Self Polarization in Storage Rings

Contents

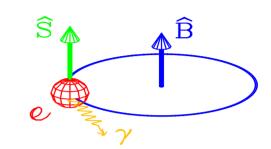
- Theoretical discovery of radiative polarization.
- Very first evidence of beam polarization (at low energy).
- Later measurements at higher energy.
- HERA collider.
- Summary

Eliana GIANFELICE (Fermilab) SPIN 2018, FERRARA



Radiative polarization

Sokolov-Ternov effect: a small amount of the radiation emitted by a e^\pm moving in the field is accompanied by $spin\ flip$.



Slightly different probabilities \rightarrow self polarization!

• Equilibrium polarization

$$|ec{P}_{
m ST} = \hat{y} P_{
m ST} \qquad |P_{
m ST}| = rac{|n^+ - n^-|}{n^+ + n^-} = rac{8}{5\sqrt{3}} = 92.4\%$$

 e^- polarization is anti-parallel to \vec{B} , while e^+ polarization is parallel to \vec{B} .

Build-up rate

$$au_{
m ST}^{-1}=rac{5\sqrt{3}}{8}rac{r_e\hbar}{m_0}rac{\gamma^5}{|
ho|^3} \quad
ightarrow \quad au_p^{-1}=rac{5\sqrt{3}}{8}rac{r_e\hbar}{m_0C}\ointrac{ds}{|
ho|^3} \;\; {
m for\; an\; \it actual\; ring}$$

Prediction of radiative polarization

A 1961 paper by Ternov, Loskutov and Korovina already predicts self-polarization

SOVIET PHYSICS JETP

VOLUME 14, NUMBER 4

APRIL, 1962

POSSIBILITY OF POLARIZING AN ELECTRON BEAM BY RELATIVISTIC RADIATION IN A MAGNETIC FIELD

I. M. TERNOV, Yu. M. LOSKUTOV, and L. I. KOROVINA

Moscow Power Institute

Submitted to JETP editor May 17, 1961

J. Exptl. Theoret. Phys. (U.S.S.R.) 41, 1294-1295 (October, 1961)

Spin flip due to radiation produced by the motion of electrons in a uniform magnetic field is considered. It is shown that an initially unpolarized beam becomes partially polarized, with the magnetic moment primarily in the direction of the field.





The famous paper by Sokolov and Ternov about electrons moving in a homogeneous constant magnetic field appeared in 1963 in russian and in 1964 in english translation

Доклады Академии наук СССР 1963. Том 153, № 5

ФИЗИКА

А. А. СОКОЛОВ, И. М. ТЕРНОВ

О ПОЛЯРИЗАЦИОННЫХ И СПИНОВЫХ ЭФФЕКТАХ В ТЕОРИИ СИНХРОТРОННОГО ИЗЛУЧЕНИЯ

(Представлено академиком Н. Н. Боголюбовым 4 VII 1963)

$$P_{\infty} = rac{n_1 - n_2}{n_1 + n_2} = 92.4\%$$

X

$$n_{1,2} = \frac{(15 \pm 8 \sqrt{3}) n_0 \mp (15 (n_{20} - n_{10}) + 8 \sqrt{3} n_0) e^{-t/\tau}}{30}.$$
 (16)

Здесь верхние знаки относятся к n_1 , а нижние к n_2 , в начальный момент времени $n_1=n_{10}$ и $n_2=n_{20}$.

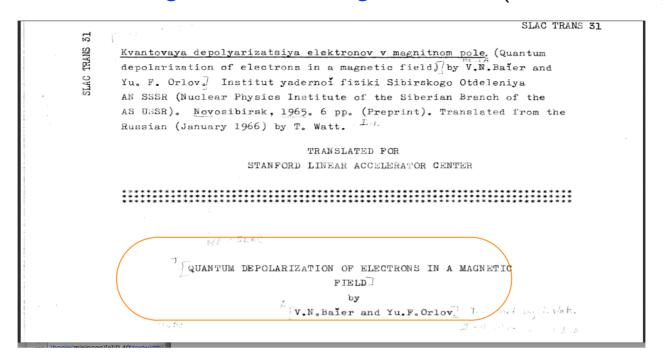
Время жизни т равно

$$\tau = \left[\frac{5\sqrt{3}}{8} \frac{\hbar}{m_0 c R} \left(\frac{E}{m_0 c^2}\right)^5 \frac{e_0^2}{m_0 c R^2}\right]^{-1} \tag{17}$$

 $\Rightarrow e^{\pm}$ -beams polarization for free!

"...it seemed that Nature had to concede a quid pro quo for the energy loss she was exacting." (B. W. Montague - 1984)

However "... Nature was being somewhat less generous..." (B. W. Montague - 1984).



By applying Thomas-BMT equation to the spin of particles circulating in an *actual* storage ring, Baier and Orlov in their 1965 paper, pointed out the existence of *depolarising resonances* activated by the stochastic nature of photon emissions (particularly important at high energy)

$$u_{spin} = m \pm m_x Q_x \pm m_y Q_y \pm m_s Q_s$$

Baier-Katkov-Strakhovenko (1970) generalized Sokolov-Ternov formulas to the case where the spin and motion direction are not everywhere perpendicular, but neglecting spin diffusion.

SOVIET PHYSICS JETP

VOLUME 31, NUMBER 5

NOVEMBER 1970

KINETICS OF RADIATIVE POLARIZATION

V. N. BAĬER, V. M. KATKOV, and V. M. STRAKHOVENKO

Novosibirsk State University

Submitted October 17, 1969

Zh. Eksp. Teor. Fiz. 58, 1695-1702 (May, 1970)

An equation describing the behavior of a spin in an external electromagnetic field is obtained in the quasiclassical approximation taking radiation effects into account. With the aid of this equation the process of radiative polarization is investigated.

$$ec{P}_{
m BKS} = \hat{n}_0 P_{
m BKS}$$
 $\hat{n}_0(s) \equiv$ periodic solution to T-BMT eq. on closed orbit

$$P_{
m BKS} = P_{
m ST} rac{\oint ds \; \hat{n}_0(s) \cdot \hat{b}(s)/|
ho|^3}{\oint ds \; [1 - rac{2}{9} (\hat{n}_0(s) \cdot \hat{v}(s))^2]/|
ho|^3} \;\;\;\; \hat{b} \equiv \hat{v} imes \dot{\hat{v}}/|\dot{\hat{v}}|$$

$$au_{
m BKS}^{-1} = rac{5\sqrt{3}}{8} \, rac{r_e \gamma^5 \hbar}{m} rac{1}{C} \oint ds \, [1 - rac{2}{9} (\hat{n}_0 \cdot \hat{v})^2]/|
ho|^3$$



The "final" step is the paper by Derbenev and Kondratenko (1973).

