

Light Dark Matter @ Accelerators
 24-28 May 2017, La Biodola – Isola d'Elba

17

GUT

SUSY

axions

WIMPs

SIMP's

ELDERs

MACHOs

sexaquarks

self-interacting
DM

Andrea Pocar
 University of Massachusetts, Amherst



AMHERST CENTER FOR FUNDAMENTAL INTERACTIONS
 Physics at the interface: Energy, Intensity, and Cosmic frontiers
 University of Massachusetts Amherst

Liquid xenon (LXe) and liquid argon (LAr)

for low-background physics

Noble liquid detectors have risen to be a leading technology in low-energy rare event searches over the past ~decade
(WIMP dark matter, $0\nu\beta\beta$ decay)

Scintillating calorimeters

- XMASS (LXe)
- DEAP-3600 (LAr)
- miniCLEAN

all growing larger

TPCs

- ZEPLIN
- XENON 10/100/1T/nT
- LUX / LZ (LXe)
- PandaX I-II
- WaRP (LAr)
- ArDM
- DarkSide-50 / DS-20k

- EXO-200
nEXO (LXe)

General features, requirements

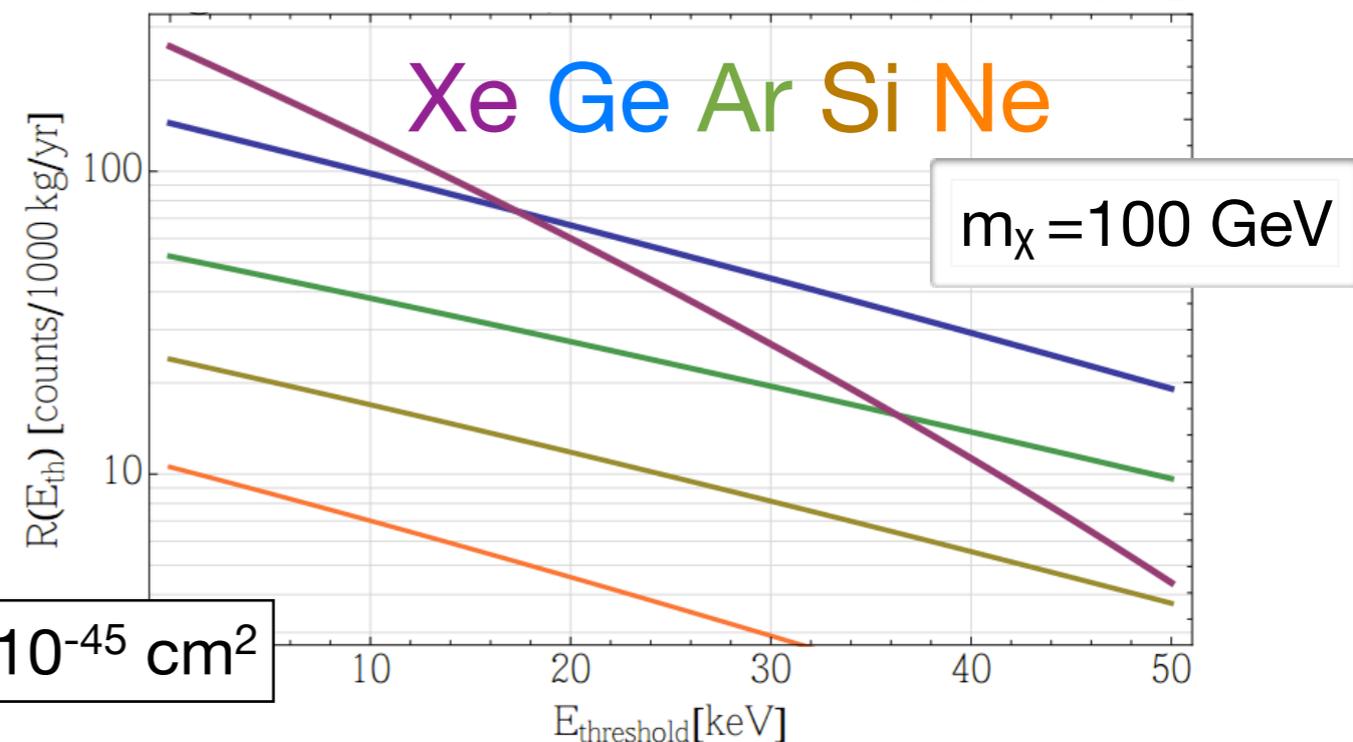
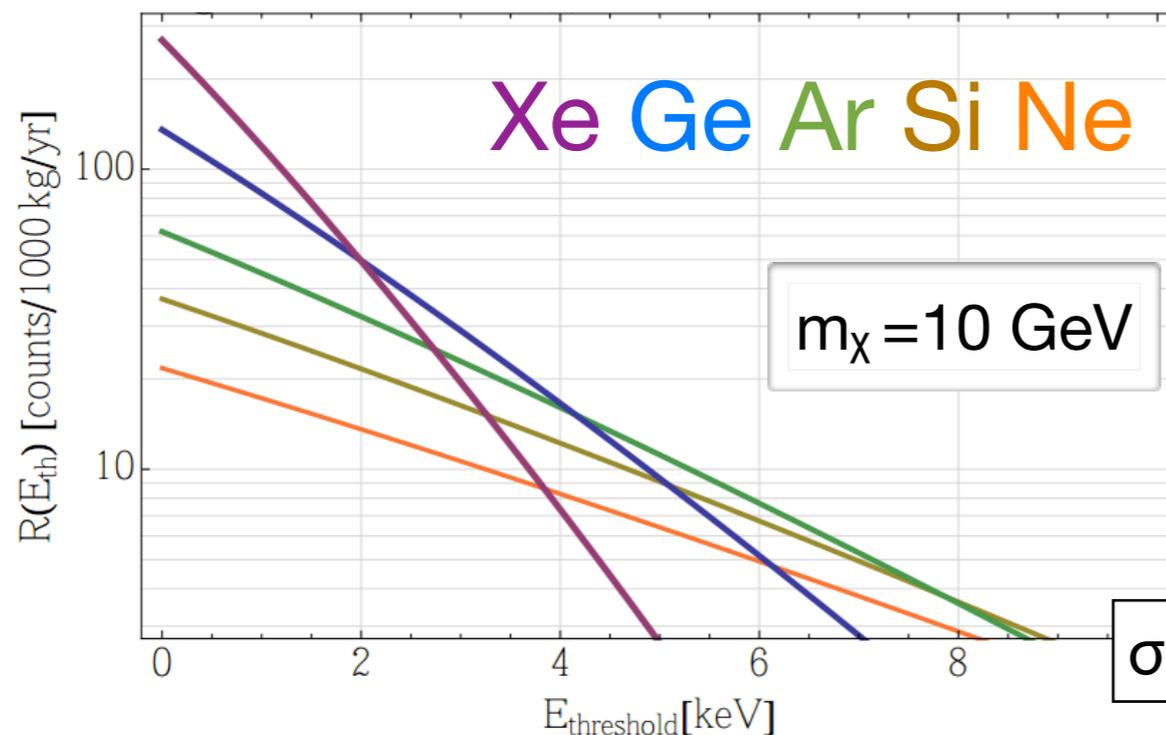
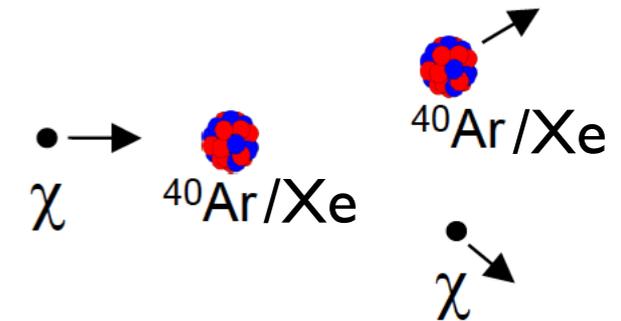
Strengths:

- Powerful combination of low energy threshold, energy resolution, event ID
- LXe/LAr are excellent scintillators
- Background discrimination (technique-specific)
- Scalability
- Inline purification

- Large ‘empty’ volume filled with clean, re-purifiable LXe/LAr
- ‘Dirty’ components pushed away from fiducial volume (self-shielding)
- Selected materials, reduce mass where possible
- Apply detector design concepts developed for low energy neutrinos to searches for WIMPs

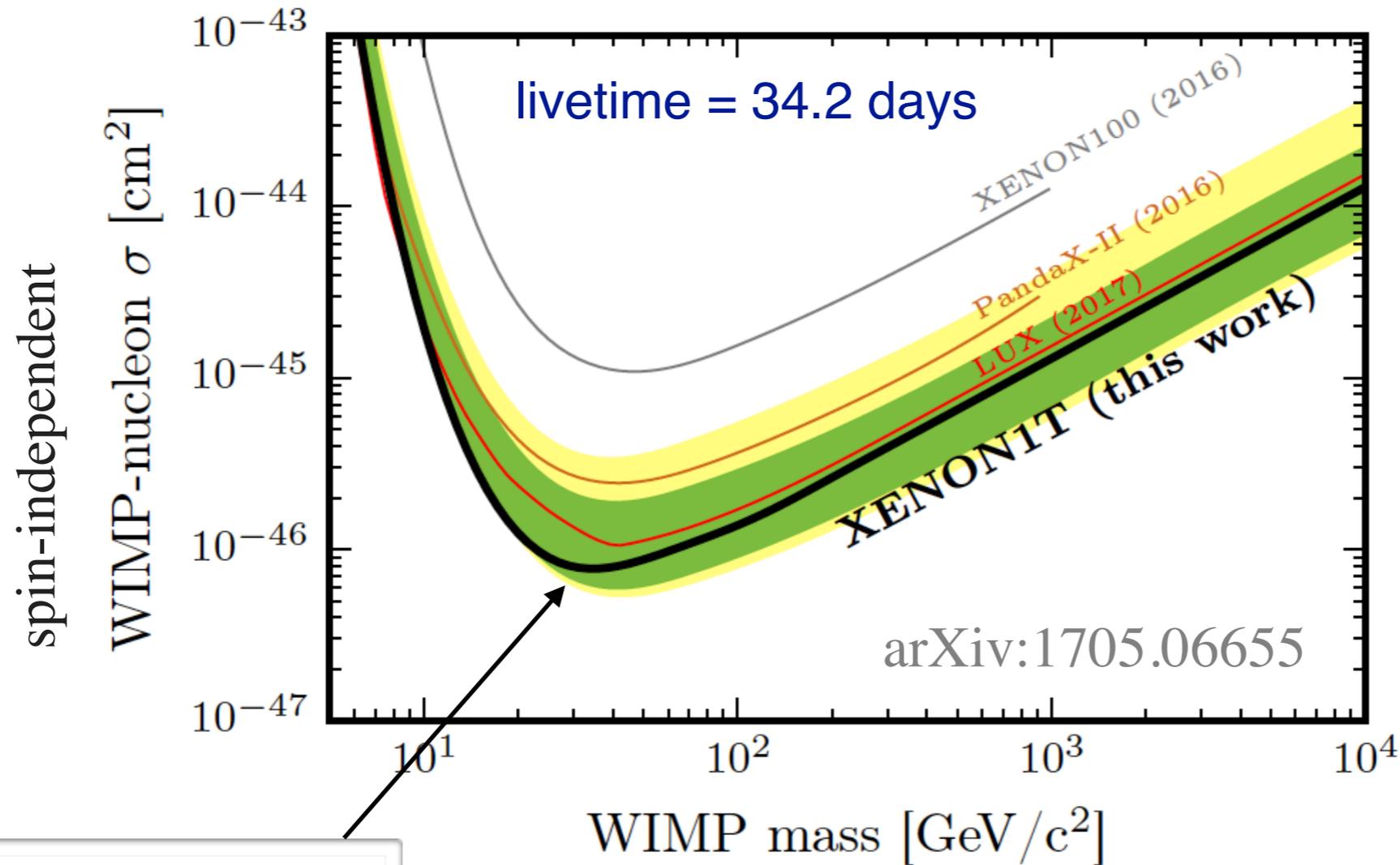
Direct dark matter search: what we measure

→ look for excess of nuclear-like recoil events



Noble liquids are very powerful WIMP-searching machines with mass larger than a few GeV/c^2

