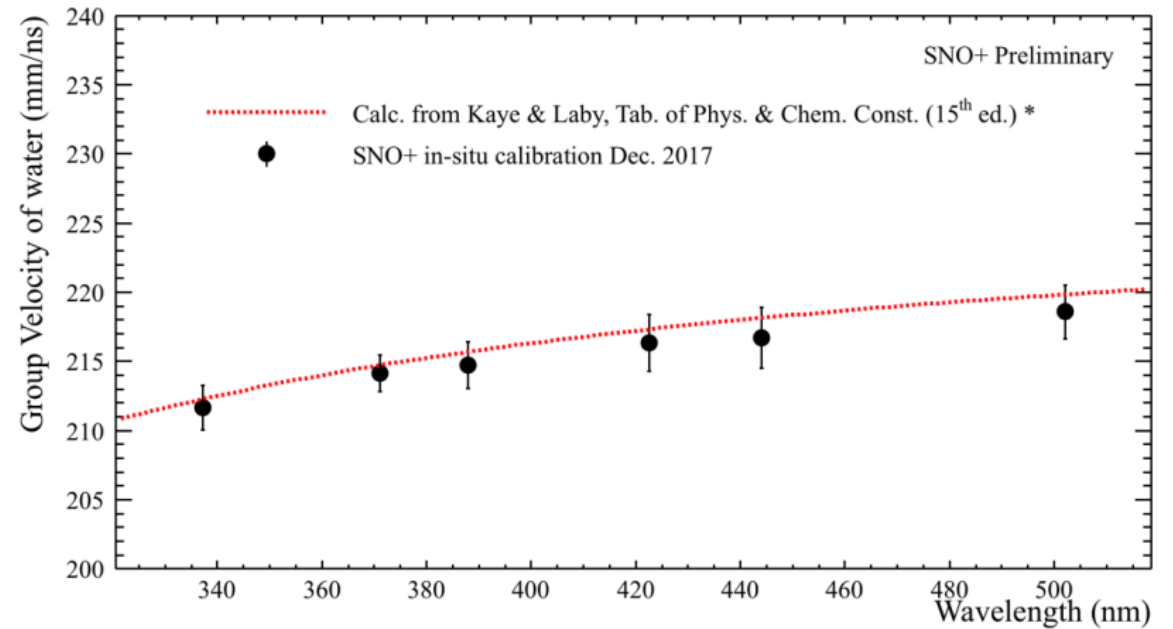
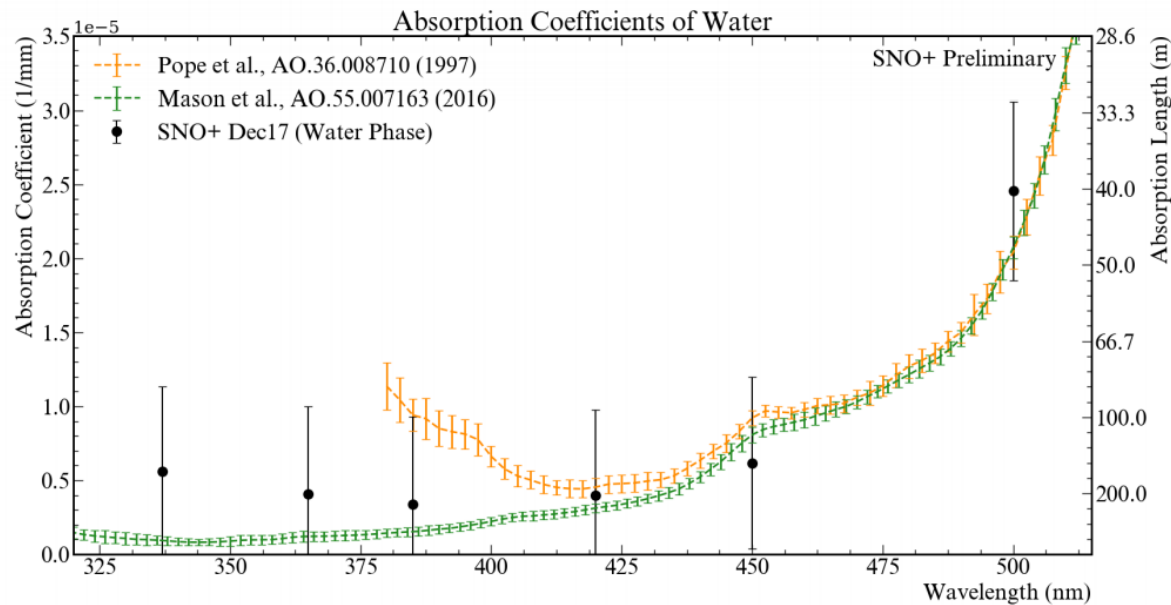
Current status

- Taking water data since early 2017
- Physics results include 114.7 days of livetime (of 235 calendar days)
 - Twice the amount of data collected in the mean time
- Extensive detector calibration with deployed and mounted sources