Polarization kinetics of particles in storage rings

Ya. S. Derbenev and A. M. Kondratenko

Institute of Nuclear Physics, Siberian Division, USSR Academy of Sciences (Submitted September 22, 1972)
Zh. Eksp. Teor. Fiz. **64**, 1918-1929 (June 1973)

A closed description of the radiative kinetics of the polarization of charged particles with arbitrary spin and magnetic moment is presented, and takes spin-orbit coupling into account. The analysis is based on an investigation of the dynamics of spin motion in inhomogeneous fields [8,10]. For ultrarelativistic electrons (positrons) the paper combines the results of a number of investigations [1-6] and contains some new effects due to spin-orbit coupling in an inhomogeneous field. The time constant and degree of equilibrium polarization of a beam in storage rings with arbitrary fields are found for nonresonant conditions. The method developed in the paper may also be applied when the perturbing electromagnetic field is related to an "external" source.

They included spin diffusion by using a semiclassical approach.

$$ec{P}_{
m DK} = \hat{n}_0 rac{8}{5\sqrt{3}} rac{\oint ds < rac{1}{|
ho|^3} \hat{b} \cdot (\hat{n} - rac{\partial \hat{n}}{\partial \delta}) >}{\oint ds < rac{1}{|
ho|^3} \left[1 - rac{2}{9} (\hat{n} \cdot \hat{v})^2 + rac{11}{18} (rac{\partial \hat{n}}{\partial \delta})^2
ight] >} \qquad \hat{b} \equiv \hat{v} imes \dot{\hat{v}}/|\hat{v}|$$

periodic solution to T-BMT eq. on c.o.

randomization of particle spin directions due to photon emission $(\delta \equiv \delta E/E)$

Polarization rate

$$au_{
m DK}^{-1} = rac{5\sqrt{3}}{8} rac{r_e \gamma^5 \hbar}{m_0 C} \oint ds < rac{1}{|
ho|^3} \Big[1 - rac{2}{9} (\hat{n} \cdot \hat{v})^2 + rac{11}{18} \Big(rac{\partial \hat{n}}{\partial \delta}\Big)^2 \Big] >$$

These formulas involve averaging over the beam distribution.

Final? There have been some disputes about the meaning of the quantity \hat{n} and $\partial \hat{n}/\partial \delta$ in the original paper.

Many authors (S. Mane, K. Yokoya, D. P. Barber, G. Hoffstätter, M. Vogt...) have contributed to give a rigorous mathematical definition.

Nowadays

- \hat{n} is understood as an *invariant spin field* i.e. a solution of T-BMT eq. satisfying the condition $\hat{n}(\vec{u};s)=\hat{n}(\vec{u};s+C)$.
- The term $\partial \hat{n}/\partial \delta$ quantifies the depolarizing effects resulting from the trajectory perturbations due to photon emission.

The computation of Derbenev-Kondratenko expressions in the general case is tricky and codes attempting to evaluate them have limitations.

Tools for radiative polarization computation

- By linearizing orbit and spin motion it is possible to calculate polarization "analytically" in terms of one turn maps. This formalism has been developed at the end of the 70s by A. Chao (SLIM) and K. Yokoya. A thick lenses version of SLIM is D. P. Barber SLICK.
 - Only linear resonances!
- S. R. Mane wrote SMILE (middle 80s) handling fully 3D spin motion in *perturbation* theory. This approach required large computing time and had convergence problem at high energy.
- SODOM by K. Yokoya (1992) computes Derbenev-Kondratenko \hat{n} and $\partial \hat{n}/\partial \delta$ using Fourier expansions. It has similar issues as SMILE.

• Instead of trying to evaluate Derbenev-Kondratenko expression, a *statistical* approach is used in SITROS by J. Kewisch (1982). Orbital motion is up to 2d order and spin motion is <u>not</u> linearized. Initially fully polarized beam is tracked and stochastic photon emission is simulated by random emission of "big photons" at user selected machine dipoles. As polarization evolves as

$$P(t) = P_{\infty}(1 - e^{-t/\tau_p}) + P(0)e^{-t/\tau_p}$$

with

$$rac{1}{ au_p} \simeq rac{1}{ au_{
m BKS}} + rac{1}{ au_{
m d}} \qquad ext{and} \qquad P_{\infty} \simeq rac{ au_p}{ au_{
m BKS}} P_{
m BKS}$$

SITROS evaluates au_p from P(t) and P_{∞} from $rac{ au_p}{ au_{
m BKS}}P_{
m BKS}$

 More recently Polymorphic Tracking Code (PTC) by E. Forest has been used for implementing this statistical approach for CEPC by Zhe Duan et al.



First measurements of radiative polarization

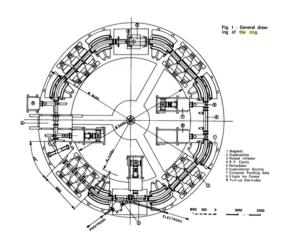
As reported by Belbeoch $et\ al.$, at 1968 USSR Nat. Conf. Part. Acc., the very first observation was in 1968 at ACO (Anneau de Collisions d'Orsay) in operation as e^+/e^- collider from 1965 to 1973.

Available on-line

STATUS REPORT ON ACO

The Orsay Storage Ring Group^{†)}
Laboratoire de l'Accélérateur Linéaire, Université de Paris-Sud, Centre d'Orsay, Orsay, France.
(presented by D. Potaux)

(8th International Conference on High-Energy Accelerators, CERN, 1971)



Polarization measurement exploited the spin dependence of Coulomb scattering crosssection.

