

# Update on $a_{\mu}^{\text{had, VP}}$ from KNT17

Alex Keshavarzi

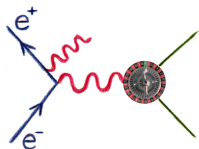
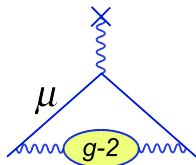
University of Liverpool

(in collaboration with Daisuke Nomura & Thomas Teubner [KNT17])

Working Group on Radiative Corrections and Generators for Low Energy Hadronic Cross Section and Luminosity

Institute of Nuclear Physics, Mainz

30<sup>th</sup> June 2017



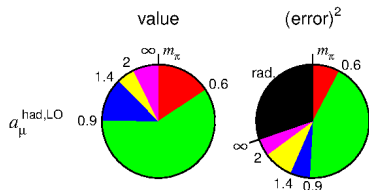
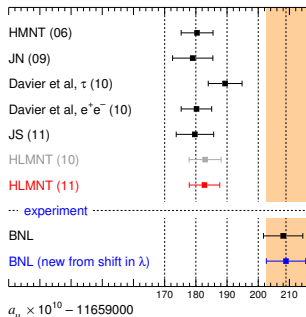
# The previous analysis... [HLMNT(11), J. Phys. G38 (2011), 085003]

<b>QED</b> contribution	11 658 471.808 (0.015) $\times 10^{-10}$	Kinoshita & Nio, Aoyama et al
<b>EW</b> contribution	15.4 (0.2) $\times 10^{-10}$	Czarnecki et al
<b>Hadronic</b> contribution		
<b>LO</b> hadronic	694.9 (4.3) $\times 10^{-10}$	HLMNT11
<b>NLO</b> hadronic	-9.8 (0.1) $\times 10^{-10}$	HLMNT11
<b>light-by-light</b>	10.5 (2.6) $\times 10^{-10}$	Prades, de Rafael & Vainshtein
<b>Theory TOTAL</b>	<b>11 659 182.8 (4.9) <math>\times 10^{-10}</math></b>	
<b>Experiment</b>	<b>11 659 208.9 (6.3) <math>\times 10^{-10}</math></b>	world avg
<b>Exp - Theory</b>	<b>26.1 (8.0) <math>\times 10^{-10}</math></b>	<b>3.3 <math>\sigma</math> discrepancy</b>

(Numbers taken from HLMNT11, arXiv:1105.3149)

- $a_\mu^{\text{had, LOVP}}$  still **dominated uncertainty**
- Potential for **improvement from experimental data and data combination**
- **New x4 accuracy measurements** planned from Fermilab and J-PARC
- ⇒ If  $a_\mu^{\text{SM}}$  &  $a_\mu^{\text{EXP}}$  improve as planned...

**g-2 discrepancy  $> 7\sigma$ !**



# Correlation and covariance matrices

⇒ **Correlated data** beginning to **dominate** full data compilation...

→ Non-trivial, **energy dependent influence** on both **mean value and error estimate**

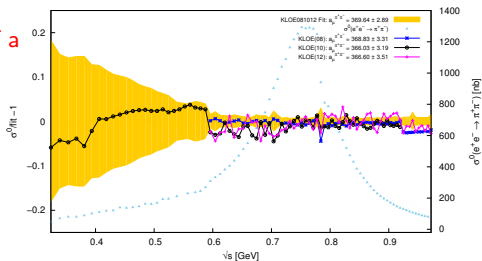
## KNT17 prescription

- Construct full covariance matrices for each channel & entire compilation  
⇒ **Framework available for inclusion of any and all inter-experimental correlations**
- If experiment does not provide matrices...  
→ Statistics occupy diagonal elements only  
→ Systematics are 100% correlated
- If experiment does provide matrices...  
→ **Matrices must satisfy properties of a covariance matrix**

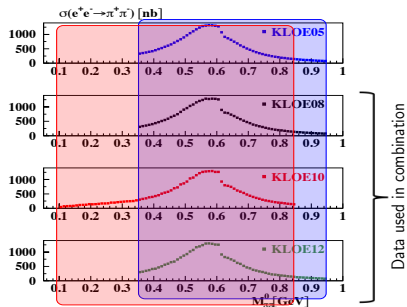
e.g. - KLOE  $\pi^+\pi^-\gamma(\gamma)$  combination  
covariance matrices update

⇒ **Originally, NOT a positive semi-definite matrix:**

**(DO NOT USE PPG14 DATA!)**



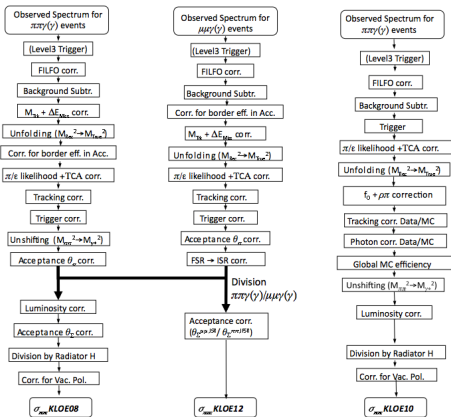
# The KLOE data sets [preliminary]



⇒ Three measurements of  $\sigma_{\pi\pi(\gamma)}^0$  by KLOE

→ KLOE08, KLOE10 and KLOE12

⇒ Overlapping energy range covering  $\rho$  region



⇒ They are, in part, highly correlated

→ must be incorporated

→ e.g. - KLOE08 and KLOE12 share the same  $\pi\pi(\gamma)$  data, with KLOE12 normalised by the measured  $\mu\mu(\gamma)$  cross section

