Cosmic rays and accelerator physics at LHCf

Oscar Adriani University of Florence & INFN Firenze

> NPQCD 2017 Pollenzo, May 22nd, 2017

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ Contents

Introduction

- LHCf @ different energies and different beams
 - Contribution to CR physics
 - Contribution to forward physics
- RHICf
- Future @ LHC



Introduction

+

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ The High Energy cosmic ray spectrum

- The spectrum falls very rapidly with energy ($\sim E^{-2.7}$)
- No direct measurements are possible for E>10¹⁵ eV (Flux< 1/m²/year)
- We have to rely on the atmospheric showers measurements



High Energy CR Showers main Observables



- X_{max} : depth of air shower maximum in the atmosphere
- RMS(X_{max}): fluctuations in the position of the shower maximum
- N_µ: number of muons in the shower at the detector level
- To go from these observables to the CR composition and energy determination passing through the hadronic interaction models is mandatory

Uncertainty of hadron interaction models

Uncertainty in the interpretation of the observables

+ The role of the accelerators experiments





Accelerator based experiments are the most powerful available tools to determine the high energy hadronic interactions characteristics → Hadronic interactions models tuning

LHC 13 TeV \rightarrow 9.10¹⁶ eV Unique opportunity to calibrate the models in the 'above knee' region

elerator physics at LHCf

How accelerator experiments can contribute?



+ Impressive coverage of the central region







General purpose detectors (ATLAS, CMS,...) cover the spatial region at low rapidity.

Special detectors to access forward particles are necessary!

- The largest detectors for particle physics
- Surrounding the LHC Interaction Points
- Covering many fundamental physics

items

 Designed for discoveries!





Cosmic rays and accelerator physics at LHCf

+ And also of the forward region!



+ LHC phase space coverage



We may profit (and we are profiting) of the very broad coverage! Dedicated forward detectors for a better measurement of the energy flow

First models tuning after the first LHC data (EPOS and QGSJET)





Significant reduction of differences btw different hadronic interaction models!!!

O. Adriani

Cosmic rays and accelerator physics at LHCf

Second models tuning after the first LHC data (Sibyll 2.3)



+ But not everything is perfect....

C



+ LHCf detector and performances

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ LHCf: location and detector layout **Detector II Detector I INTERACTION POINT** Tungsten Tungsten GSO GSO IP1 (ATLAS) **GSO** bars Silicon ustrips **Front Counter** Front Counter 140 m 140 m 8 cm 6 cm π^0 П **INCOMING NEUTRAL** PARTICLE BEAM 44X₀, 1.6 λ_{int} Energy resolution: < 5% for photons 30% for neutrons Position resolution: **Arm#1 Detector** $< 200 \,\mu$ m (Arm#1) **Arm#2 Detector** 20mmx20mm+40mmx40mm 25mmx25mm+32mmx32mm $40 \,\mu$ m (Arm#2) **4 X-Y GSO Bars tracking layers 4 X-Y Silicon strip tracking layers** Pseudo-rapidity range: $\eta > 8.7$ @ zero Xing angle $\eta > 8.4 @ 140 urad$

+ A brief LHCf photo-history



Dec-Jul 2010 0.9TeV& 7TeV pp Detector removal





Dec 2012- Feb 2013 5TeV/n pPb, 2.76TeVpp (Arm2 only) Detector removal



May-June 2015 13 TeV dedicated pp Detector removal



November 2016 8 Tev p-Pb

+ Event category in LHCf



O. Adriani

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ Event category in LHCf



+ π^0 reconstruction



+ γγ invariant mass distribution



О.