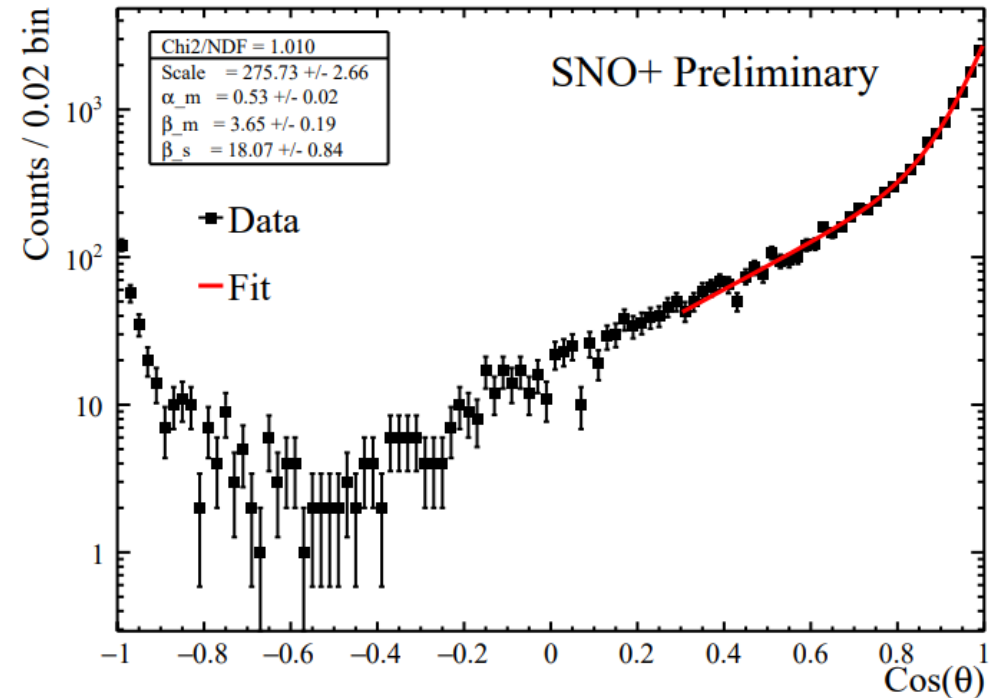
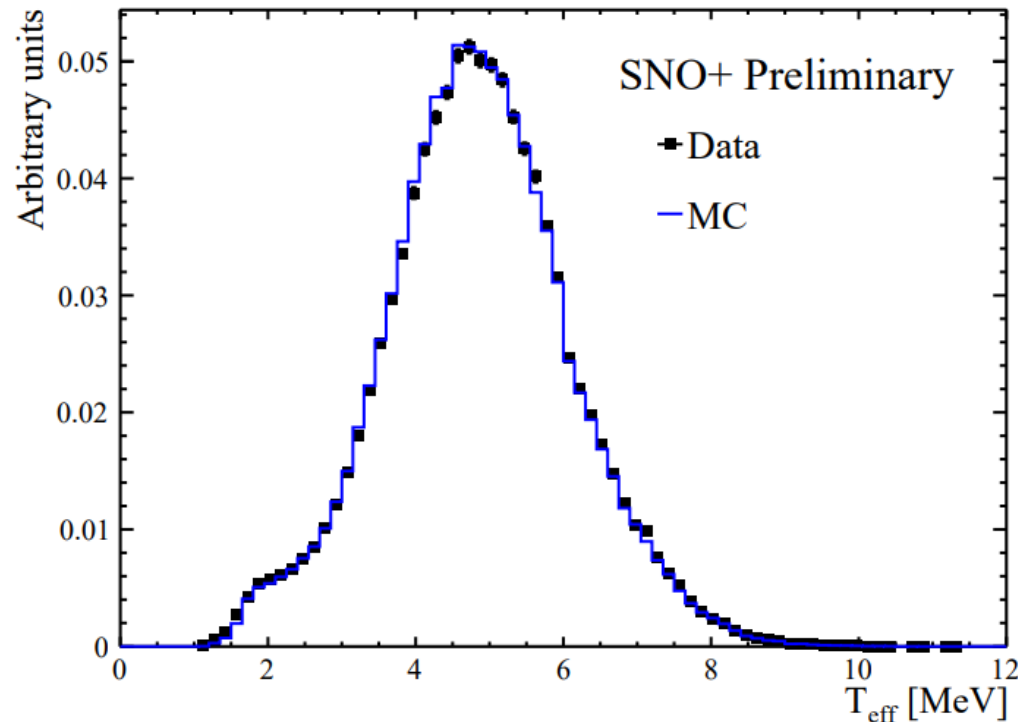
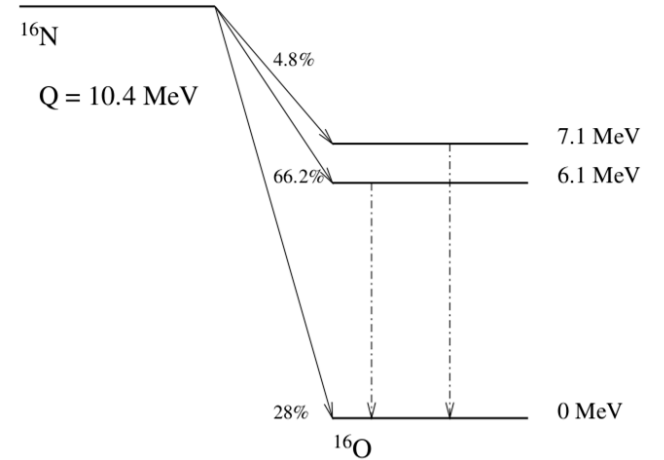
Optical calibration

- Isotropic light source deployed at multiple positions/wavelengths
- Full analysis of optical response in water conducted
 - Attenuation, group velocity, relative angular acceptance of optical sensors
- LED/laser systems mounted and being tested to do constant monitoring



Detector response

- Response calibrated using an ^{16}N source
 - Producing two gammas \rightarrow Compton scatter e^-
 - Data used to characterize energy response and fit algorithms (position, direction)



Solar flux measurement in water

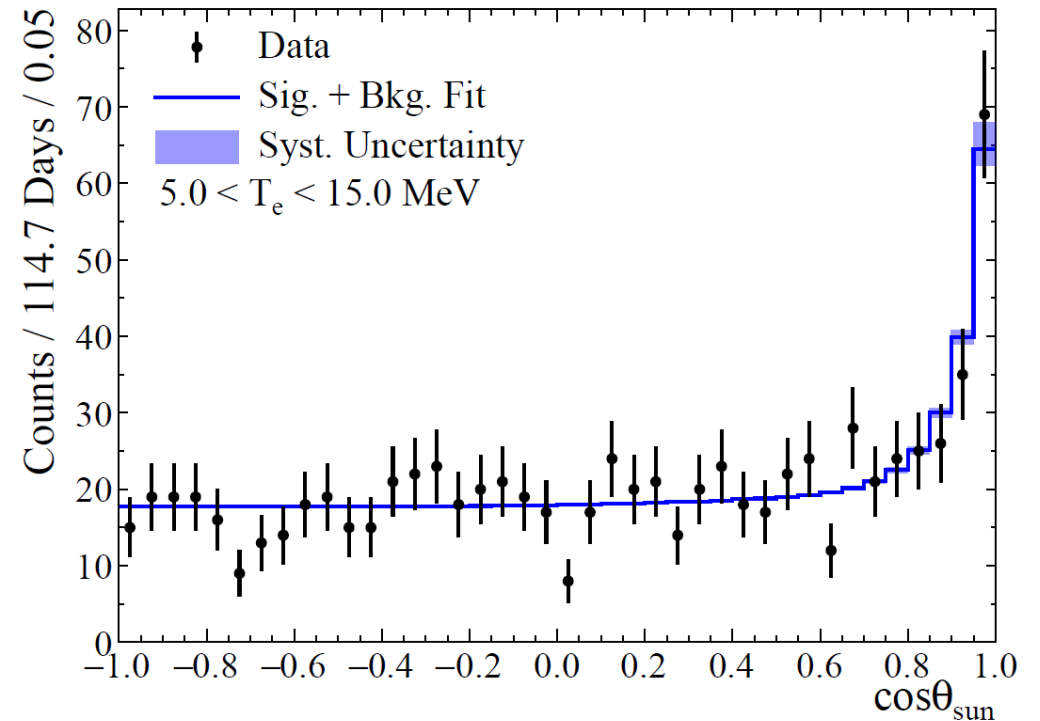
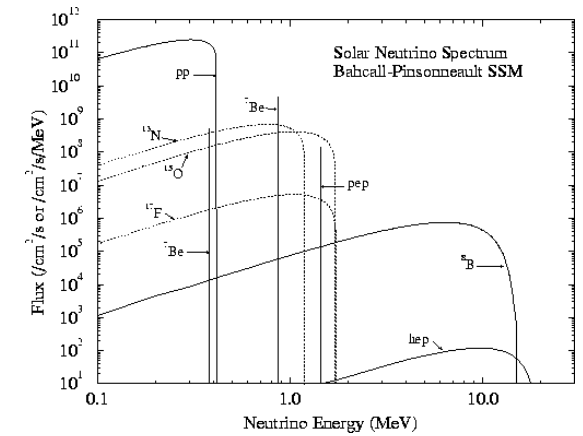
- Neutrinos from ^8B observed

