

UPDATE ON THE BACKGROUND ESTIMATE AND VETO STUDIES

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STAWELL
UNDERGROUND
PHYSICS LAB



OUTLOOK



- Cosmogenic activation
- Background measurements in the PoP phase
- CIS effect on veto efficiency

COSMIC ACTIVATION



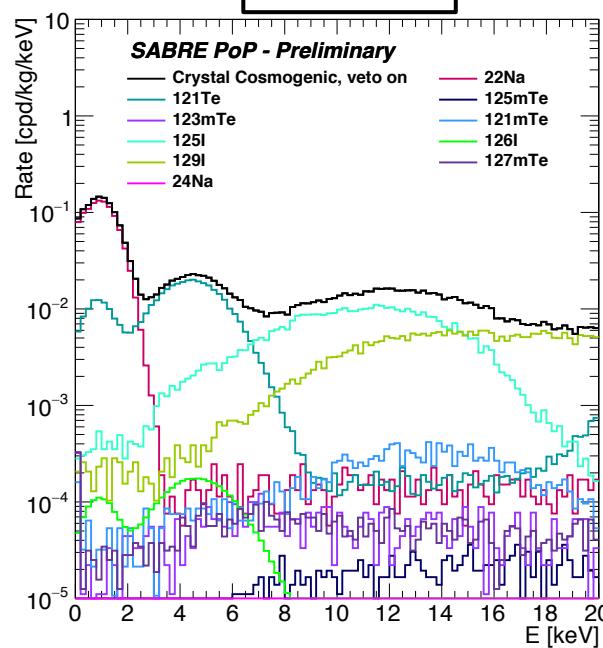
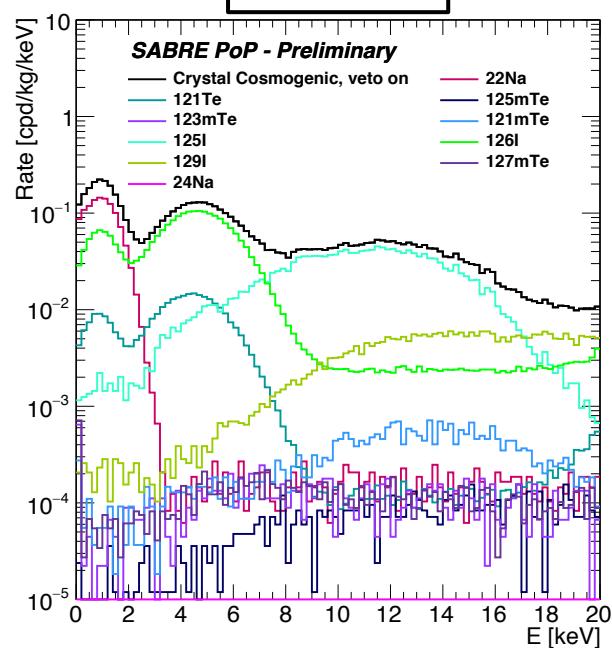
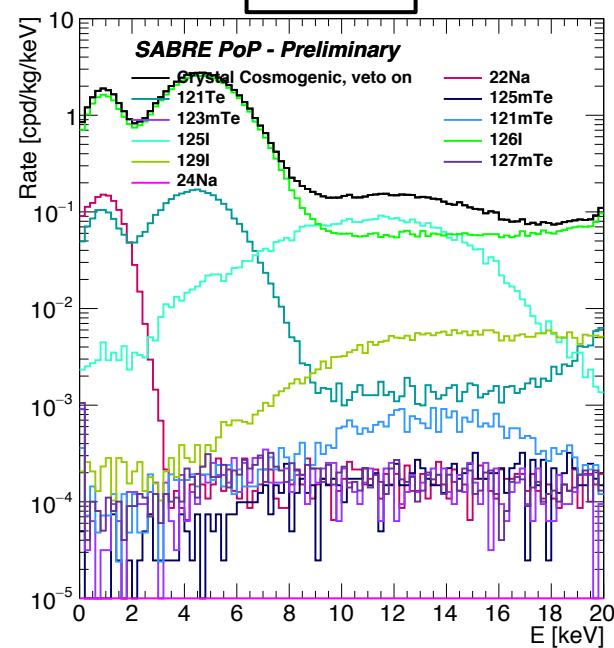
| Isotope | Half-life | Comment |
|---------------------------|-----------|-------------------------------|
| ^{22}Na | 2.602 y | Peaks at 1-2 keV |
| ^{121}Te | 16.78 d | Peaks at 3-7 keV |
| $^{121\text{m}}\text{Te}$ | 154 d | Regenerates ^{121}Te |
| ^{125}I | 59.40 d | Dominant above 7 keV |
| ^{126}I | 12.93 d | Disappears quickly |

