

The image shows the Fermi Gamma-ray Space Telescope in space. The satellite is a rectangular box with two large, blue, rectangular solar panel arrays extended outwards. It is positioned in the upper center of the frame. The background is a deep space scene with a bright, glowing nebula or galaxy in the upper right, featuring vibrant reds, oranges, and purples. A bright, white star or distant galaxy is visible in the upper left, with a faint, curved, glowing trail of light extending from it. The overall lighting is dramatic, with the bright celestial objects illuminating the scene.

Fermi

Gamma-ray Space Telescope

The Fermi Large Area Telescope: Highlights from the first year on orbit

S.Rainò

Università and INFN Bari

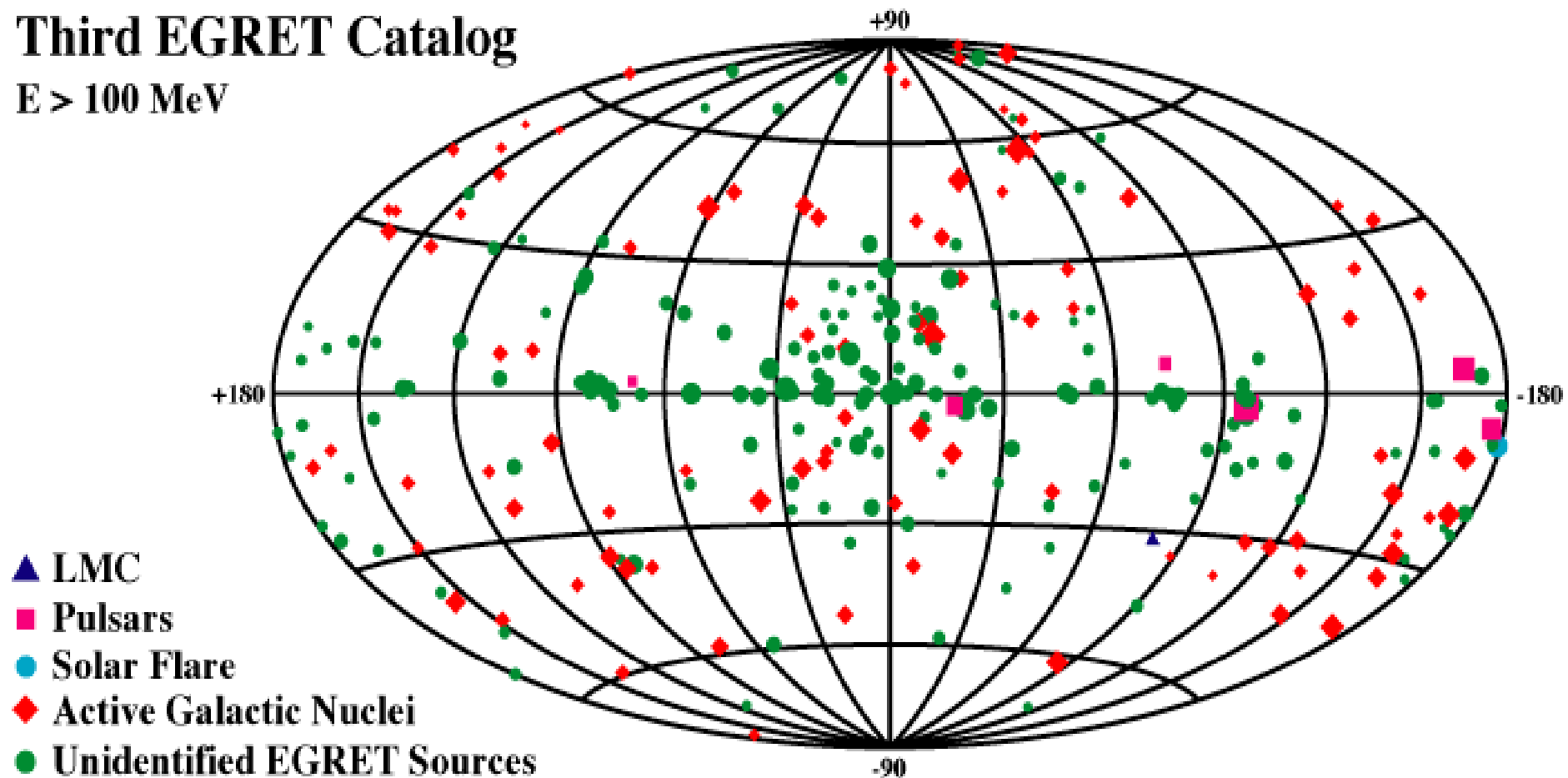
on behalf of the Fermi LAT Collaboration

Before FERMI: the EGRET Era

Data from April 5, 1991 to October 3, 1995

Third EGRET Catalog

$E > 100 \text{ MeV}$



An overview of Fermi

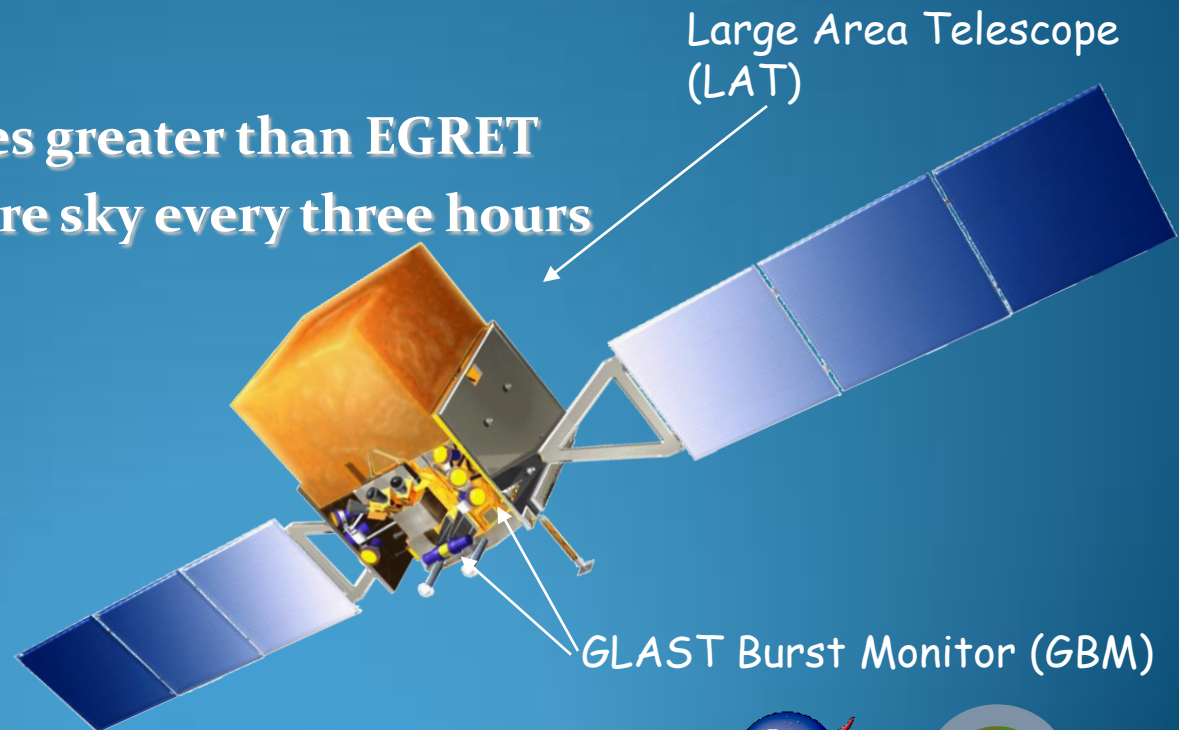
Fermi is an International Science Mission exploring the gamma-ray sky by means of its two main instruments:

- **GLAST Burst Monitor (GBM) : 8 keV to 40 MeV**
- **Large Area Telescope (LAT) : 20 MeV to > 300 GeV**

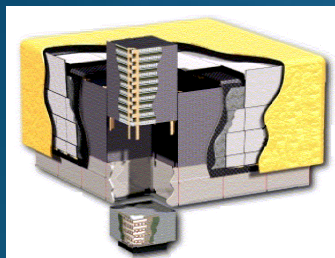
Huge energy range: including largely unexplored band for a total of >7 energy decades!

Strategy:

- **Sensitivity : >10 times greater than EGRET**
- **Survey mode \Rightarrow entire sky every three hours**



LAT: Large Area Telescope

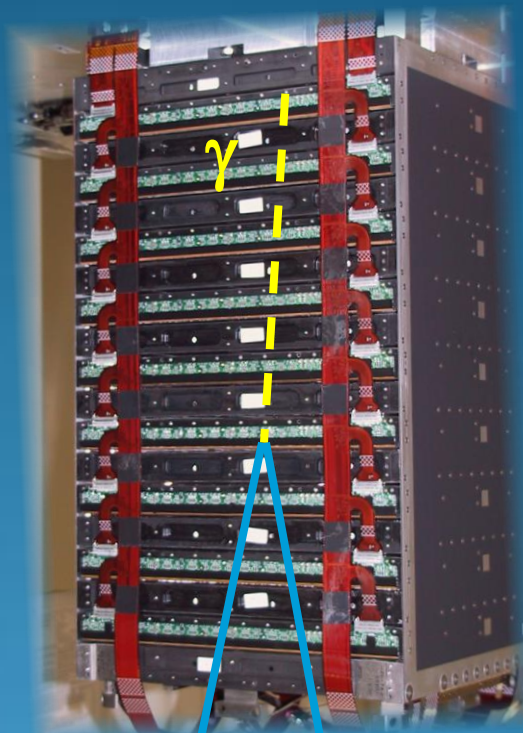
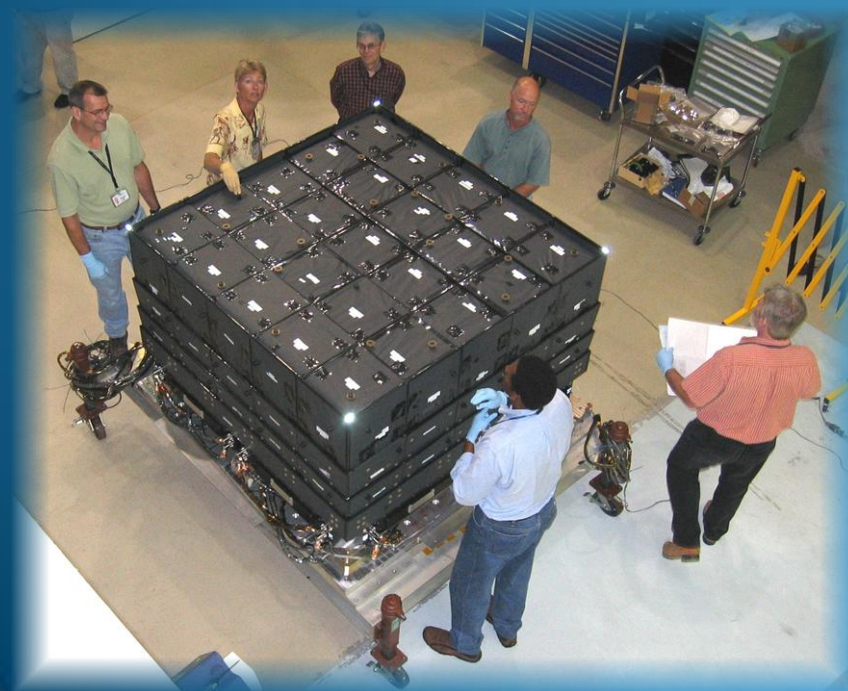


LAT:

