



DAMPE space mission and recent results

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On Behalf of the DAMPE collaboration

Gran Sasso Science Institute (GSSI)

& INFN Laboratori Nazionali del Gran Sasso

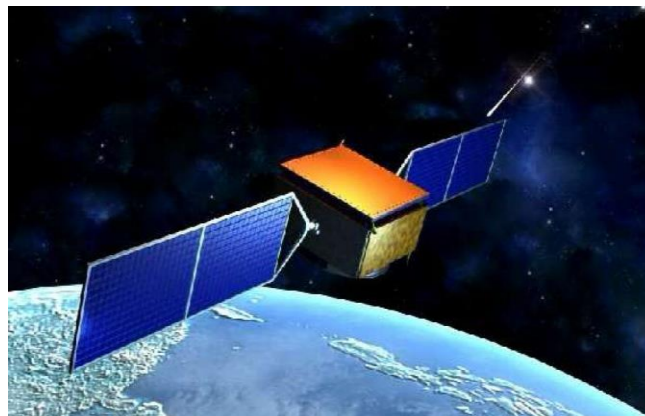
Speaker: Z. Wang



WIN2019 The 27th International Workshop on weak Interactions and Neutrinos

Overview

- Introduction of the collaboration and scientific goals of DArk Matter Particle Explorer (DAMPE)
- DAMPE structure and functionality parameters
- Recent results concerning spectra of electron, proton and helium
- Brief introduction on our analysis on proton + helium spectrum



The Collaboration



Launched on December 17th 2015, DAMPE has been collecting CR data for more than 3 years!

- **CHINA**

- Purple Mountain Observatory, CAS, Nanjing
- University of Science and Technology of China, Hefei
- Institute of High Energy Physics, CAS, Beijing
- National Space Science Center, CAS, Beijing
- Institute of Modern Physics, CAS, Lanzhou



- **ITALY**

- INFN Perugia and University of Perugia
- INFN Bari and University of Bari
- INFN Lecce and University of Salento
- INFN LNGS and Gran Sasso Science Institute



- **SWITZERLAND**

- University of Geneva

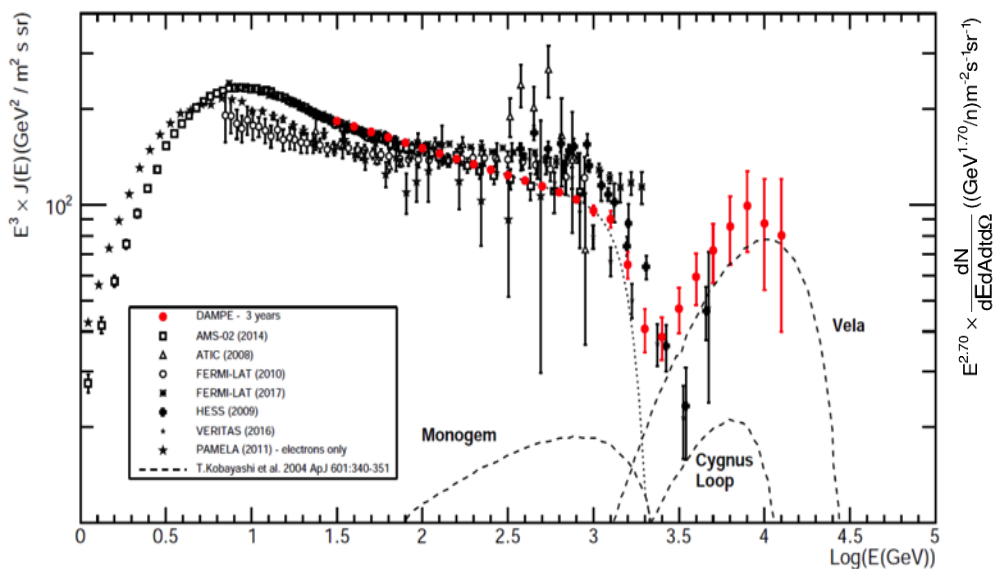


The scientific goals

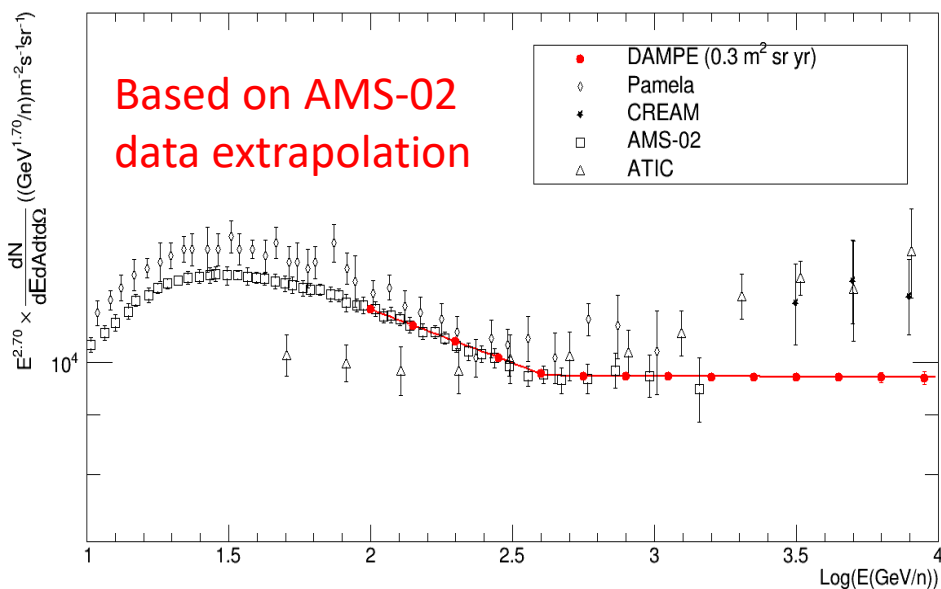
High energy particle detection in space

- Study of the cosmic-ray electron spectra
- Study of cosmic-ray protons and nuclei: spectrum and composition
- High energy gamma ray astronomy
- Search for dark matter signatures in lepton spectra