- Flux

$$\Phi_{\text{sB}} = 5.95^{+0.75}_{-0.71}(\text{stat.})^{+0.28}_{-0.30}(\text{syst.}) \times 10^6 \text{cm}^{-2} \text{s}^{-1}.$$

consistent with SNO

- Fit in direction of $\cos(\theta_{\text{sun}})$, backgrounds are flat



Phys. Rev. D 99, 012012 (2019)

Solar flux measurement in water

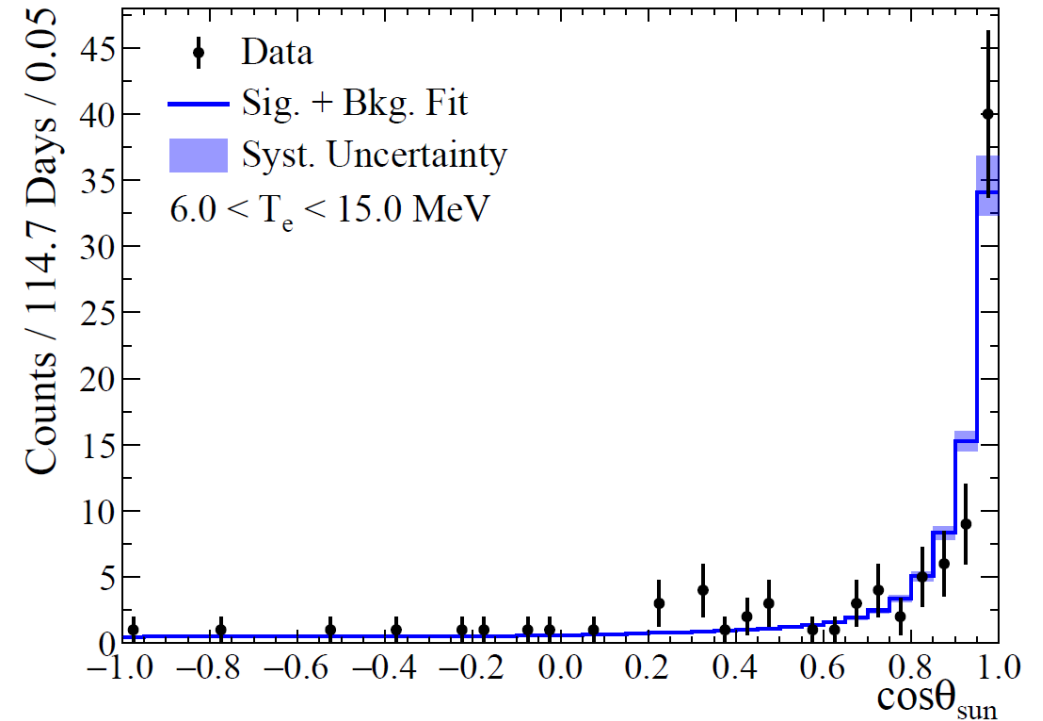
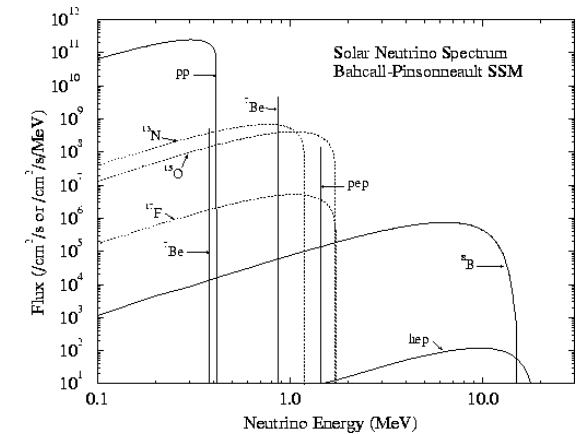
- Neutrinos from ^8B observed

- Flux

$$\Phi_{s_B} = 5.95_{-0.71}^{+0.75}(\text{stat.})_{-0.30}^{+0.28}(\text{syst.}) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}.$$

consistent with SNO

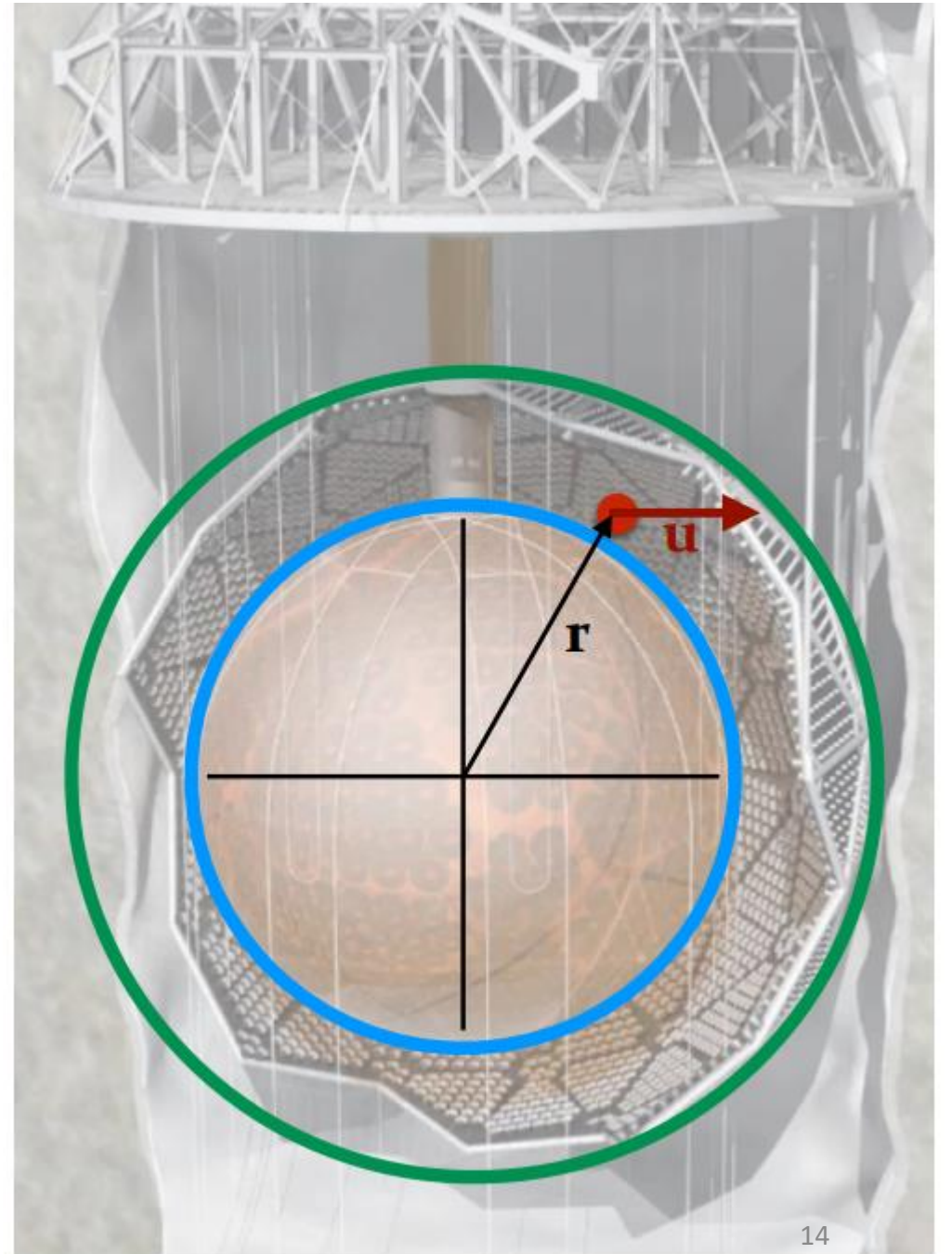
- Fit in direction of $\cos(\theta_{\text{sun}})$,
backgrounds are flat



