

# Cross-section measurements by means of laser beams

(and future experiments @ ELI-NP?)

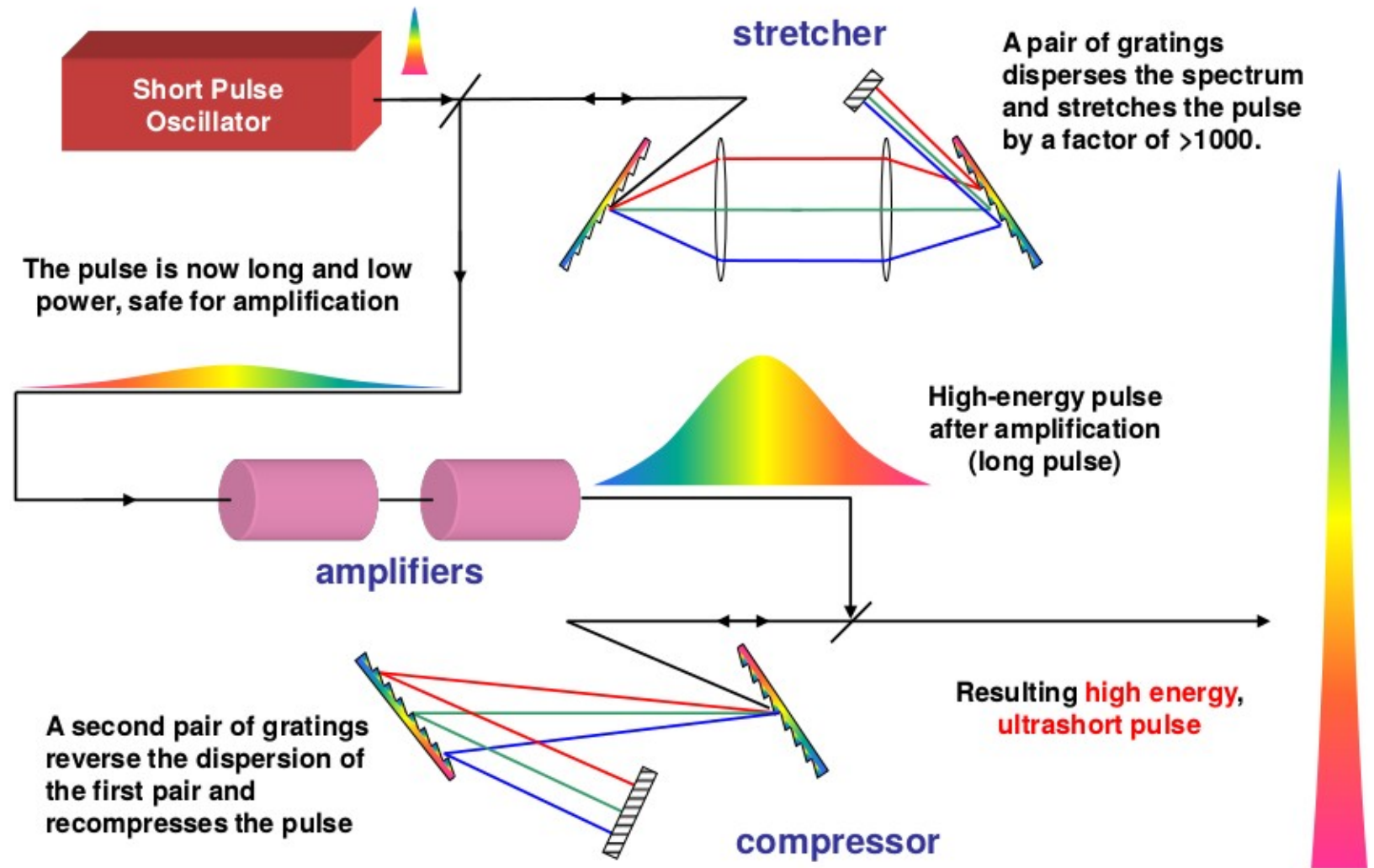


**Dario Lattuada**  
**ELI-NP, LNS**

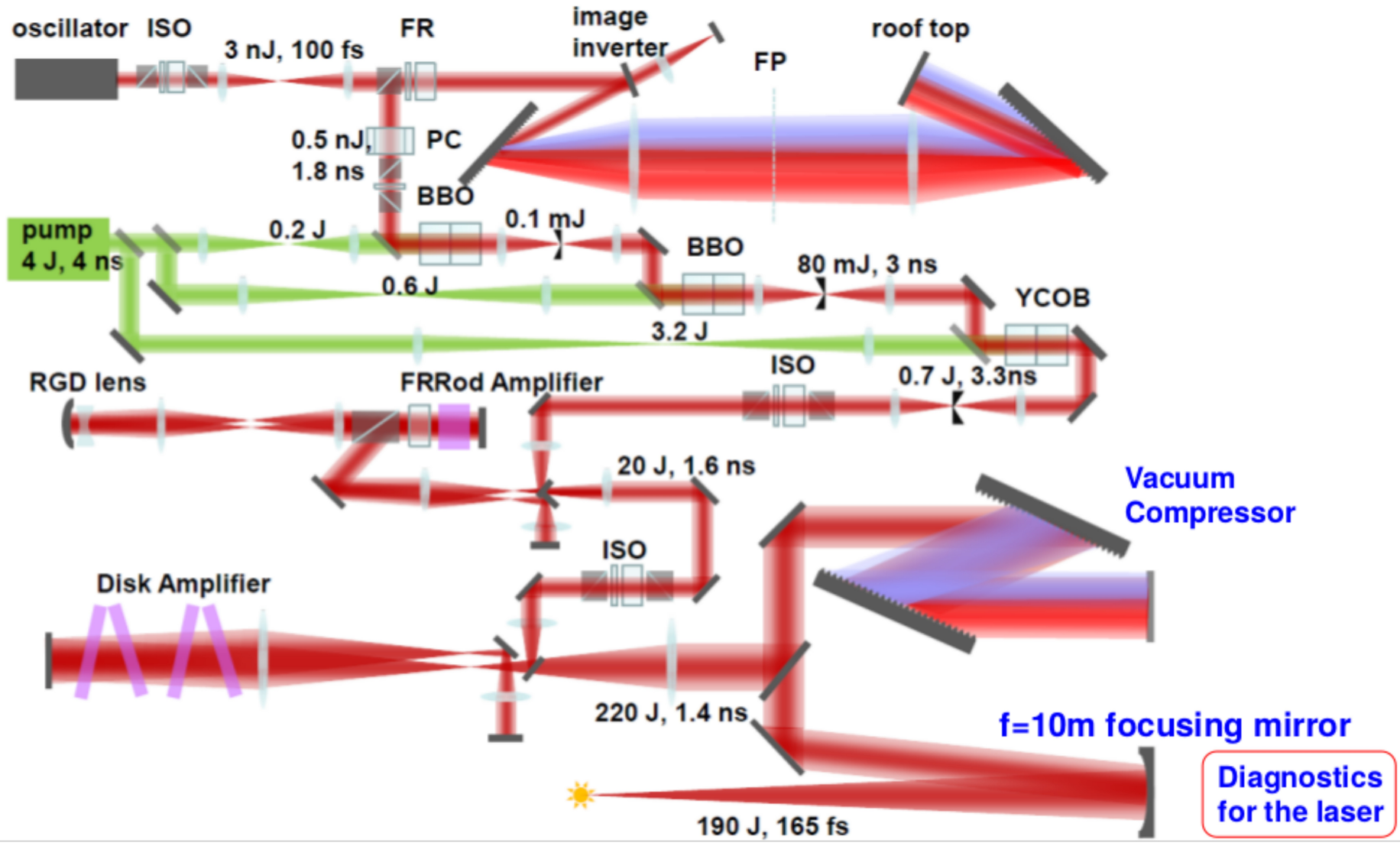
# Outlines

- Laser-cluster experiment @ UT PW Laser
- Cross-section measurements (results)
- Future experiments: UT and ELI-NP
- Photonuclear reactions
- ELI-NP and the GBS
- ELISSA

# Chirped Pulse Amplification (CPA) is the basic technique for modern ultra-intense, ultrafast lasers



# Texas Petawatt laser uses three OPCPA and two Nd:glass amplifier stages and delivers a 190J laser pulse



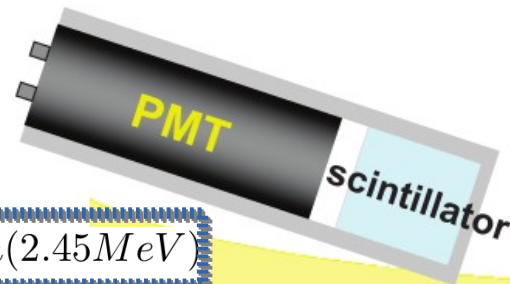
# UT Petawatt Laser

## Nuclear fusion from laser-cluster interaction

<sup>3</sup>He mixture)

D<sub>2</sub> gas tank  
(high backing pressure)

Supersonic Nozzle



10<sup>2</sup>J released in ~ 10<sup>-13</sup> s

High power  
laser pulse in

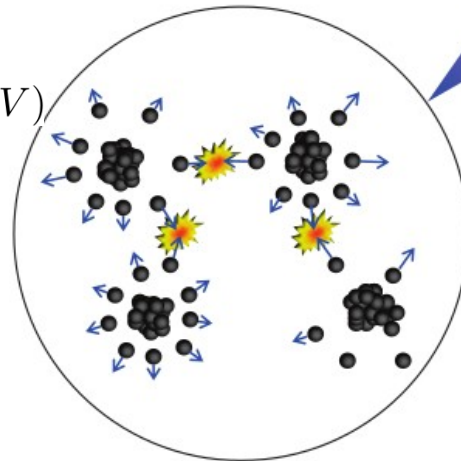


deuterium ions

Kinetic Energy < 10<sup>2</sup> keV

Density ~ 10<sup>18</sup> atoms/cm<sup>3</sup>

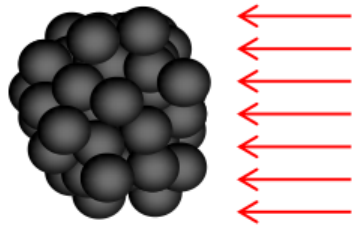
10<sup>5</sup>-10<sup>7</sup> neutrons per shot



- Most of the laser pulse energy is absorbed by the atomic clusters.
- Clusters experience Coulomb explosion after electrons escape.
- DD fusion occurs, and 2.45MeV fusion neutrons are produced.

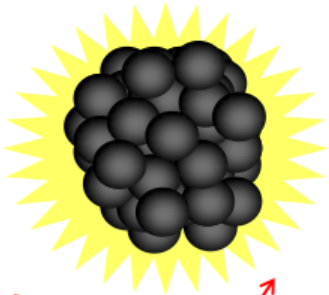
(and D<sup>3</sup>He)

# Coulomb Explosion



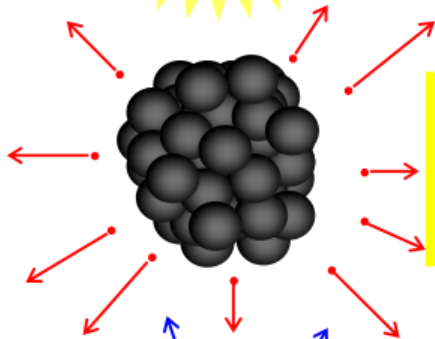
## Step 1

Clusters are irradiated by high intensity laser pulse ( $\sim 10^{16} \sim 10^{18}$  W/cm<sup>2</sup>).



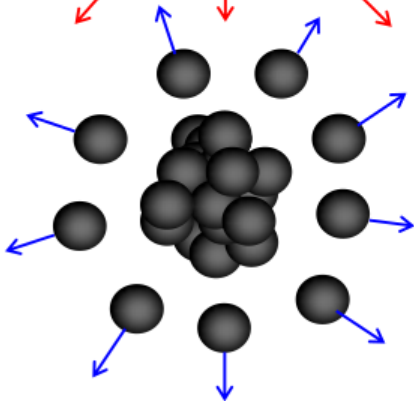
## Step 2

Laser pulse energy is first absorbed by electrons via heating mechanisms such as rapid collisional heating.



## Step 3

Electrons escape from the cluster and leave positive charge build-up on the cluster.