DAMPE expected electron+positron spectrum



DAMPE expected proton spectrum



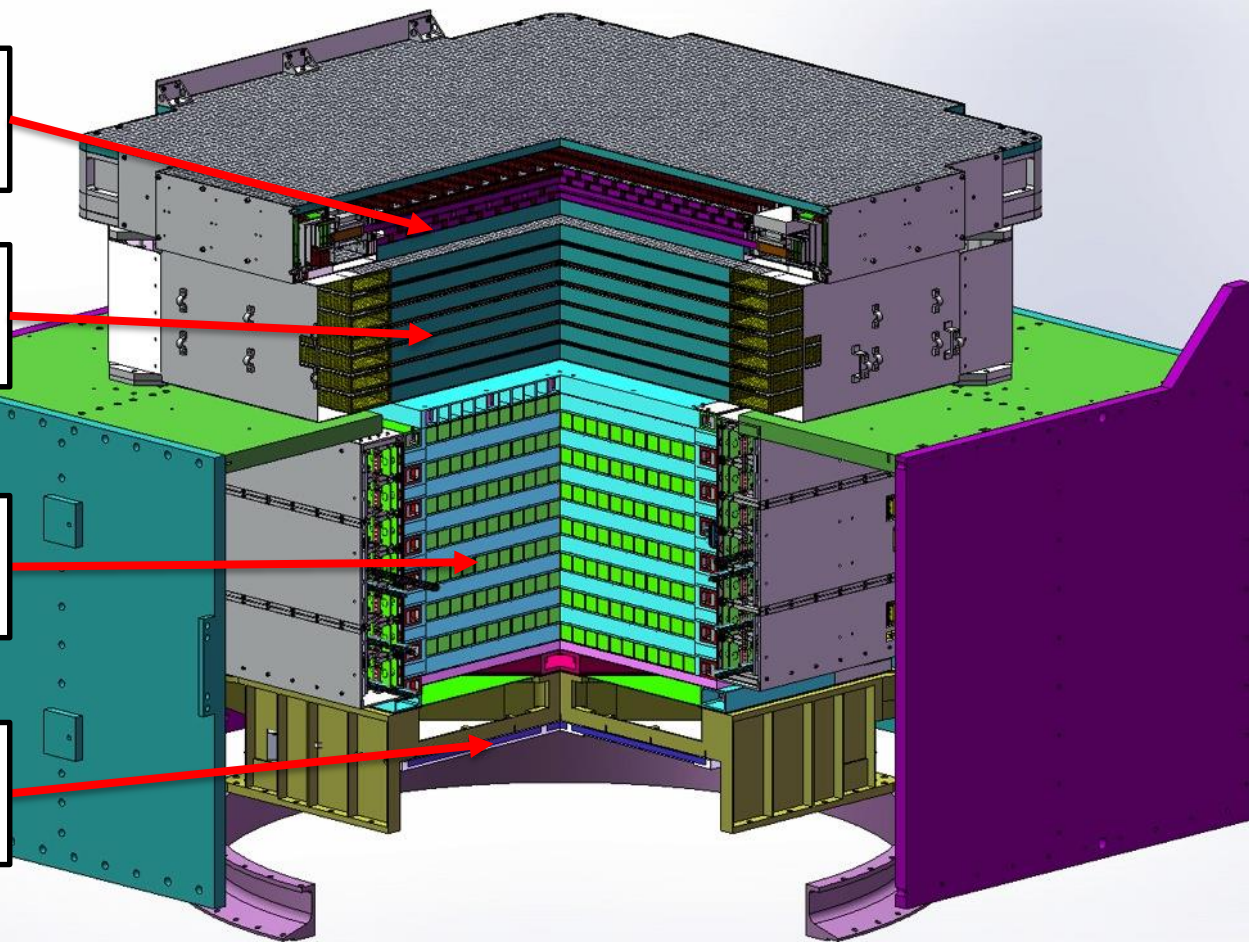
The detector structure

Plastic Scintillator Detector
(PSD)

Silicon-Tungsten Tracker
(STK)

BGO Calorimeter
(BGO)

Neutron Detector
(NUD)



PSD: Charge measurement; Identify electron and γ -ray;

STK: Tungsten converter (pair production); Precise tracking (silicon strips);

BGO: Energy measurement; e/p separation;

NUD: Hadron rejection;

The detector parameters

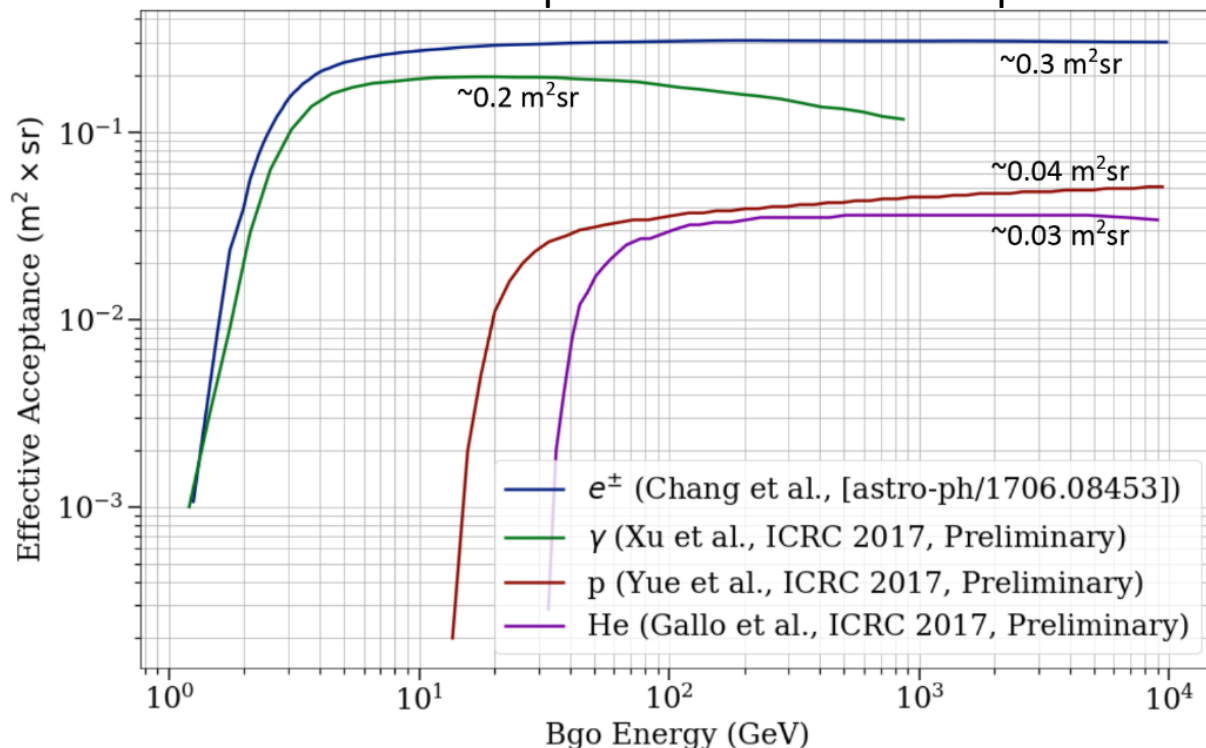
DAMPE main parameters

| Parameter | Value |
|--|------------------------------------|
| Energy range of γ -rays/electrons | 5 GeV–10 TeV |
| Energy resolution ^a of γ -rays/electrons | $\leq 1.5\%$ at 800 GeV |
| Energy range of protons/heavy nuclei | 50 GeV–100 TeV |
| Energy resolution ^a of protons | $\leq 40\%$ at 800 GeV |
| Effective area at normal incidence (γ -rays) | 1100 cm ² at 100 GeV |
| Geometric factor for electrons | 0.3 m ² sr above 30 GeV |
| Photon angular resolution ^b | $\leq 0.2^\circ$ at 100 GeV |
| Field of View (FoV) | ~ 1.0 sr |

