

# The hadronic vacuum polarisation contribution to the muon $g - 2$

Alex Keshavarzi

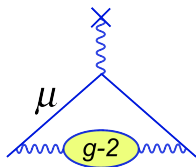


(in collaboration with Daisuke Nomura & Thomas Teubner [KNT17])

LFC17: Old and New Strong Interactions from LHC to Future Colliders

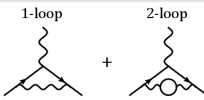


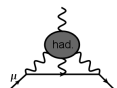
ECT, Trento, Italy

14<sup>th</sup> September 2017



# The Standard Model (SM) contributions to $a_\mu$

$$a_\mu^{\text{SM}} = a_\mu^{\text{QED}} + a_\mu^{\text{EW}} + a_\mu^{\text{had, VP}} + a_\mu^{\text{had, LbL}} + a_\mu^{\text{BSM??}}$$

|             |   |  |  |
|-------------|---|--|--|
| <b>QED</b>  | <p>1-loop      2-loop</p>  | <p>Known to <b>five-loop</b><br/>(12,672 diagrams)</p>                     | <p>99.99% of <math>a_\mu^{\text{SM}}</math>    0.001% of <math>\delta a_\mu^{\text{SM}}</math></p>     |
| <b>EW</b>   |                            | <p>Known to <b>two-loop</b><br/>(with <math>m_H</math> known)</p>          | <p>0.0001% of <math>a_\mu^{\text{SM}}</math>    0.2% of <math>\delta a_\mu^{\text{SM}}</math></p>      |
| <b>HVP</b>  |                            | <p><b>Non-perturbative</b><br/>(experimental input)</p>                    | <p>0.006% of <math>a_\mu^{\text{SM}}</math>    <b>73%</b> of <math>\delta a_\mu^{\text{SM}}</math></p> |
| <b>HLbL</b> |                            | <p><b>Non-perturbative</b><br/>(data input + model/lattice)</p>            | <p>0.0001% of <math>a_\mu^{\text{SM}}</math>    27% of <math>\delta a_\mu^{\text{SM}}</math></p>       |
| <b>BSM</b>  | <p>???????</p>  | <p><math>\Rightarrow</math> 0.0002% of <math>a_\mu^{\text{exp}}</math></p> |  |

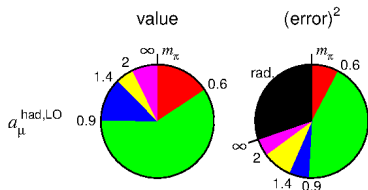
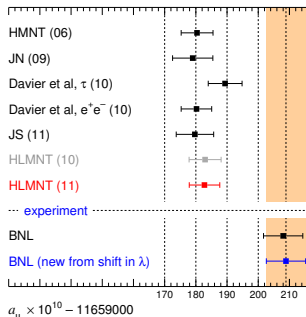
# The previous analysis... [HLMNT(11), J. Phys. G38 (2011), 085003]

|                              |  |  |
|------------------------------|--|--|
| <b>QED</b> contribution      | 11 658 471.808 (0.015) $\times 10^{-10}$               | Kinoshita & Nio, Aoyama et al              |
| <b>EW</b> contribution       | 15.4 (0.2) $\times 10^{-10}$                           | Czarnecki et al                            |
| <b>Hadronic</b> contribution |  |  |
| <b>LO</b> hadronic           | 694.9 (4.3) $\times 10^{-10}$                          | HLMNT11                                    |
| <b>NLO</b> hadronic          | -9.8 (0.1) $\times 10^{-10}$                           | HLMNT11                                    |
| <b>light-by-light</b>        | 10.5 (2.6) $\times 10^{-10}$                           | Prades, de Rafael & Vainshtein             |
| <b>Theory TOTAL</b>          | <b>11 659 182.8 (4.9) <math>\times 10^{-10}</math></b> |  |
| <b>Experiment</b>            | <b>11 659 208.9 (6.3) <math>\times 10^{-10}</math></b> | world avg                                  |
| <b>Exp - Theory</b>          | <b>26.1 (8.0) <math>\times 10^{-10}</math></b>         | <b>3.3 <math>\sigma</math> discrepancy</b> |

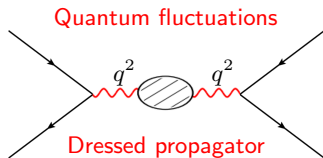
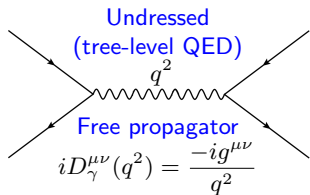
(Numbers taken from HLMNT11, arXiv:1105.3149)

- $a_\mu^{\text{had, LOVP}}$  still **dominated uncertainty**
- Potential for **improvement from experimental data and data combination**
- **New x4 accuracy measurements** planned from Fermilab and J-PARC
- ⇒ If  $a_\mu^{\text{SM}}$  &  $a_\mu^{\text{EXP}}$  improve as planned...

**g-2 discrepancy  $> 7\sigma$ !**



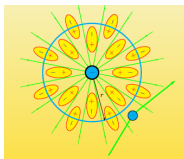
# The vacuum polarisation (VP)



$$iD_\gamma^{\mu\nu}(q^2) = \text{wavy line with blob} = \text{wavy line} + \text{wavy line with blob} + \text{wavy line with two blobs} + \dots = \frac{-ig^{\mu\nu}}{q^2(1 - \Pi_\gamma(q^2))}$$

(sum of all 1PI hadronic 'blobs'...)

