Workshop on Chirality, Vorticity and Magnetic Field in Heavy Ion Collisions GGI, Arcetri, Firenze 19-22 Mar 2018



Impact of the vortical pre-equilibrium stage of relativistic heavy ion collision on photon production

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ELECTROMAGNETIC PROBES

Once produced, **photons** do not suffer further interaction with the medium due to their electromagnetic nature ($\alpha \ll \alpha_{\rm S}$)

DIRECT PHOTONS

emerge directly from a particle collision represent less than 10% of all detected photons



Experiments can not distinguish among the different sources

Theoretical models can be used to identify these sources and their relative importance in the spectrum

BOLTZMANN TRANSPORT EQUATION

In order to simulate the temporal evolution of the fireball we solve the **Boltzmann equation** for the parton distribution function **f**(**x**,**p**) $+ gQF^{\mu\nu}p_{\mu}\partial^{p}_{\nu}) f = \mathcal{C}[f]$ $(p_{\mu}\partial^{\mu}$ Free streaming Field interaction **Collision** integral e.g. color-electric field deviations from ideal hydro $\mathcal{C}[f] = \frac{1}{2E_1} \int \frac{d^3 p_2}{(2\pi)^3 2E_2} \frac{1}{\nu} \int \frac{d^3 p_1'}{(2\pi)^3 2E_1'} \frac{d^3 p_2'}{(2\pi)^3 2E_1'} (f_1' f_2' - f_1 f_2)$ $\times |\mathcal{M}_{12 \to 1'2'}| (2\pi)^4 \delta^{(4)} (p_1' + p_2' - p_1 - p_2)$ Δx

> <u>TEST PARTICLES METHOD</u> to map the phase space

STOCHASTIC METHOD to simulate collisions

Follow the entire dynamical fireball evolution within one single theoretical approach

Xu and Greiner, PRC 79 (2009) 014904 Greco et al., PLB 670 (2009) 325 Bratkovskaya, et al., NPA 856 (2011) 162 Ruggieri et al., PRC 89 (2014) 054914 Plumari et al., PRC 92 (2015) 054902

BOLTZMANN TRANSPORT EQUATION

In order to permit photon production we add to the collision integral of the Boltzmann equation processes with a photon in the final state

 $(p_{\mu}\partial^{\mu} + gQF^{\mu\nu}p_{\mu}\partial^{p}_{\nu})f =$

QCD Compton scattering



 $d\sigma^{Compton}$ $\pi\alpha\alpha_s \; u^2 + s^2$ $3s^2$ dtus $d\sigma^{annihil}$ $8\pi\alpha\alpha_s\;u^2+t^2$ $9s^2$ dtut

Quark-antiquark annihilation

 $C[f] = C_{22}[f] + C_{23}[f] + \dots$





MOTIVATIONS Starting time $t_0 = 0^+$ fm/c

SELF-CONSISTENT SOLUTION OF BOLTZMANN AND MAXWELL EQUATIONS

 $(p_{\mu}\partial^{\mu} + gQ_{jc}F^{\mu\nu}p_{\mu}\partial^{p}_{\nu})f_{jc} = p_{0}\frac{\partial}{\partial t}\frac{dN_{jc}}{d^{3}xd^{3}p} + \mathcal{C}[f] \qquad \qquad \frac{dE}{d\tau} = -j_{M} - j_{D}$

Florkowski and Ryblewski, PRD 88 (2013) 034028 Ruggieri, Puglisi, Oliva, Plumari, Scardina and Greco, PRC 92 (2015) 064904



MOTIVATIONS Starting time $t_0 = 0^+$ fm/c

SELF-CONSISTENT SOLUTION OF BOLTZMANN AND MAXWELL EQUATIONS

$$(p_{\mu}\partial^{\mu} + gQ_{jc}F^{\mu\nu}p_{\mu}\partial^{p}_{\nu})f_{jc} = p_{0}\frac{\partial}{\partial t}\frac{dN_{jc}}{d^{3}xd^{3}p} + \mathcal{C}[f] \qquad \qquad \frac{dE}{d\tau} = -j_{M} - j_{D}$$

longitudinal chromo-electric fields decay in gluon pairs and quark-antiquark pairs by SCHWINGER MECHANISM