- Neutron rate 440 times higher at 12 km then surface**
- 1 day flight ~ 1 year surface in exposure
- 8.8 half-lives to reduce the contaminations to 1 day of surface exposure

0 days

60 days*

180 days



* ^{121}Te regeneration not considered

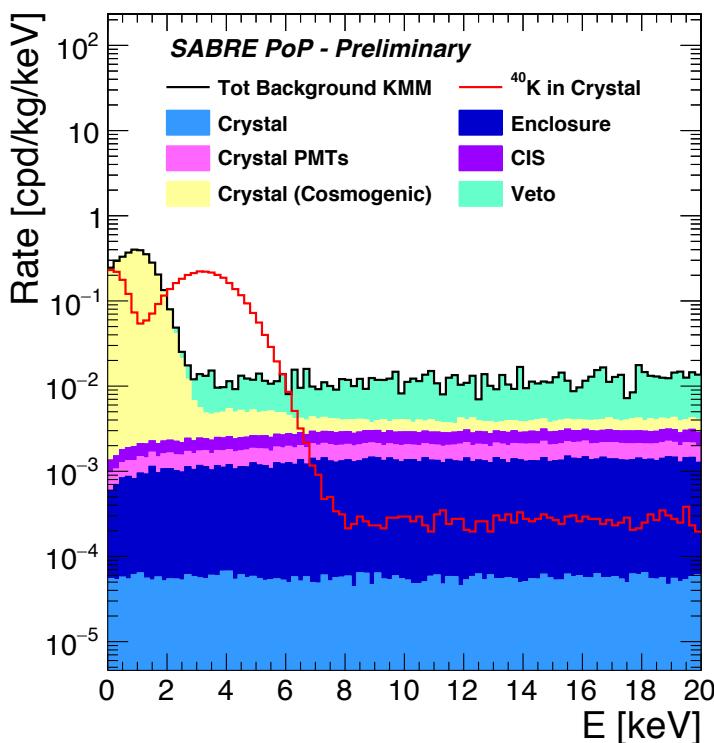
POP SENSITIVITY



- The PoP phase allows the measurement of the crystal contaminations
- What is the precision expected?

$E(\text{Scintillator}) \in [1280, 1640] \text{ keV}$, $E(\text{Crystal}) \in [2, 4] \text{ keV}$

40K measurement mode



| Isotope | Expected rate [cpd/kg/keV] |
|------------------------------|----------------------------|
| ^{40}K in Crystal | 1.91E-01 |
| Crystal (other radiogenic) | 5.85E-05 |
| Crystal (cosmogenic 60 days) | 4.13E-02 |
| Crystal PMTs | 5.78E-04 |
| Enclosure | 1.14E-03 |
| CIS | 9.70E-04 |
| Veto | 6.27E-03 |
| Total background | 5.1 E-02 |

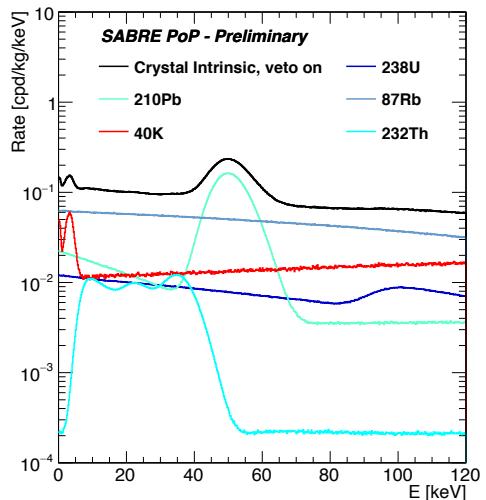
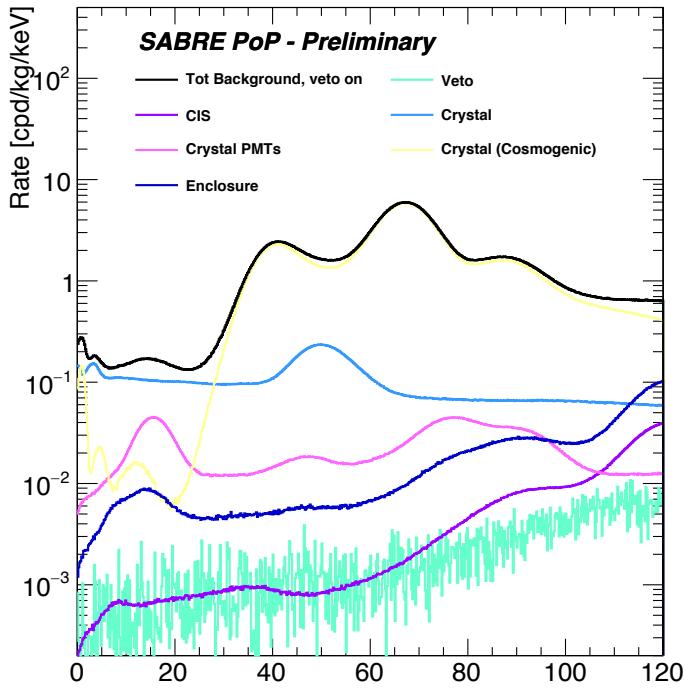
After 90 days:

$S=172$, $B=46$
 $\text{SQRT}(S+B)/S=8.6\%$
equivalent to 0.86 ppb of ^{40}K

Can do a template fit to data to measure ^{40}K and (if enough low energy sensitivity) ^{22}Na

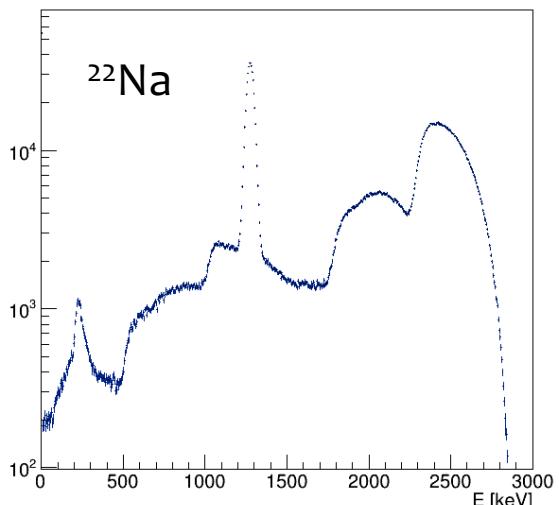
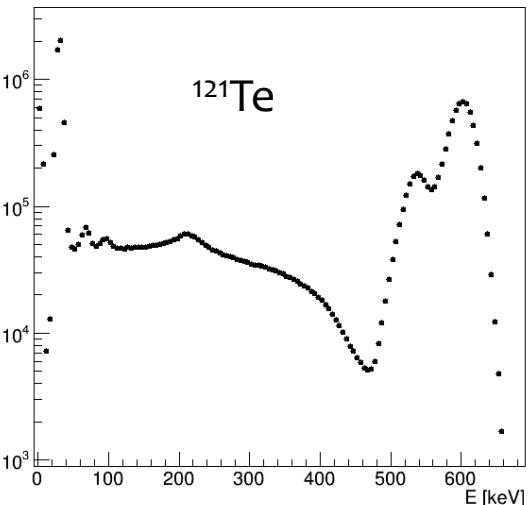
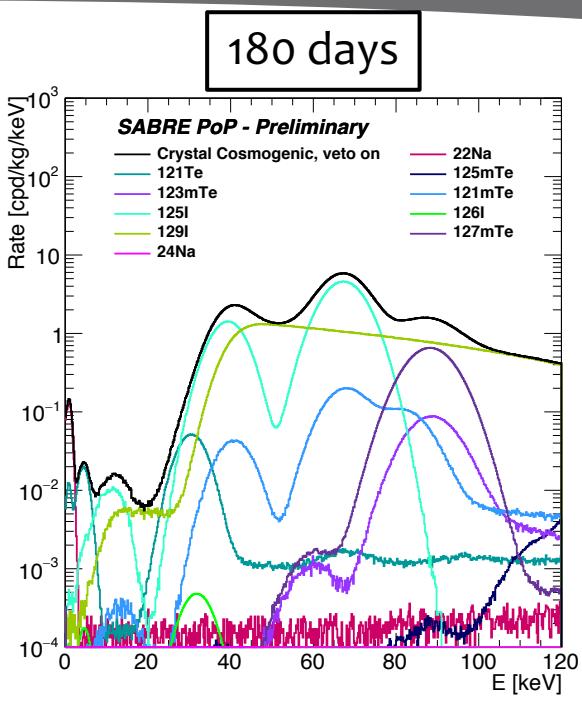
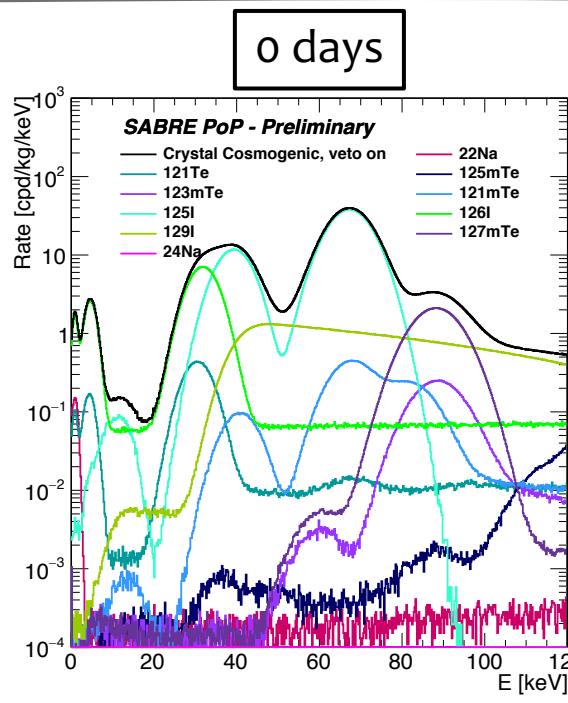
Can we measure other contaminations?