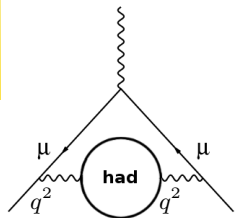
Virtual loop(s) = VP = charge screening =



⇒ Running QED coupling:

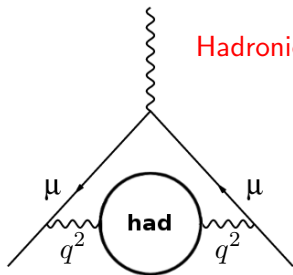
$$\alpha \rightarrow \alpha(q^2) = \frac{\alpha}{1 + 4\pi\alpha\text{Re}\Pi_\gamma(q^2)} = \frac{\alpha}{1 - \Delta\alpha(q^2)}$$

where  $\Delta\alpha(q^2) = \Delta\alpha_{\text{lep}}(q^2) + \Delta\alpha_{\text{had}}^{(5)}(q^2) + \Delta\alpha(q^2)_{\text{top}}$



# The hadronic vacuum polarisation (HVP)

Hadronic 'blob' contains sum of all hadronic states



$$\gamma^* \rightarrow \pi^0 \gamma \rightarrow \gamma^*$$

$$\gamma^* \rightarrow \pi^+ \pi^- \rightarrow \gamma^*$$

$$\gamma^* \rightarrow K^+ K^- \rightarrow \gamma^*$$

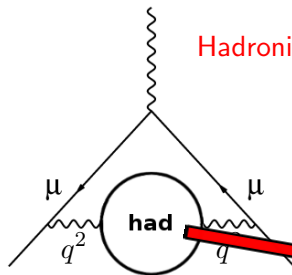
⋮

⋮

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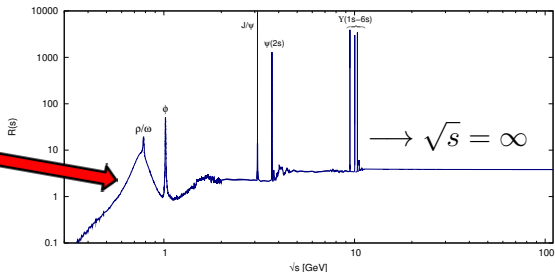
⋮

# The hadronic vacuum polarisation (HVP)



Hadronic 'blob' contains sum of all hadronic states

$$\begin{aligned} \gamma^* &\rightarrow \pi^0 \gamma \rightarrow \gamma^* \\ \gamma^* &\rightarrow \pi^+ \pi^- \rightarrow \gamma^* \\ \gamma^* &\rightarrow K^+ K^- \rightarrow \gamma^* \\ &\vdots \\ &\vdots \\ &\vdots \end{aligned}$$



→ Requires knowing entire hadronic spectrum  
( $s_{\text{th}} \leq \sqrt{s} \leq \infty$ )

⇒ Then, how do we calculate  $a_{\mu}^{\text{had,VP}}$ ?

# Dispersion relation (analyticity) and the optical theorem

So far, we know we need to relate

- The hadronic vacuum polarisation tensor,  $\Pi_{\text{had}}(q^2)$
- The hadronic spectrum

⇒ We do this in **two steps**...

## 1) Analyticity

Relate real part of  $\Pi(s)$  to its imaginary part

$$\text{had.} = \int \frac{ds}{\pi(s-q^2)} \text{Im} \text{had.}$$

$$\Pi_{\text{had}}(q^2) = \frac{q^2}{\pi} \int_{s_{\text{th}}}^{\infty} ds \frac{\text{Im}\Pi_{\text{had}}(s)}{s(s-q^2-i\epsilon)}$$

→ ...convolute with one loop contribution from coupling of virtual photon to muon...

$$a_{\mu}^{\text{had,LOVP}} = \frac{1}{4\pi^3} \int_{s_{\text{th}}}^{\infty} ds \sigma_{\text{had},(\gamma)}^0(s) K(s)$$

## 2) Optical theorem

Relate imaginary part of  $\Pi(s)$  to total  $e^+e^- \rightarrow$  hadrons cross section

$$2 \text{Im} \text{had.} = \sum_{\text{had.}} \int d\Phi \left| \text{had.} \right|^2$$

$$\text{Im}\Pi_{\text{had}}(s) = \left( \frac{s}{4\pi\alpha} \right) \sigma_{\text{had}}(s)$$

# HVP dispersion integral input

$$a_\mu^{\text{had,LOVP}} = \frac{1}{4\pi^3} \int_{s_{\text{th}}}^{\infty} ds \sigma_{\text{had},(\gamma)}^0(s) K(s)$$

**Important points** to note:

⇒ Can also define **dispersive integral** for NLOHVP with identical data input

⇒ Integral has **1/s dependence** ( $K(s) = \frac{m_\mu^2}{3s} \tilde{K}(s)$ )

⇒ Must **undress cross section of VP effects**:  $\sigma_{\text{had}} \rightarrow \sigma_{\text{had}}^0$

⇒ Cross section must **include FSR effects**:  $\sigma_{\text{had}}^0 \rightarrow \sigma_{\text{had},(\gamma)}^0$

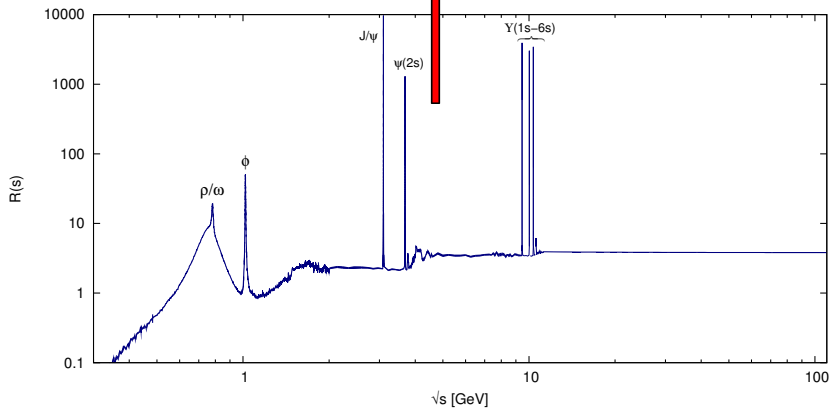
⇒ Can conventionally **define hadronic R-ratio**:  $R(s) = \frac{\sigma_{\text{had},(\gamma)}^0(s)}{\sigma_{\text{pt}}(s)} \equiv \frac{\sigma_{\text{had},(\gamma)}^0(s)}{4\pi\alpha^2/3s}$

$$a_\mu^{\text{had,LOVP}} = \frac{\alpha^2}{3\pi^2} \int_{s_{\text{th}}}^{\infty} \frac{ds}{s} R(s) K(s)$$



