

Nuclear Shapes

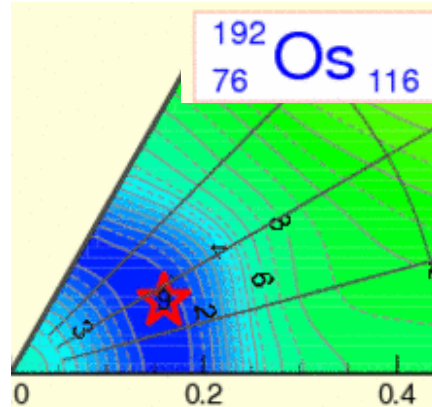
Lecture IV

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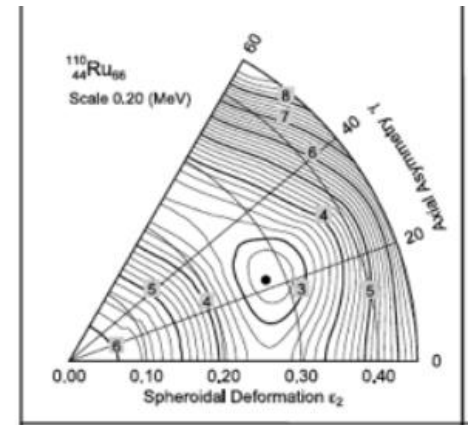
Selected Topics in Nuclear and Atomic Physics
Fiera di Primiero
2.-6. October 2017





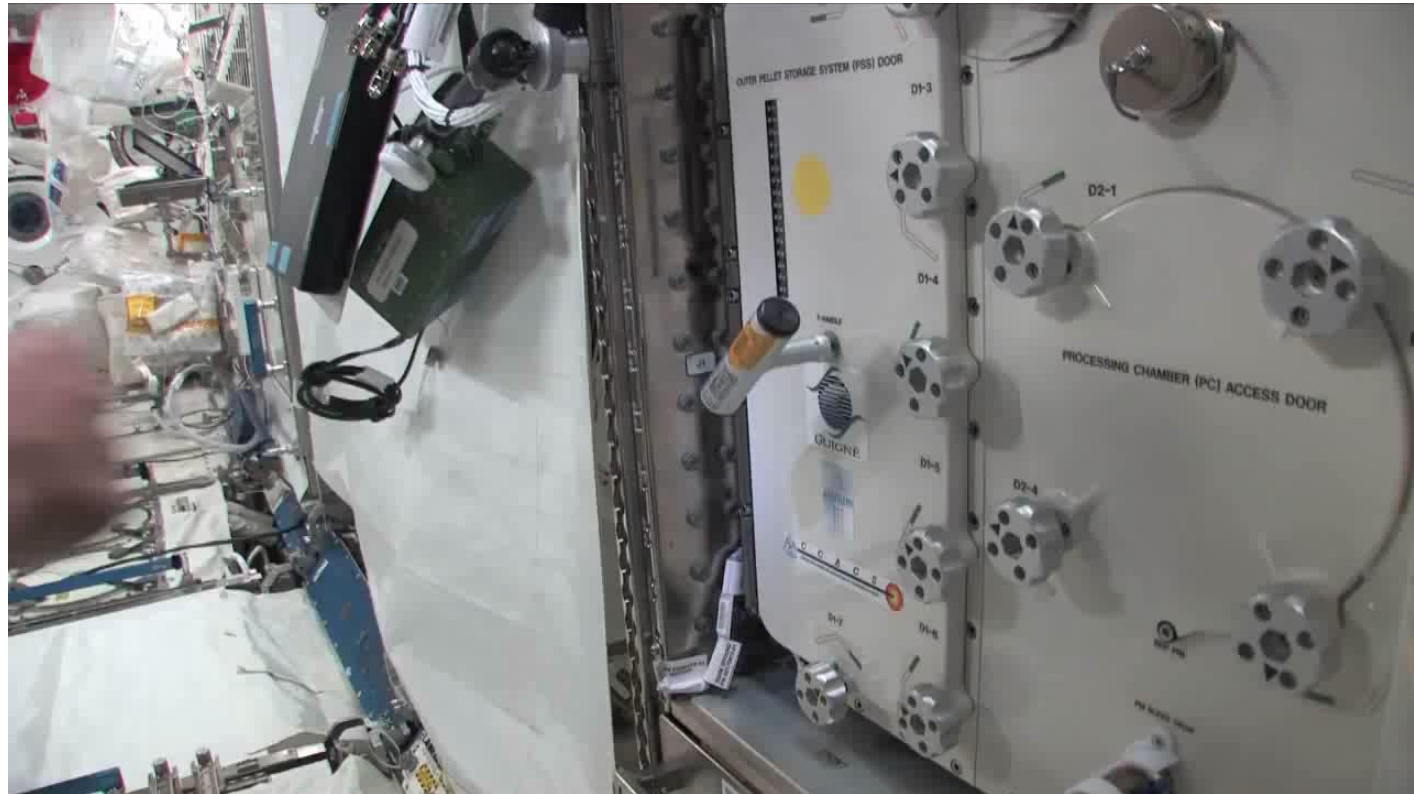


gamma-soft



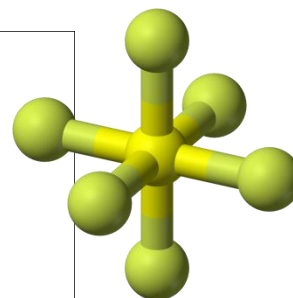
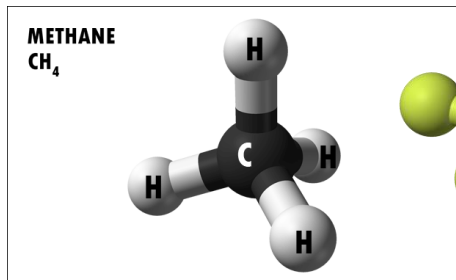
more gamma-rigid

what happens if we rotate a triaxial body?



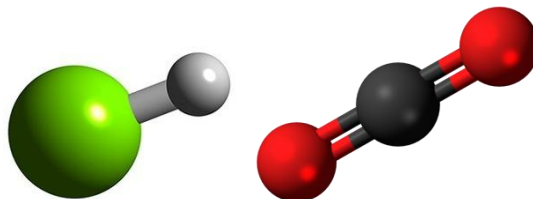
Molecular spectroscopy

spherical rotor: CH_4 , SF_6



$$\mathfrak{I}_a = \mathfrak{I}_b = \mathfrak{I}_c$$

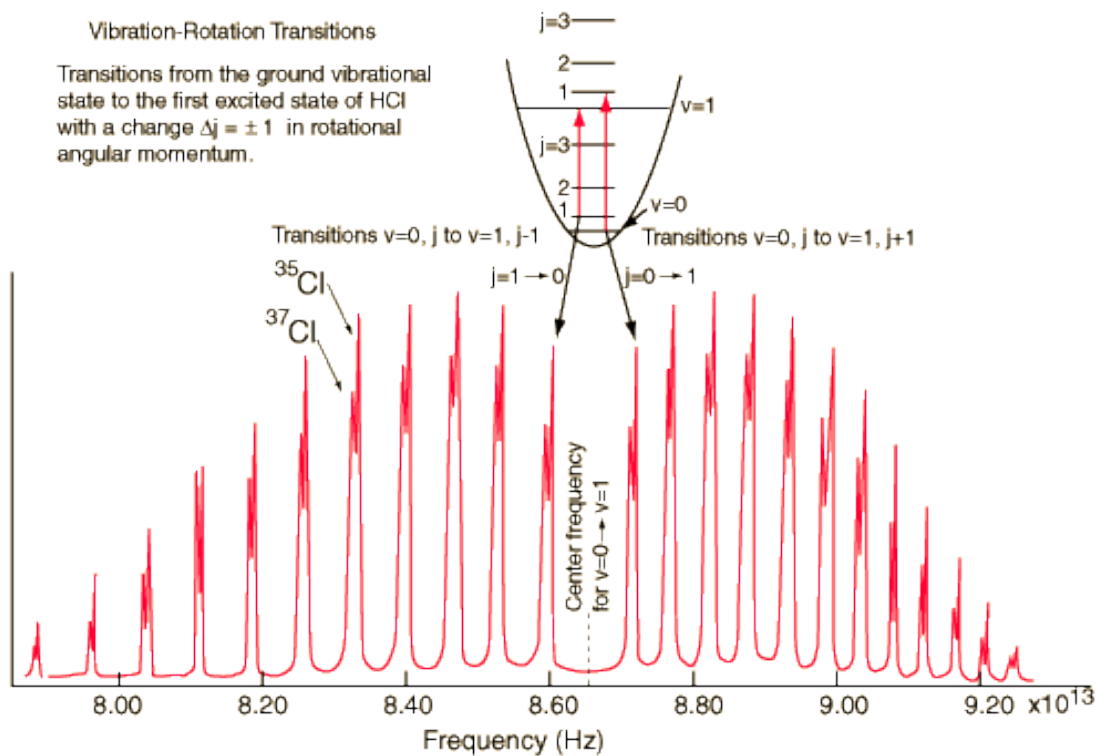
linear rotor: HCl , CO_2



$$\mathfrak{I}_a = 0, \quad \mathfrak{I}_b = \mathfrak{I}_c$$

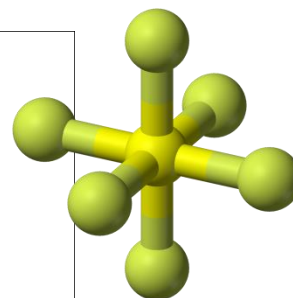
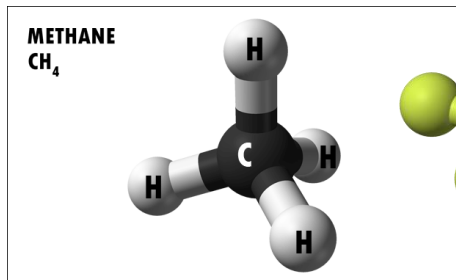
Vibration-Rotation Transitions

Transitions from the ground vibrational state to the first excited state of HCl with a change $\Delta j = \pm 1$ in rotational angular momentum.



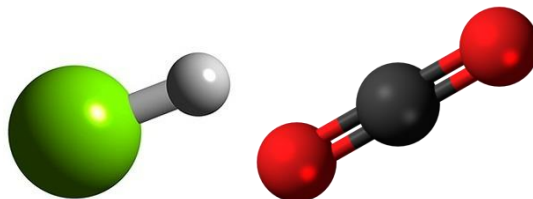
Molecular spectroscopy

spherical rotor: CH_4 , SF_6



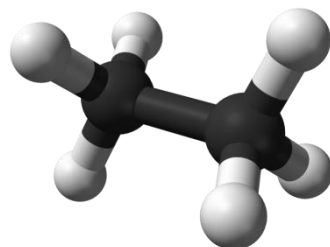
$$\mathfrak{I}_a = \mathfrak{I}_b = \mathfrak{I}_c$$

linear rotor: HCl , CO_2



$$\mathfrak{I}_a = 0, \quad \mathfrak{I}_b = \mathfrak{I}_c$$

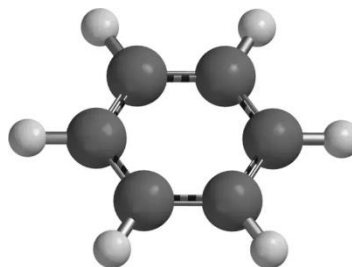
symmetric prolate rotor: C_2H_6



$$\mathfrak{I}_a < \mathfrak{I}_b = \mathfrak{I}_c$$

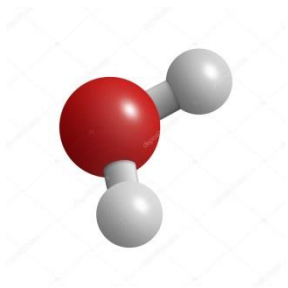
$$E_{J,K} = \frac{J(J+1)\hbar^2}{2\mathfrak{I}_b} + \left(\frac{1}{2\mathfrak{I}_a} - \frac{1}{2\mathfrak{I}_b} \right) K^2 \hbar^2$$