Comparison with AMS-02 and Fermi LAT

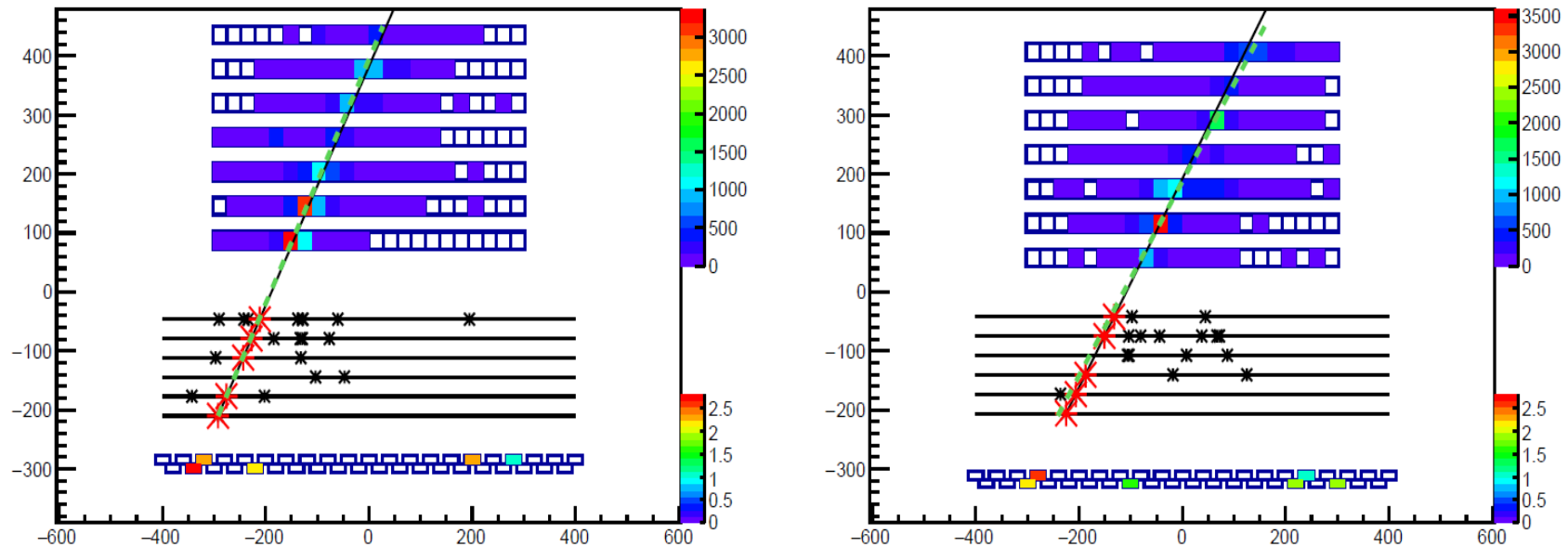
| | DAMPE | AMS-02 | Fermi LAT |
|--|--|---------------|-----------|
| e/ γ Energy res.@100 GeV (%) | 1.2 | 3 | 10 |
| e/ γ Angular res.@100 GeV (deg) | 0.2 | 0.3 | 0.1 |
| e/p discrimination | 10^5-10^6 | $10^5 - 10^6$ | 10^3 |
| Calorimeter thickness (X_0) | 32 | 17 | 8.6 |
| Geometrical accep. (m ² sr) | 0.3 | 0.09 | 1 |

DAMPE effective acceptance for different CR particles



An example for one triggered event

A proton event, seen from XZ (left) and YZ (right) view. The event releases 40.518 GeV energy inside the BGO calorimeter.

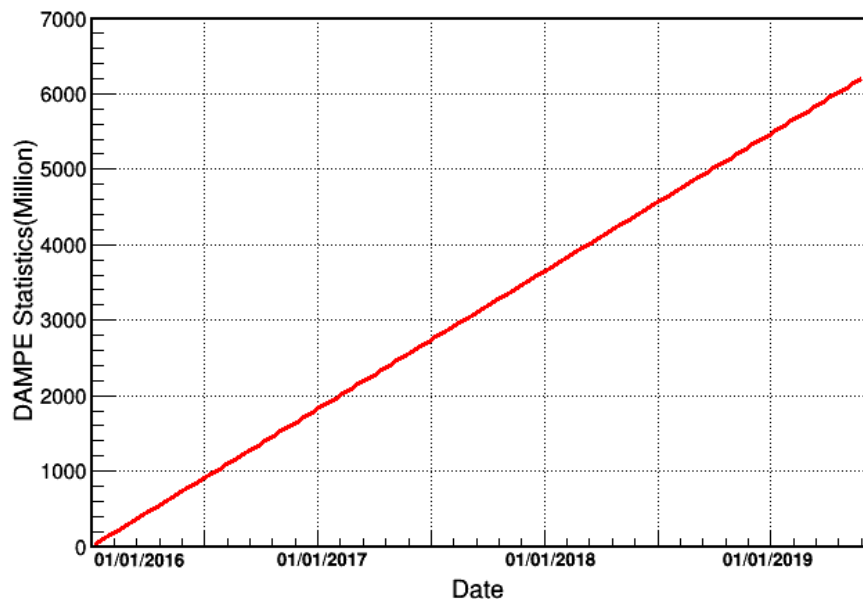


Green dash line: BGO track, Black line: STK track.

Track, energy deposition and charge measurement information are combined together to reconstruct the primary information of the events. (Including the energy, category and direction etc.)

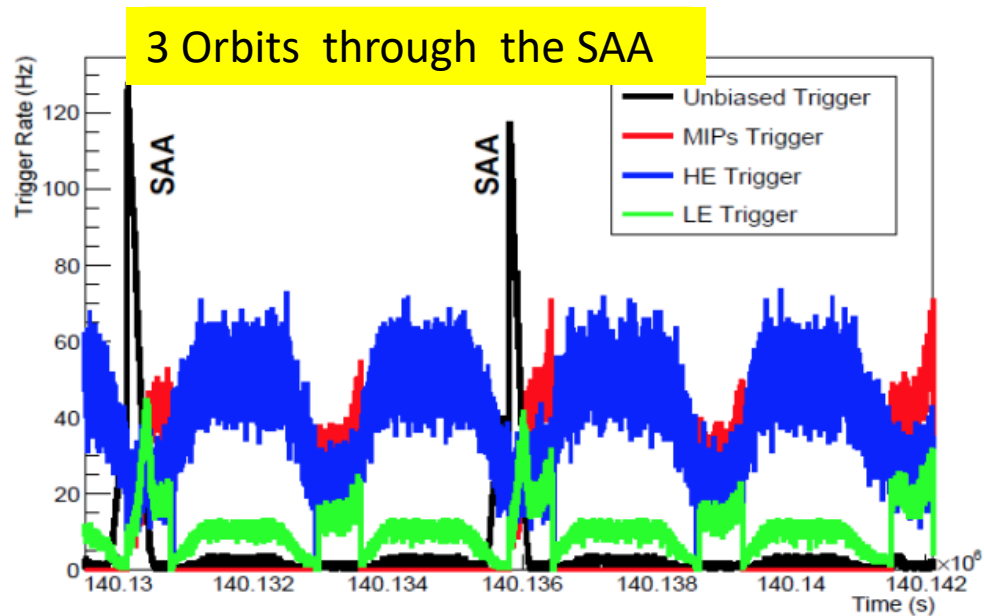
Present status

DAMPE statistics



Up to now, DAMPE has collected more than 6×10^9 events. On average, 5 million events are collected for each day.