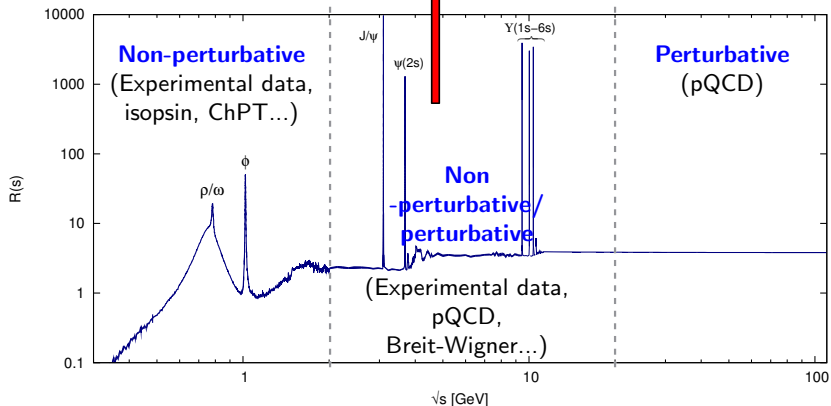
## Hadronic cross section input

$$a_{\mu}^{\text{had,LOVP}} = \frac{\alpha^2}{3\pi^2} \int_{s_{th}}^{\infty} \frac{ds}{s} R(s) K(s), \text{ where } K(s) = \frac{m_{\mu}^2}{3s} \tilde{K}(s)$$



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**Must build full hadronic cross section/ $R$ -ratio...**

# Building the hadronic $R$ -ratio

$$\underline{m_\pi \leq \sqrt{s} \leq 2 \text{ GeV}}$$

- Input **experimental hadronic cross section data\***
- **Combine all available data** in exclusive hadronic final states ( $\pi^+\pi^-$ ,  $K^+K^-$ , ...)
- **Sum  $\sim 35$  exclusive channels**
- Robust **treatment of experimental errors**
- Estimate missing data input (**isospin relations, ChPT...**)

$$\underline{2 \leq \sqrt{s} \leq 11.2 \text{ GeV}}$$

- Can **use experimental inclusive  $R$  data\*** or **pQCD**
- Must use **data at quark flavour thresholds**
- **Combine all available  $R$  data**
- Robust treatment of **experimental errors**
- **Include narrow resonances**

$$\underline{11.2 \leq \sqrt{s} < \infty \text{ GeV}}$$

- **Calculate  $R$  using pQCD (rhad)**

## \* $\sigma_{\text{had}}$ experiments

- KLOE
- BaBar
- SND
- CMD-(2/3)
- KEDR
- BESIII
- More old and new...

## Combining experimental data

### Question:

What are the main points of concern when correcting, combining and integrating experimental data to evaluate  $a_{\mu}^{\text{had, VP}}$ ?

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  - ...how to best **amalgamate large amounts of data from different experiments**

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  - ...the correct implementation of **correlated uncertainties** (statistical and systematic)

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  - ...finding a solution that is **free from bias**



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- ⇒ When **combining data**...
  - ...how to best amalgamate large amounts of data from different experiments
  - ...the correct implementation of correlated uncertainties (statistical and systematic)
  - ...finding a solution that is free from bias
- ⇒ The **reliability of the integral** and **error estimate**

# Combining experimental data

## Question:

What are the main points of concern when correcting, combining and integrating experimental data to evaluate  $a_{\mu}^{\text{had, VP}}$ ?

- ⇒ Radiative corrections of data and the corresponding error estimate
- ⇒ When **combining data**...
  - ...how to best amalgamate large amounts of data from different experiments
  - ...the correct implementation of correlated uncertainties (statistical and systematic)
  - ...finding a solution that is free from bias
- ⇒ The reliability of the integral and error estimate
- ⇒ The choices when **estimating unmeasured hadronic final states**

# Radiative corrections (for data that is not $\sigma_{\text{had},(\gamma)}^0$ )

## Vacuum polarisation corrections

⇒ Fully updated, self-consistent VP routine: [vp\_knt\_v3\_0]

→ Cross sections undressed with full photon propagator (must include imaginary part),  $\sigma_{\text{had}}^0(s) = \sigma_{\text{had}}(s) |1 - \Pi(s)|^2$

⇒ If correcting data, apply corresponding radiative correction uncertainty

→ Take  $\frac{1}{3}$  of total correction per channel as conservative extra uncertainty

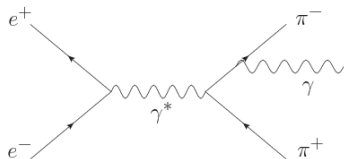
## Final state radiation corrections

⇒ For  $\pi^+\pi^-$ , include through sQED approximation

[Eur. Phys. J. C 24 (2002) 51, Eur. Phys. J. C 28 (2003) 261]

⇒ For higher multiplicity states, difficult to estimate correction

∴ Apply conservative uncertainty



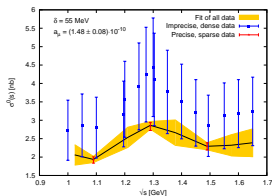
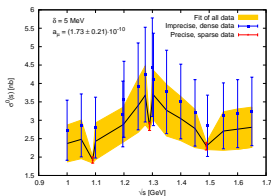
**Need new, more developed tools to increase precision here**

(e.g. - CARLOMAT 3.1 [Eur.Phys.J. C77 (2017) no.4, 254 ]?)

# Clustering data

⇒ Re-bin data into *clusters*

Better representation of data combination through adaptive clustering algorithm



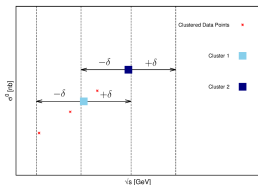
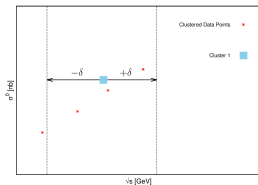
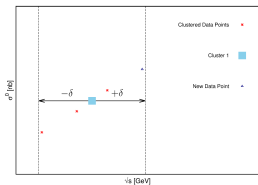
→ More and more data ⇒ risk of **over clustering**

⇒ loss of information on resonance

→ Scan cluster sizes for **optimum solution** (error,  $\chi^2$ , check by sight...)

