

Detection of Microwave Photons with SC Qubits for Axion Searches

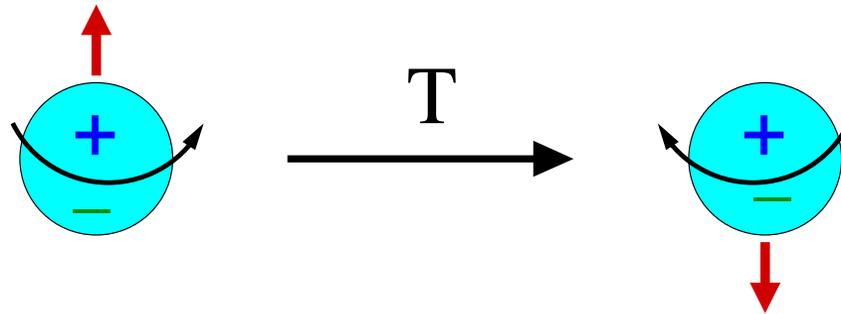
C.Gatti – LNF INFN

F.Chiarello – CNR-IFN

Axion: CP violation in QCD

$$\mathcal{L}_{QCD} = \dots + \theta \frac{\alpha_s}{8\pi} G_{a\mu\nu} \tilde{G}_a^{\mu\nu}$$

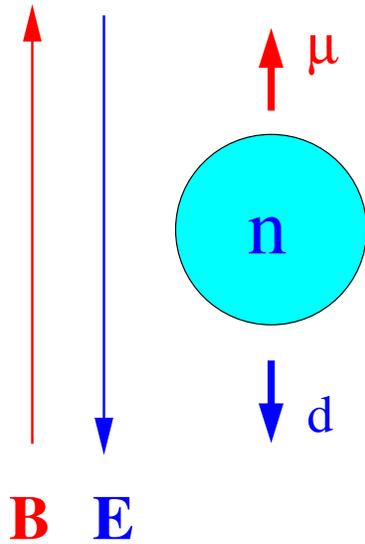
This term violates T (CP) symmetries and induces a neutron electric dipole moment (EDM)



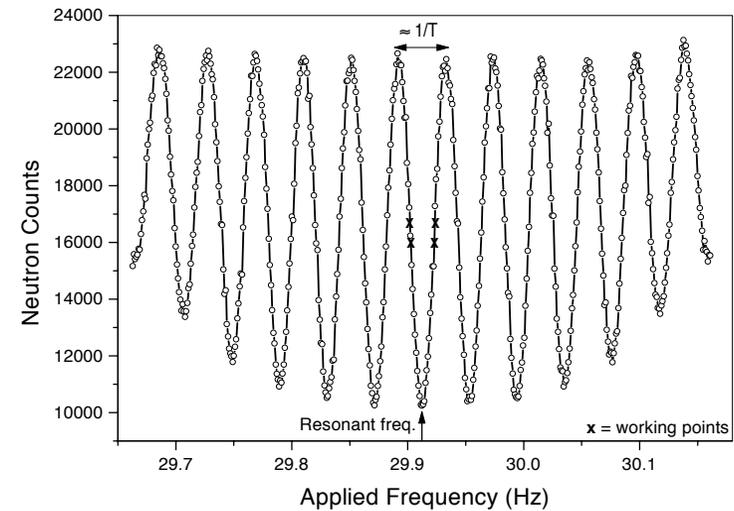
$$d_n \simeq \theta \frac{m_q}{M_n} \frac{e}{M_n} \sim \theta \times 10^{-3} e \text{ GeV}^{-1} \sim \theta \times 10^{-15} e \text{ cm}$$

Axion: Neutron EDM

$$h\nu = |2\mu_n B \pm 2d_n E|$$



PRL 82(5) (1999) p.904B

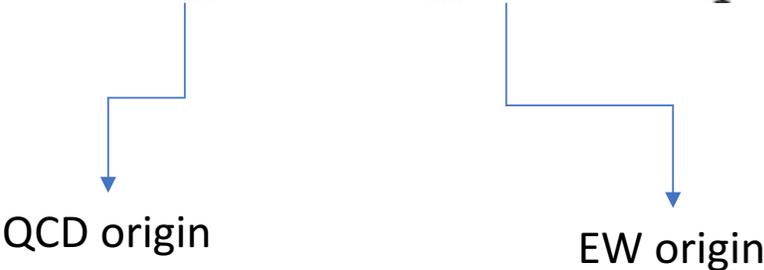


$$d_n < 2.9 \times 10^{-26} e \text{ cm}$$

$$\theta < 10^{-10}$$

Axion: $\theta=0$?

Diagonalization of quark mass matrix (M_q) has the effect of shifting the value of θ :

$$\theta = \theta_0 + \arg \det M_q$$


The diagram shows the equation $\theta = \theta_0 + \arg \det M_q$. Below θ_0 , a blue arrow points down to the text "QCD origin". Below $\arg \det M_q$, a blue arrow points down to the text "EW origin".

Strong CP problem: why the combination of QCD and EW parameters is so small?

