# Correlations between charge radii, E0 transitions and M1 strength

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Correlations between

Charge radii and EO transitions (Wood et al.) Charge radii and summed M1 strength (Heyde et al.) EO transitions and summed M1 strength Application in the rare-earth nuclei

2PTN-9, Padua, May 2018

#### Coexistence or collective?

Origin of EO transitions in nuclei:

Mixing of coexisting configurations with different shapes (Heyde & Wood);

Between  $\beta$ -vibrational states in the geometric collective model (Reiner).

In a geometric framework EO strength should rise in the transition from spherical to deformed.

## Operators in the IBM

The charge radius operator:

$$\hat{T}(r^2) = \langle r^2 \rangle_{\text{core}} + \alpha N_{\text{b}} + \frac{\eta}{N_{\text{b}}} \hat{n}_d$$

The EO operator:

$$\hat{T}(E0) = \left(e_{\rm n}N + e_{\rm p}Z\right)\frac{\eta}{N_{\rm b}}\hat{n}_d$$

The M1 operator:

$$\hat{T}(\mathrm{M1}) = \sqrt{\frac{3}{4\pi}} \left( g_{\nu} \hat{L}_{\nu} + g_{\pi} \hat{L}_{\pi} \right)$$

F. Iachello and A. Arima, The Interacting Boson Model

# Application to rare-earth nuclei

Application to even-even nuclei with Z=58-74. Procedure:

Determine IBM hamiltonian from spectra with special care to the spherical-to-deformed transitional region. Determine coefficients  $\alpha$  and  $\eta$  in T(r<sup>2</sup>) from isotope and isomer shifts.

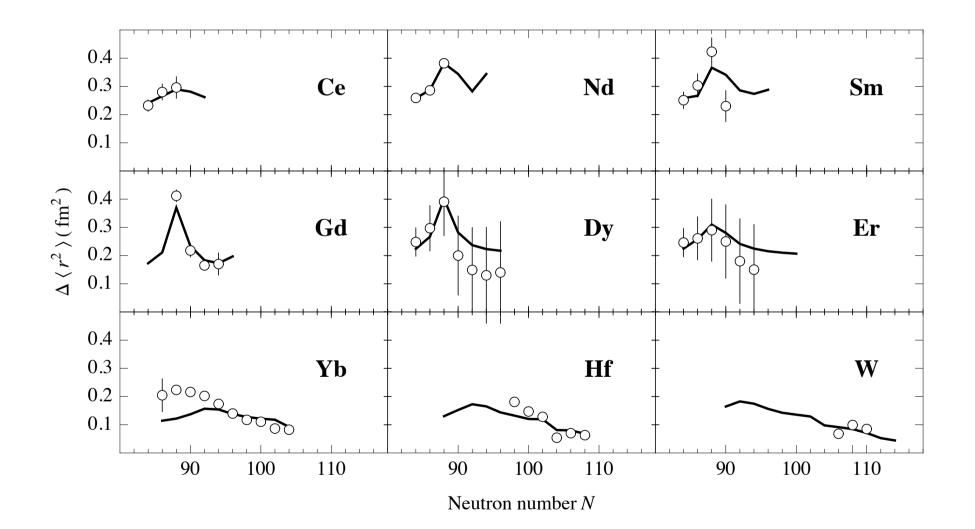
Calculate  $\rho$ (EO) values.

## Isotope shifts

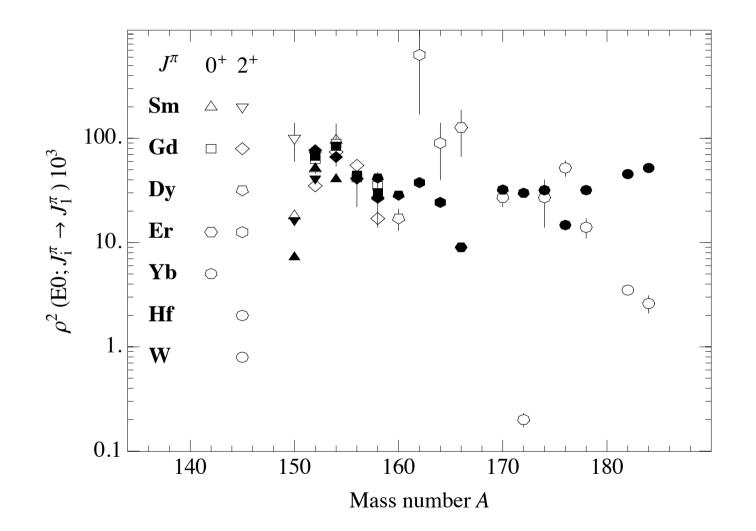
Isotopes shifts depend on the coefficients  $\alpha$  and  $\eta$ :