Fresh from the press: new results from XENON1T



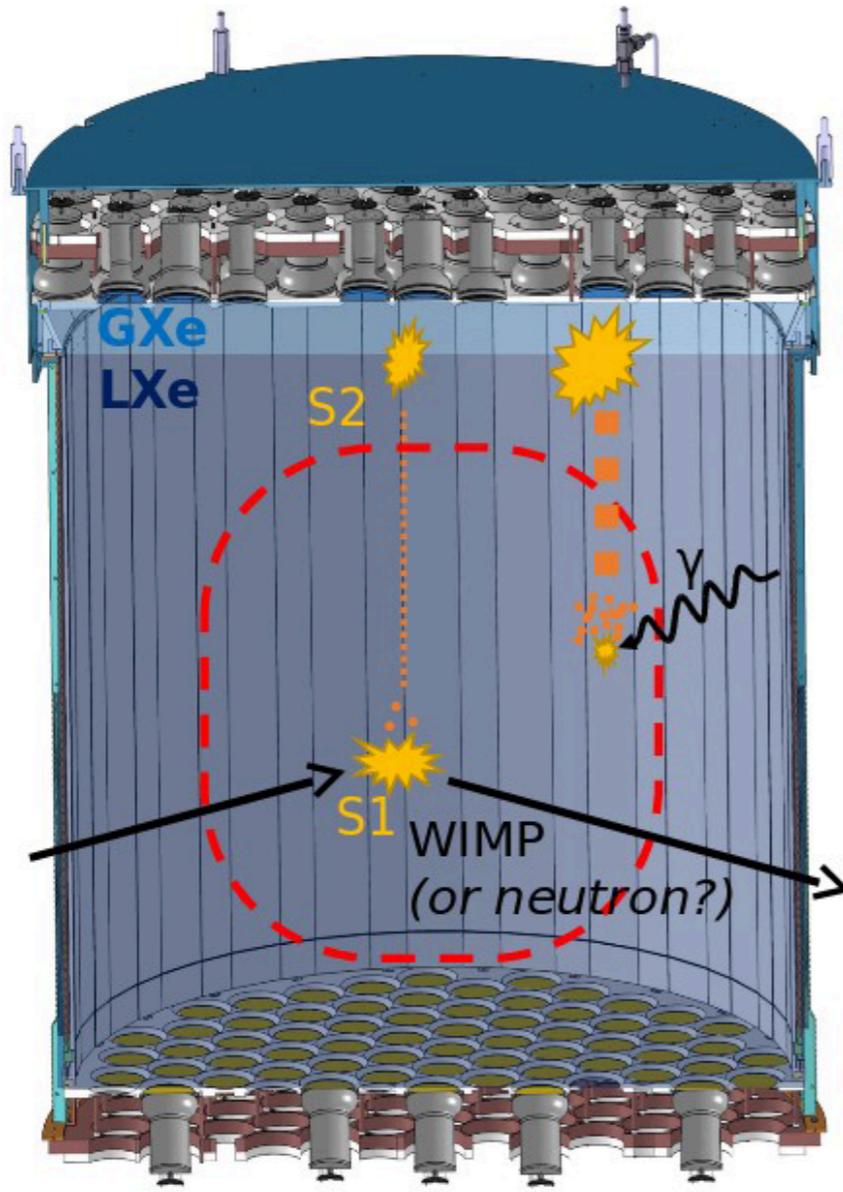
fiducial mass:
 1042 ± 12 kg

$7.7 \times 10^{-47} \text{ cm}^2$

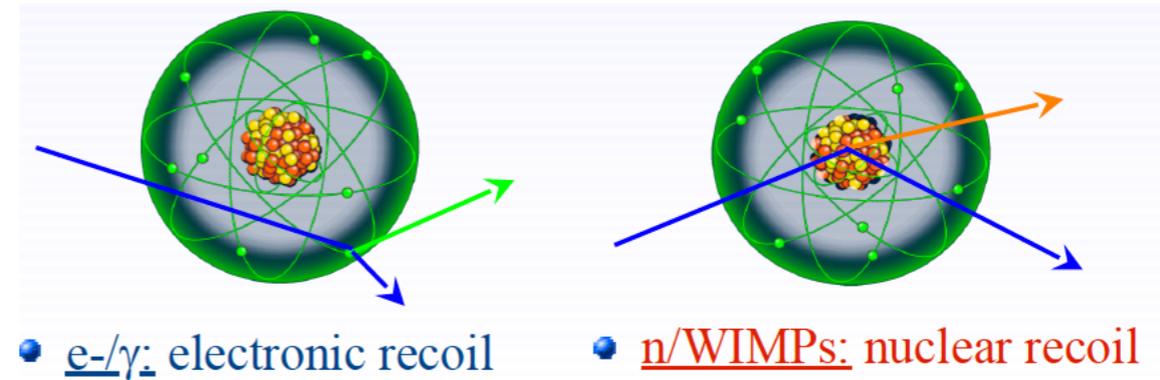
dual phase LXe: current leaders for WIMP searches

Dual-phase LXe TPC

adapted from
Gianmarco Bruno and Ethan Brown



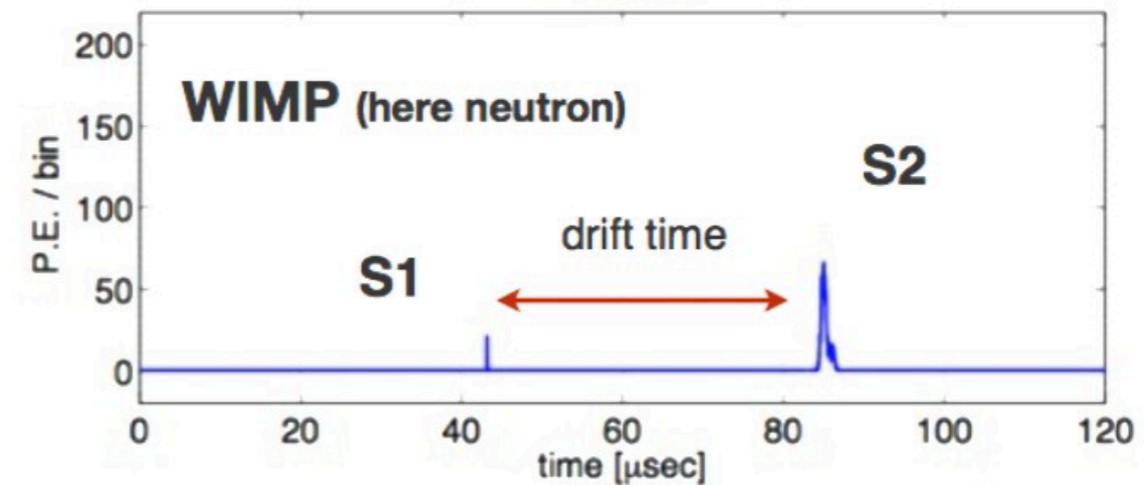
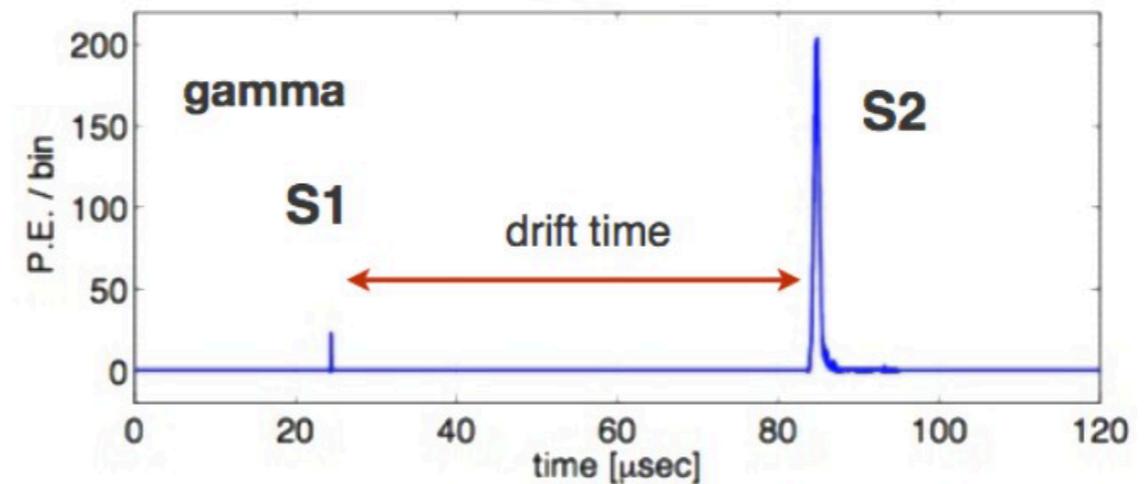
- S2/S1 ratio provides particle ID
- Top-array hit pattern provides the vertex position in the X-Y plane
- The time difference provides the depth