The counting rate \dot{n} is proportional to the square of the current, to the inverse of the volume of the bunch (V = $\Delta X \times \Delta Z \times \Delta \ell$) and to the cross section for Coulomb scattering within the bunch, integrated over the acceptance of the system :

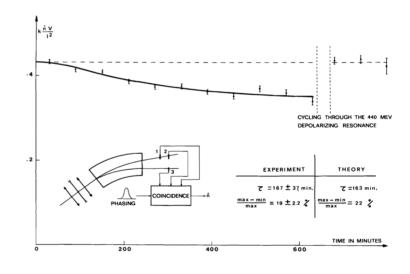
$$\dot{n} = k \frac{I^2}{\Delta X \times \Delta Z \times \Delta \ell} \times \sigma$$

If polarization occurs the normalized counting rate ${\tt Y}$

$$Y = \dot{n} \frac{\Delta X \Delta Z \Delta \ell}{T^2}$$

should exhibit a typical variation with time of the form :

$$Y = a - b \left(1 - e^{-t/\tau}\right)^2$$
.



536 MeV positrons

- Polarization was observed in 1972 also at VEPP-2.
- Polarization close to 92% (!) was later measured at
 - ACO (1973)
 - VEPP-2M (1976) at 500 MeV

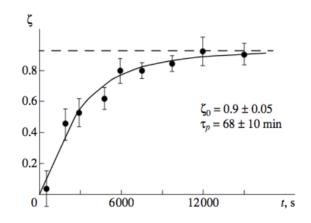


Fig. 3. The polarization buildup in VEPP-2M (1976).

(from Yu. M. Shatunov, Polarized Beams at Storage Rings, 2006)

Far from resonances, spin diffusion is negligible at low energy!

Beam energy calibration

At VEPP-2M the method of beam energy measurement by *resonant depolarization* was first invented.

Particle Accelerators 1980 Vol.10 pp.177-180 0031-2460/80/1003-0177\$06.50/0 © Gordon and Breach, Science Publishers, Inc.
Printed in the United States of America

ACCURATE CALIBRATION OF THE BEAM ENERGY IN A STORAGE RING BASED ON MEASUREMENT OF SPIN PRECESSION FREOUENCY OF POLARIZED PARTICLES*

YA. S. DERBENEV, A. M. KONDRATENKO, S. I. SEREDNYAKOV, A. N. SKRINSKY, G. M. TUMAIKIN, and YU. M. SHATUNOV

Institute of Nuclear Physics, Siberian Division USSR Academy of Sciences, Novosibirsk 90, USSR

(Received January 29, 1979)

A method is described for measuring the particle energy in an electron-positron storage ring by means of resonance depolarization by a high frequency field. The measurement accuracy is discussed taking into account energy spread and synchrotron oscillations. It is found that in practice the limitation in accuracy is due to the irregular pulsations of the magnetic guide field. As a result, the electron beam energy in the storage ring VEPP-2M has been measured with an accuracy of ±2.10⁻⁵.

They used a *longitudinal* magnetic field for depolarizing the beam.

Depolarization occurs when spin precession and field frequency are in resonance

$$u_{spin} = rac{f_{exc}}{f_{rev}} + k$$

ullet $u_{spin}=a\gamma$ (in a planar ring, w/o solenoids) $o E_{beam}=ig[k\pmrac{f_{exc}}{f_{rev}}ig]rac{E_0}{a}$



Later on polarization was observed at

• SPEAR (SLAC) where a Compton polarimeter was built for SPEAR2

(A. Chao, Polarization of a stored electron beam, 1981)

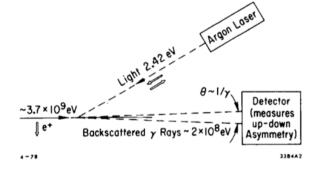


Fig. 4. Schematic diagram of SLAC-Wisconsin laser polarimeter. 18

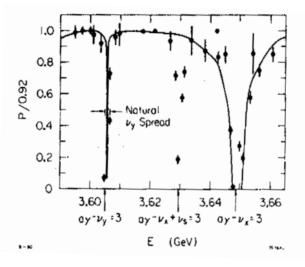
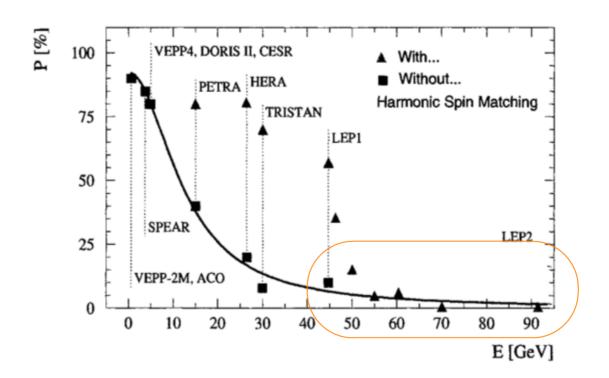


Fig. 8 Comparison of calculation and measurements for SPEAR. Agreement near the two linear sideband resonances $a\gamma - v_y = 3$ and $a\gamma - v_x = 3$ is acceptable. The third resonance $a\gamma - v_x + v_s = 3$, however, is missed by the calculation.

(R. F. Schwitters, Experimental review of beam polarization in high energy e^+e^- storage rings, 1979)

• DORIS II (beam energy calibration), PETRA (where "harmonic spin-matching" was invented), CESR, TRISTAN, LEP (beam energy calibration).

Polarization became ...fashionable!



(R. Assmann et al., Spin dynamics in LEP, Osaka 2000)

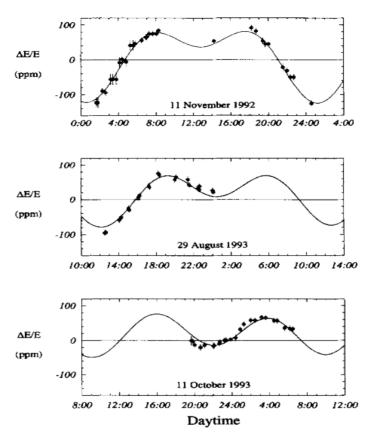


Fig. 3. The evolution of the relative beam energy variation due to tides is shown as a function of time for three periods with stable beam conditions. The solid line is calculated using the CTE tide model with the average coefficient from Eq. (4). The top picture corresponds to full-moon, the bottom picture to a time close to half-moon. Relative beam energy variations of up to 220 ppm are observed on November 11th 1992.