symmetric oblate rotor: C_6H_6

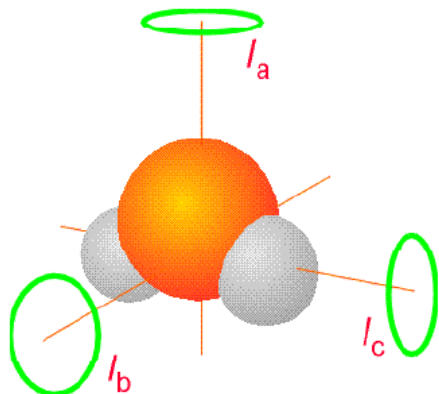


$$\mathfrak{I}_a = \mathfrak{I}_b < \mathfrak{I}_c$$

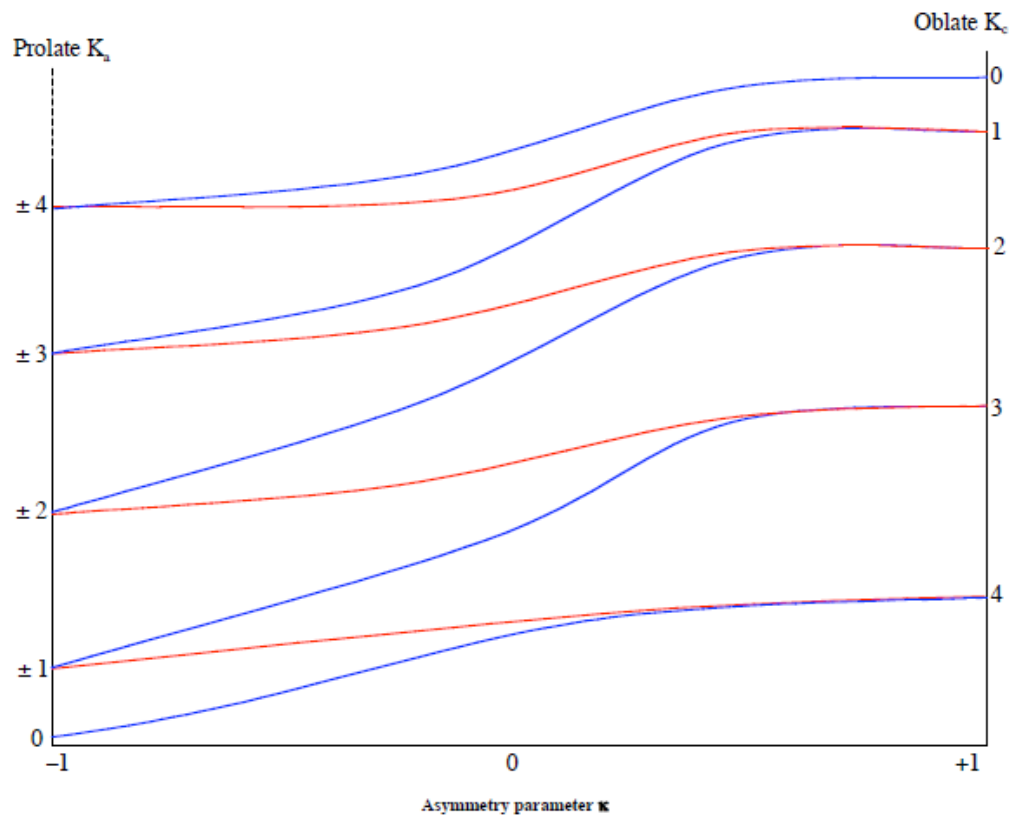
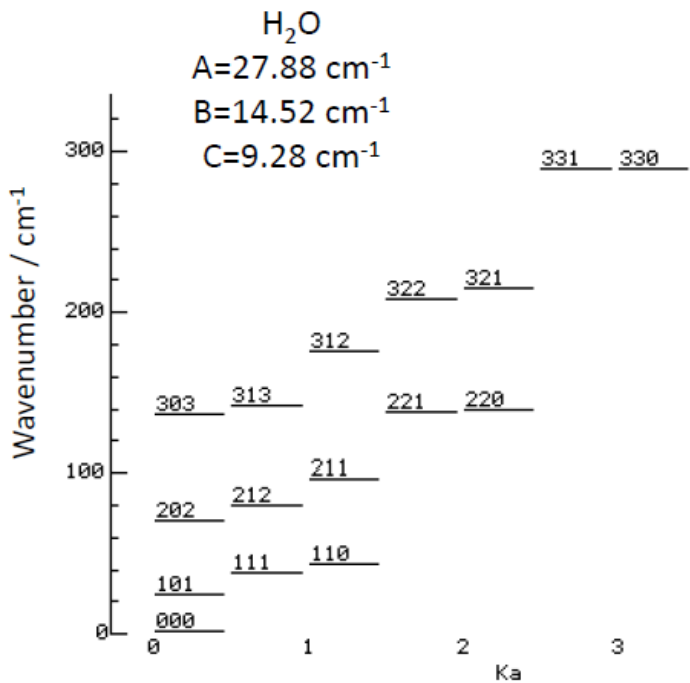
asymmetric rotor: H₂O



$$\mathfrak{I}_a < \mathfrak{I}_b < \mathfrak{I}_c$$



states classified by (JK_aK_c)



$$\text{asymmetry parameter: } \kappa = \frac{2B-A-C}{A-C}$$

Triaxial nuclei and wobbling

Bohr and Mottelson apply this for triaxial nuclei:

$$\mathfrak{I}_a > \mathfrak{I}_b > \mathfrak{I}_c$$

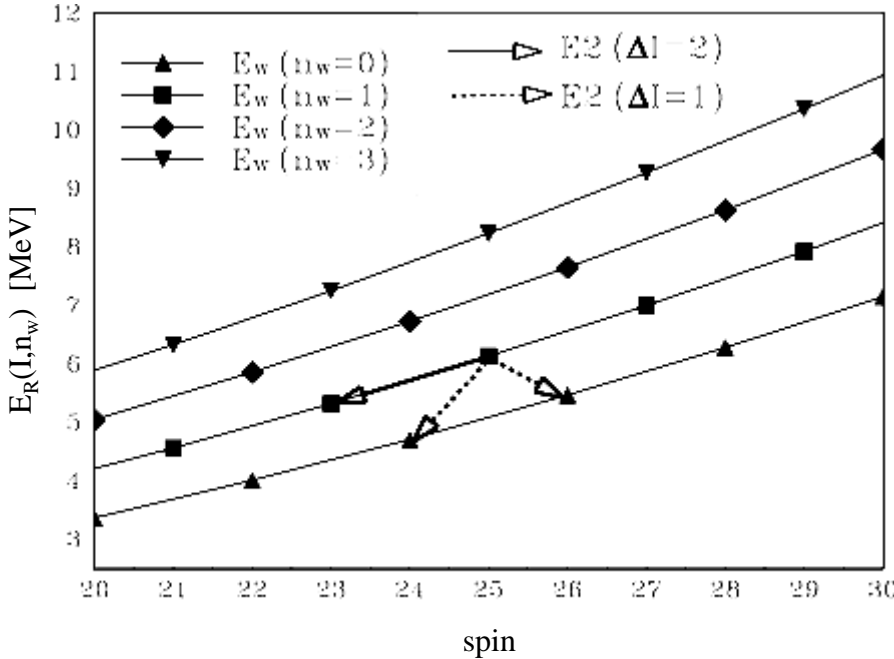
high spin: $I \approx I_x \gg 1$

$$E(I, n_w) = \frac{I(I + 1)}{2\mathfrak{I}_x} + \hbar\omega_w(n_w + \frac{1}{2})$$

n_w wobbling phonon number
 ω_w wobbling frequency

$$\hbar\omega_w = \frac{I}{\mathfrak{I}_a} \sqrt{\frac{(\mathfrak{I}_a - \mathfrak{I}_b)(\mathfrak{I}_a - \mathfrak{I}_c)}{\mathfrak{I}_b\mathfrak{I}_c}}$$

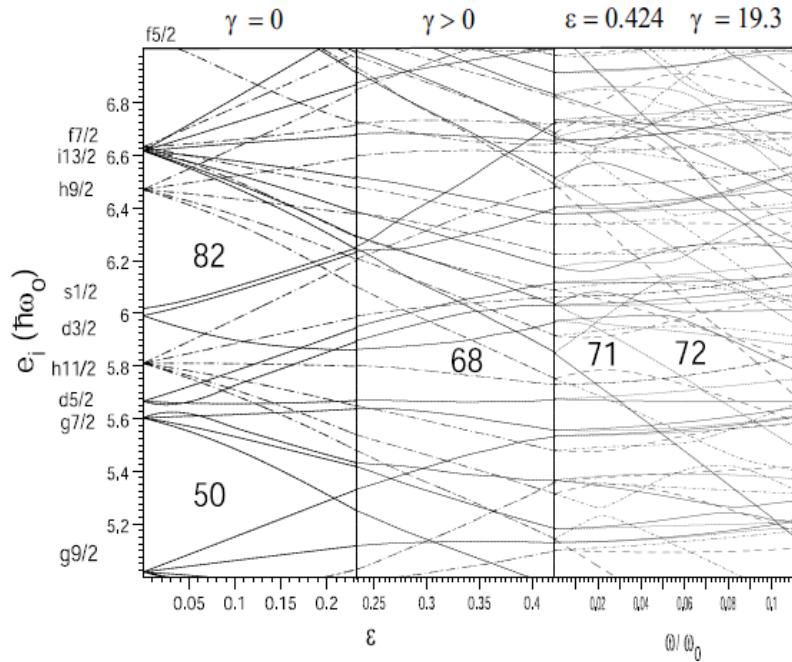
Bohr & Mottelson, Vol. 2, p.190 ff



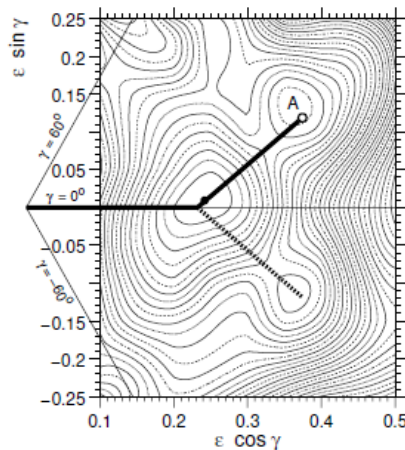
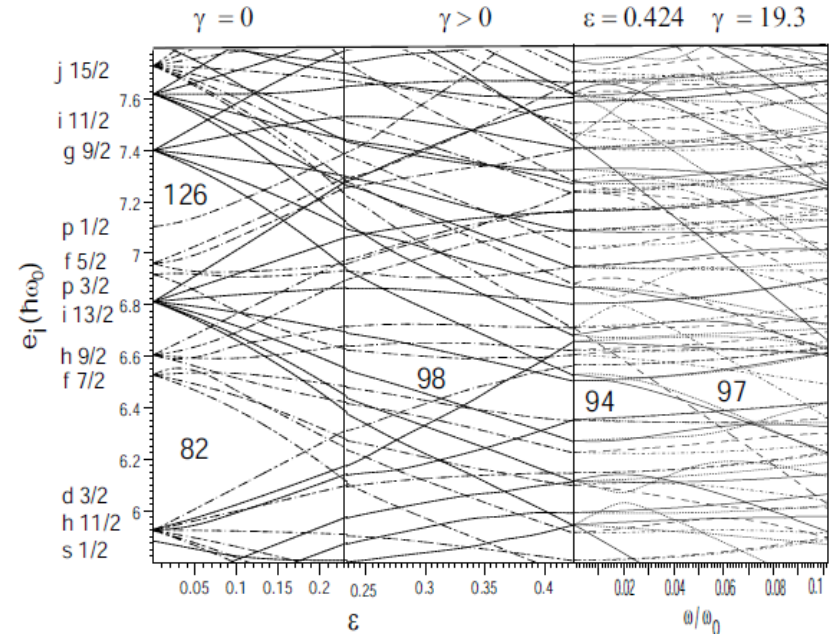
- Family of bands with similar rotational properties.
- Each band characterized by the wobbling phonon number n_w .
- Collective E2 inter-band decay compete with in-band transitions.

Where can we find nuclei with triaxial shape (at high spin) ?

Single Proton Levels



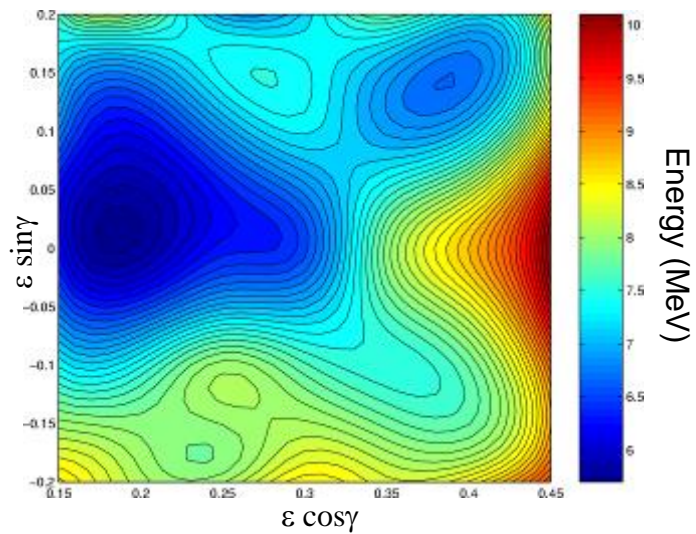
Single Neutron Levels



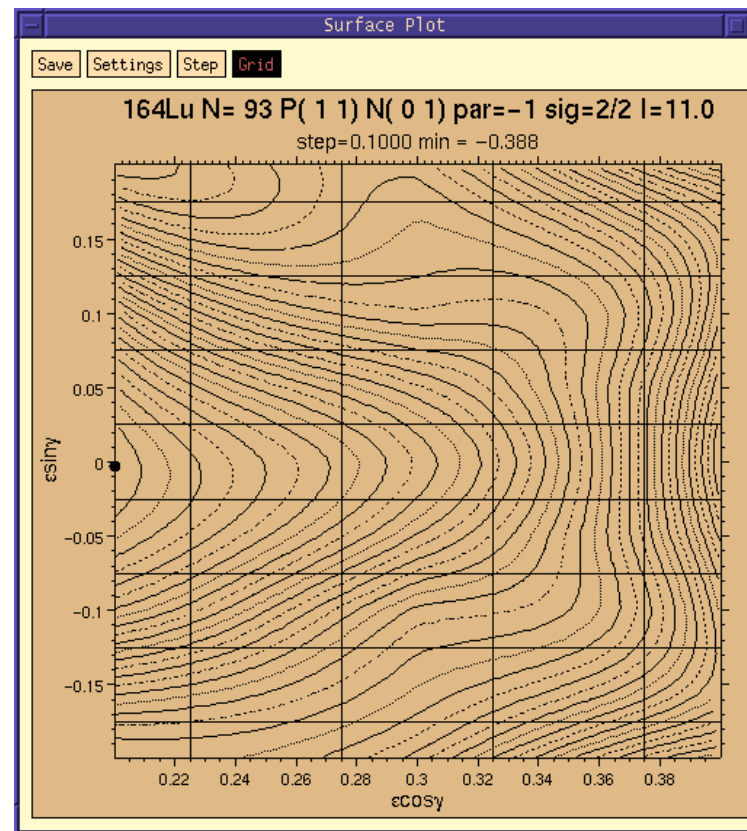
R. Bengtsson, H. Ryde
Eur. Phys. J. A 22, 355 (2004)

Strongly deformed triaxial shapes
stabilized by shell gaps in the region
 $Z \approx 71$ and $N \approx 94$ (^{165}Lu)

