

Search for
 γ -ray
emission in
the NLS1
galaxy IRAS
20181-2244

Michele
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Search for γ -ray emission in the NLS1 galaxy IRAS 20181-2244

Michele Frezzato

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Supervisor: Stefano Ciroi, Francesco Di Mille

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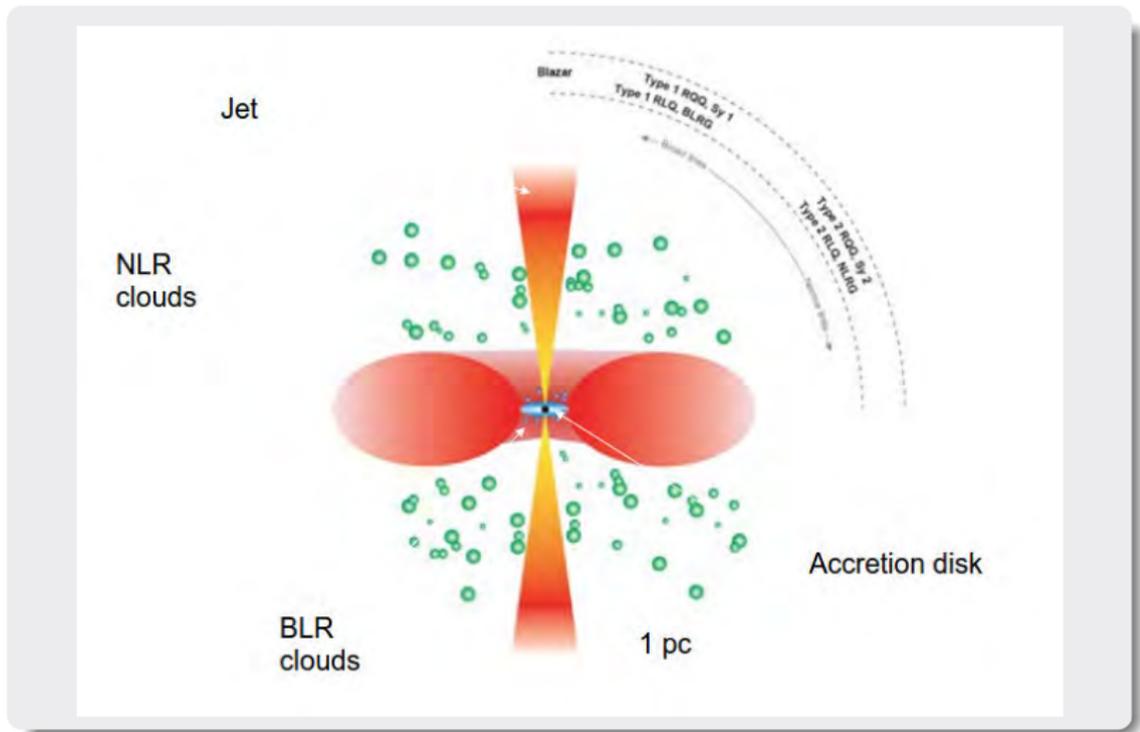
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Narrow Line Seyfert 1 are classified according to their unusual spectral properties [Osterbrock+87]:

- Active Galactic Nuclei
- Type 1 AGN: $H\beta/[O III] < 3$ and presence of strong Fe II multiplets
- However the FWHM of $H\beta$ is lower than 2000 km s^{-1}

Have been subdivided into

- Radio Loud NLS1
- Radio Quiet NLS1

On the basis of the Radio Loudness parameter [Kellerman+ 89]:

$$R = \frac{F_{5\text{GHz}}}{F_B}$$

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Summary

For many years it was believed that radio-loud AGN were hosted only by elliptical galaxies [Laor00, Chiaberge11].

Radio-loudness is commonly interpreted as the presence of powerful relativistic jet, usually found in systems harbouring a very massive black hole.

Due to the $M_{BH}-\sigma_*$ relation they are found more commonly found in elliptical galaxies, hence leading likely to a selection effect.

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γ detection in NLS1

FERMI-LAT led to the discovery of γ -ray emission coming from a flat-spectrum radio-loud narrow-line Seyfert 1 [Abdo+09a, Abdo+09c]

Another clue of the presence of relativistic jets in NLS1.

There may be a correlation between the environment and the presence of jets in NLS1.

In particular radio-loudness can be the byproduct of merging events [Chiaberge+15]

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- It is a RLNLS1 likely hosted in a spiral galaxy
- Preliminary analysis of IR photometry seem to indicate a merger event
- Previous (unpublished) FERMI PASS7 analysis of the target found a tentative γ detection ($9 < TS < 25$)

It's worth a try!

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Summary

All the available photon data available of the source were retrieved, along with the spacecraft data.

A region of interest of 30° around the source was then selected with the task `gtselect`.

Only PASS 8 SOURCE class events were selected (`evclass = 128`), taking both front and back events (`evtype=3`).

