



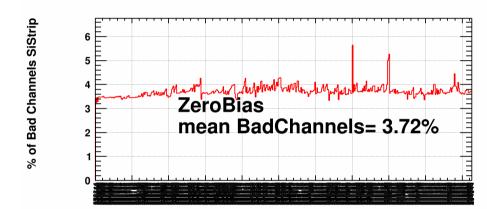
# Strip Performance, calibration/alignment & commissioning: plans for 2017 Gaëlle Boudoul On behalf of the Tracker Project

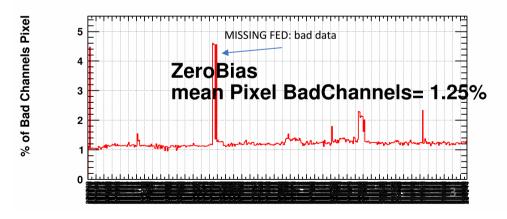
## Looking back over 2016 ...



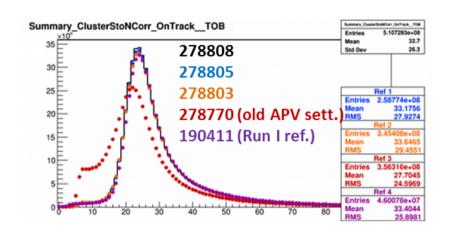
## Tracker data taking efficiency (pixel + Strips)

- Data losses due to Tracker: ~785 pb<sup>-1</sup> (~2% of the luminosity)
  - Pixel data-taking inefficiencies: ~183 pb<sup>-1</sup>
  - Strip data-taking inefficiencies: ~220 pb<sup>-1</sup>
  - Useless Tracker data: 382 pb<sup>-1</sup>
    - Mostly due to Strip (316 pb<sup>-1</sup>) and Pixel (66 pb<sup>-1</sup>) detectors not in global run while fixing problems
- Causes of the largest losses
  - Cooling plant valve and sensor failures
  - Pixel power supply and Strip ACDC converters (aka MAOs)
  - FED crate power supply
  - Pixel optical link baseline instabilities
  - Strip FEDs in "disconnected state"
  - Pixel stuck in "SoftErrorRecovery"

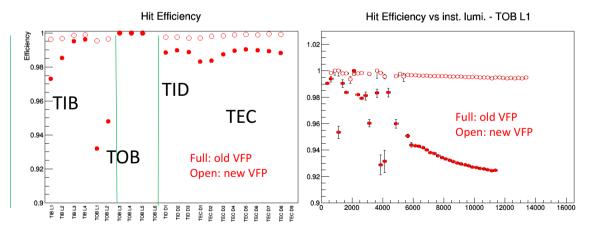


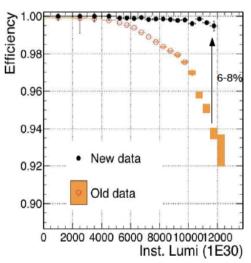


#### Strip Performance after the VFP change



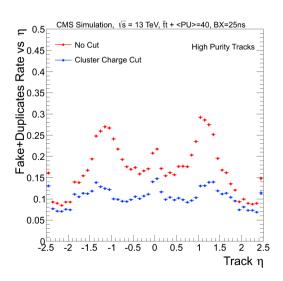
- Hit signal restored
- Hit efficiency restored
- Track length and efficiency restored
  - Since 2016H era
  - Previous data re-reconstructed with a less demanding track reconstruction for the double sided modules and a different btagging track selection

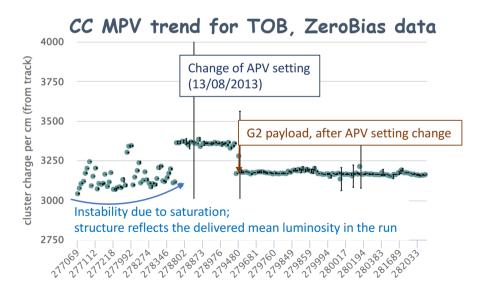




#### Strip Calibrations

- Strip gain calibration
  - Stability of the cluster charge distribution needed because of the "cluster charge cut" used to reduce the out of time pileup
    - Reduced fake rate, faster reco
  - Scale factors computed from opto-link gain measurements and with collision data (G1 and G2)
    - Virtually impossible to calibrate with collision data with the old VFP setting!!
    - Everything OK with VFP=0





## Looking forward 2017...



#### Plans for 2017 - Strips Detector Calibration

- Smooth running ©
- Recalibrate detector before data taking
  - Laser Gain
  - FFD thresholds
  - APV baseline
  - APV pulse shape
  - Pedestal and noise
- Recalibrations **over the year**:
  - Pedestal and noise thresholds
    - $\rightarrow$  will try to move to 1x month (was 4/year)
  - Laser Gains + FED thresholds (need to be done together): 3x in 2016
    → need to increase frequency to cope smoothly with radiation effects
  - In addition: FED threshold measurement intimately related to G1 gain
     → aim to move this into an O2O procedure to have detector conditions and offline reconstruction in sync by construction.
    - See next slide