⇒ Scanning/**sampling by varying bin widths**

→ Clustering algorithm now **adaptive to points at cluster boundaries**



# Correlation and covariance matrices

⇒ **Correlated data** beginning to **dominate** full data compilation...

→ Non-trivial, **energy dependent influence** on both **mean value and error estimate**

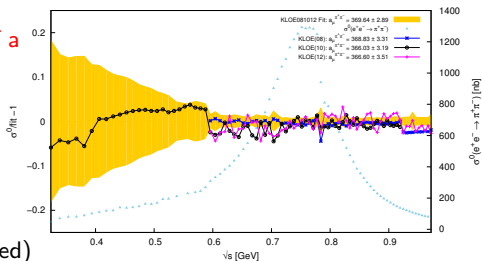
## KNT17 prescription

- Construct full covariance matrices for each channel & entire compilation  
⇒ **Framework available for inclusion of any and all inter-experimental correlations**
- If experiment does not provide matrices...  
→ Statistics occupy diagonal elements only  
→ Systematics are 100% correlated
- If experiment does provide matrices...  
→ **Matrices must satisfy properties of a covariance matrix**

e.g. - KLOE  $\pi^+\pi^-\gamma(\gamma)$  combination  
covariance matrices update

⇒ **Originally, NOT a positive semi-definite matrix:**

(Corrected in separate work to be published)



# Systematic bias and use of the data/covariance matrix

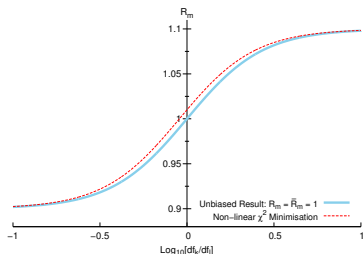
⇒ Data is re-binned using an **adaptive clustering algorithm**

⇒ Iterative fit of covariance matrix as defined by data → **D'Agostini bias**

[Nucl.Instrum.Meth. A346 (1994) 306-311]

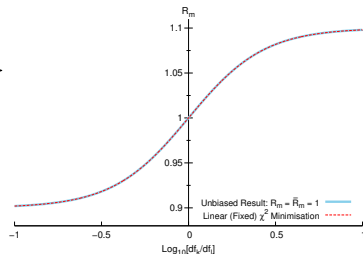
## HLMNT11

⇒ Non-linear  $\chi^2$  minimisation fitting  
nuisance parameters  
→ **Penalty trick bias**



## KNT17

⇒ **Fix the covariance matrix** in an  
iterative  $\chi^2$  minimisation  
→ **Free from bias**



**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**

# Linear $\chi^2$ minimisation

⇒ Redefine clusters to have **linear cross section**

→ **Fix covariance matrix with linear interpolants** at each iteration  
(extrapolate at boundary)

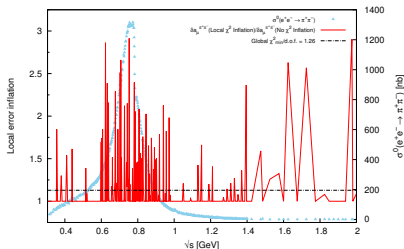
$$\chi^2 = \sum_{i=1}^{N_{\text{tot}}} \sum_{j=1}^{N_{\text{tot}}} (R_i^{(m)} - \mathcal{R}_m^i) \mathbf{C}^{-1}(i^{(m)}, j^{(n)}) (R_j^{(n)} - \mathcal{R}_n^j)$$

⇒ **Through correlations and linearisation**, result is the minimised solution of all neighbouring clusters

→ ...and **solution is the product of the influence of all correlated uncertainties**

⇒ The **flexibly of the fit** to vary due to the energy dependent, correlated uncertainties benefits the combination

→ ...and any data tensions are reflected in a **local  $\chi^2$  error inflation**

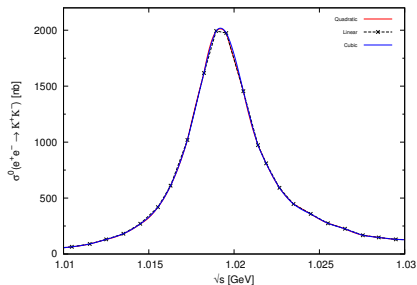
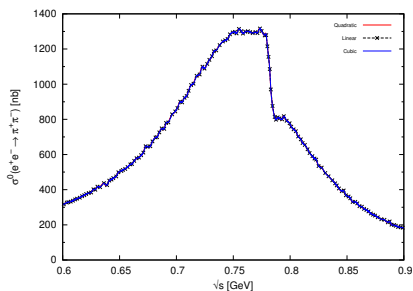


# Integration

⇒ Trapezoidal rule integral

→ Consistency with linear cluster definition

→ High data population ∴ **Accurate estimate from linear integral**



→ Higher order polynomial integrals give **(at maximum)** differences of  $\sim 10\%$  of error

⇒ Estimates of error non-trivial at **integral borders**

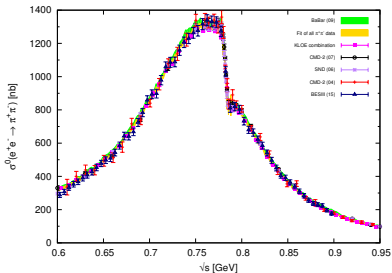
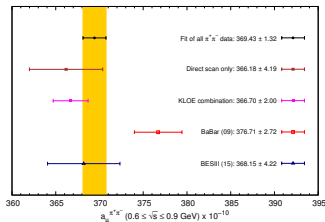
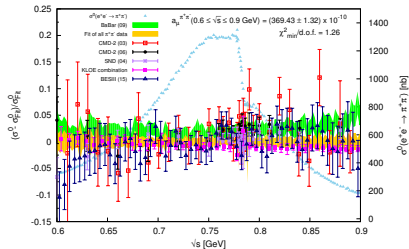
→ **Extrapolate/interpolate covariance matrices**