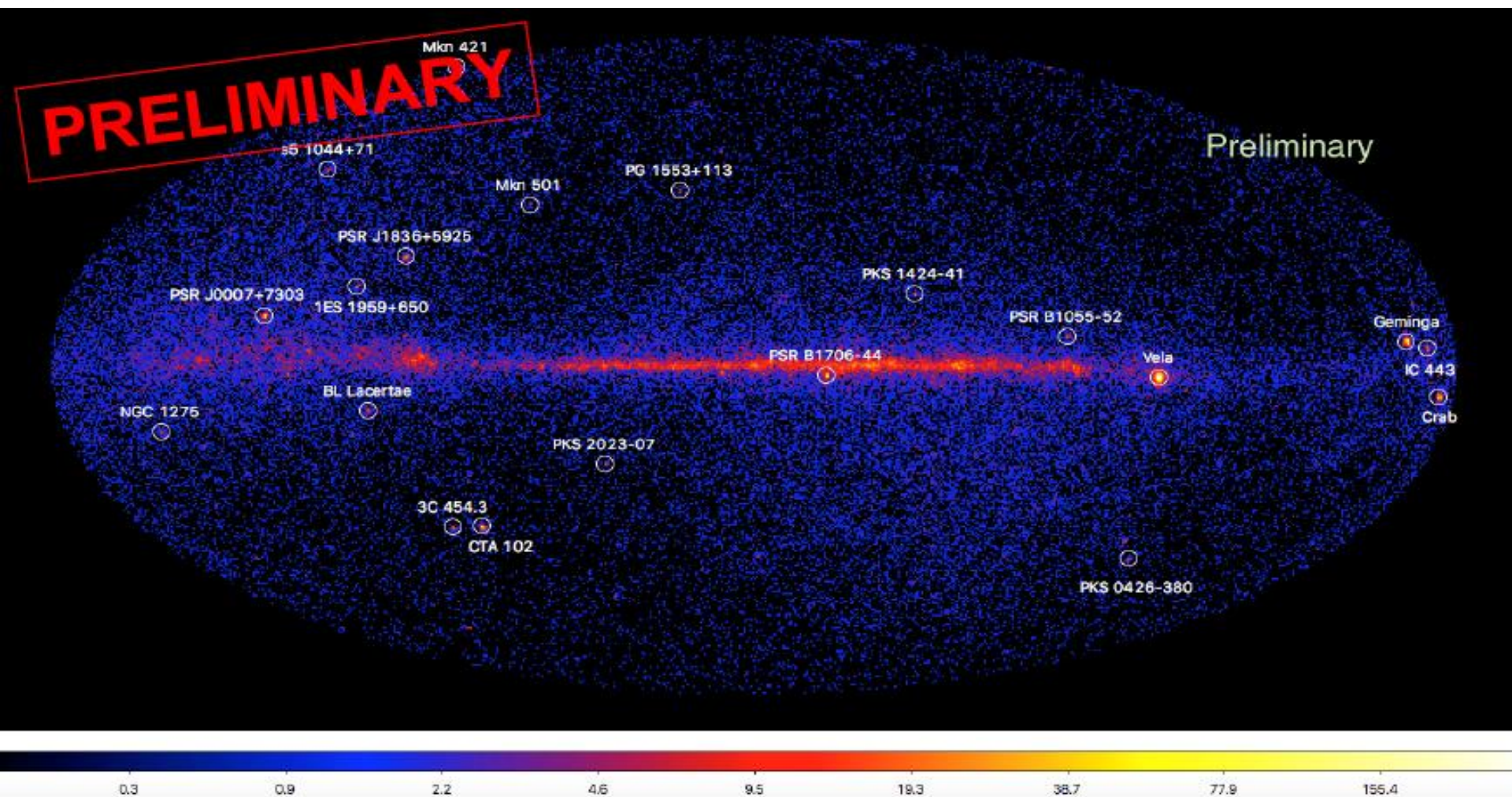
DAMPE trigger rates



High energy trigger rate: 20-60 Hz

The data taking progress is still going on steadily.

The gamma ray sky



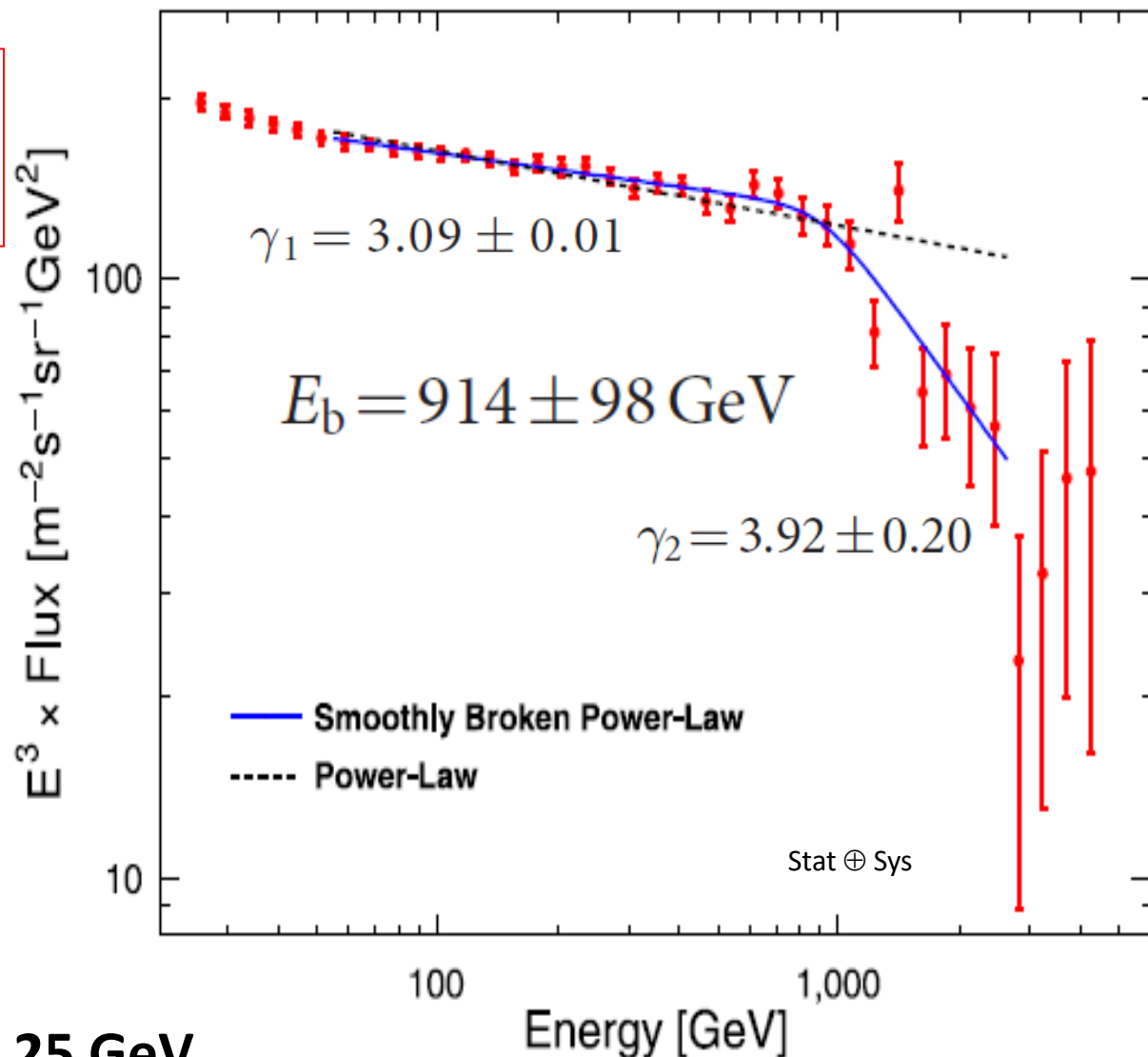
510 days counts map. Mollweide projection, $0.5^\circ \times 0.5^\circ$ pixels
 $E > 2\text{GeV}$ 90,000 events O(20) sources detected

The DAMPE ($e^+ + e^-$) spectrum

LETTER nature
 International weekly journal of science
 10.1038/nature24475
 Direct detection of a break in the teraelectronvolt cosmic-ray spectrum of electrons and positrons
 DAMPE Collaboration*