Triaxial superdeformation

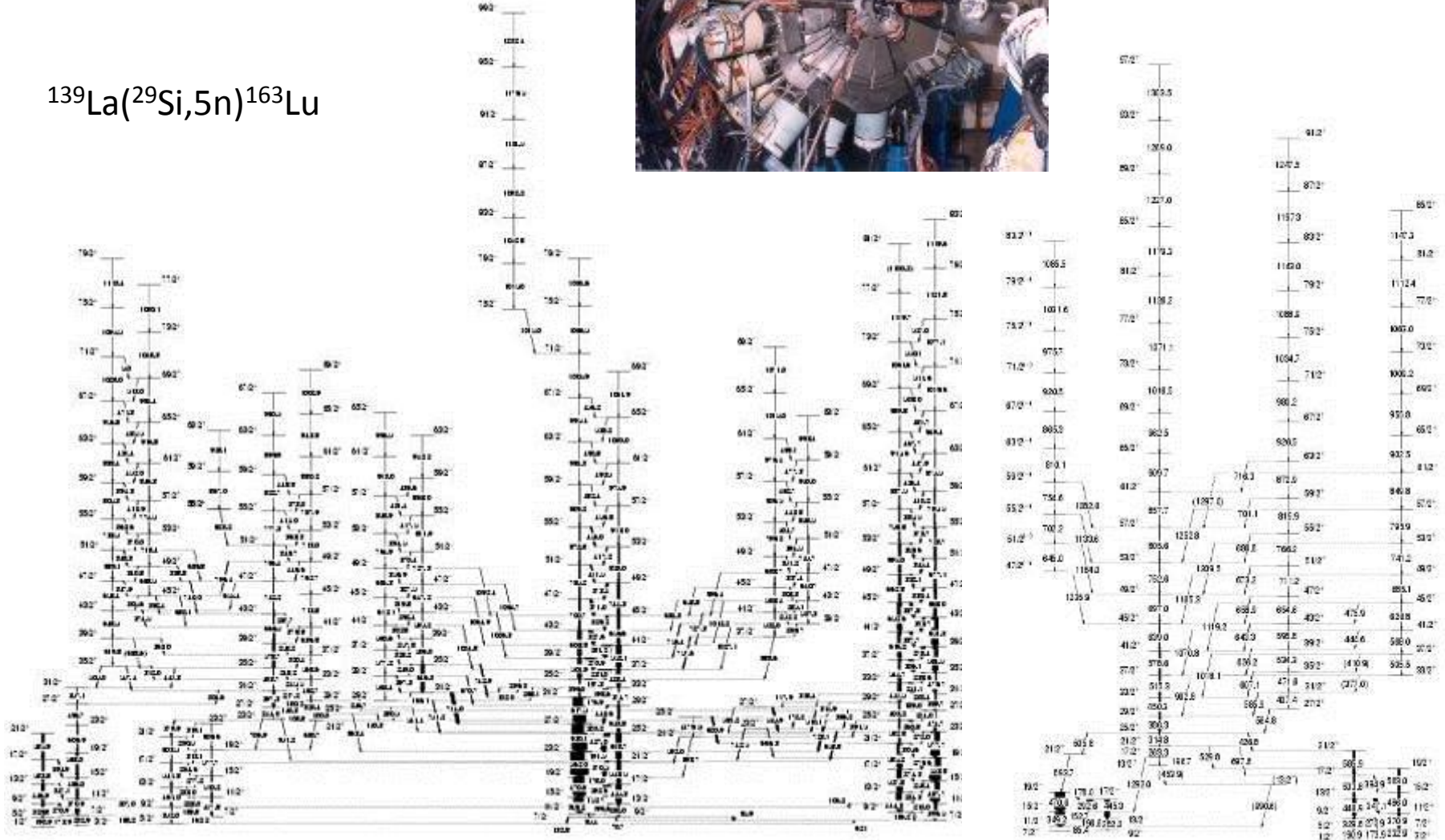
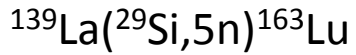


Potential energy
surface for ^{163}Lu
at $I=57/2$



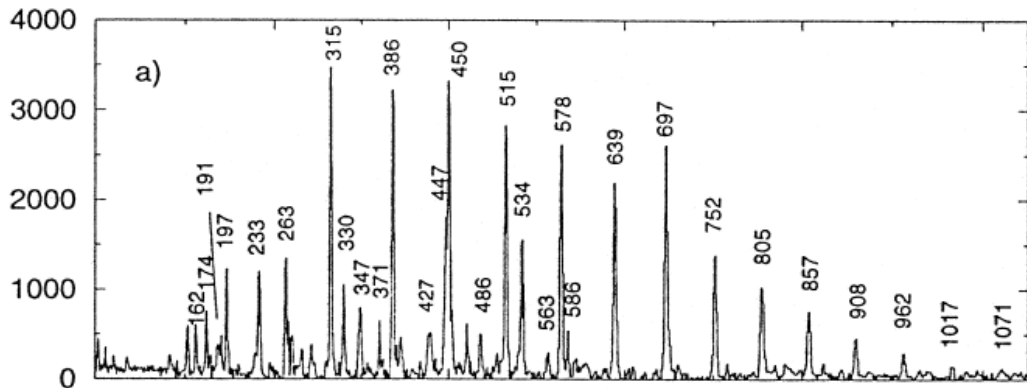
Level scheme of ^{163}Lu

277 levels
489 gammas

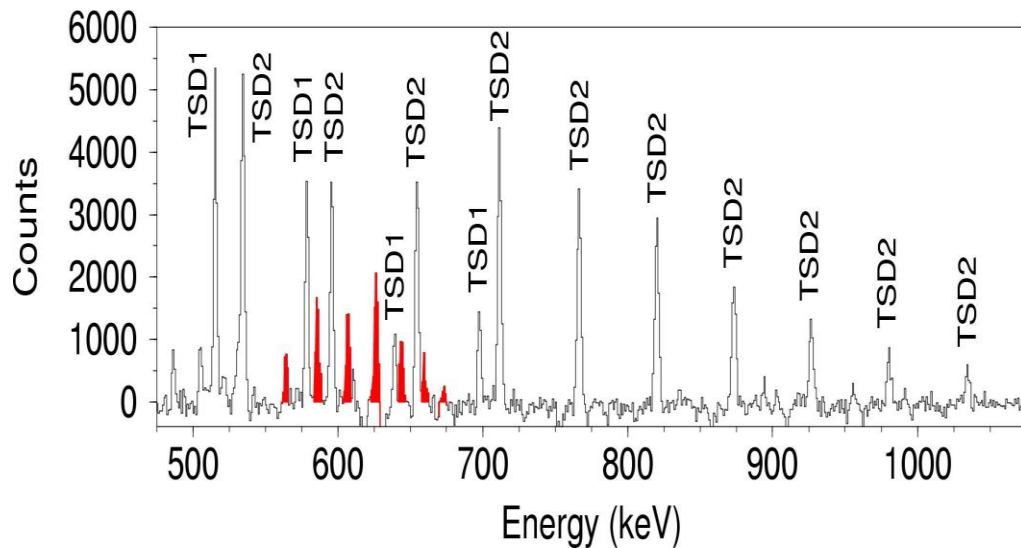


D.R. Jensen et al., Phys. Rev. Lett. 89, 142503 (2002)
Nucl. Phys. A 703, 3 (2002)

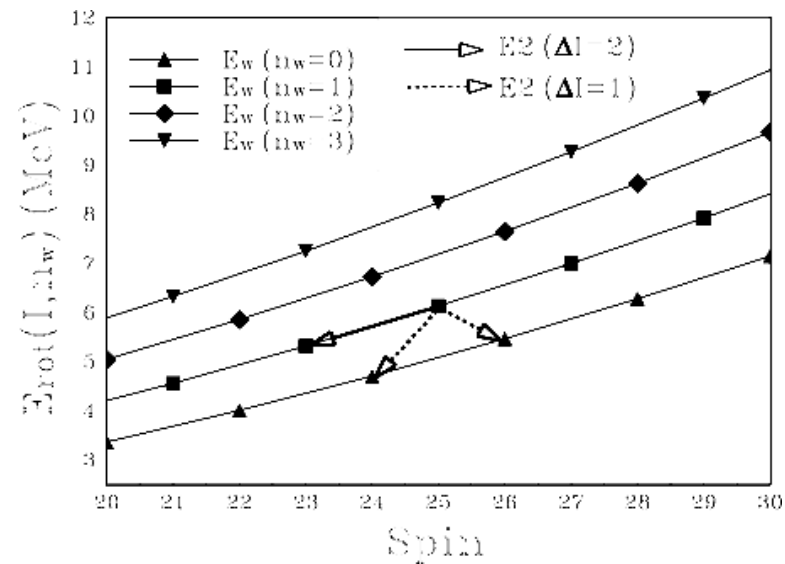
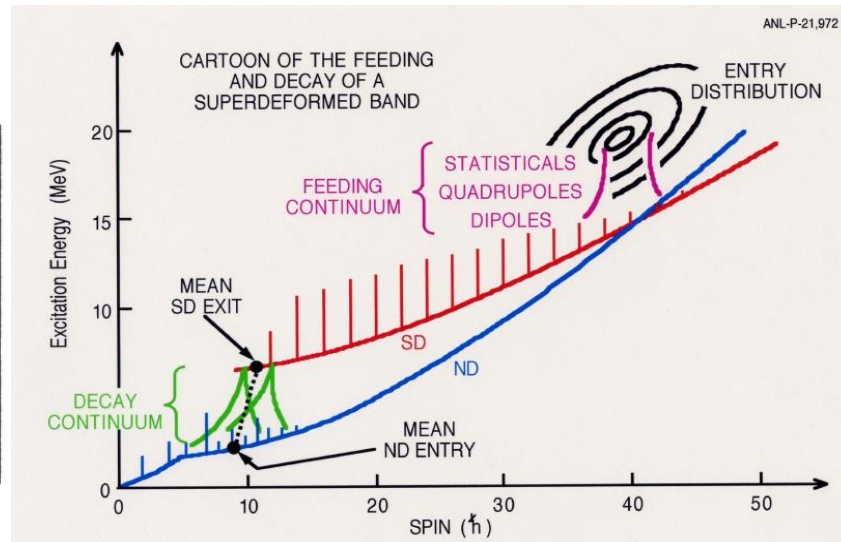
Triaxial superdeformed (TSD) bands in ^{163}Lu



J.Domscheit et al., Nucl. Phys. A 660, 381 (1999)



D.R.Jensen et al., Nucl. Phys. A 703, 3 (2002)



Level scheme of ^{163}Lu

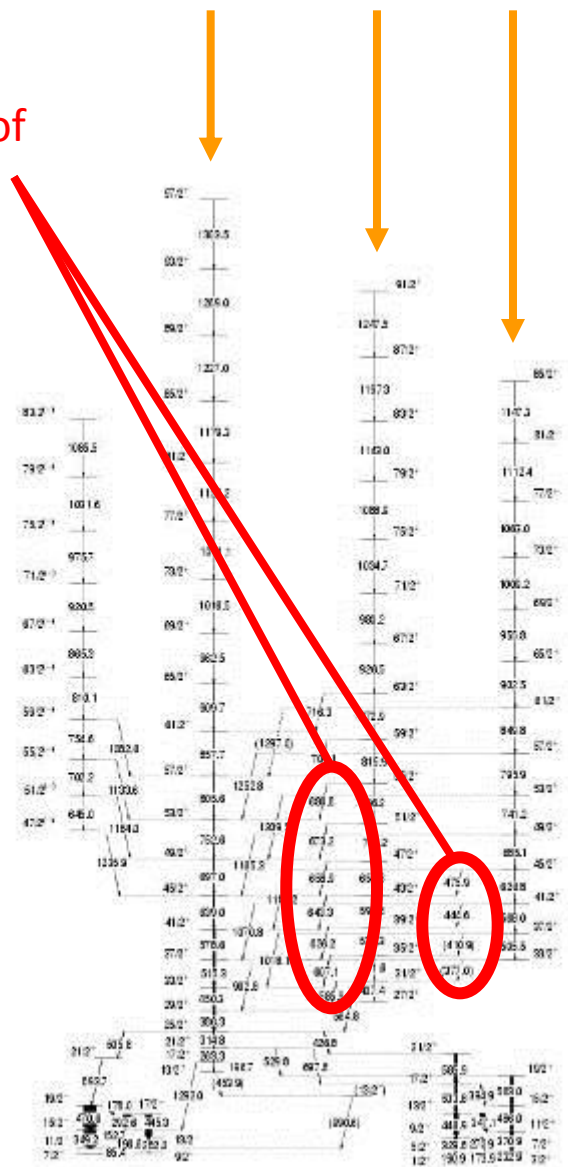
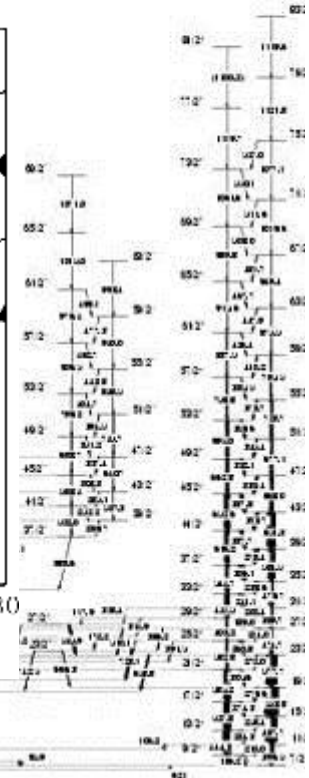
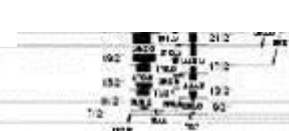
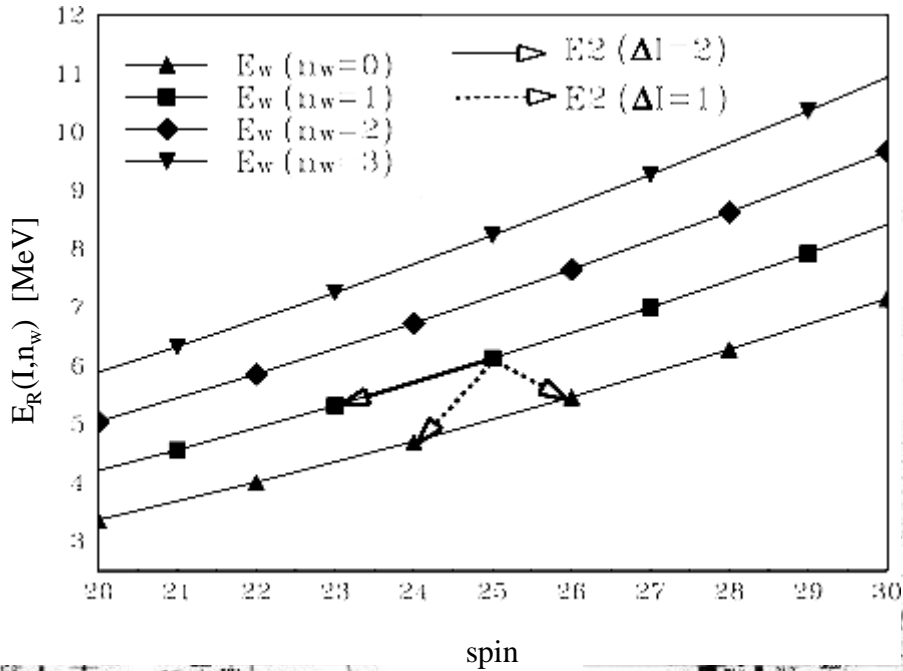
277 levels
489 gammas