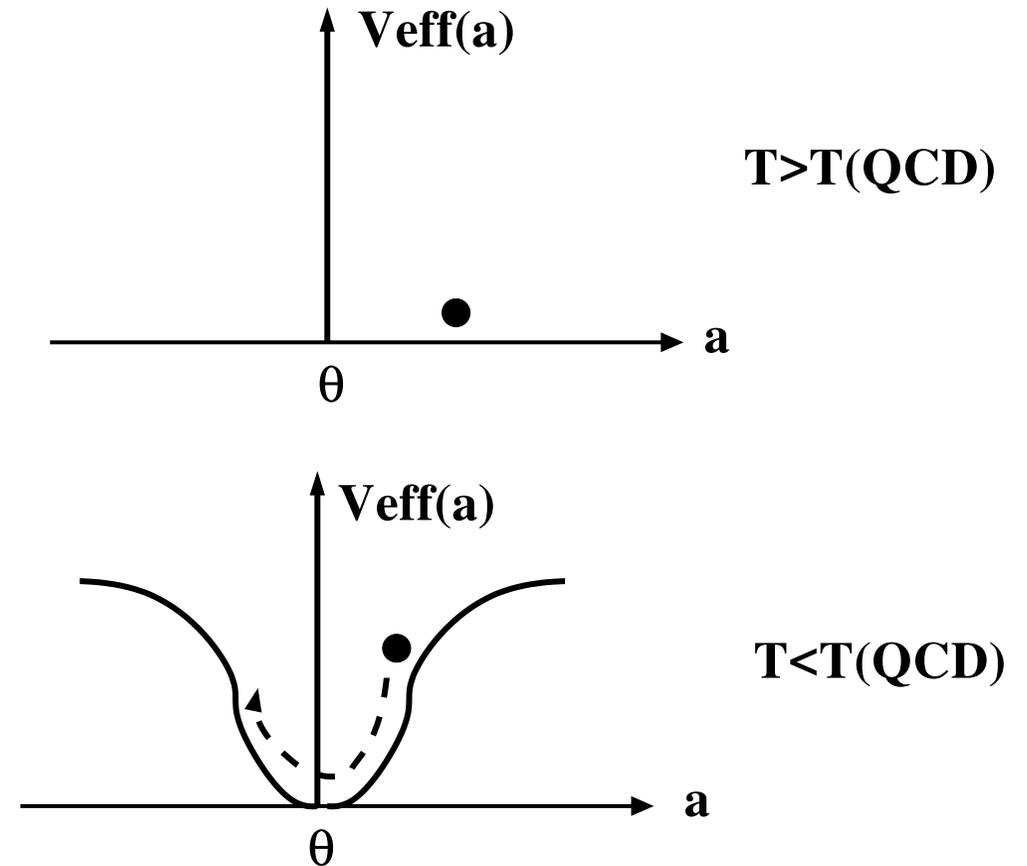
Axion: The solution

$$\mathcal{L}_{QCD} = \dots + \frac{\alpha_s}{8\pi} \left(\theta - \frac{a}{f_a} \right) G_{a\mu\nu} \tilde{G}_a^{\mu\nu}$$

“a” is a new scalar field.

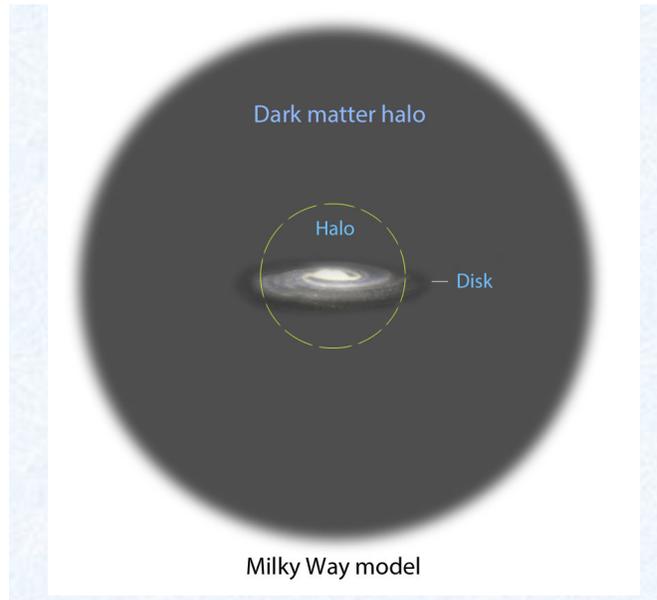
When the temperature approaches the QCD scale the potential V_{eff} turns on and the axion acquires a mass.

(Asztalos Ann.Rev.Part. Sci. 2006 56: 293-326)



Peccei Quinn Weinberg Wilczek

Axions: CDM



$$\rho \simeq 0.3 \text{ GeV}/\text{cm}^3$$

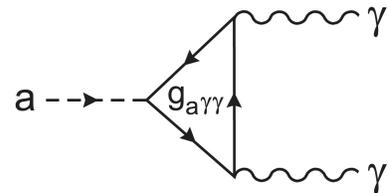
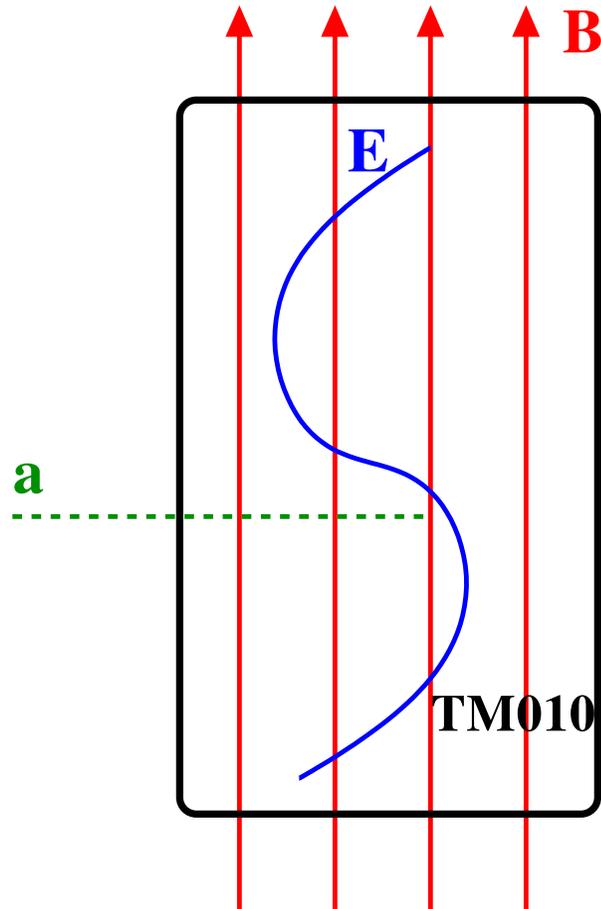
$$n_a \simeq 3 \times 10^{12} \left(\frac{100 \mu\text{eV}}{m_a} \right) 1/\text{cm}^3$$

