The TRACE spectrometer at SPES

A. Goasduff Dipartimento di Fisica e Astronomia - INFN Padova

SPES one-day Workshop

Probing Fundamental Symmetries and Interactions by low energy excitation with RIBs Feb. 1, 2018





Università degli Studi di Padova



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- High counting rate capabilities
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\Longrightarrow The GHT collaboration

GHT Collaboration Agreement

Introduction

GHT (acronym for GASPARD, HYDE and TRACE, in reference to the corresponding initial projects) is an international collaboration aimed to develop a new detector for optimal study of reactions using low and intermediate energy mew type of compact, highly segmented, silicon arrey, fully integrated within next generation gamma detectors such as AGATA and PARIS. Such new type of Silicon-based array is also meant to offer state-off-the art particle identification to

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- Data acquisition system and integration with other arrays





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Intermediate stage between preams and digitizers: PLAS (R.Aliaga, A.Gadea) Input: Output:

- Different polarity and signal range
- Sampling at 200 MS/s
- Common Trigger Request signal

• 32 inputs with independent trigger R.J. Aliaga et al., NIM A 800 (2015)

- Generates timestamp for pulses
- Synchronizable with each other and/or GTS
- Low noise (11.5 ENOB spec)
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- 200 MS/s digitizers
- Compatible with AGATA/GALILEO GTS system
- Compatible with the new Trigger Processor
- Customization possible at the firmware level



CSIC





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Design and detectors of GASPARD-TRACE arrays:

- Elastic and inelastic scattering
- Transfer reaction with very light targets (d, ^{3,4}He, ...)
- Cluster transfer / Incomplete fusion with weakly bound ions (^{6,7}Li, ⁹Be)
- Multi-nucleon transfer with light ions
- Heavy-, light-nuclei break-up
- Cluster decay
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SPES beams are well suited for direct reaction studies: Energy, intensity, emittance, purity

Direct reaction studies with TRACE



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Lol for r-process studies at SPES using transfer reactions S. D. Pain et al.



Prog. Part. Nucl. Phys. 86 (2016) 86.

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- Transfer reactions can provide input (energies, quantum numbers and spectroscopic factors) to calculations of direct and resonant neutron capture cross sections
- Recent advances in surrogate reaction theory

TABLE 1: Example experiments that could be performed with projected Phase 1 beams from SPES. In each case, data from the (d_2p) , (d_1) and (d_1) the reactions could be acquired simultaneously. In the case of experiments motivated primarily be constraining n-capture cross sections, the (d_2p) reaction of foremost interest, but data on pickup reactions would also be acquired in such a measurement.

| Beam | Projected intensity | Reactions | Primary motivation |
|---------------------------------|---|----------------------------------|----------------------|
| ^{80,81} Ge | 8×10^{4} | (d,t) (d, ³ He) | structure |
| 81 Ge | 1×10^{4} | (d,p) (d,t) (d, ³ He) | n-capture |
| 78,80,81 Ga | 8×10^4 , 1.5×10^4 , 3.5×10^3 | (d,p) (d,t) (d, ³ He) | n-capture |
| ⁸⁴ Se | 7×10^{4} | (d,t) (d, ³ He) | structure |
| ^{129,131} Sn | 8.7×10^6 , 1.7×10^6 | (d,p) (d,t) (d, ³ He) | n-capture |
| ¹³⁰ Sn | 4×10^{6} | (d,t) (d, ³ He) | structure |
| ^{132}Sb | 9×10^{5} | (d,p) (d,t) (d, ³ He) | structure |
| ¹³⁴ Sb | 1.5×10^{4} | (d,p) (d,t) (d, ³ He) | n-capture |
| $^{132,134,136,138}\mathrm{Te}$ | 2×10^7 , 5.8×10^6 , 2.7×10^5 , 1.1×10^4 | (d,p) (d,t) (d, ³ He) | structure, n-capture |
| 137 Xe | 4×10^{4} | (d,p) (d,t) (d, ³ He) | n-capture |
| 138,140,142 Xe | $5.6 \times 10^{6}, 3.4 \times 10^{5}, 1.8 \times 10^{4}$ | $(d,p) (d,t) (d,^{3}He)$ | structure, n-capture |