Phys. Rev. D 99, 012012 (2019)

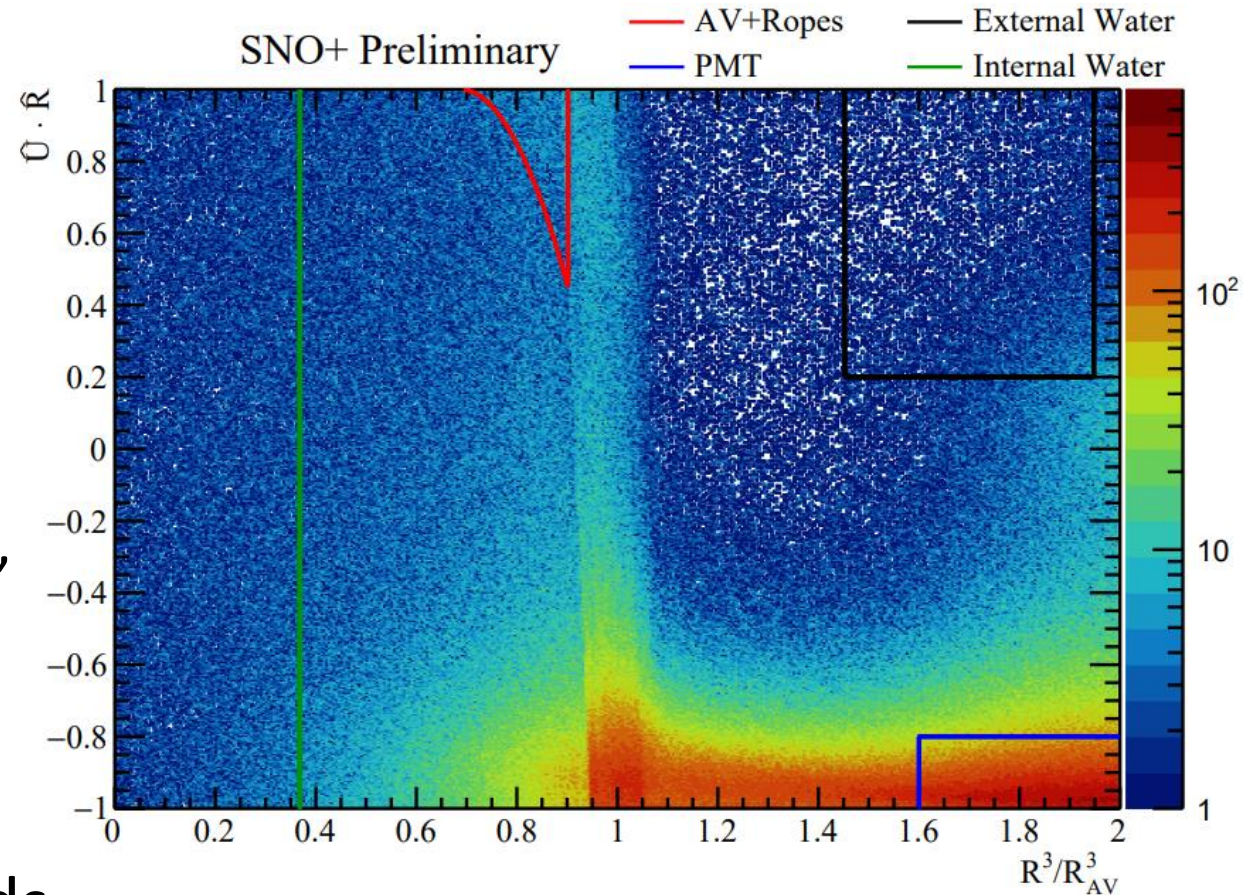
Background levels

- Internal backgrounds
 - Intrinsic radioactivity of water in the vessel
- External backgrounds
 - Intrinsic radioactivity in vessel, ropes, PMTs and water
 - Will not change in between phases
- Observables to identify backgrounds
 - Energy, position, isotropy and direction



Background levels

- Internal backgrounds
 - Intrinsic radioactivity of water in the vessel
- External backgrounds
 - Intrinsic radioactivity in vessel, ropes, PMTs and water
 - Will not change in between phases
- Observables to identify backgrounds
 - Energy, position, isotropy and direction



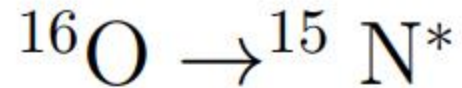
External background results

- Multiple analyses constrain the backgrounds
 - Fit to spectral shapes, counting events within a region
- Results are consistent, indicate external background at expectation

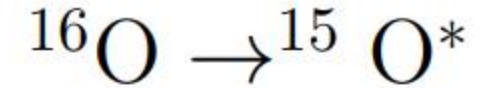
Background source	Results (observed / expected) for latest period analyzed	
	$z > 0$ (upper hemisphere)	$z < 0$ (lower hemisphere)
Acrylic vessel + rope system	2.2 ± 0.08 (stat) $^{+2.4}_{-1.9}$ (syst)	1.3 ± 0.08 (stat) $^{+1.0}_{-0.9}$ (syst)
External water	0.6 ± 0.06 (stat) $^{+1.9}_{-0.6}$ (syst)	1.0 ± 0.07 (stat) $^{+3.3}_{-1.0}$ (syst)
PMTs	1.2 ± 0.02 (stat) $^{+1.1}_{-0.5}$ (syst)	