- modular - 4x4 array
- 3 tons - 650 watts

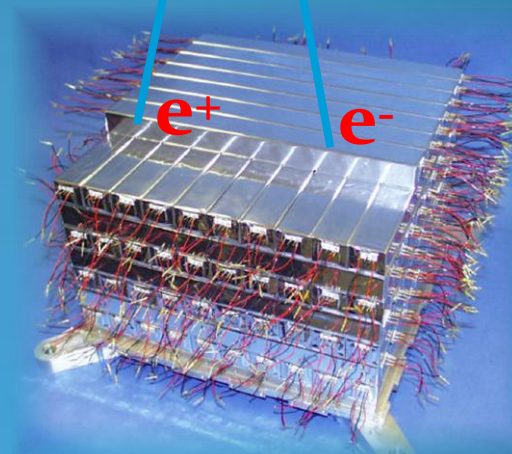
Anti-Coincidence (ACD):

- Segmented (89 tiles + 8 ribbons)
- Self-veto @ high energy limited
- 0.9997 detection efficiency



Tracker/Converter (TKR):

- Si-strip detectors
- ~80 m² of silicon (total)
- W conversion foils
- 1.5 X₀ on-axis
- 18XY planes
- ~10⁶ digital elx chans
- Highly granular
- High precision tracking

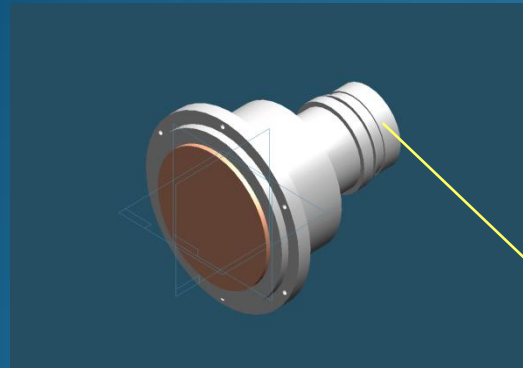


Calorimeter (CAL):

- 1536 CsI(Tl) crystals
- 8.6 X₀ on-axis
- large elx dynamic range (2MeV-60GeV per xtal)
- **Hodoscopic** (8x12)
- Shower profile recon
- leakage correction

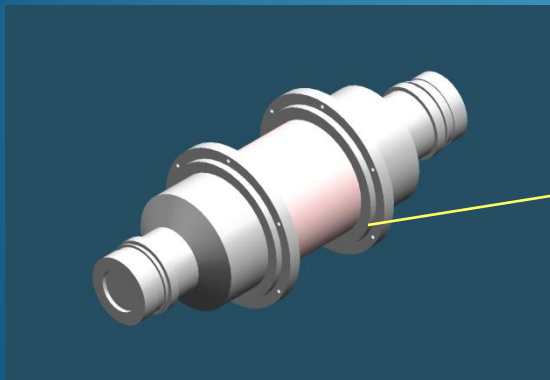
GBM: Gamma-ray Burst Monitor

(12) Sodium Iodide (NaI) Scintillation Detectors

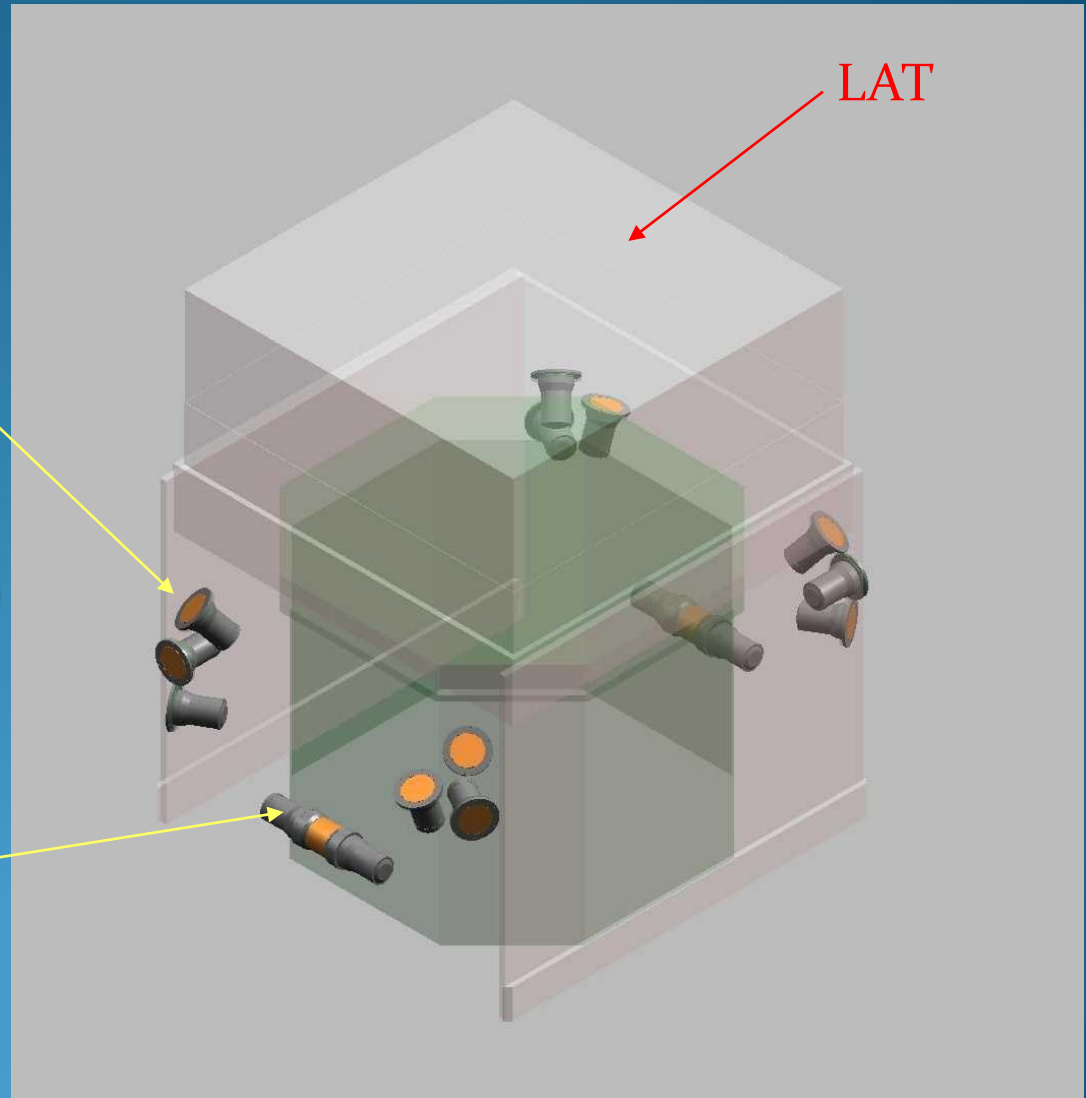


– spectral coverage: 8 keV – 1 MeV

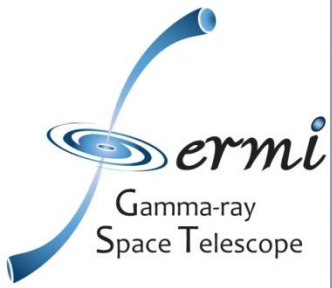
(2) Bismuth Germanate (BGO) Scintillation Detectors



– spectral coverage: 150 keV – 40 MeV



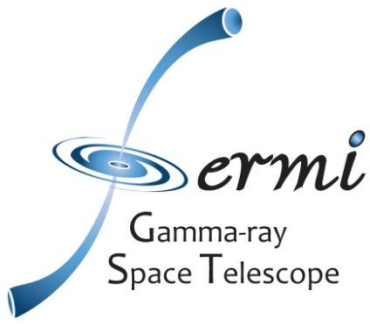




Cape Canaveral

June 11, 2008



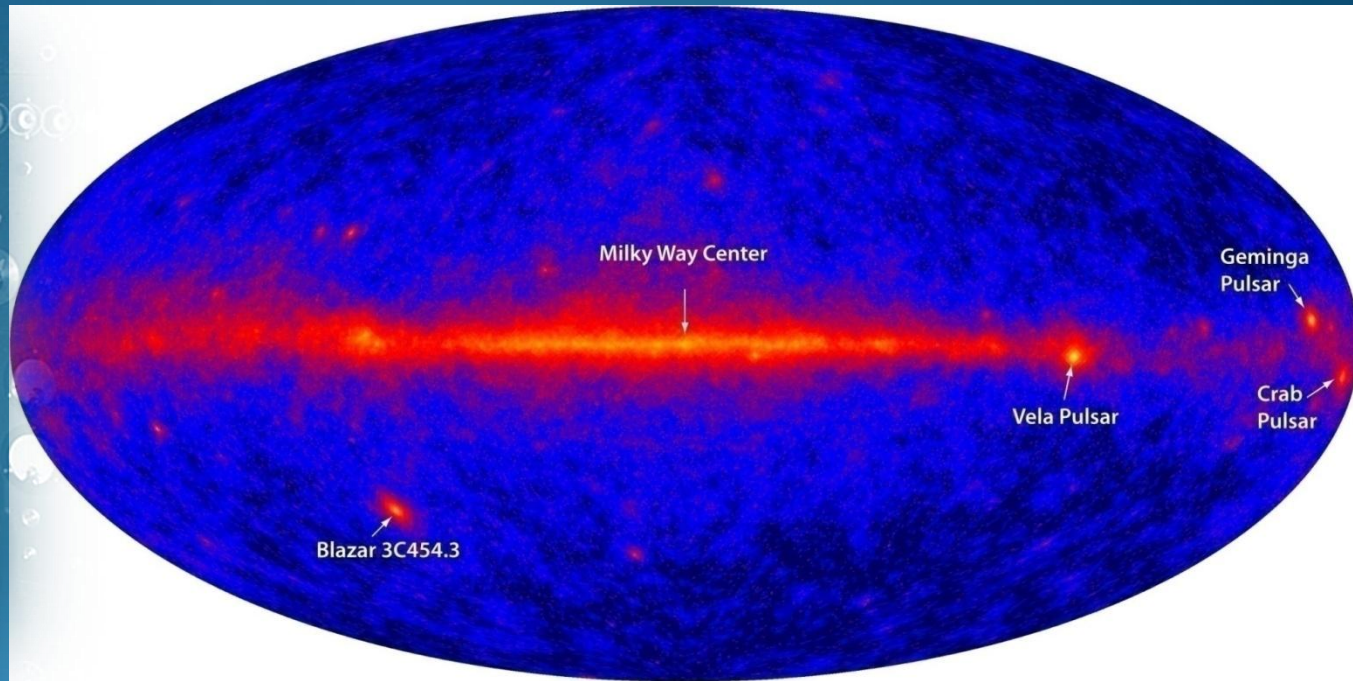


The Fermi Era

- 11 June 2008 at 12:05PM EDT: Launch from Cape Canaveral Air Station
- Circular orbit, 565 km altitude (96 min period), 25.6 deg inclination.
- June 24, 2008 : LAT turn on
- August 11, 2008: Launch & Early Orbit (L&EO) phase ended
- August 26, 2008: Mission name changed to Fermi
- August 25, 2009: Fermi Data public