Wobbling scenario requires inter-band transitions with:

- $\Delta I = 1$
- E2 character
- phonon rules

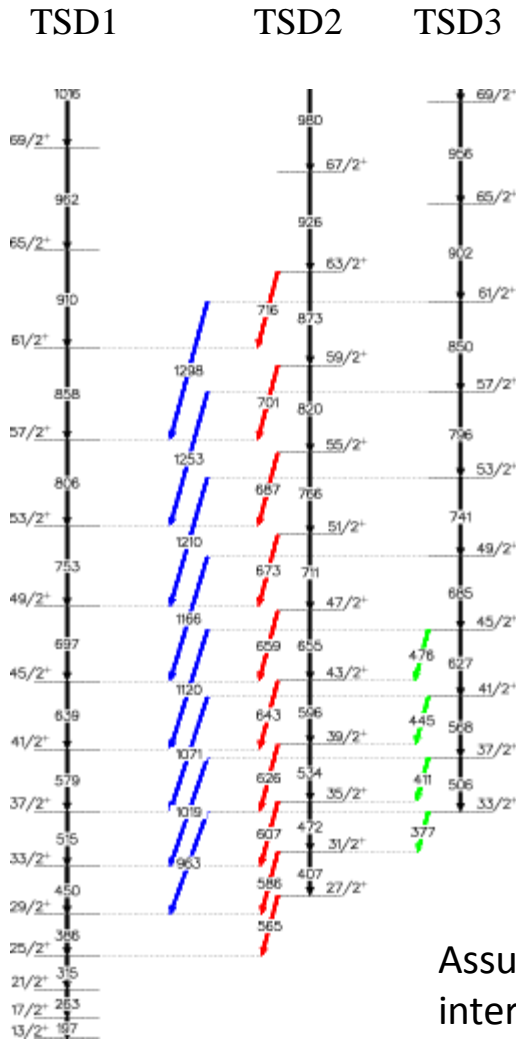
properties of inter-band transitions

triaxial superdeformed wobbling bands

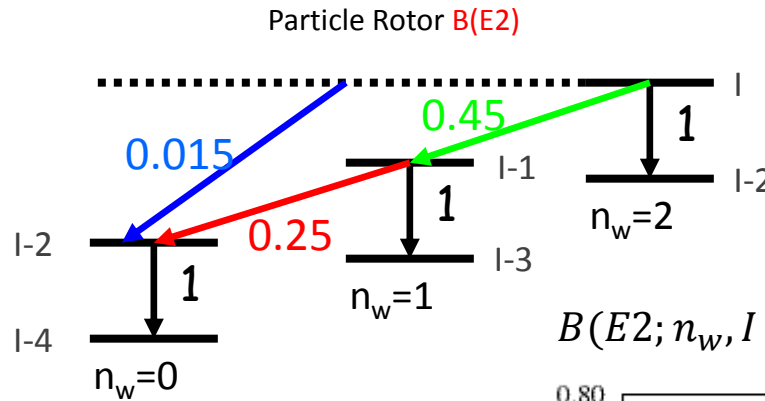
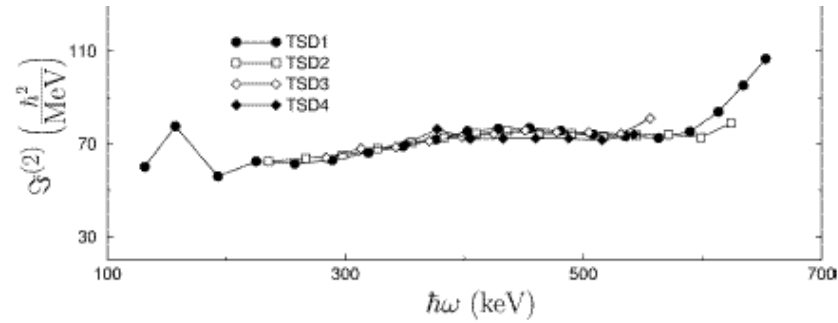


D.R. Jensen et al., Phys. Rev. Lett. 89, 142503 (2002)
Nucl. Phys. A 703, 3 (2002)

Evidence for the wobbling mode in ^{163}Lu

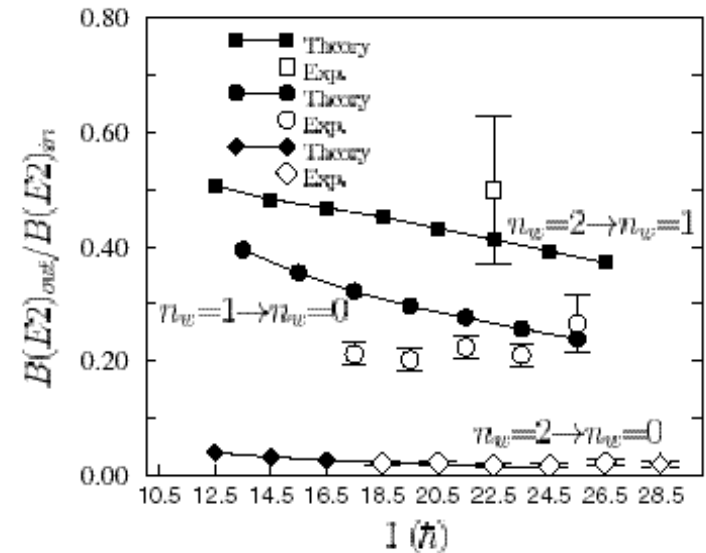


Family of bands with very similar rotational properties.



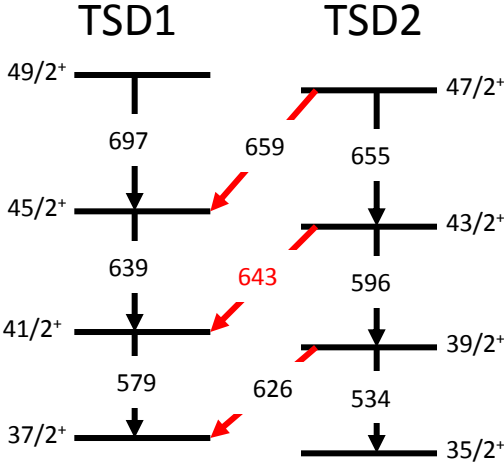
$$B(E2; n_w, I \rightarrow n_w - 1, I - 1) \propto \frac{n_w}{I}$$

Assuming E2 character for the inter-band transitions, they follow the phonon rule.



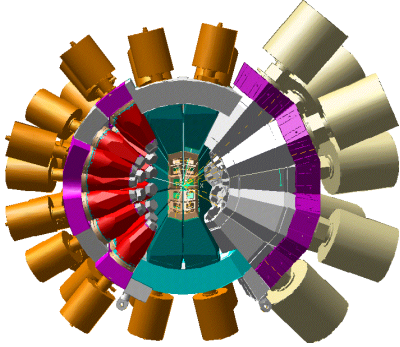
Are they E2?

Properties of inter-band transitions

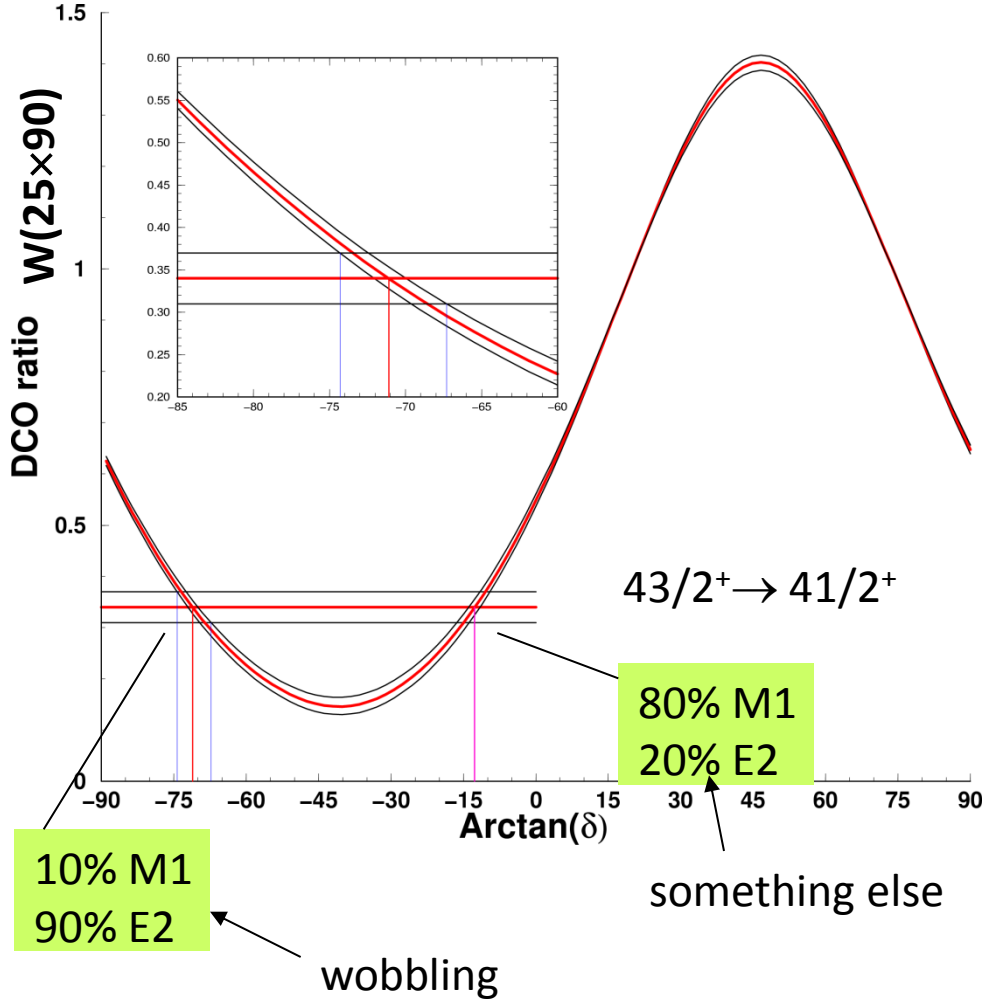


angular distribution (DCO ratio) for inter-band transitions

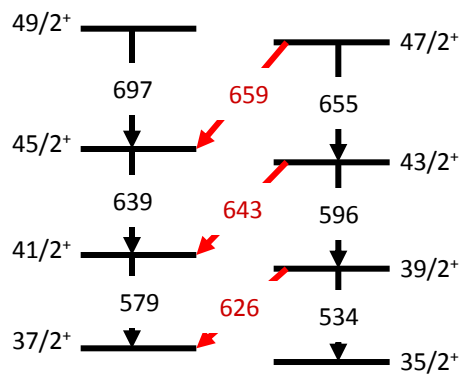
Two possible solutions



beam axis as quantization axis
probability of emission forward/backward relative to emission at 90°

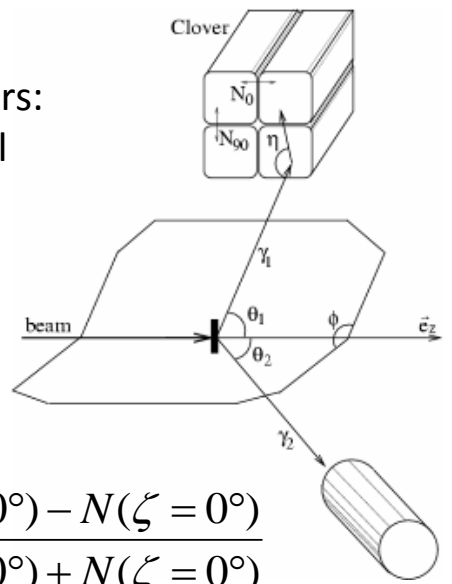


Polarization measurement in ^{163}Lu



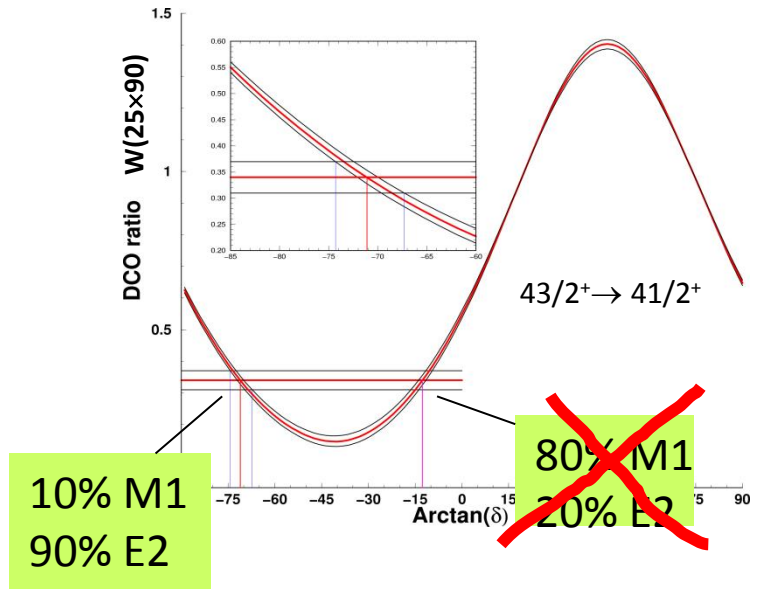
	E_γ	$A = \frac{N(90^\circ) - N(0^\circ)}{N(90^\circ) + N(0^\circ)}$	
E2	579	0.10 ± 0.03	positive
	697	0.13 ± 0.03	
	386	0.06 ± 0.05	
	534	0.05 ± 0.04	
M1	349	-0.11 ± 0.05	negative
inter-band	607	0.05 ± 0.05	positive ⇒ electric
	626	0.12 ± 0.05	
	643	0.11 ± 0.05	
	659	0.17 ± 0.09	
	673	0.18 ± 0.09	