Invisible nucleon decay search

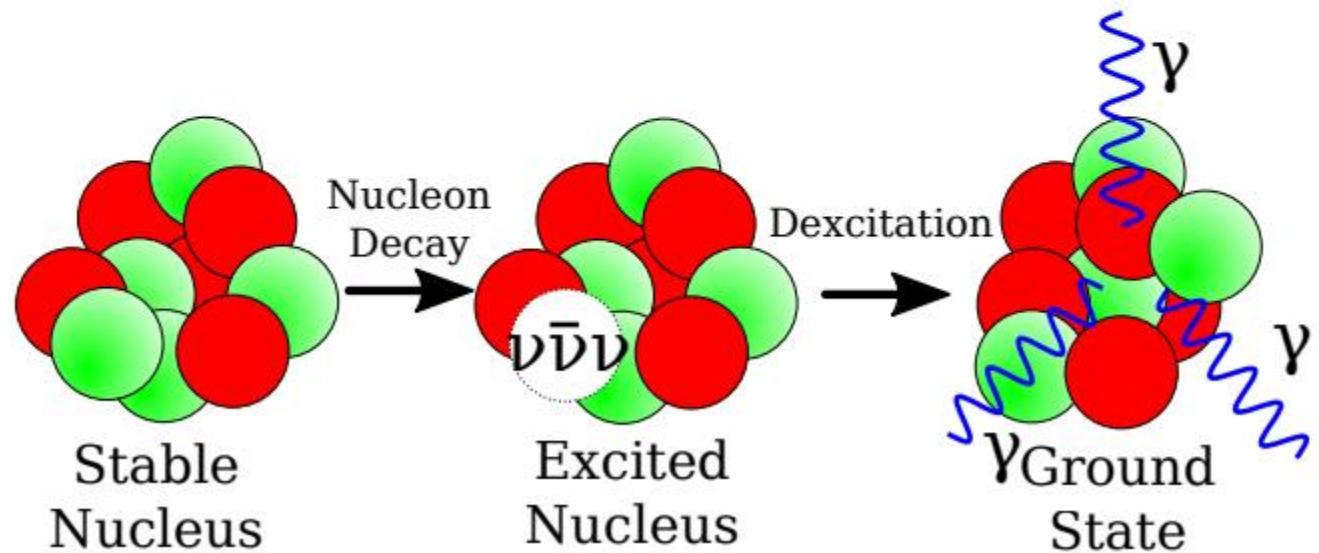
- Never observed baryon-number violating process
- Theories propose invisible decay modes (e.g. $n \rightarrow 3\nu$)
- Decay could be observed indirectly with gammas



3%: 9.932 MeV γ
41%: 6.32 MeV γ



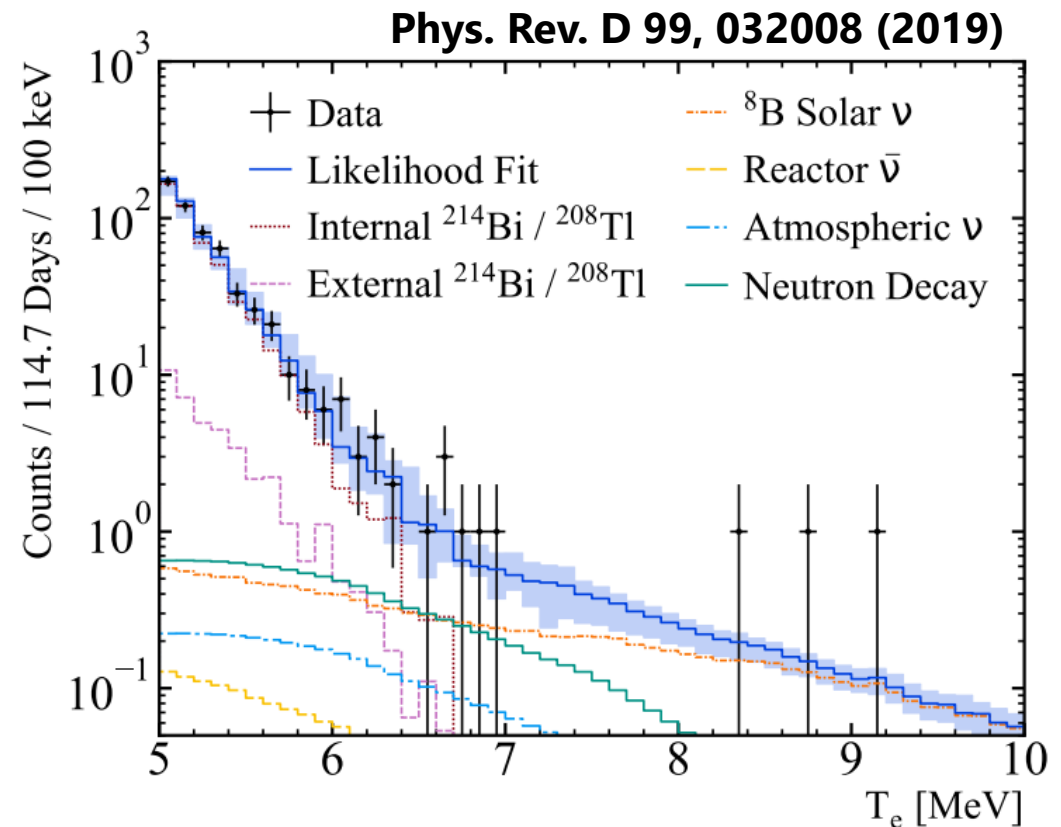
44%: 6.18 MeV γ



Results from nucleon decay search

- Selection based on straight cuts to remove backgrounds
- Two analysis performed: cut & count and likelihood fit
- No excesses, only limits

Data set	Observed events	Expected events	
1	1	$1.17^{+4.60}_{-0.05}$	$+1.33_{-0.39}$
2	2	$2.35^{+4.62}_{-0.40}$	$+3.44_{-0.81}$
3	4	$3.47^{+4.60}_{-0.15}$	$+3.11_{-0.96}$
4	8	$3.37^{+4.60}_{-0.17}$	$+2.70_{-0.98}$
5	1	$1.46^{+4.60}_{-0.13}$	$+2.17_{-0.60}$
6	6	$5.84^{+7.40}_{-2.31}$	$+2.68_{-0.62}$
Total	22	$17.65^{+12.68}_{-2.36}$	$+6.51_{-1.85}$



Limits on nucleon decay

- Results and comparison with existing limits

	Spectral analysis	Counting analysis	Existing limits
n	$2.5 \times 10^{29} \text{ y}$	$2.6 \times 10^{29} \text{ y}$	$5.8 \times 10^{29} \text{ y}$ [KamLAND]
p	$3.6 \times 10^{29} \text{ y}$	$3.4 \times 10^{29} \text{ y}$	$2.1 \times 10^{29} \text{ y}$ [SNO]
pp	$4.7 \times 10^{28} \text{ y}$	$4.1 \times 10^{28} \text{ y}$	$5.0 \times 10^{25} \text{ y}$ [Borexino]
pn	$2.6 \times 10^{28} \text{ y}$	$2.3 \times 10^{28} \text{ y}$	$2.1 \times 10^{25} \text{ y}$ [Treyak et al.]
nn	$1.3 \times 10^{28} \text{ y}$	$0.6 \times 10^{28} \text{ y}$	$1.4 \times 10^{30} \text{ y}$ [KamLAND]

Phys. Rev. D 99, 032008 (2019)

Next: scintillator fill

- Cover-gas system in place to seal the vessel
- Radon monitor in place to follow activity
- Liquid scintillator purification plant being tested
 - About 1.8 tonnes injected in the vessel thus far
- Fill had to be halted several months due to a leak in the distillation column
- Leak has been repaired – about to restart fill very soon