• e^-/γ : electronic recoil

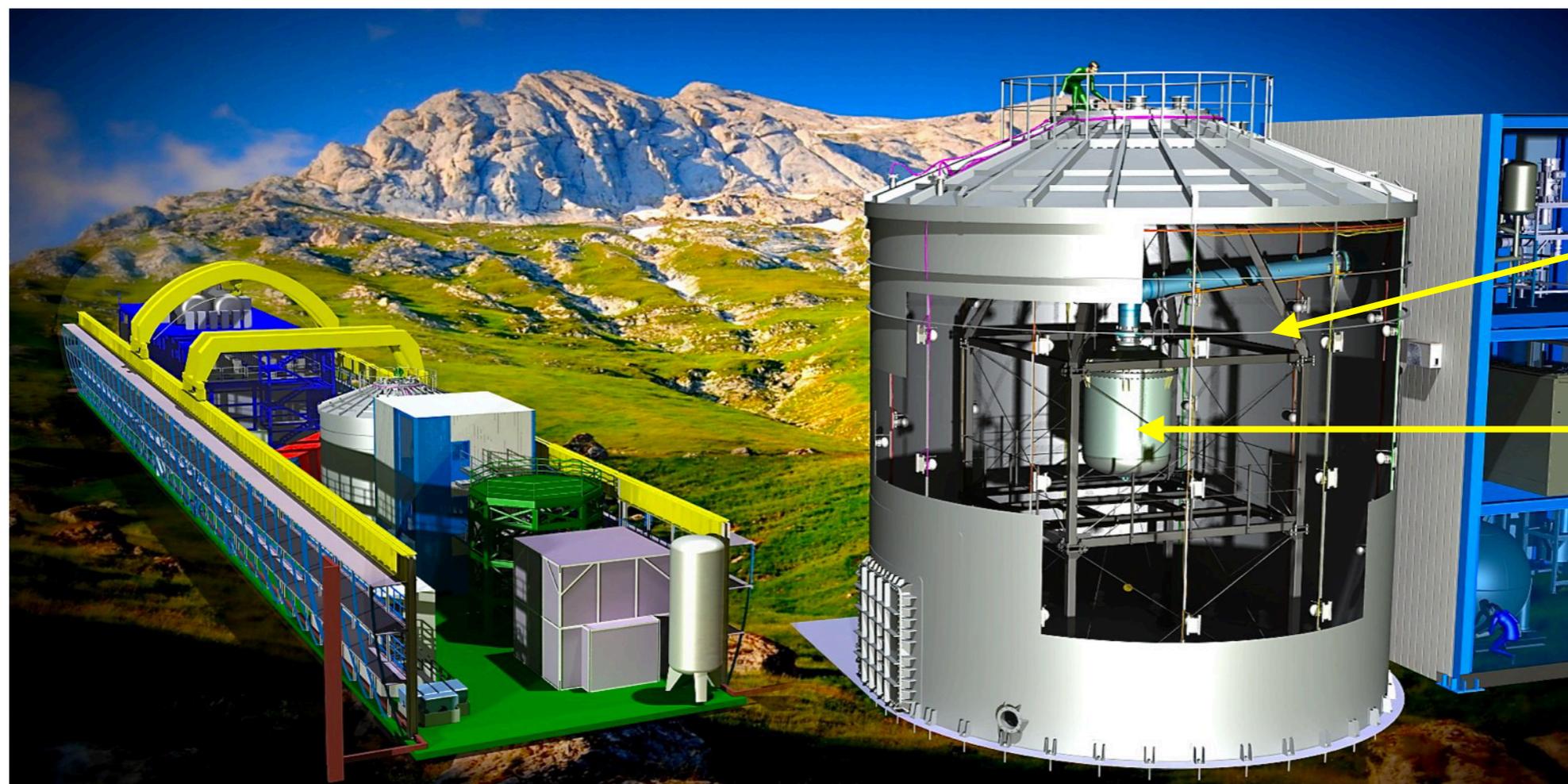
• $n/WIMPs$: nuclear recoil

$$(S2/S1)_{NR} < (S2/S1)_{ER}$$



XENON1T at LNGS

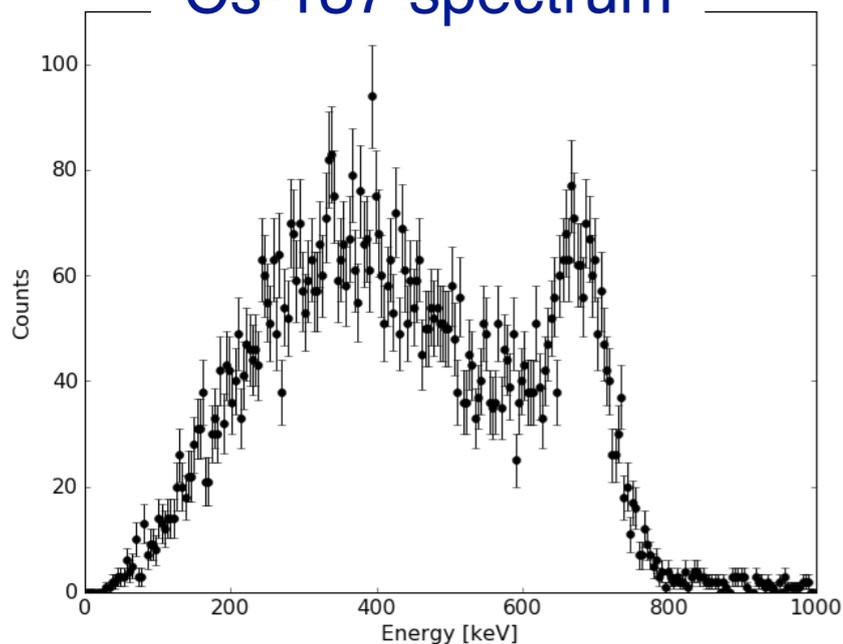
adapted from
Gianmarco Bruno



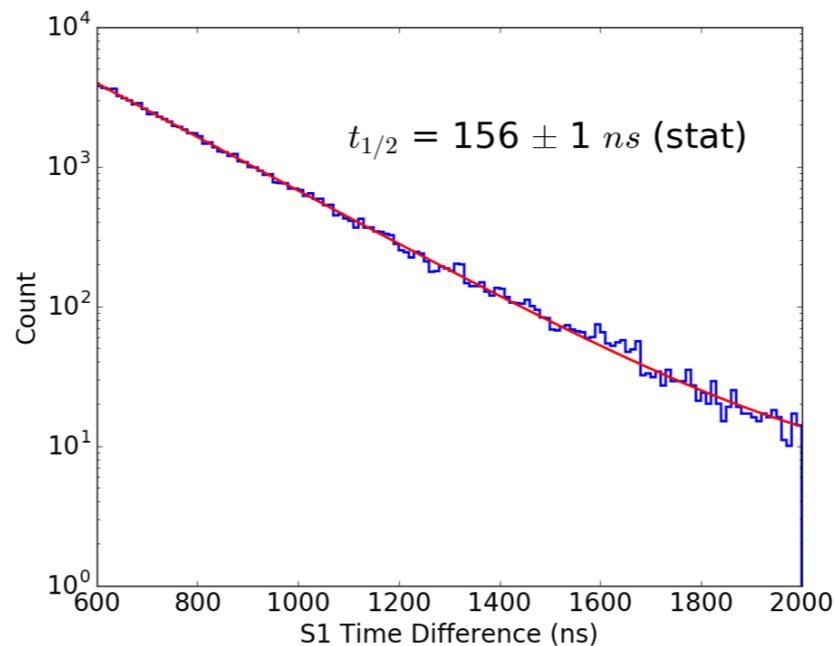
active water shield

TPC

Cs-137 spectrum



XENON1T: Kr83m 7/2+ Halflife



LY is x2 higher
than XENON100

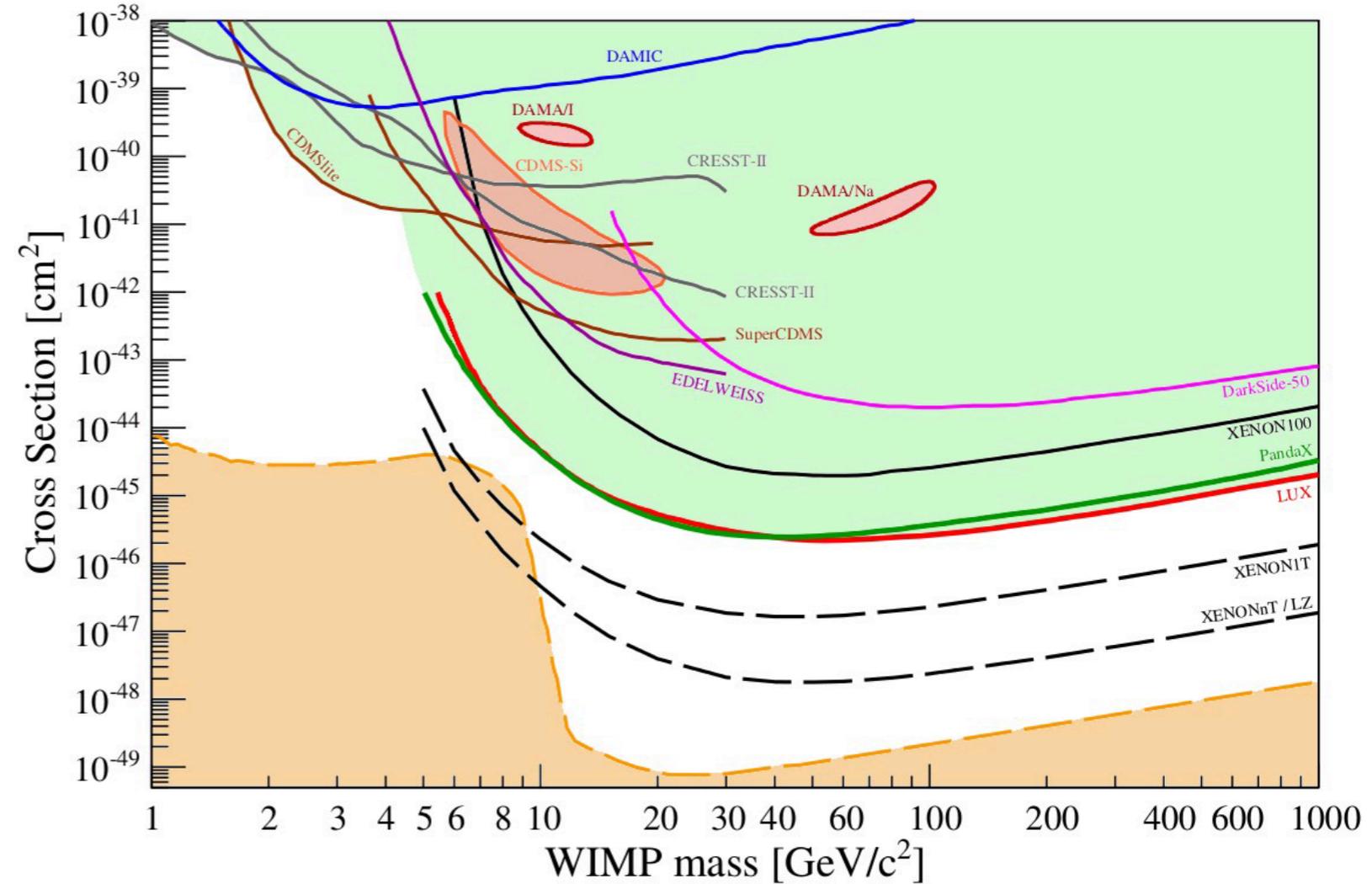
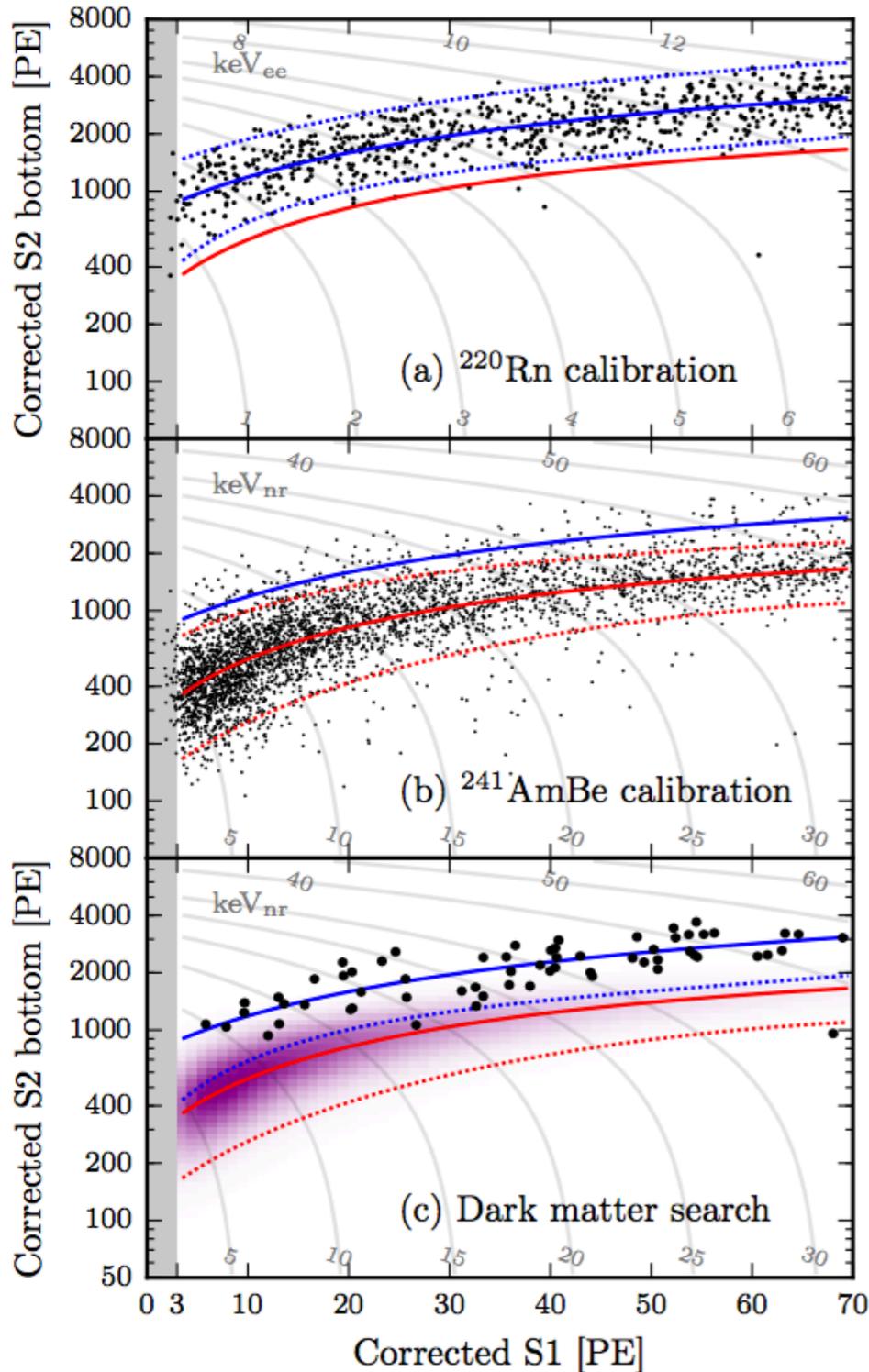
$$(1.93 \pm 0.25) \times 10^{-4} \text{ events/kg/d/keV}_{ee}$$

XENON1T \rightarrow nT sensitivity projections

adapted from
Gianmarco Bruno
and Marc Schumann



Updated from: arXiv:1512.07501 JCAP04(2016)027



- 2 t·y exposure
- energy interval: 4-50 keV
- ER rejection: 99.5%
- NR acceptance: 50%
- 0.2 ppt $^{\text{nat}}\text{Kr}$, 10 $\mu\text{Bq/kg}$ ^{222}Rn

The XENON detector genealogy

adapted from G. Bruno

(2007)



XENON10
Total Xe: 25 kg
Target: 14 kg
Fiducial: 5.4 kg
Limit: $\sim 10^{-43}$

(2015)



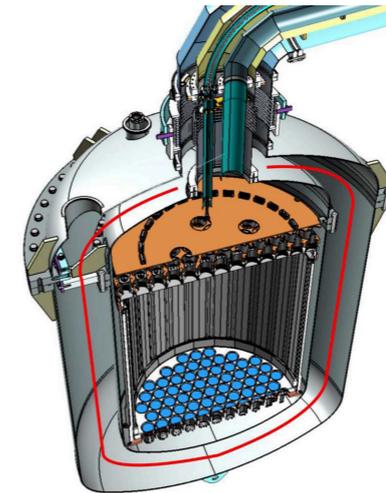
XENON100
Total Xe: 162 kg
Target: 62 kg
Fiducial: 34/48 kg
Limit: $\sim 10^{-45}$

(2019)



XENON1T
Total Xe: 3.5 ton
Target: 2 ton
Fiducial: 1 ton
Limit: $\sim 10^{-47}$

(2022?)



XENONnT
Total Xe: 7.5 ton
Target: 6 ton
Fiducial: 4.5 ton
Limit: $\sim 10^{-48}$

(2028?)



DARWIN
Total Xe: 50 ton
Target: 40 ton
Fiducial: 30 ton
Limit: $\sim 10^{-49}$

- Outer vessel sized for upgraded detector
- 200 more PMTs are needed
- Total Xe mass: 7.5 t
- Rn removal system
- Projected to start science in 2019

neutrino
“floor”

PandaX program

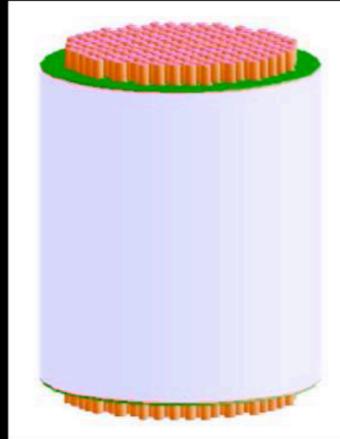
adapted from Jianglai Liu, Pheno 2017



PandaX-I: 120 kg
DM experiment
2009-2014



PandaX-II: 500 kg
DM experiment
2014-2018

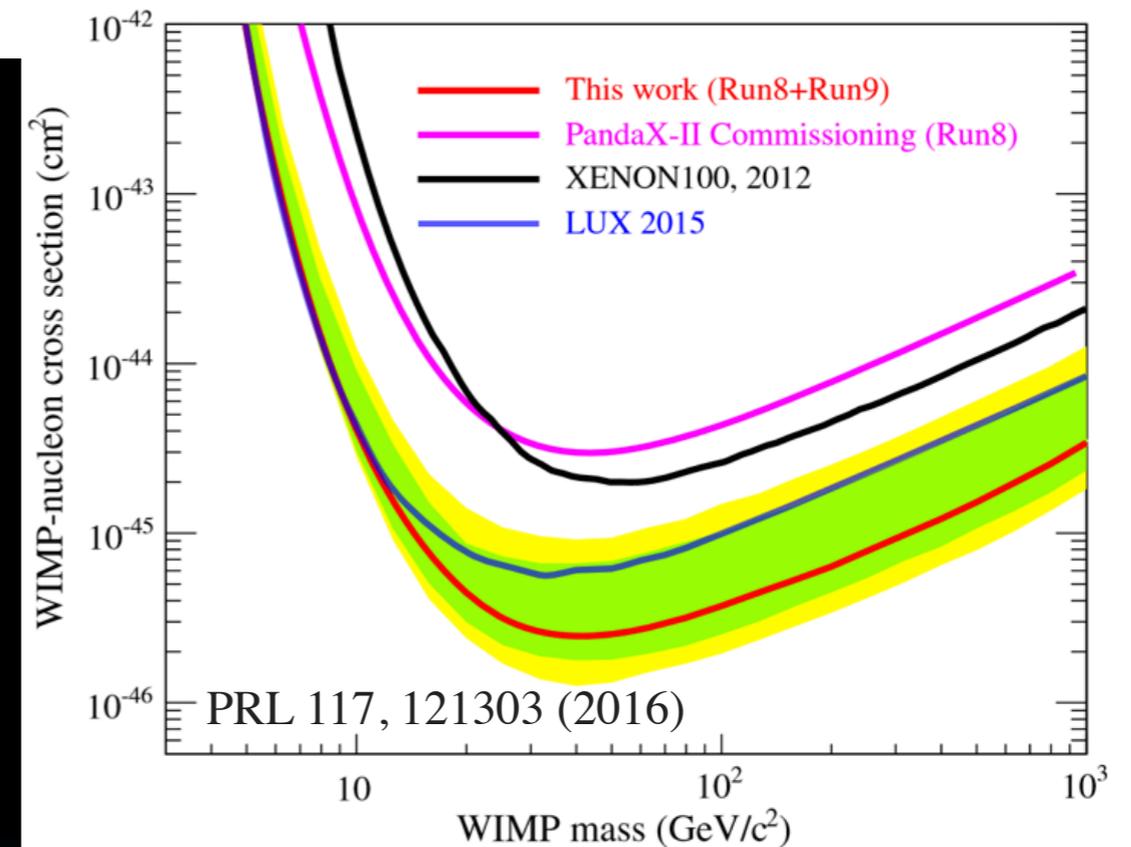


PandaX-xT:
multi-ton (~4-T)
DM experiment
Future

CJPL-I

CJPL-II

 **PANDA X** = Particle and Astrophysical Xenon Experiments



33 tonne-day
exposure



- PandaX-II running after extended downtime
- Preparing hall at CJPL-2 for PandaX-4T with water tank (~ 10^{-47} cm², commissioning ca. 2020)
- Goal: 30T experiment → neutrino floor