# Updating the KLOE $\pi^+\pi^-\gamma(\gamma)$ data sets [preliminary]

⇒ For **KLOE08**, **KLOE10** and **KLOE 12**,

→ Analysis data is **not rounded**

→ **Updated precision** of input parameters and fundamental constants

⇒ For **KLOE08** and **KLOE10**,

→ **Updated VP correction** (FJ03VP → FJ16VP)

→ VP correction now applied using both **real and imaginary parts**

$$a_{\mu}^{\pi^+\pi^-}(\text{KLOE08}, 0.35 \leq s' \leq 0.95 \text{ GeV}^2) = (386.6 \pm 0.5_{\text{stat}} \pm 3.3_{\text{sys}}) \times 10^{-10}$$

$$\text{Before: } a_{\mu}^{\pi^+\pi^-}(\text{KLOE08}, 0.35 \leq s' \leq 0.95 \text{ GeV}^2) = (387.2 \pm 0.5_{\text{stat}} \pm 3.4_{\text{sys}}) \times 10^{-10}$$

$$a_{\mu}^{\pi^+\pi^-}(\text{KLOE10}, 0.1 \leq s' \leq 0.85 \text{ GeV}^2) = (477.8 \pm 2.2_{\text{stat}} \pm 6.7_{\text{sys}}) \times 10^{-10}$$

$$\text{Before: } a_{\mu}^{\pi^+\pi^-}(\text{KLOE10}, 0.1 \leq s' \leq 0.85 \text{ GeV}^2) = (478.5 \pm 2.0_{\text{stat}} \pm 6.7_{\text{sys}}) \times 10^{-10}$$

⇒ For **KLOE12**,

→ **Corrected** for flaw in error determination

$$a_{\mu}^{\pi^+\pi^-}(\text{KLOE12}, 0.35 \leq s' \leq 0.95 \text{ GeV}^2) = (385.1 \pm 1.2_{\text{stat}} \pm 2.3_{\text{sys}}) \times 10^{-10}$$

$$\text{Before: } a_{\mu}^{\pi^+\pi^-}(\text{KLOE12}, 0.35 \leq s' \leq 0.95 \text{ GeV}^2) = (385.1 \pm 1.1_{\text{stat}} \pm 2.7_{\text{sys}}) \times 10^{-10}$$

# KLOE $\pi^+\pi^-\gamma(\gamma)$ correlations [preliminary]

## KLOE08 and KLOE10

Statistics - **no correlation**

Systematics - **luminosity, radiator function and vacuum polarisation correction**

(Note: more correlations now included after revised analysis. Waiting for approval from KLOE collaboration.)

## KLOE08 and KLOE12

Statistics - **unfolding and unshifting**

Systematics - **all uncertainties that enter from shared  $\pi^+\pi^-\gamma(\gamma)$  data are correlated**

## KLOE10 and KLOE12

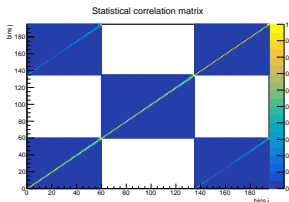
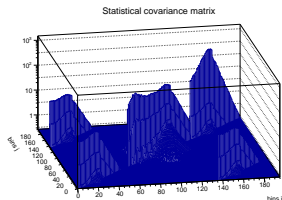
Statistics - **no correlation**

Systematics - **no correlation**

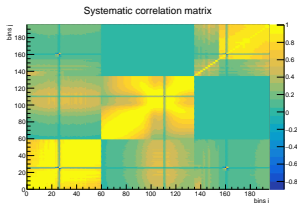
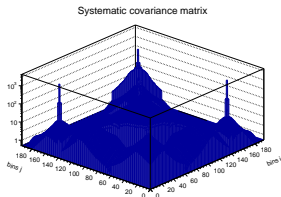
(Note: KLOE10 and KLOE12 are now correlated for systematic uncertainties after revised analysis. Waiting for approval from KLOE collaboration.)

# The KLOE $\pi^+\pi^-\gamma(\gamma)$ combination covariance matrices [preliminary]

Statistical covariance and correlation matrix:



Systematic covariance and correlation matrix:



# Systematic bias and use of the data/covariance matrix

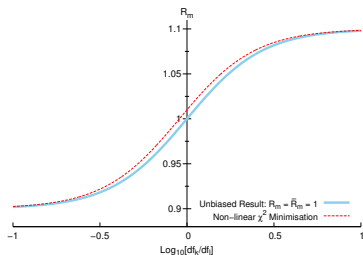
⇒ Data is re-binned using an **adaptive clustering algorithm**

⇒ Iterative fit of covariance matrix as defined by data → **D'Agostini bias**

[Nucl.Instrum.Meth. A346 (1994) 306-311]

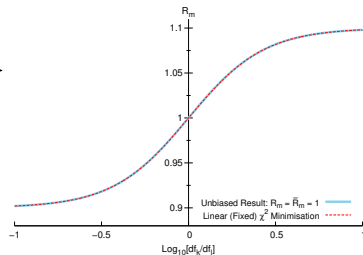
## HLMNT11

⇒ Non-linear  $\chi^2$  minimisation fitting  
nuisance parameters  
→ **Penalty trick bias**



## KNT17

⇒ **Fix the covariance matrix** in an  
iterative  $\chi^2$  minimisation  
→ **Free from bias**



**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**



# Linear $\chi^2$ minimisation

⇒ Redefine clusters to have **linear cross section**

→ **Fix covariance matrix with linear interpolants** at each iteration  
(extrapolate at boundary)