(R. Arnaudon et al., Effects of terrestrial tides on the LEP beam energy, 1995)

Radiative polarization at HERAe

HERA was a 6.3 km p/e^{\pm} collider operating in Hamburg from 1992 to 2007



- 2 collider experiments: H1 and Zeus
- ullet 2 single beam experiments: Hermes (e^\pm) and Hera-B (p)



e^{\pm} polarization at HERA

The e^\pm ring was conceived from the beginning on for delivering beam longitudinal polarization

- planar geometry
- large number of BPMs and orbit correctors
- spin rotators for longitudinal polarization
 - large number of independently powered quadrupoles for spin matching















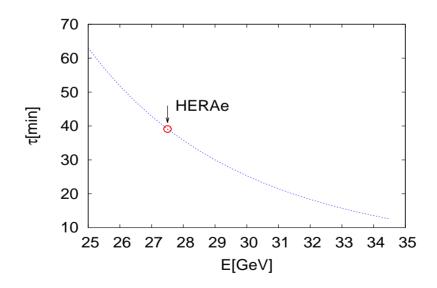


The e^\pm beam polarization relied on Sokolov-Ternov effect. Polarization rate

$$au_p^{-1} = rac{5\sqrt{3}}{8} rac{r_e \gamma^5 \hbar}{m_0 C} \oint rac{ds}{|
ho_b|^3}$$

which for Hera- e^{\pm} with $ho_b \simeq$ 608 m gives

$$au_p[\mathrm{sec}] = rac{0.37 imes 10^{11}}{E[\mathrm{GeV}]^5}$$



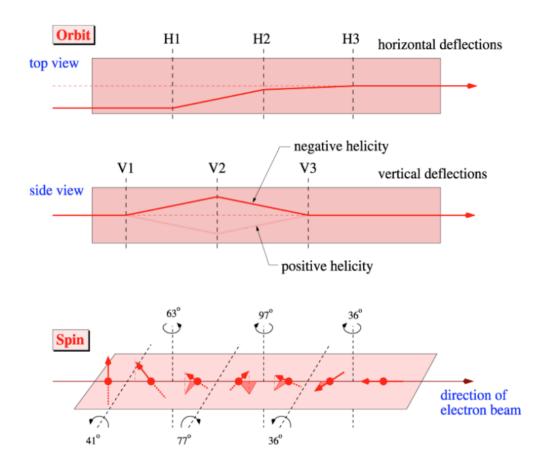
HERAe Mini-rotator

Several designs considered. Final choice: Steffen-Buon "mini-rotator".

- Chain of interleaved horizontal and vertical bending magnets.
- Short enough (\simeq 56 m): no quadrupoles needed!
- Spin helicity changed by flipping the sign of the vertical bending magnets:
 - magnets mounted on remotely controlled jacks for adjusting their elevation w/o entering the tunnel.
- Designed to operate between 26.8 and 35 GeV by changing the magnet settings:
 - energy changes larger then ± 100 MeV required a manual machine re-alignment.



Orbit and \hat{n}_0 at 27.5 GeV, $\hat{y} \simeq$ 200 mm







HERAe Mini-rotator in place on the vertically movable jacks



















Rotator spin matching

A flat machine is *spin transparent*. The presence of rotators breaks the transparency:

- $\hat{n}_0 \neq \hat{y}$ between rotator pairs leads to spin diffusion in the involved quadrupoles;
- vertical dispersion introduced by vertical bending magnets or by solenoids (through coupling) leads to a finite vertical beam dimension \sim spin diffusion in the quadrupoles of the whole ring.

Spin transparency for *linear* spin/orbit motion may be recovered by *spin matching*.

It consists in designing the focusing structure so that the spin direction does not depend on the particle orbital coordinates. For the HERAe mini-rotator with

- $D_y \neq 0$ only at the rotators
- ullet \hat{n}_0 non vertical only between the rotator pair

assuming symmetric focusing around IP and arc center \sim 5 additional optics conditions which in terms of Twiss functions write

$$\int\limits_{IP}^{s_R}ds\,K\sqrt{eta_x}\cos\mu_x=0 \qquad \int\limits_{IP}^{s_R}ds\,\hat{n}_0\cdot\hat{s}\,K\sqrt{eta_y}\cos\mu_y=0 \ \int\limits_{IP}ds\,e^{i(\psi\pm\mu_y)}K\sqrt{eta_y}=0 \qquad \int\limits_{IP}^{s_R}ds\,D_xK=0$$

 s_R = rotator entrance

 $\psi =
u_s imes$ cumulative bending angle

Spin diffusion due to random errors and counter-measures

In a real machine spin transparency is also randomly broken by magnet misalignments

- Quadrupoles
 - vertical displacement

$$* D_y \neq 0$$

$$* \delta \hat{n}_0 \neq 0$$

– roll

*
$$D_y \neq 0$$
 if $D_x \neq 0$

- * betatron motion coupling $\sim \epsilon_y \neq 0$
- Horizontal bending magnets
 - roll $\leadsto D_u \neq 0$ and $\delta \hat{n}_0 \neq 0$



The most dangerous is the vertical misalignment of quadrupoles:

- Usual closed orbit correction cured spurious vertical dispersion at HERAe. LEP used a *dispersion free* correction.
- ullet tilt of \hat{n}_0 wrt nominal direction required special care!

 \hat{n}_0 tilt due to dipolar errors on the design orbit (with $\psi(s)=2\pi
u_s s/C$):

$$\delta\hat{n}_0(s) = rac{ie^{i[\psi(s)-\pi
u]}}{2\sin\pi
u} \int\limits_{s-C}^s ds' e^{-i\psi(s')} f(s')$$

with f "spin-orbit" function

$$egin{aligned} egin{aligned} \Re(f) \ \Im(f) \end{aligned} &= \mathsf{L} \left[\mathsf{F} ec{y} - rac{1}{B
ho} egin{pmatrix} \Delta B_s(1+a) \ \Delta B_x(1+a\gamma) \ \Delta B_y(1+a\gamma) \end{aligned}
ight] \end{aligned}$$

 $\Delta \vec{B}=$ extra fields along the design c.o., $\mathbf{L}=\mathbf{2}\times\mathbf{3}$ matrix of \hat{m} and $\hat{\ell}$ components in the orbital reference system, \mathbf{F} is a $\mathbf{3}\times\mathbf{6}$ (energy dependent) matrix containing the nominal fields and \vec{y} is the 6-dimensional c.o.