Clover detectors as Compton polarimeters: horizontal vs. vertical scattering



$$P = \frac{A}{Q} = \frac{1}{Q} \frac{N(\zeta = 90^\circ) - N(\zeta = 0^\circ)}{N(\zeta = 90^\circ) + N(\zeta = 0^\circ)}$$

Evidence for the wobbling mode in nuclei
⇒ triaxial deformation



S.W. Ødegård et al., Phys. Rev. Lett. 86, 5866 (2001)

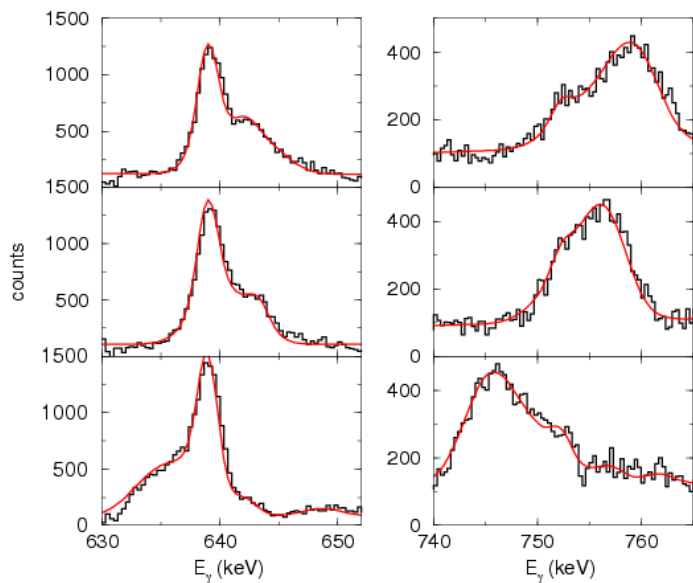
Can we say more about the shape?
 ⇒ DSAM lifetime measurement

target with backing

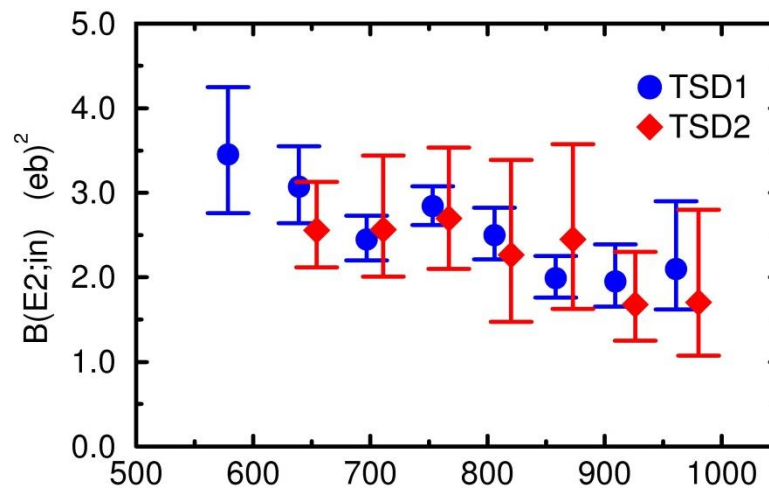
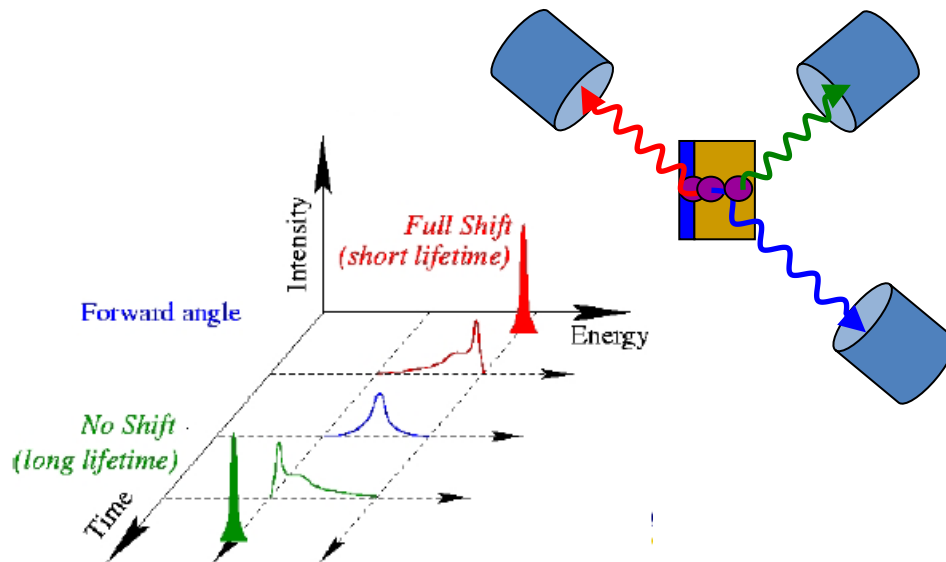
gamma rays are emitted

- with full recoil velocity
- slowed down
- finally stopped

characteristic lineshape

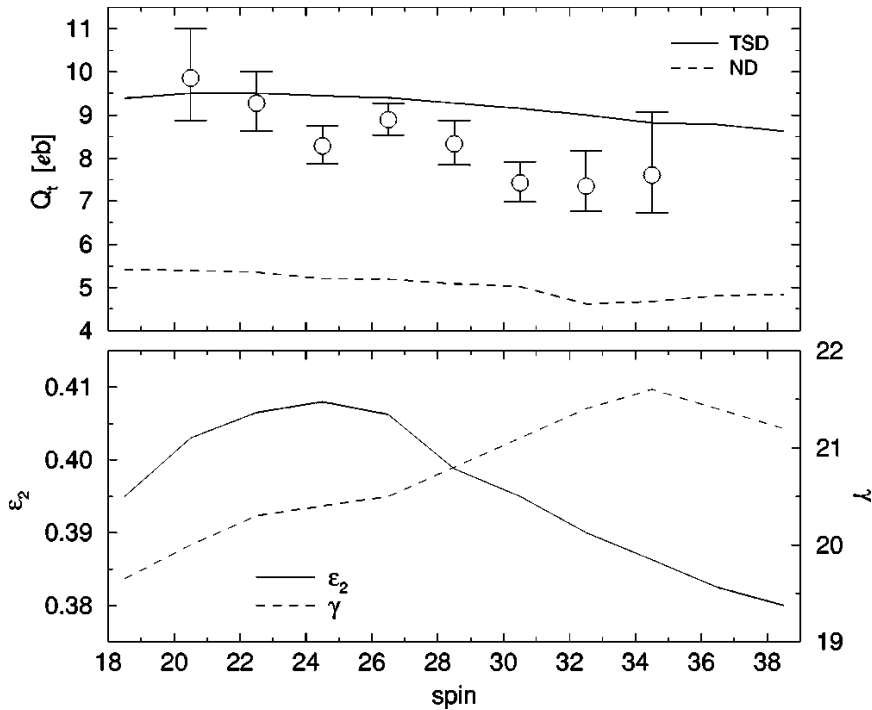


A. G3rgeren et al.,
 Phys. Rev. C 69, 031301(R) (2004)

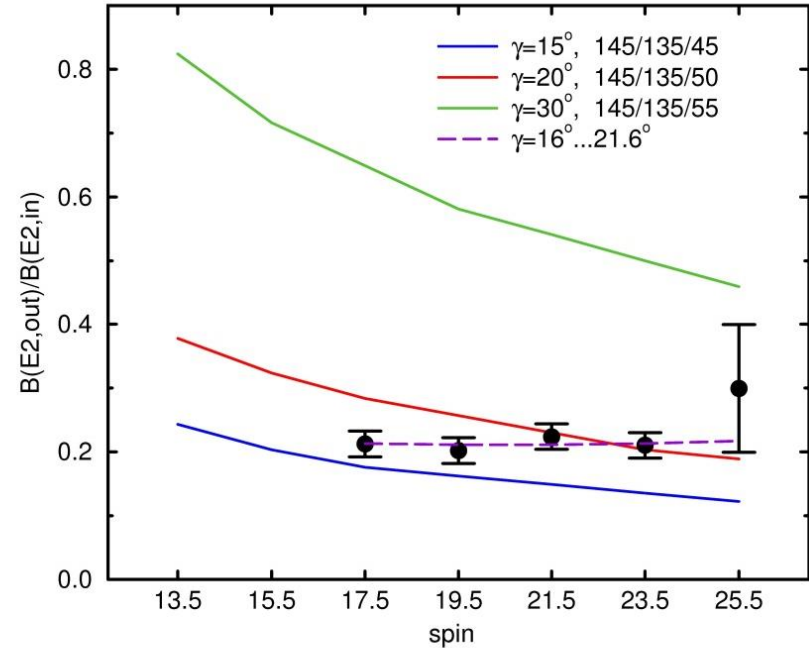


very similar in-band $B(E2)$ values
 for $n_w = 0$ and $n_w = 1$ bands

Deformation of the wobbling bands: triaxial superdeformation



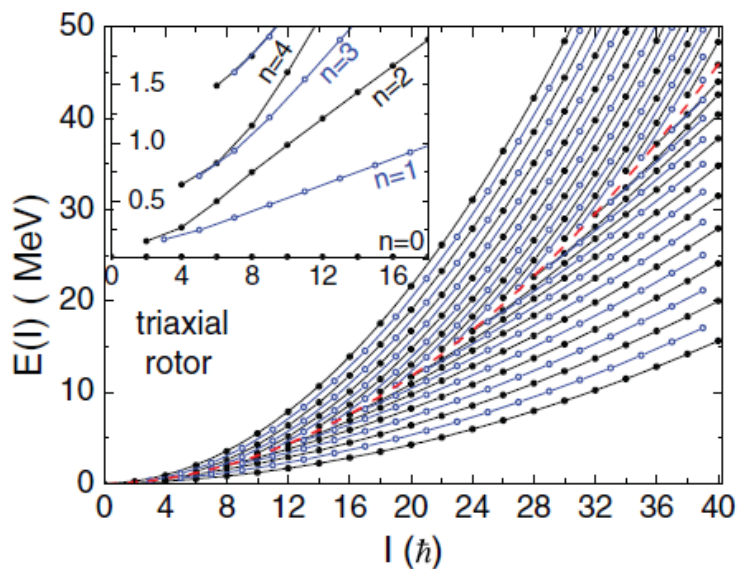
in-band $B(E2)$ values consistent with second minimum in PES



inter-band transition strengths suggest increase in γ deformation with spin

A. G3rger et al.,
Phys. Rev. C 69, 031301(R) (2004)

Wobbling frequency



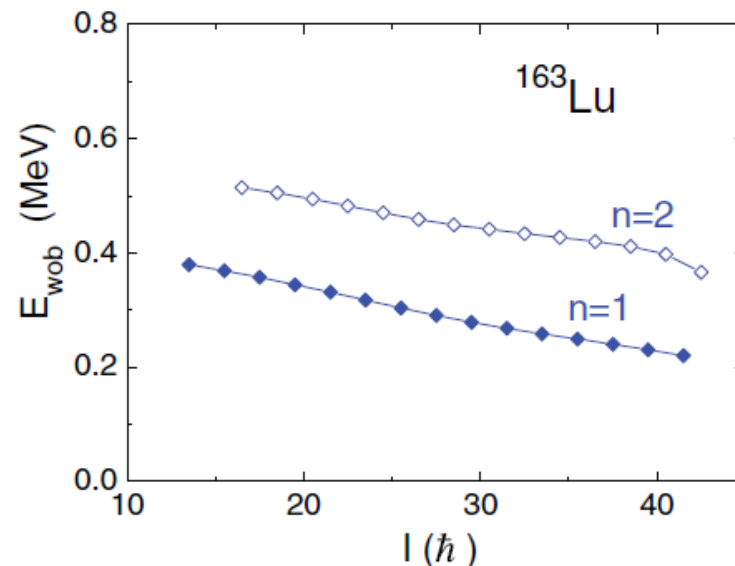
band structure for a triaxial rotor

S.Frauendorf and F.Dönau
PRC 89, 014322 (2014)

wobbling frequency should increase with spin

$$\hbar\omega_w = \frac{I}{\mathfrak{I}_a} \sqrt{\frac{(\mathfrak{I}_a - \mathfrak{I}_b)(\mathfrak{I}_a - \mathfrak{I}_c)}{\mathfrak{I}_b \mathfrak{I}_c}}$$

wobbling frequency decreases experimentally



what about the odd particle?