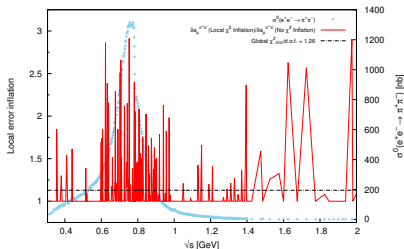
$$\chi^2 = \sum_{i=1}^{N_{\text{tot}}} \sum_{j=1}^{N_{\text{tot}}} (R_i^{(m)} - \mathcal{R}_m^i) \mathbf{C}^{-1}(i^{(m)}, j^{(n)}) (R_j^{(n)} - \mathcal{R}_n^j)$$

⇒ **Through correlations and linearisation**, result is the minimised solution of all neighbouring clusters

→ ...and **solution is the product of the influence of all correlated uncertainties**

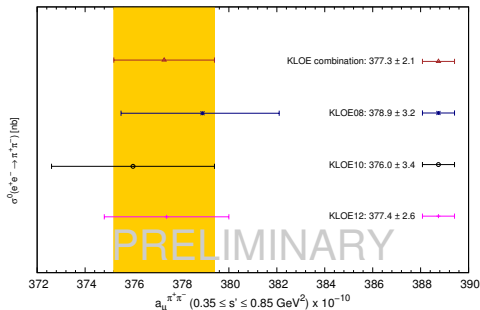
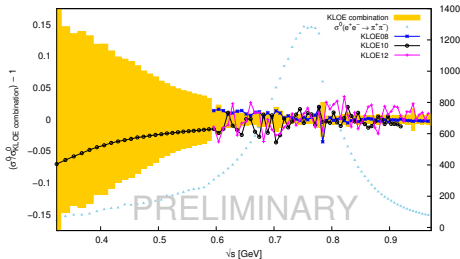
⇒ The **flexibly of the fit** to vary due to the energy dependent, correlated uncertainties benefits the combination

→ ...and any data tensions are reflected in a **local  $\chi^2$  error inflation**



# The resulting KLOE $\pi^+\pi^-\gamma(\gamma)$ combination [preliminary]

⇒ Combination of KLOE08, KLOE10 and KLOE12 gives 85 distinct bins between  $0.1 \leq s \leq 0.95 \text{ GeV}^2$



→ Covariance matrix now correctly constructed

⇒ a **positive semi-definite matrix**

→ **Non-trivial influence of correlated uncertainties** on resulting mean value

$$a_{\mu}^{\pi^+\pi^-}(0.1 \leq s' \leq 0.95 \text{ GeV}^2) = (489.9 \pm 2.0_{\text{stat}} \pm 4.3_{\text{sys}}) \times 10^{-10}$$

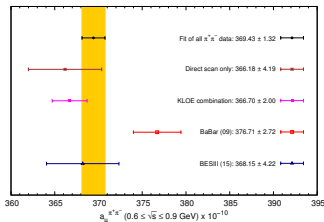
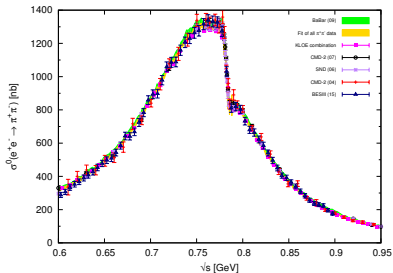
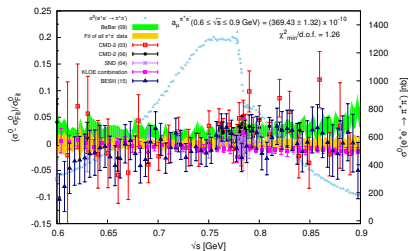
→ **Reduction in uncertainties** within uncertainties of individual measurements...

...and **emanate smallest contributing stat/sys uncertainties**

# $\pi^+\pi^-$ channel [preliminary]

⇒ Large improvement for  $2\pi$  estimate

→ BESIII [Phys.Lett. B753 (2016) 629-638] and KLOE combination provide **downward influence** to mean value



⇒ Correlated & experimentally corrected  $\sigma_{\pi\pi(\gamma)}^0$  data now entirely **dominant**

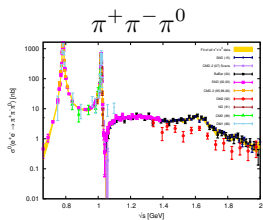
$$a_\mu^{\pi^+\pi^-} (0.305 \leq \sqrt{s} \leq 2.00 \text{ GeV}):$$

$$\text{HLMNT11: } 505.77 \pm 3.09$$

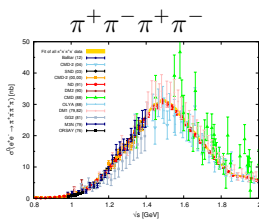
$$\text{KNT17: } 502.85 \pm 1.93$$

(no radiative correction uncertainties)

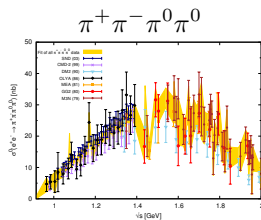
## Other notable exclusive channels



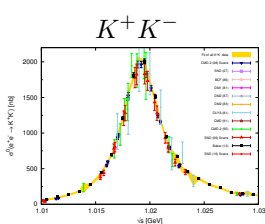
HLMNT11:  $47.51 \pm 0.99$   
 KNT17:  $47.68 \pm 0.70$