Harmonic expansion

$$\delta \hat{n}_0(s) = -irac{L}{2\pi}\sum_krac{f_k}{k-
u_s}e^{i2\pi ks/C}$$

- ullet \hat{n}_0 tilt increases linearly with energy
- ullet \hat{n}_0 tilt is more sensitive to the harmonics of the spin-orbit function f close to u_s .

The correction consists in compensating these harmonics by powering some of the usual correction coils. In practice using *closed orbit bumps* in the arc cells allows to correct $\delta \hat{n}_0$ w/o perturbing the orbit everywhere \sim harmonic bumps









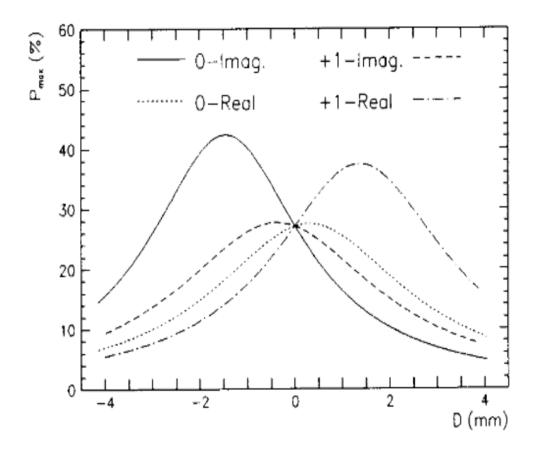






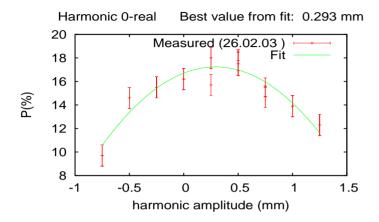
Correction done empirically by scanning the most important harmonic components.

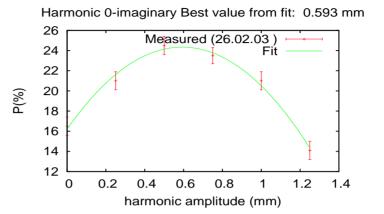
Expected effect on polarization of harmonic bumps scans:

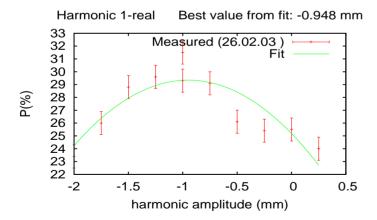


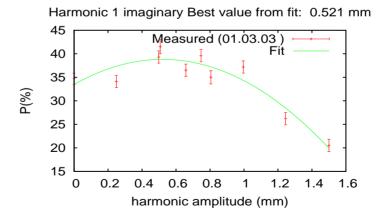


...and real life:

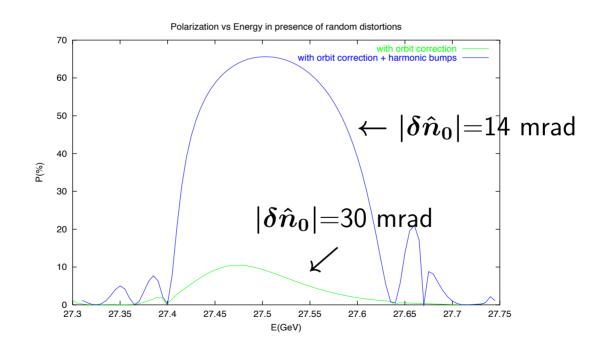








Expected effect of harmonic bumps on HERAe polarization (simulation with δQ_{rms}^y =0.3 mm):





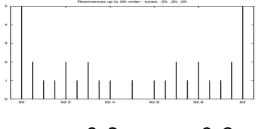
Beam parameters choice

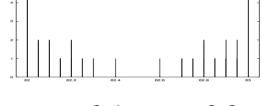
Spin diffusion is larger when the spin-orbit resonance conditions are met

$$u_s = m \pm m_x Q_x \pm m_y Q_y \pm m_s Q_s$$

→ Small fractional parts of betatron tunes are convenient.

Resonances position

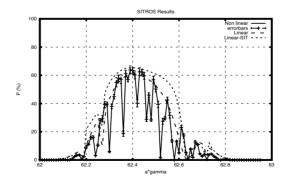


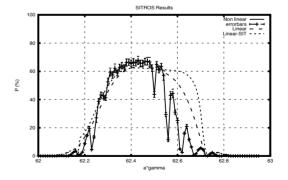


$$q_x = 0.2$$
 $q_y = 0.3$



SITROS simulations





$$q_x = 0.2$$
 $q_y = 0.3$

$$a_{u} = 0.3$$

$$q_x = 0.1$$
 $q_y = 0.2$

$$q_y = 0.2$$













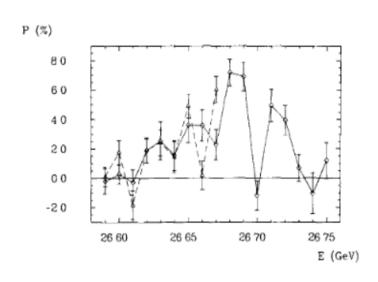


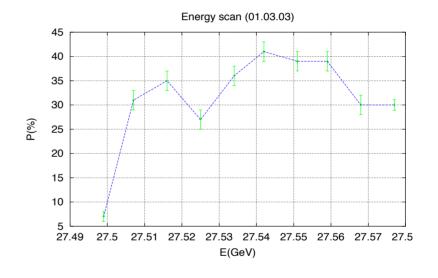






- Same order resonances are not equally strong
- Beam energy not exactly known
- \rightarrow Beam energy scan allows finding best spot for polarization.