# $\pi^+\pi^-$ channel [preliminary]

⇒ Large improvement for  $2\pi$  estimate

→ BESIII [Phys.Lett. B753 (2016) 629-638] and KLOE combination provide **downward influence** to mean value



⇒ Correlated & experimentally corrected  $\sigma_{\pi\pi(\gamma)}^0$  data now entirely **dominant**

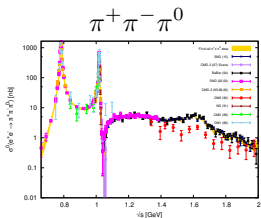
$$a_{\mu}^{\pi^+\pi^-} (0.305 \leq \sqrt{s} \leq 2.00 \text{ GeV}):$$

$$\text{HLMNT11: } 505.77 \pm 3.09$$

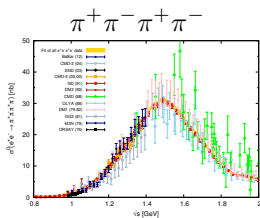
$$\text{KNT17: } 502.85 \pm 1.93$$

(no radiative correction uncertainties)

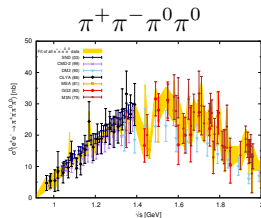
## Other notable exclusive channels



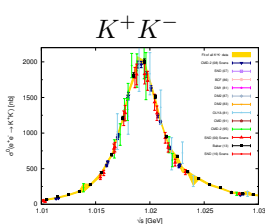
HLMNT11:  $47.51 \pm 0.99$   
 KNT17:  $47.92 \pm 0.70$



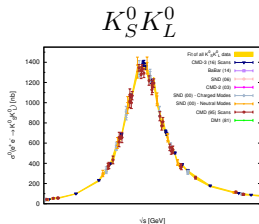
HLMNT11:  $14.65 \pm 0.47$   
 KNT17:  $15.18 \pm 0.14$



HLMNT11:  $20.37 \pm 1.26$   
 KNT17:  $20.07 \pm 1.19$



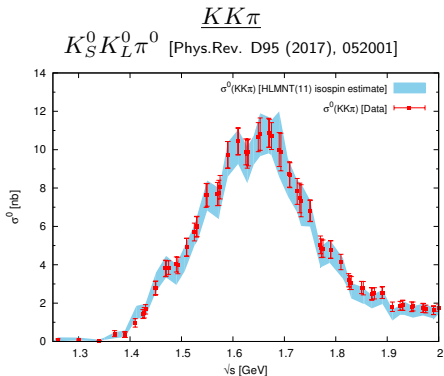
HLMNT11:  $22.15 \pm 0.46$   
 KNT17:  $22.79 \pm 0.25$



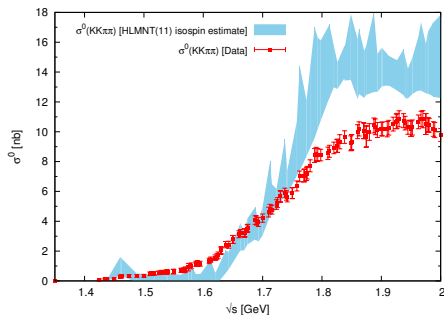
HLMNT11:  $13.33 \pm 0.16$   
 KNT17:  $13.04 \pm 0.12$

# $KK\pi$ , $KK\pi\pi$ and isospin

⇒ New BaBar data for  $KK\pi$  and  $KK\pi\pi$   
 removes reliance on isospin (only  $K_S^0 = K_L^0$ )



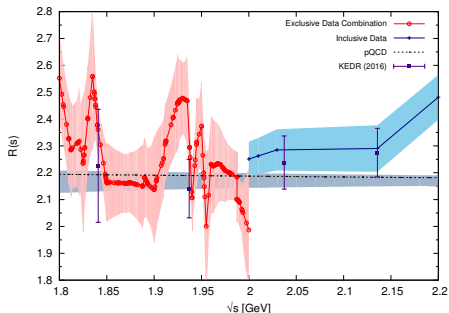
$KK\pi\pi$   
 $K_S^0 K_L^0 \pi^+ \pi^-$  [Phys.Rev. D80 (2014), 092002]  
 $K_S^0 K_S^0 \pi^+ \pi^-$  [Phys.Rev. D80 (2014), 092002],  
 $K_S^0 K_L^0 \pi^0 \pi^0$  [Phys.Rev. D95 (2017), 052001]  
 $K_S^0 K^\pm \pi^\mp \pi^0$  [arXiv:1704.05009]



⇒ **But**, still reliant on isospin estimates for  $\pi^+ \pi^- 3\pi^0$ ,  $\pi^+ \pi^- 4\pi^0$ ,  $KK3\pi$ ...

# Inclusive

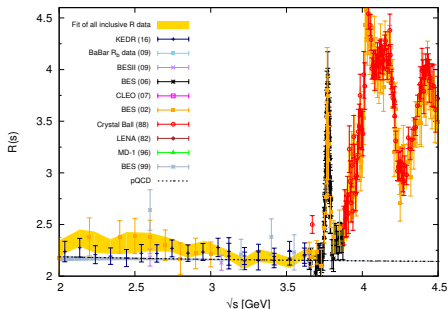
⇒ **New KEDR inclusive  $R$  data** ranging  $1.84 \leq \sqrt{s} \leq 3.05$  GeV [Phys.Lett. B770 (2017) 174-181]  
and  $3.12 \leq \sqrt{s} \leq 3.72$  GeV [Phys.Lett. B753 (2016) 533-541]