LHCf physics results

÷

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ LHCf Data Taking and Analysis matrix

	Proton E _{LAB} (Ev)	Photon (EM shower)	Neutron (hadron shower)	Π ⁰ (EM shower)	
Test beam at SPS		NIM. A 671, 129–136 (2012) JINST 12P03023(2017)	JINST 9 P03016 (2014) (2014)P03016		
p-p at 900GeV	$4.3 x 10^{14}$	Phys. Lett. B 715, 298-303 (2012)			
p-p at 7TeV	2.6x10 ¹⁶	Phys. Lett. B 703, 128–134 (2011)	Phys. Lett. B 750, 360-366 (2015)	Phys. Rev. D 86, 092001 (2012)+ Phys. Rev. D 94, 032007(2016) Type II	Runl
p-p at 2.76TeV	4.1x10 ¹⁵			Phys. Rev. C 89, 065209 (2014)+	Dum (
p-Pb at 5.02TeV	1.3×10^{16}			032007(2016) Type II	KUNZ
p-p at 13TeV	9.0×10^{16}	Submitted to PLB	Preliminary results		Run3
p-Pb at 8.1 TeV	3.6×10^{16}	Run completed in November 2016			Run4

O. Adriani

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

γ energy spectra at 13 TeV

η >10.94



QGSJET II-04: overall good agreement **EPOS-LHC**: overall good agreement **DPMJET 3.06**: overall higher flux **SIBYLL 2.3**: overall lower flux **PYTHIA 8.212**: higher flux above 3 TeV

O. Adriani

+ γ energy spectra at 13 TeV

8.81< η <8.99



QGSJET II-04: overall lower flux EPOS-LHC: higher flux above 3-4 TeV DPMJET 3.06: overall higher flux SIBYLL 2.3: higher flux above 2 TeV PYTHIA 8.212: higher flux above 3 TeV

Cosmic rays and accelerator physics at LHCf

+ Photon spectra – Feynman Scaling



Feynman scaling: differential cross section as a function of $X_{_{\rm F}}$ independent of \sqrt{s} for $X_{_{\rm F}}$

Feynman scaling holds within systematic uncertainties

Cosmic rays and accelerator physics at LHCf

+ Preliminary ARM2 unfolded neutron spectra @ 13 TeV Differential production cross section



O. Adria

+ Feynman scaling in neutron production cross-section



Feynman scaling hypothesis holds within the error bars Consistency is good especially in the region $0.2 < x_F < 0.75$

+ Measurement of interesting quantities for CR Physics



+ LHCf @ pPb 5.02 TeV: π^0 analysis

(Soft) QCD : **Ultra peripheral collisions :** central and peripheral collisions virtual photons from rel. Pb collides a proton **Central collisions** Peripheral collisions impact parameter : b Pb proton $b > R_p + R_{\rm Pb}$ Pb $b \ll R_p + R_{\rm Pb}$ $b \sim R_p + R_{\rm Pb}$ proton Momentum distribution of the UPC induced secondary particles is estimated as rest frame 1. energy distribution of virtual photons is estimated by the Weizsacker Williams approximation. 2. photon-proton collisions are simulated by the SOHIA model (E_{γ} > pion threshold). 3. produced mesons and baryons by γ -p collisions are boosted along the proton beam. Dominant channel to forward π^0 is [mb/GeV Ed³ cv/dp³ [mb/GeV² 00 UPCs(py 2eV_E__300et GeV<E < 10GeV)GeV<Ė <100GeV DPMJET3 $\gamma + p \rightarrow \Delta(1232) \rightarrow p + \pi^0$ Ed³σ/dp³ [100GeV<E_<1000000GeV Comparison Break down About half of the observed π^0 with soft-QCD of UPC 10 may originate in UPC, another half is from soft-QCD. 0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.1 0.2 0.3 0.4 0.5 0.6 0.7 P_T [GeV/c] P_T [GeV/c] O. Adriani I UNCILLO, IMAY DD , DUI

Cosmic rays and accelerator physics at LHCf

+ LHCf @ pPb 5.02 TeV: π^0 pT spectra



0......

+ Nuclear modification factor



zo, May 22nd, 2017

LHCf & ATLAS

+

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ ATLAS-LHCf combined data taking

- Trigger sharing with ATLAS at ~100-500 Hz in p+p (500 Hz in 2016 p+Pb)
- Off-line event matching
- Internal note (p+Pb 2013)
 - ATL-PHYS-PUB-2015-038
- Important to separate the contributions due to diffractive and non-diffractive collisions
 - It makes more easy improving the hadronic interaction models





+ Diffractive studies

MC studies

- Contributions on forward photon/ neutron spectra from diffractive/nondiffractive collisions.
- Event-selection by the central particle production to separate these events