if the odd particle aligns with the largest Mol:

$$E = \frac{(I - j)^2}{2\mathfrak{I}} + \hbar\omega_w \left(n_w + \frac{1}{2} \right)$$

energies are shifted, but wobbling frequency should still increase with spin

⇒ longitudinal wobbling

Wobbling motion

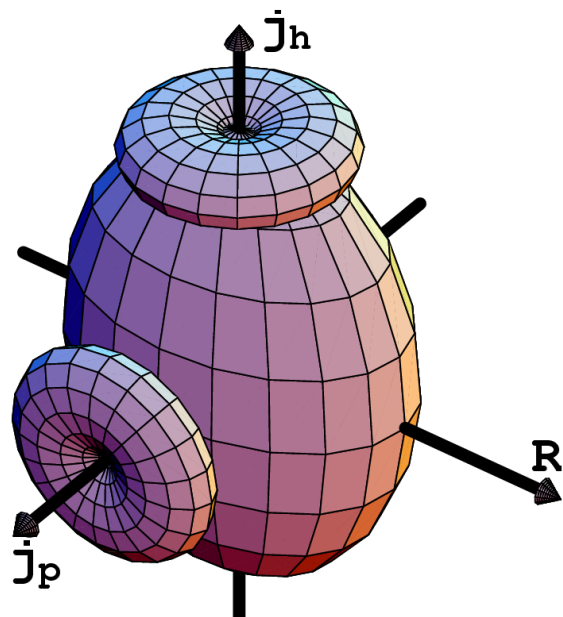
triaxial nucleus with short, medium, and long axis

typically it is: $\mathfrak{J}_m > \mathfrak{J}_s > \mathfrak{J}_l$

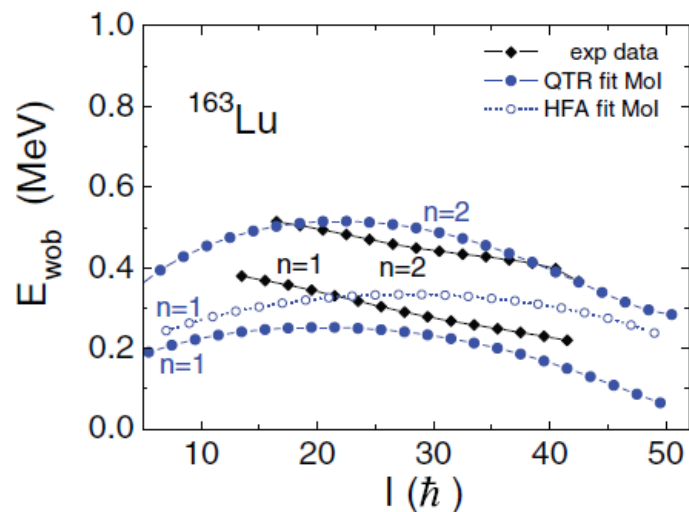
Nucleus	ε	$\gamma(\text{deg})$	Model	\mathcal{J}_m	\mathcal{J}_s	\mathcal{J}_l
^{163}Lu	0.4	20	fit	64	56	13
	0.4	20	hydrodyn	68	29	8
	0.4	20	cranking	59	51	13

- high- j particle aligns with short axis (maximum overlap for density distribution)
- high- j hole aligns with long axis
- collective rotation about medium axis

quasi-particle triaxial rotor calculation
(with *frozen alignment* approximation)



⇒ transverse wobbling



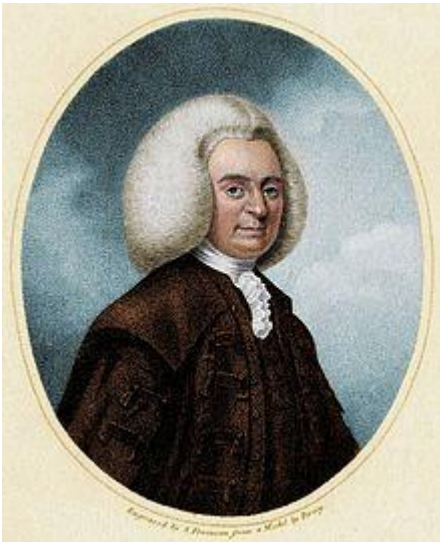
MacLaurin shapes

What happens if we spin
a liquid drop ?

- equatorial diameter 12756 km
- polar diameter 12714 km

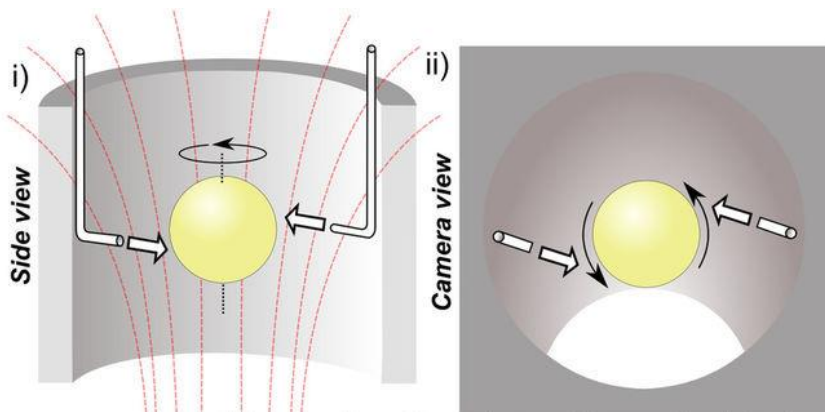
It becomes oblate !

MacLaurin shape
after Colin MacLaurin
(1698-1746)



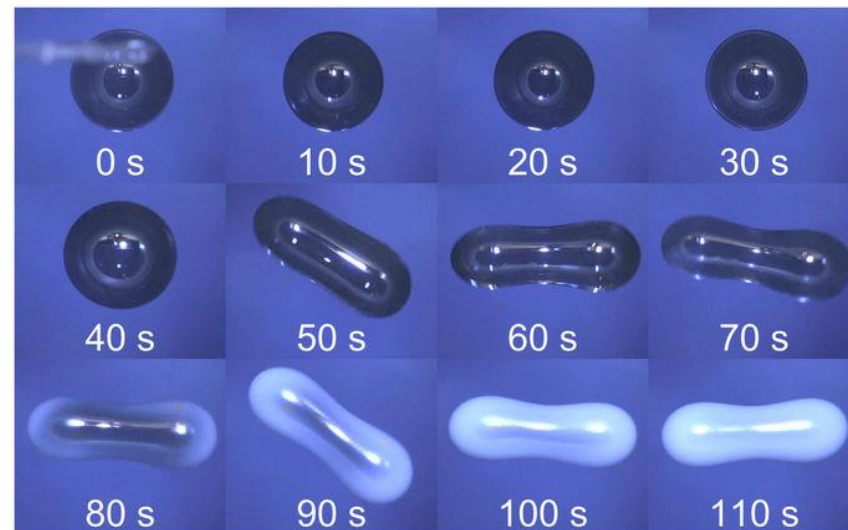
But what if we spin really fast ?

Toy model: spinning wax levitated in magnetic field

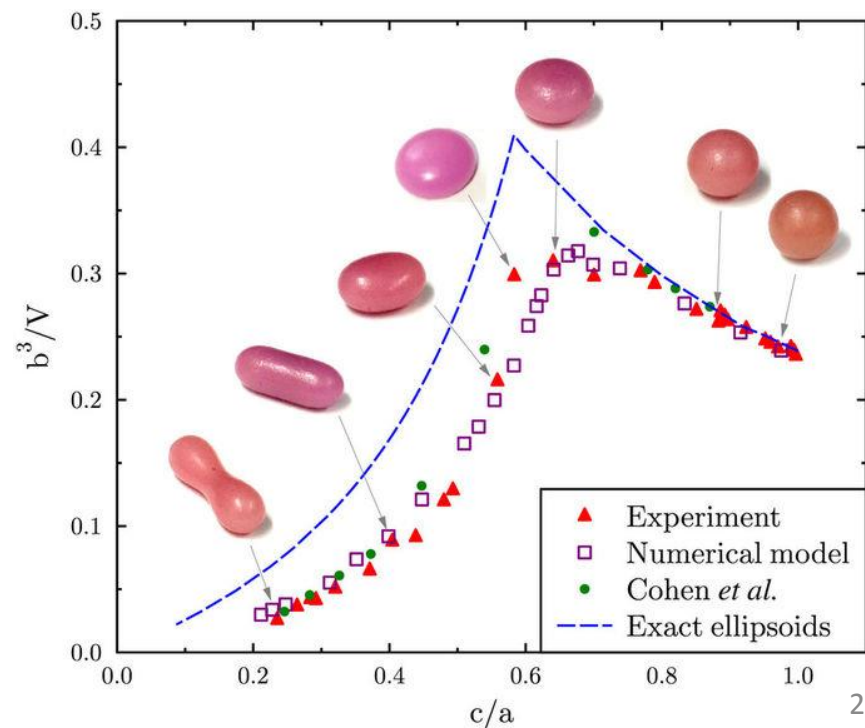


diamagnetic levitation in 18.5 T magnetic field rotated by air flow from two nozzles

K.A. Baldwin et al., Sci. Rep. 5, 7660 (2015)



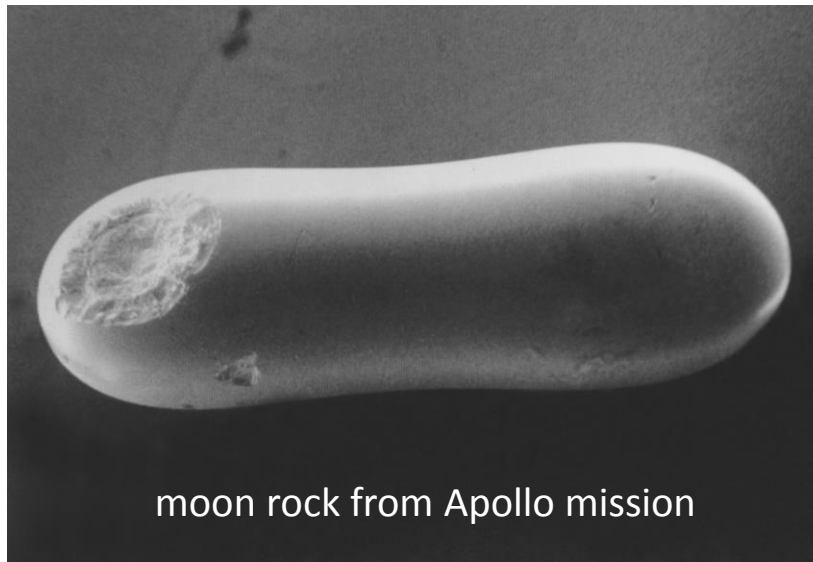
b) increasing angular momentum



Jacobi shapes



Carl Gustav Jacob Jacobi
(1804 - 1851)



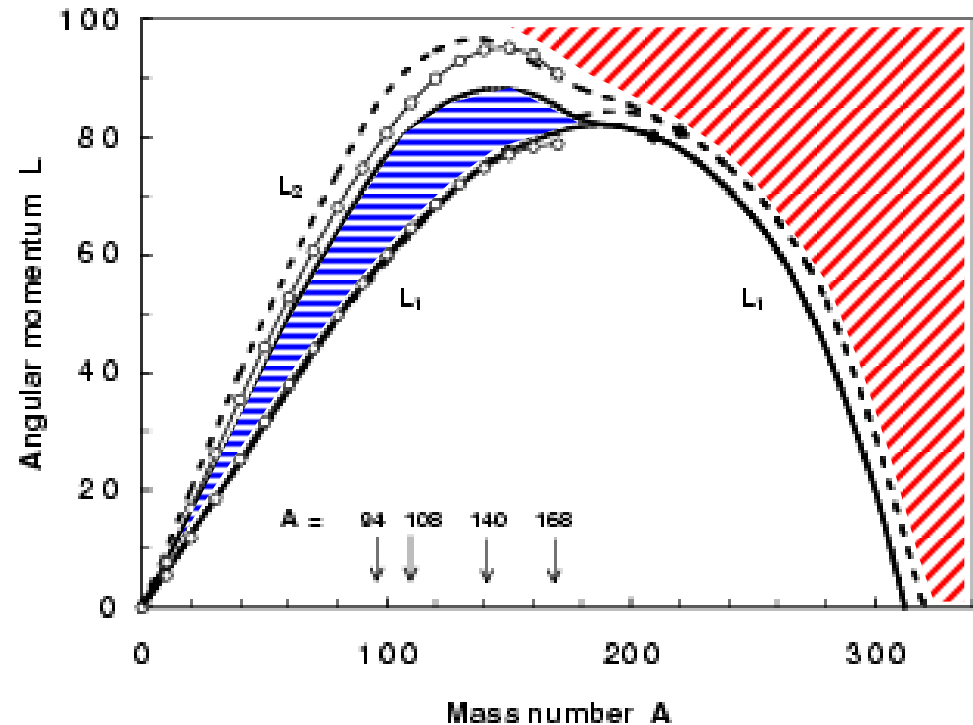
moon rock from Apollo mission

What happens
if you spin
even faster?

Fission!

Nuclear liquid drop model
(no shell effects)

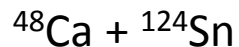
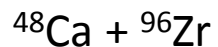
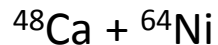
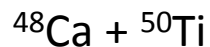
- heavy nuclei: fission for $L > L_1$
- Jacobi transition for $L_1 < L < L_2$
largest spin window around $A \approx 100$



W.D. Myers and W.J. Swiatecki
Acta Phys. Pol. B 32, 1033 (2001)

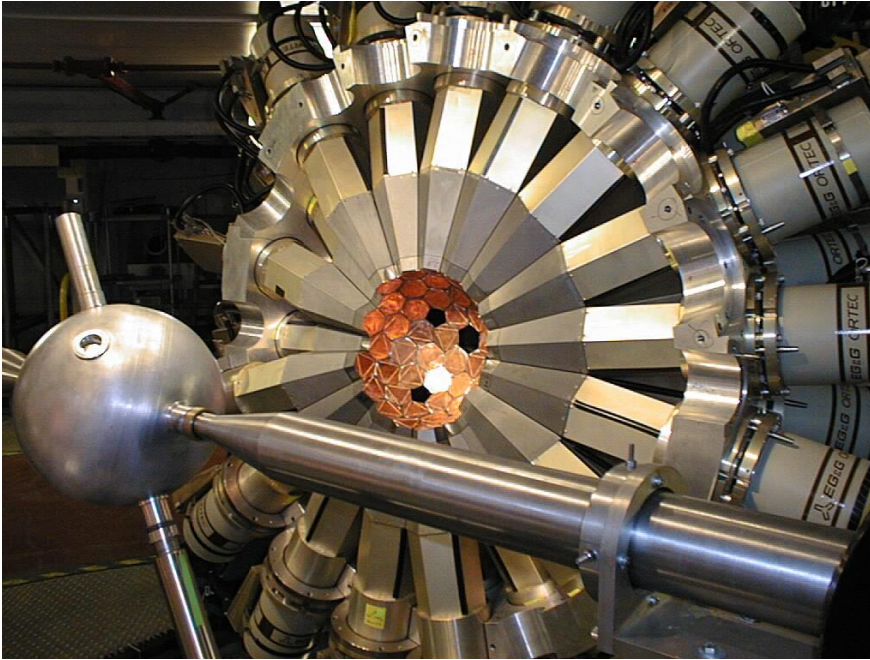
What is the signature of a Jacobi transition in nuclei ?