## Step 4

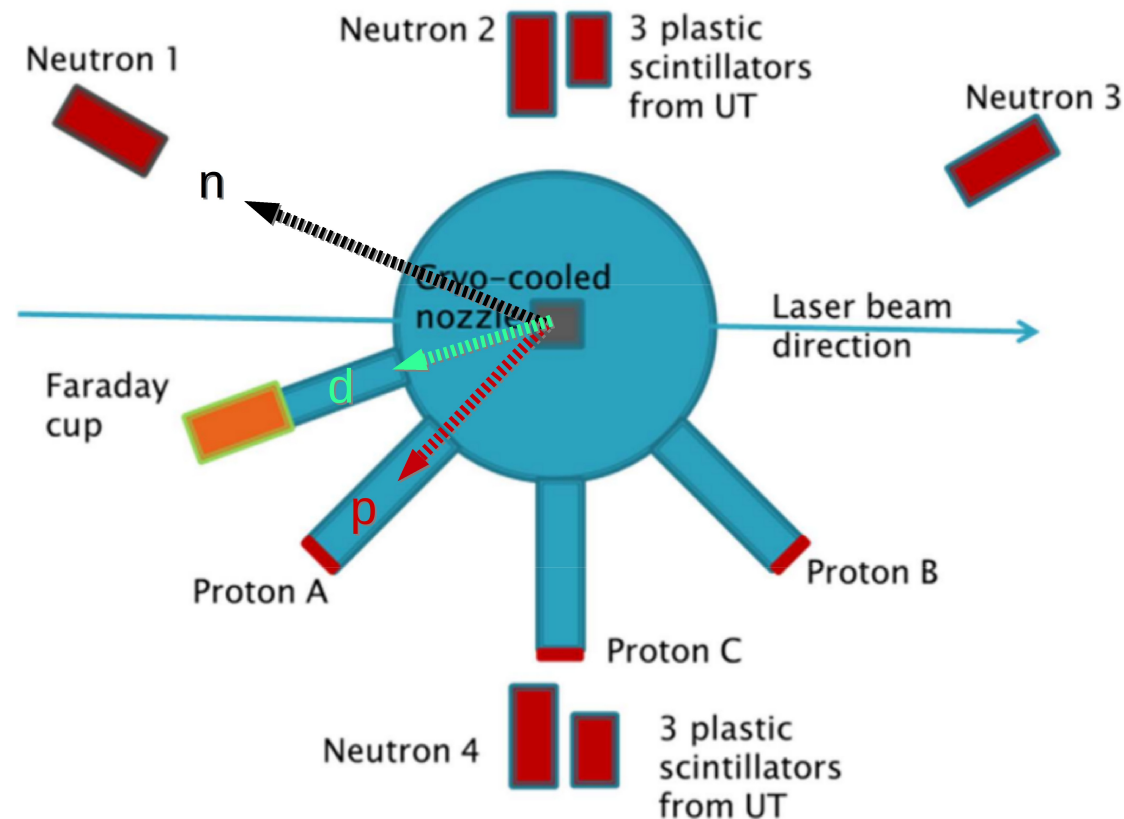
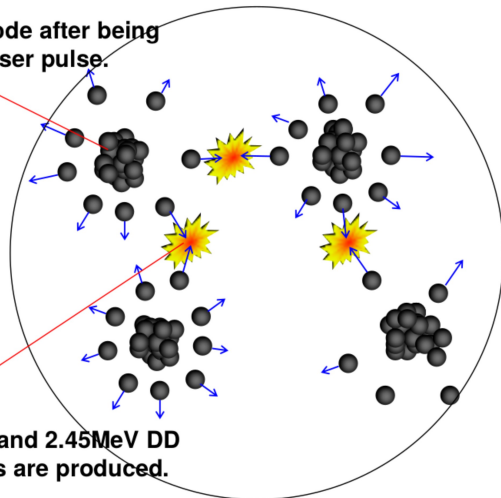
The cluster “explodes” and deuterons acquire multi-keV kinetic energy.

# The 2011 experiment

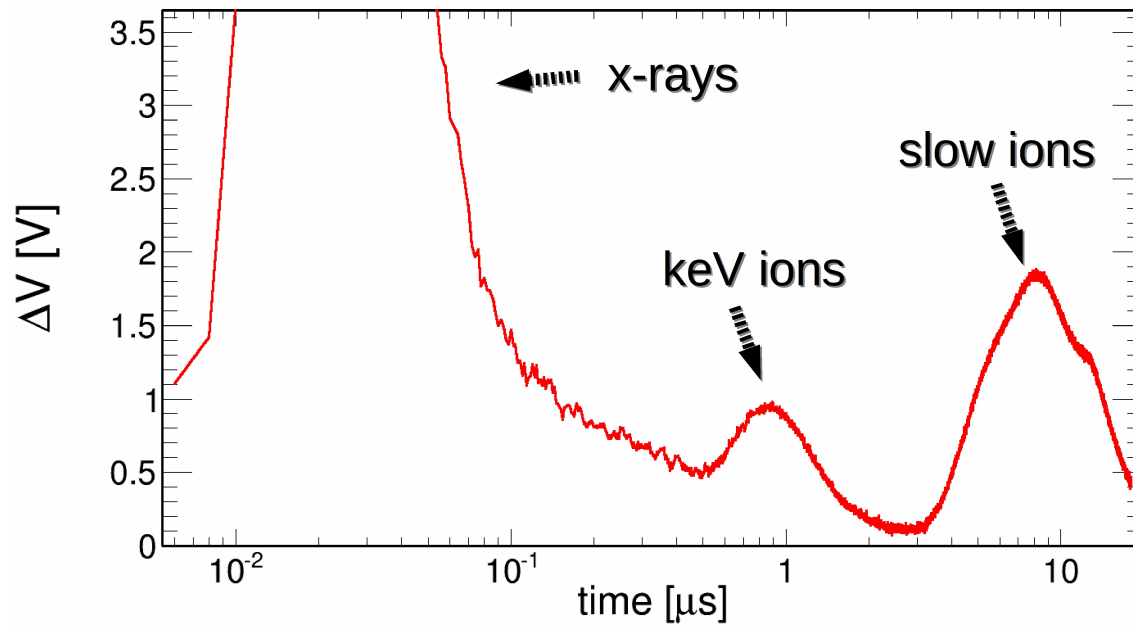


$$Y_n{}_{BB} = l\rho_D \int \frac{dN}{dE} \sigma_{BB}(E) dE$$

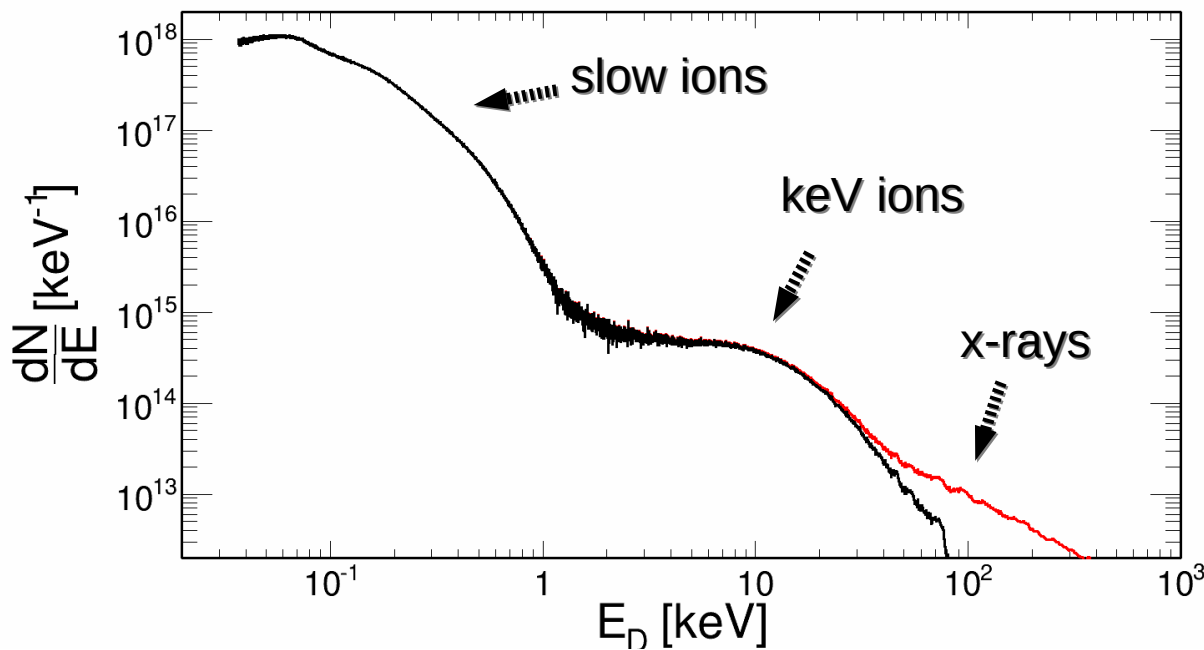
Clusters explode after being heated by a laser pulse.



# The 2011 experiment



$$\frac{d^2 N}{dt d\Omega} = \frac{\Delta V}{qe R_\Omega \Delta \Omega}$$



$$\frac{d^2 N}{dE d\Omega} = \frac{s^3}{m_D v_D^3 \pi r_F^2} \frac{\Delta V}{qe R_\Omega}$$