Very forward photon energy spectra predicted by four models with **total/diffractive/nondiffractive**

- Total: Very similar spectra in EPOS,QGSJET and SIBYLL (LHCf alone)
- Diffractive/Non-diffractive: Very big difference between models (ATLAS-LHCf)

Zhou et al., Eur.Phys.J. C77 (2017) no.4, 212

sics at LHCf

+ Diffractive studies

Event selection for
Diffractive/Non-diffractive
by using N_{charged} with
p_T>100MeV in | η |<2.5

Expected efficiencies

By using ATLAS-tracker information, We can separate diffractive/non-diffractive events with high efficiency and purity



Cosmic rays and accelerator physics at LHCf

+ Physics cases with Atlas jointly taken data

- In p+p collisions
 - Forward spectra of Diffractive/ Nondiffractive events
 - Measurement of proton-π collisions

Both are important for preciseunderstanding of CR air shower development



<u>р-п measurement at LHC</u>

Leading neutron can be tagged by LHCf detectors -> total cross section multiplicity measurement



In p+Pb collisions

Measurement of UPC in the forward region.
The future....

+

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ The future @ RHIC: From the Large Hadron Collider to the Longisland Hadron Collider



LHCf Arm2 detector in the LHC tunnel









+ RHICf detector acceptancemmx20mm and 40mmx40mm)



+ \sqrt{s} scaling, or breaking?

LHCf 2.76TeV and 7TeV data shows scaling of forward π^0



O. Adriani

Pollenzo, May 22nd, 2017

+ RHICf commissioning

- RHIC starts first RUN2017 collision on 20-Feb
- RHICf observed shower signal (PMT coincidence) and tuned timing
- Common operation (RHICf triggers STAR) tested and common data successfully recorded at STAR (analysis of physics correlation on going)
- Data taking: last week of June





RHICf calorimeter PMT signals and ADC Gate after timing



Bunch ID of RHICf trigger recorded at "STAR" Two abort gaps correctly identified O. Adriani

+ Diffractive vs. non diffractive at $\eta > 8.2$ with $\sqrt{s}=510$ GeV p+p collisions



PYTHIA 8 simulation

BLUE: inclusive spectra expected by RHICf only

RED: diffractive only ("RHICf + no central track in STAR" will be similar => TBC) BLACK: non diffractive ("RHICf + >=1 central track in STAR" => TBC)

+ The Near-Far Future at LHC

- The most promising future at LHC for LHCf involve the proton-light ion collisions
- To go from p-p to p-Air is not so simple....
- Comparison of p-p, Pb-Pb and p-Pb is useful, but model dependent extrapolations are anyway necessary
- Direct measurements of p-O or p-N could significantly reduce some systematic effects
- Still make sense to take data if intermediate ion (like Ar) will be available



Conclusions

- In the last few years the importance of accelerator based measurements useful for Cosmic Ray physics came up very clearly, in addition to the 'standard' physics case
- LHC is the ideal laboratory for these studies
- Many important measurements have already been done
 - Significant improvement of EPOS_LHC, QGSJET-04 and Sibyll_2.3 hadronic interaction models
- LHCf provided many precise results on forward γ , n and π^0 with different collision's conditions
- Joint analysis with Atlas is on-going for diffractive/non diffractive events selection
- RHICf will take data next month

÷

Backup slides

+

Cosmic rays and accelerator physics at LHCf

Pollenzo, May 22nd, 2017

+ LHCf-Atlas: photons





v 22nd, 2017

+ RHICf beam condition proposal

- RHICf DAQ speed is limited to 1kHz
- Collision pile up cannot be resolved
- Small angular dispersion is preferred
- Beam Proposal
 - 510GeV p+p collisions
 - $\beta^* = 10m$
 - Radial (horizontal) polarization; 0.4-0.5
 - $\varepsilon = 20$ mm mrad, $I_b = 2 \times 10^{11}$, $n_{b-colliding} = 100$, $n_{b-noncolliding} = 20$ (nominal)
 - Luminosity=1.1 10^{31} cm⁻²s⁻¹
- Operation
 - Few days for physics and few days for contingency
 - π⁰ (double tower event) enhanced and single shower prescaled triggers are used simultaneously
 - Trigger exchange with PHENIX
 - Stay at the garage position not to interfere ZDC when RHICf does not take data