Tellurium process systems installed

- TeA and TeDiol plants moving to commissioning

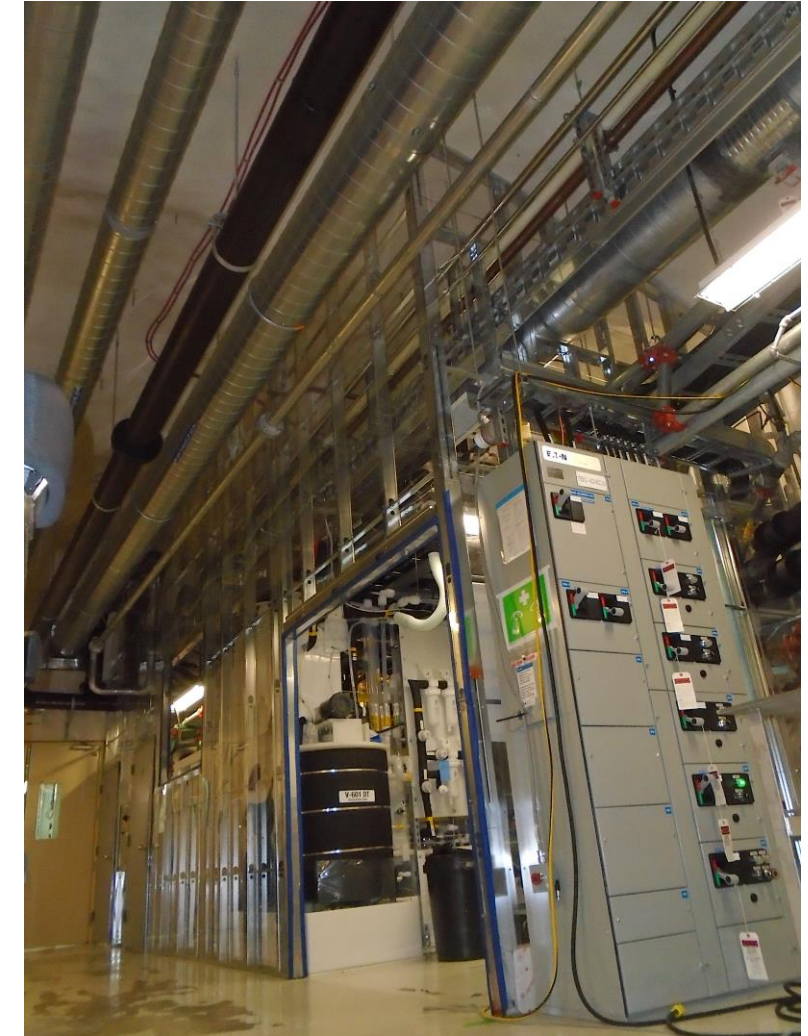
2016



2017



2018



Telluric acid plant

Tellurium process systems installed

- TeA and TeDiol plants moving to commissioning



2018

TeDiol plant

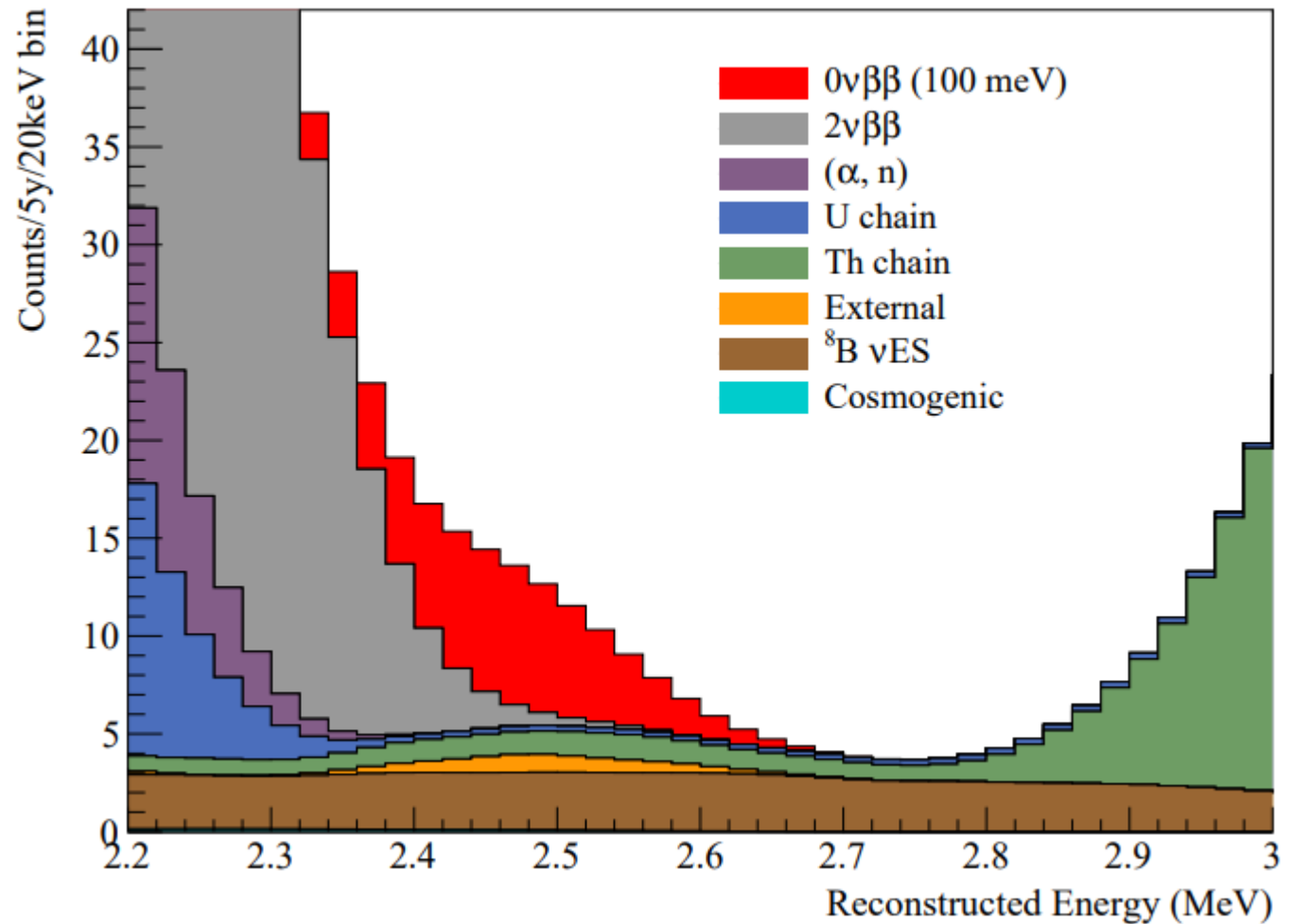