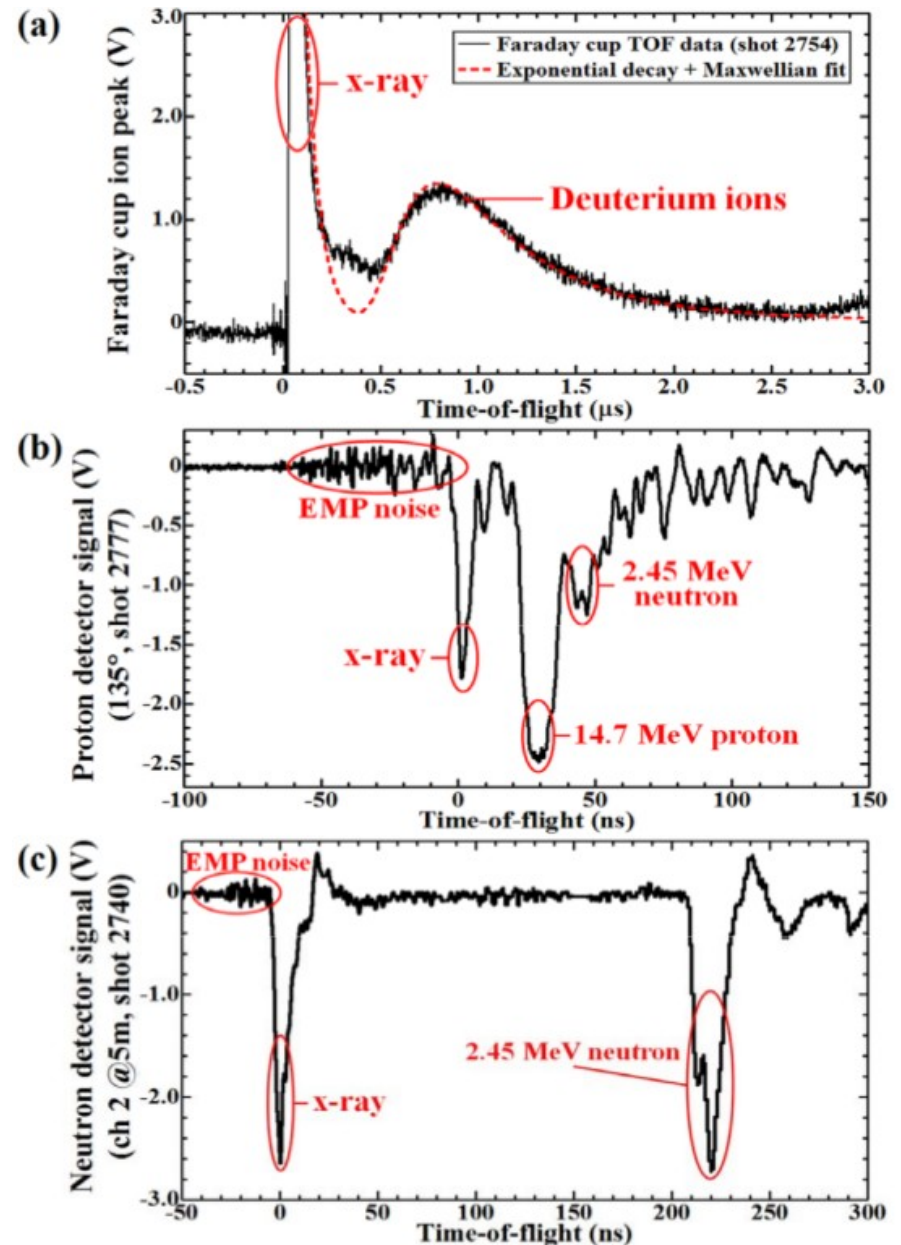
# The 2011 experiment

$$Y_{n,BB} = l\rho_D \int \frac{dN}{dE} \sigma_{BB}(E) dE$$

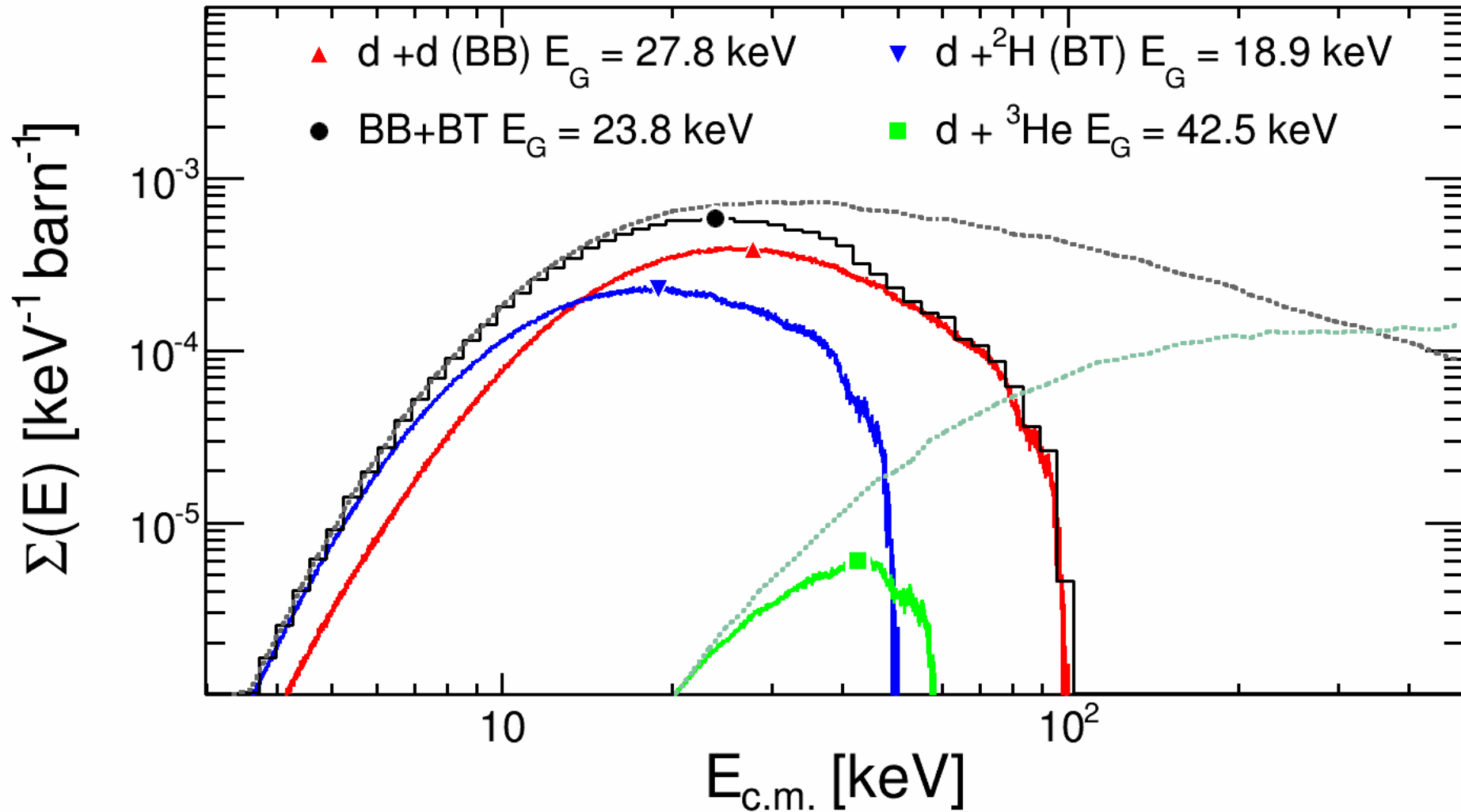
$$\sigma(E) = \frac{S(E)}{E} \exp(-2\pi\eta(E))$$

$$2\pi\eta = b/\sqrt{E}$$

$$b = 0.9898 Z_i Z_j \sqrt{A} \text{ MeV}^{1/2}$$

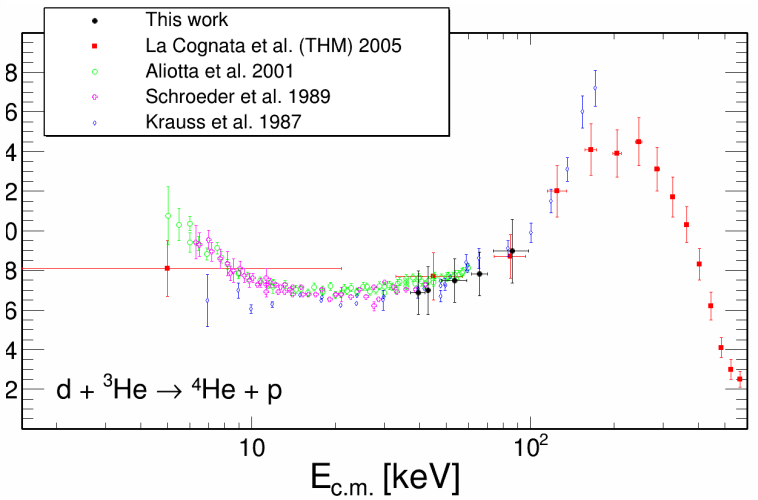
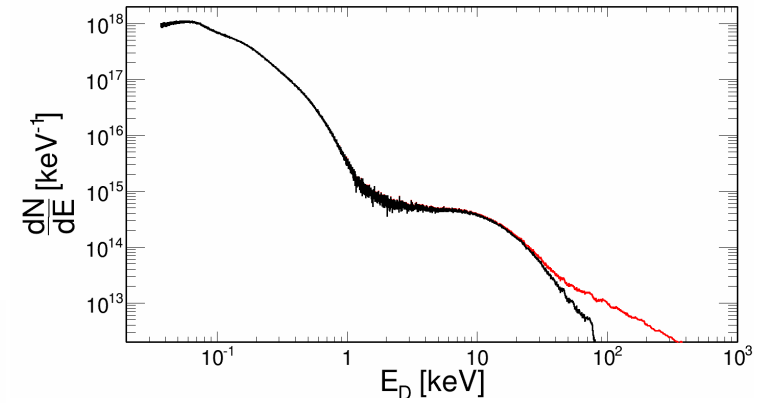
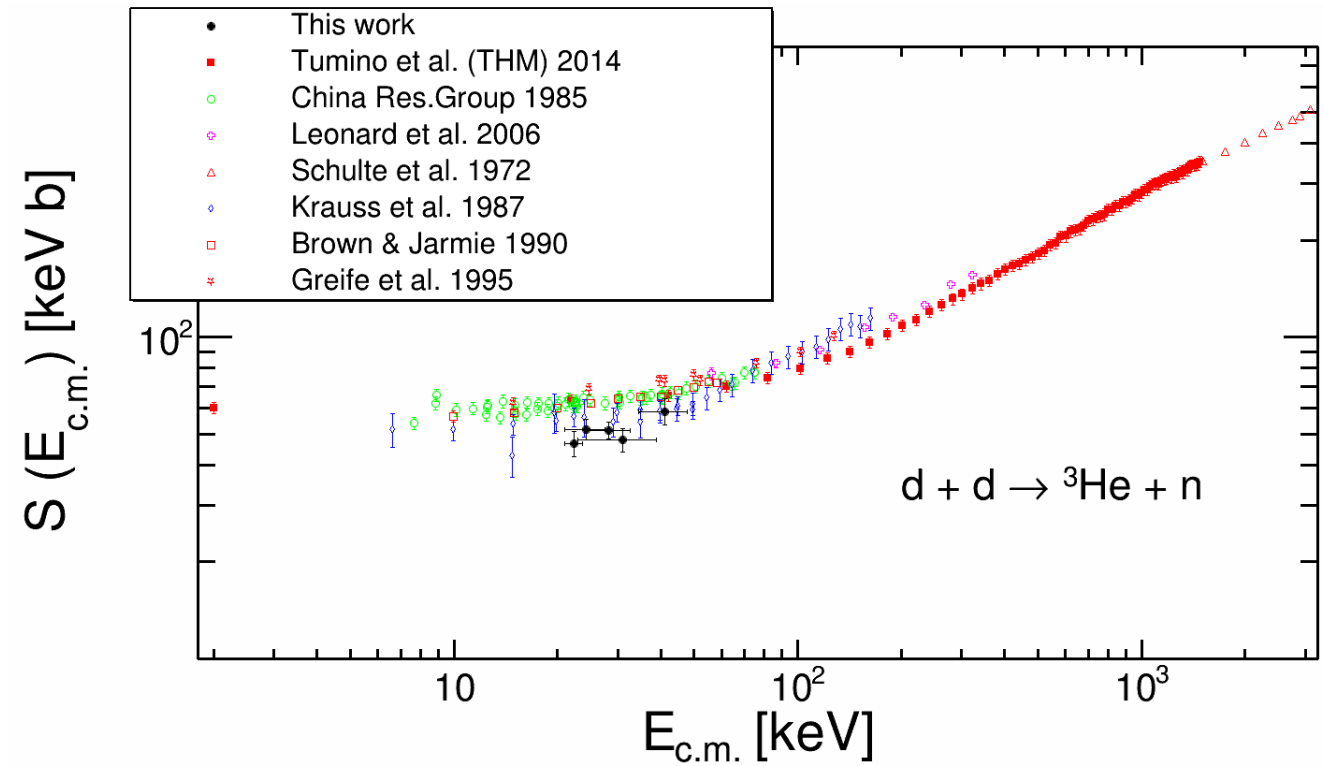