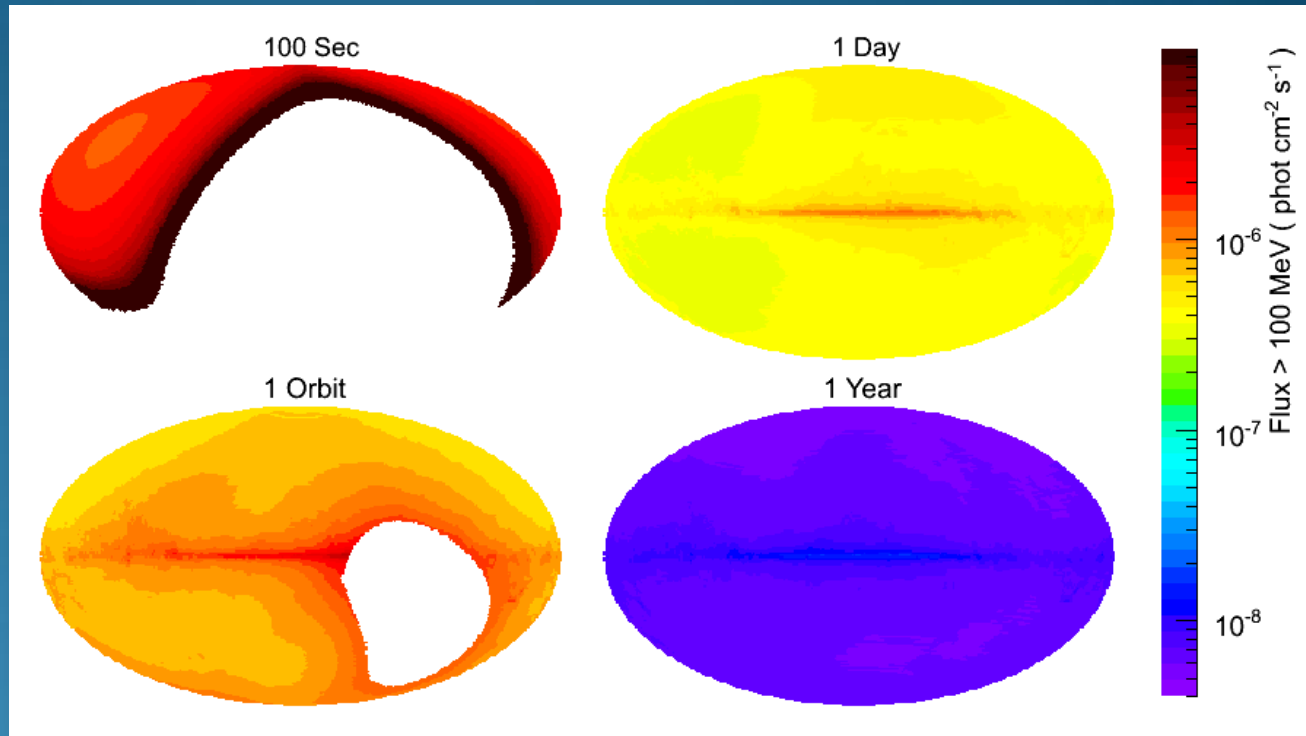
... few weeks after launch: First Light!



The new name, Fermi – Gamma-ray Space Telescope, honors Prof. Enrico Fermi (1901 - 1954), a pioneer in high-energy physics.

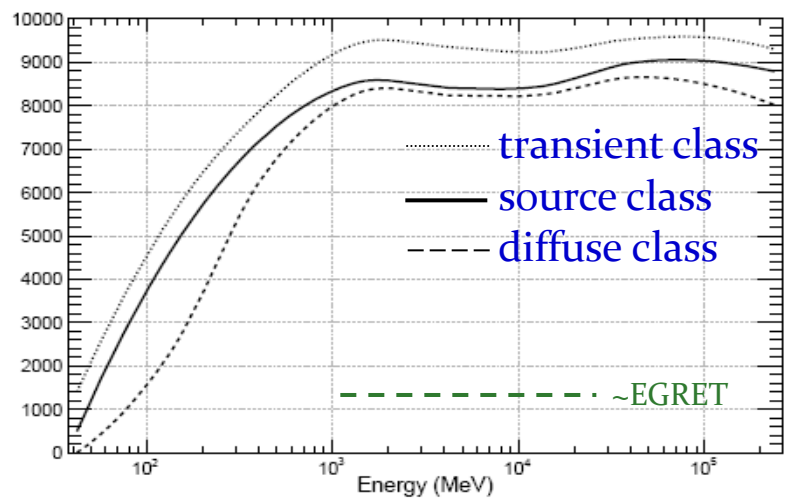
Fermi Operating modes

- Primary observing mode is **Sky Survey**
- Full sky every 2 orbits (3 hours)**
- Uniform exposure, with each region viewed for ~30 min every 2 orbits
- Best serves majority of science, facilitates multi-wavelength observation planning
- EGRET sensitivity reached in days
- Pointed observations when appropriate** (Target of Opportunity pointing)
- Autonomous re-pointing for onboard GRB detections

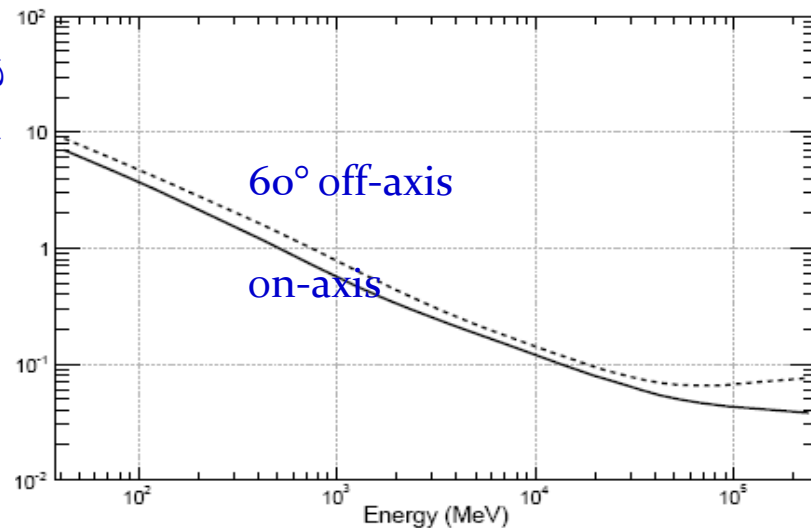


Fermi-LAT Instrument Performance

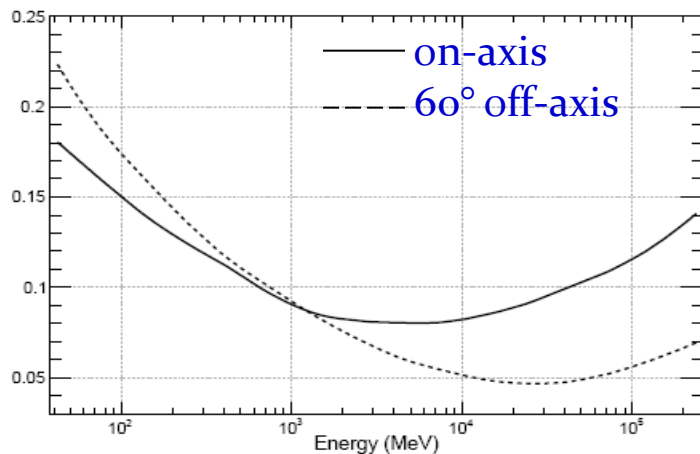
Effective area (cm^2)
Normal incidence



Point Spread Function
68% containment (deg)



Energy dispersion
68% cont



The Large Area Telescope on the Fermi
Gamma-ray Space Telescope

Atwood, W. B. et al. 2009, ApJ, 697, 1071

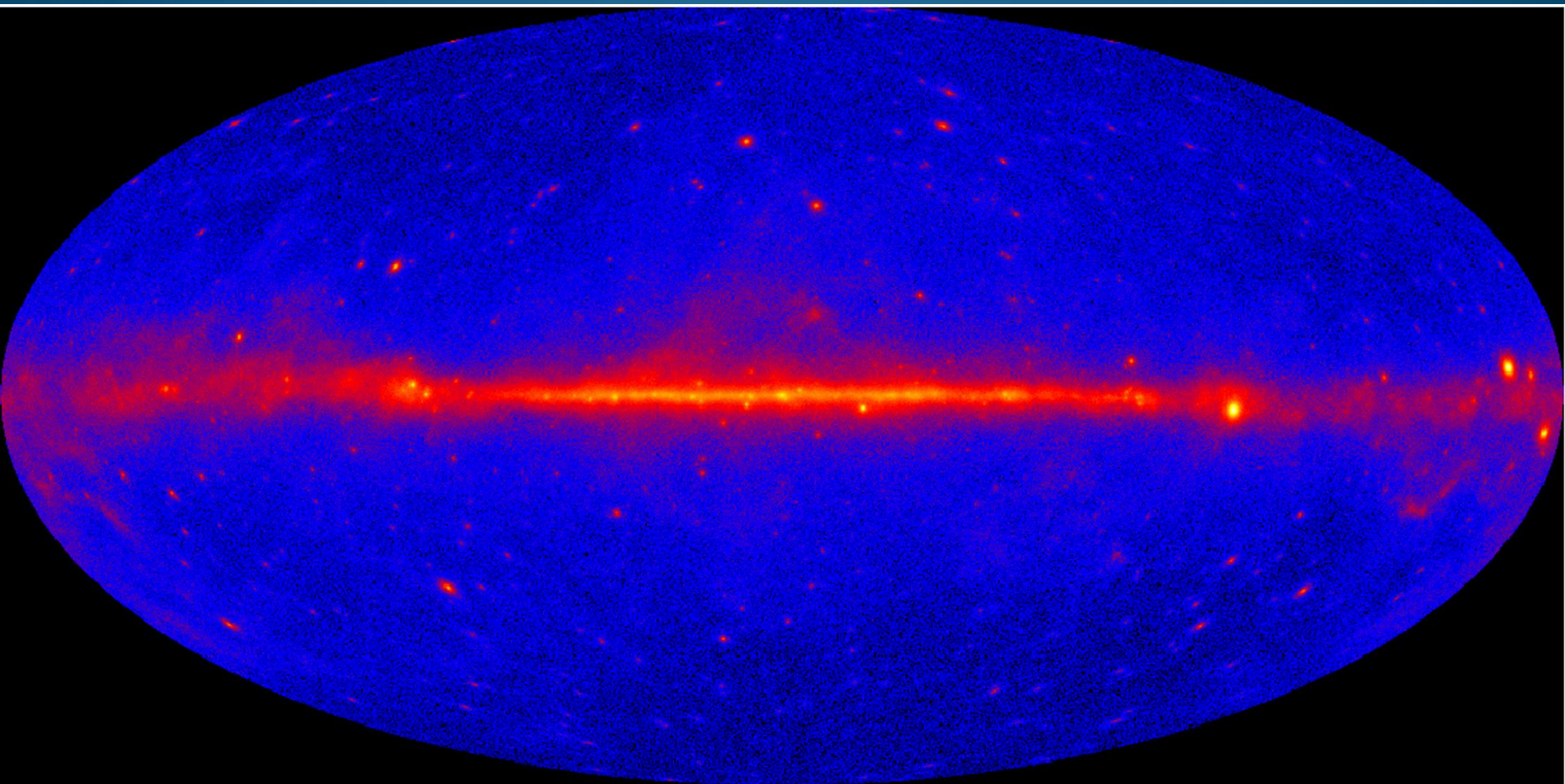
LAT science program

- Active galactic nuclei
- Gamma ray bursts
- Supernova remnants
- Pulsars
- Solar system objects
- Galaxies, clusters of galaxies, X-ray binaries
- Unidentified sources/new populations
- Study of diffuse gamma-ray emission
- Cosmic-ray acceleration & propagation
- Study of Extra-galactic background light (EBL)
- Search for Particle Dark matter/ tests of new physics
- Test Quantum Gravity (?)