LUX \rightarrow LZ (LUX + Zeplin) program at Sanford Lab

adapted from Attila Dobi



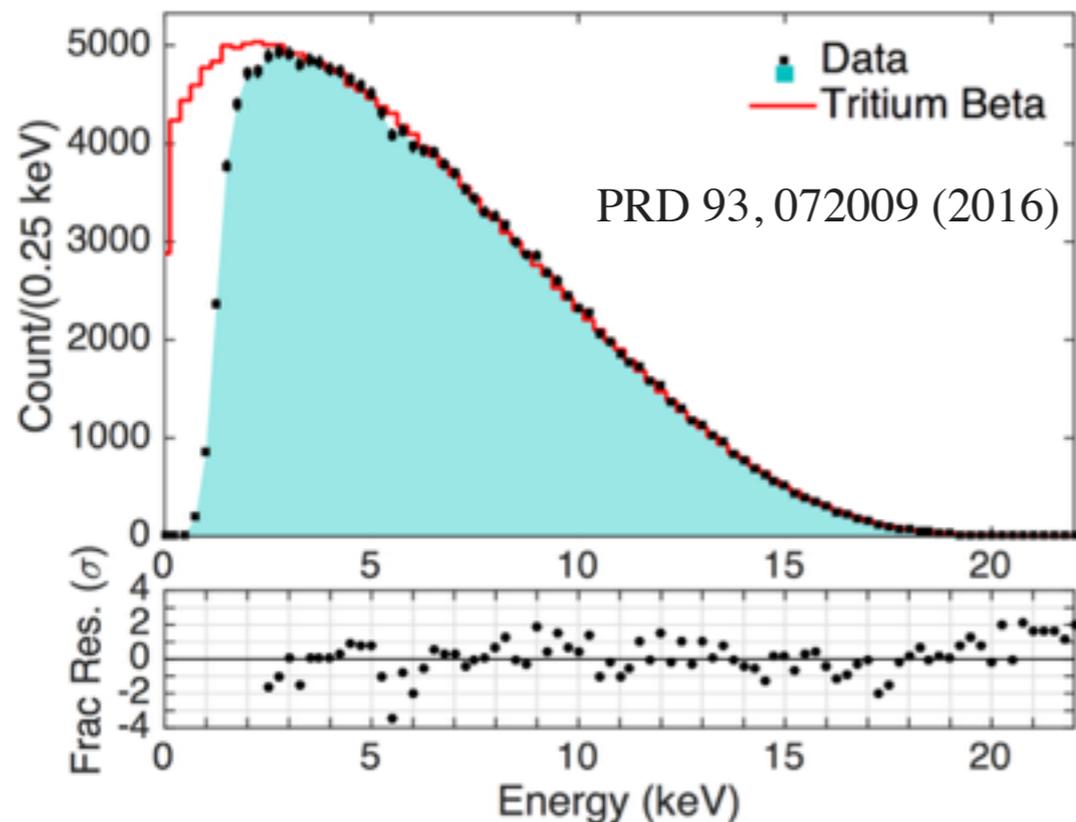
- LUX decommissioned in 2016 (~400 live days, 3.4×10^4 kg d exposure)
- 2nd best SI limits

PRL 118, 021303 (2016)

- innovative calibrations and Kr-85 bg reduction and control

- LUX \rightarrow LZ

Tritium Beta Spectrum Measured in LUX

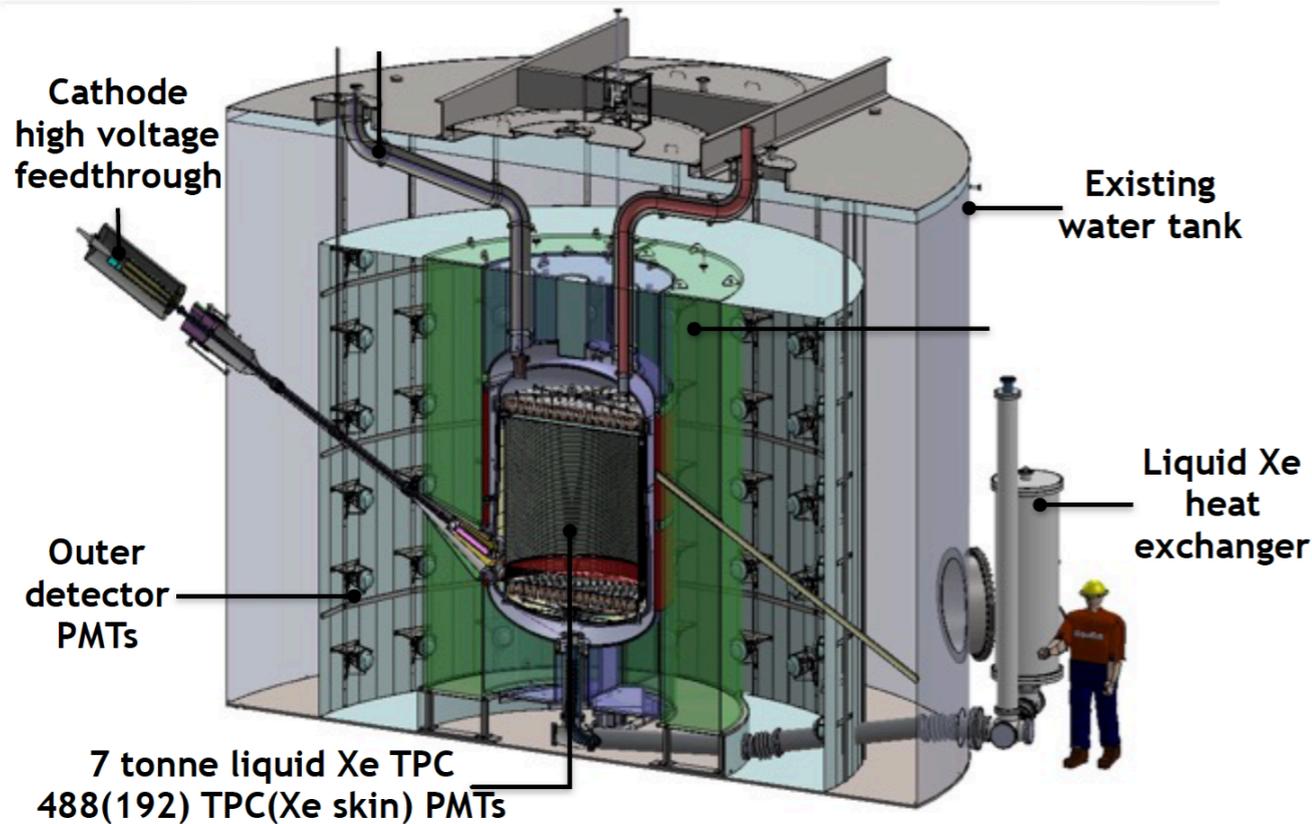


Kr-85: best LUX batch 200 ppq

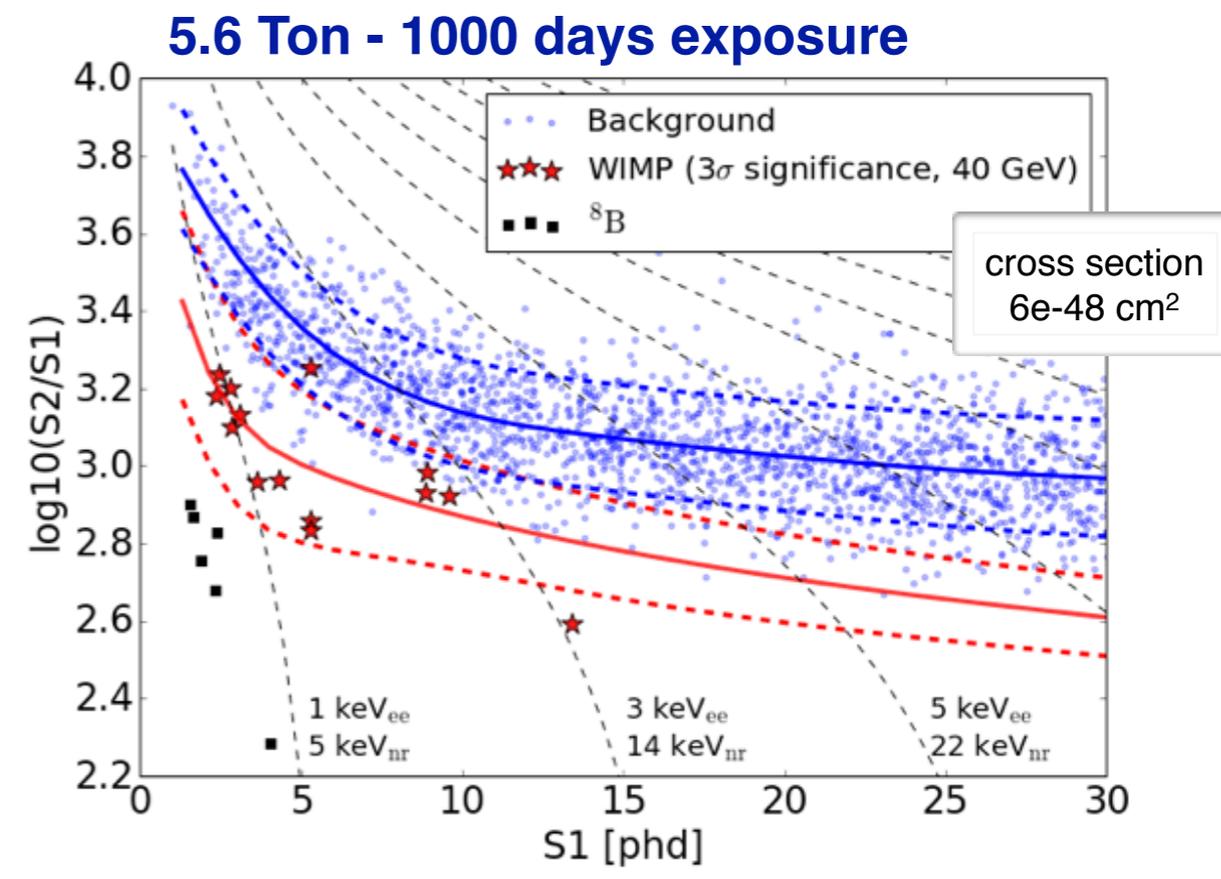
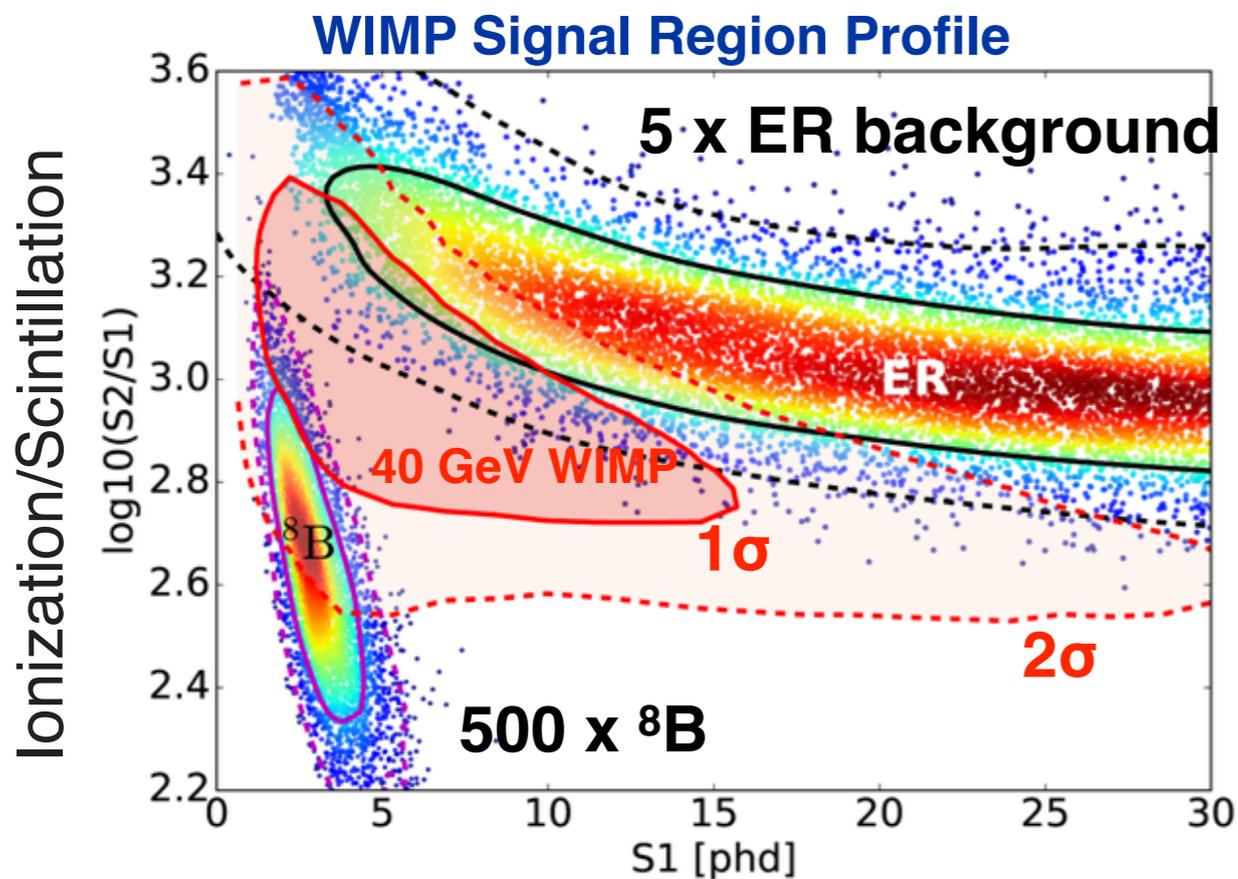