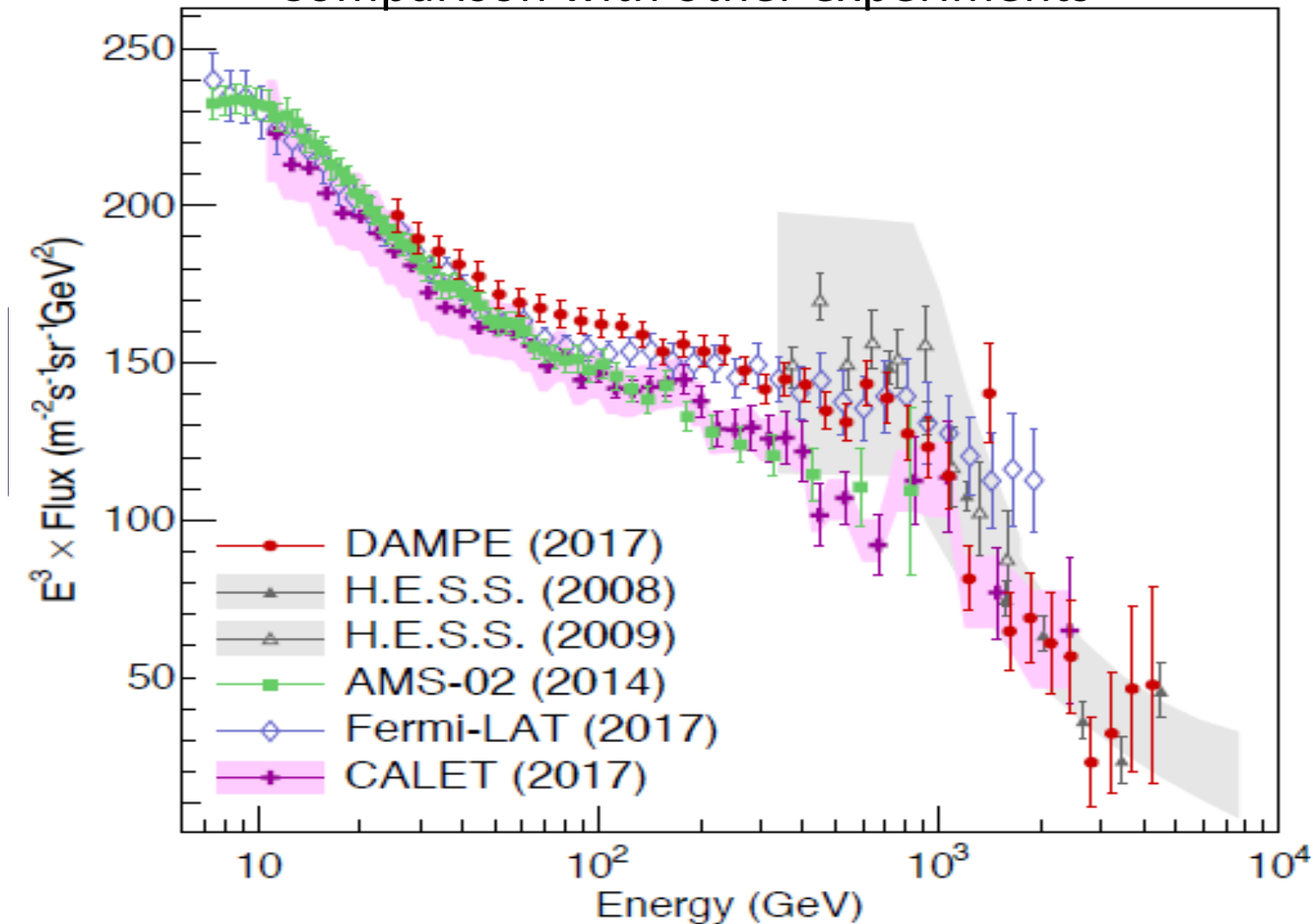
**First Direct Evidence for
 a spectral break in the
 all-electron spectrum at
 0.9 TeV**

- 530 days
- 2.8 billions CR events
- 1.5 million CREs above 25 GeV

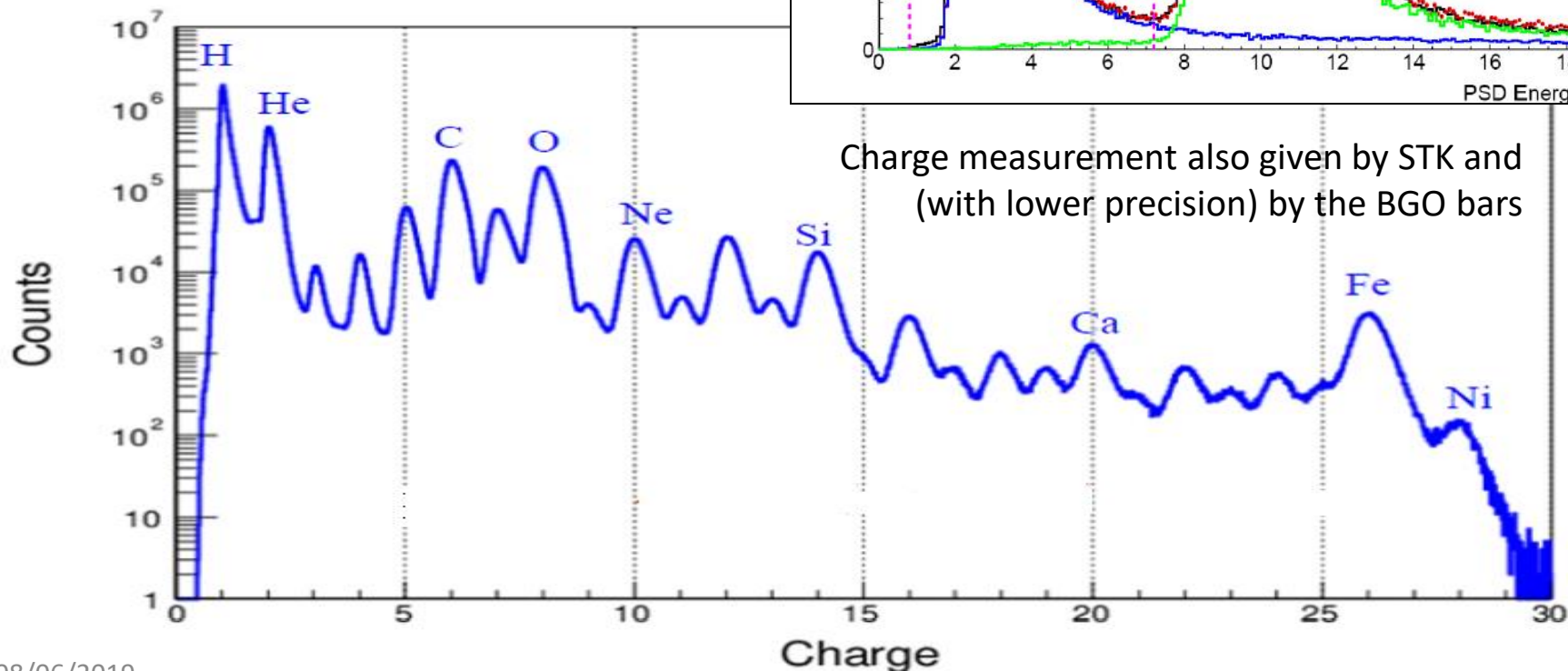
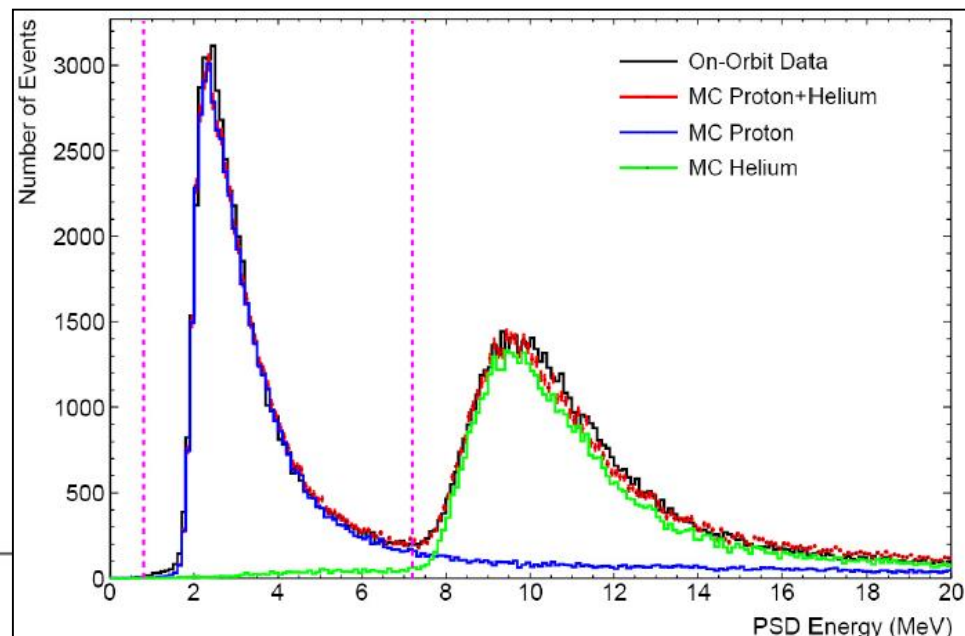
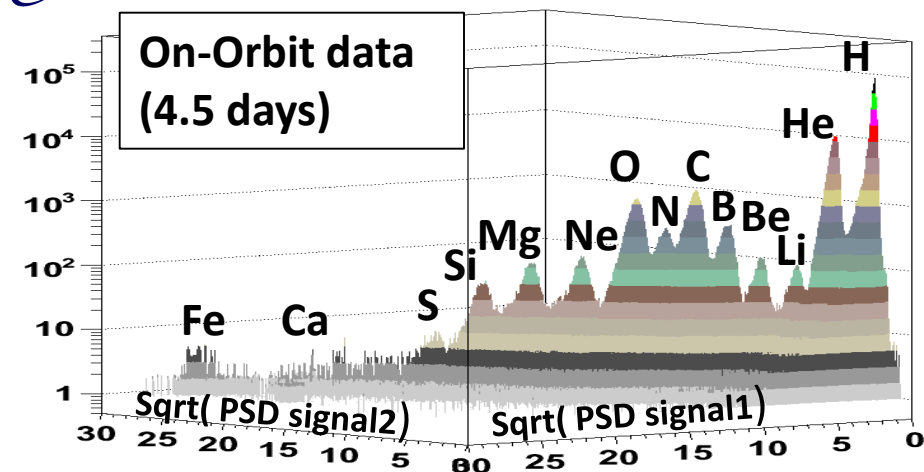