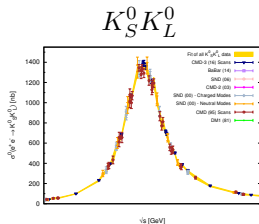
HLMNT11:  $14.65 \pm 0.47$   
 KNT17:  $15.18 \pm 0.14$



HLMNT11:  $20.37 \pm 1.26$   
 KNT17:  $20.07 \pm 1.19$



HLMNT11:  $22.15 \pm 0.46$   
 KNT17:  $22.76 \pm 0.22$



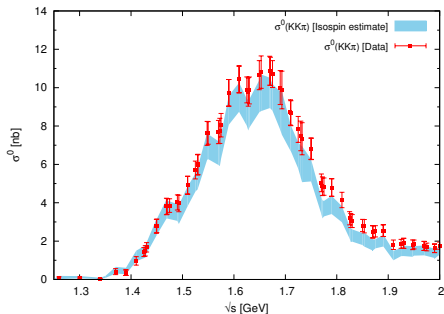
HLMNT11:  $13.33 \pm 0.16$   
 KNT17:  $13.09 \pm 0.12$

# $KK\pi$ , $KK\pi\pi$ and isospin

⇒ New BaBar data for  $KK\pi$  and  $KK\pi\pi$   
 removes reliance on isospin (only  $K_S^0 = K_L^0$ )

## $KK\pi$

$K_S^0 K_L^0 \pi^0$  [Phys.Rev. D95 (2017), 052001]



HLMNT11:  $2.77 \pm 0.15$

KNT17:  $2.83 \pm 0.14$

⇒ **But**, still reliant on isospin estimates for  $\pi^+\pi^-3\pi^0$ ,  $\pi^+\pi^-4\pi^0$ ,  $KK3\pi\dots$

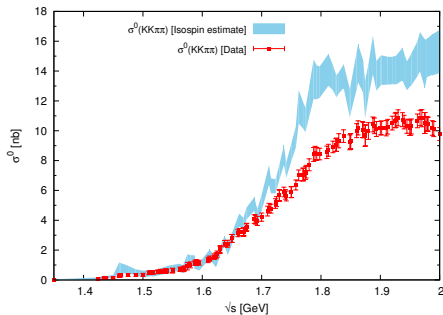
## $KK\pi\pi$

$K_S^0 K_L^0 \pi^+ \pi^-$  [Phys.Rev. D80 (2014), 092002]

$K_S^0 K_S^0 \pi^+ \pi^-$  [Phys.Rev. D80 (2014), 092002],

$K_S^0 K_L^0 \pi^0 \pi^0$  [Phys.Rev. D95 (2017), 052001]

$K_S^0 K^\pm \pi^\mp \pi^0$  [arXiv:1704.05009]

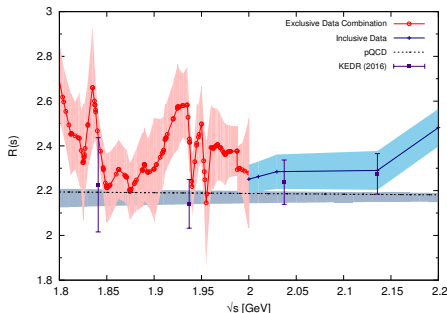


HLMNT11:  $3.31 \pm 0.58$

KNT17:  $2.42 \pm 0.09$

# Inclusive

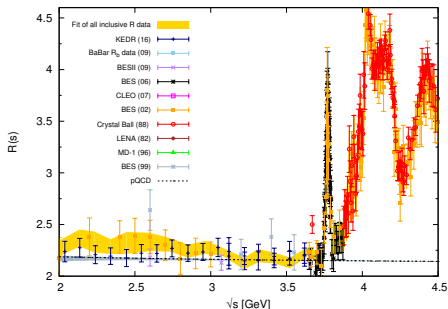
⇒ **New KEDR inclusive  $R$  data** ranging  $1.84 \leq \sqrt{s} \leq 3.05$  GeV [Phys.Lett. B770 (2017) 174-181]  
and  $3.12 \leq \sqrt{s} \leq 3.72$  GeV [Phys.Lett. B753 (2016) 533-541]



$a_{\mu}^{\text{had, LOVP}}(1.84 \leq \sqrt{s} \leq 2.00 \text{ GeV})$ :

pQCD :  $6.42 \pm 0.03$

Data :  $6.88 \pm 0.25$



$a_{\mu}^{\text{had, LOVP}}(2.60 \leq \sqrt{s} \leq 3.73 \text{ GeV})$ :

pQCD (inflated errors) :  $10.82 \pm 0.38$

Data :  $11.20 \pm 0.14$

⇒ **Choose to adopt entirely data driven estimate from threshold to 11.2 GeV**

# KNT17 $a_\mu^{\text{had, VP}}$ and $\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$ update [preliminary]

$$\underline{(g-2)_\mu}$$

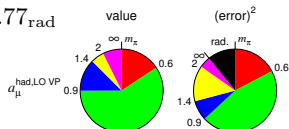
$$\text{HLMNT(11): } 694.91 \pm 4.27$$

$$\text{This work: } a_\mu^{\text{had, LOVP}} = 692.23 \pm 1.26_{\text{stat}} \pm 2.02_{\text{sys}} \pm 0.31_{\text{vp}} \pm 0.70_{\text{fsr}}$$