$a_{\mu}^{\text{had, LOVP}}(1.84 \leq \sqrt{s} \leq 2.00 \text{ GeV})$ :

pQCD :  $6.42 \pm 0.03$

Data :  $6.88 \pm 0.25$



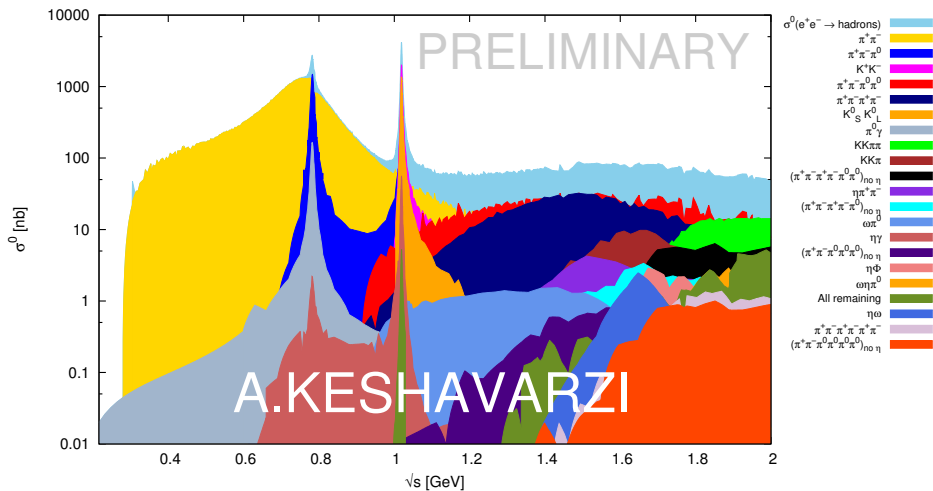
$a_{\mu}^{\text{had, LOVP}}(2.60 \leq \sqrt{s} \leq 3.73 \text{ GeV})$ :

pQCD (inflated errors) :  $10.82 \pm 0.38$

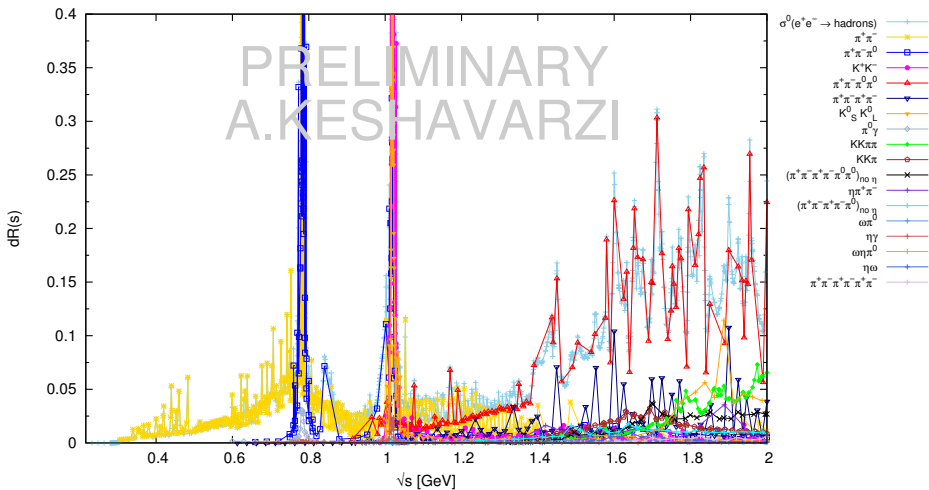
Data :  $11.20 \pm 0.14$

⇒ **Choose to adopt entirely data driven estimate from threshold to 11.2 GeV**

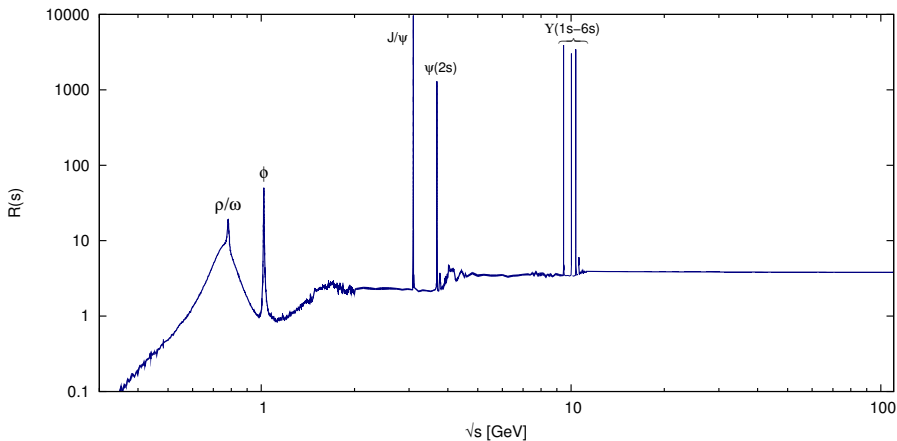
## Contributions to mean value below 2GeV



## Contributions to uncertainty below 2GeV



$R(s)$  for  $m_\pi \leq \sqrt{s} < \infty$



⇒ Full compilation data set for hadronic  $R$ -ratio to be made available soon...

⇒ ...complete with full covariance matrix

# KNT17 $a_\mu^{\text{had, VP}}$ and $\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$ update [preliminary]

$$\underline{(g-2)_\mu}$$

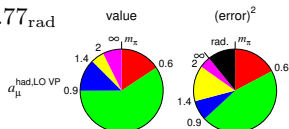
$$\text{HLMNT(11): } 694.91 \pm 4.27$$

$$\text{This work: } a_\mu^{\text{had, LOVP}} = 692.23 \pm 1.26_{\text{stat}} \pm 2.02_{\text{sys}} \pm 0.31_{\text{vp}} \pm 0.70_{\text{fsr}}$$

$$= 692.23 \pm 2.42_{\text{exp}} \pm 0.77_{\text{rad}}$$

$$= 692.23 \pm 2.54_{\text{tot}}$$

$$a_\mu^{\text{had, NLOVP}} = -9.83 \pm 0.04_{\text{tot}}$$



$$\underline{\Delta\alpha(M_Z^2)}$$

$$\text{HLMNT11: } (276.26 \pm 1.38_{\text{tot}}) \times 10^{-4}$$

$$\text{This work: } \Delta\alpha_{\text{had}}^{(5)}(M_Z^2) = (276.06 \pm 0.39_{\text{stat}} \pm 0.64_{\text{sys}} \pm 0.08_{\text{vp}} \pm 0.82_{\text{fsr}}) \times 10^{-4}$$

$$= (276.06 \pm 0.76_{\text{exp}} \pm 0.83_{\text{rad}}) \times 10^{-4}$$

$$= (276.06 \pm 1.13_{\text{tot}}) \times 10^{-4}$$

$\Rightarrow$  Accuracy better than 0.4%  
(uncertainties include all available correlations)