LZ at Sanford Lab

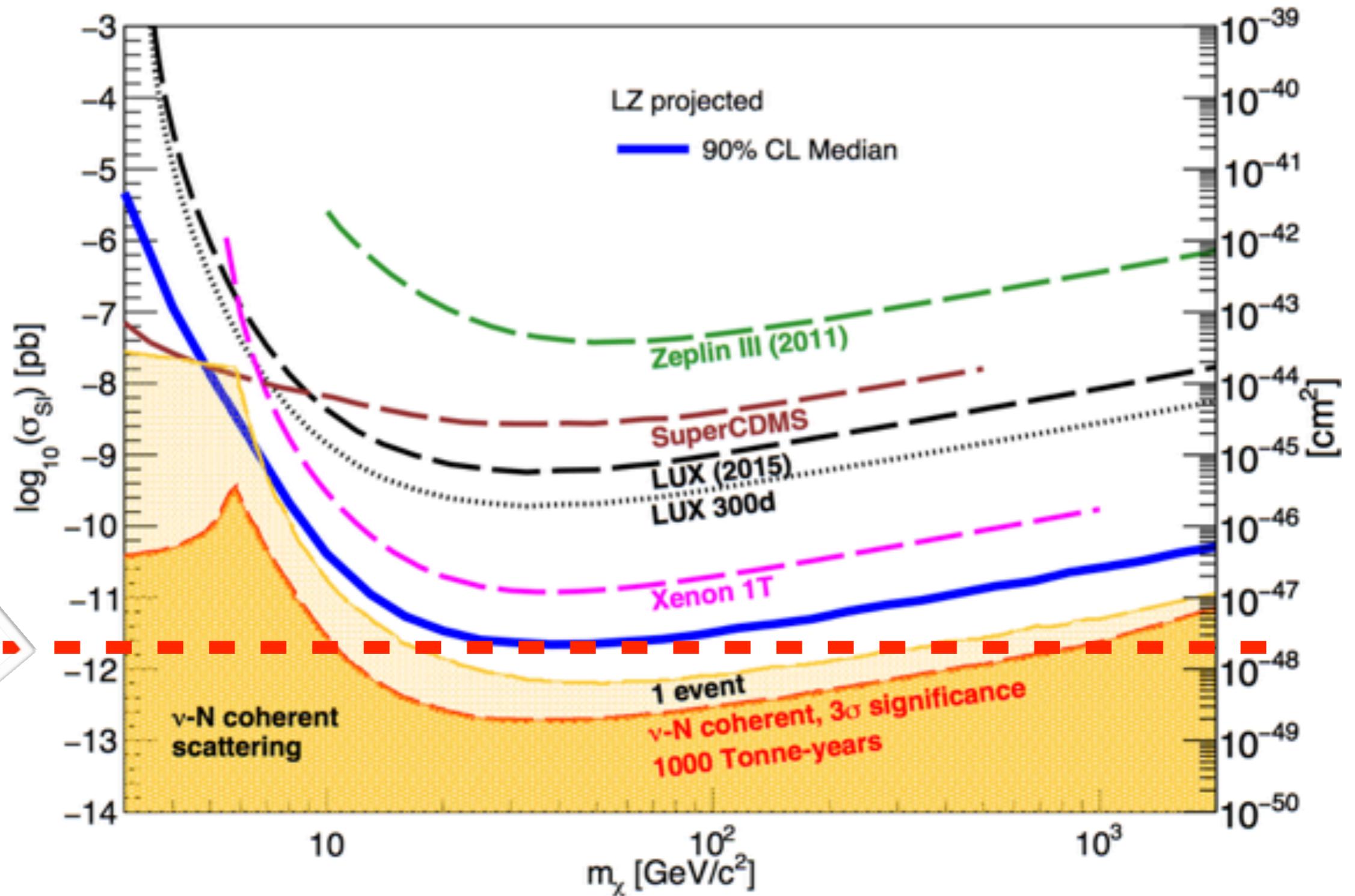
adapted from Attila Dobi



- 10 tons LXe / 7 active / 5.6 fiducial
- Triple veto system
- Gd-loaded scintillator n veto



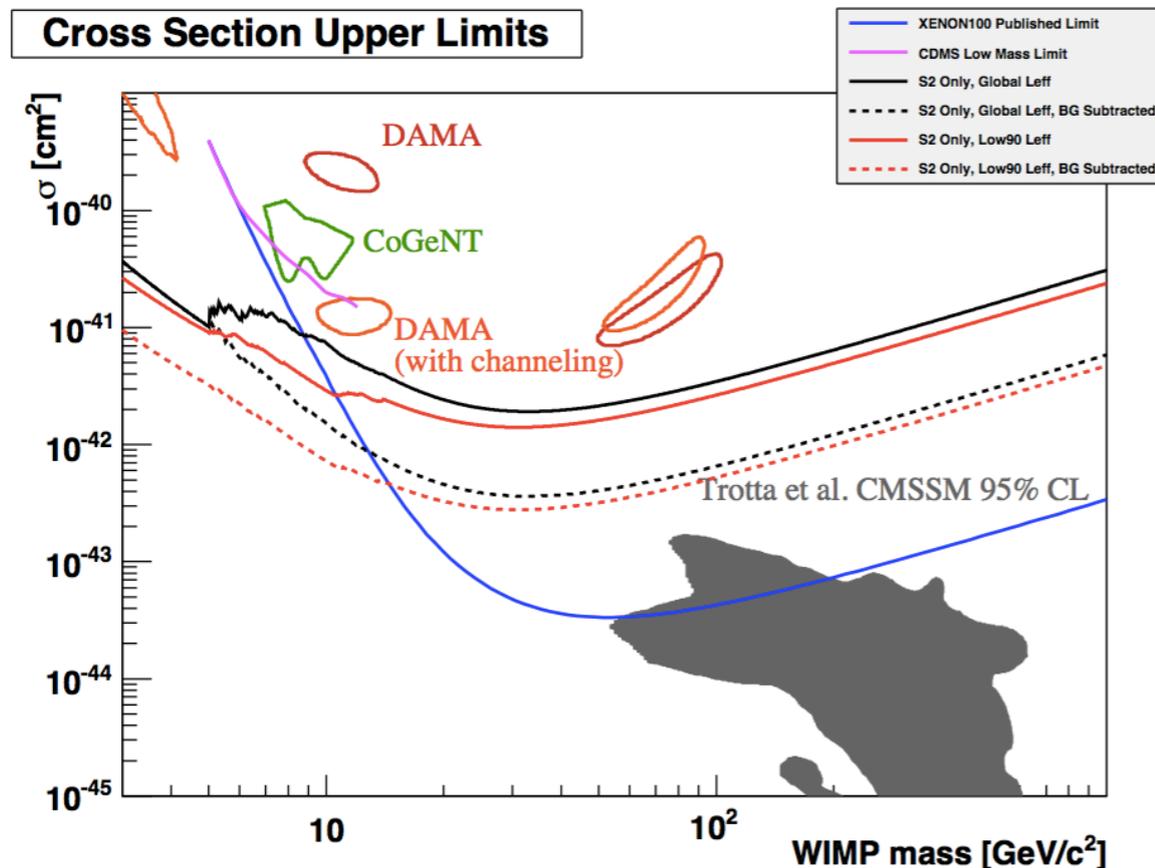
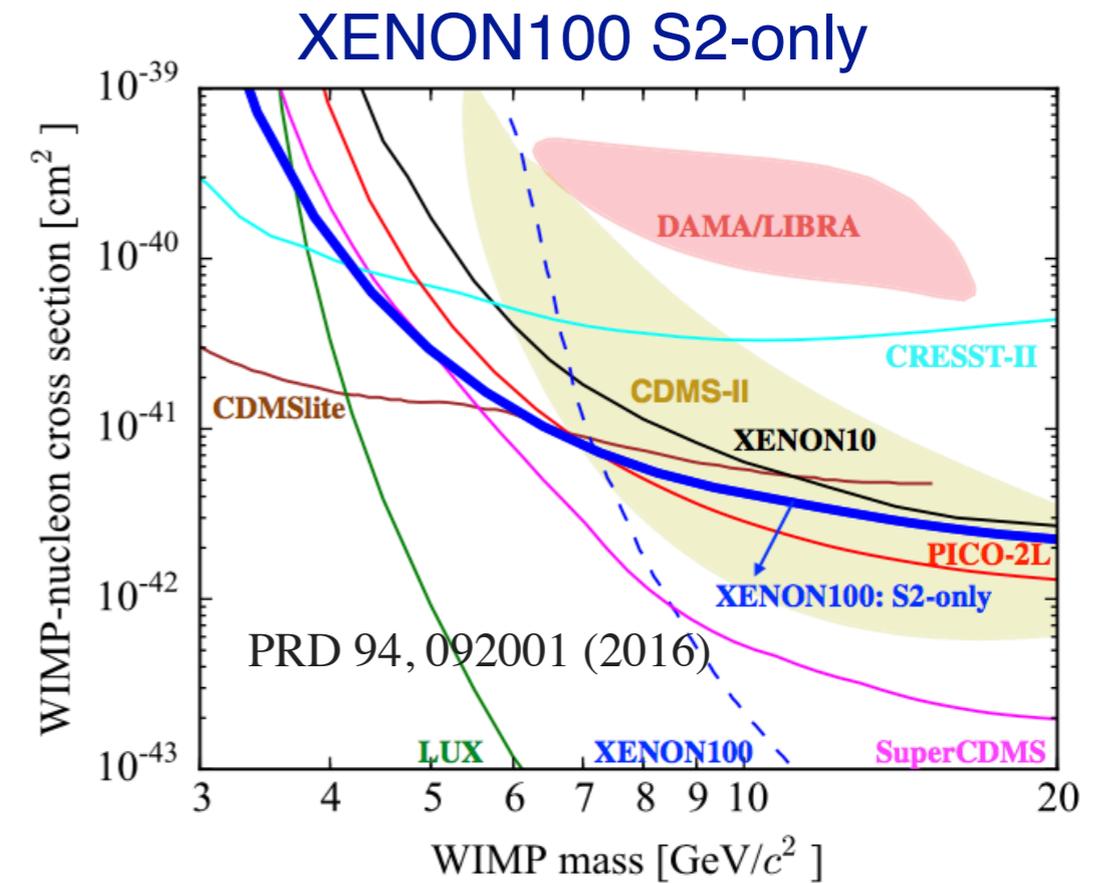
LZ projected sensitivity (SI, 5.6 Tonnes, 1000 live days)



Low-mass WIMPs

adapted from
Maria Elena Monzani

	S1 + S2	S2 - only
Nuclear Recoils	<ul style="list-style-type: none"> • Vanilla WIMPs 	<ul style="list-style-type: none"> • Light(er) WIMPs • Asymmetric Dark Matter
Electron Recoils	<ul style="list-style-type: none"> • \simkeV axion-like particles 	<ul style="list-style-type: none"> • subGeV hidden sector models