INTRINSIC BACKGROUNDS



- Cosmogenic background dominates the higher energies
- ^{210}Pb could be measured at the 46 keV peak which contrast the well in the cosmogenic background (resolution important)
- ^{87}Rb gives exponential falling background, probably hard to measure
- U and Th too small

COSMOGENIC BACKGROUNDS

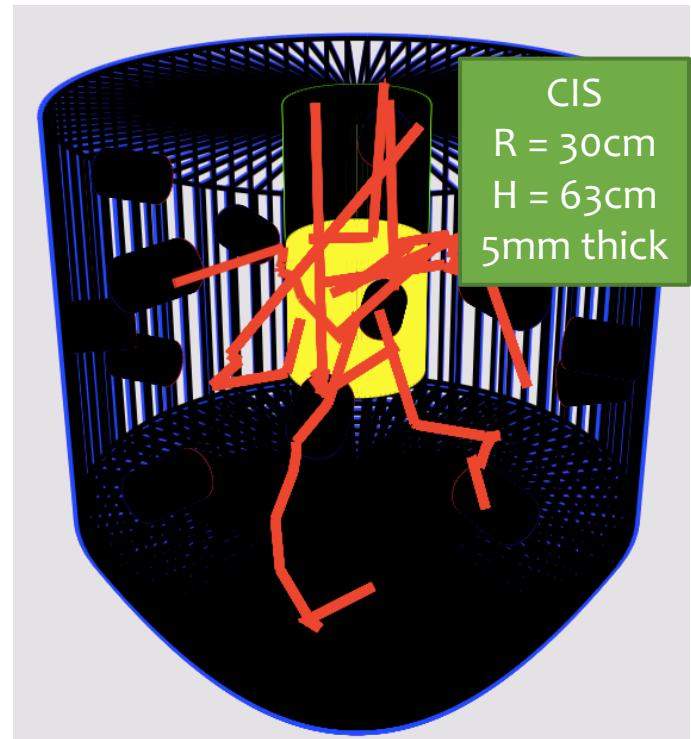
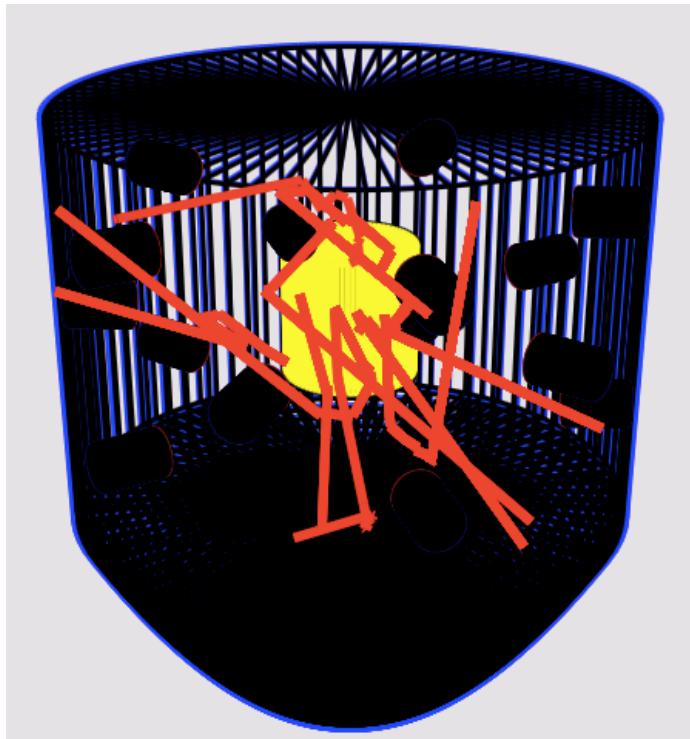