Full KNT17 VP package  
[vp\_knt\_v3\_0.f]  
available for use



# KNT17 vs. DHMZ17 vs. FJ17 [preliminary]

⇒ Different data treatment/methods produce **very different results**

| Channel $\sqrt{s} \leq 1.8$ GeV   | KNT17             | DHMZ17            | FJ17              |
|-----------------------------------|-------------------|-------------------|-------------------|
| $\pi^+\pi^-$                      | $502.73 \pm 1.94$ | $507.14 \pm 2.58$ |                   |
| $\pi^+\pi^-2\pi^0$                | $17.82 \pm 0.99$  | $18.03 \pm 0.54$  |                   |
| $2\pi^+2\pi^-$                    | $14.00 \pm 0.20$  | $13.68 \pm 0.31$  |                   |
| $K^+K^-$                          | $22.75 \pm 0.26$  | $22.81 \pm 0.41$  |                   |
| $K_S^0K_L^0$                      | $13.03 \pm 0.20$  | $12.82 \pm 0.24$  |                   |
| Total HVP $\sqrt{s} < \infty$ GeV | $692.23 \pm 2.54$ | $693.1 \pm 3.4$   | $689.43 \pm 3.25$ |

⇒ Between  $1.8 \leq \sqrt{s} \leq 2$  GeV, KNT use **data**, DHMZ use **pQCD**

BUT, **pQCD** =  $8.30 \pm 0.09$ , **KNT data** =  $8.42 \pm 0.29$ , **DHMZ data** =  $7.71 \pm 0.32$

⇒ **DHMZ17 use correlated systematics differently in determination of the mean value**

→ Determining  $\pi^+\pi^-$  **using only local weighted average** gives  $508.91 \pm 2.84$

→ **Much better agreement** when neglecting the effect of correlated uncertainties on the mean value

# KNT17 $a_\mu^{\text{SM}}$ update [preliminary]

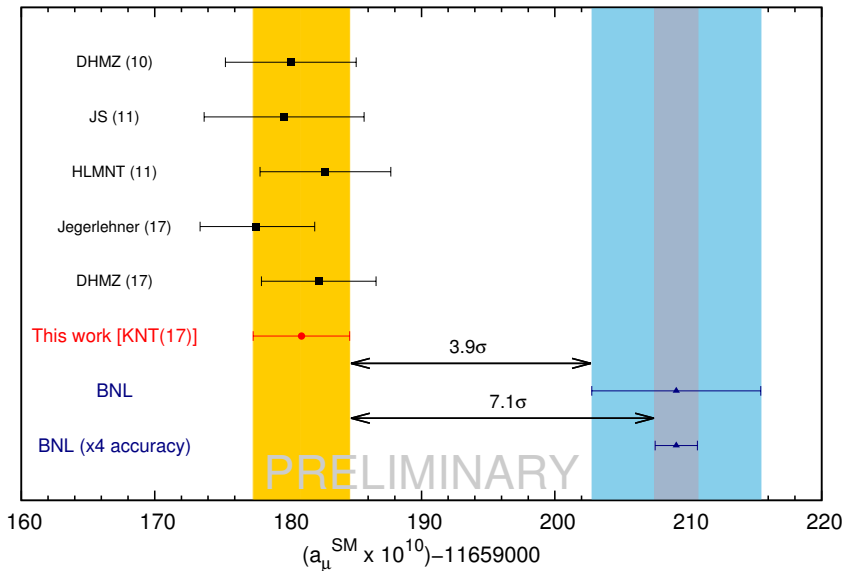
|          | <u>2011</u>        | → | <u>2017</u>        |                                      |
|----------|--------------------|---|--------------------|--------------------------------------|
| QED      | 11658471.81 (0.02) | → | 11658471.90 (0.01) | [Phys. Rev. Lett. 109 (2012) 111808] |
| EW       | 15.40 (0.20)       | → | 15.36 (0.10)       | [Phys. Rev. D 88 (2013) 053005]      |
| LO HLbL  | 10.50 (2.60)       | → | 9.80 (2.60)        | [EPJ Web Conf. 118 (2016) 01016]     |
| NLO HLbL |                    |   | 0.30 (0.20)        | [Phys. Lett. B 735 (2014) 90]        |

|          | <u>HLMNT11</u> | → | <u>KNT17</u>  |                                |
|----------|----------------|---|---------------|--------------------------------|
| LO HVP   | 694.91 (4.27)  | → | 692.23 (2.54) | this work                      |
| NLO HVP  | -9.84 (0.07)   | → | -9.83 (0.04)  | this work                      |
| NNLO HVP |                |   | 1.24 (0.01)   | [Phys. Lett. B 734 (2014) 144] |

|              |                    |   |                    |           |
|--------------|--------------------|---|--------------------|-----------|
| Theory total | 11659182.80 (4.94) | → | 11659181.00 (3.62) | this work |
| Experiment   |                    |   | 11659209.10 (6.33) | world avg |
| Exp - Theory | 26.1 (8.0)         | → | 28.1 (7.3)         | this work |

|                |              |   |              |           |
|----------------|--------------|---|--------------|-----------|
| $\Delta a_\mu$ | 3.3 $\sigma$ | → | 3.9 $\sigma$ | this work |
|----------------|--------------|---|--------------|-----------|

# KNT17 $a_\mu^{\text{SM}}$ update [preliminary]



# Conclusions

- ✓ Can use **dispersive techniques** to calculate **HVP** and **HLbL** contributions to  $a_\mu$
- ✓ Many necessary changes made **in order to improve data combination**
- ⇒ When **combining data...**
  - ✓ ...all **covariance matrices** are **correctly constructed** with a **framework that can accommodate any available correlations**
  - ✓ ...employ a **linear  $\chi^2$  minimisation** that has been shown to be **free from bias**
- ✓ New method shows **improvements in all channels** due to **increased fit flexibility**
- ✓ **Less reliance on isospin** for estimated states with more measured final states
- ✓  $a_\mu^{\text{had,LOVP}}$  **accuracy better than 0.4%**
- ✓  $(g-2)_\mu$  **theory initiative** to collaborate to produce **single estimate** for  $a_\mu^{\text{SM}}$