#### Gain calibration: G1 factor

#### G1 gain factor: status of the tools and operations in 2016

- Single CMSSW analyzer that produces the G1 payload from the online tickmark run data;
  - tickmark data retrieved through WBM interface;
  - analysis of signal height versus noise: identification of problematic channels
- Run manually when notified by the online crew that new tickmark runs are available and processed.
- Full Alca Validation of the payload required
  - $\rightarrow$  in short : a lot of manual and human interventions (including central teams) , we would like to improve and reduce the load

#### • 2017: Milestones for G1 improvements:

- access online validation infos to identify faulty channels
- Transfer the measurements using the "O2O" procedure
- define set of DQM histograms to be checked by shifters;
- Goal is to set this workflow in production by end of April

#### Plans for 2017 - special runs with beam

#### Timing scan early in the year for validation

- 3-5 hours with beam
- Pixels in the run and efficient
- ~ 1 kHz of ZeroBias on disk

#### Bias voltage scans

- Full scans 2x per year
  - 1.5 2 hours
  - Pixels in the run and efficient
  - >~ 1 kHz of ZeroBias on disk
- Small scan 1x month (during physics)

### APVE emulator and latency

- The APVe fulfills three main functions
  - Emulation of the APV pipeline to prevent buffer overflow in the APVs
  - Transmission of read out pipeline address for L1As to tracker FEDs via b-channel data for data integrity checks
  - Blocking of triggers in certain positions in the APV readout cycle as large noise is induced in certain pipeline locations
- The full APVe functionality was moved from a separate electronic board to TCDS after LS1
- Veto mask was recommissioned and correctly placed in the TCDS CPM in 2016

- With the APVe in the CPM, there is no transmission and decoding delay when the APVe asserts busy to prevent buffer overflow
- → With the APVe in the CPM, busy threshold could maybe be lowered
  - This by itself will mainly benefit deadtime
  - Measurements were made at the end of run 1 to see what maximum latency we can tolerate without too much increase in deadtime
    - with APVe in CPM these measurements could be revisited
- -> Could mean that latency of APVs MIGHT be shifted IF necessary for L1 trigger upgrade
  - Has to be done with a lot of care
    - If the APV buffer overflows, data will be garbage until next resync

#### Event Size in the Strip Tracker

- No problem with input data rate into FEDs
  - Non-sparsified data from APV chips
  - → data rate = constant
- Strip tracker event size expected to increase ~linearly with pile-up
  - No problem with output rate from FED point of view  $\rightarrow$  output rate up to 640 MB/s
- Some problems observed in 2016 with increased event size in TID
  - Caused increase in dead time because of backpressure from DAQ
- Total event size for scenarios under consideration by high lumi WG (estimates based on #PV not PU)

	1.8E34 (PU 52)	2.0E34 (PU 60)	2.25E34 (PU 70)
Min-Max per FED	0.6-2.2kB/ev/FED	0.7-2.5kB/ev/FED	0.75-2.9kB/ev/FED
Total	~850 kB	~960 kB	~1090 kB

#### Shift Operation

- Strips DOC remains single point of contact for strips operations as before
  - Reports in daily run meetings
  - Reports in strips operations meetings
  - Performs (shared with pixels) comprehensive monitoring of system
  - check list (DAQ, DCS, DQM) at start of physics fill (extended working hours)
  - Will see if taking of calibration runs can be integrated into DOC duties during 2017
- Offline Shift Crew and PFG
  - · Check the data quality during operation and follow up closely on any unexpected feature observed in data
  - Responsible to provide a "daily summary" on data quality to RunCoordination
    - Daily coordination meeting at 9:00 AM betwwen Offline shifters and TkDOCs
  - Documentation deeply revised in 2015 and 2016
  - · Offline Shift leaders are reporting DPG meetings and attending-/reporting to online strips meeting
  - · Ofline Shifts: A fraction of Offline shifts already assigned to qualified remote centers
  - DQM on call is the contact for any features observed by Online CMS DQM crew at P5 and ensure expertise for the tracker Offline crew
- The Offline Shift Crew take care also of the certification of Prompt-Reco data (and Reprocessing if needed)

### Offline PFG

- Plans for 2017:
  - improve collaboration/communication with online
  - make the feedback more frequent and effective
  - post slides with a set of relevant plots (to be reviewed for 2017) each 12 hours (twice as 2016)
- Improve the training of shifters:
  - introduce a tutorial for expert shifters (SL and on-call)
- make the on-call role more effective:
  - report at the DQM meeting (Friday), work more on documentation and shift training
- Shifts during MWGRs will mostly focus on setting up, testing procedures and tools
  - Tracker virtual machines, new release area for shifters, test of the scripts, revision of instructions