1991 very first energy scan

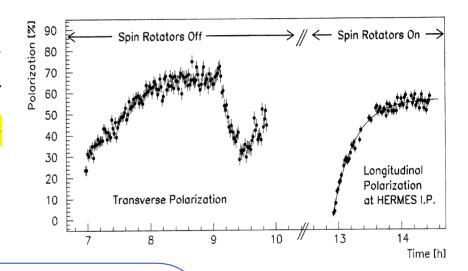
A 2003 scan with 3 rotators pairs



HERAe polarization milestones

- November 1991: Compton polarimeter in HERA West brought into operation and first observation of transverse electron beam polarization, $\simeq 10\%$, at 26.66 GeV w/o dedicated optimizations.
- 1992: H1 and ZEUS start data taking; after re-alignment of some machine magnets and correction of beam ellipse tilt, polarization increased to $\simeq 18\%$.
- June 1992: after dedicated machine tuning (energy and harmonic bumps scans, tunes optimization) 60% polarization achieved routinely!
 - Approval of HERMES and installation during 1993/1994 shut down of a pair of Buon-Steffens spin rotators around IP East.

- May 1994:
 - machine energy increased to 27.5 GeV for operating the rotators
 - dedicated polarization optimization with rotators still off \sim 65% polarization.
 - However still some people were worried...
- May 4: Rotators turned on, polarization reached 56% and 65% after some re-tuning! First time achievement of longitudinal polarization in a high energy storage ring!



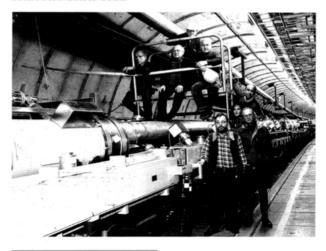






Nouvelles des Laboratoires

L'équipe du rotateur de spin au collisionneur électron-proton HERA de DESY. Les aimants rotateurs de spin spéciaux pour l'anneau d'électrons sont visibles sur la droite



DESY Succès de la polarisation longitudinale des électrons

e 4 mai l'anneau d'électrons de 6,3 km du collisionneur électron-proton HERA à DESY, Hambourg, a maintenu en circulation un faisceau d'électrons longitudinalement polarisés c'est-à-dire dont les électrons individuels, comparables à de petites toupies, avaient leurs spins alignés sur la direction de leur mouvement et de même sens ou de sens contraire.

La polarisation longitudinale de ces électrons rend l'asymétrie gauchedroite inhérente aux interactions faibles plus faciles à observer et ouvre la porte à de nouvelles mesures de précision.

C'est la première fois qu'un tel faisceau d'électrons de haute énergie a été produit: normalement les faisceaux d'électrons sont polarisés transversalement, c'est-à-dire avec leurs

spins orientés perpendiculairement à la direction du mouvement.

Des aimants rotateurs de spin spéciaux sont utilisés pour ces manipulations: ils ont été installés dans l'anneau HERA l'hiver dernier après avoir été mis au point sous la direction du regretté Klaus Steffen en collaboration avec Jean Buon de Saclay.

Ces aimants communiquent une série d'impulsions horizontales et verticales qui renversent les spins mais ensemble ne produisent aucune déflexion nette de l'orbite des électrons.

De nombreux sceptiques avaient pensé que cette polarisation serait difficile sinon impossible à obtenir du fait de la sensibilité potentielle de ces exercices délicats à de petites perturbations du faisceau. La première session a produit un degré de polarisation de 55%

Les aimants rotateurs de spin font pivoter la polarisation naturelle transversale (verticale) des électrons: dans le mouvement circulaire de ces derniers, leur spin, comme l'aiguille d'une boussole, tend à s'aligner sur le champ magnétique de l'anneau de stockage. Cependant même cette polarisation naturelle est à la merci des résonances dépolarisantes qui ont tendance à

déranger les spins. La disponibilité d'électrons polarisés longitudinalement est comme un lever de rideau pour l'expérience HERMES (décembre 1993, page 19) utilisant une cible interne dans le faisceau d'électrons d'HERA pour étudier le rôle du spin du quark dans celui des nucléons. HERMES doit commencer à saisir des données l'an prochain.

Des aimants rotateurs de spin supplémentaires seront installés dans le faisceau d'électrons d'HERA de sorte que les grandes expériences ZEUS et H1 pourront aussi bénéficier de la

Noyaux magiques Etain-100 étincelant

De même que le tableau périodique des éléments chimiques reflète le remplissage des couches successives d'électrons orbitaux, les nombres dits magiques de la physique nucléaire correspondent à des couches fermées de 2, 8, 20, 28, 50, 82, 126, ... neutrons et/ou protons. Plus fortement liés que les autres novaux, ils sont les analogues nucléaires des gaz inertes. Les noyaux "doublement magiques" possèdent des couches fermées de neutrons aussi bien que de protons. Les exemples présents dans la nature sont l'hélium-4 (2 protons et 2 neutrons), l'oxygène-16 (8 et 8), le calcium-40 (20 et 20) et le calcium-48 (20 et 28). L'étain radioactif-132 (50+82) a été largement étudié.

Dans cette liste les noyaux "recherchés" comprennent l'oxygène 28 (8 et 20), le nickel 78 (28 et 50) et l'étain-100 (50 et 50). Ce dernier vient d'être observé dans des études sur ions lourds au laboratoire français GANIL, Caen, et au laboratoire allemand GSI, Darmstadt. Bien que ces observations confirment comme prévu la stabilité de l'étain-100, il faudra attendre de disposer d'intensités plus élevées pour en étudier réellement les propriétés physiques.

C'est le laboratoire Bévalac à Berkeley qui a lancé la production de noyaux exotiques par fragmentation de

Courrier CERN, juillet/août 1994





















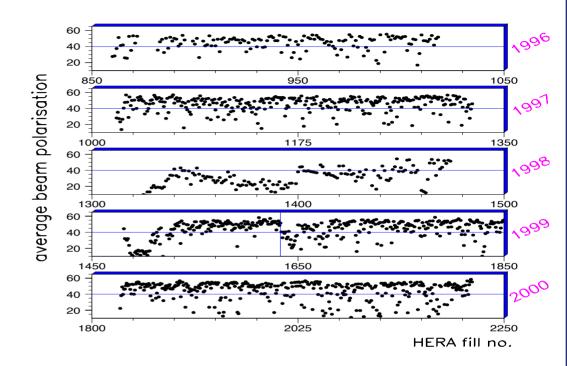






 HERMES started data taking in 1995.