Florkowski and Ryblewski, PRD 88 (2013) 034028 Ruggieri, Puglisi, Oliva, Plumari, Scardina and Greco, PRC 92 (2015) 064904

fast isotropization for small viscosity



Pre-equilibrium photons During classical field decay



m photons eld decay Thermal QGP photons

During thermal QGP evolution



No net distinction between photons produced in the pre-equilibrium stage and after thermalization



MOTIVATIONS

Starting time

 $t_0 = 0^+ \text{ fm/c}$

Oliva, Ruggieri, Plumari, Scardina, Peng and Greco, PRC 96, 014914 (2017)

Pre-equilibrium photons During classical field decay

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During thermal QGP evolution

No net distinction between photons produced in the pre-equilibrium stage and after thermalization



Oliva, Ruggieri, Plumari, Scardina, Peng and Greco, PRC 96, 014914 (2017)



MOTIVATIONS

Starting time

 $t_0 = 0^+ \text{ fm/c}$

The early stage shines:

pre-equilibrium photons comparable in number with those emitted from the equilibrated QGP during its whole lifetime.

RELATIVISTIC HEAVY ION COLLISIONS





made by Jonah Bernhard

- Initial phase strongly anisotropic and not thermalized
- Intense magnetic field
 eB_y ~ 5-50 m_π²
 - Vorticity due to the large orbital angular momentum J_y ~ 10⁵-10⁶ ħ

WHICH IS ITS IMPACT ON PHOTON OBSERVABLES?

3+1D EXPANSION simulating **RHIC and LHC collisions**

EQUILIBRIUM INITIAL CONDITIONS given by Glauber distribution in x_{T} -space thermal spectrum in p_T -space **Th-Glauber simulations**

RHIC Starting time: $t_0 = 0.6$ fm/c

For 20-40% centrality class

- Multiplicity of about 1700 particles
- Initial eccentricity of about 0.33







Miller et al., ARNPS 57 (2007) 205



... is trasferred \mathcal{X} to the plasma as a shear flow in the longitudinal direction \boldsymbol{Z}

Snellings, Sorge, Voloshin, Wang and Xu, PRL 84, 2803 (2000) Becattini, Piccinini e Rizzo, PRC 77, 024906 (2008) Csernai, Magas and Wang, PRC 87, 034906 (2013) Becattini et al, EPJ C 75, 406 (2015) Deng and Huang, PRC 93, 064907 (2016) Jiang, Lin and Liao, PRC 94, 044910 (2016); PRC 95, 049904 (2017)

Au-Au collisions @ RHIC 200 AGeV - b=7 fm



ESTIMATED ANGULAR MOMENTUM OF THE PLASMA



Longitudinal velocity profile at $|\eta_s| < 1$ and averaged over all y-layers

0.2 (a) (V_{1z}) 0.1 V37 $\langle V_{4z} \rangle$ 0.0 Au+Au, b=10fm \sqrt{s} =200GeV -0.1 $\eta < 1.0$ _0.2 __10 -5 0 5 10 x (fm)

In qualitative agreement with Deng and Huang, PRC 93, 064907 (2016)

Au-Au collisions @ RHIC 200 AGeV - b=7 fm

$$\omega_{y}(\boldsymbol{x},t) = \frac{\partial v_{x}}{\partial z} - \frac{\partial v_{z}}{\partial x} \quad [\text{fm}^{-1}]$$

2

[Lu]

×

-2

-4



y-component of the classical vorticity field in the reaction plane (y=0)

In agreement with Jiang, Lin and Liao, PRC 94, 044910 (2016); PRC 95, 049904 (2017)