# The 2011 experiment

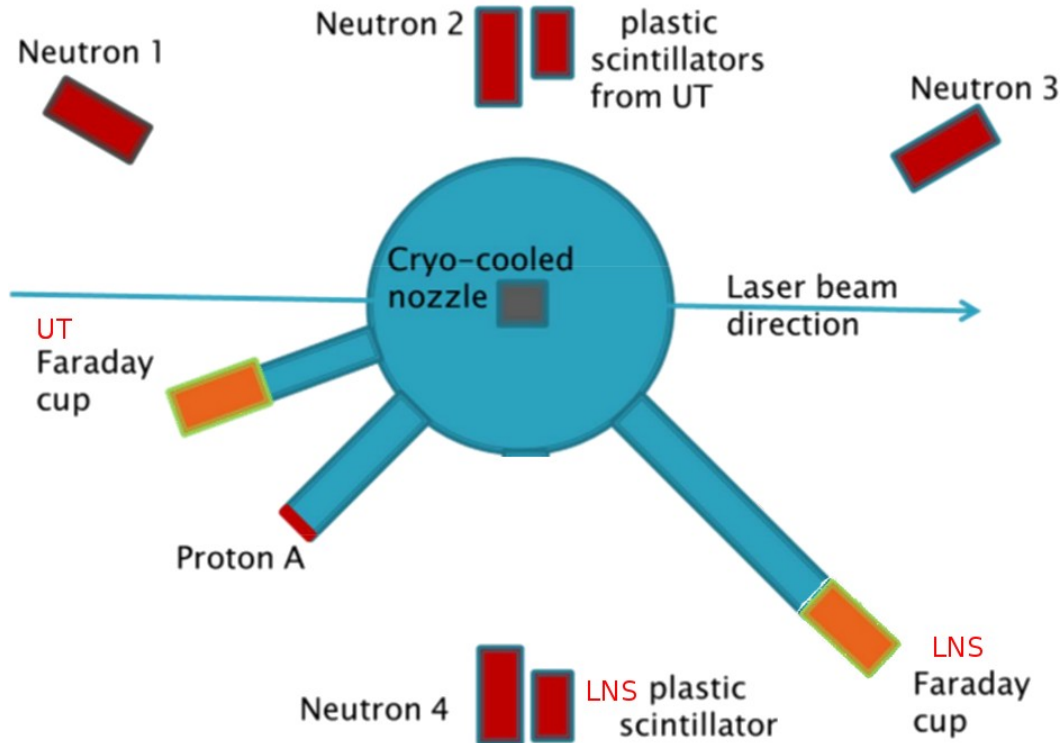
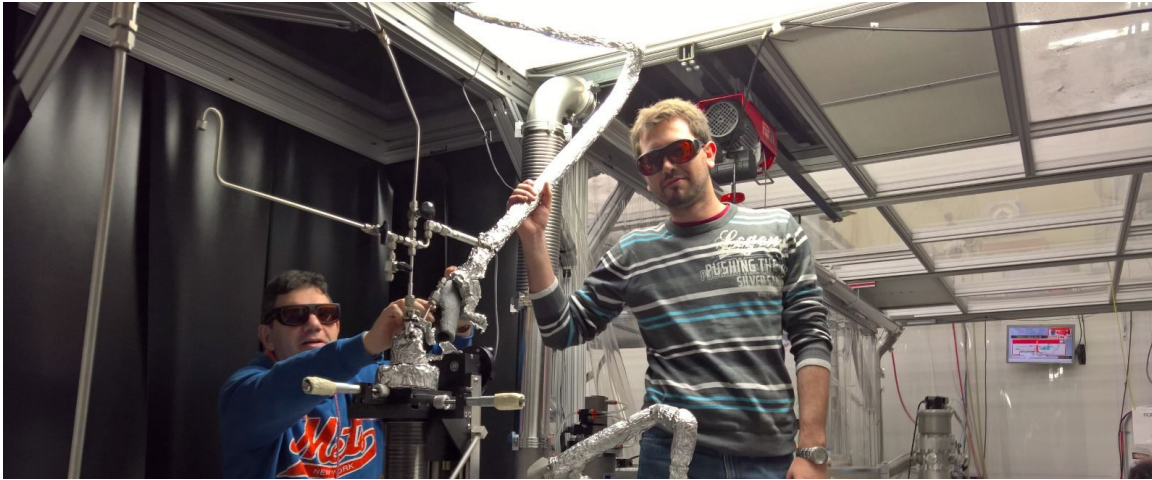


## Model-independent determination of the astrophysical $S$ factor in laser-induced fusion plasmas



D. Lattuada, M. Barbarino, A. Bonasera, W. Bang, H. J. Quevedo, M. Warren, F. Consoli, R. De Angelis, P. Andreoli, S. Kimura, G. Dyer, A. C. Bernstein, K. Hagel, M. Barbui, K. Schmidt, E. Gaul, M. E. Donovan, J. B. Natowitz, and T. Ditmire  
 Phys. Rev. C **93**, 045808 – Published 19 April 2016



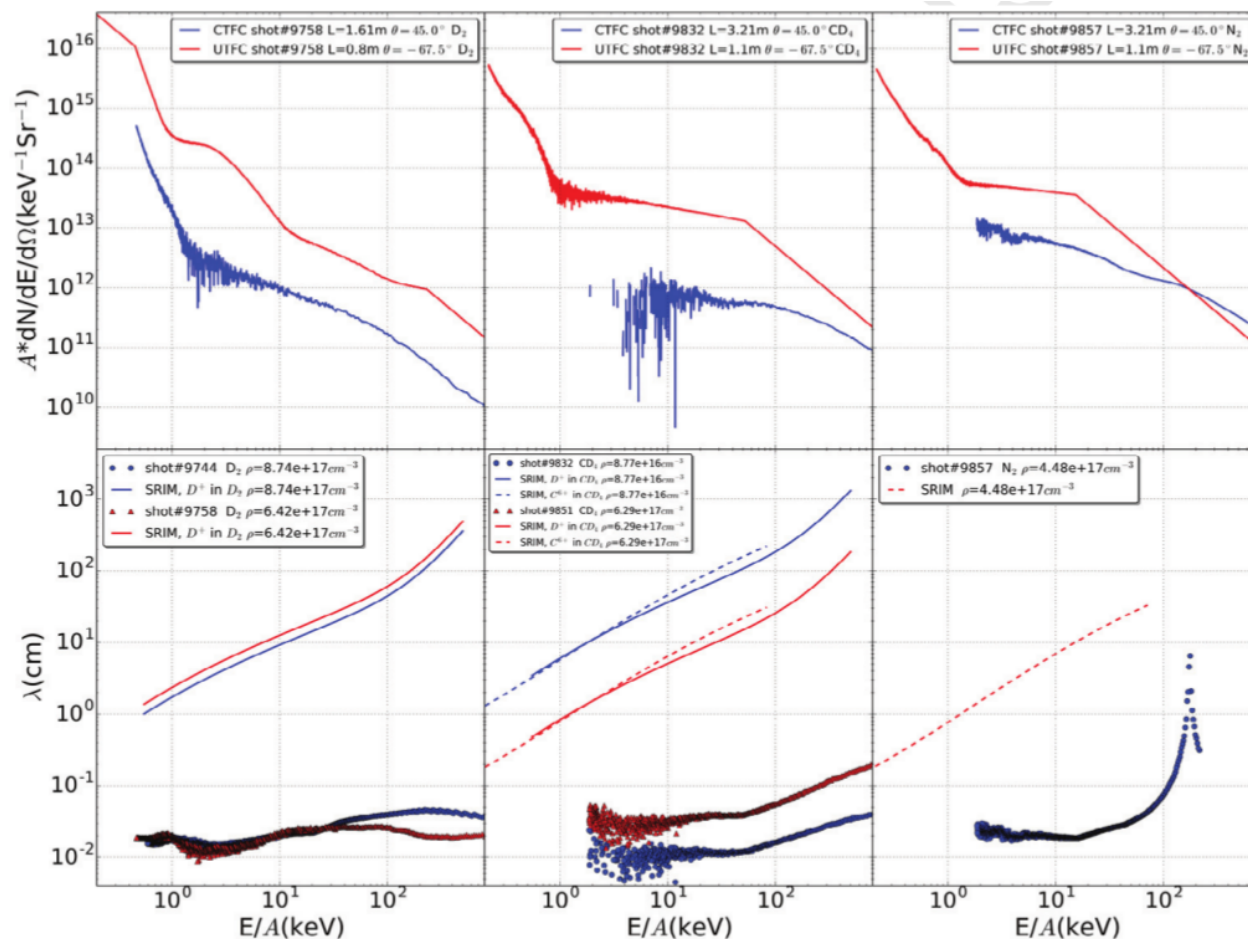
# The 2016 experiment



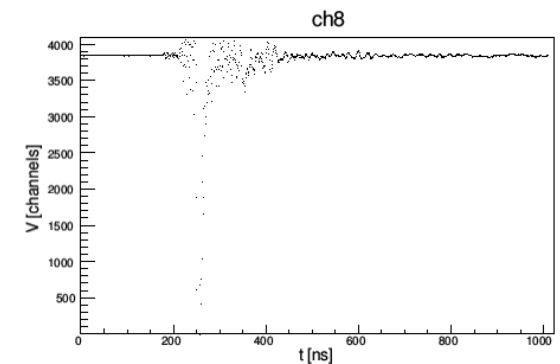
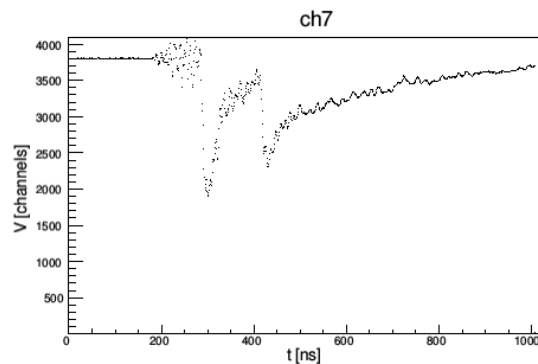
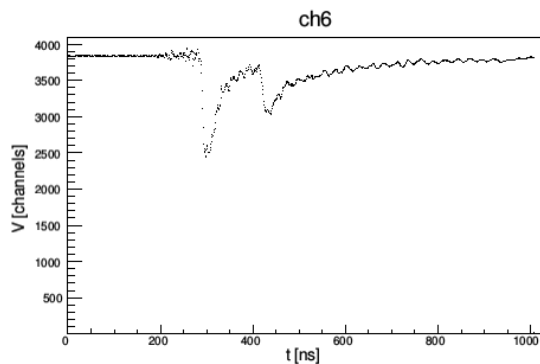
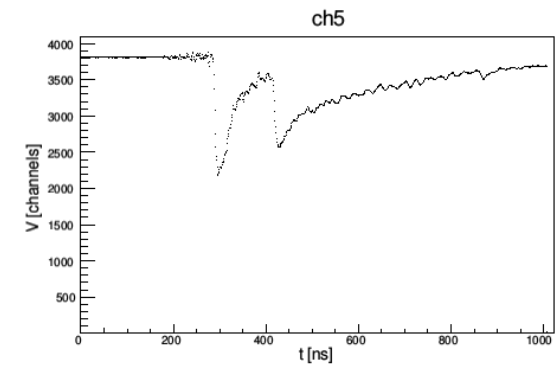
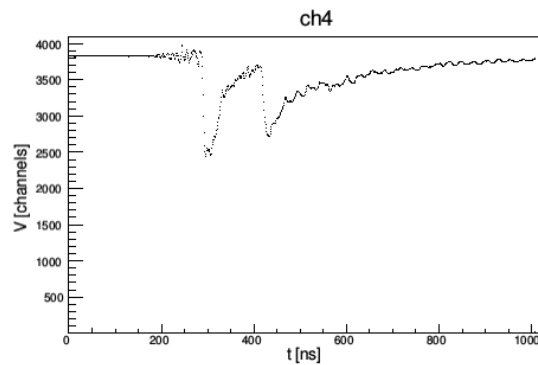
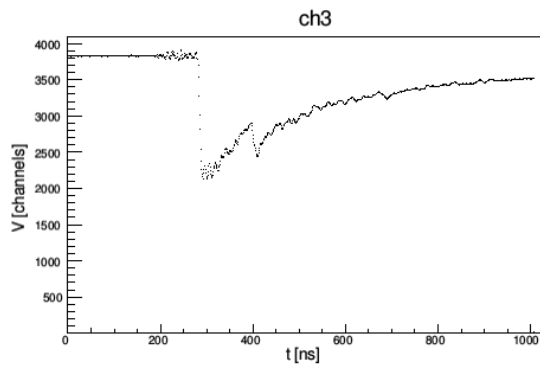
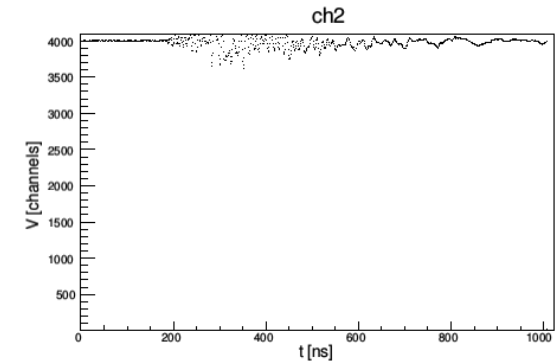
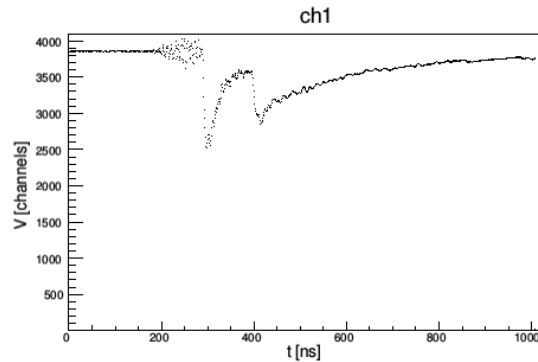
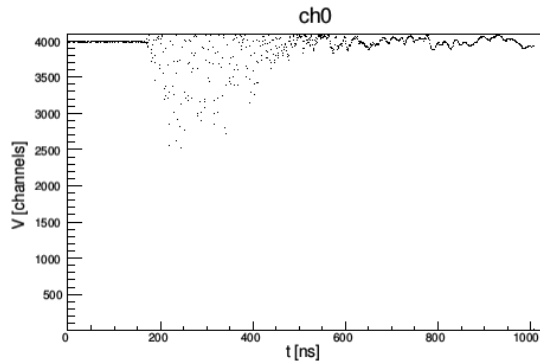
# Range of plasma ions in cold cluster gases near the critical point

G. Zhang<sup>a, 1</sup>, H.J. Quevedo<sup>b</sup>, A. Bonasera<sup>a, c</sup>,  , M. Donovan<sup>b</sup>, G. Dyer<sup>b</sup>, E. Gaul<sup>b</sup>, G.L. Guardo<sup>c</sup>, M. Gulino<sup>c, d</sup>, M. La Cognata<sup>c</sup>, D. Lattuada<sup>c</sup>, S. Palmerini<sup>e, f</sup>, R.G. Pizzone<sup>c</sup>, S. Romano<sup>c</sup>, H. Smith<sup>b, O</sup>, Trippella<sup>e, f</sup>, A. Anzalone<sup>c</sup>, C. Spitaleri<sup>c</sup>, T. Ditmire<sup>b</sup>

- Measurement of the ion range in cold gases
- It strongly depends on initial gas T and P
- “SRIM + homogenous gases” fails to reproduce
- Range increase when near the critical point gas/liquid..