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- Transfer reactions can provide input (energies, quantum numbers and spectroscopic factors) to calculations of direct and resonant neutron capture cross sections
- Recent advances in surrogate reaction theory
- Final r-process abundances are sensitive to neutron-capture cross sections during freeze-out



Phys. Rev. C 79 (2009) 045809

Nuclear Astrophysics studies with SPES

• β-decay station and associated detectors (see talk of A. Gottardo)

- Measurement of the decay characteristics of nuclei around *A* = 90 relevant to the *r*-process nucleosynthesis [T. Kurtukian-Nieto et al.]
- Lols for measurements at SPES on β-decay properties of nuclei belonging to the *s*-process path [S. Cristallo et al.]
- Study of beta-decay properties of neutron-rich isotopes approaching the r-process path [D. Testov et al.]

Reaction studies with TRACE

- Lols transfer reaction measurements at SPES for r-process nucleosynthesis [S.D. Pain et al.]
- Measurement of astrophysical relevant reactions induced by α , protons and neutrons at the Gamow peak using the Trojan Horse method [M.La Cognata. et al.]
- Direct Reactions at SPES: Shell Evolution and Nuclear Astrophysics around Z ~ 50 and N ~ 82 [D. Mengoni et al.]
- Measurements at SPES of n-capture cross sections on radioactive nuclei interesting for *s*-process nucleosynthesis [O. Trippella et al.]



Mass number

Clustering is a phenomenon existing at all scales:

• Strongly bound cluster: α



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Phys. Rev. C 83 (2011) 054319

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How to study these states:

- Cluster transfer reaction (⁶Li,d), (⁷Li,t), ...
- Breakup into the clusters
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SPES LOI: Search for deformed oblate structures in ⁹⁶Y by γ -spectroscopy and cluster transfer reactions with a ⁹⁵Sr SPES beam. B. Fornal, S. Leoni ...

MUGAST+AGATA @ GANIL + SPIRAL1 beams

The MUGAST array:

- 5 GASPARD trapezoidal detectors backward
- 2 TRACE square DSSSD around 90 degrees
- 4 MUST2 telescopes forward



The physics campaign:

- Shell-evolution far from stability
- Shape-coexistence
- Reaction dynamics
- Nuclear astrophysics
- ...



Possible experimental campaign in 2019

Letters of intent for the MUGAST campaign

• Shell structure evolution & deformation:

- Mapping of neutron orbitals around N = 28
- Oblate driving force in n-deficient nuclei above ⁵⁶Ni
- Shape transition along and across N = 28
- Interplay of single-part and collective structures in ⁴⁶Ca
- Shell evolution toward the island of inversion
- Island of Inversion and shape coexistence in ^{30,31}Mg
- ⁷⁵Kr: Shape coexistence in characterization in light Kr

Neutron-proton pairing:

np-pairing in fp-shell

Astrophysics:

- Breakout from hot CNO to rp-process
- Explosive H-burning in Novae
- s-process ⁷⁹Se(n,γ)
- s-process 60 Fe(n, γ)

• Reaction dynamics:

Space-time characterization of emitting sources in HI collisions

F.Flavigny, O.Sorlin et al. A.Goasduff, D.Mengoni, et al. L.Fortunato, D.Mengoni et al. S.Leoni et al. A.Matta, W.Catford, N.Orr, et al. B.Fernandez-Dominguez et al. A.Matta, W.Catford, N.Orr, et al

M. Assié et al.

C.Diget et al. N. de Séréville, F. Hammache et al. G.de Angelis et al. A.Matta, W.Catford, N.Orr, et al.

G. Verde, et al.

The protoTRACE array:

- 8 segmented TRACE prototypes
- Barrel/cube configuration
- GALILEO digitizers @ 100 MS/s

- Shell-evolution near the stability
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Timeline



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THANK YOU FOR YOUR ATTENTION