$$\beta_a \sim 10^{-3}$$

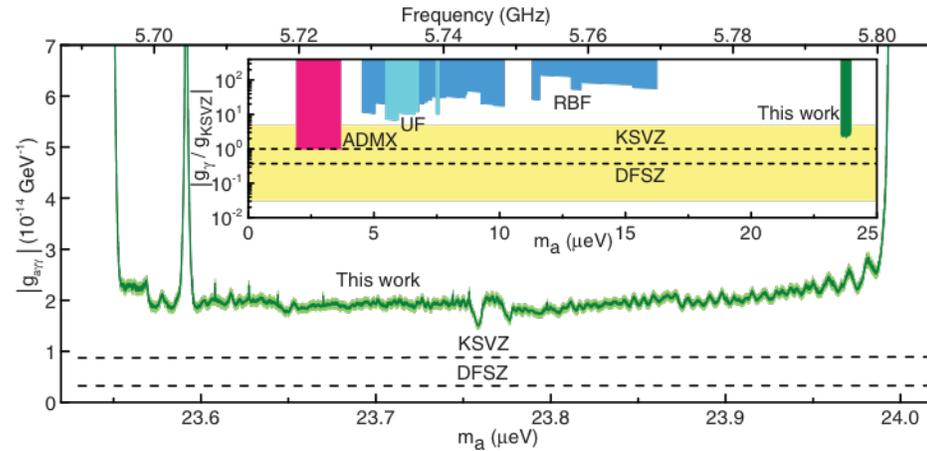
$$\frac{\delta\omega}{\omega} \sim 10^{-6} \quad \frac{\omega}{2\pi} = 24 \left(\frac{m_a}{100 \mu\text{eV}} \right) \text{GHz}$$

$$a = a_0 \cos \omega t$$

Galactic Axion Searches: ADMX-HF



$$P_S = \left(g_\gamma^2 \frac{\alpha^2}{\pi^2} \frac{\rho_a}{\Lambda^4} \right) \left(\omega_c B_0^2 V C_{mnl} Q_L \frac{\beta}{1+\beta} \right)$$

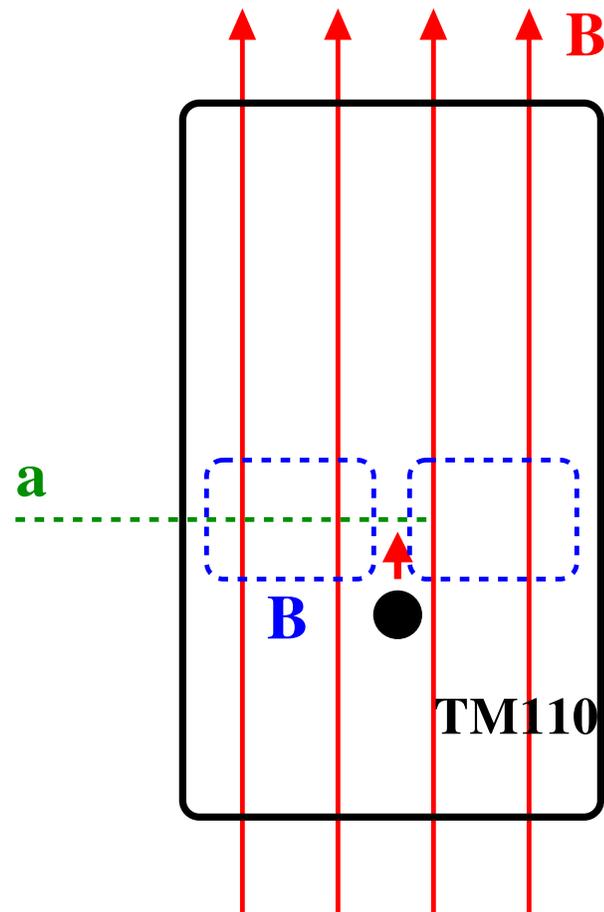


Phys. Rev. Lett. **118**, 061302

$$P_S \simeq 5 \times 10^{-24} \text{ W}$$

$$R_S \simeq 1 \text{ Hz}$$

Galactic Axion Searches: Quax



Electron coupling

$$\frac{g_p \hbar}{2m} \sigma \cdot \nabla a$$

$$P_S \sim 5 \times 10^{-26} \text{ W}$$

$$R_S \sim 1 \times 10^{-3} \text{ Hz}$$

$$m_a = 200 \mu\text{eV}$$

Physics of the Dark Universe 15 (2017), 135-141

Amplifier vs photon counter

$$P_{min} = k_B T_S \sqrt{\frac{\Delta\nu}{\tau}} \text{ W} \quad \longrightarrow \quad P_{min} = \hbar\omega \sqrt{\frac{\Delta\nu}{\tau}} \text{ W}$$