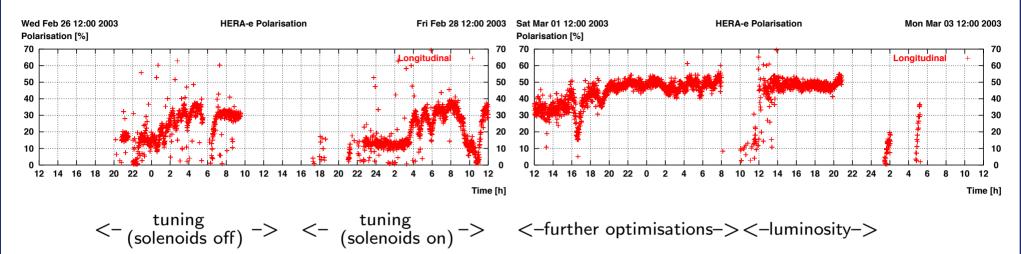
1996: Longitudinal polarimeter in HERA East, bunch-by-bunch measurement!

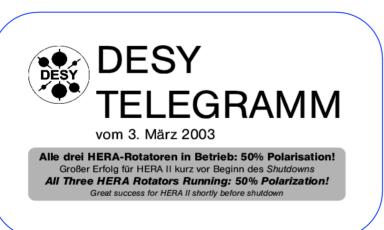


(E. Aschenauer courtesy)

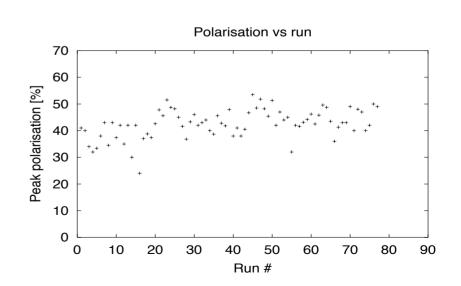
• 2 more pairs of spin rotators installed for H1 and ZEUS during the 2000/2001 shut down for the machine luminosity upgrade, while the compensating solenoids were removed because of lack of space!

February 2003: First polarization with 3 rotators pairs!





Achieved peak polarization in 2004:





Finding a trade-off between luminosity and polarization wasn't easy!



SPIN IS IN

Figure1: Bryan Montague's cartoon from the 1980 Spin Symposium in Lausanne



















Reliability: sharpen your skills!

- Beam parameters choice
 - Experimental Bkg tuning (local bumps, betatron tunes...)
 - Maximizing luminosity (betatron tunes...)
- Limited polarization tuning possibility during luminosity operation: keeping (vertical) closed orbit below 1 mm rms value w/o affecting the luminosity \rightsquigarrow "lumicor".
- Meticolous book keeping
 - record of golden orbit;
 - tracking of beam energy changes due to different horizontal corrector settings;
 - run-to-run tracking of harmonic bump settings;
 - analysis of harmonic content of changed vertical corrector settings.



Summary

We have reviewed the milestones of Radiative Polarization in storage rings:

- Prediction by Sokolov-Ternov (1963), after 1961 Korovina et al. paper.
- Very first observations at ACO (1968) and VEPP-2 by exploiting loss rate dependence on polarization.
 - Large polarization was observed at low energy, w/o special corrections.
 - In PETRA at 15 GeV dedicated orbit correction was successfully experimented.
- Use of beam polarization for precise energy calibration at VEPP-2M, DORIS II, LEP1 with the tides observation!
- The unique experience of the HERA p/e^{\pm} collider where longitudinal lepton beam polarization was an integral part of the physics program.
 - Now the adventure continues in the US with the EIC design pursued by BNL and JLab!



THANKS!



























BACK UP SLIDES

















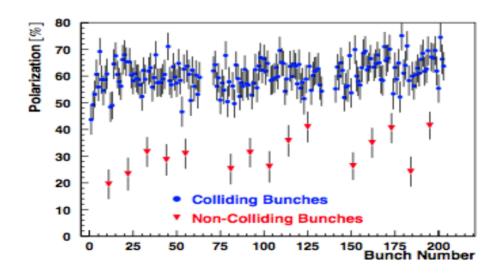






Beam-beam effects on HERAe Polarization

- ullet Extra non-linear lens (o tune shift and spread)
- Emittance increase
- Stochastic kicks

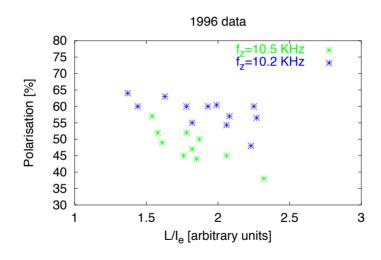




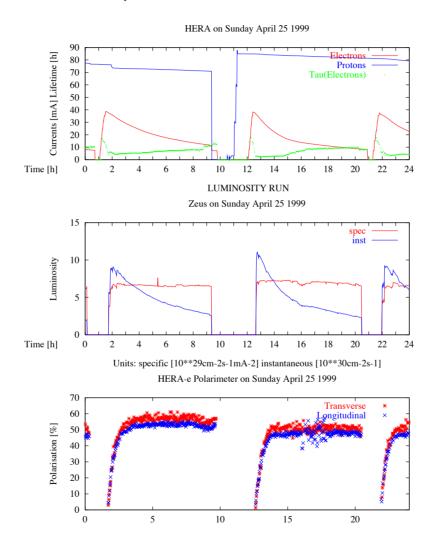
1996: First hints of beam-beam effects on HERAe polarization when

- $\beta_z^p = 0.7 \text{ m} \rightarrow 0.5 \text{ m}$
- $I_p \simeq 70$ mA.

A Q_y change by -6×10^{-3} allowed to recover polarization!