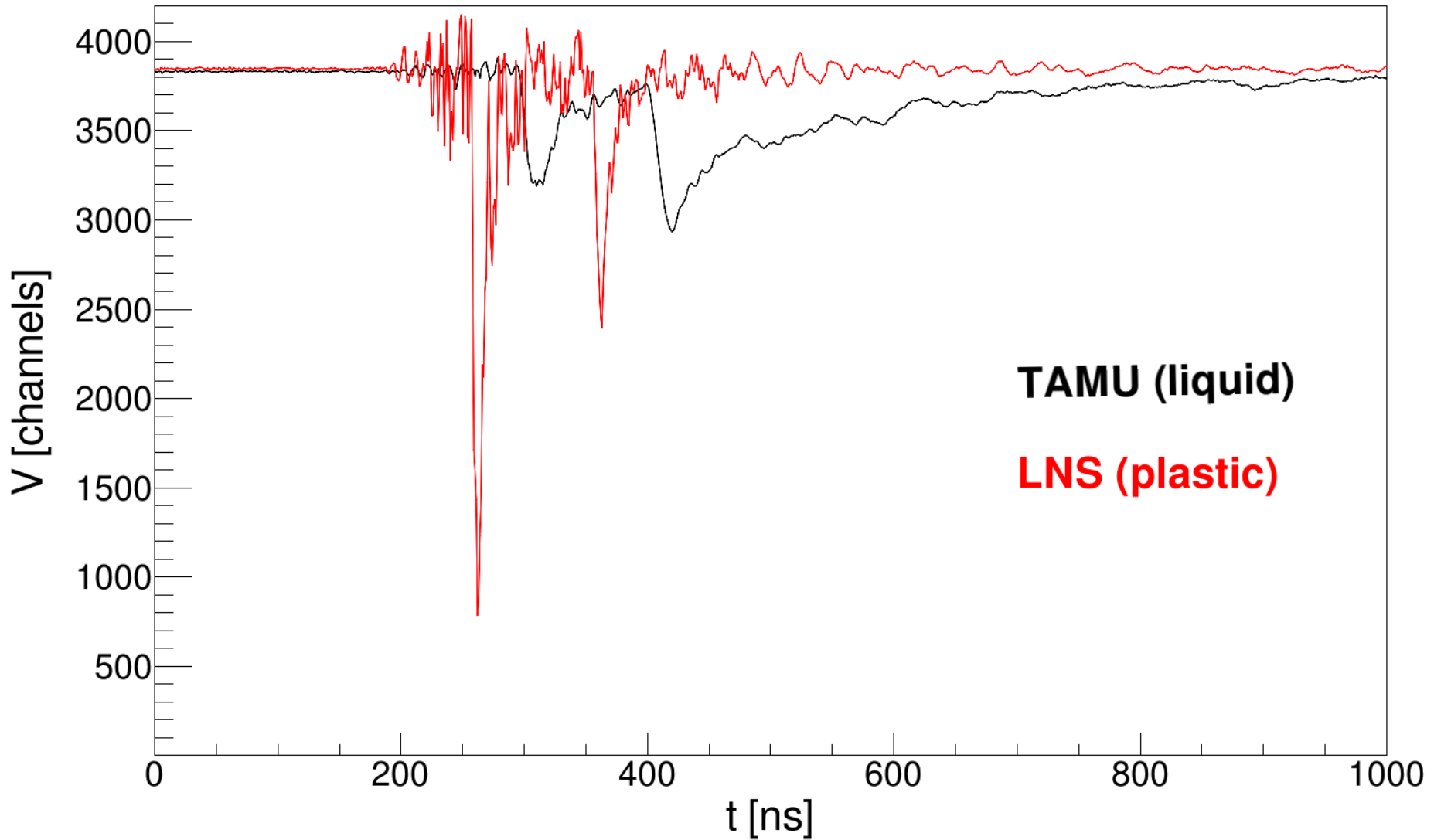
# The 2016 experiment



# The 2016 experiment

Ongoing analysis

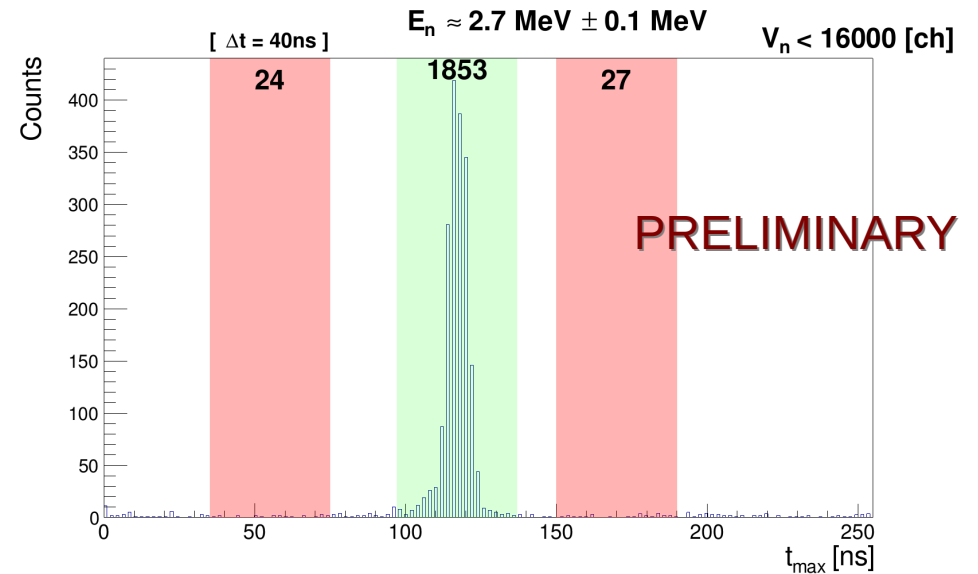
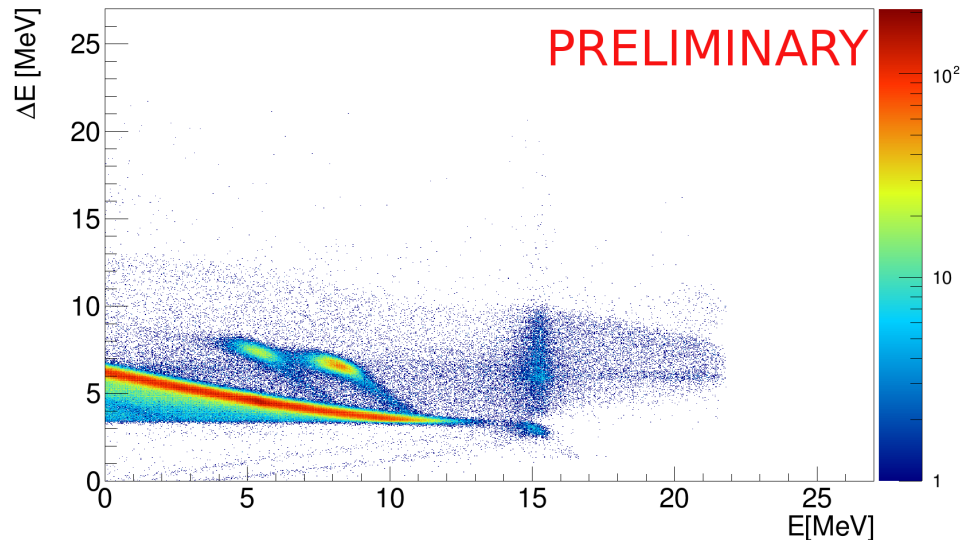
**UT9849**



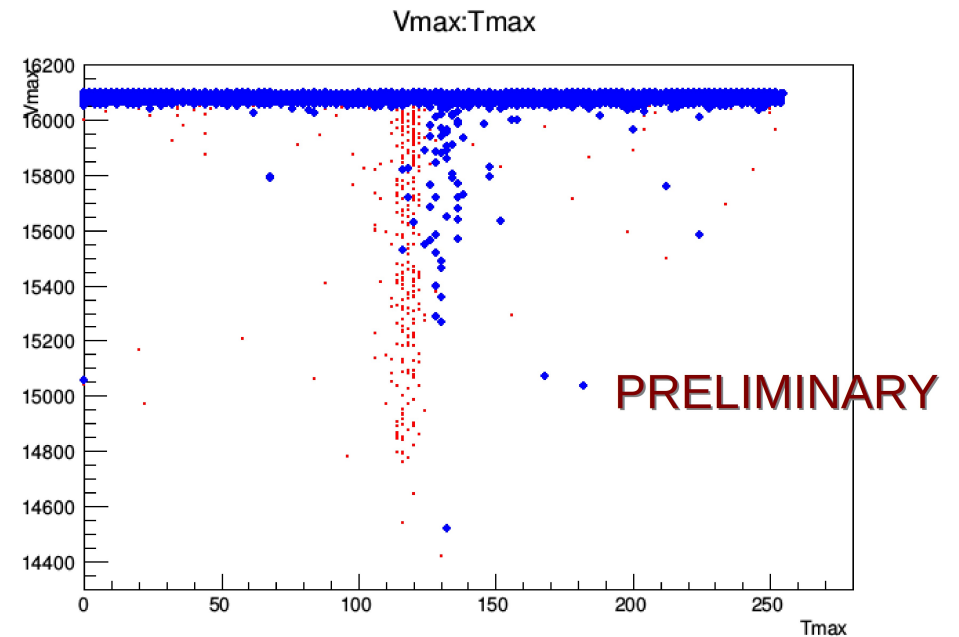
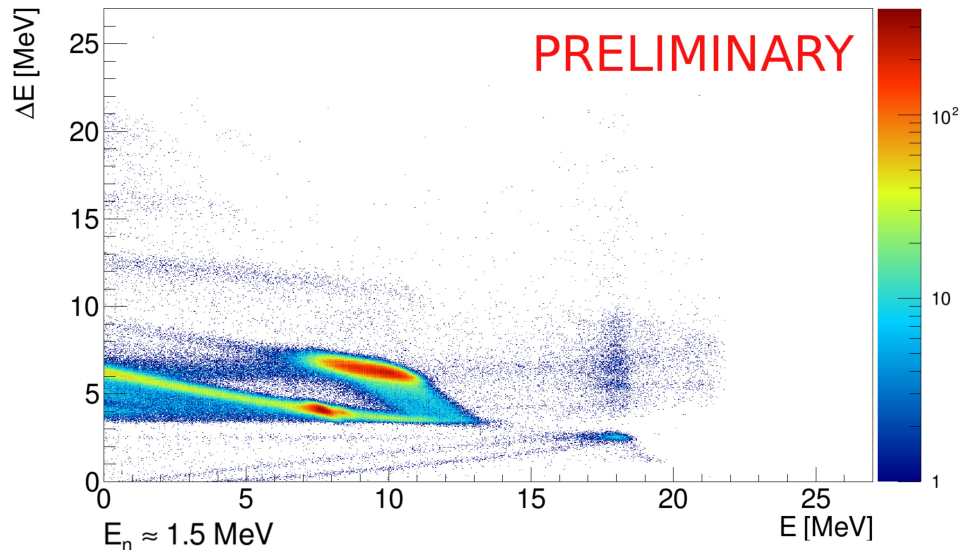
# efficiency tests @LNS

Ongoing analysis

${}^7\text{Li}(p,n){}^7\text{Be}$  @ 18 MeV,  $\theta_{\text{tel}} = 3.7^\circ \pm 0.3^\circ$



${}^7\text{Li}(p,n){}^7\text{Be}$  @ 20 MeV,  $\theta_{\text{tel}} = 4.7^\circ \pm 0.3^\circ$





# Conclusions & 2018+ experiments?

- It is possible to accelerate ions to multi-keV energies using ultrashort laser pulses in laser-cluster experiments
- It is possible to calculate  $S(E)$  without assuming thermal equilibrium (and MB )
- It is possible to extend the measurements to higher and lower energies by tuning the nozzle T and P .
- Ion range is not reproduced by SRIM, with 2-3 orders of magnitude discrepancy
- Critical point... (preliminary, submitted)
  
- Proposal UT 2018 ?
  