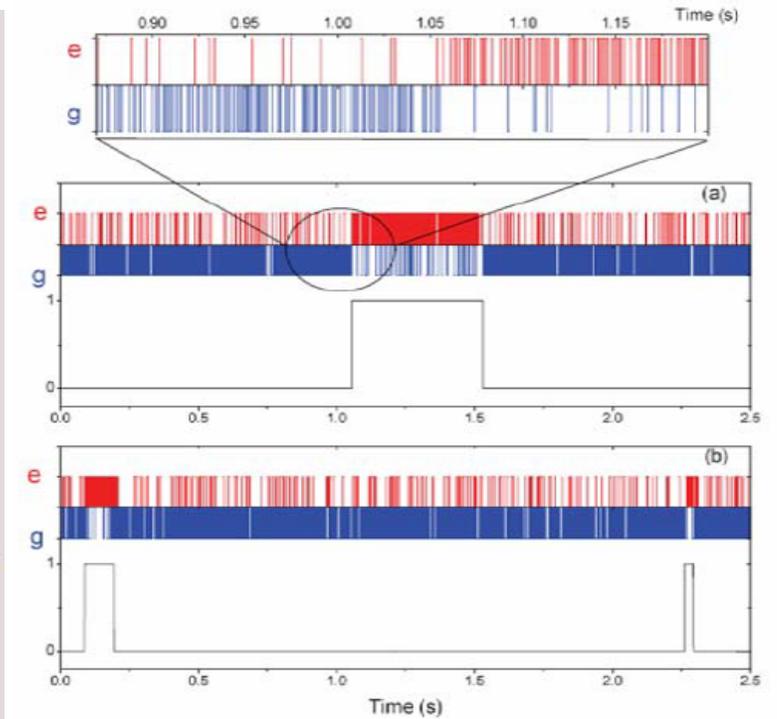
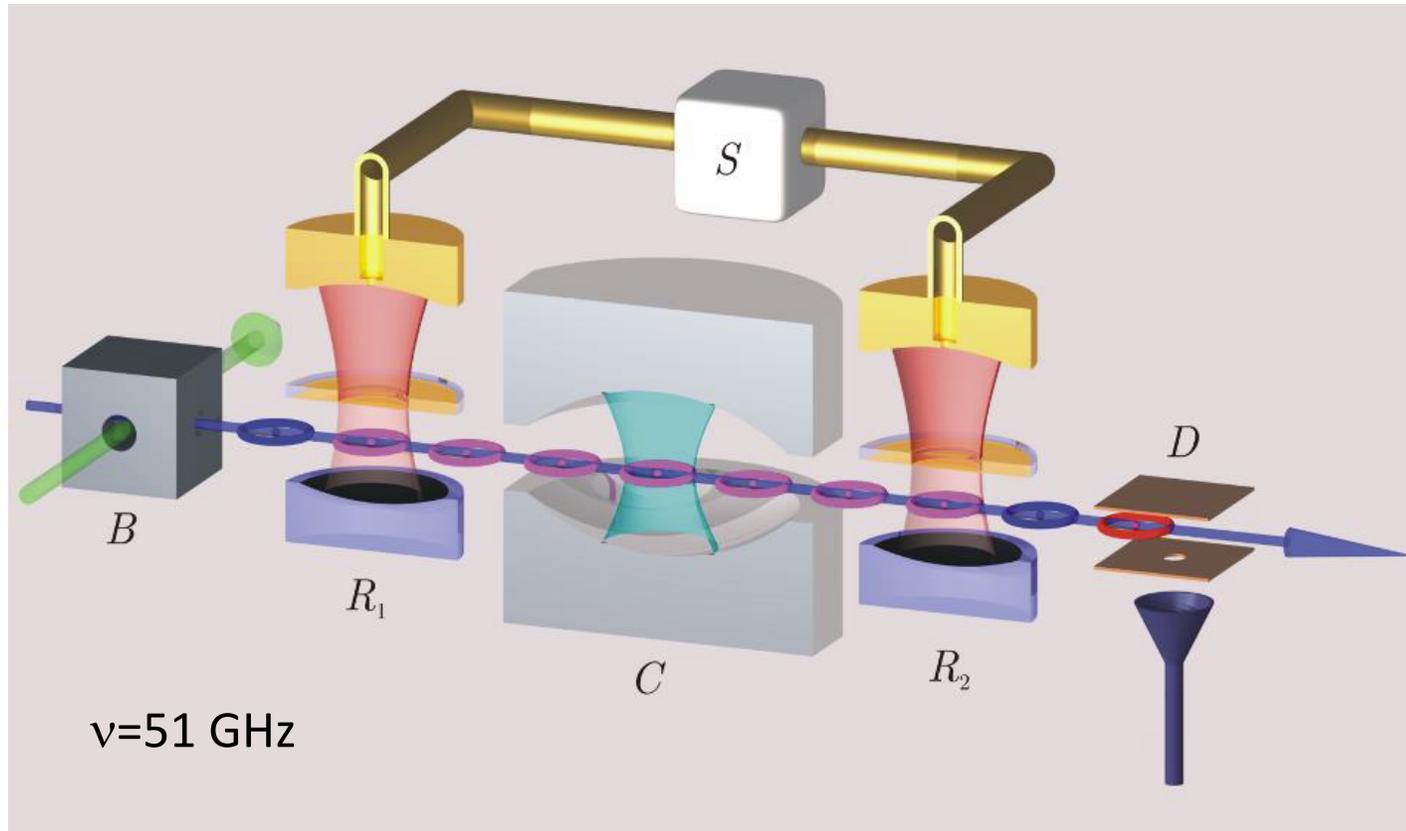
Linear amplifiers add noise to the signal, at least $\hbar\omega$. Caveas PR D26 (1982)

$$\tau \sim 10^4 \text{ s}$$

$$P_{min}^{Quax} = 4 \times 10^{-23} \text{ W}$$

Need a single photon counter to evade standard quantum limit (SQL)

Birth and Death of a Photon



Nature 446, 297-300 (2007)

Cavity QED

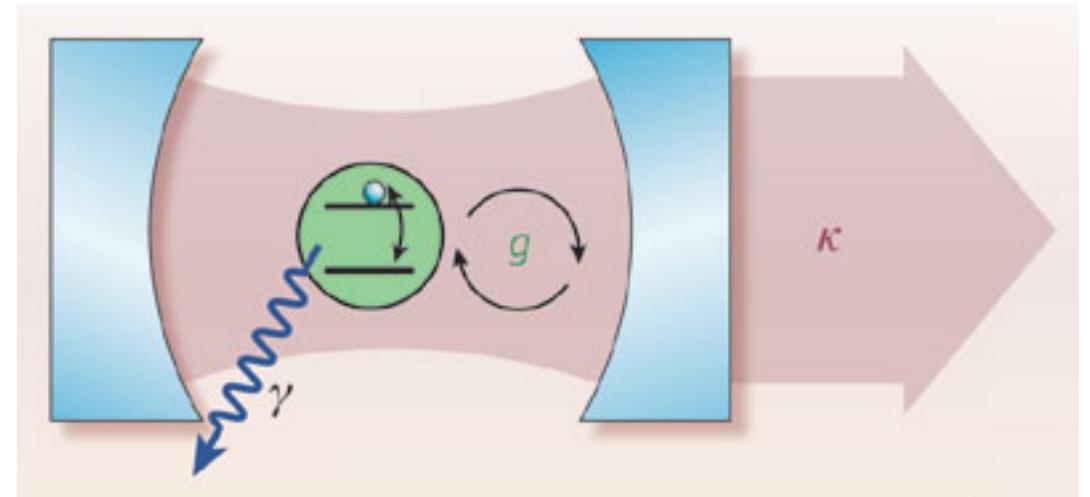
Two level atom interacting via dipole coupling with a cavity is described by the Jaynes-Cummings hamiltonian:

$$H_{JC} = \hbar\omega_c(a^\dagger a + 1/2) + \hbar\frac{\omega_a}{2}\sigma_z + \hbar g(a^\dagger\sigma^- + a\sigma^+)$$

Photons in cavity

Two level atom

Dipole interaction



D.I.Schuster "Circuit QED" PhD Thesis

Cavity QED

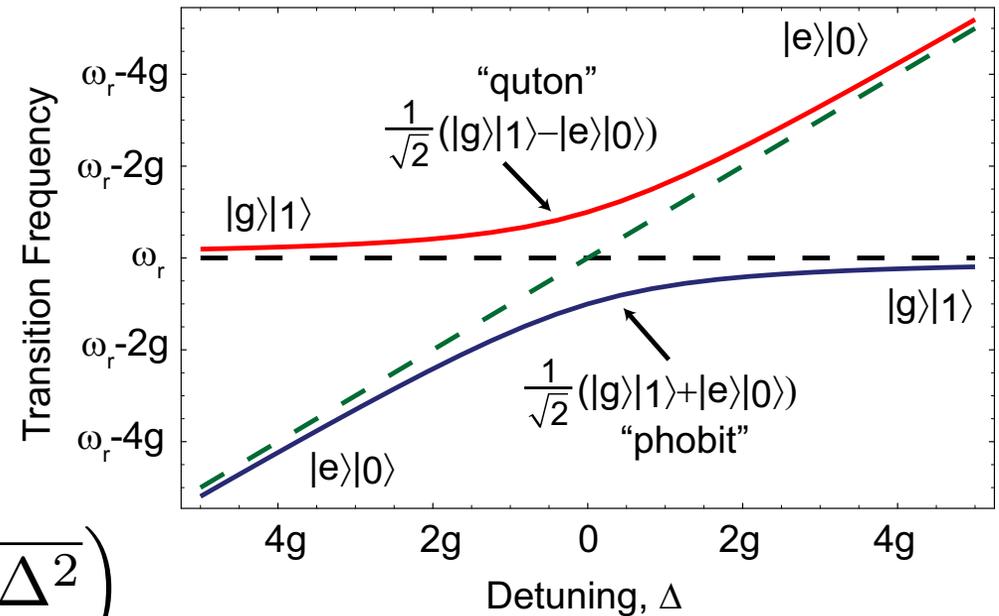
$$|-, n\rangle = \cos \theta_n |g, n\rangle - \sin \theta_n |e, n-1\rangle$$

$$|+, n\rangle = \sin \theta_n |g, n\rangle + \cos \theta_n |e, n-1\rangle$$

$$\theta_n = \frac{1}{2} \arctan \left(\frac{2g\sqrt{n}}{\Delta} \right)$$

$$\Delta = \omega_a - \omega_c$$

$$E_{\pm, n} = \frac{\hbar}{2} \left((n-1)\omega_c + \omega_a \pm \sqrt{4ng^2 + \Delta^2} \right)$$

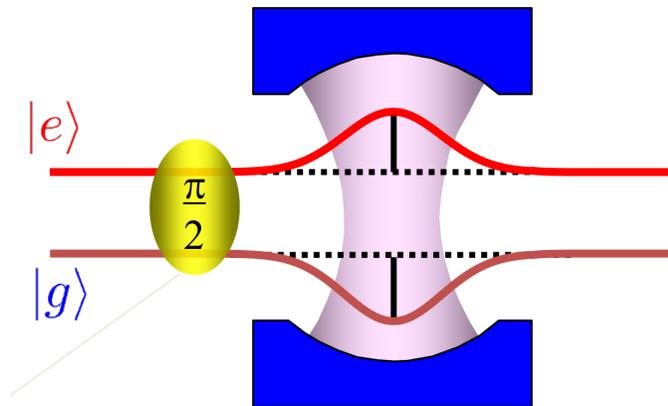


In the strong coupling limit ($\Delta \ll g$) photon is continuously absorbed and re-emitted inside the cavity with the atom oscillating between ground and excited state.

Cavity QED: Dispersive Limit ($\Delta \gg g$)

$$H_{JC} \simeq \hbar\omega_c(a^\dagger a + 1/2) + \frac{\hbar}{2} \left(\omega_a + \frac{2g^2}{\Delta} a^\dagger a + \frac{g^2}{\Delta} \right) \sigma_z$$

The electric field in the cavity produces a Stark-shift in the atom energy levels.



Dispersive limit allow quantum non destructive (QND) measurement of photons in the cavity