LHCf @ pp 7 TeV: neutron analysis

Motivations:

- Inelasticity measurement k=1-p_{leading}/p_{beam}
- Muon excess at Pierre Auger Observatory
 Cosmic rays experiment measure PCR energy from muon
 number at ground and florescence light
 20-100% more muons than expected have been observed





Number of muons depends on the energy fraction of produced hadron

- Muon excess in data even for Fe primary MC!!!!
- EPOS predicts more muons due to larger baryon production, even if it is not sufficient to reproduce the experimental data

importance of baryon measurement!!!



+ Inclusive neutron spectra (7 TeV pp)



Very large high energy peak in the η >10.76 (predicted only by QGSJET) \rightarrow Small inelasticity in the very forward region!

O. Adriani

+ Type II π^0 in pp 7 TeV collisions

Present LHCf results are based on the Type-I π^0 events. Improved π^0 reconstruction, Type-II, is now ready for use in analysis.



+ π^0 energy spectra (for different p_T bins)



- DPMJET and PYTHIA are harder than LHCf $p_T < 1.0$ GeV, although compatible at low p_T and low E.
- QGSJET II gives good agreement at $0 < p_T < 0.2$ GeV and $0.8 < p_T < 1.0$ GeV.
- EPOS 1.99 agrees with LHCf at $0.4 < p_T < 0.8$ GeV. LHCf prefers EPOS 1.99 than EPOS LHC.

+ $\pi^0 p_T$ spectra (for different rapidity bins)



O. Adriani

Cosmic rays and accelerator physics at LHCf

+ 2015 updated LHC operation schedule Start LHC commissioning LHCf run LHCfsremoval with beam May Jùne Apr 14 16 17 18 19 20 21 23 24 26 Wk 15 22 30 Easter Mon c Whit Mo 25 13 18 20 27 11 22 8 15 Tu Special physicrun **Recommissioning with** TS1 Injector TS We beam Machine checkout Th Ascension ay 1st May Fr Sa Su

From M. Lamont, LMC Meeting, 15/04/15

- 8 weeks beam commissioning
- 5 days special physics at beta* = 19 m (VdM, LHCf, TOTEM & ALFA)
- Start TS1 15th June. 24 hour technical stop in SPS in parallel followed by SPS scrubbing.

+ DATA vs MC : comp. 900GeV/7TeV

• None of the model nicely agrees with the LHCF data

• Here we plot the ratio MC/Data for the various models



+ DATA : 900GeV vs 7TeV



✓ Normalized by the number of entries in $X_F > 0.1$

✓ No systematic error is considered in both collision energies.

Good agreement of X_F spectrum shape between 900 GeV and 7 TeV. →weak dependence of $< p_T >$ on E_{CMS}

$$rac{1}{\sigma_{
m inel}} rac{d\sigma_{\gamma}}{dX_{
m F}}\Big|_{\eta <
m limited} \propto rac{1}{\sigma_{
m inel}} rac{d\sigma_{\gamma}}{p_{
m T} dp_{
m T} dX_{
m F}} \langle p_{
m T}
angle dp_{
m T}$$

Pollenzo, May 22nd, 2017



π^{0} analysis at $\sqrt{s}=7TeV$



1. Thermodynamics (Hagedron, Riv. Nuovo Cim. 6:10, 1 (1983)) $rac{d^2 - c^2}{dp^3} = A \cdot \exp(-\sqrt{p_{
m T}^2 c^2} + m_{\pi^0}^2 c^4/T)^2$ $\sigma_{\rm inel}$ $rac{\pi m_{\pi^0}c^2T}{2}rac{K_2(m_{\pi^0}c^2/T)}{K_{3/2}(m_{\pi^0}c^2/T)}$ 2. Numerical integration actually up to the $2\pi p_{\mathrm{T}}^2 f(p_{\mathrm{T}}) dp_{\mathrm{T}}$ upper bound of histogram

 $2\pi p_{\mathrm{T}}$]

- Systematic uncertainty of LHCf data is 5%.
- Compared with the UA7 data ($\sqrt{s}=630$ GeV) and MC simulations (QGSJET, SIBYLL, EPOS).
- Two experimental data mostly appear to lie along a common curve
 - \rightarrow no evident dependence of $< p_T >$ on E_{CMS}.
- Smallest dependence on ECMS is found in EPOS and it is consistent with LHCf and UA7.
- Large E_{CMS} dependence is found in SIBYLL

elerator physics at LHCf

05.4578).