$$= 692.23 \pm 2.42_{\text{exp}} \pm 0.77_{\text{rad}}$$

$$= 692.23 \pm 2.54_{\text{tot}}$$

$$a_\mu^{\text{had, NLOVP}} = -9.83 \pm 0.04_{\text{tot}}$$



$$\underline{\Delta\alpha(M_Z^2)}$$

$$\text{HLMNT11: } (276.26 \pm 1.38_{\text{tot}}) \times 10^{-4}$$

$$\text{This work: } \Delta\alpha_{\text{had}}^{(5)}(M_Z^2) = (276.06 \pm 0.39_{\text{stat}} \pm 0.64_{\text{sys}} \pm 0.08_{\text{vp}} \pm 0.82_{\text{fsr}}) \times 10^{-4}$$

$$= (276.06 \pm 0.76_{\text{exp}} \pm 0.83_{\text{rad}}) \times 10^{-4}$$

$$= (276.06 \pm 1.13_{\text{tot}}) \times 10^{-4}$$

$\Rightarrow$  Accuracy better than 0.4%  
(uncertainties include all available correlations)

Full KNT17 VP package  
[vp\_knt\_v3\_0.f]  
available for use

# KNT17 vs. DHMZ17 vs. FJ17 [preliminary]

⇒ Different data treatment/methods produce **very different results**

Channel $\sqrt{s} \leq 1.8$ GeV	KNT17	DHMZ17	FJ17
$\pi^+\pi^-$	$502.73 \pm 1.94$	$507.14 \pm 2.58$	
$\pi^+\pi^-2\pi^0$	$17.80 \pm 0.99$	$18.03 \pm 0.55$	
$2\pi^+2\pi^-$	$14.00 \pm 0.19$	$13.68 \pm 0.31$	
$K^+K^-$	$22.70 \pm 0.25$	$22.81 \pm 0.41$	
$K_S^0K_L^0$	$13.08 \pm 0.14$	$12.81 \pm 0.24$	
Total HVP $\sqrt{s} < \infty$ GeV	$692.23 \pm 2.54$	$693.1 \pm 3.4$	$689.43 \pm 3.25$

⇒ Between  $1.8 \leq \sqrt{s} \leq 2$  GeV, KNT use **data**, DHMZ use **pQCD**

BUT, **pQCD** =  $8.30 \pm 0.09$ , **KNT data** =  $8.42 \pm 0.29$ , **DHMZ data** =  $7.71 \pm 0.32$

⇒ **DHMZ17 may not use correlated systematics in determination of the mean value**

→ Determining  $\pi^+\pi^-$  **using only local weighted average** gives  $508.91 \pm 2.84$

→ **Much better agreement** when neglecting the effect of correlated uncertainties on the mean value



# KNT17 $a_\mu^{\text{SM}}$ update [preliminary]

	<u>2011</u>	→	<u>2017</u>	
QED	11658471.81 (0.02)	→	11658471.90 (0.01)	[Phys. Rev. Lett. 109 (2012) 111808]
EW	15.40 (0.20)	→	15.36 (0.10)	[Phys. Rev. D 88 (2013) 053005]
LO HLbL	10.50 (2.60)	→	9.80 (2.60)	[EPJ Web Conf. 118 (2016) 01016]
NLO HLbL			0.30 (0.20)	[Phys. Lett. B 735 (2014) 90]

	<u>HLMNT11</u>	→	<u>KNT17</u>	
LO HVP	694.91 (4.27)	→	692.23 (2.54)	this work
NLO HVP	-9.84 (0.07)	→	-9.83 (0.04)	this work

NNLO HVP			1.24 (0.01)	[Phys. Lett. B 734 (2014) 144]
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Theory total	11659182.80 (4.94)	→	11659181.00 (3.62)	this work
Experiment			11659209.10 (6.33)	world avg
Exp - Theory	26.1 (8.0)	→	28.1 (7.3)	this work

$\Delta a_\mu$	3.3 $\sigma$	→	3.9 $\sigma$	this work
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# Conclusions

- ✓ Many necessary changes made since HLMNT11 in order to improve data combination
- ✓ KLOE combination covariance matrices correctly constructed to incorporate all available correlations
- ⇒ When combining data...
  - ✓ ...all covariance matrices are correctly constructed with a framework that can accommodate any available correlations
  - ✓ ...employ a linear  $\chi^2$  minimisation that has been shown to be free from bias
- ✓ New method shows improvements in all channels due to increased fit flexibility
- ✓ Less reliance on isospin for estimated states with more measured final states
- ✓  $a_{\mu}^{\text{had,LOVP}}$  accuracy better than 0.4%
- ✓ Differences comparing with other analyses due to different treatment and combination of data

# Extra Slides

# Vacuum polarisation corrections

- ⇒ Fully updated, self-consistent VP routine: [vp\_knt\_v3\_0]
  - Cross sections undressed with full photon propagator (must include imaginary part),  $\sigma_{\text{had}}^0(s) = \sigma_{\text{had}}(s)|1 - \Pi(s)|^2$
- ⇒ Applied to all dressed experimental data in all channels
  - Accurate to  $\mathcal{O}(1\text{‰})$  precision
- ⇒ If correcting data, apply corresponding radiative correction uncertainty
  - Take  $\frac{1}{3}$  of total correction per channel as conservative extra uncertainty
- ⇒ Influence/need for VP corrections has changed over time
  - Less prominent in some dominant channels
- ⇒ Undressing of narrow resonances must be done *excluding the contribution from the resonance*
  - ...or would double count contribution

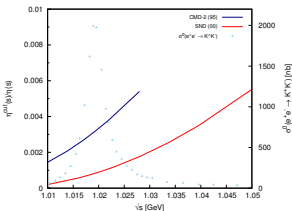
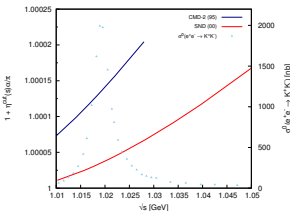
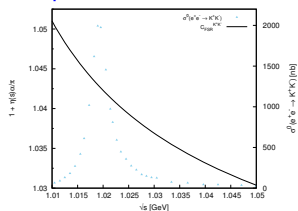
# Final state radiation corrections

⇒ For  $\pi^+\pi^-$ , FSR more frequently included

→ If not, must include through sQED approximation [Eur. Phys. J. C 24 (2002) 51,

Eur. Phys. J. C 28 (2003) 261]

⇒ For  $K^+K^-$ , is there available phase space for the creation of hard photons?