Huge increment in capabilities

Draws the interest of both the High Energy Particle Physics and High Energy Astrophysics communities.

1 year Fermi sky

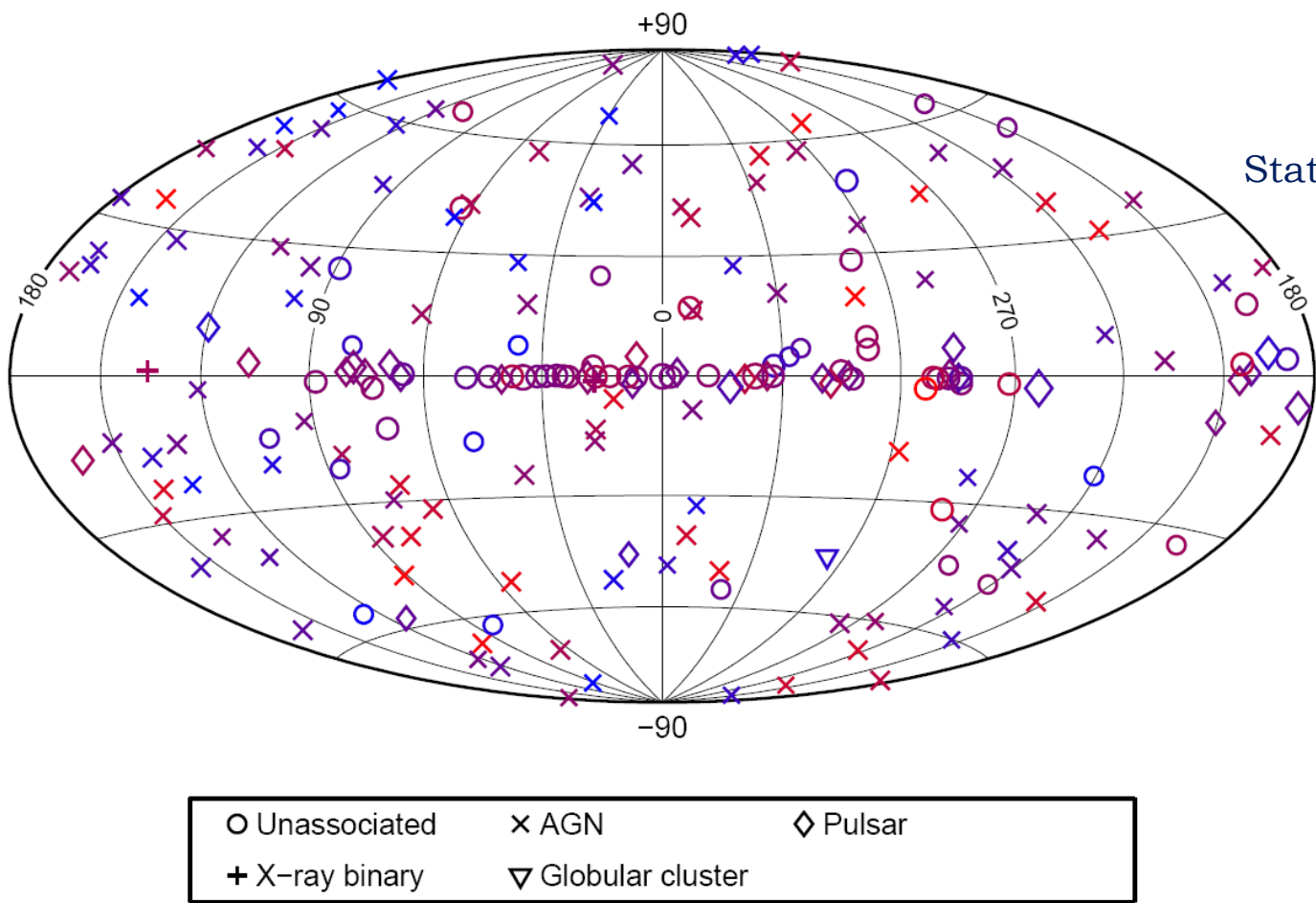


1 year Fermi sky. $E > 200$ MeV in $0.3^\circ/\text{pixel}$



Fermi LAT Bright γ -ray Source List (OFGL)

Fermi Large Area Telescope Bright Gamma-ray Source List
2009, ApJS, 183, 46



3 months LAT data
 $E > 100$ MeV
205 detections with
Statistical significance $> 10 \sigma$

γ -ray bursts in LAT + GBM

So far in the LAT: 12 GRBs, from 8 keV to tens of GeV

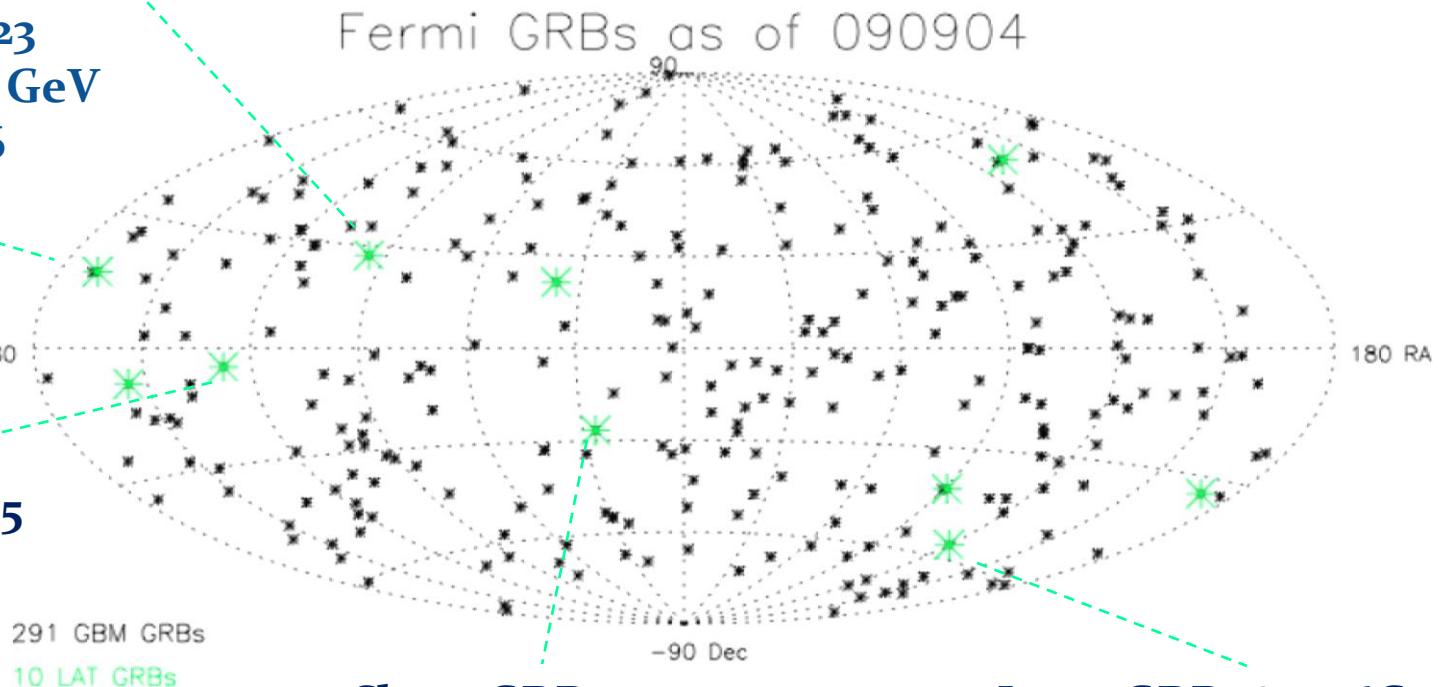
In **1 year** from GBM turn on: 252 GRBs, 138 in the LAT FoV

(see E.Moretti and E.Bissaldi talks)

GRB090902B

Long GRB090323
(200 s) radio to GeV
afterglow, $z=3.6$

Long GRB080825
afterglow



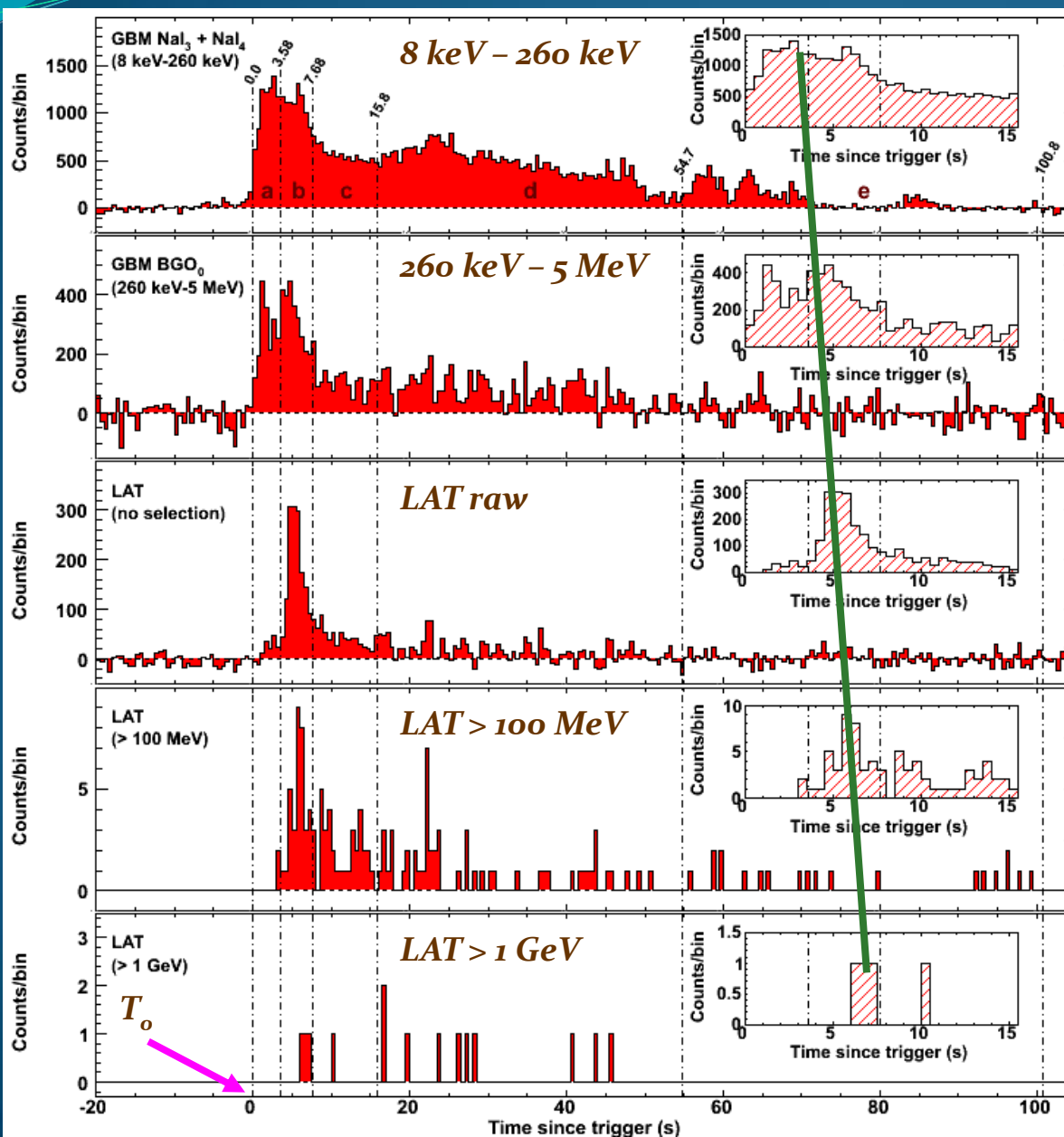
Short GRB090510
intense, $z=0.9$

Long GRB080916C
intense, $z=4.35$, to 13 GeV

Example of a long burst: GRB 080916C

- First high-energy GRB (>100 MeV) with known redshift
- Largest sample >100 MeV
 - 14 events >1 GeV
 - High energy photon (E = 13.2 GeV after 16.5 s) from GRB