- ELI-NP 2020 ? Laser-cluster interaction (RA3)  
photonuclear reactions with SD (ELISSA in RA4)

# Thank you

UT2016 experiment:

A. Anzalone<sup>1</sup>, A. Bonasera<sup>1,2</sup>, T. Ditmire<sup>3</sup>, M. Donovan<sup>3</sup>, G. Dyer<sup>3</sup>, E. Gaul<sup>3</sup>, G. L. Guardo<sup>4</sup>, M. Gulino<sup>1,5</sup>, M. La Cognata<sup>1</sup>, D. Lattuada<sup>4</sup>, S. Palmerini<sup>6,7</sup>, R. G. Pizzone<sup>1</sup>, H. J. Quevedo<sup>3</sup>, S. Romano<sup>1</sup>, H. Smith<sup>3</sup>, O. Trippella<sup>6,7</sup>, C. Spitaleri<sup>1</sup>, G. Zhang<sup>2</sup>

1) Laboratori Nazionali del Sud-INFN, via S. Sofia 64, 95123 Catania, Italy.

2) Cyclotron Institute, Texas A&M University, 77843 College Station, TX.

3) Center for High Energy Density Science, C1510, University of Texas at Austin, Austin, Texas 78712, USA

4) Extreme Light Infrastructure – Nuclear Physics, Magurele, Bucharest, Romania.

5) Libera Università' Kore, 94100 Enna, Italy

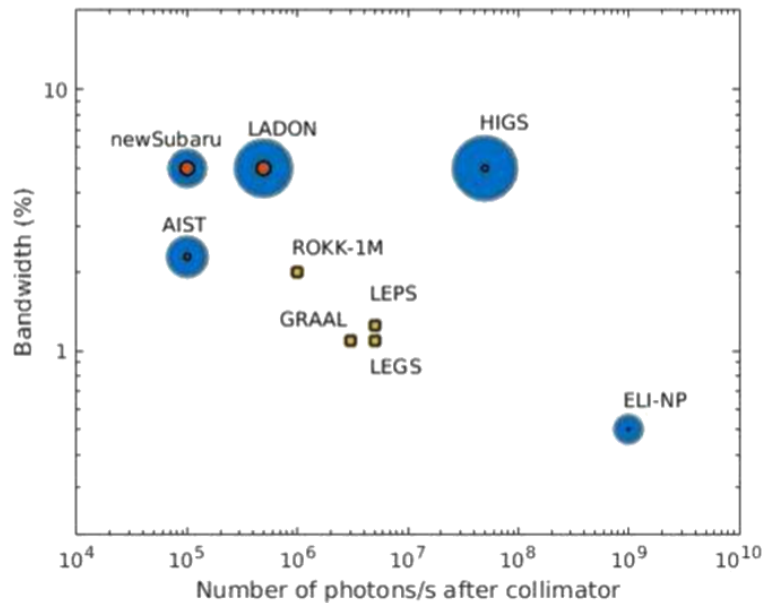
6) Department of Physics and Geology, University of Perugia, Via A. Pascoli, 06123 Perugia, Italy

7) Istituto Nazionale di Fisica Nucleare, Section of Perugia, Via A. Pascoli, 06123 Perugia, Italy

# ELI-NP laser and GBS

Extreme Light Infrastructure - Nuclear Physics

10 PW Laser & Gamma Beam System ..



**Table 1.** The parameters of the gamma beams at ELI-NP Gamma Beam System (GBS).

Gamma beam parameters	Value
Energy [MeV]	0.2-19.5
Spectral density [photons/s/eV]	$>0.5 \cdot 10^3$
Bandwidth [%]	$\leq 0.5$
Peak brilliance [photons/s·mm <sup>2</sup> ·mrad <sup>2</sup> ·0.1% bdw]	$10^{20}-10^{23}$
Pulse length rms [ps]	0.7-1.5
Linear polarization [%]	$>95$
Macro repetition rate [Hz]	100
Number of pulses/macropulse	32
Pulse-to-pulse separation [ns]	16

# ELI-NP & Nuclear Astrophysics



➤ **Big Bang Nucleosynthesis and Li-problem**

${}^7\text{Li}(\gamma, t){}^4\text{He}$  reaction



➤ **Si-burning in stars and presupernova phase**

${}^{24}\text{Mg}(\gamma, \alpha){}^{20}\text{Ne}$  reaction

${}^{28}\text{Si}(\gamma, p){}^{27}\text{Al}$  reaction

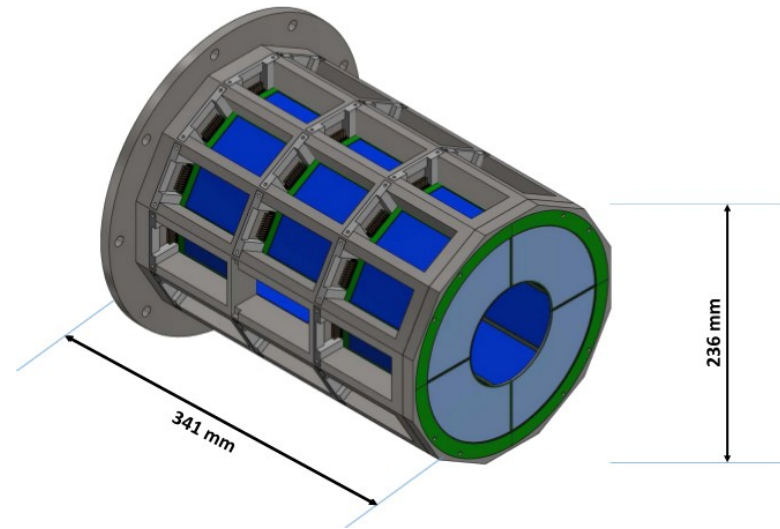


➤ **p-process (production of proton rich nuclei)**

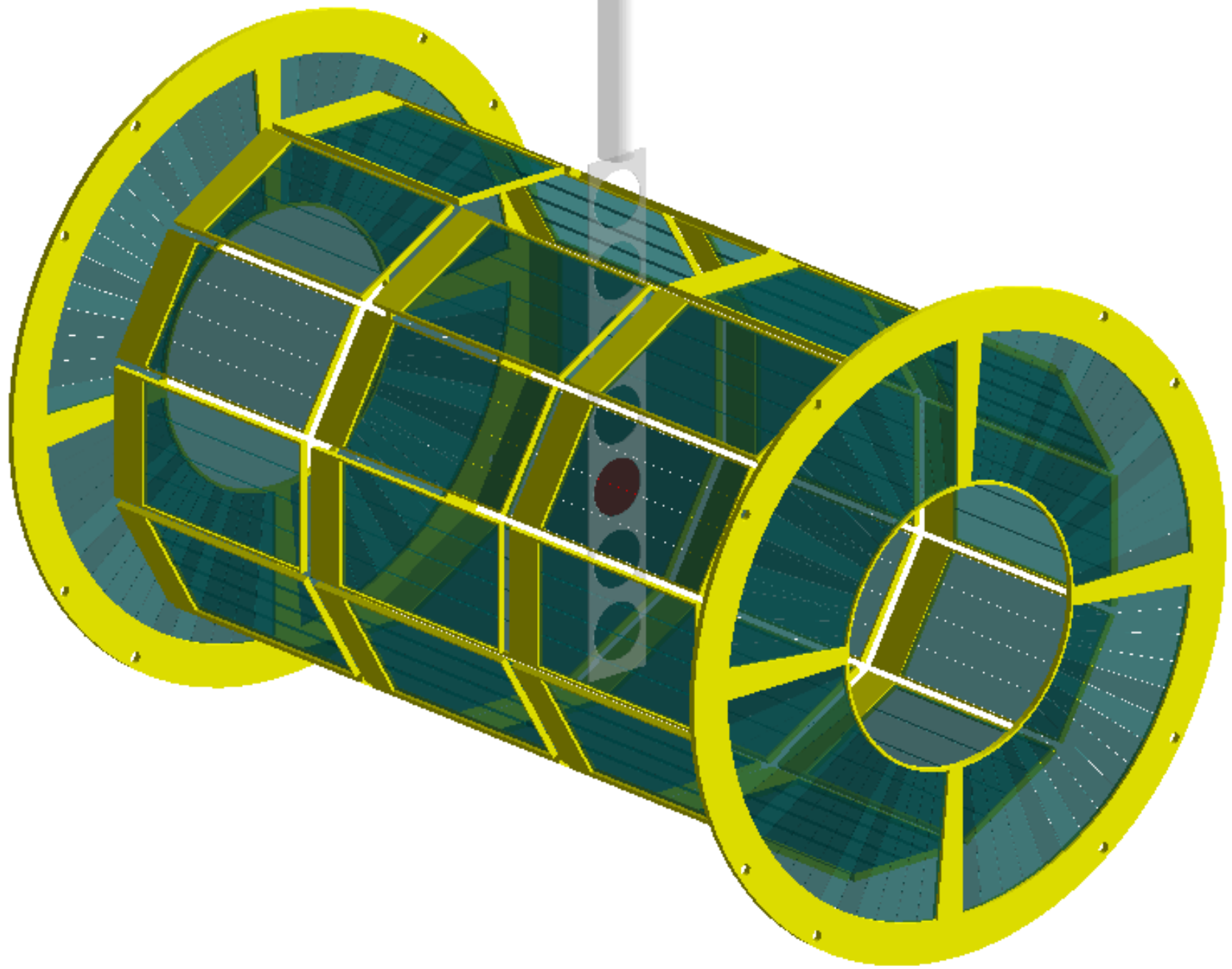
${}^{96}\text{Ru}(\gamma, \alpha){}^{92}\text{Mo}$  reaction