⇒ Choose to no longer apply FSR correction for  $K^+K^-$

⇒ For higher multiplicity states, difficult to estimate correction

∴ Apply conservative uncertainty

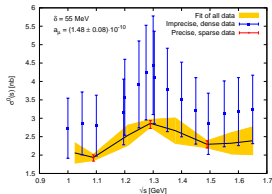
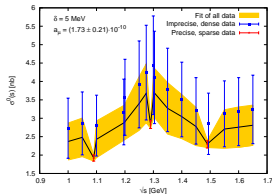
**Need new, more developed tools to increase precision here**

(e.g. - CARLOMAT 3.1 [Eur.Phys.J. C77 (2017) no.4, 254]?)

# Clustering data

⇒ Re-bin data into *clusters*

Better representation of data combination through adaptive clustering algorithm



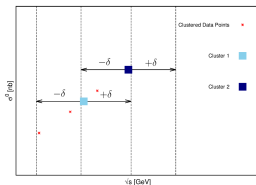
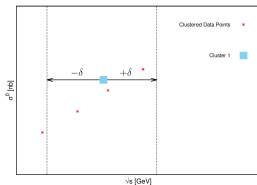
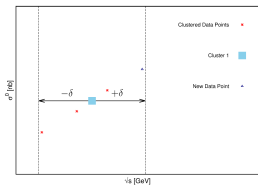
→ More and more data ⇒ risk of **over clustering**

⇒ loss of information on resonance

→ Scan cluster sizes for **optimum solution** (error,  $\chi^2$ , check by sight...)

⇒ Scanning/**sampling by varying bin widths**

→ Clustering algorithm now **adaptive to points at cluster boundaries**



# Fixing the covariance matrix [JHEP 1005 (2010) 075, Eur.Phys.J. C75 (2015), 613]

⇒ Apply a procedure to **fix the covariance matrix**

$$C_I(i^{(m)}, j^{(n)}) = C^{\text{stat}}(i^{(m)}, j^{(n)}) + \frac{C^{\text{sys}}(i^{(m)}, j^{(n)})}{R_i^{(m)} R_j^{(n)}} R_m R_n ,$$

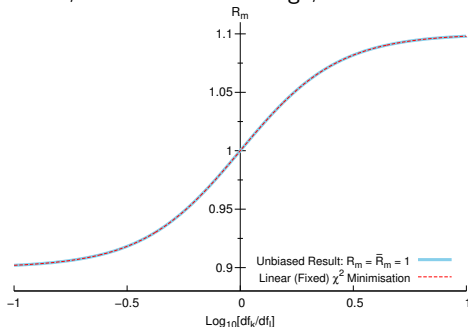
in an **iterative  $\chi^2$  minimisation** method that, to our best knowledge, is **free from bias**

⇒ Fixing with theory value **regulates influence**

⇒ Can be shown from toy models to be **free from bias**

⇒ **Swift convergence**

⇒ Comparison with past results shows **HLMNT11 estimates are largely unaffected**



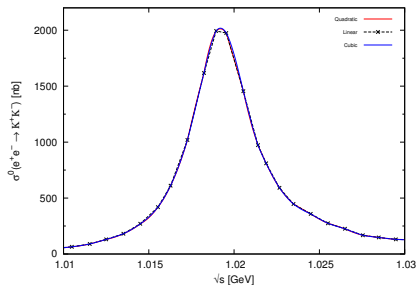
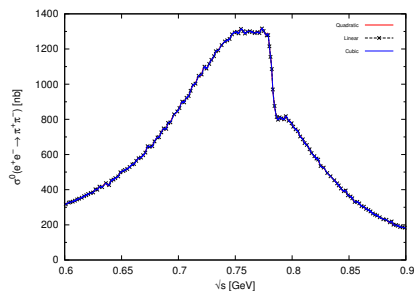
**Allows for increased fit flexibility and full use of energy dependent, correlated uncertainties**

# Integration

⇒ Trapezoidal rule integral

→ Consistency with linear cluster definition

→ High data population  $\therefore$  Accurate estimate from linear integral



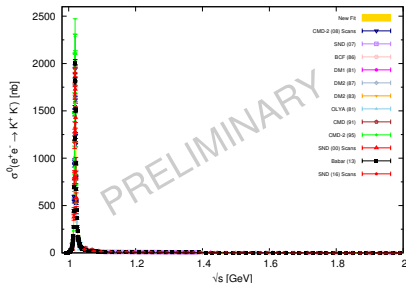
→ Higher order polynomial integrals give (at maximum) differences of  $\sim 10\%$  of error