The all-electron spectrum

Comparison with other experiments



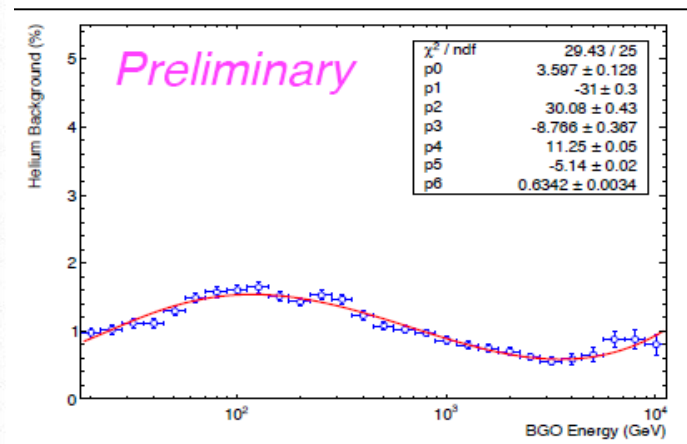
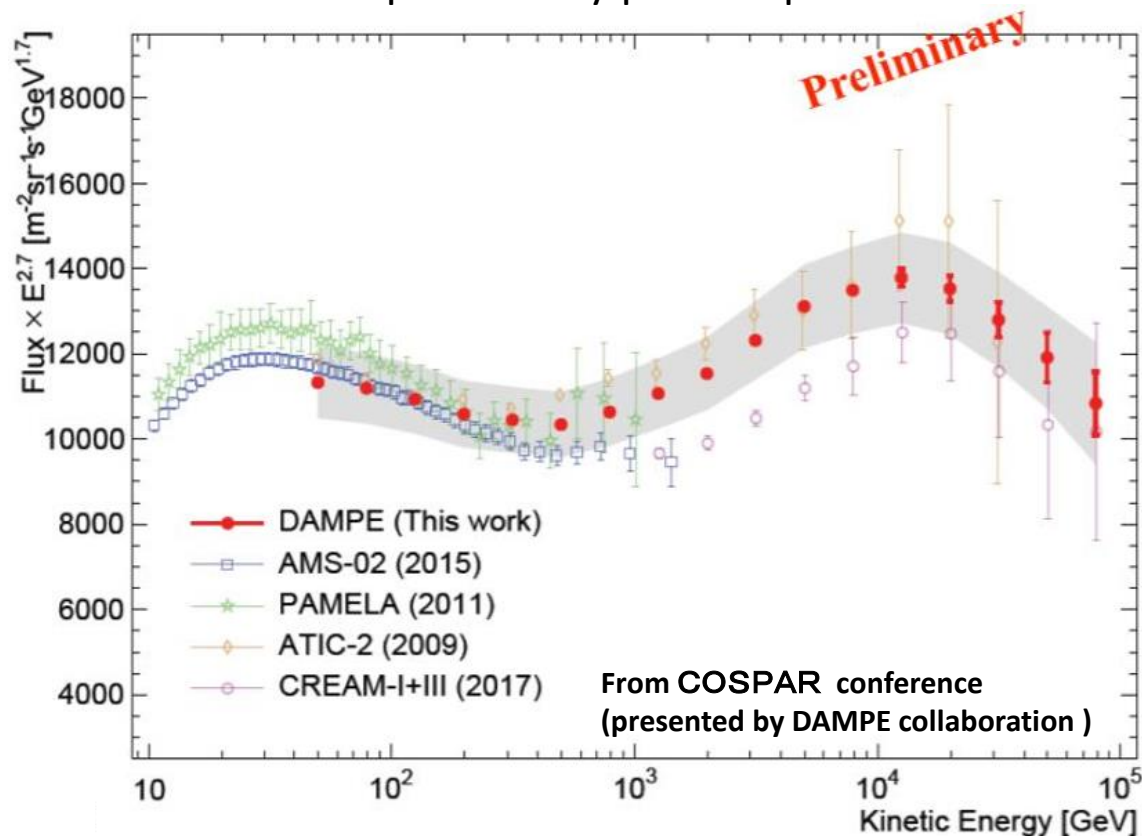
Nuclei ID with PSD