+ Muon excess at Pierre Auger Obs.





O. .



Auger hybrid analysis

- event-by-event MC selection to fit FD data (top-left)
- comparison with SD data vs MC (topright)
- muon excess in data even for Fe primary MC

EPOS predicts more muon due to larger baryon production

=> importance of baryon measurement

Pierog and Werner, PRL 101 (2008) 171101

ccelerator physics at LHCf



+ Common trigger with ATLAS



MC impact parameter vs. # of particles in ATLAS LUCID

- LHCf forced to trigger ATLAS
- Impact parameter may be determined by ATLAS
- Identification of forward-only events

O. Adriani

+ Analysis of hadron production in p-p Data Collisions at 13 TeV 12 July 2015, 22:32-1:30 (3 hours) *p-p*, √s = 13 TeV **Beam Center** Fill # 3855 Arm2 Hadron-like Estimated using 2 $\mu=0.01$ (>1 TeV) 0.8 D fit on high energ $\int Ldt = 0.19 \ nb^{-1}$ y hadron hitmap di 10 $\sigma_{ine} = 78.53 \text{ mb}$ stribution ۲ [mm] 50 0 20 30 -10 10 X [mm] 40 8.81 < n < 8.99

Event selection criteria: software trigger

at least 3 consecutive layers with deposit above threshold $dE>dE^{thr}$

PID selection

 $L_{2D}>L_{2D}^{thr}$ where L_{2D} is a variable related to shower longitudinal profile

pseudorapidity acceptance

O. Adriani PLB 750 (2015) 360-366 rays and accelerator physics at LHCf 3 different pseudorapidity regions. Pollenzo, May 22nd, 2017

30

20

10

-10

-10

0

8.99 < h < 9.22

> 10.76

20

30

X [mm]

10

Same as 7 TeV analysis

Charged particle distribution in pseudorapidity



(



+ Photon reconstruction



Cosmic rays and accelerator physics

+ π^0 mass peak

 $\Delta m_{\gamma\gamma}/m_{\gamma\gamma} \sim 3.5\%$



+ Neutron reconstruction

Performance for 1.5 TeV neutrons: $\Delta E/E \sim 35\%-40\%$ $\Delta x \sim 1mm$

And....

Detector performance is also interaction model dependent!

Unfolding is essential to extract physics results from the measured spectra

Physics measurement important to try to solve the 'Muon eccess' observed from the ground based HECR experiments



+ Reconstructed ARM2 hadron energy spectra @ 13 TeV Events / N_{ine} / dE



DPMJET 3.04 have very different shape and yield

+ LHCf @ pPb 5.02 TeV: π^0 spectra @ p-remnant side





- The LHCf results in p-Pb (filled circles) show good agreement with DPMJET and EPOS.
- The LHCf results in p-Pb are clearly harder than the LHCf results in p-p at 5.02TeV (shaded area)
- which are interpolated from the results at 2.76TeV and 7TeV. O. Adriani

+ Very forward neutral particle spectra: neutrons

- Even larger differences wrt γ!
- 30% energy resolution is not taken into account
- But unfolding works well!



+ What happens if we off-line combine ATLAS and LHCf?

- ATLAS0: no charged particles in the $|\eta| < 2.5$ and $p_t > 0.1$ GeV/c
- ATLAS2: >1 charged particles in the $|\eta| < 2.5$ and $p_t > 0.1$ GeV/c
- Central activity selection enhance the differences btw models
- Could be used to tune different components of the models



+ √s scaling : a key for extrapolation beyond the LHC

All π⁰ expected from models (0.5TeV, 14TeV and 50TeV)



(green triangle in the phase space plot) Cosmic rays and accelerator physics at LHCf

LHCf single photon

Pollenzo, May 22nd, 2017

LHC 7TeV p+p collision LHC 900GeV p+p collision