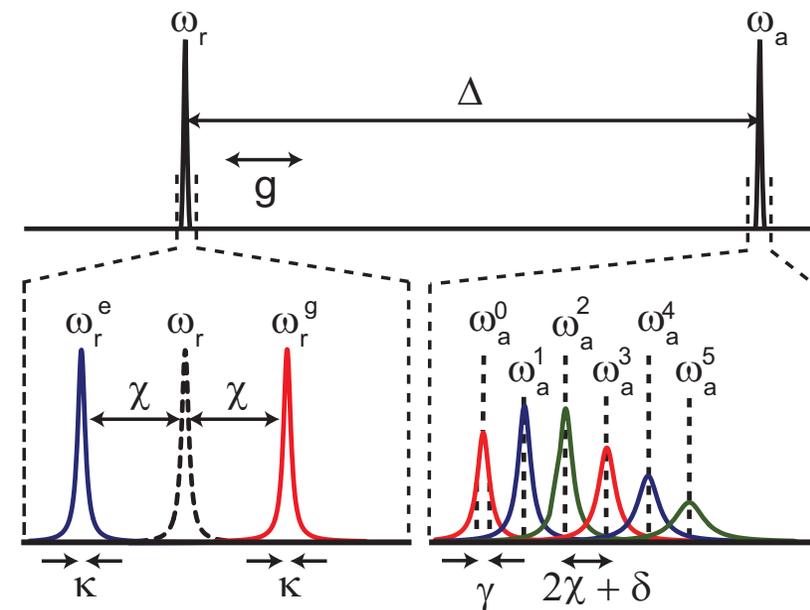
S.Deleglise <https://agenda.infn.it/conferenceOtherViews.py?view=standard&confId=12992>

Cavity QED: Strong Dispersive Limit

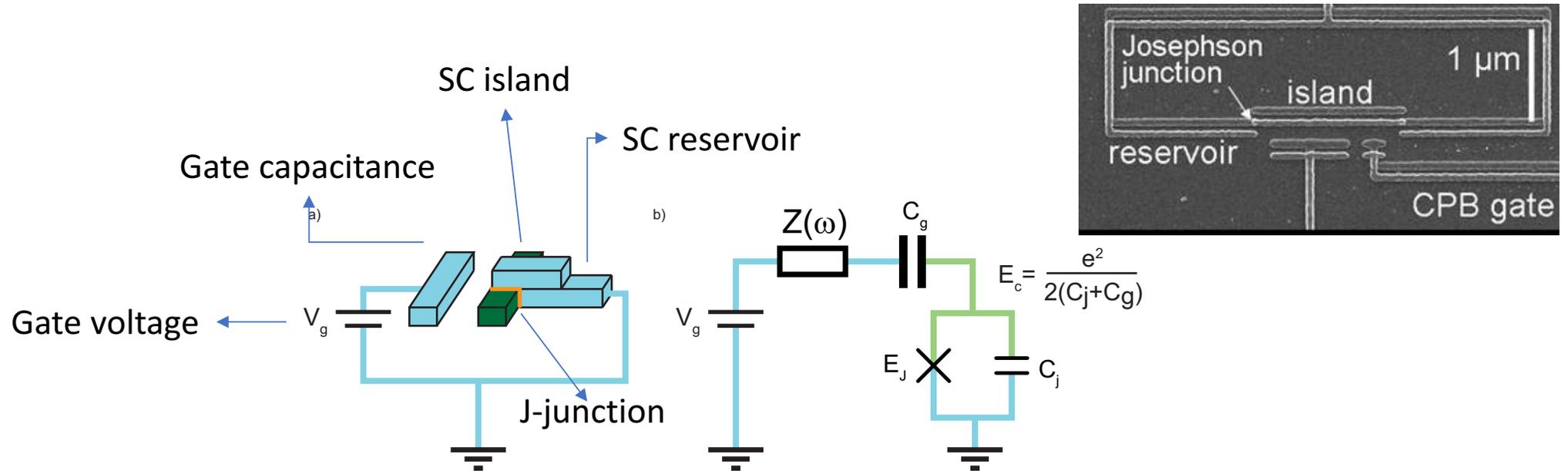
In the strong dispersive limit the energy shift is larger than natural width of atom (κ) and cavity mode (γ):

$$g^2 / \Delta \ll \gamma, \kappa$$

Photon number in the cavity are measured by performing spectroscopy on the atom to measure the energy shift.



Artificial atoms: The Cooper Pair Box



CPB consists of a SC island connected to a reservoir via a Josephson junction. The gate voltage induces tunneling of Cooper pairs.

V. Bouchiat Phys. Scrip. T76 (1998) 165-170

The Cooper Pair Box

$$H_{CPB} = 4E_C \sum_n (n - n_g/2)^2 |n\rangle\langle n| - \frac{E_J}{2} \sum_n (|n\rangle\langle n+1| + |n+1\rangle\langle n|)$$

n excess of Cooper pairs in the island

E_C electrostatic energy to charge a single electron

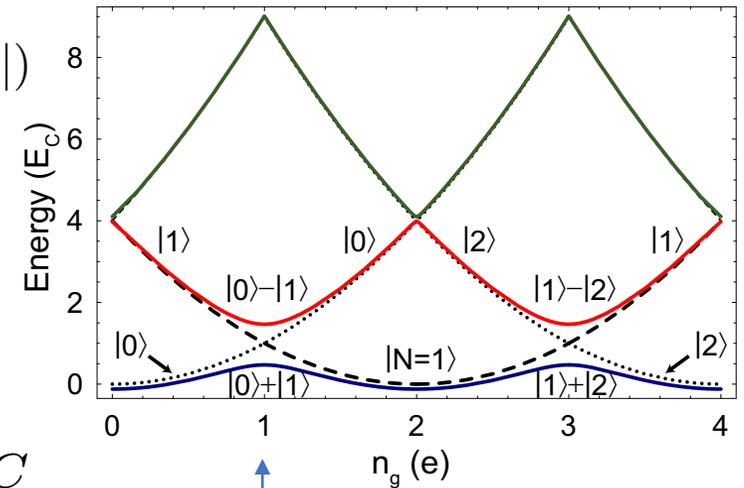
E_J Josephson energy

For $n_g=1$ the circuit behaves as a two state atom:

$$E_J \ll E_C$$

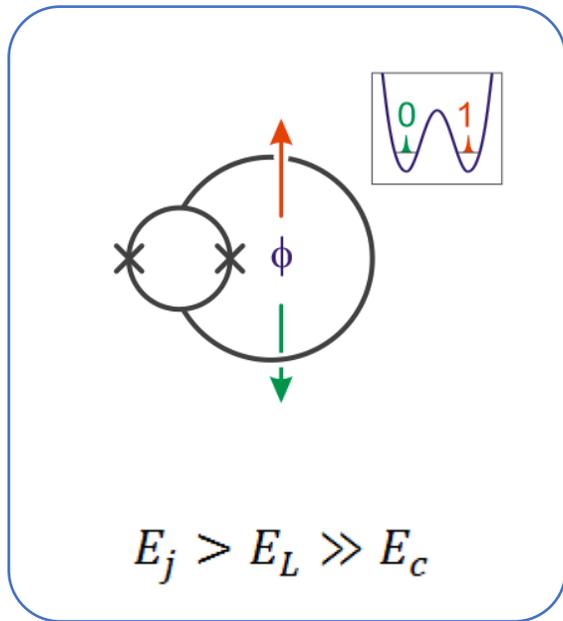
$$|+\rangle = (|0\rangle - |1\rangle) / \sqrt{2}$$

$$|-\rangle = (|0\rangle + |1\rangle) / \sqrt{2}$$

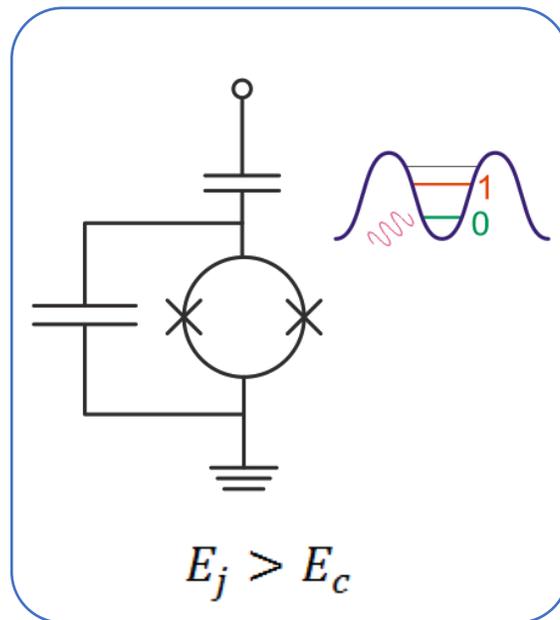


Superconducting Qubits

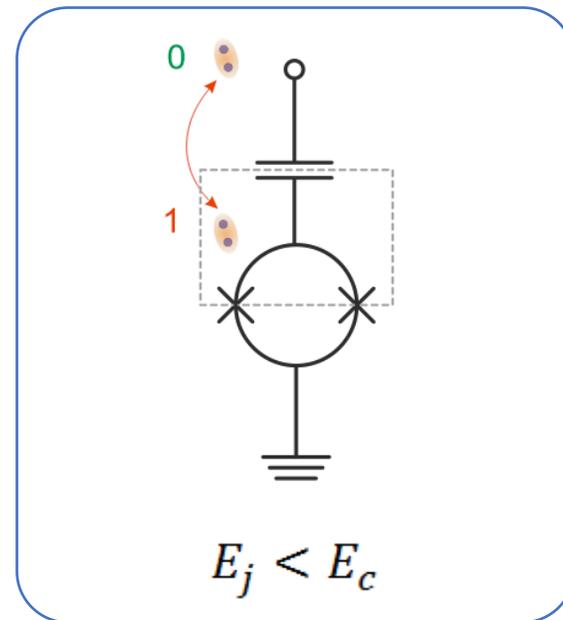
Flux Qubit



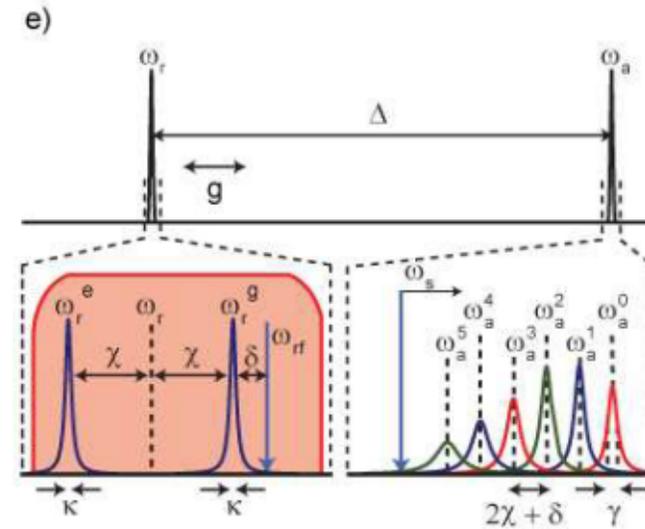
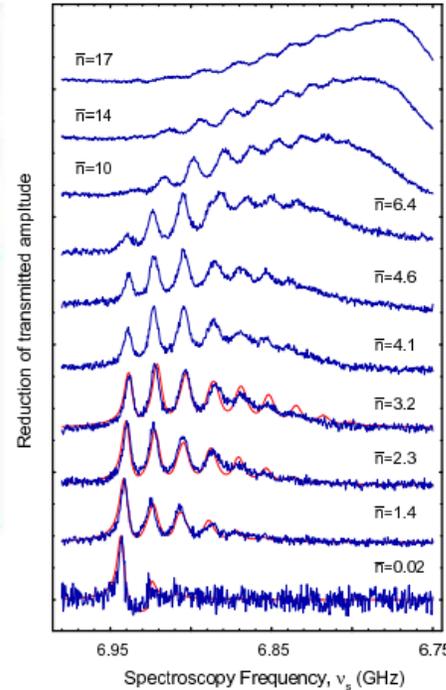
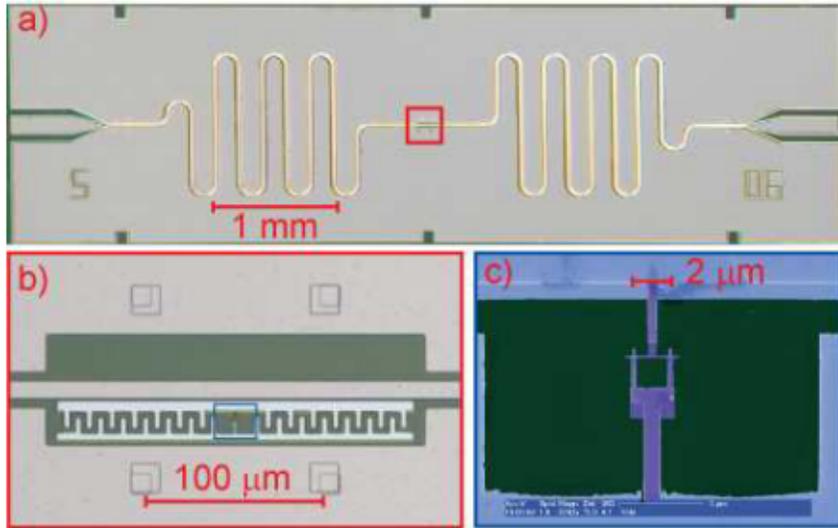
Transmon Qubit