- sharp decrease of frequency with increasing angular momentum (giant backbend of the moment of inertia)
- frequency of collective rotation is related to the E2 γ -ray energy: $\hbar\omega = \frac{1}{2}E_\gamma$
- many rotational bands at high spin quasi-continuous transitions
- measure the energy of the quasi-continuous 'E2 bump' as a function of angular momentum
- series of experiments with Gammasphere

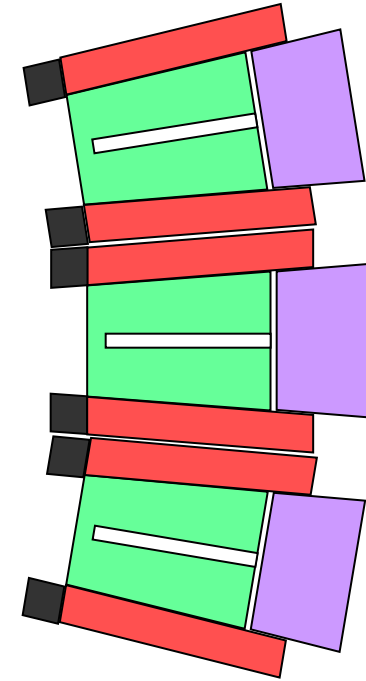


as neutron rich as possible:
 \Rightarrow higher fission barrier

Measuring angular momentum with Gammasphere

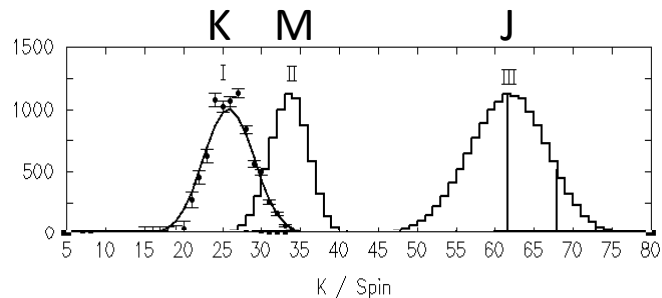


108 Compton-suppressed
HPGe detectors



108 Ge detectors
 $6 \times 108 = 648$ BGO detectors

increase in false veto signals
reduced Ge efficiency
but very high granularity

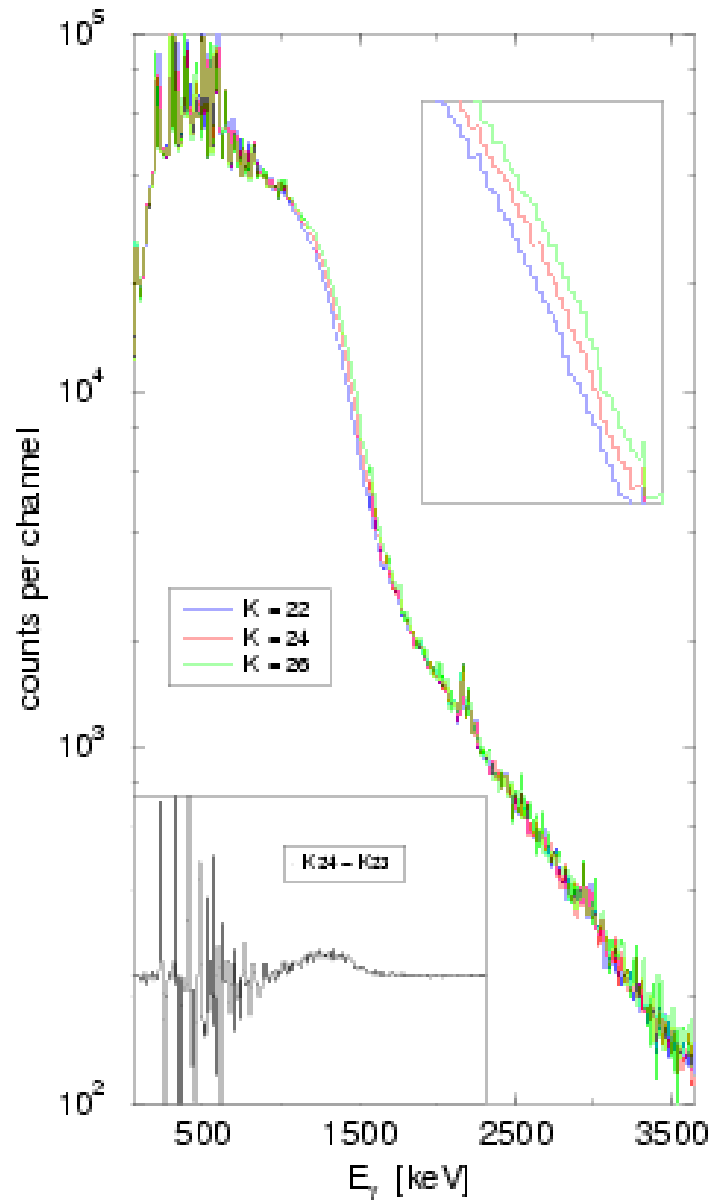


K = number of hits = fold

M = γ rays emitted = multiplicity (from response function)

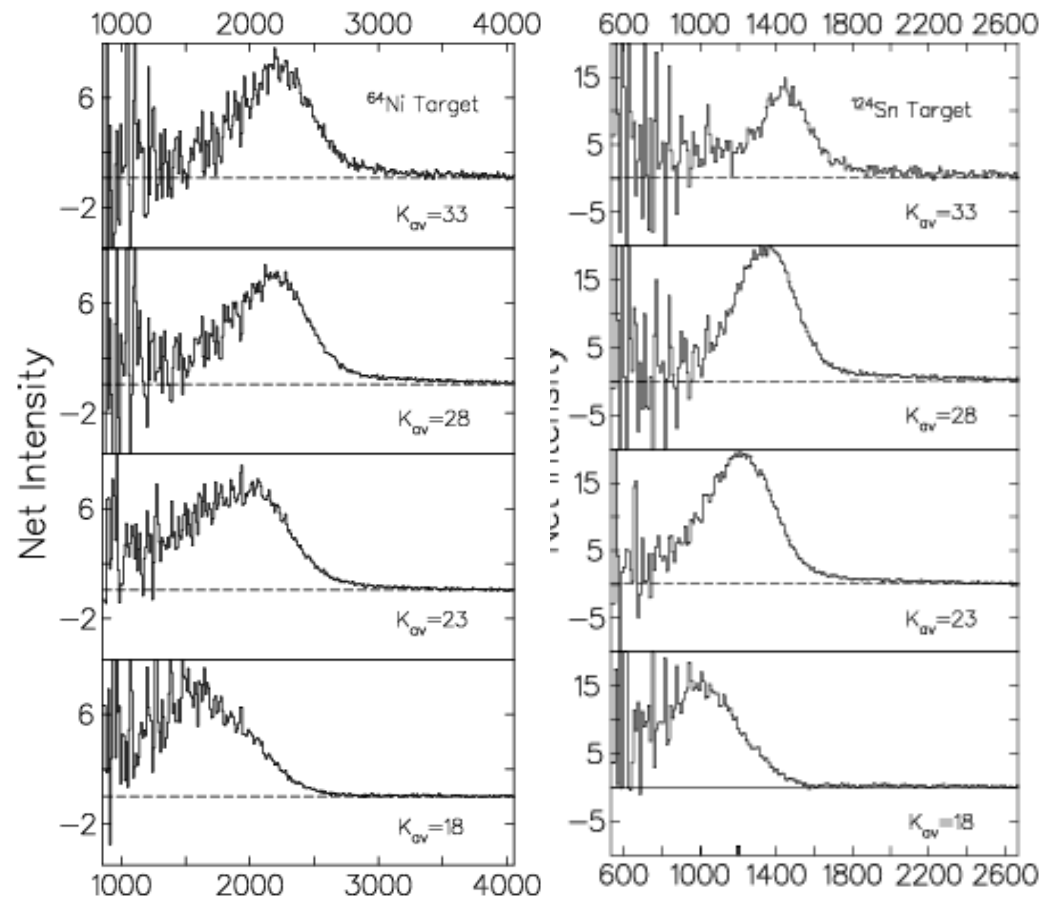
J = initial angular momentum (from angular distribution)

The E2 bump



Incremental spectra:

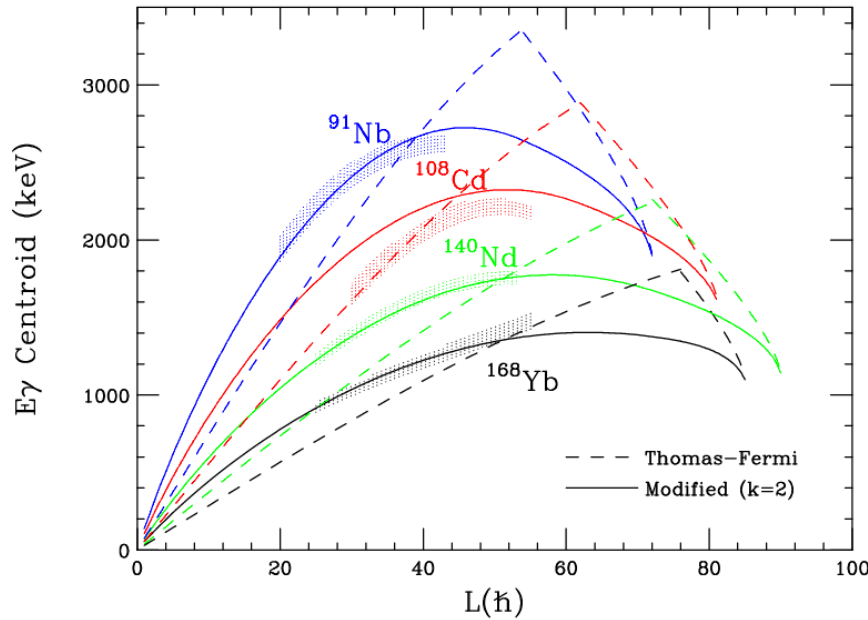
Multiplicity ($K_{av}-1$) gated spectrum
subtracted from ($K_{av}+1$) spectrum



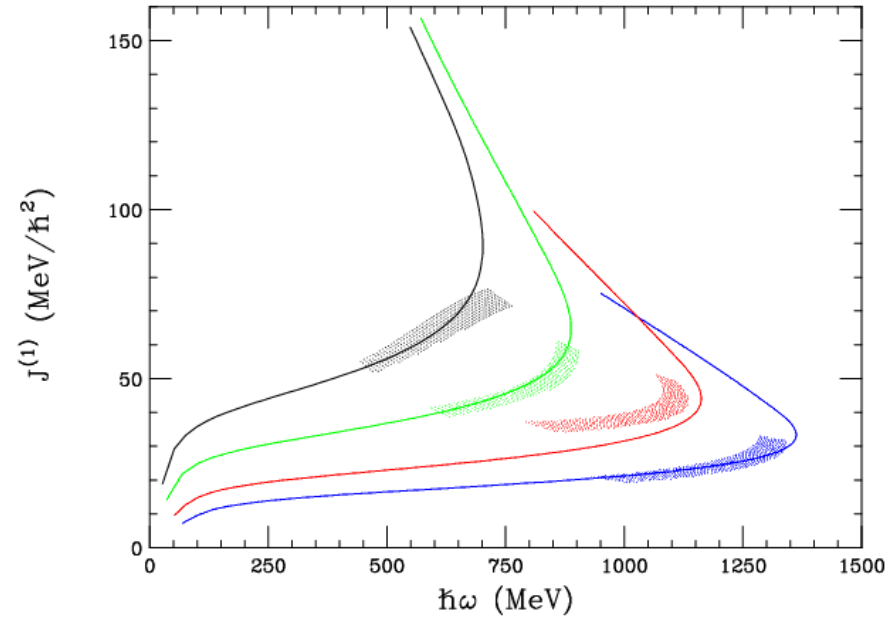
K measures the angular momentum

E2 bump measures rotational frequency

Comparison to liquid drop calculations



D. Ward et al., Phys. Rev. C 66, 024317 (2002)



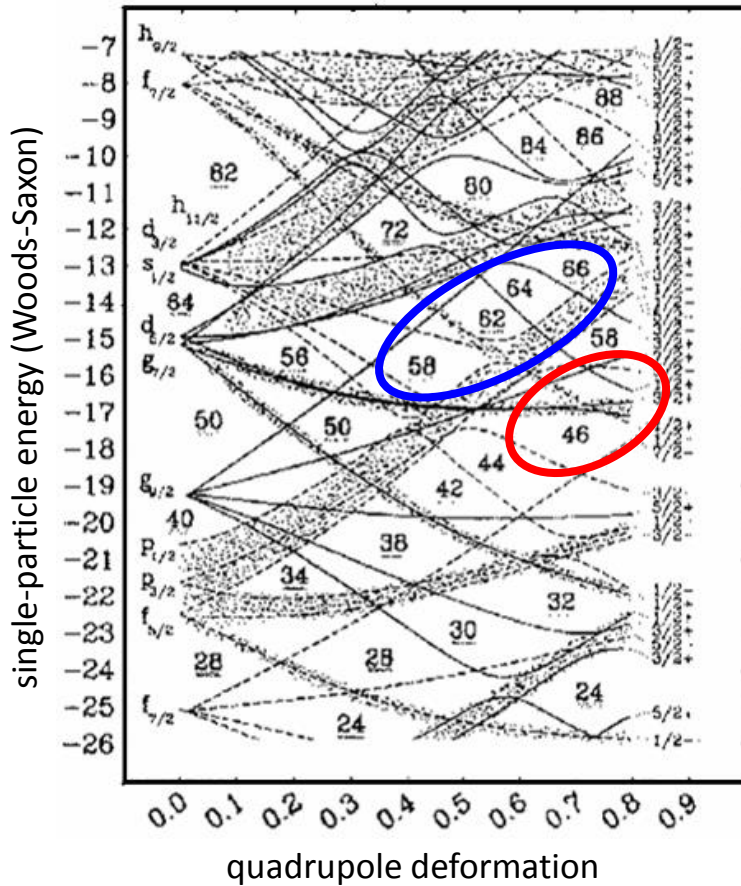
two modifications:

- lower effective moment of inertia at low spin due to pairing
- no collective rotation about axially symmetric (Maclaurin) shapes in nuclei, instead, collective rotations are associated with (mostly) prolate shapes
→ no sharp transition caused by breaking of axial symmetry, but smooth transition

shell effects?