2019

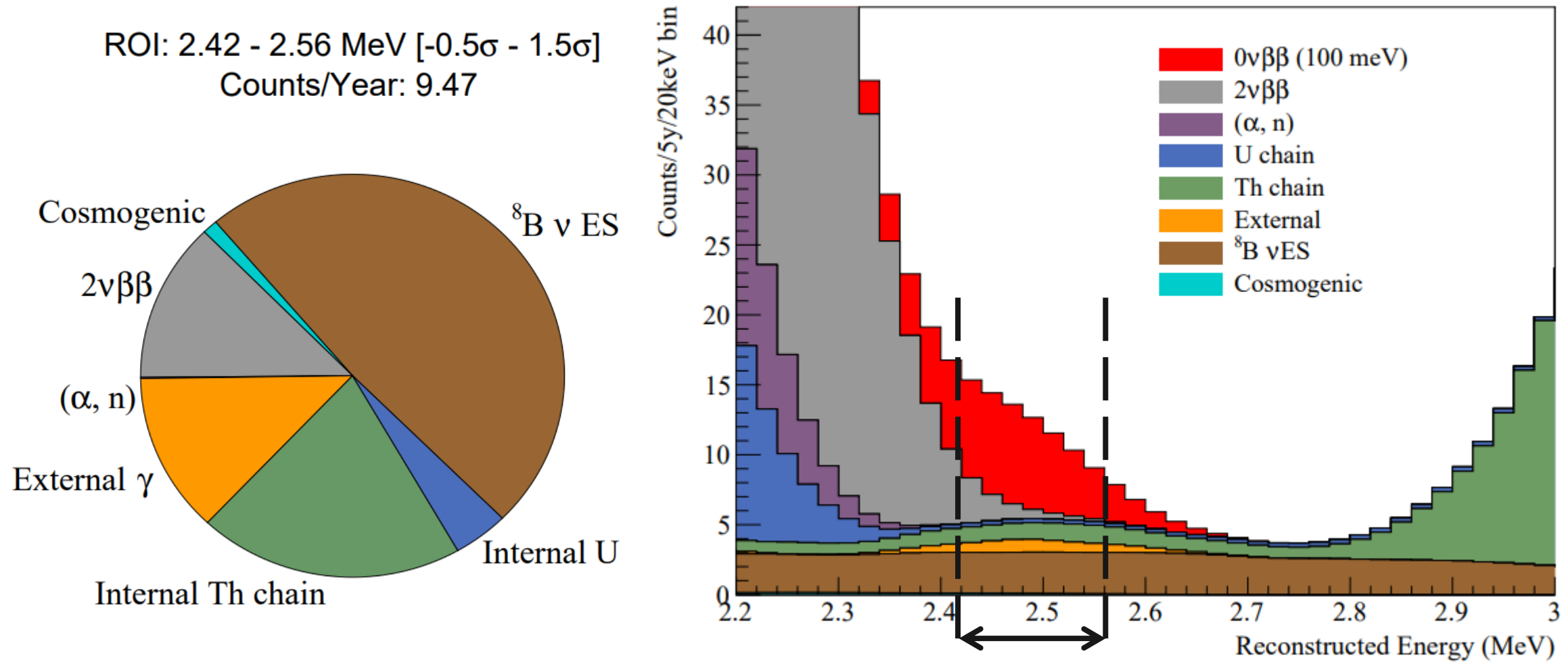


Majorana neutrino search

- Signal of Majorana neutrinos in the energy spectrum
- Using best knowledge of background levels
- Two analyses planned
 - Cut&count
 - Likelihood

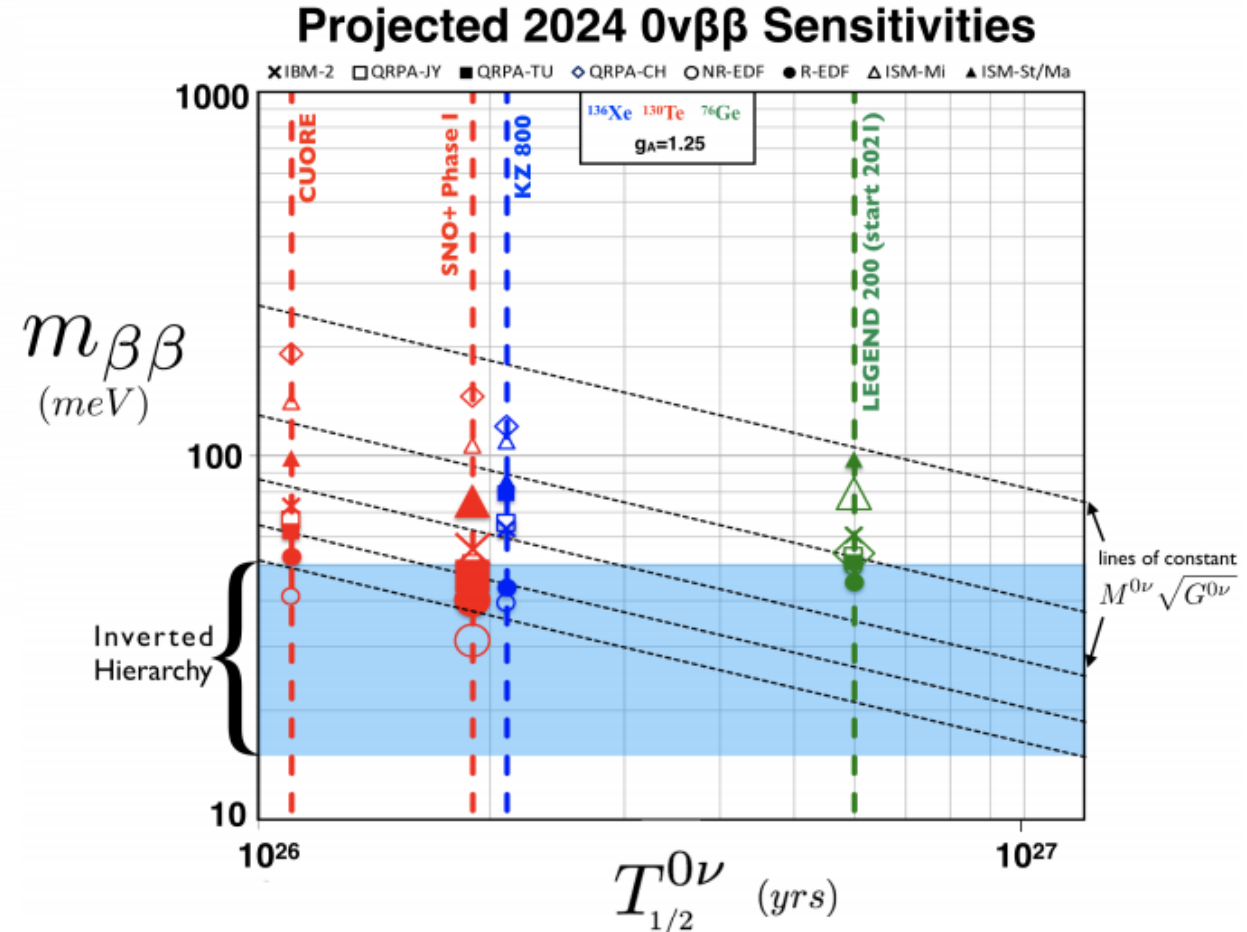
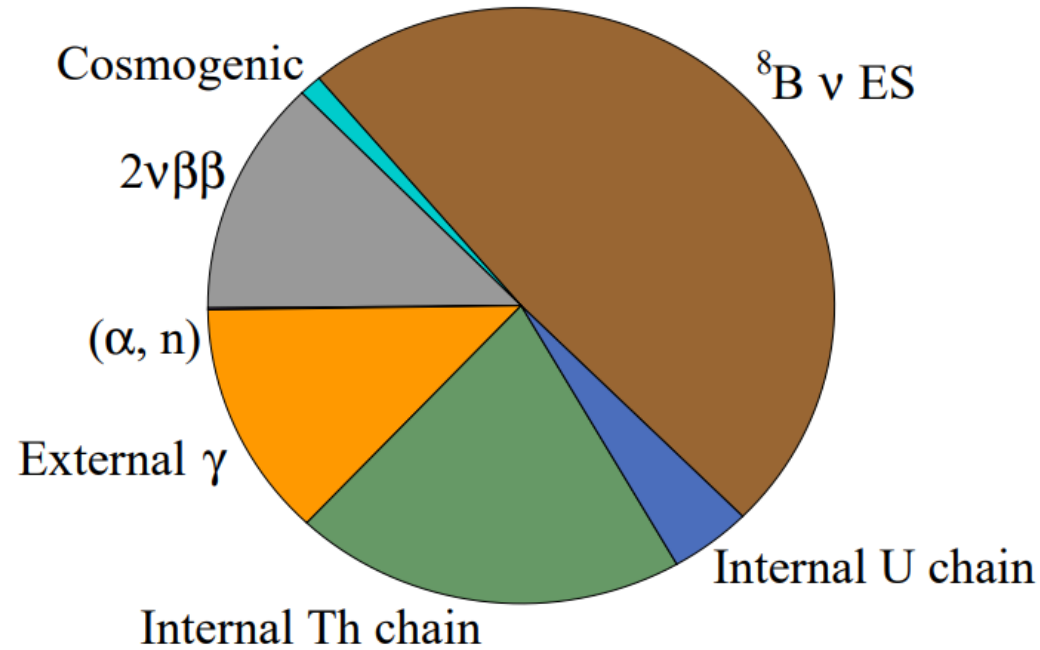