⇒ Estimates of error non-trivial at integral borders

→ Extrapolate/interpolate covariance matrices



## Kaon FSR study



BUT  $K^+K^-$  cross section is totally dominated by  $\phi$  resonance

⇒ No phase space for creation of hard real photons at  $\phi$

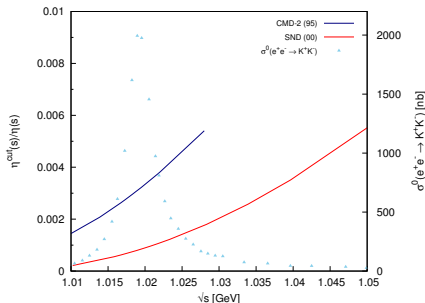
Inclusive FSR correction is large over-correction →

∴ No longer apply FSR correction

Inclusive FSR correction was previously applied to  $K^+K^-$  cross section

KLN theorem requires all virtual and soft corrections necessarily included in given cross section

∴ Only hard real radiation is left to be corrected for



## Properties of a covariance matrix

Any covariance matrix,  $\mathcal{C}_{ij}$ , of dimension  $n \times n$  must satisfy the following requirements:

- As the diagonal elements of any covariance matrix are populated by the corresponding variances, all the diagonal elements of the matrix are positive. Therefore, the trace of the covariance matrix must also be positive

$$\text{Trace}(\mathcal{C}_{ij}) = \sum_{i=1}^n \sigma_{ii} = \sum_{i=1}^n \text{Var}_i > 0$$

- It is a symmetric matrix,  $\mathcal{C}_{ij} = \mathcal{C}_{ji}$ , and is, therefore, equal to its transpose,  $\mathcal{C}_{ij} = \mathcal{C}_{ij}^T$
- The covariance matrix is a positive, semi-definite matrix,

$$\mathbf{a}^T \mathcal{C} \mathbf{a} \geq 0 ; \mathbf{a} \in \mathbf{R}^n,$$

where  $\mathbf{a}$  is an eigenvector of the covariance matrix  $\mathcal{C}$

- Therefore, the corresponding eigenvalues  $\lambda_{\mathbf{a}}$  of the covariance matrix must be real and positive and the distinct eigenvectors are orthogonal

$$\mathbf{b}^T \mathcal{C} \mathbf{a} = \lambda_{\mathbf{a}}(\mathbf{b} \cdot \mathbf{a}) = \mathbf{a}^T \mathcal{C} \mathbf{b} = \lambda_{\mathbf{b}}(\mathbf{a} \cdot \mathbf{b})$$

$$\therefore \text{if } \lambda_{\mathbf{a}} \neq \lambda_{\mathbf{b}} \Rightarrow (\mathbf{a} \cdot \mathbf{b}) = 0$$

- The determinant of the covariance matrix is positive:  $\text{Det}(\mathcal{C}_{ij}) \geq 0$

# Tests of reliability of $f_k$ method

## Did the $f_k$ method incur a bias?

Compare  $f_k$  method and fixed matrix method with **only multiplicative normalisation uncertainties**.

→ If we see **differences** in mean value, then **bias previously influenced the fit**.

→ **Previous results unreliable**

→ If we see **no differences** in mean value, then **bias did not influence fit** (any change comes from improved treatment of systematics)

→ **Previous results reliable**

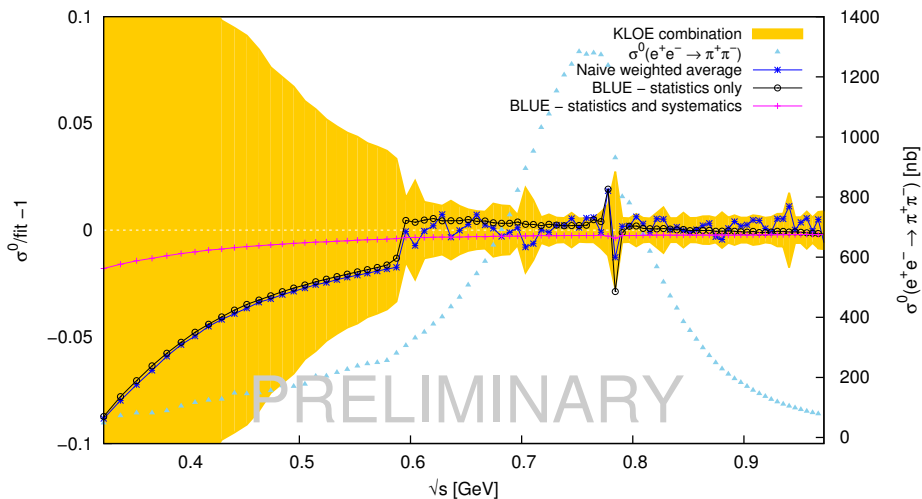
*Example -  $\pi^+\pi^-$*

Set 1 - CMD-2(06) (0.7% Systematic Uncertainty), Set 2 - CMD-2(06) (0.8% Systematic Uncertainty), Set 3 - SND(04) (1.3% Systematic Uncertainty)

From 0.37 → 0.97 GeV

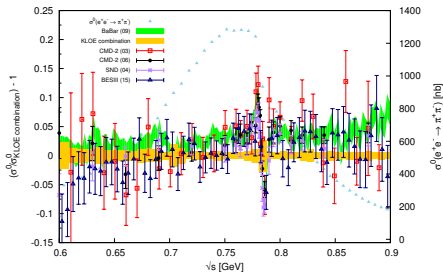
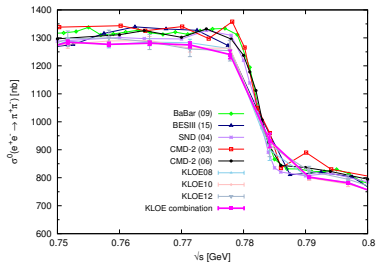
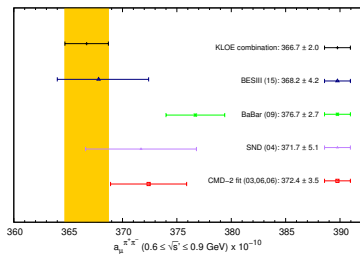
Fit Method:	$f_k$ method		Fixed matrix method		
Channel	$a_\mu$	$\chi^2_{\min}/\text{d.o.f.}$	$a_\mu$	$\chi^2_{\min}/\text{d.o.f.}$	Difference
$\pi^+\pi^-$	$481.42 \pm 4.26$	1.10	$481.42 \pm 4.05$	1.02	0.00