### Some operation items to be pursued in 2017

- Main goal: make our life easier, be faster when reaction is needed, better diagnose, reduce the load on experts
- Revive the script to send SMS to experts in case the Tracker is not giving data
- HistoricDQM:
  - has to be made faster, more "user-friendly", and more effective:
  - possibility to select the plots to be shown, improve the selection of axis limits, visualize fill bundaries, lenght of runs
- TkCommissioner: tool developed for online and has proven to be very useful for offline PFG crew too
  - Understanding feature of data taking, already in place but still needs some developments to allow prompt offline feedback
- Over the past (~10) years, a dedicated package in CMSSW is available, containing many macros, analyzers
  - Plan is to continue to document those tools (DPGAnalysis/SiStriTools package) in a more "shifter-friendly" way .
  - Give a set of "ready-to-use" cfg and examples of use cases on the basis of 2016 experience (APV shot studies, TID/TEC noise in cosmics, ...)
- Improve the tools for the monitoring of bad components and the robutness of the algorithm
- Exploit information from spy data runs

# Strip Offline Gain Calibration (also known as G2 calibration)

#### Main goal for 2017: exploit the multi run harvesting procedure at PCL

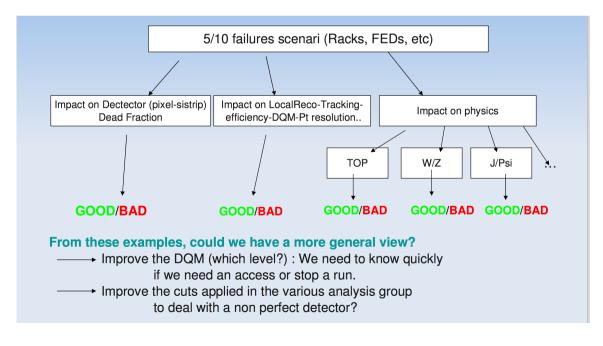
- Reduce the time delay to deliver updates;
- Reduce the workload of the experts
  - Exploit automatic machinery that doesn't need support from manual intervention;
  - Gradually involve the offline shifters for feedback after any updates
- The multirun harvesting is being validated
  - Thanks to PPD teams (in particular Piotr) for the great help and the framework development!
  - Payloads produced 'manually' have been shown recently to be the same as the ones from this new technique

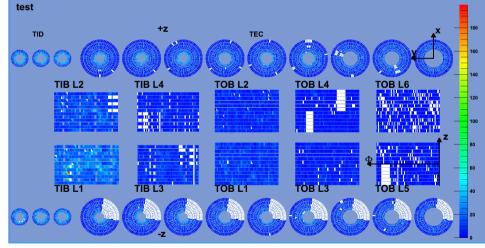
#### Simulation plans

- Introducing a realistic simulation of the APV behavior and saturation effect depending on APV settings
  - Require major rework on our simulation code
- Data MC comparisons should be pursued and simulation should be (fine) tuned
  - Example: Ensuring the out of time PU is correctly modeled with the increase of PU at LHC.

#### Failure Scenarios

- Idea: Identify some realistic failure scenarios and study their impact on reconstruction and physics analysis to ease for the future the decision on how to react in case of
- We prepared a list of them based on the operation experience over the past years and prepared the corresponding payloads





Example: Power Supply Rack Failure: Missing regions in TEC- D1->9 + TOB L4-£5

#### Longer term plan

- Strips DAQ and DPG took note of the request from the HI community to increase the L1 rate for the 2018 running
- In HI running there is an efficiency loss due to APV baseline deformations which cause problems in Zero Suppression
  - (N.B. this effect is still present also with lower VFP)
  - Needs change in FED firmware to have more elaborate ZS algorithm (akin to baseline follower which has been used in HLT during HI runs for ZS)
  - Will start discussion with HI community and tracker firmware developers
- Even if the target is 2018 HI data taking, this should be adressed already in 2017

#### Outlook

- Strips Online & offline Group are actively preparing the data taking
  - To be ready for CMS physics measurements and discoveries to come!
- Goals:
  - Provide the best calibrations of the detector for high quality physics data
  - Provide the best possible data/MC agreement
  - Reduce the load in 'every day ' operation
    - Simplify worklows, procedures
    - Allow more time to anticipate, avoid, or eventually fix any bad surprises
  - Be prepared for an 'unlikely case of emergency ' (\*)
    - · Failure scenario, tools for diagnose and feedback, efficient shift crew
  - Enjoy running the Strip detector, calibrating it, and eventually presenting results of the performance of it

## Backup

## Remember, CRAFT Workshop 2009 Torino