Charge measurement also given by STK and
(with lower precision) by the BGO bars

Proton flux measurement

DAMPE preliminary proton spectrum



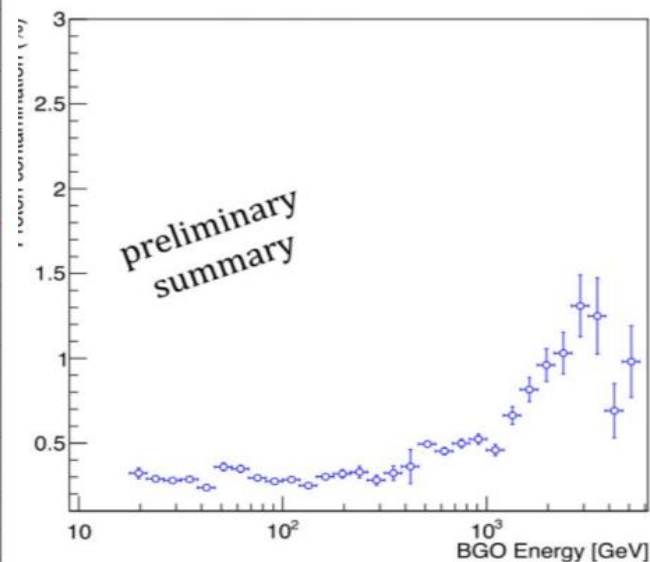
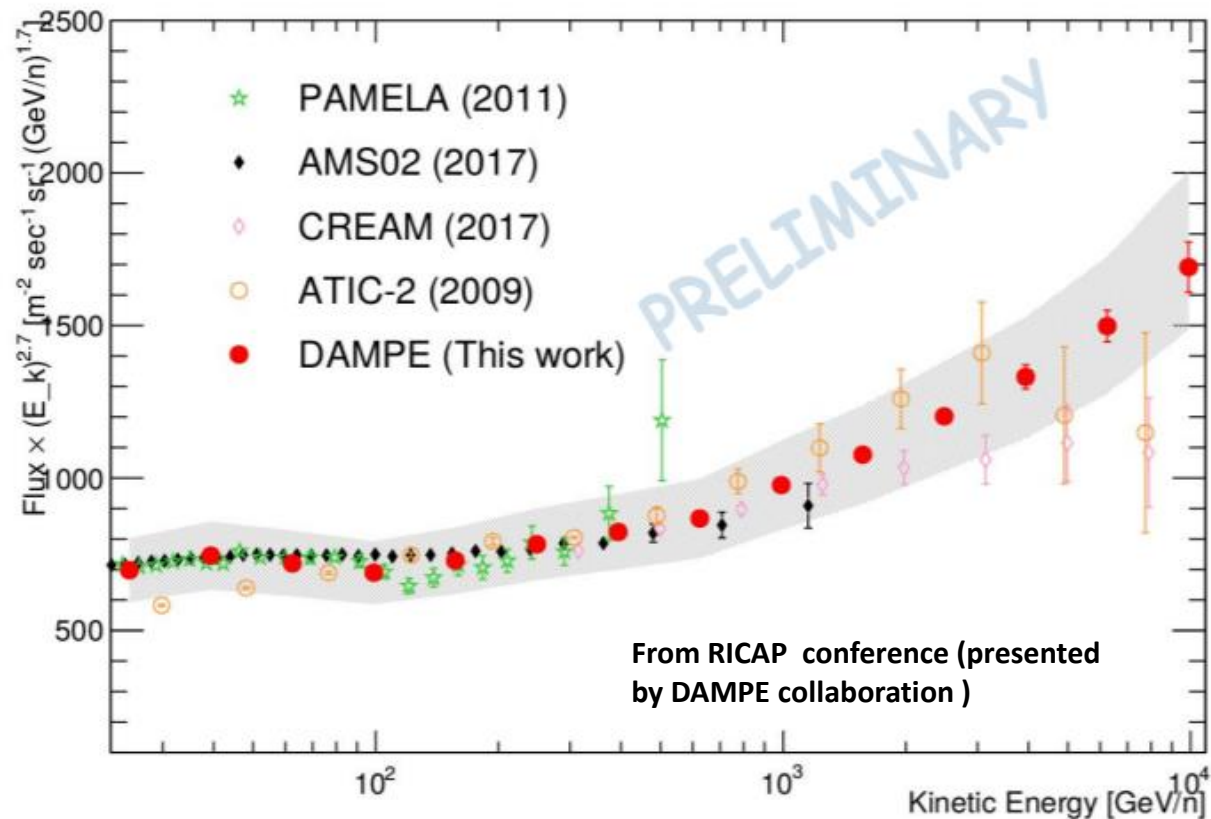
He contamination < 2%

These results:

- Confirm the spectral hardening around 300 GeV observed by ATIC/CREAM/PAMELA/AMS-02/CALET
- **Reveal a spectral softening above ~10TeV**

Helium flux measurement

DAMPE preliminary helium spectrum

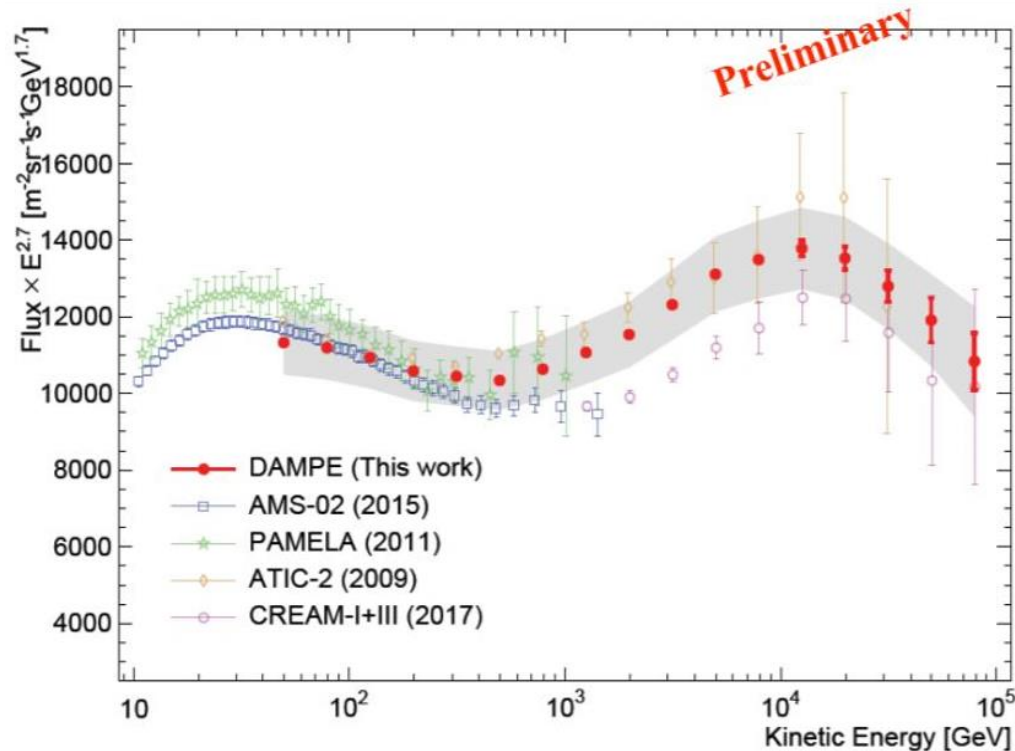


Proton contamination < 1.5%

More precise analysis on helium spectrum is in progress...

p + He spectrum

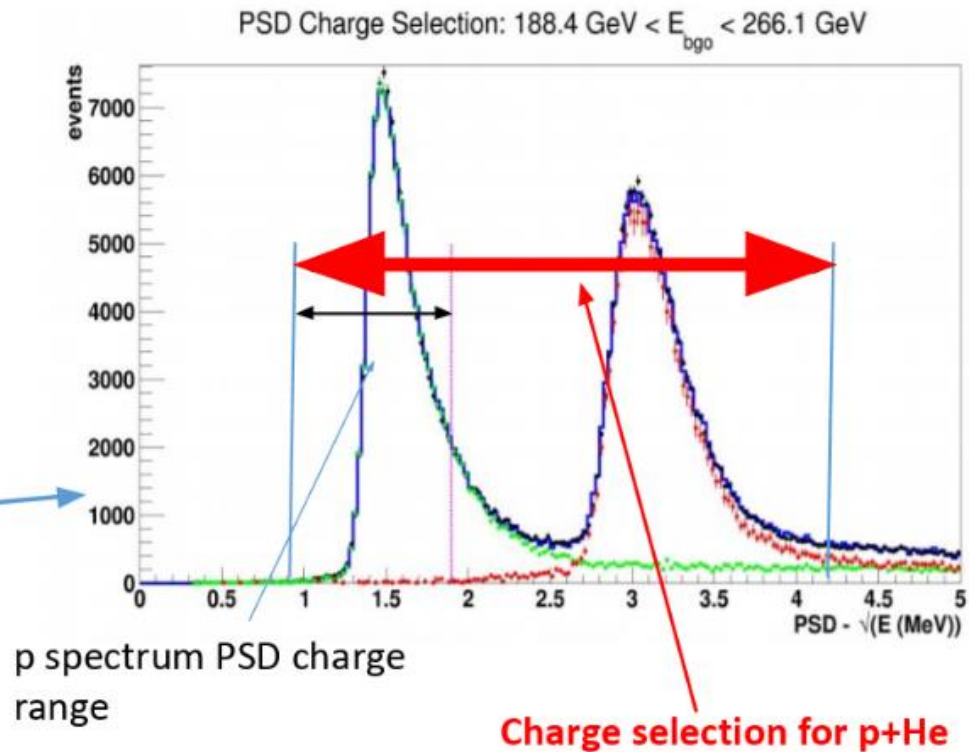
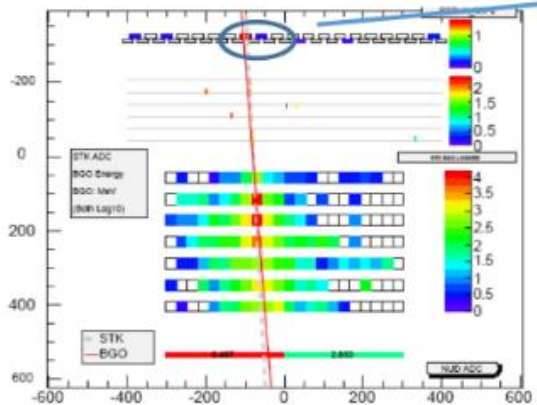
Why do we study the p + He spectrum?