- ➊ ^{125}I from the double peak at 30 keV and 70 keV
- ➋ ^{121}Te and ^{22}Na from higher energy bands

INSERTION SYSTEM



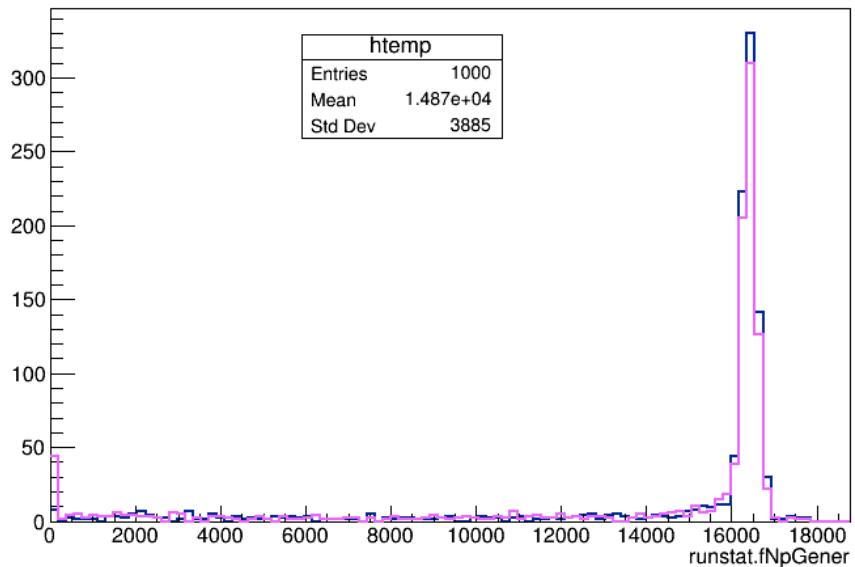
- Is the dead volume of the CIS significantly reducing the veto efficiency?
- Compared solution with/without CIS



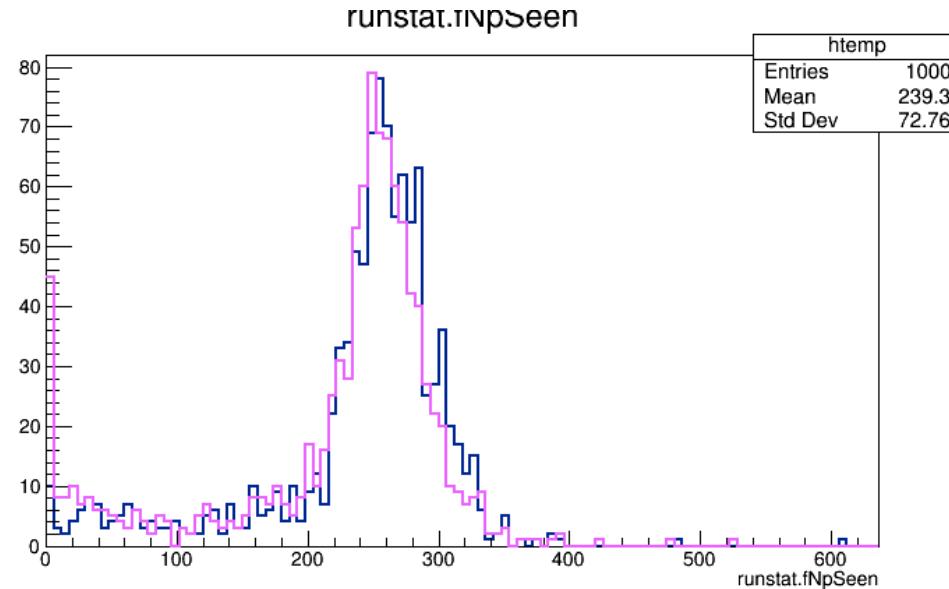
LIGHT YIELDS



Generated photons



Observed photons



Backup