





# First preliminary results of CTR measurement in AWAKE runs June/September 2017

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## The Zoo of Plasma Wake-field Accelerators



... started from pioneer paper "Laser Electron Accelerator" by T.Tajima and J.Dawson Phys. Rev. Lett. 43, 267 – Published 23 July 1979

Laser Beat-Wave WFA (~1 ns) Two frequencies laser pulse (pulse train)

**Self-Modulated Laser WFA (~1 ns)** Raman forward scattering instability in a long laser pulse

Laser WFA (~0.1 ps) Short intense laser pulse

**Particle Bunch WFA** Short intense particle bunch ~ 1ps proton bunch does not exist !



~1ns

### Self-Modulated Particle Bunch WFA

Long bunch experience transverse self-modulation instability

## **Scope of AWAKE proof-of-principle experiment**

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# **AWAKE Physics: Principle**

Ionization front is co-propagating with a short laser pulse and seeds Self Modulation Instability (SMI)  $\tau_{laser} \sim 100 \text{ fs} \ll \tau_{wake} \sim 3 \text{ ps}$ 



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## **AWAKE Physics: Principle**

Ionization front is co-propagating with a short laser pulse and seeds Self Modulation Instability (SMI)



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## The aims of CTR diagnostics are:

- To measure a relative or absolute CTR signal strength
- To measure a carrier frequency of CTR signal or its harmonics
- To show that it is close to an expected plasma frequency
- With our AWAKE parameters we expect f<sub>CTR</sub> = 90 290 GHz



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In AWAKE case large-screen and far-field conditions are easily fulfilled for microwave CTR due to a small emitter size – foot print of a bunch Coulomb EM-field at modulation frequency is finite ( < 20 mm diameter)





## **CTR : Simulation**

dW/dOmega, kW/strad



- **•** Donut-shape spatial pattern of a CTR beam, not 1/γ cone
- At peak, pulse fluence is 2-10 µJ/sr => 5-25 nJ/cm<sup>2</sup> at 20 cm
- At peak, pulse power is 5-30 kW/sr => 13-75 W/cm<sup>2</sup> at 20 cm









- Amount of harmonics of plasma frequency is essential and measurable
- Note a different scale on vertical axis!

Spectral energy angular distribution,  $\mu$ J/sr, integrated over 20 GHz band



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## **CTR Diagnostics: Setup**



## Key components:

- 150mm travel UHV translator
- Screen is a Ø100mm Al-coated Si wafer
- Ø100mm quartz view-port
- 4" off-axis parabola





## **CTR Diagnostics: Layout**









## Three waveguide-integrated Schottky diodes with horn antennas Three different bands, for quick CTR signal check-up







**Typical CTR signal** from wide-band Schottky diodes



- Single shot diagnostics for quick CTR signal check
- **CTR signal detected at harmonics (power not calibrated)**
- Preliminary !!! Modulation of p<sup>+</sup> is "nonlinear", hinted by presence of CTR harmonia ۲

Harmonics (?) are visible in the wide-band diodes





• Signal has a large scatter – source is not understood, laser/proton overlap jitter?

In this particular example:

- Plasma frequency about ~172 GHz
- Laser-proton delay 125ps (TiSa is in the 1<sup>st</sup> quarter of a proton bunch)

#### Raw WR's signal

#### ... and binned with a bunch charge

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• Signal has a large scatter – source is not understood, laser/proton overlap jitter?

In this particular example:

- Bunch charge 2.5~3.5e+11
- Laser-proton delay 125ps (TiSa is in the 1<sup>st</sup> quarter of a proton bunch)



#### Raw WR's signal

#### ... and binned with a plasma frequency



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- Microwave transfer line length ~ 15 m
- Brings CTR signal to a shielded room
- Rectangular overmoded waveguide WR-90 (cut-off at 6.5 / 13 GHz)

See Falk's talk today, next

- Output custom horn antenna
- 2 RF beam splitters to share a signal between different detectors



Courtesy of F. Braunmueller (MPP)





- Intermediate frequency (IF) above ~10 GHz is amust for correct measurement
- Larger bandwidth of a detector facilitates search for apriori unknown CTR frequency



Numerical example of heterodyne mixing with different IF





- disadvantage

- Also a heterodyne system, based on free space ACST Schottky diode
- RF local oscillator is replaced by a photocurrent beat of two tunable CW lasers directly in the Schottky diode chip (custom development)
- Almost no band limitation : covers 50÷1000 GHz (we need 90÷290 GHz) advantage
- Less sensitive (factor ~100) than conventional heterodyne receivers







- Given a low sensitivity of the photo-mixer, multiple events had to be Preliminary !!! accumulated for every LO frequency point
- Rb density  $7.7 \cdot 10^{14}$  cm<sup>-3</sup>, plasma frequency  $f_p = 249.2$  GHz
- Measured CTR carrier frequency is  $f_{CTR} = 250.5 \text{ GHz}$



#### **ACST2** raw FFT signal

#### ... and after processing



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- We observed a strong CTR signal on all detectors on a plasma frequency as well as on its harmonics
- Signal scatter is large, no clear correlation revealed for now
- WR's detectors are not power-calibrated, difficult to interpret the data, further understanding is required
- ACST photo-mixer demonstrated a promising performance, however IF signal is pretty low. New hardware is on the way.
- Measured a CTR carrier frequency is in agreement with a frequency of OTR streak camera FFT and with a plasma frequency calculated from Rb density.
- Overall, CTR diagnostics might be in particular useful at higher Rb density, when streak-camera resolution drops down





# Thank you!

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