$$\Delta \left\langle r^2 \right\rangle \equiv \left\langle r^2 \right\rangle_{0_1^+}^{(A+2)} - \left\langle r^2 \right\rangle_{0_1^+}^{(A)} = \left| \alpha \right| + \frac{\eta}{N_{\rm b}} \left( \left\langle \hat{n}_d \right\rangle_{0_1^+}^{(A+2)} - \left\langle \hat{n}_d \right\rangle_{0_1^+}^{(A)} \right)$$

## Isotope shifts



$$ho^2$$
 values



# $\rho^{\rm 2}$ values

Isotope	Transition	J 0	$ ho^2(E0)  imes 10^3$				
	$740 \rightarrow 0$		Th1 <sup>a</sup>	Th2 <sup>b</sup>	Th3 <sup>c</sup>	Expt. <sup>d</sup>	
						18	2
	$1046 \rightarrow 334$	2	16	13		100	40
<sup>152</sup> Sm	$685 \rightarrow 0$	0	52	52	72	51	5
	$811 \rightarrow 122$	2	41	41	77	69	6
	$1023 \rightarrow 366$	4	29	29	84	88	14
	$1083 \rightarrow 0$	0	2	2		0.7	0.4
	$1083 \rightarrow 685$	0	47	47		22	9
<sup>154</sup> Sm	$1099 \rightarrow 0$	0	41	49		96	42
<sup>152</sup> Gd	$615 \rightarrow 0$	0	68	68		63	14
	$931 \rightarrow 344$	2	77	77		35	3
<sup>154</sup> Gd	$681 \rightarrow 0$	0	84	102		89	17
	$815 \rightarrow 123$	2	66	80		74	9
	$1061 \rightarrow 361$	4	38	46		70	7
<sup>156</sup> Gd	$1049 \rightarrow 0$	0	44	64		42	20
	$1129 \rightarrow 89$	2	41	59		55	5
<sup>158</sup> Gd	$1452 \rightarrow 0$	0	30	51		35	12
	$1517 \rightarrow 79$	2	27	45		17	3
<sup>158</sup> Dy	$1086 \rightarrow 99$	2	42	70		27	12
<sup>160</sup> Dy	$1350 \rightarrow 87$	2	28	56		17	4
<sup>162</sup> Er	$1171 \rightarrow 102$	2	38	64		630	460
<sup>164</sup> Er	$1484 \rightarrow 91$	2	24	48		90	50
<sup>166</sup> Er	$1460 \rightarrow 0$	0	9	20		127	60
<sup>170</sup> Yb	$1229 \rightarrow 0$	0	32	72		27	5
<sup>172</sup> Yb	$1405 \rightarrow 0$	0	30	76		0.2	0.03
<sup>174</sup> Hf	$900 \rightarrow 91$	2	32	71		27	13
<sup>176</sup> Hf	$1227 \rightarrow 89$	2	15	38		52	9
<sup>178</sup> Hf	$1496 \rightarrow 93$	2	32	72		14	3
$^{182}W$	$1257 \rightarrow 100$	2	45	77		3.5	0.3
$^{184}W$	$1121 \rightarrow 111$	2	52	75		2.6	0.5