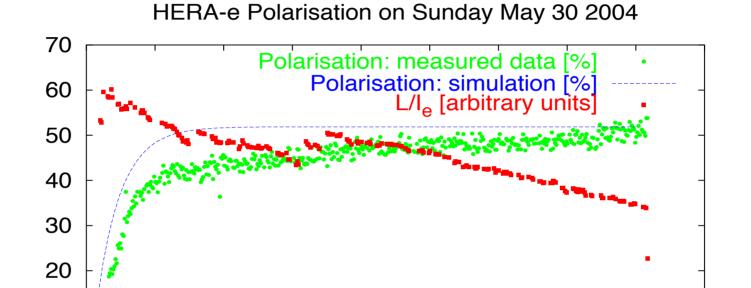


Larger the luminosity, smaller the polarization!





Hints of beam-beam effect on polarization: polarization grows more slowly than expected from Sokolov-Ternov effect:

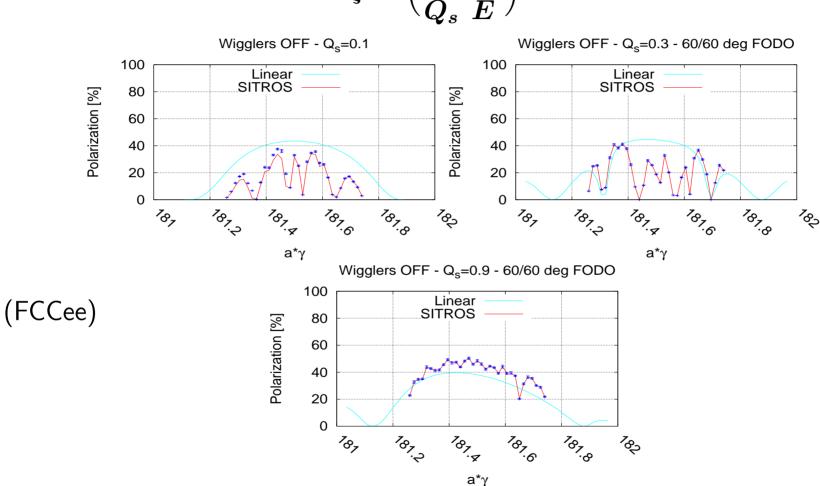


Time [h]

Importance of synchrotron tune

Energy spread enhancement of synchrotron side bands

$$\xi = \left(rac{a\gamma}{Q_s}rac{\sigma_E}{E}
ight)^2$$



Rotator spin matching

A flat machine is *spin transparent*. The presence of rotators breaks the transparency.

- Vertical dispersion introduced by vertical bending magnets or by solenoids (through coupling) leads to a finite vertical beam dimension → spin diffusion in the whole ring quadrupoles
- $\hat{n}_0 \neq \hat{y}$ between rotator pairs leads to spin diffusion in the involved quadrupoles

Spin transparency may be recovered by *spin matching*: it consists in designing the focusing structure (quadrupoles) so that globally the spin direction does not depend on the orbital coordinates.

 $G_{2 \times 6}$ matrix relating spin orientation wrt \hat{n}_0 to the 6 orbital coordinates

$$egin{pmatrix} \Deltalpha \ \Deltaeta \end{pmatrix} = G\,ec{y}$$

In a flat machine:

$$G_x \equiv \underline{0}$$
 $G_s \equiv \underline{0}$ $G_y
eq \underline{0}$ but $y = 0$



With rotators

$$y \neq 0$$

and

 G_x and $G_s
eq \underline{0}$ between rotator pairs

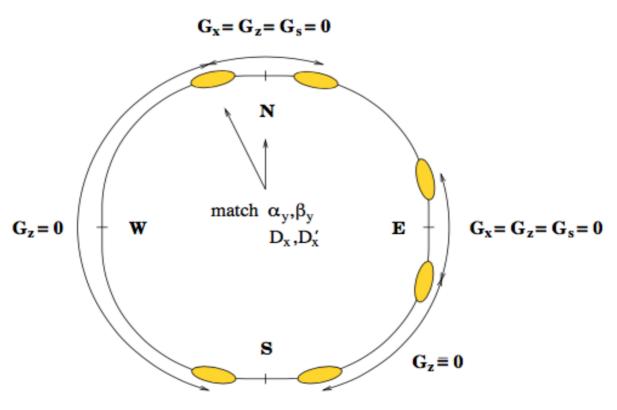


Figure 6.7: Spin matching conditions in the HERA upgrade lattice. The shaded ellipses represent the rotators.

(from M. Berglund PhD Thesis)

In terms of Twiss functions the conditions for spin transparency are

$$\Delta_{\pm x} = \Delta_{\pm y} = \Delta_{\pm s} = 0$$

$$egin{align} \Delta_{\pm x, \pm y}(s) &= (a\gamma + 1) rac{e^{\mp i \mu_{x,y}}}{e^{2i\pi(
u \pm Q_{x,y})} - 1} rac{[-D \pm i (lpha D + eta D')]_{x,z}}{\sqrt{eta_{x,y}}} J_{\pm x, \pm y}(s) \ & \Delta_{\pm s} = (a\gamma + 1) rac{e^{\pm i \mu_{s}}}{e^{2i\pi(
u \pm Q_{s})} - 1} J_{s} \ \end{matrix}$$

with

$$J_{\pm x,\pm y} = \int\limits_s^{s+L} ds' (\hat{m}_0 + i\hat{l}_0) \cdot egin{cases} \hat{y}\sqrt{eta_x} \ \hat{x}\sqrt{eta_y} \end{pmatrix} K e^{\pm i\mu_{x,y}} \ J_s = \int\limits_s^{s+L} ds' (\hat{m}_0 + i\hat{l}_0) \cdot (\hat{y}D_x + \hat{x}D_y) K \end{cases}$$

$\delta \hat{n}_0$ correction for a nominally planar ring

$$\delta \hat{n}_0(s) = -irac{1+a\gamma}{2\sin\pi
u_s}e^{i[\psi(s)-\pi
u_s]}\int\limits_{s-C}^s\,ds'\,y_{co}''e^{-i
u_s\phi_B(s')}$$

 $\phi_B \equiv$ cumulative bending angle.

If in addition the machine is made out only of arcs where $\nu_s \phi_B(s) \simeq 2\pi \nu_s s/L$, the main contributions come from the closed orbit harmonics near to ν_s .

At LEP those harmonics were extracted from the BPMs reading and a "deterministic" correction applied.

The method used at HERA is general and accounts also for the contribution of the horizontal closed orbit to the tilt of \hat{n}_0 .