${}^{74}\text{Se}(\gamma, p){}^{73}\text{As}$  reaction

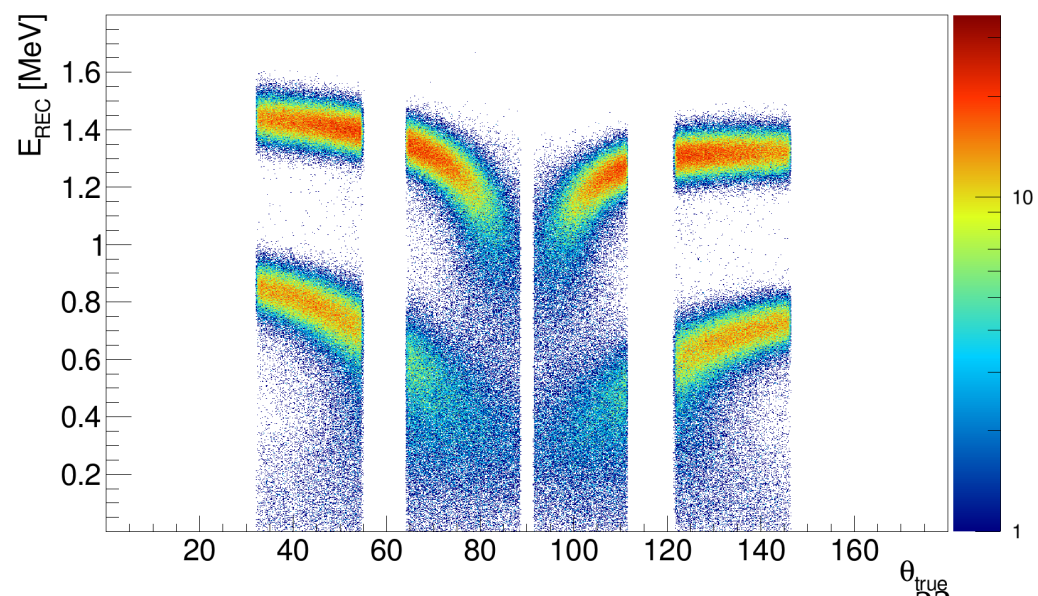
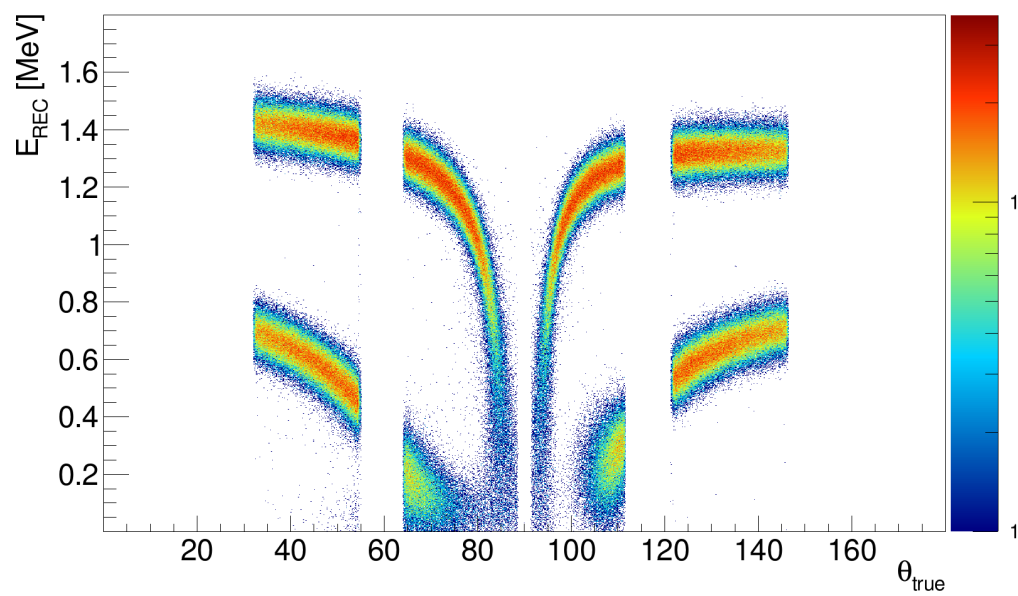
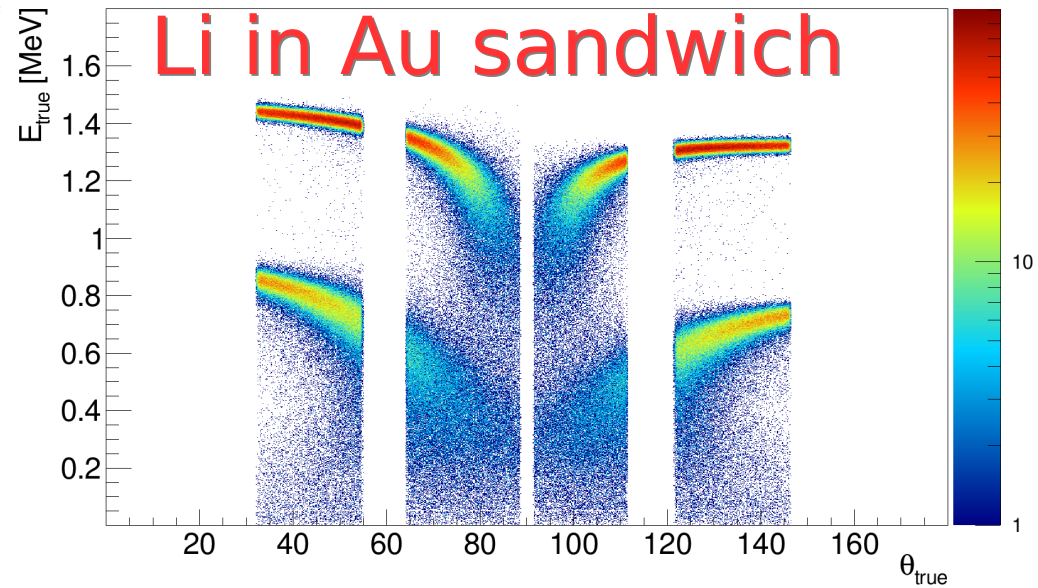
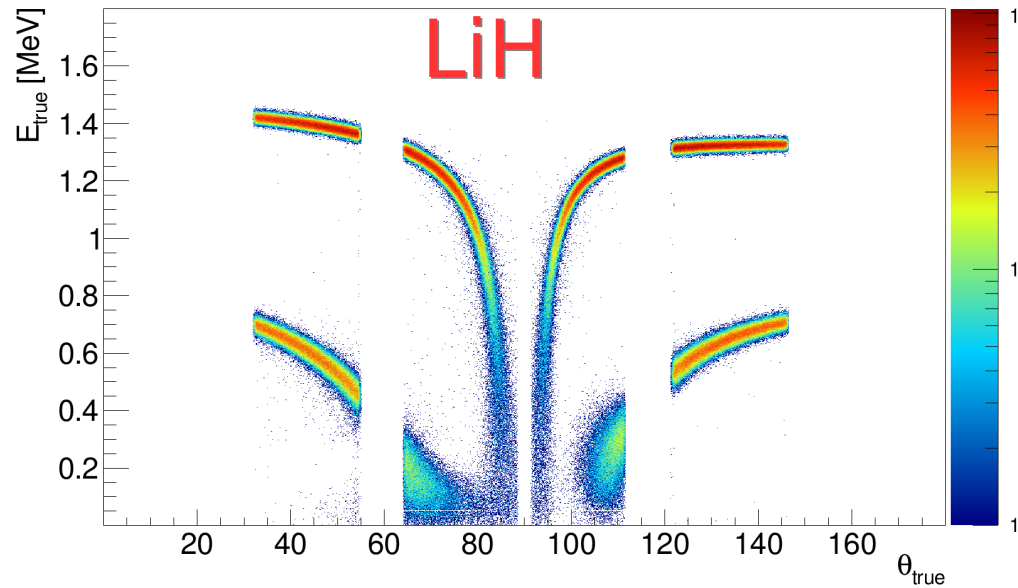
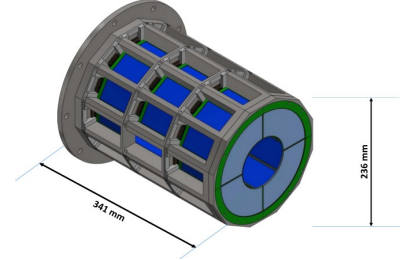
# Extreme Light Infrastructure Silicon Strip Array



- ♦ 3 rings of 12 position-sensitive X3 silicon-strip detectors (minus 1)  
Energy resolution (FWHM)  $\sim 0.3\%$   
Angular resolution 1 mm or  $\sim 0.4$  deg
- ♦ 2 end cap detectors made up of 4 QQQ3 DSSSD  
Energy resolution (FWHM)  $\sim 0.3\%$   
Angular resolution 3 mm or  $\sim 0.8$  deg



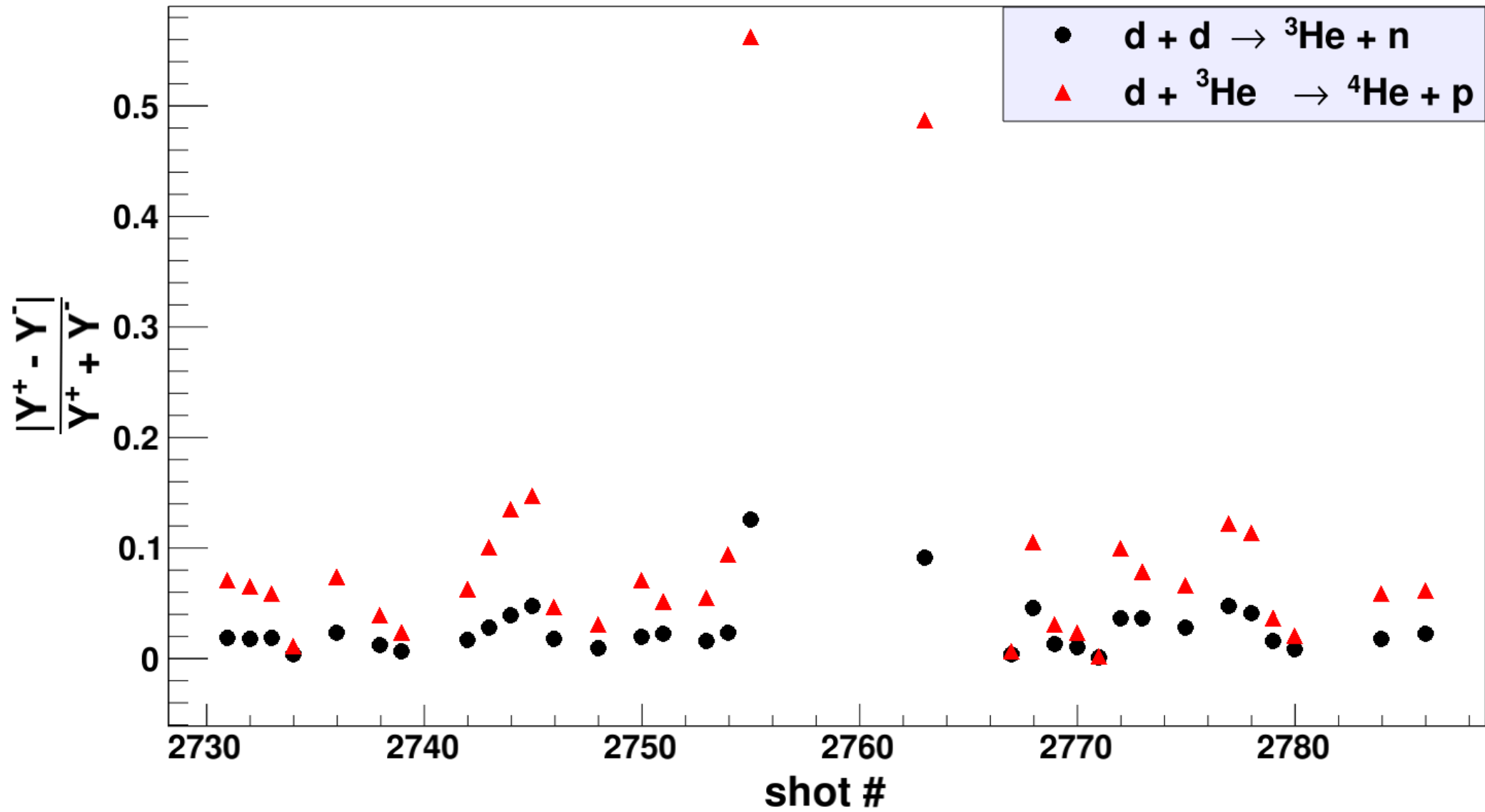
# ${}^7\text{Li}(\gamma, t){}^4\text{He}$ @ 5 MeV