*Abdo et al.,
Science 323, 1688 (2009)*



About GRBs :

⦿ Talks:

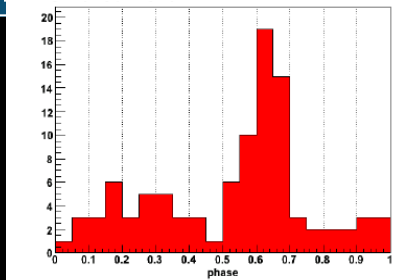
- ◆ The Fermi observation of gamma-ray bursts - E. Moretti
- ◆ GRB090510 – F. de Palma
- ◆ The Fermi Gamma-Ray Burst Monitor: results from the first year – E. Bissaldi
- ◆ Fermi and other Gamma-ray tests for Lorentz Invariance Violation and Quantum Gravity Models - F. Stecker

⦿ Posters:

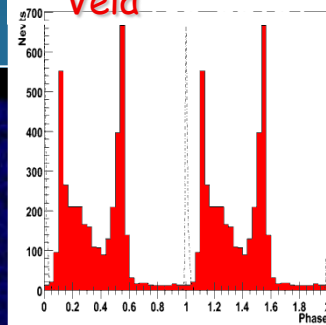
- ◆ Fermi high energy observation of GRB090217 - S. Cutini

Where we started ...EGRET pulsars

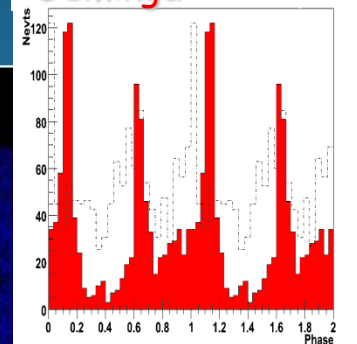
PSR B1951+32



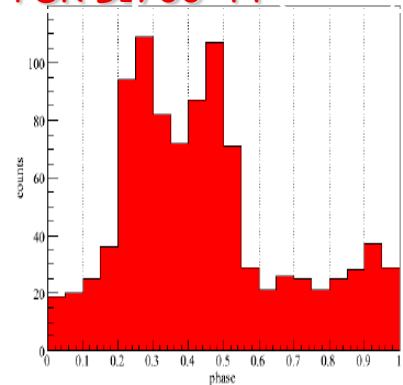
Vela



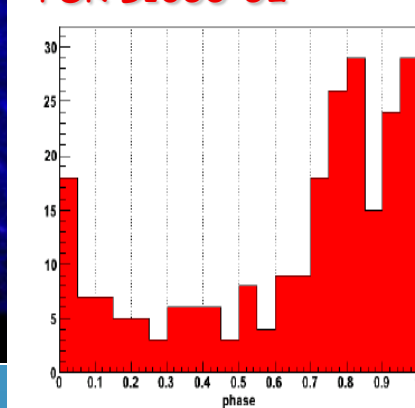
Geminga



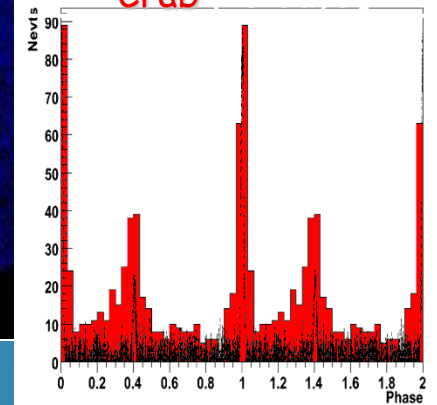
PSR B1706-44



PSR B1055-52

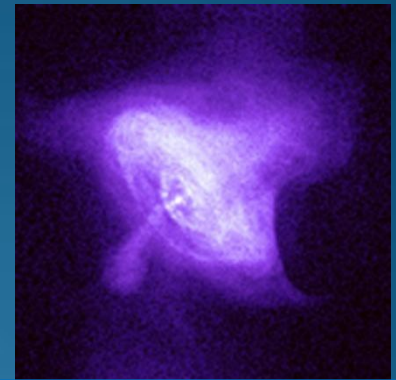


Crab



Galactic sources

- 6 EGRET pulsars → 46 Fermi pulsars in the first Fermi pulsar catalog with 6 months data(submitted to ApJ)
 - ◊ 16 new young, turbulent, radio loud pulsars
 - ◊ 8 old, radio loud, millisecond pulsars
 - ◊ 16 young pulsars, pulsing in gamma-rays alone
- Globular clusters
- Pulsar Wind Nebula
- Supernova remnants
- X-ray binaries
- So far, EGRET unidentified are pulsars.



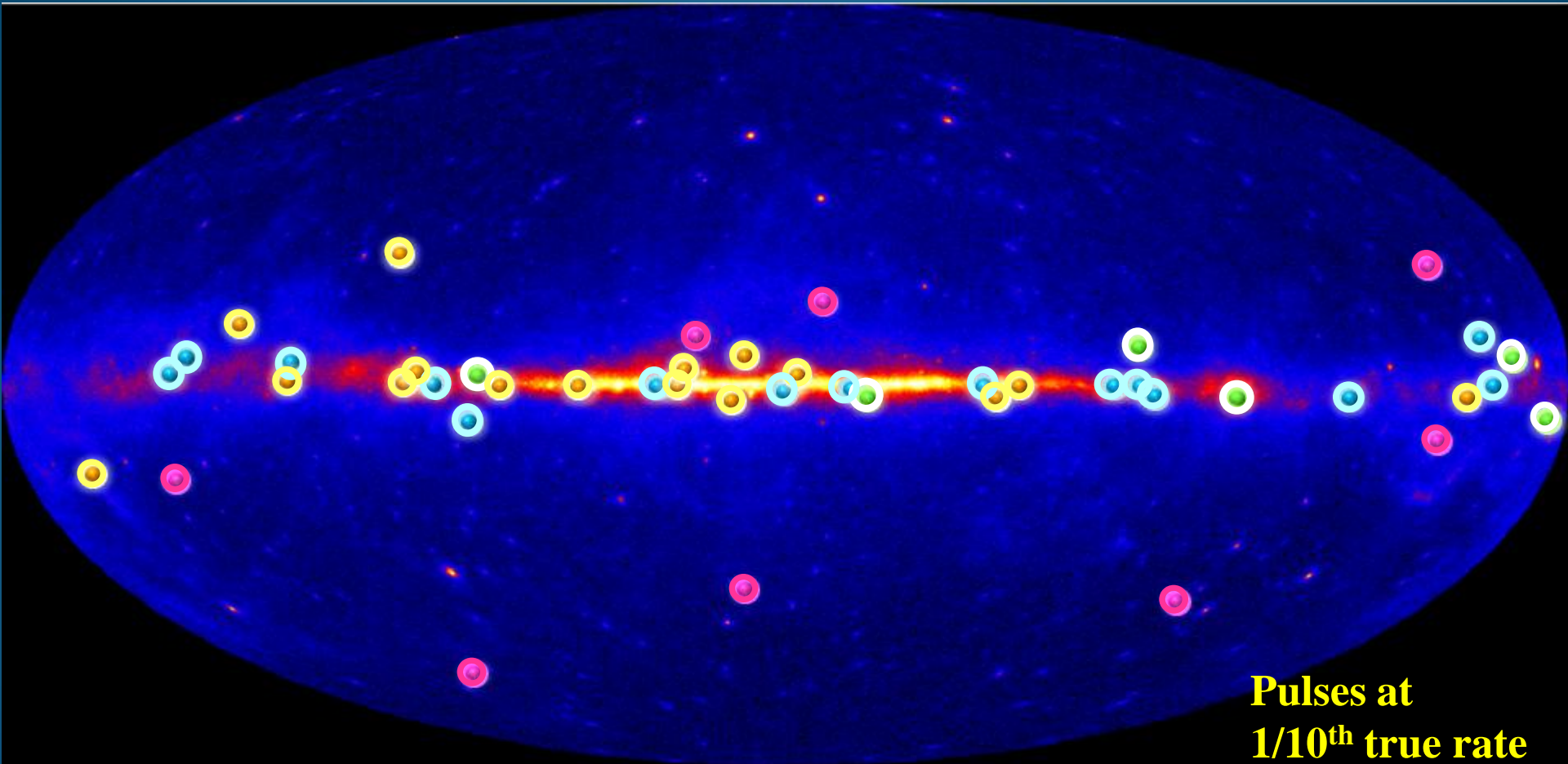
« Pulsar Wind Nebula = PWN »
Chandra X-ray image
Pulsar in the middle

Crab supernova remnant
(optical, Hubble)
Supernova seen in 1054



(see F. Giordano talk)

The “pulsating” sky



**Pulses at
1/10th true rate**

Fermi Pulsar Detections

- New pulsars discovered in a blind search
- Millisecond radio pulsars
- Young radio pulsars
- Pulsars seen by Compton Observatory EGRET instrument

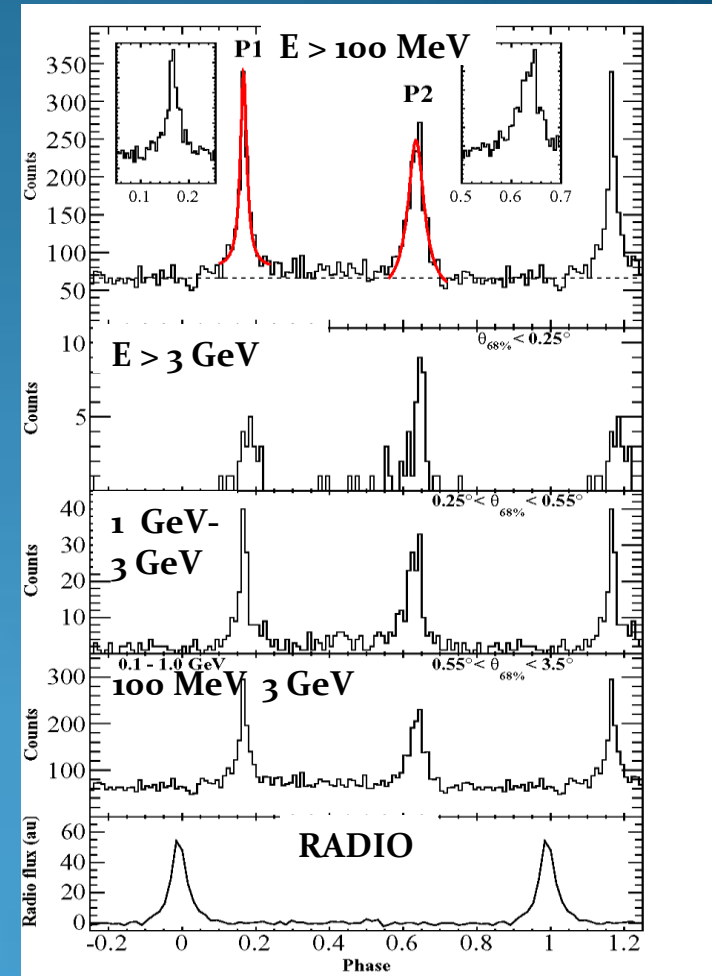
About LAT pulsars

- Generally (but not always), pulse profiles have 2 peaks, separated by ≥ 0.2 of rotational phase.
- Generally (but not always), gamma peak offset from radio.