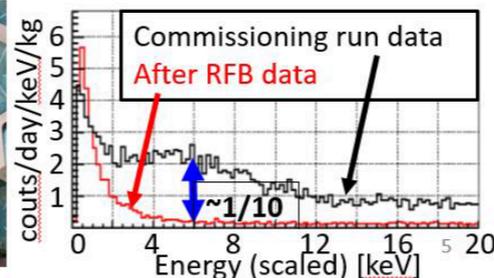
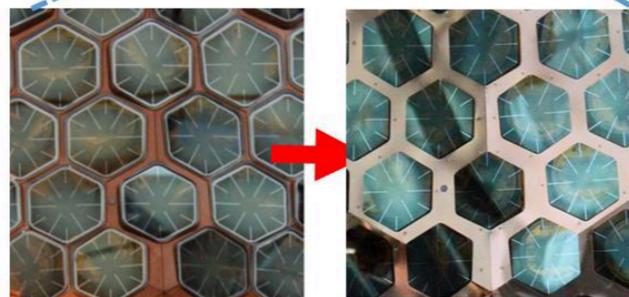
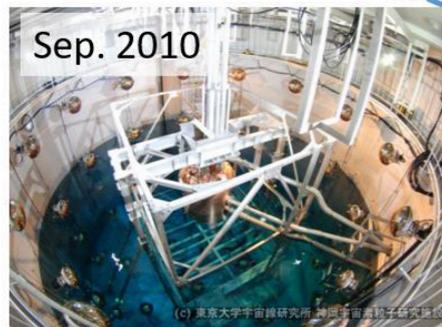
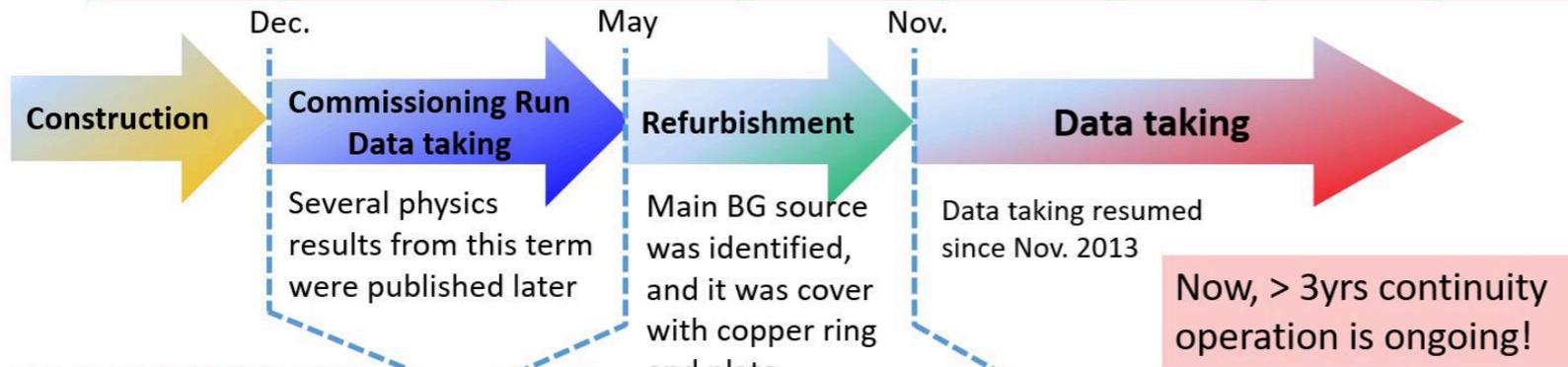
- no z-coordinate
- limited by rejection
- very sensitive to impurities
- extraction efficiency
- single electron background hard to model

Single phase LXe: XMASS at Kamioka

adapted from A. Takeda
XeSAT 2017



- 835 kg of LXe, 630 PMTs
- relative simplicity
- limited discrimination
- Key requirements:
 - low external background
 - low radon
 - fiducialization



- issues with PMT bg
- early attempt at PSD in small-scale setup
- annual modulation study

LXe \rightarrow LAr

- complementary searches at higher masses
 - different systematic effects
 - heading for the neutrino floor
- \rightarrow Argon is inexpensive and relatively “easy” to purify to levels required for DM searches - true for O₂, N₂, etc. and **also for radon**
 - \rightarrow Singlet/triplet ratio and lifetimes in argon allow extremely good scintillation PSD (β/γ vs nuclear recoil rejection of 10¹⁰) – low background single phase (scintillation-only) detector possible
 - \rightarrow TPC also exploiting charge collection (S1/S2) increases background rejection (β/γ vs recoil + position reconstruction)
 - \rightarrow ³⁹Ar – approx. 1 Bq of β decays per kg of argon – must be reduced or rejected in analysis

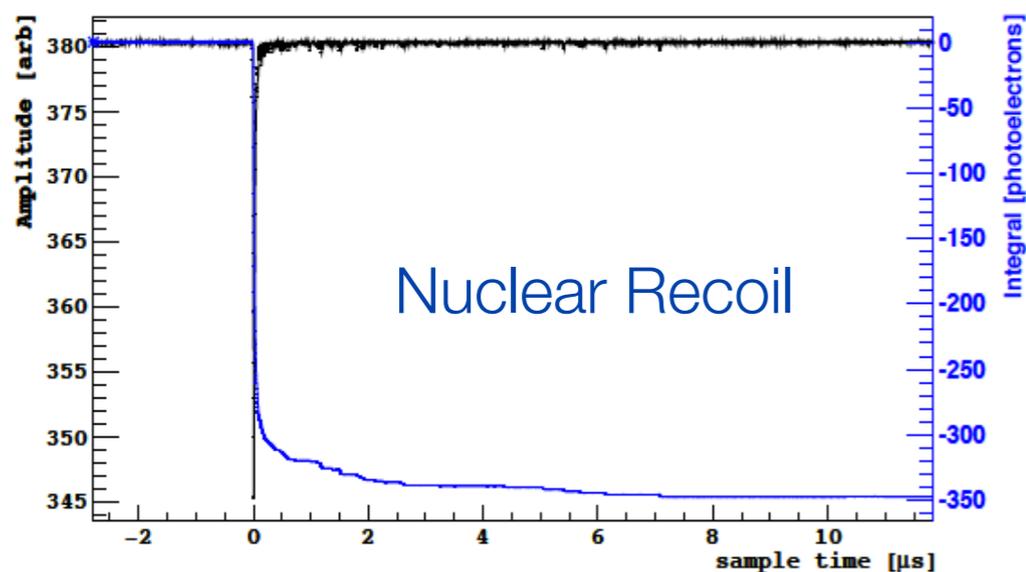
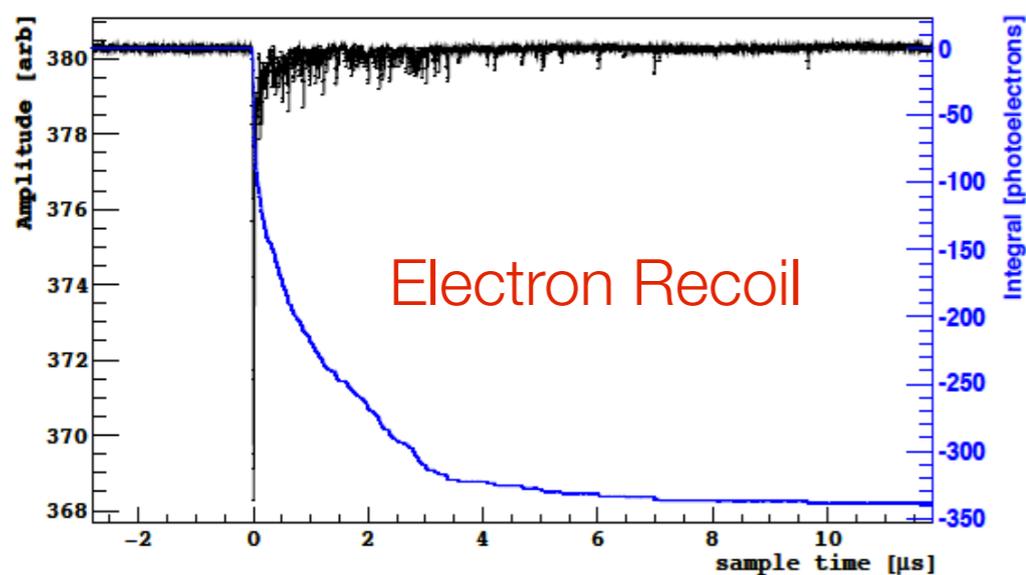
Pulse Shape Discrimination (S1) in LAr

Boulay and Hime, Astropart. Phys. **25**, 179 (2006)

Electron and nuclear recoils produce different excitation densities in the argon, leading to different **ratios of singlet and triplet excitation states**

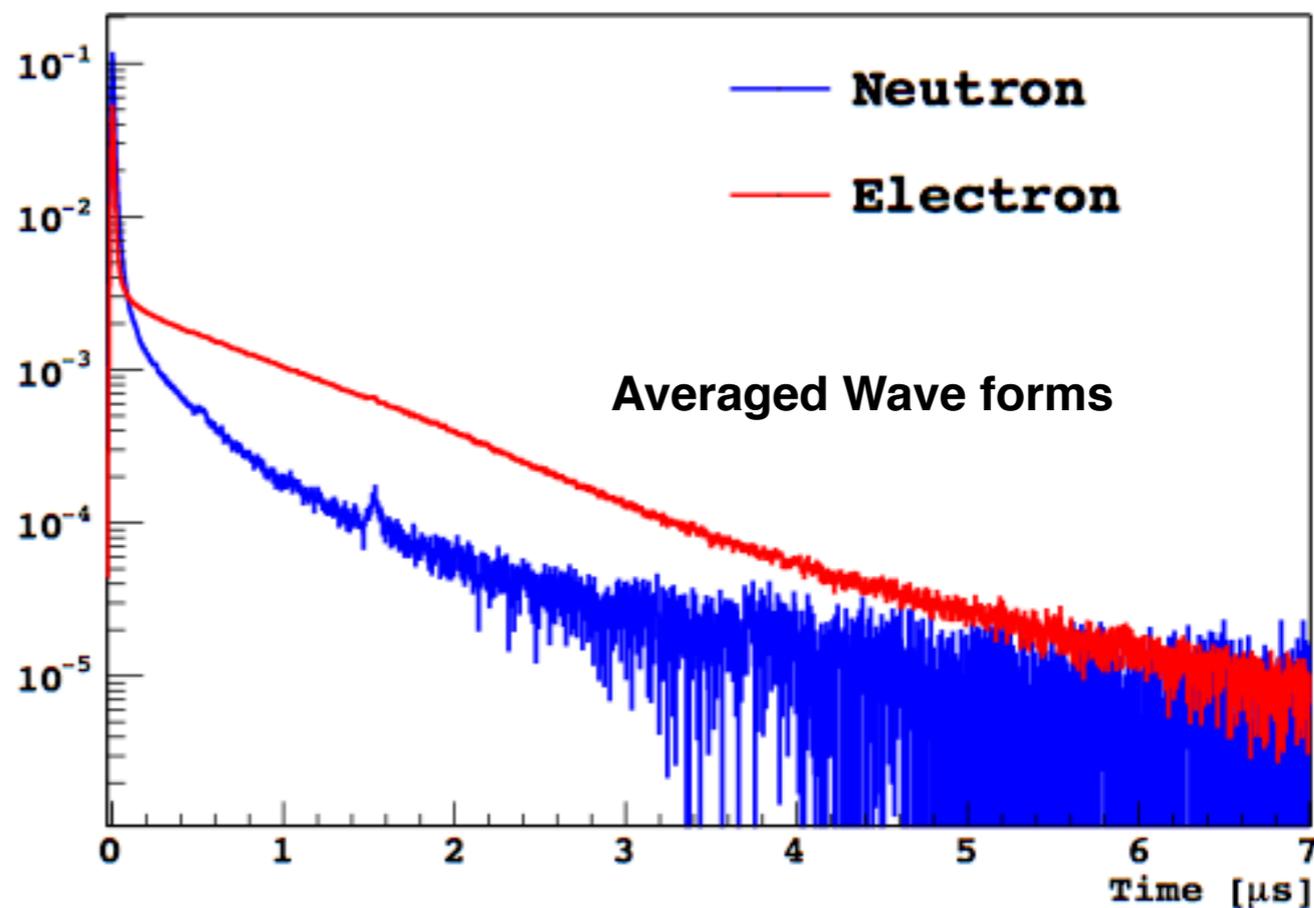
$\tau_{\text{singlet}} \sim 7 \text{ ns}$

$\tau_{\text{triplet}} \sim 1500 \text{ ns}$



PSD parameter

F90: Ratio of detected light in the first 90 ns, compared to the total signal
~ Fraction of singlet states



World-wide LAr collaboration

Researchers from

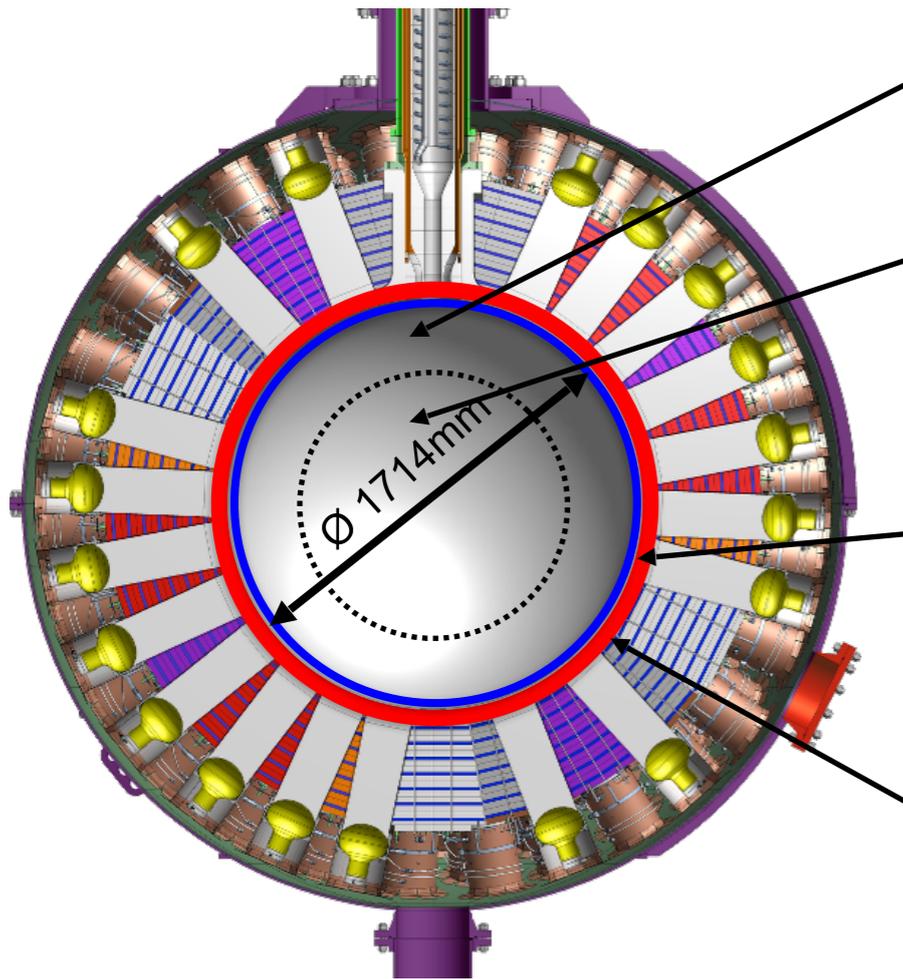


planning to collaborate on future program:

- Completion of current science and R&D programs by each collaboration (DS-50, DEAP-3600, MiniCLEAN, ArDM)
- Joint collaboration on DS-20K at LNGS, including Low Radioactivity Argon (operation starting 2021) and SiPM photodetectors
- Joint collaboration on future multi-hundred-tonne LAr detector, site TBD (mid-2020's)

Single phase LAr: DEAP-3600

adapted from Fabrice Retière

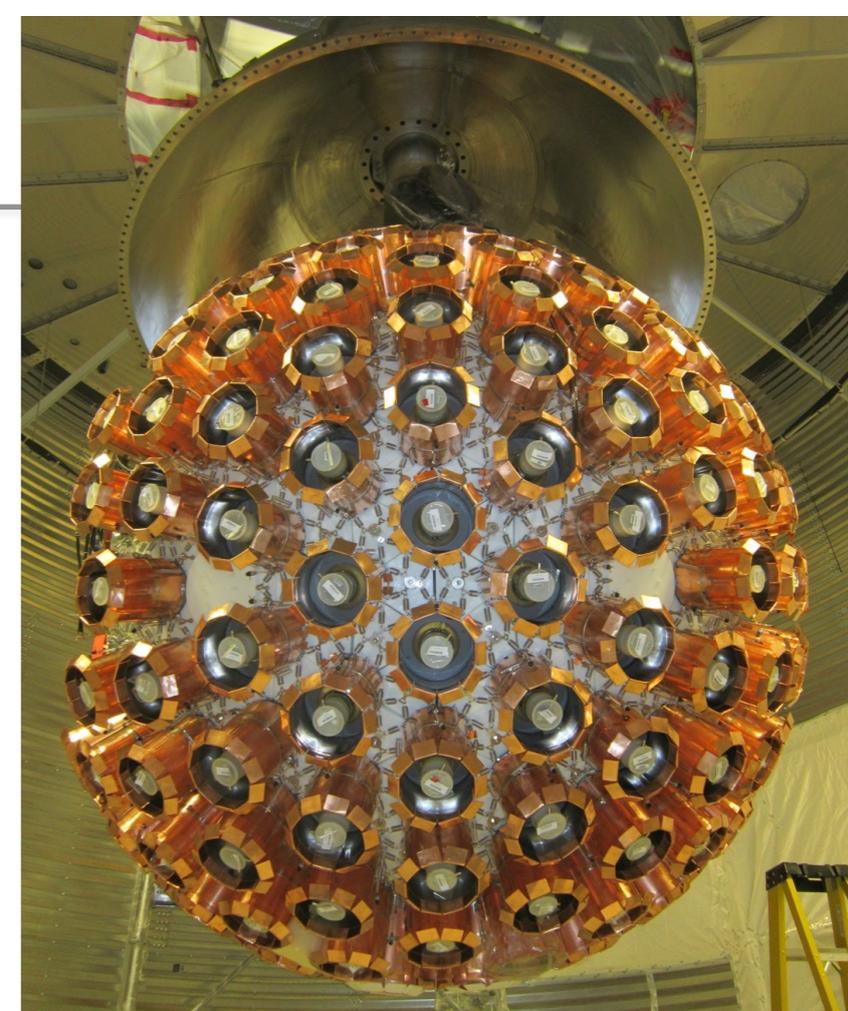


3600 kg of Liquid Argon

1000 kg Fiducial mass

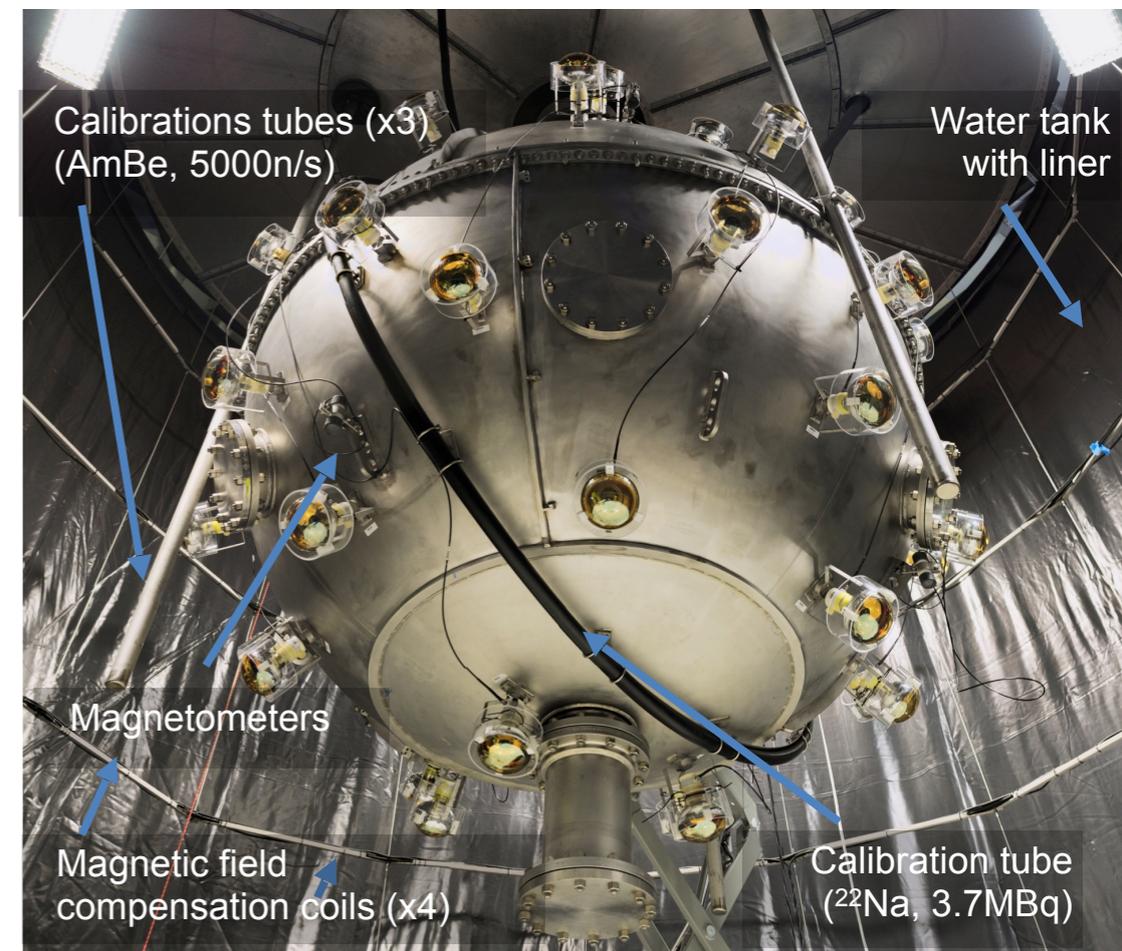
Wavelength shifter
(distilled TPB)

Vessel



WIMP search

Distinguish nuclear from electron recoils by Pulse Shape Discrimination (PSD) of the scintillation signal
(very powerful in Argon)





- Room-temperature PMTs
- Ultra-low bg acrylic
- Exposure in line with XENON1T
- Very high statistics PSD

Background	Fiducial No. Events in Energy ROI – 3 live years
Neutrons	<0.2
Surface α 's	<0.2
^{39}Ar β 's (natural argon)	<0.2

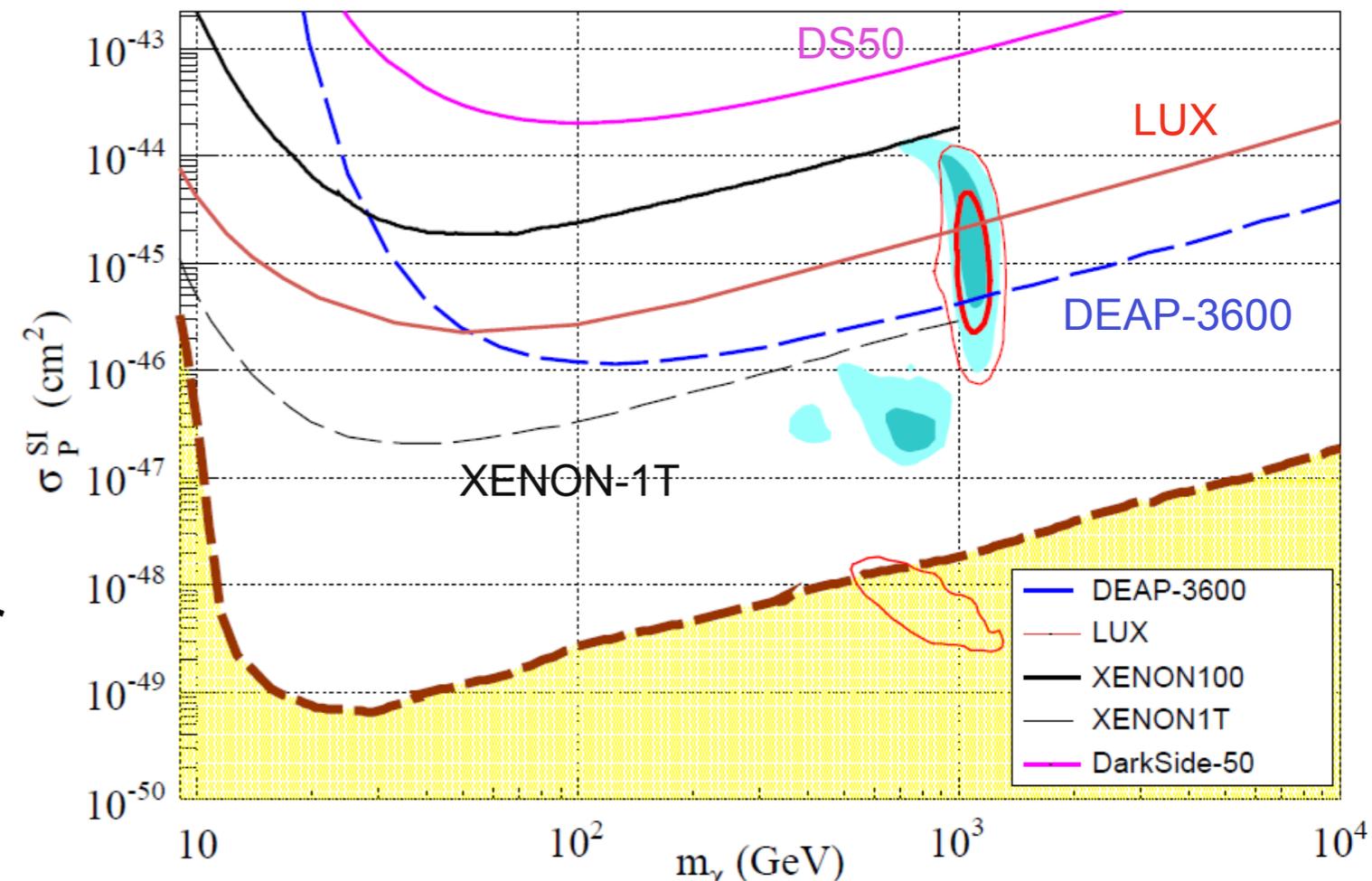
designed for
1-tonne fiducial mass
3 live years

DEAP-3600 status

adapted from Mark Boulay



- Detector filled since Nov 1, 2016
- Collecting DM search data so far $> 0.5E6$ kg-days raw exposure
- So far stable performance good light yield
- Taking physics and calibration data
- plan: continue data collection for ~ 4 years
- Analysis from the 1st fill data ongoing physics publication expected this year

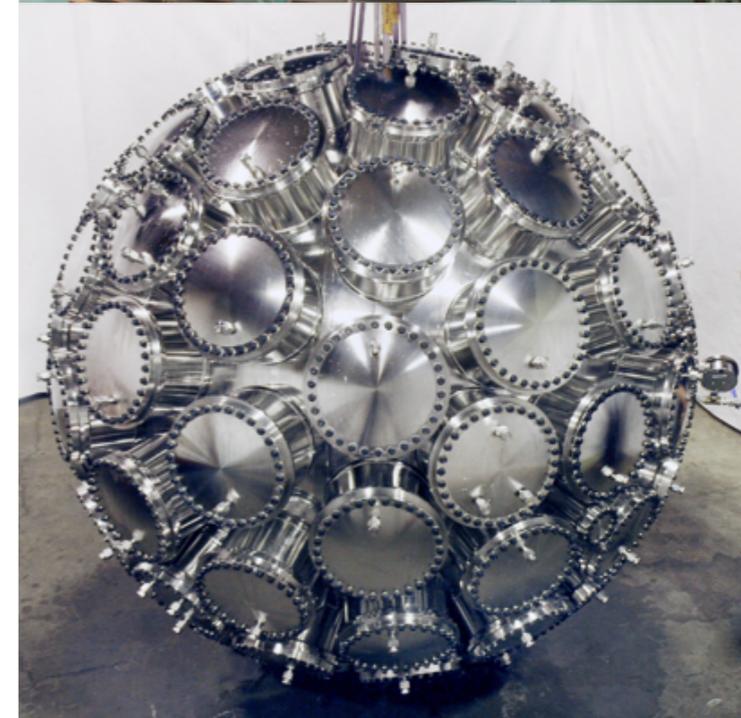


- ▶ Current Project Status
 - Exceptional purity: Gas triplet time constant $>3.6 \mu\text{s}$
 - Detector atmospheric liquid argon fill underway