A maximum zentith angle of 90° was set in order to remove earth limb contamination

The good time intervals (GTI) were obtained from the spacecraft data with the task `gtmktime`, constraining the data quality to be > 0

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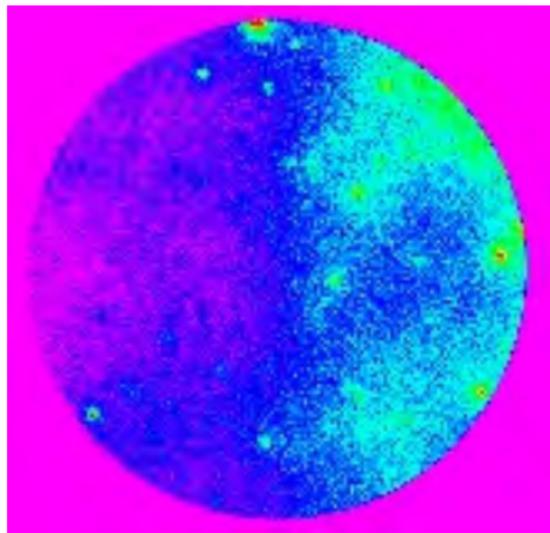
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Count map

The count map of the ROI was created using `gtbin` with a scale of 0.2 degrees/pixel.



Similarly the counts cube was obtained, composed of 37 logarithmically spaced energy bins.

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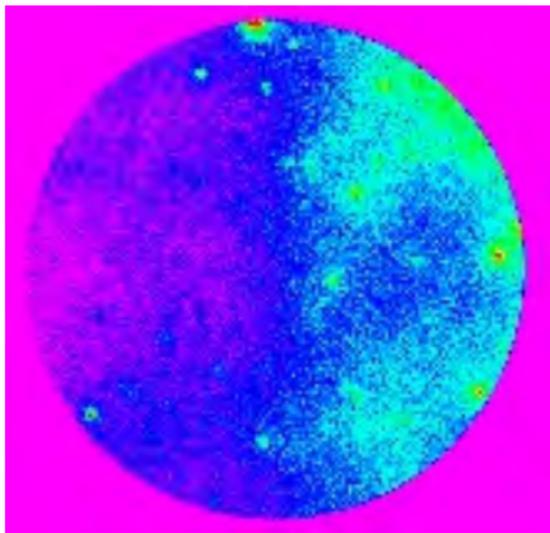
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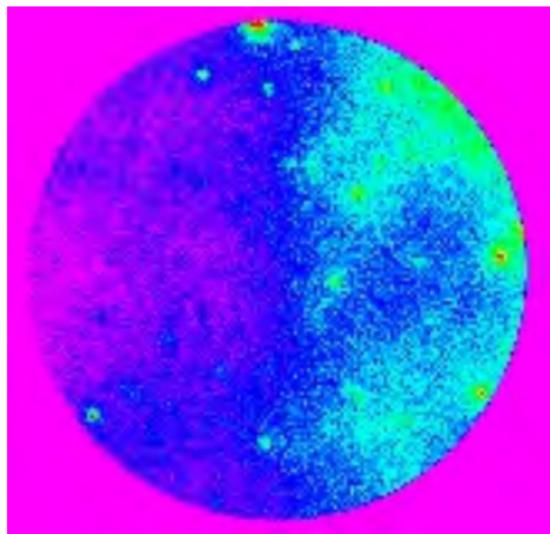
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Summary

The source model is a xml file containing the supposed location and spectral shape of the sources in the ROI. It was created using `Make3FGLxml.py`, which add the sources from the 3FGL catalog.

Two different models, A and B, were chosen for the analysis:

- A: IRAS 20181-2244 is modeled as a power law in the center of the ROI with index and normalization left as free parameters
- B: no additional source was added

`Make3FGLxml.py` takes into account all the 3FGL sources in the ROI as well as in 10 additional degrees around it.

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Last steps of the analysis:
Spacecraft data \rightarrow Calculate livetimes - `gtltcube`