Pulsed gamma-rays from PSR J2021+3651 with the Fermi Large Area Telescope
Abdo et al. ApJ700, 1059 (2009)

Peak separation = 0.468 ± 0.002

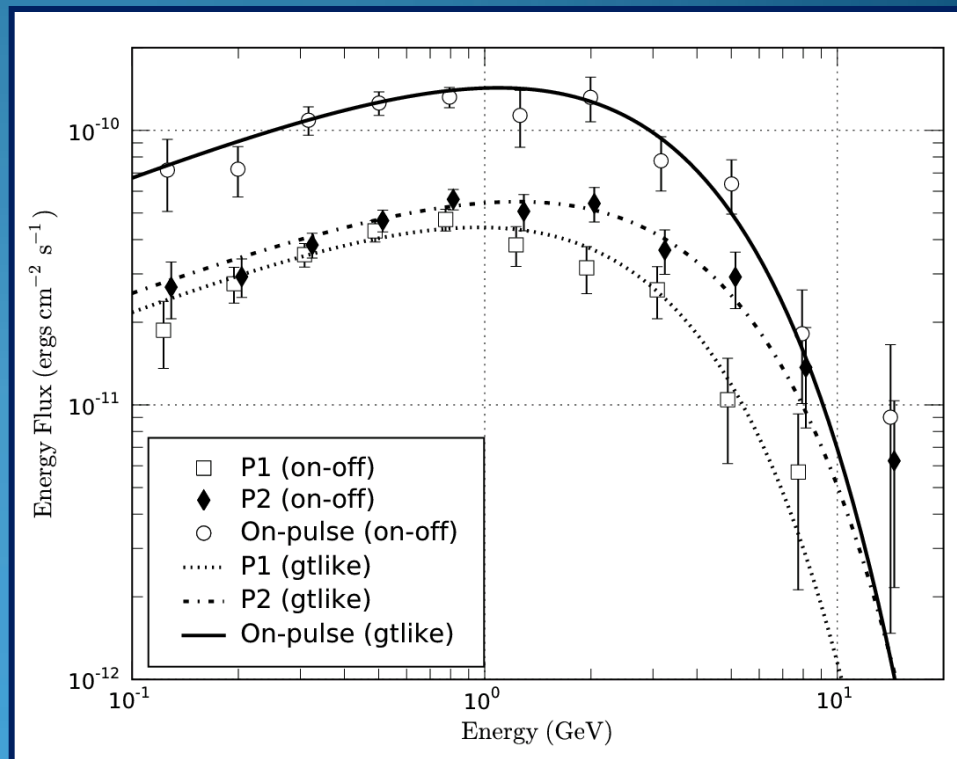
Radio lag = 0.162 ± 0.004



About LAT pulsars

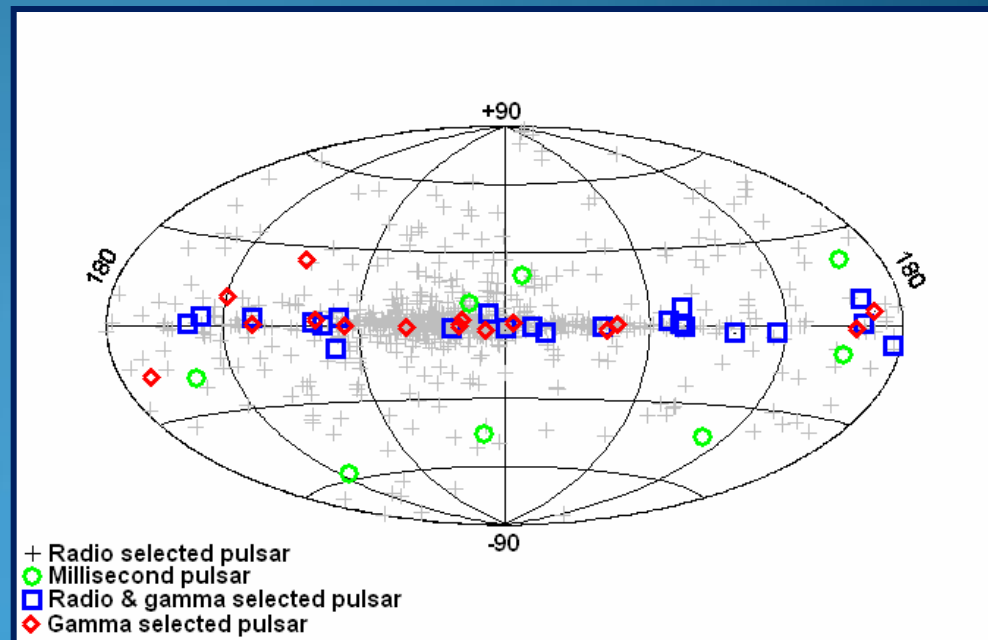
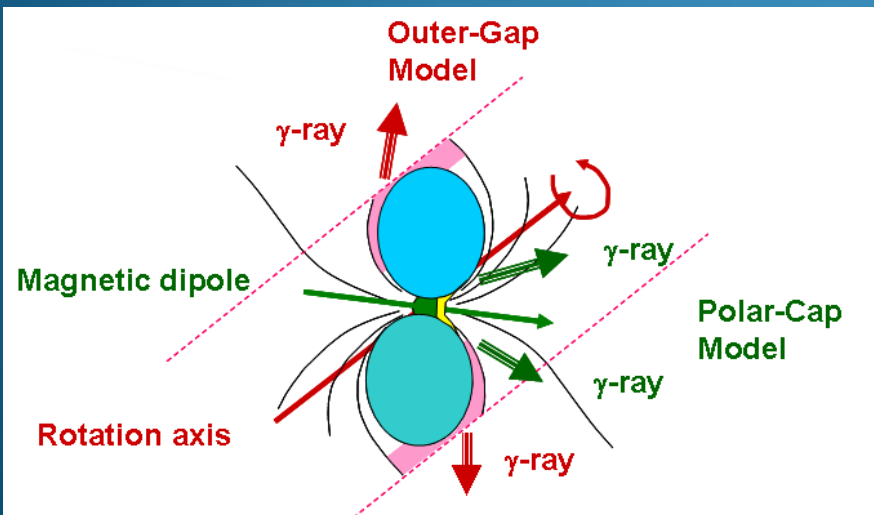
- Generally (but not always), pulse profiles have 2 peaks, separated by ≥ 0.2 of rotational phase.
- Generally (but not always), gamma peak offset from radio.
- Energy spectra well described by a power-law with exponential cut-offs, with cut-off energy at ~ 1 to ~ 5 GeV.

*Pulsed gamma-rays from PSR J2021+3651
with the Fermi Large Area Telescope
Abdo et al. ApJ700, 1059 (2009)*



About LAT pulsars

- Generally (but not always), pulse profiles have 2 peaks, separated by ≥ 0.2 of rotational phase.
- Generally (but not always), gamma peak offset from radio.
- Energy spectra well described by a power-law with exponential cut-offs, with cut-off energy at ~ 1 to ~ 5 GeV.
- Favors outer magnetospheric emission.

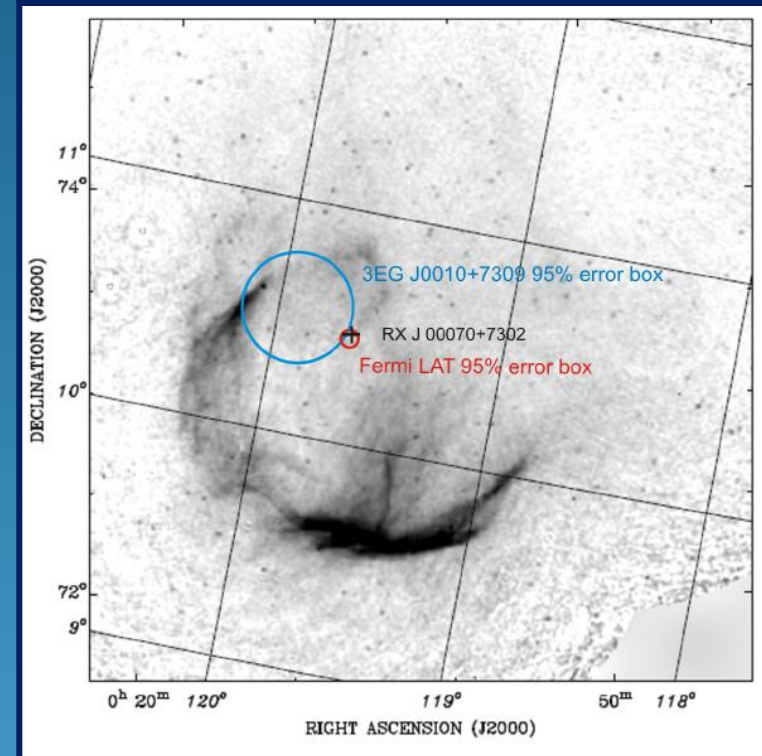
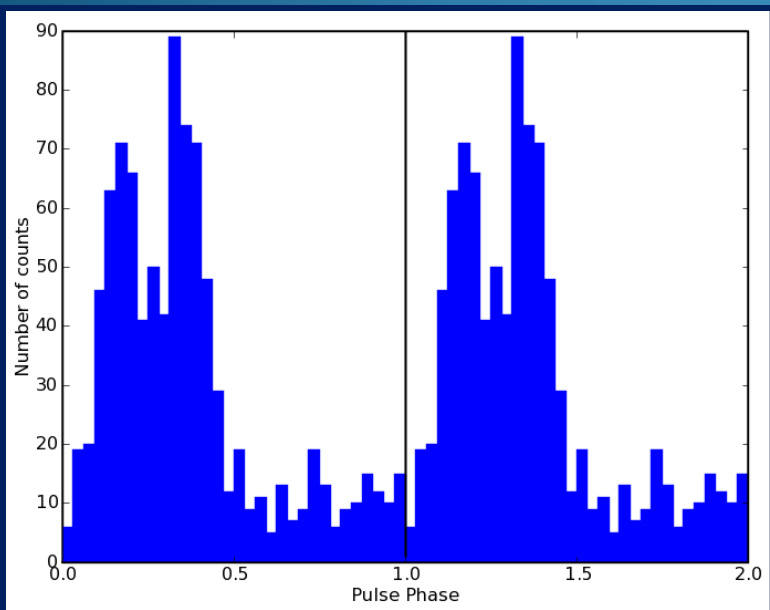


The CTA 1 discovery: the first gamma-ray only pulsar

Discovery of γ pulsation: July 10-11 2008

*"The Fermi Gamma Ray Space Telescope discovers the pulsar in the young Galactic Supernova Remnant CTA1
Science 21 November 2008: Vol. 322. no. 5905, pp. 1218 - 1221*

- ◇ $l = 119.652$, $b = 10.468$;
- ◇ Period ~ 316 ms, $\dot{P} \sim 3.6^\circ\text{E-13}$
- ◇ $\Delta\Phi = 0.2$
- ◇ Distance ~ 1.4 kpc



About Galactic sources:

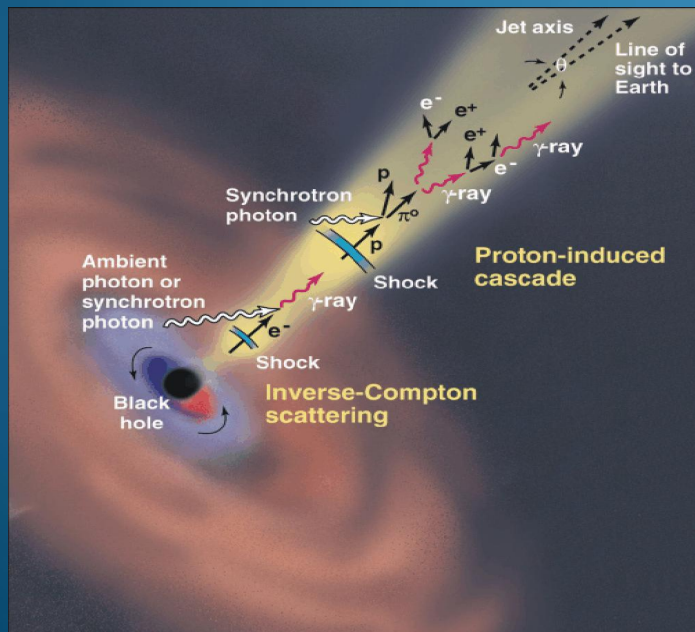
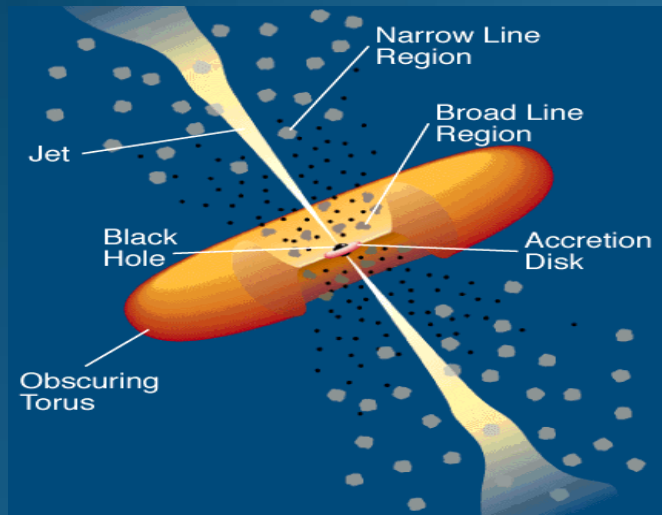
⊙ Talks:

- ◆ High energy Pulsars detection with Fermi LAT - F.Giordano

⊙ Posters:

- ◆ Observation of the Crab pulsar and Nebula with the Fermi Large Area Telescope - M.H.Grondin, M. Lemoine-Goumard, F.Loparco, M.N.Mazziotta

Active Galactic Nuclei (AGN)



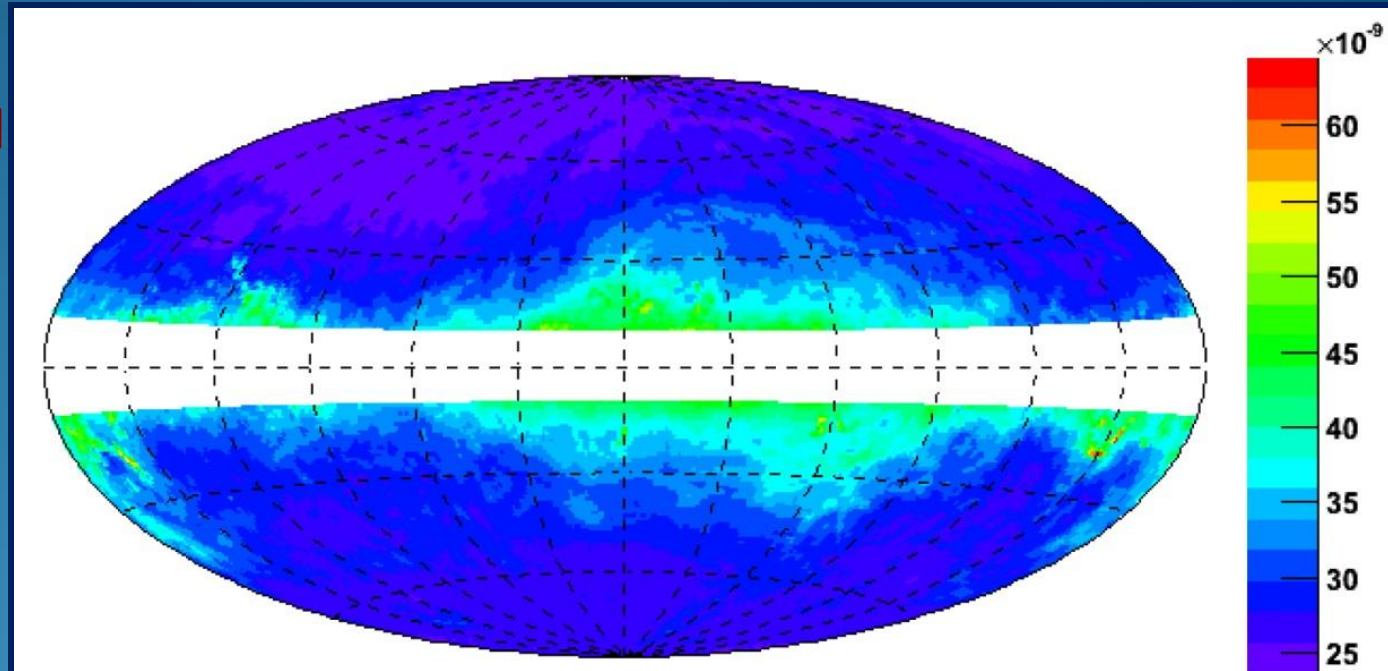
- Active galactic nuclei (AGN) are galaxies with extraordinarily luminous cores powered by super massive black holes
- In the standard model of AGN, cold material close to the central black hole forms an accretion disc
- At least some accretion discs produce jets, twin highly collimated and fast outflows that emerge in opposite directions from close to the disc
- Blazars are objects emitting non-thermal radiation across the entire electromagnetic spectrum from a relativistic jet that is viewed closely along the line of sight

(see P.Giommi talk)

The first Fermi LAT AGNs Source List

**Bright AGN Source List from the First Three Months of the
Fermi Large Area Telescope All-Sky Survey**
ApJ 700, 597 (2009)

3-months flux [$\text{ph cm}^{-2} \text{s}^{-1}$]



132 sources with $|b| > 10 + \text{TS} > 100$

7 pulsars

11 unidentified

114 associated with AGNs

LAT Bright AGN Sample (LBAS)

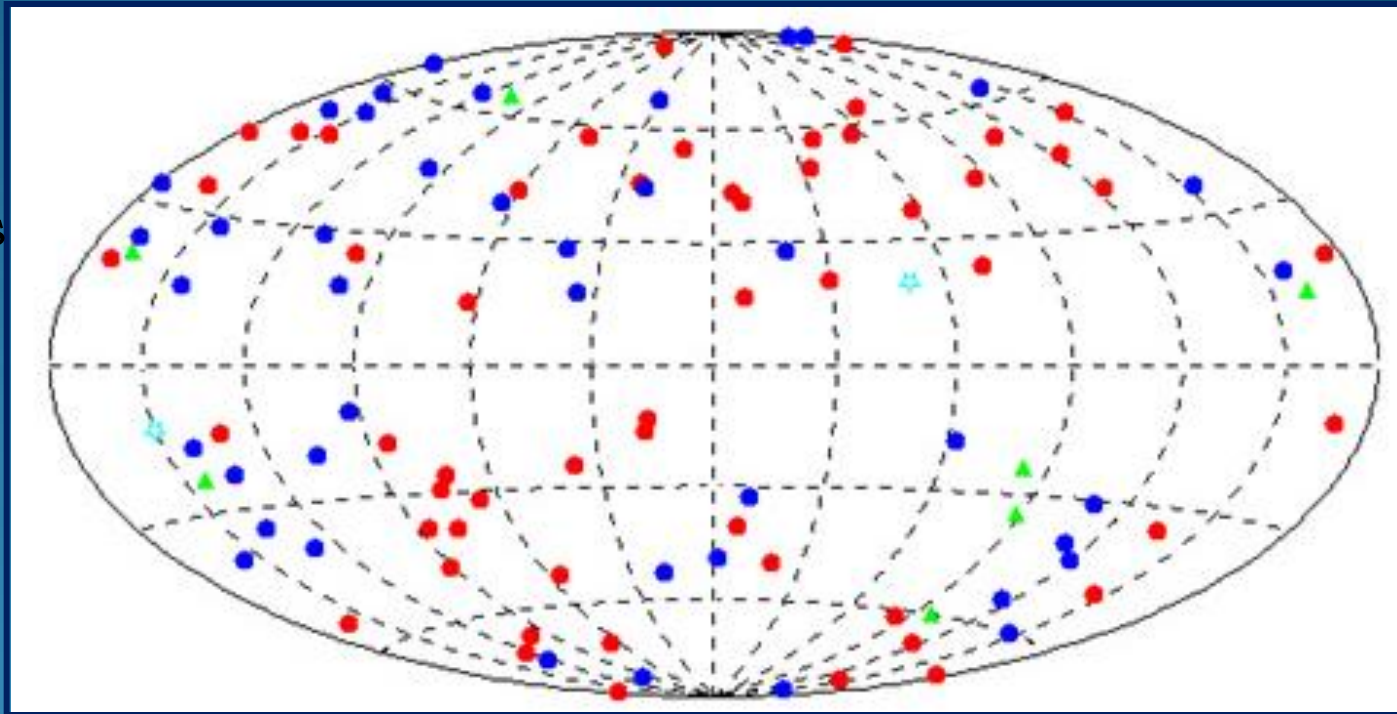
High confidence
associations for
106 AGN (LBAS):

→ 2 Radio Galaxies

- Centaurus A
- NGC 1275

→ 104 Blazars:

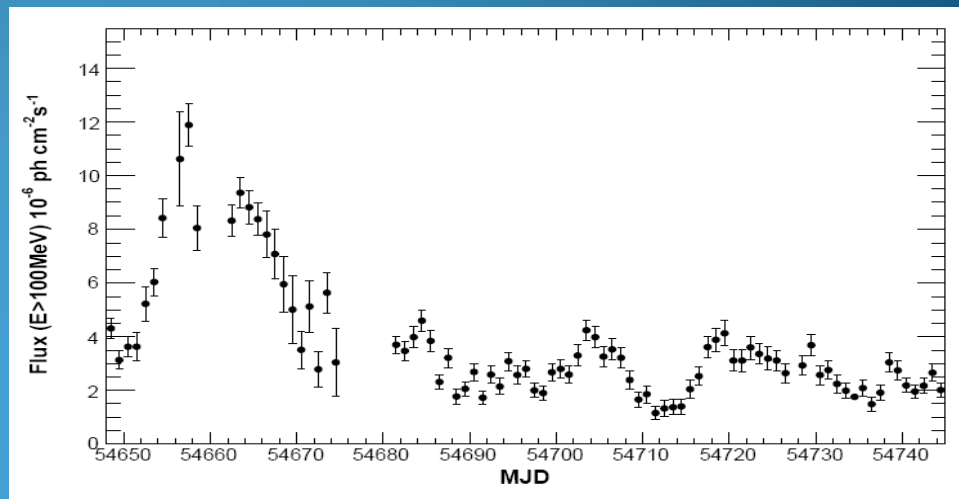
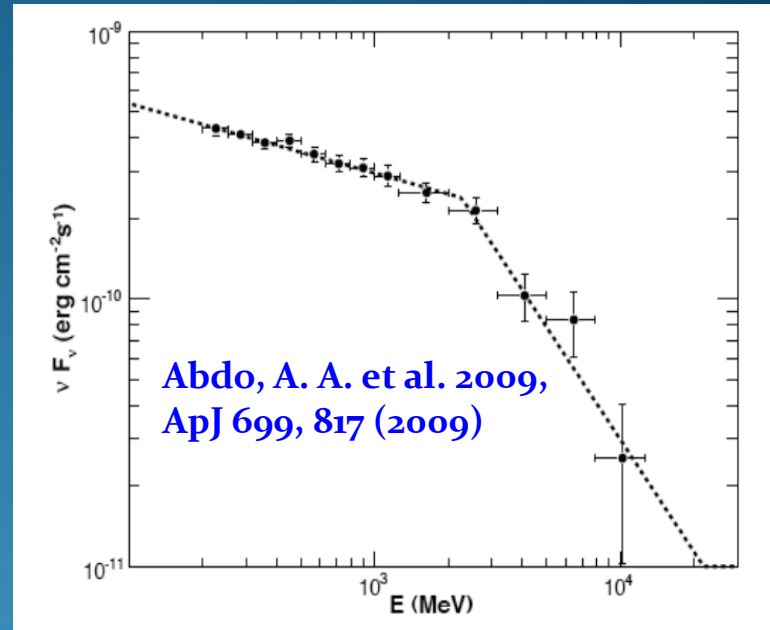
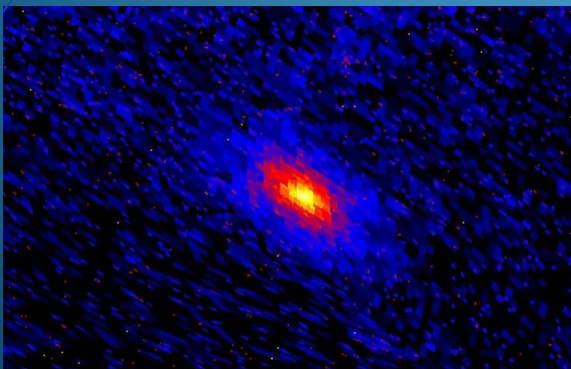
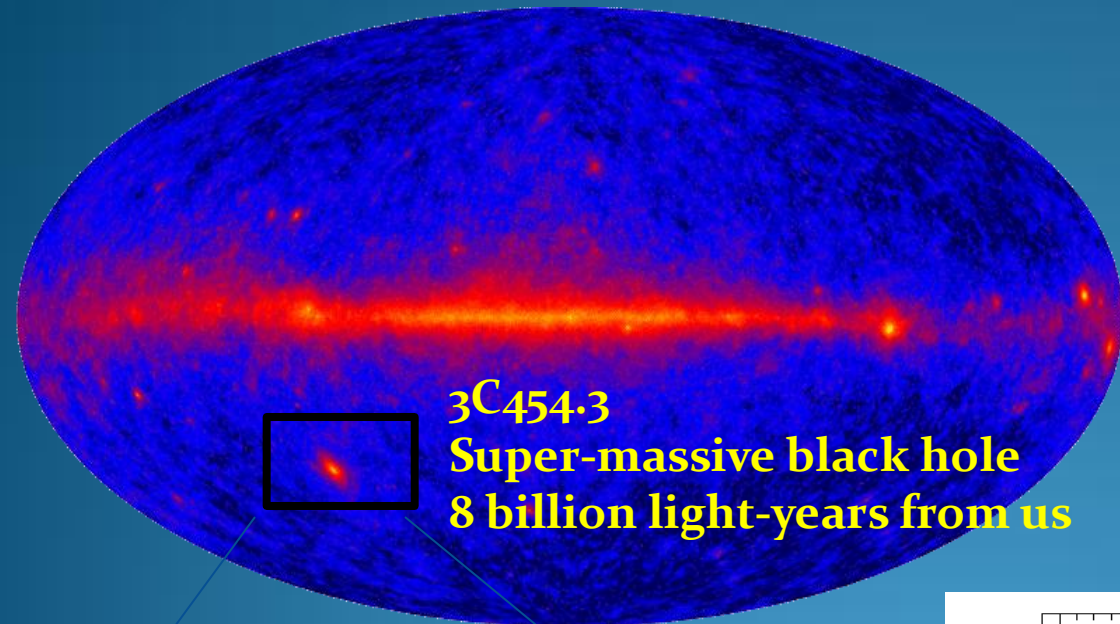
- 58 FSRQ
- 42 BLLac
- 4 Unknown



4 NEW BLAZARS DISCOVERED on the basis of the LAT detections
(present in the CRATES catalog)

3C454.3

The brightest gamma-ray extra-galactic source observed in the first 3 months Fermi-LAT survey

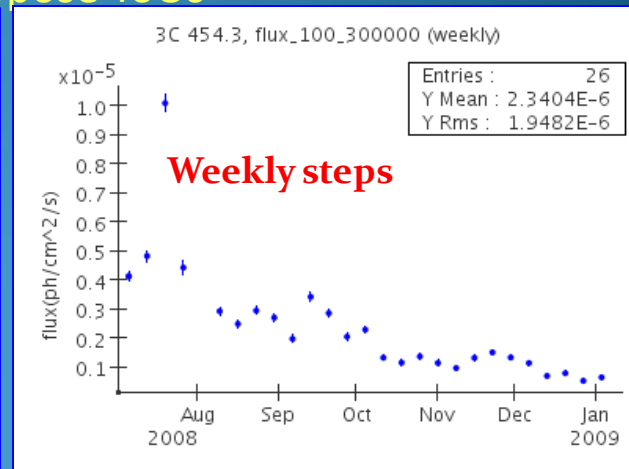
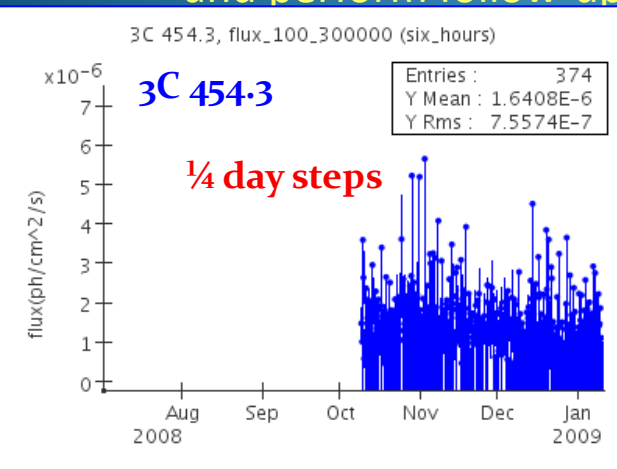


The flaring and variable sky

~50 Astronomers Telegrams: Discovery of new gamma-ray blazars, flares from known gamma-ray blazars, galactic plane transients

Source Monitoring activities

- Automated Science Processing (ASP)
 - Transient detection: Find all point sources in data from each epoch (6hr, day, week).
 - Follow-up monitoring: Run full likelihood analysis on detected sources, and for the “Data Release Plan” (DRP) sources
 - 1×10^{-6} ph cm $^{-2}$ s $^{-1}$ threshold (daily) for public release of non-DRP
- Flare Advocates:
 - LAT scientists from Galactic and Extragalactic groups examine ASP pipeline output and perform follow-up analyses, produce ATels, and propose ToOs



FERMI GAMMA-RAY SKY

WEDNESDAY, JUNE 3, 2009

Fermi LAT Weekly Report N. 52

Covered period: 2009.May.25 - 2009.May.31

- Candidate blazar **4C31.03** (see [ATel #2054](#)) seen in day timescales with flux levels reaching $0.8e-6$ ph/cm²/s.
- **PKS 1510-089** remains in the $1e-6$ to $2e-6$ daily flux range (>100 MeV)
- **PKS 1502+106** shows a steady trend with daily fluxes (>100 MeV) around $1e-6$ ph/cm²/s.
- **3C 454.3** showed consistent daily flux levels (>100 MeV) just below $1e-6$ ph/cm²/s.

LAT DATA

[LAT Monitored Source List Light Curves](#)

[LAT Bright Source List](#)

[Browse interface to monitored source data](#)

BLOG ARCHIVE

▼ 2009 (8)

▼ June (1)

[Fermi LAT Weekly Report N. 52](#)

► May (4)

► April (3)

CONTRIBUTORS

[Flare Advocate](#)

About extra-galactic sources:

☉ Talks:

- ◆ General AGN - P.Giommi
- ◆ Diagnostic of the peak frequencies in the SED of Fermi blazars –
A. Tramacere
- ◆ Fermi-Swift synergetic campaign on the new gamma-ray blazar PKS
1502+106 - S.Ciprini

☉ Posters:

- ◆ Analysis of the Fermi SED - D.Gasparrini
- ◆ Fermi-LAT Flare Advocate Activity during the First Year of Mission -
S.Ciprini

And even more

☉ Talks:

- ◆ Search for Dark Matter in the sky – A. Morselli
- ◆ Indirect Search for Dark Matter from the Milky Way center with the LAT detector on-board Fermi– V. Vitale
- ◆ Measurement of the high energy cosmic-ray electron spectrum with the Fermi Large Area Telescope - F.Loparco
- ◆ Fermi measurements of diffuse gamma-ray emission: results at the first-year milestone - L. Tibaldo

Conclusions


- The Fermi Gamma-Ray Space Telescope has been performing very well and stably for the first year of operations
- Photon data are public since August 25, 2009
 - Join the fun at <http://fermi.gsfc.nasa.gov/ssc/>
- Wealth of results in γ -ray astrophysics
 - ~ 50 pulsars detected, many only in γ -rays
 - many flaring active galaxies observed
 - 12 GRBs at high energy
 - ... and a lot of others results not mentioned

2009 Fermi Symposium

2-5 November 2009
Washington DC

if you want to get more information,
please join to the Fermi mailing lists.

<http://fermi.gsfc.nasa.gov/ssc/resources/newsletter/>



Fermi
Gamma-ray Space Telescope

The Symposium is being held at the
Hyatt Regency on Capitol Hill
in Washington DC

Fermi Symposium, 2-5 November 2009

Local Organizing Committee:

- Neil Johnson (NRL) co-chair
- Dave Thompson (GSFC) co-chair
- Awia Abdo (NRCC/NRL)
- Sandra Barres (USRA/GSFC)
- Orkan Cok (UMBC/GSFC)
- Teddy Cheung (GSFC)
- Chul Gwon (NRL)
- Elizabeth Hays (GSFC)
- Tom Langanstien (Stanford)
- Peter Michelson (Stanford)
- Alex Moiseev (CRESST/GSFC)
- Gerry Share (UMd/NRL)
- Mark Strickman (NRL)
- Vassilis Vasileiou (UMBC/GSFC)
- Michael Wolff (NRL)
- Lucy Zhou (Stanford)

International Science Organizing Committee:

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- R. Bellazzini (Pisa)
- R. Blandford (Stanford/KIPAC)
- E. Bloom (SLAC)
- F. Caraveo (INAF-IAGF-Milano)
- Y. Connaughton (UA Huntsville)
- C. Dermer (NRL)
- N. Gehrels (GSFC)
- J. Grenier (MPE)
- I. Grenier (Laboratoire AIM, Saclay)
- D. Huter (LIR)
- B. Jannus (FONC)
- S. Johnston (ATNF)
- N. Kawai (Tokyo)
- P. Michelson (Stanford)
- A. Mészáros (BU)
- J. McEnery (GSFC)
- J. Ormer (Denver)
- W. Paciesas (UA Huntsville)
- A. Readhead (Caltech)
- S. Ritz (GSFC)
- J. Urrutia (NRAD)
- S. Wagner (Heidelberg)

<http://fermi.gsfc.nasa.gov/science/symposium2009/>