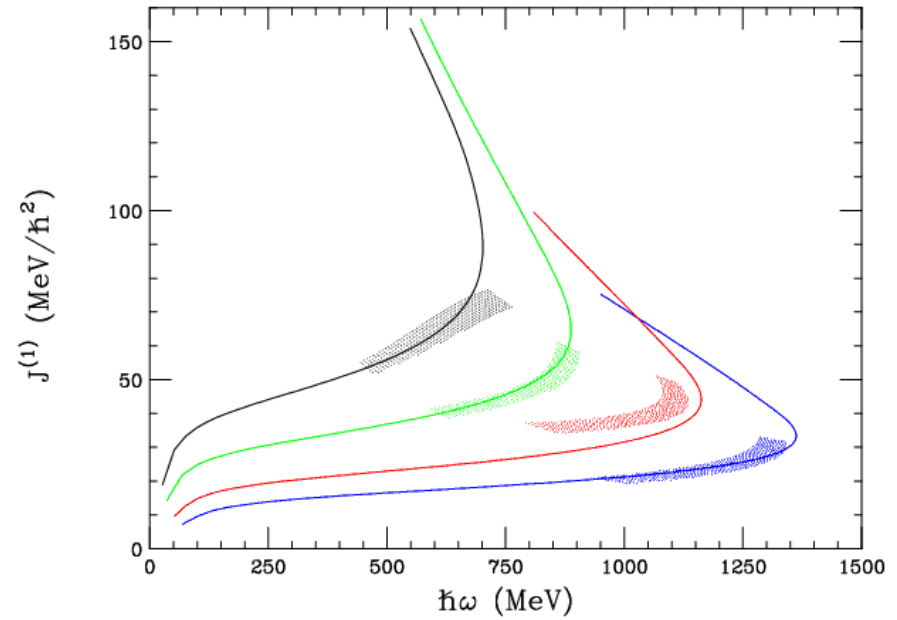
Extreme deformation in ^{108}Cd

- indication of Jacobi transition at very high spin ($I > 50 \hbar$)



- shell gaps at large deformation for $Z \approx 48$ and $N \approx 60$

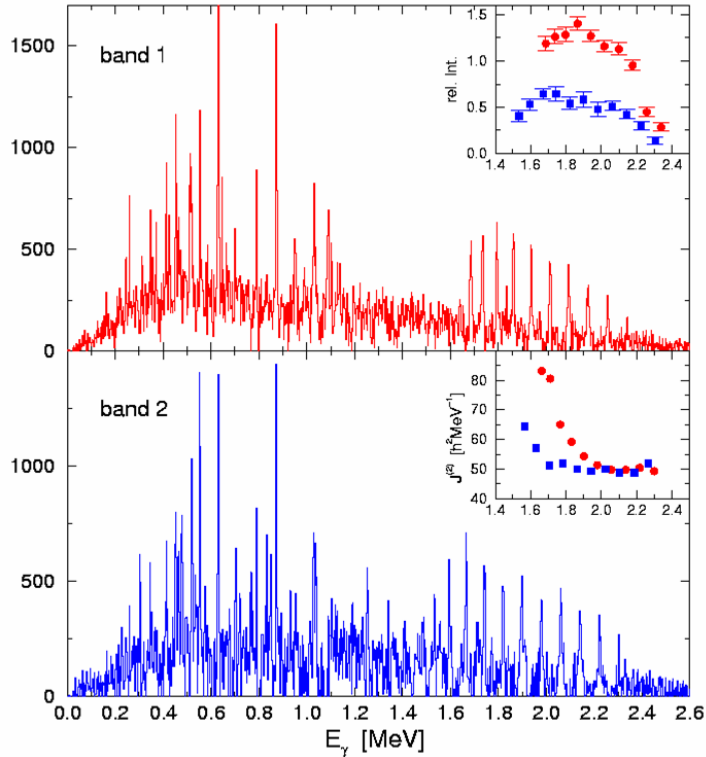
D. Ward et al., Phys. Rev. C 66, 024317 (2002)



Very deformed structures in ^{108}Cd

$^{64}\text{Ni}(^{48}\text{Ca},4n)^{108}\text{Cd}$

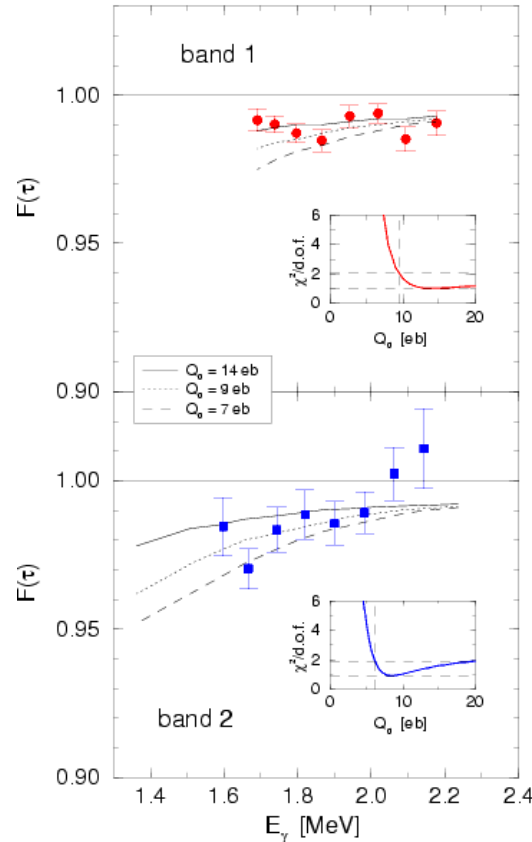
GammaSphere @ LBNL / ANL



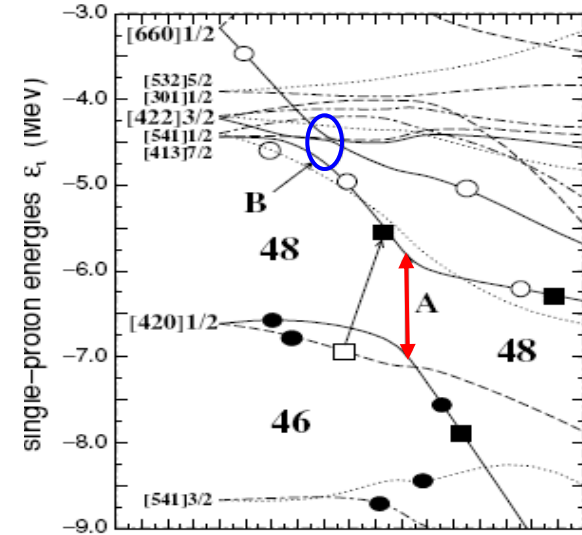
γ -ray multiplicity
 ➤ spin range 40-60 \hbar

R.M. Clark et al., PRL 87, 202502 (2001)

A. Gorgen et al., PRC 65, 027302 (2002)



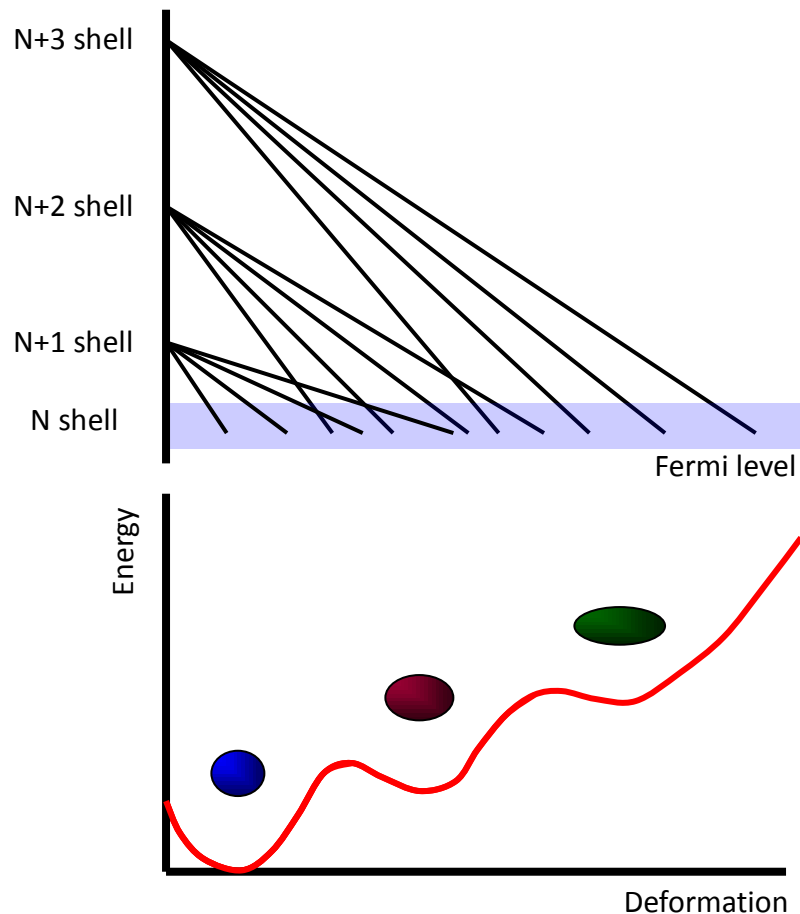
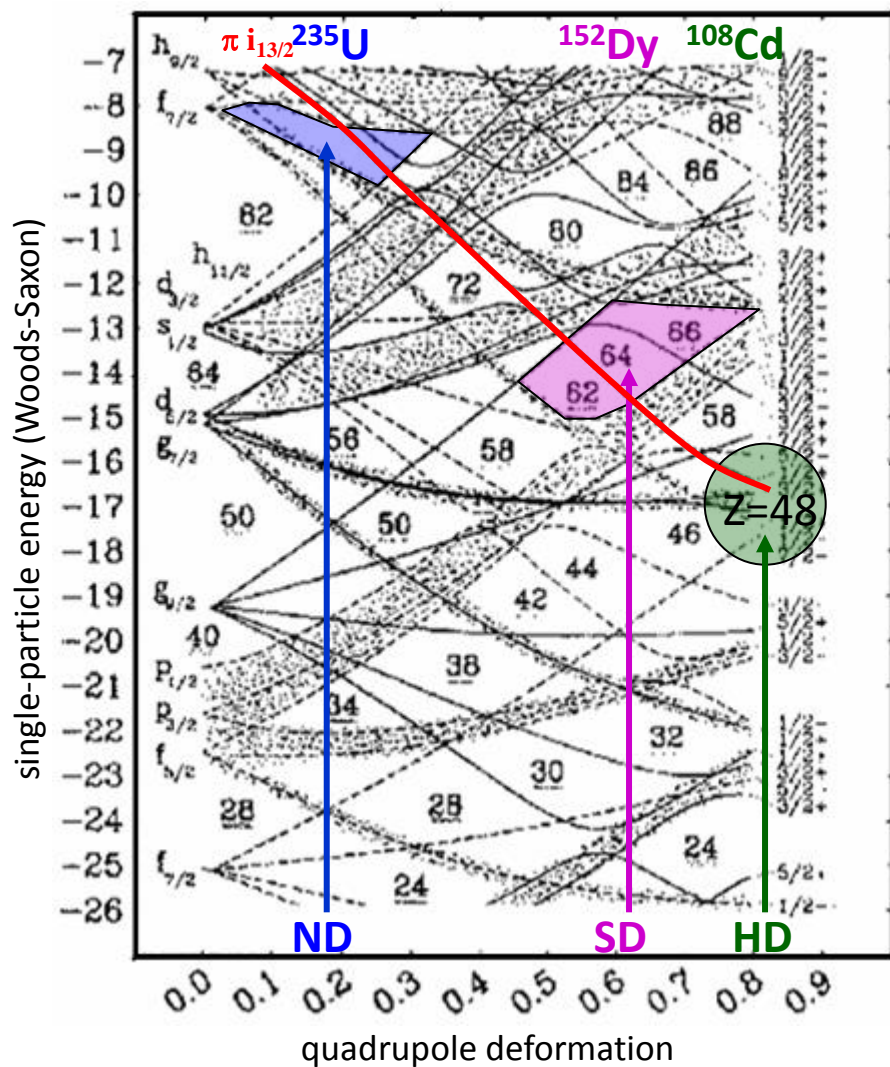
residual Doppler shifts
 ➤ lower limit on Q_t , $\beta_2 \geq 0.6$



A.V. Afanasjev, S. Frauendorf,
 PRC 72, 031301 (2005)

cranked RMF calculations
 ➤ evidence for occupation
 of proton $i_{13/2}$ orbital

Intruder orbitals



- (N+1) intruder
⇒ normal deformed, e.g. ^{235}U
- (N+2) super-intruder
⇒ Superdeformation, e.g. ^{152}Dy
- (N+3) hyper-intruder occupied in ^{108}Cd
⇒ Hyperdeformation ?