## Comparison of KLOE combination methods [preliminary]

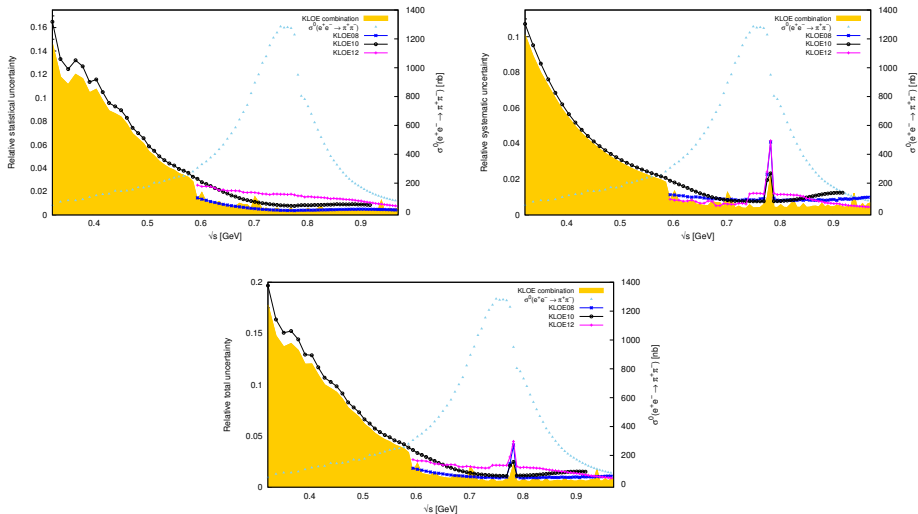


# Comparison of KLOE combination with other experiments

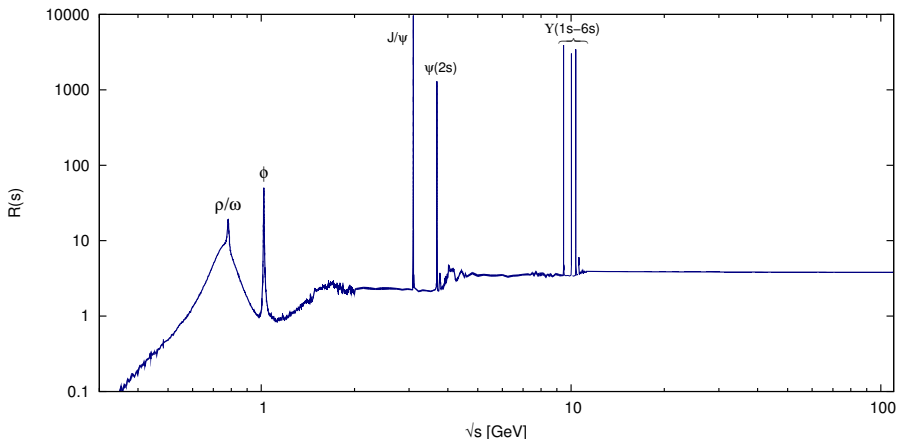
[preliminary]



## KLOE combination errors [preliminary]



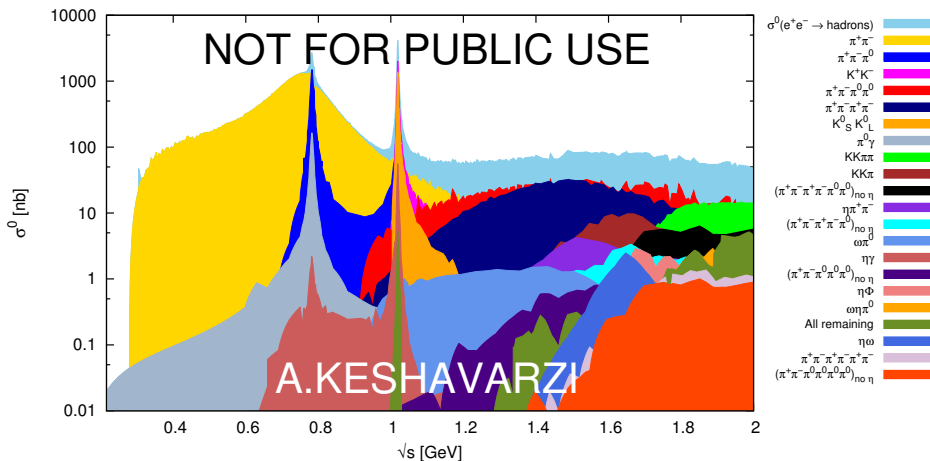
$$R(s) \text{ for } m_\pi \leq \sqrt{s} < \infty$$



⇒ Full compilation data set for hadronic  $R$ -ratio to be made available soon...

⇒ ...complete with full covariance matrix

## Contributions to mean value below 2GeV





## Contributions to uncertainty below 2GeV