Majorana neutrino search



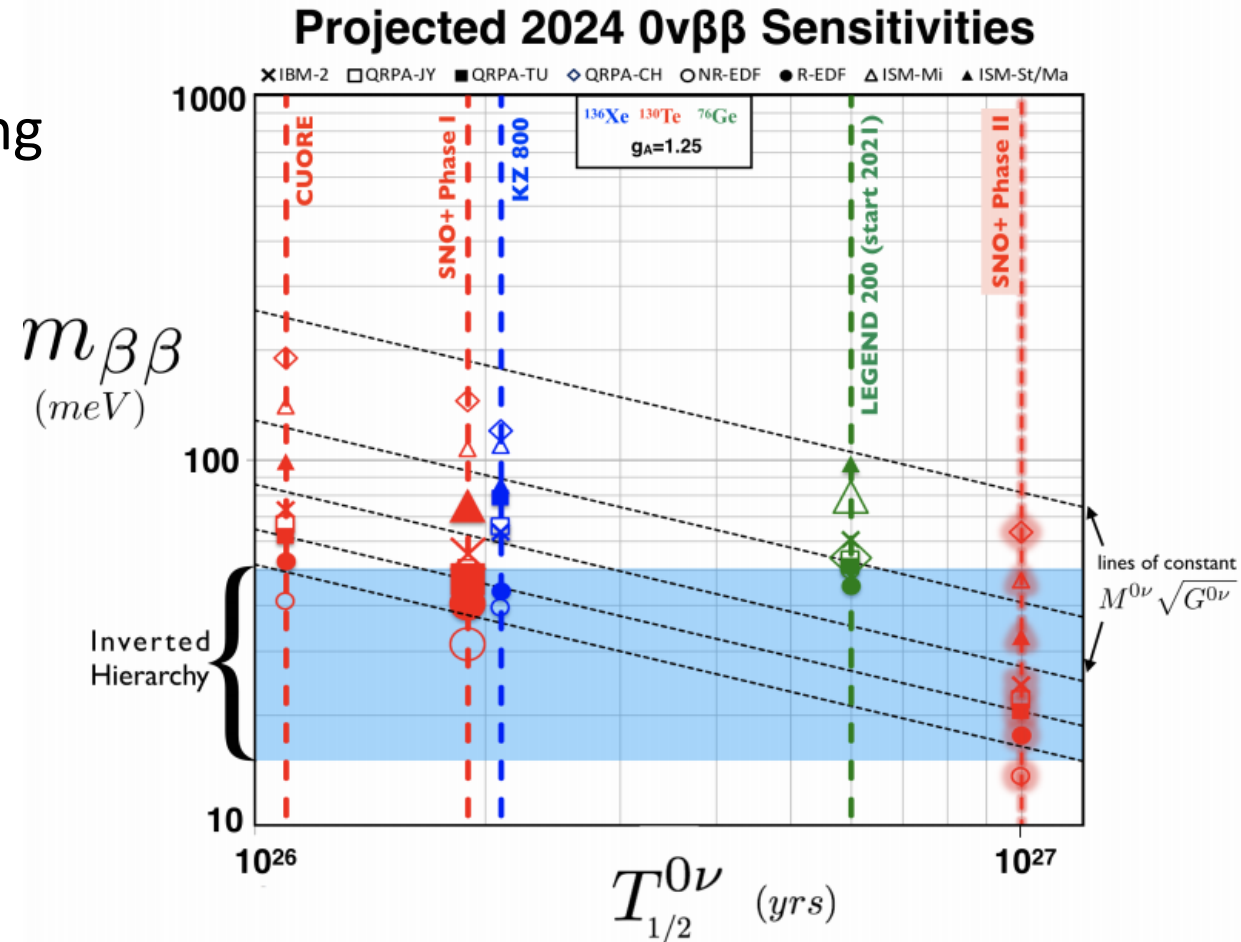
Majorana neutrino search

ROI: 2.42 - 2.56 MeV $[-0.5\sigma - 1.5\sigma]$
 Counts/Year: 9.47



Towards the future

- SNO+ Phase I is 0.5% Te-loaded
 - Main backgrounds aren't from loading
 - Test-bed for a multi-ton experiment
- Telluric acid loading R&D
 - Chemistry studies to increase light yield
- Detector upgrade path
 - High QE PMTs being studied
 - Possible to replace PMT focusing



Summary

- Water phase wrapping up
 - Calibration systems tested
 - Detector response understood
 - Modeling of sub-leading effects with low background data
- Results from the water phase out
 - Nucleon decay limits and observation of solar neutrinos published
 - Working on neutron capture and reactor antineutrinos in water
- Scintillator phase to begin soon
 - Loading liquid scintillator this year



University of Alberta
Queen's University
Laurentian University
TRIUMF
SNOLAB

Boston University
Brookhaven National Laboratory
University of California Berkeley
Lawrence Berkeley National Laboratory
University of Chicago
University of Pennsylvania
University of California Davis



LIP Coimbra
LIP Lisboa

UNAM



TU Dresden



Oxford University
Queen Mary University of London
University of Liverpool
University of Sussex
University of Lancaster

Thank you for your attention