**Eagerly awaiting new FNAL result!**

# Extra Slides

# Fixing the covariance matrix [JHEP 1005 (2010) 075, Eur.Phys.J. C75 (2015), 613]

⇒ Apply a procedure to **fix the covariance matrix**

$$C_I(i^{(m)}, j^{(n)}) = C^{\text{stat}}(i^{(m)}, j^{(n)}) + \frac{C^{\text{sys}}(i^{(m)}, j^{(n)})}{R_i^{(m)} R_j^{(n)}} R_m R_n ,$$

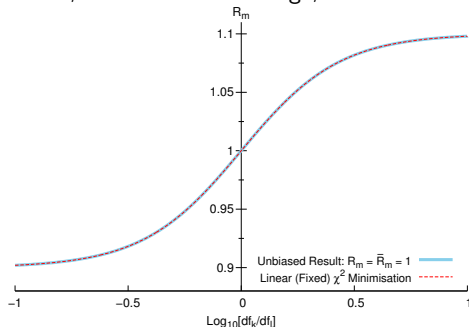
in an **iterative  $\chi^2$  minimisation** method that, to our best knowledge, is **free from bias**

⇒ Fixing with theory value **regulates influence**

⇒ Can be shown from toy models to be **free from bias**

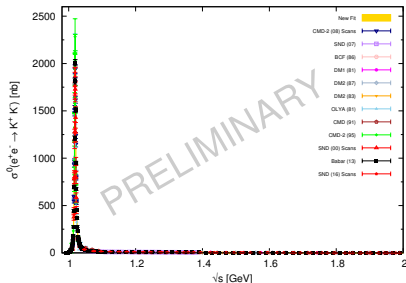
⇒ **Swift convergence**

⇒ Comparison with past results shows **HLMNT11 estimates are largely unaffected**



**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**

## Kaon FSR study



BUT  $K^+K^-$  cross section is totally dominated by  $\phi$  resonance

⇒ No phase space for creation of hard real photons at  $\phi$

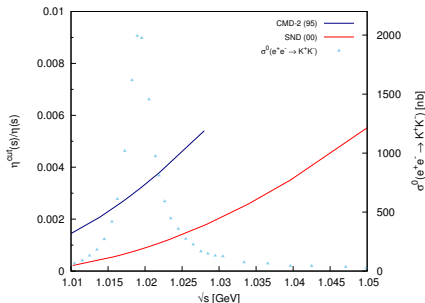
Inclusive FSR correction is large over-correction →

∴ No longer apply FSR correction

Inclusive FSR correction was previously applied to  $K^+K^-$  cross section

KLN theorem requires all virtual and soft corrections necessarily included in given cross section

∴ Only hard real radiation is left to be corrected for



## Properties of a covariance matrix

Any covariance matrix,  $\mathcal{C}_{ij}$ , of dimension  $n \times n$  must satisfy the following requirements:

- As the diagonal elements of any covariance matrix are populated by the corresponding variances, all the diagonal elements of the matrix are positive. Therefore, the trace of the covariance matrix must also be positive

$$\text{Trace}(\mathcal{C}_{ij}) = \sum_{i=1}^n \sigma_{ii} = \sum_{i=1}^n \text{Var}_i > 0$$

- It is a symmetric matrix,  $\mathcal{C}_{ij} = \mathcal{C}_{ji}$ , and is, therefore, equal to its transpose,  $\mathcal{C}_{ij} = \mathcal{C}_{ij}^T$
- The covariance matrix is a positive, semi-definite matrix,

$$\mathbf{a}^T \mathcal{C} \mathbf{a} \geq 0 ; \mathbf{a} \in \mathbf{R}^n,$$

where  $\mathbf{a}$  is an eigenvector of the covariance matrix  $\mathcal{C}$

- Therefore, the corresponding eigenvalues  $\lambda_{\mathbf{a}}$  of the covariance matrix must be real and positive and the distinct eigenvectors are orthogonal

$$\mathbf{b}^T \mathcal{C} \mathbf{a} = \lambda_{\mathbf{a}}(\mathbf{b} \cdot \mathbf{a}) = \mathbf{a}^T \mathcal{C} \mathbf{b} = \lambda_{\mathbf{b}}(\mathbf{a} \cdot \mathbf{b})$$

$$\therefore \text{if } \lambda_{\mathbf{a}} \neq \lambda_{\mathbf{b}} \Rightarrow (\mathbf{a} \cdot \mathbf{b}) = 0$$

- The determinant of the covariance matrix is positive:  $\text{Det}(\mathcal{C}_{ij}) \geq 0$



# Tests of reliability of $f_k$ method

## Did the $f_k$ method incur a bias?

Compare  $f_k$  method and fixed matrix method with **only multiplicative normalisation uncertainties**.

→ If we see **differences** in mean value, then **bias previously influenced the fit**.

→ **Previous results unreliable**

→ If we see **no differences** in mean value, then **bias did not influence fit** (any change comes from improved treatment of systematics)

→ **Previous results reliable**

*Example -  $\pi^+\pi^-$*

Set 1 - CMD-2(06) (0.7% Systematic Uncertainty), Set 2 - CMD-2(06) (0.8% Systematic Uncertainty), Set 3 - SND(04) (1.3% Systematic Uncertainty)

From 0.37 → 0.97 GeV

| Fit Method:  | $f_k$ method      |                               | Fixed matrix method |                               |            |
|--------------|-------------------|-------------------------------|---------------------|-------------------------------|------------|
| Channel      | $a_\mu$           | $\chi^2_{\min}/\text{d.o.f.}$ | $a_\mu$             | $\chi^2_{\min}/\text{d.o.f.}$ | Difference |
| $\pi^+\pi^-$ | $481.42 \pm 4.26$ | 1.10                          | $481.42 \pm 4.05$   | 1.02                          | 0.00       |