exposure map - `gtexpcube2`



source map - `gtsrcmaps`



likelihood - `gtlike (NEWMINUIT)`

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Model A

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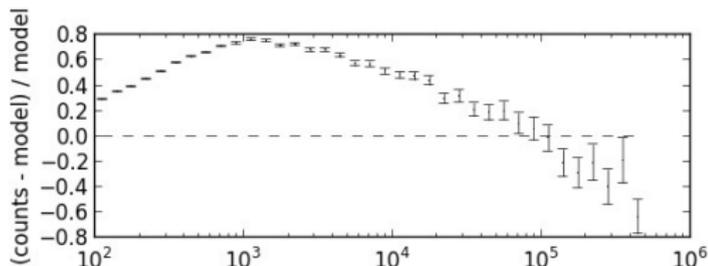
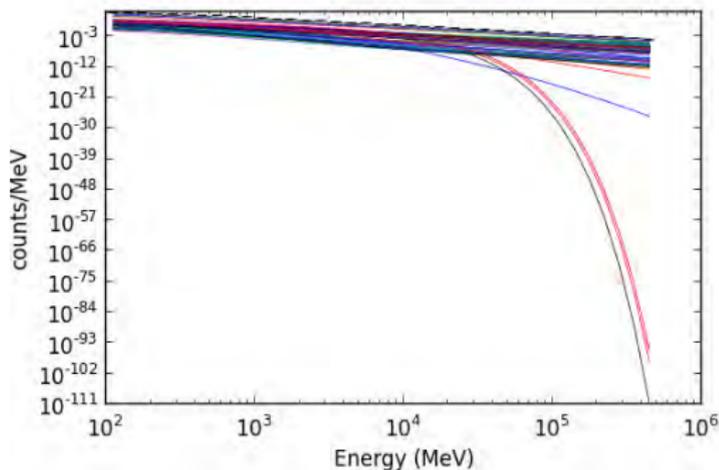
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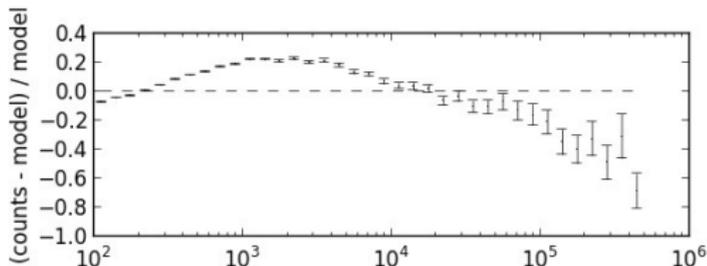
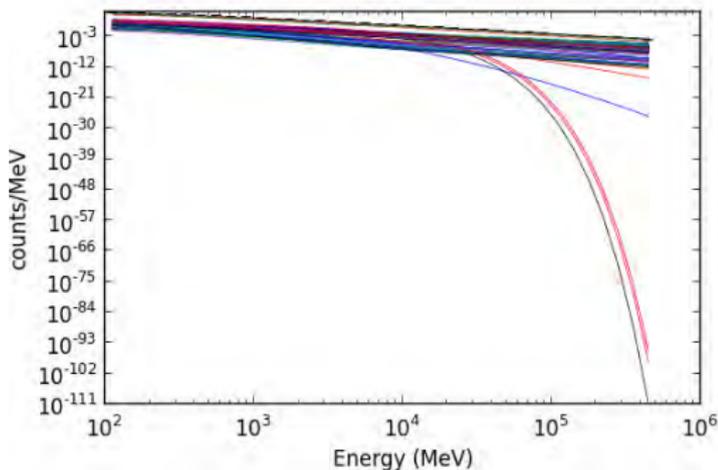
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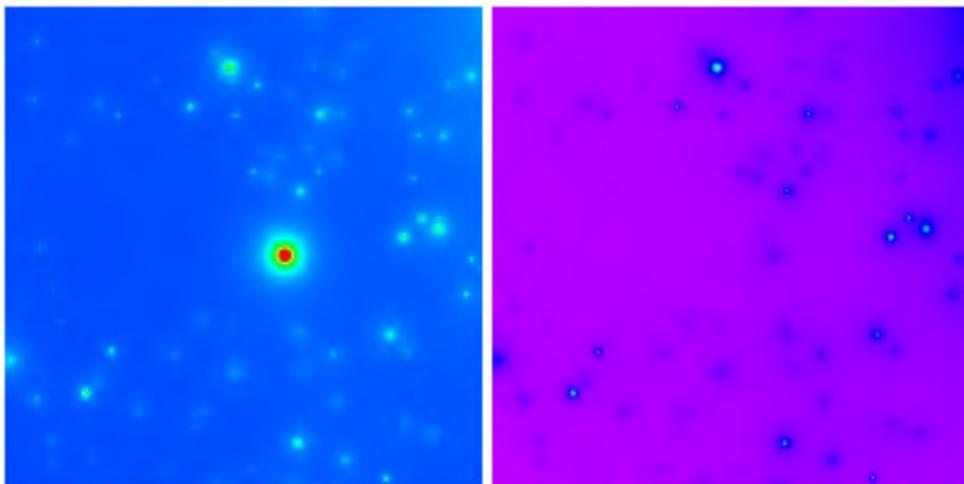
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- The results of the binned analysis seem to suggest IRAS20181 has no γ -ray emission
- Tweaking the source model to achieve convergence may help
- Proceed with the unbinned analysis on monthly time intervals

However...

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Frezzato

Introduction

The Unified
Model of AGN
Narrow Line
Seyfert 1
IRAS
20181-2244

Data Reduction

Data retrieval
and selection
Count map
Source Model
Binned
likelihood

Summary