VORTICITY FIELD IN THE FIREBALL DECREASES DURING TEMPORAL EVOLUTION



Au-Au collisions @ RHIC 200 AGeV - b=7 fm

$$\overline{\omega_{y}}(\boldsymbol{x},t) = \frac{\int d^{3}x \, w(\boldsymbol{x},t) \omega_{y}(\boldsymbol{x},t)}{\int d^{3}x \, w(\boldsymbol{x},t)}$$

⊳=7fm |ղ|<4

b=7fm |ŋ|<1

b=9fm |n|<4

b=9fm |n|<1

y-component of the vorticity averaged in $|\eta_s| < 1$ and over the full transverse plane with the weighting function *w*





0.12

0.10

0.08

0.06 0.04

0.02

0.00

0

2

Time (fm/c)

 $\omega_{v} > | (fm^{-1})$

 \leq



Au-Au collisions @ RHIC 200 AGeV - b=7 fm

Inclusion of vorticity necessary to reproduce the measured $v_1(\eta)$ of charged particles

QGP DIRECTED FLOW



$$\nu_1 = \left\langle \frac{p_x}{p_T} \right\rangle$$

ARE PHOTONS INFLUENCED BY THE VORTICAL STRUCTURE OF QUARK-GLUON PLASMA?



Experimental data: STAR Collaboration, PRL 101 (2008) 252301

IMPACT OF VORTICITY ON PHOTONS

Au-Au collisions @ RHIC 200 AGeV - b=7 fm



Au-Au collisions @ RHIC 200 AGeV - b=7 fm



Photon directed flow reflects the asymmetry of the system at emission time

PRELIMINARY

CLEAR EFFECT OF VORTICITY ON PHOTON DIRECTED FLOW

CONCLUSIONS...

Fireball evolution described by means **of relativistic kinetic theory** including the effect of a **finite angular momentum**.

Impact of vorticity on photon production from the quarkgluon plasma:

- ✓ no visible effect in momentum spectrum and in elliptic flow of photons,
- ✓ photon directed flow comparable in magnitude to that of charged particles.



...AND OUTLOOKS

✓ Extend the description of vorticity in the early stage before quark-gluon plasma equilibration.

✓ Investigate the effect of the initial electromagnetic fields on photon production.



Thank you for your attention!

Implementing the full photon production rate



NLO processes give important contribution to the rate because of collinear enhancements Arnold, Moore and Yaffe, JHEP 0112 (2001) 009

$$\frac{d\sigma^{Compton}}{dt} = -\Phi(T)\frac{\pi\alpha\alpha_s}{3s^2}\frac{u^2+s^2}{us}$$
$$\frac{d\sigma^{annihil}}{dt} = \Phi(T)\frac{8\pi\alpha\alpha_s}{9s^2}\frac{u^2+t^2}{ut}$$



Bremsstrahlung

We use a model function, $\Phi(T)$, tuned in order to reproduce the AMY rate at a given temperature

Implementing the full photon production rate



> Fair agreement between AMY rate and the one implemented in our collision integral in a broad range of temperature and photon momentum

AMY: Arnold, Moore and Yaffe, JHEP 0112 (2001) 009

We can follow photon production consistently since the very first moments after the collision regardless of the fact that the system is in local equilibrium or not

PHOTONS at RHIC: impact of pre-equilibrium

the lifetime of the early stage is at most one tenth of the full QGP lifetime in the fireball



The EARLY STAGE is QUITE BRIGHT NO DARK AGE in uRHICs

> At RHIC Lifetime of QGP lasts about 5-6 fm/c

In $\sim 1/10$ of its lifetime QGP produces $\sim 1/3$ of the photons it produces during the full evolution

PHOTONS at RHIC: impact of pre-equilibrium



Photon spectrum from QGP is dominated by the early stage photons in the transverse momentum region $p_T > 1.5 \text{ GeV}$

Oliva et al., PRC 96 (2017) 014914