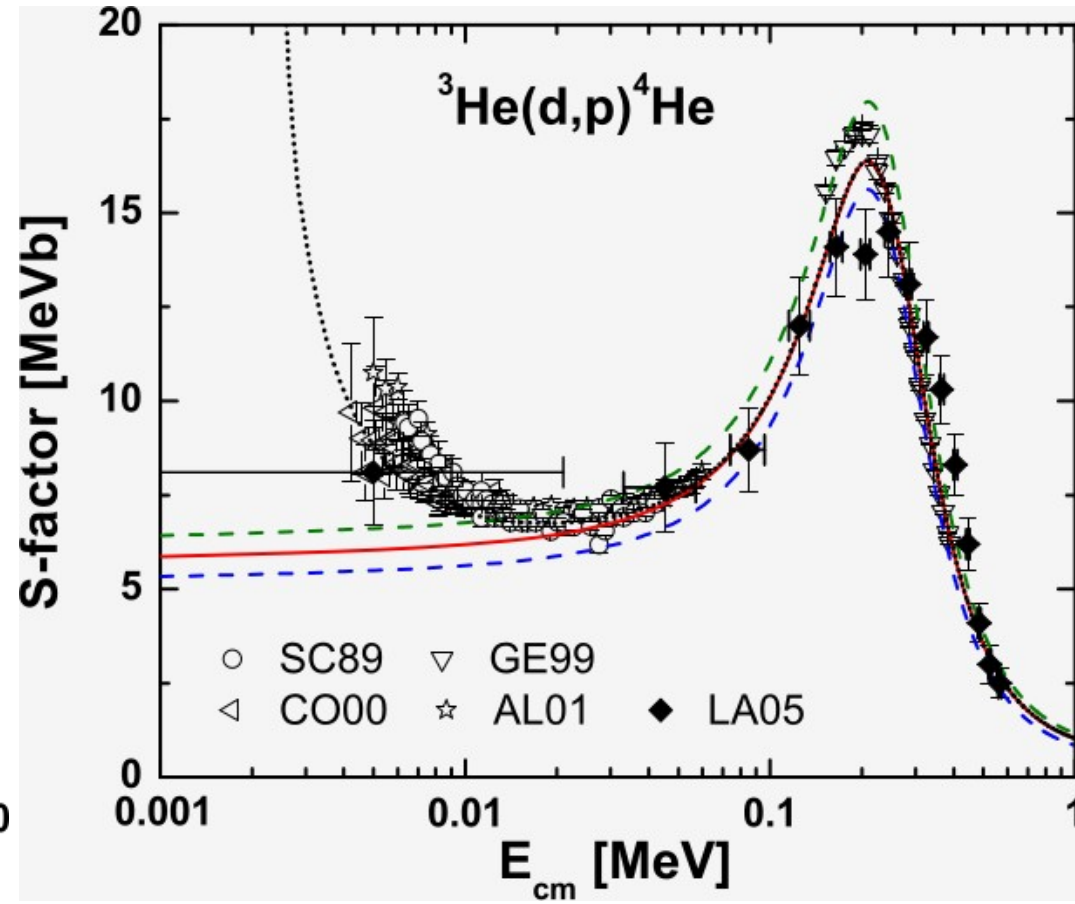
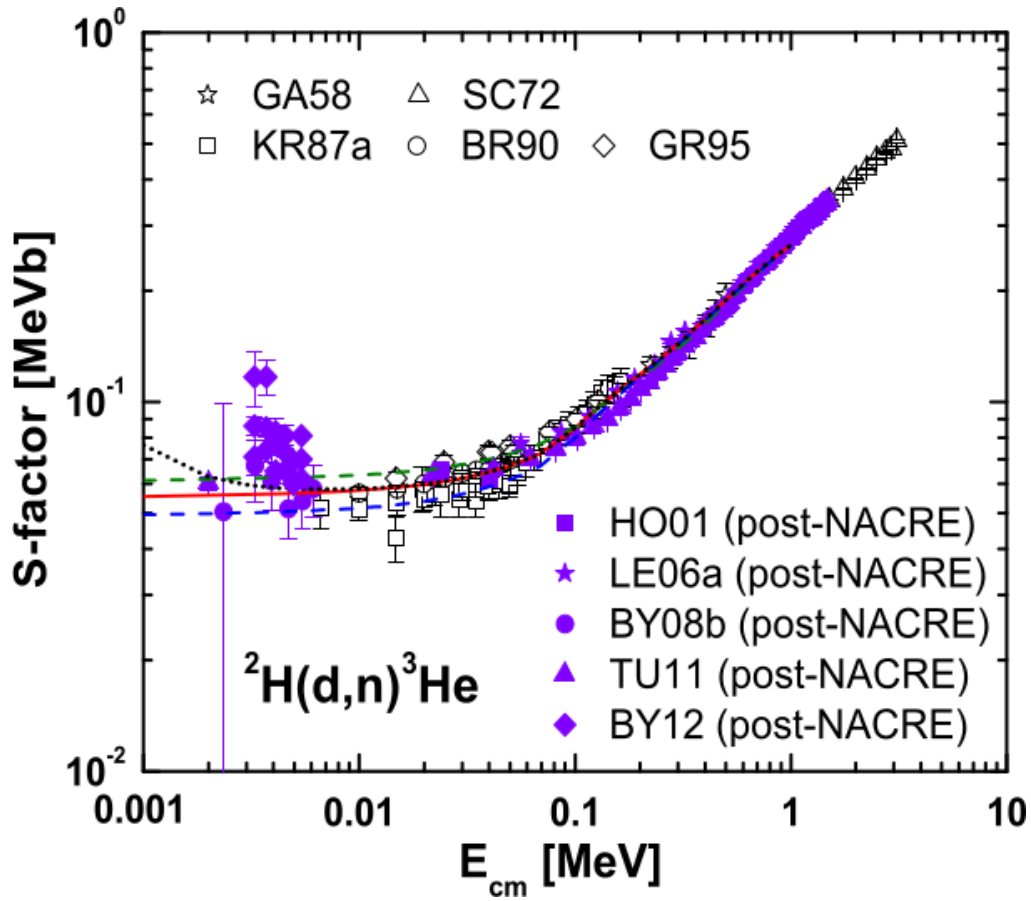
# Conclusions



# Backup slides?



# NACRE - II



$$Y = \frac{\rho_1 \int \frac{dN}{dE} S(E) \exp(-2\pi\eta(E)) v \tau dE}{1 + \delta_{12}}$$

# Physical Properties

Under solid (grey), liquid (blue) and vapor states (white) along the equilibrium curves

## GENERAL PROPERTIES

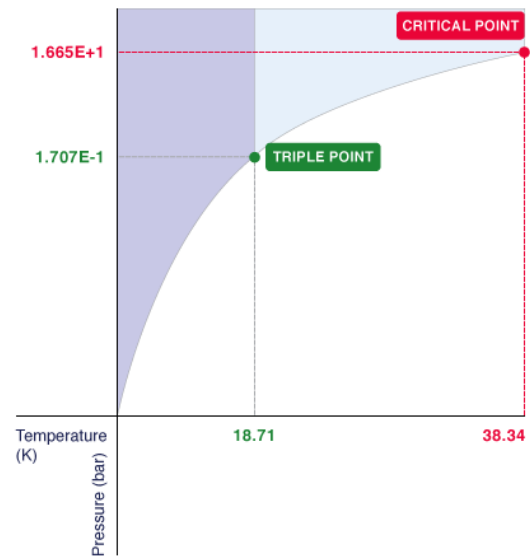
## SOLID PHASE

## LIQUID PHASE

## GAS PHASE

(P)  log(P)

DOWNLOAD



Molecular weight	4.028 g/mol
Content in air	/
<b>Critical Point</b>	
Temperature	- 234.81 °C
Pressure	16.653 bar
Density	69.797 kg/m <sup>3</sup>
<b>Triple Point</b>	
Temperature	- 254.44 °C

D<sub>2</sub>  
Deuterium

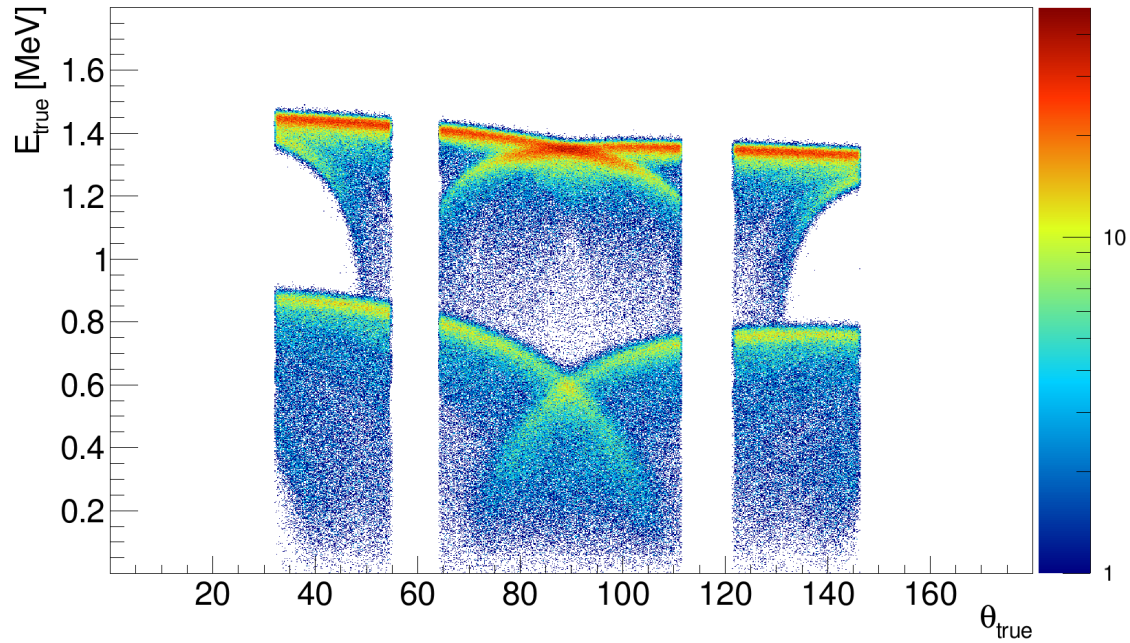
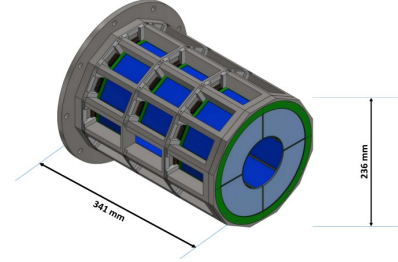


CAS Number 7782-39-0  
UN1957 (gas)

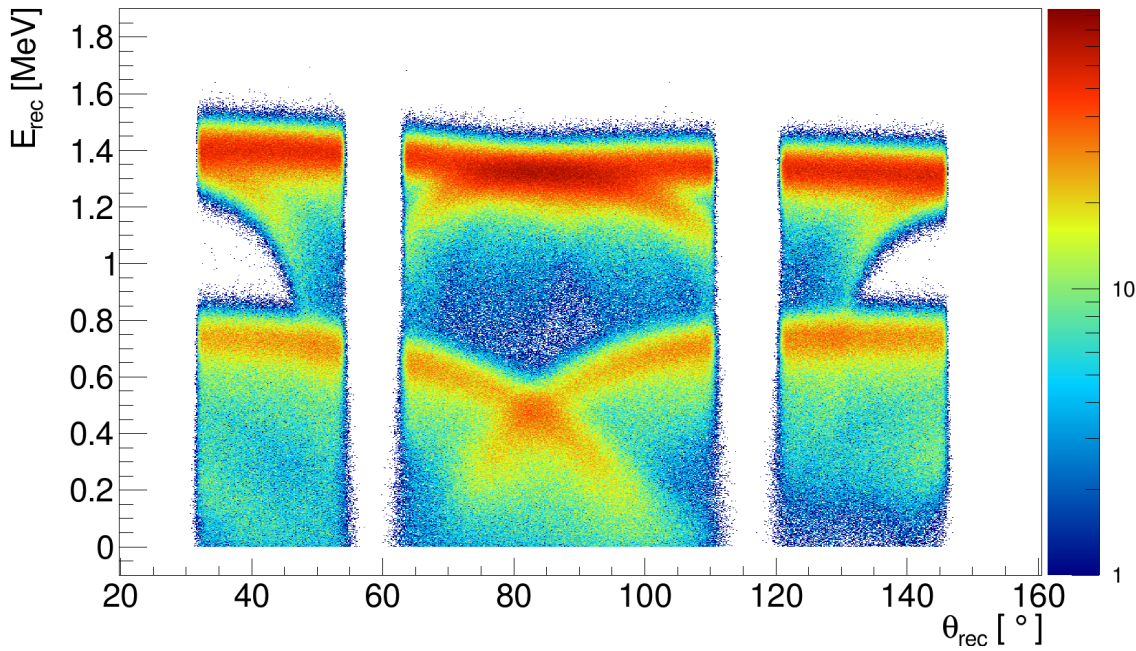
- Molecule
- Properties
- Liquid / Gas Volume
- Applications
- Safety

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Highlight All Match Case Whole Words 1 of 1 match Reached end of page, continued from top



By rotating the target ( $35^\circ$  here) we can minimize (remove) the effect of blind kinematics due to the straggling of ejectiles inside target



Analysis ongoing.