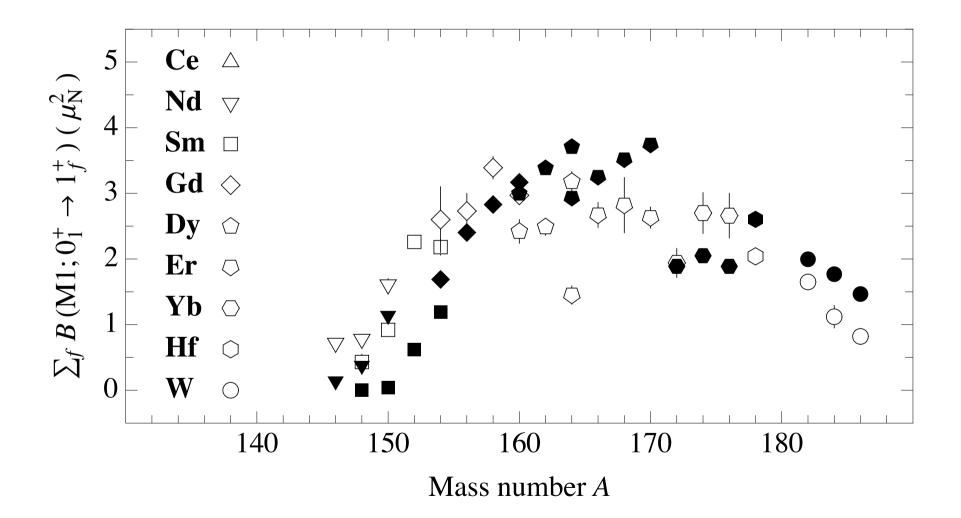
## Summed B(M1) strength

Ginocchio proved the following M1 sum rule:

$$\sum_{f} B\left(M1; 0_{1}^{+} \to 1_{f}^{+}\right) = \frac{3}{4\pi} \left(g_{\nu} - g_{\pi}\right)^{2} \frac{6N_{\nu}N_{\pi}}{N_{b}\left(N_{b} - 1\right)} \left\langle \hat{n}_{d} \right\rangle_{0_{1}^{+}}$$

Summed M1 strength to the scissors state is known in many rare-earth nuclei.

## Summed B(M1) strength



N. Pietralla et al., Phys. Rev. C 52 (1995) 2317(R)

## Correlation S(M1)-<r2>

Rewrite expressions for  $\langle r^2 \rangle$  and S(M1):

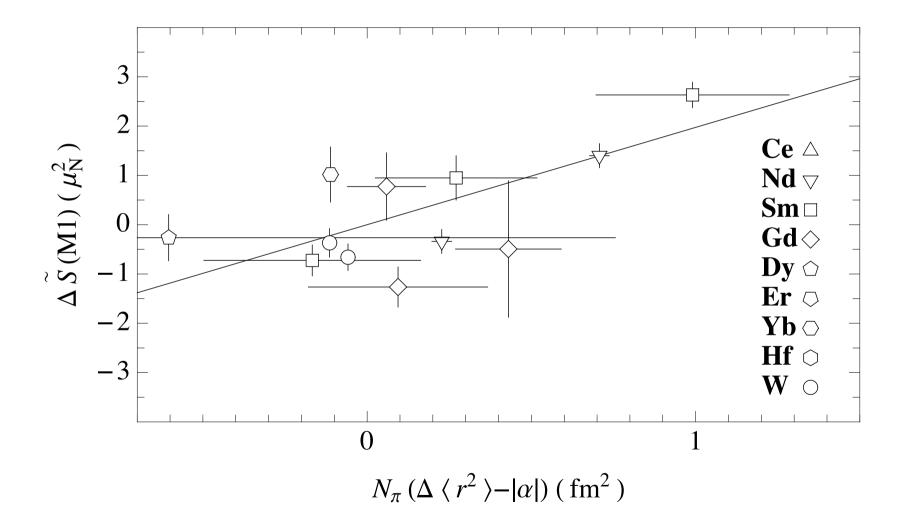
$$\begin{split} \Delta \left\langle r^2 \right\rangle &- \left| \alpha \right| = \frac{\eta}{N_{\rm b}} \left( \left\langle \hat{n}_d \right\rangle_{0_1^+}^{(A+2)} - \left\langle \hat{n}_d \right\rangle_{0_1^+}^{(A)} \right) \\ \tilde{S}\left( \mathbf{M} \mathbf{1} \right) &= \frac{N_{\rm b} - 1}{N_{\nu}} \sum_f B\left( \mathbf{M} \mathbf{1}; \mathbf{0}_1^+ \longrightarrow \mathbf{1}_f^+ \right) = \frac{9}{2\pi} \left( g_\nu - g_\pi \right)^2 \frac{N_\pi}{N_{\rm b}} \left\langle \hat{n}_d \right\rangle_{\mathbf{0}_1^+} \end{split}$$