Cooper Pair Box Qubit



Resolving Photon Number States in SC Circuit

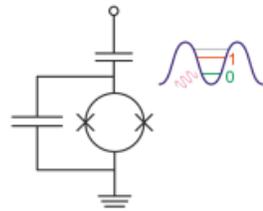


$$\nu_c = 5.7 \text{ GHz}$$

$$Q \approx 22,000$$

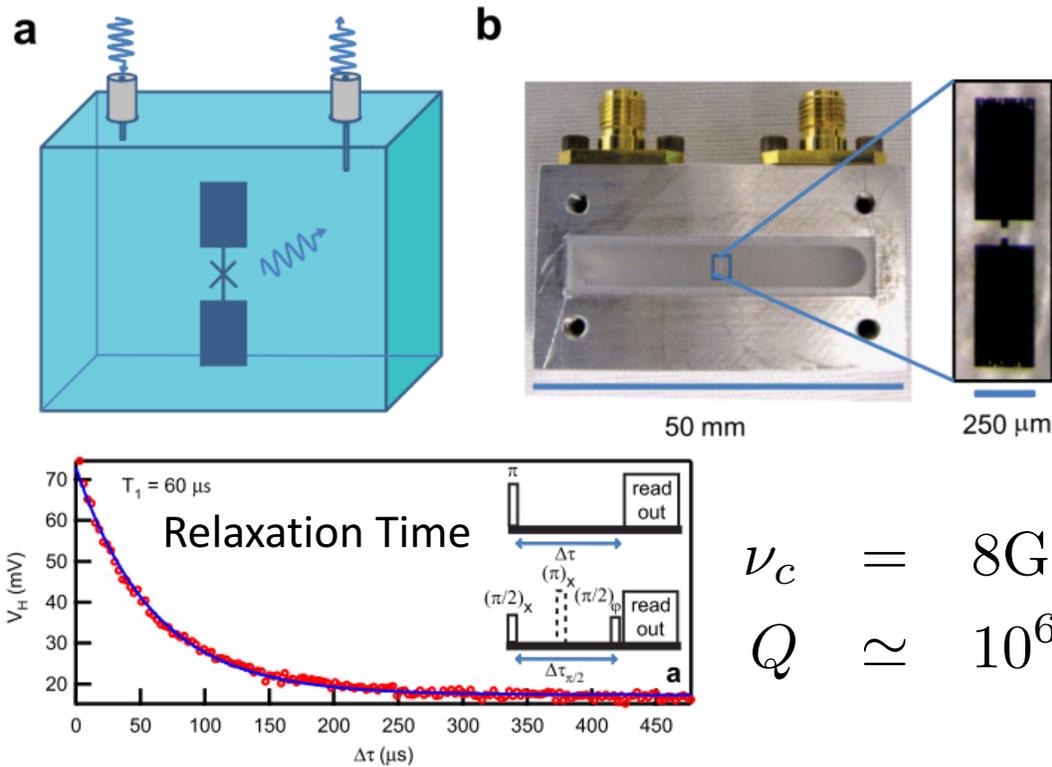
$$\tau_c \approx 4 \mu\text{s}$$

$$\tau_a \approx 0.5 \mu\text{s}$$



D. I. Schuster et al., *Nature* **445**, 515–518 (2007).

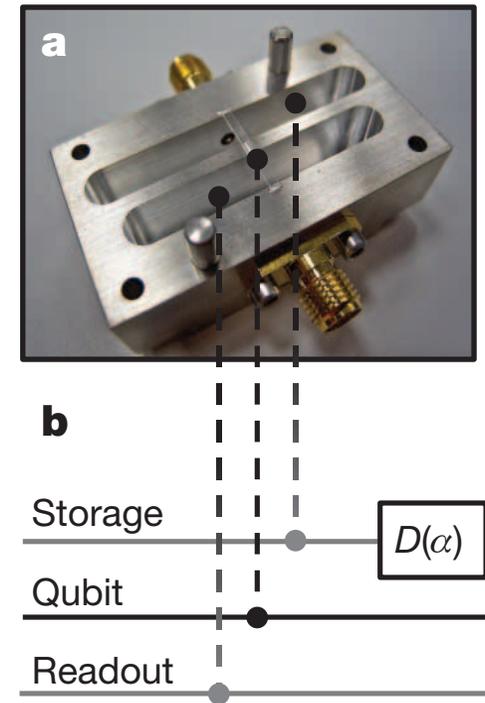
3D Cavities



$$\nu_c = 8\text{GHz}$$

$$Q \approx 10^6$$

H. Paik, Phys. Rev. Lett. **107** (2011).



L.Sun, Nature 511 (2014), 444

Conclusions

There is a rising interest in axions searches ... (no Susy at LHC?)

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Single photon counter in microwave regime required to reach ultimate sensitivity

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Single photon counter in microwave regime required to reach ultimate sensitivity

SC Qubits may be the solution ...

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Rising interest in axions searches ... (no Susy at LHC?)



Single photon counter in microwave regime required to reach ultimate sensitivity

Detecting Axion Dark Matter with Superconducting Qubits

SC Qubits may be the solution ... R&D already started for ADMX-HF!

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