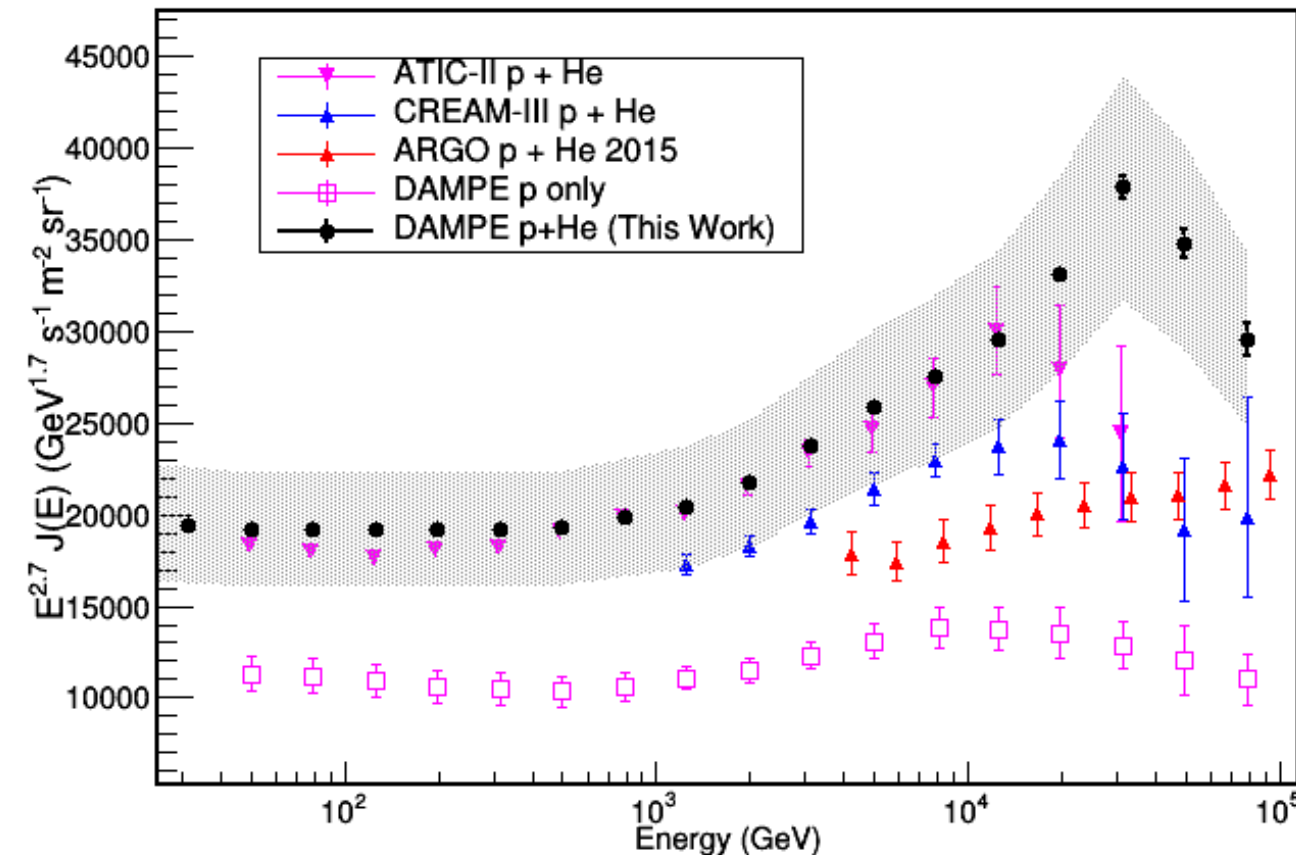
- A spectral softening at 10-15 TeV in proton spectrum?
- Crosscheck for p and He individual spectra
- Negligible background from other nuclei
- No effects of p \leftrightarrow He misidentification
- Compare the p + He spectrum with measurements from ground-based experiments at the highest energies

p + He spectrum

Comparing with the p spectrum analysis, the charge selection range is much “wider” in p + He spectrum analysis, which means larger statistics and less contamination from heavier nuclides.



p + He spectrum

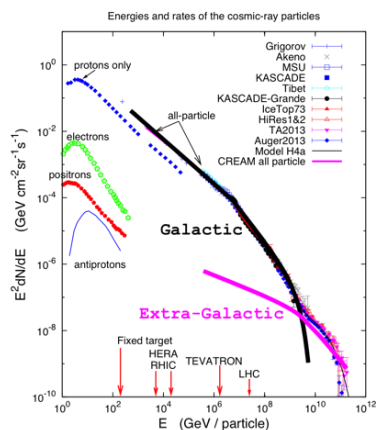


- A softening at ~ 30 TeV!
- Only part of the systematic errors (shadow area) are included.
- Future work on p + He spectrum will be mainly focused on the systematic error assessment.

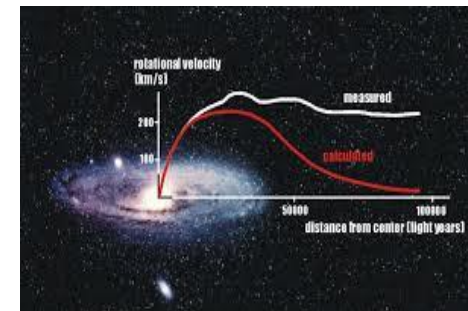
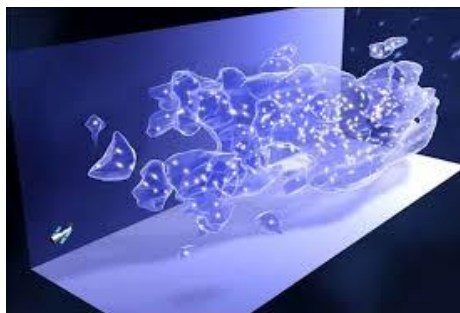
Summary

- DAMPE has been working extremely well for ~ 3.5 years. All the sub-detectors are in stable status.
- The electron + positron spectrum at TeV energies has been shown, together with the preliminary light nuclei spectra.
- Future works for DAMPE will focus on acquiring more precise nuclei spectra

More results are expected in the future.



Thanks for your attention!



BACK UP

CALET Proton spectrum