We obtain the relation

$$\Delta \tilde{S}(M1) \equiv \tilde{S}(M1)^{(A+2)} - \tilde{S}(M1)^{(A)}$$
$$= \frac{9}{2\pi} \frac{(g_v - g_\pi)^2}{\eta} N_\pi (\Delta \langle r^2 \rangle - |\alpha|)$$

K. Heyde et al., Phys. Lett. B 312 (1993) 267

# Correlation $\Delta S(M1) - \Delta < r^2 >$



## Correlation $S(M1) - \rho(E0)$

In well-deformed nuclei [SU(3)]:  $\langle 0_{1}^{+} | \hat{n}_{d} | 0_{1}^{+} \rangle = \frac{4N_{b}(N_{b}-1)}{3(2N_{b}-1)}$   $|\langle 0_{\beta}^{+} | \hat{n}_{d} | 0_{1}^{+} \rangle| = \left[\frac{8(N_{b}-1)^{2}N_{b}(2N_{b}+1)}{9(2N_{b}-3)(2N_{b}-1)^{2}}\right]^{1/2}$ We obtain the relation (for large  $N_{b}$ )

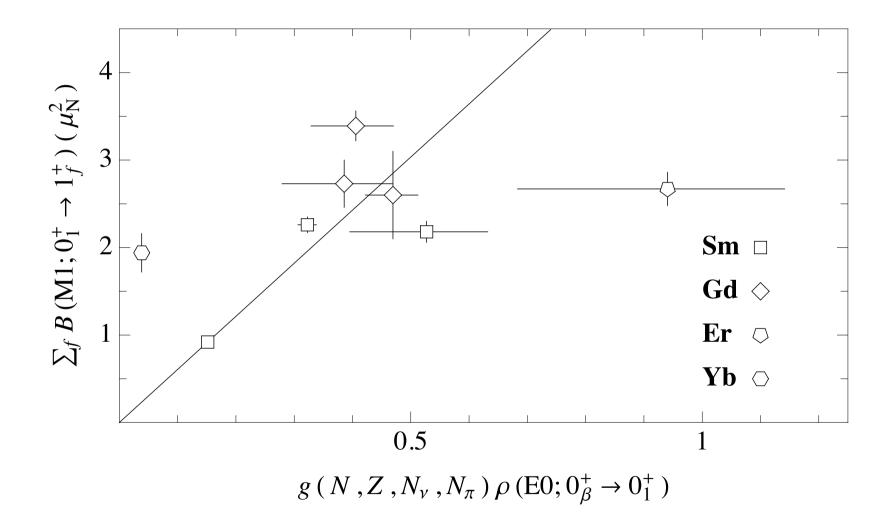
$$B(M1;0_{1}^{+} \to 1_{1}^{+}) \approx \frac{9}{\pi} (g_{v} - g_{\pi})^{2} \frac{r_{0}^{2}}{\eta} g(N,Z,N_{v},N_{\pi}) \rho(E0;0_{\beta}^{+} \to 0_{1}^{+})$$

with

$$g(N,Z,N_{v},N_{\pi}) = \frac{e(N+Z)^{2/3}}{e_{n}N + e_{p}Z} \frac{N_{v}N_{\pi}}{\sqrt{2N_{b}}}$$

P. Van Isacker, Nucl. Data Sheets 120 (2014) 119

## Correlation $S(M1) - \rho(E0)$



## Conclusions

Consistent treatment of charge radii and EO transitions assuming the same effective charges. An additional correlation exists with summed M1 strength, which can be related to charge radii and EO transitions.

Outlook:

 $S(M1)-\rho$  (EO) correlation for transitional nuclei. Need for  $\Delta < r^2 >$  through shape transition.