- ▶ Run plan (2017/18)
 - 3 months - fill and final commissioning
 - 3 months - baseline technical demonstration
 - light yield, background levels, position reconstruction,...
 - 6 months - ^{39}Ar spiked data
 - Pulse Shape Discrimination R&D at 10^{-10} level

- ▶ Lead by Pacific Northwest National Lab since 2014
 - Completion of detector construction (2015)
 - Cool down and commissioning (2016)
 - Favorable Operations Readiness Review (2016/17)
 - Leads project management, operations, data analysis & ^{39}Ar spike program

- ▶ Informs technology choices for 100+ ton experiment



ArDM — a ton-scale LAr TPC for DM research



ETH
Eidgenössische Technische Hochschule Zürich
Swiss Federal Institute of Technology Zurich

ETH Institute for Particle Physics

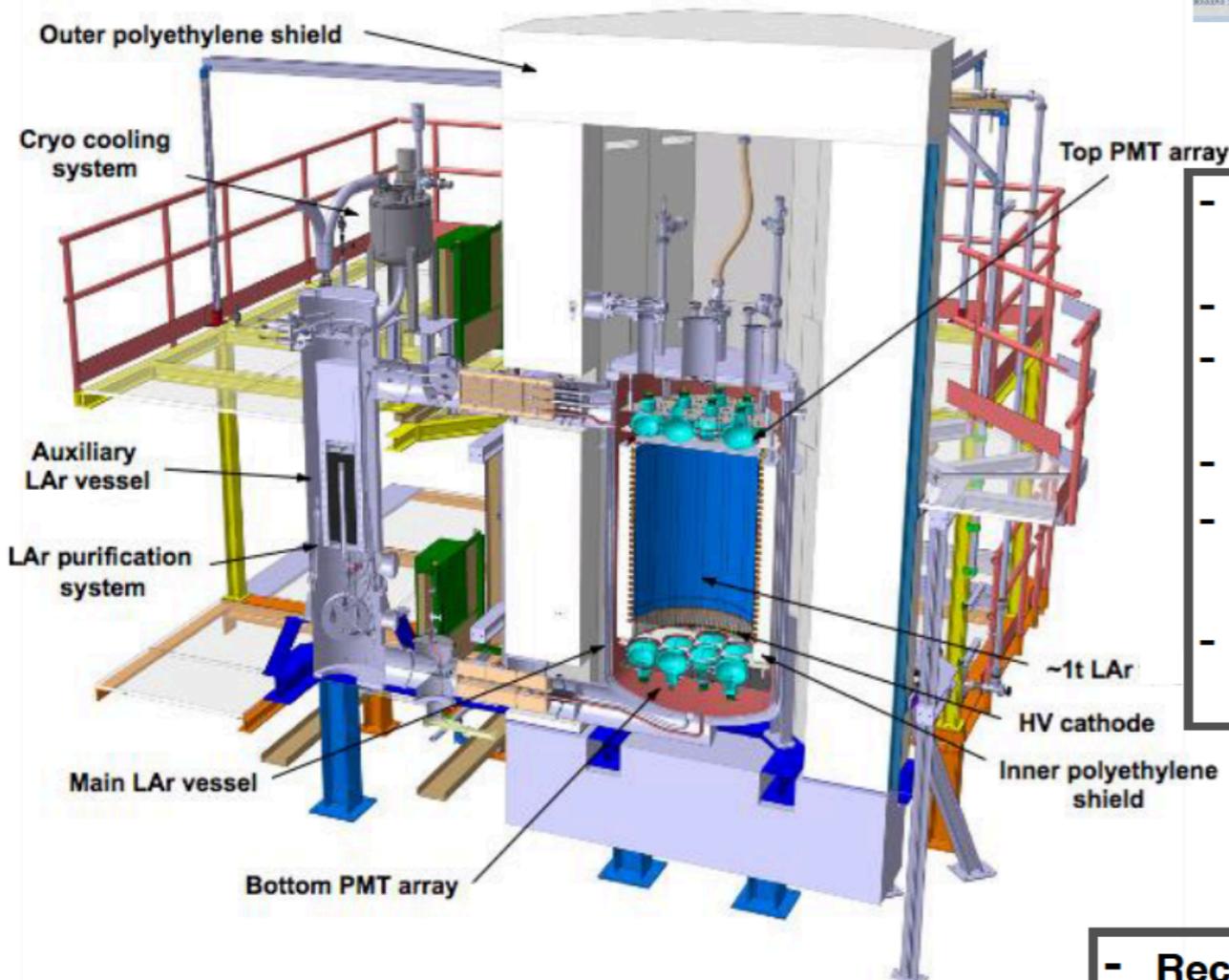
GOBIERNO DE ESPAÑA
MINISTERIO DE ECONOMÍA Y COMPETITIVIDAD

Ciemat
Centro de Investigaciones Energéticas, Medioambientales y Tecnológicas

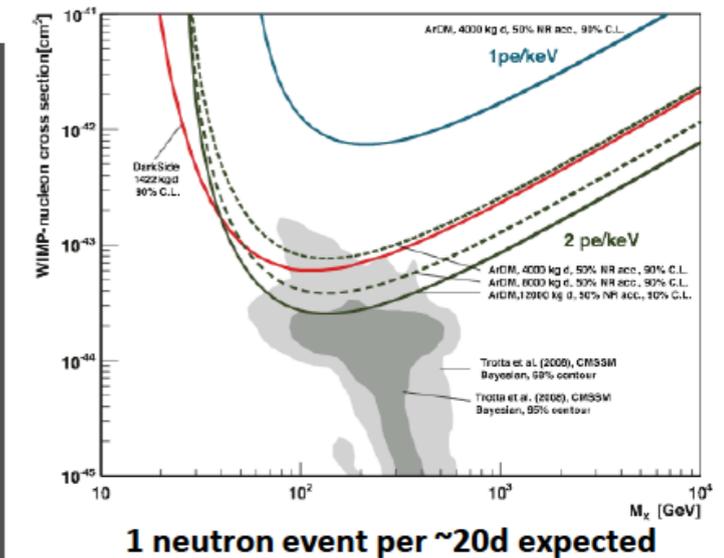


ETHZ led collaboration with CIEMAT, LSC, CERN

- Contribution to searches of higher mass WIMPs!
- Design parameters / developments for future LAr facilities!

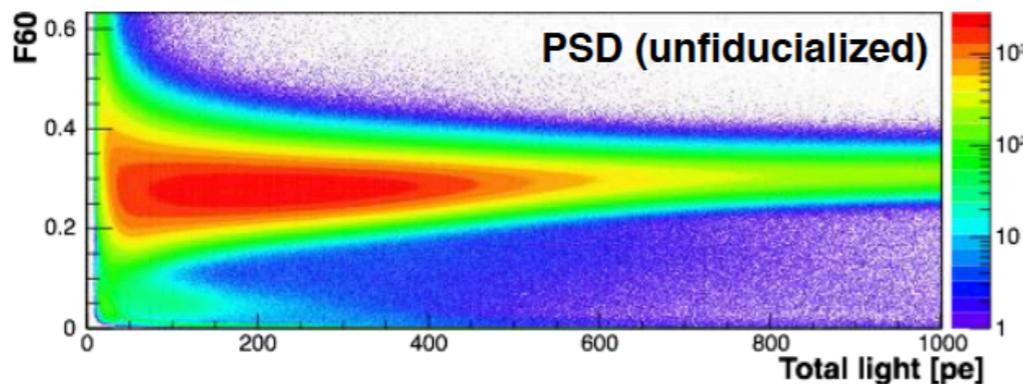


- Experiment commissioned in single phase in 2015
- 3×10^9 triggers recorded
- Low background goal confirmed
- Detector upgraded in 2016
- Double phase run planned for 2017
- Main tasks: Verification of sensitivity and neutron IA

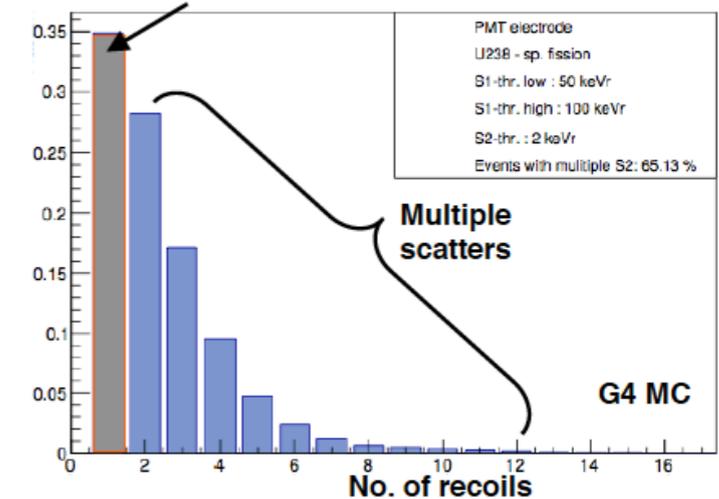


Commissioning of the ArDM experiment... JCAP, 03, 003 (2017)

- Recent integration of ArDM in the DS project
- Combining efforts towards future G2 and G3 facilities



Single scatter neutrons (indistinguishable BG)



³⁹Ar depleted argon research planned for 2018 and beyond

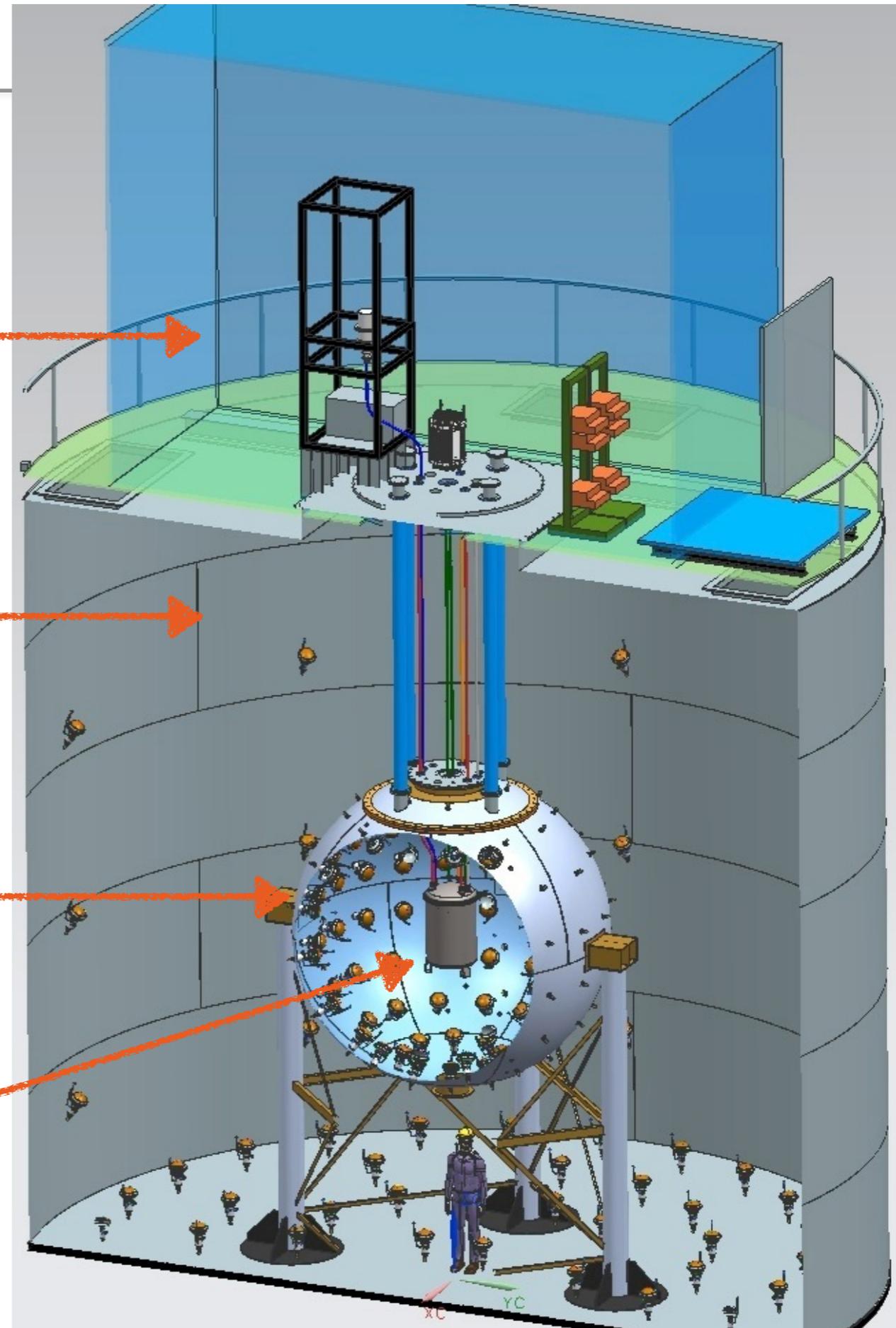
DarkSide 50 at LNGS

Radon-free **Clean Room**
(Rn levels < 10 mBq/m³)

1,000-tonne Water-based Cherenkov
Cosmic Ray Veto

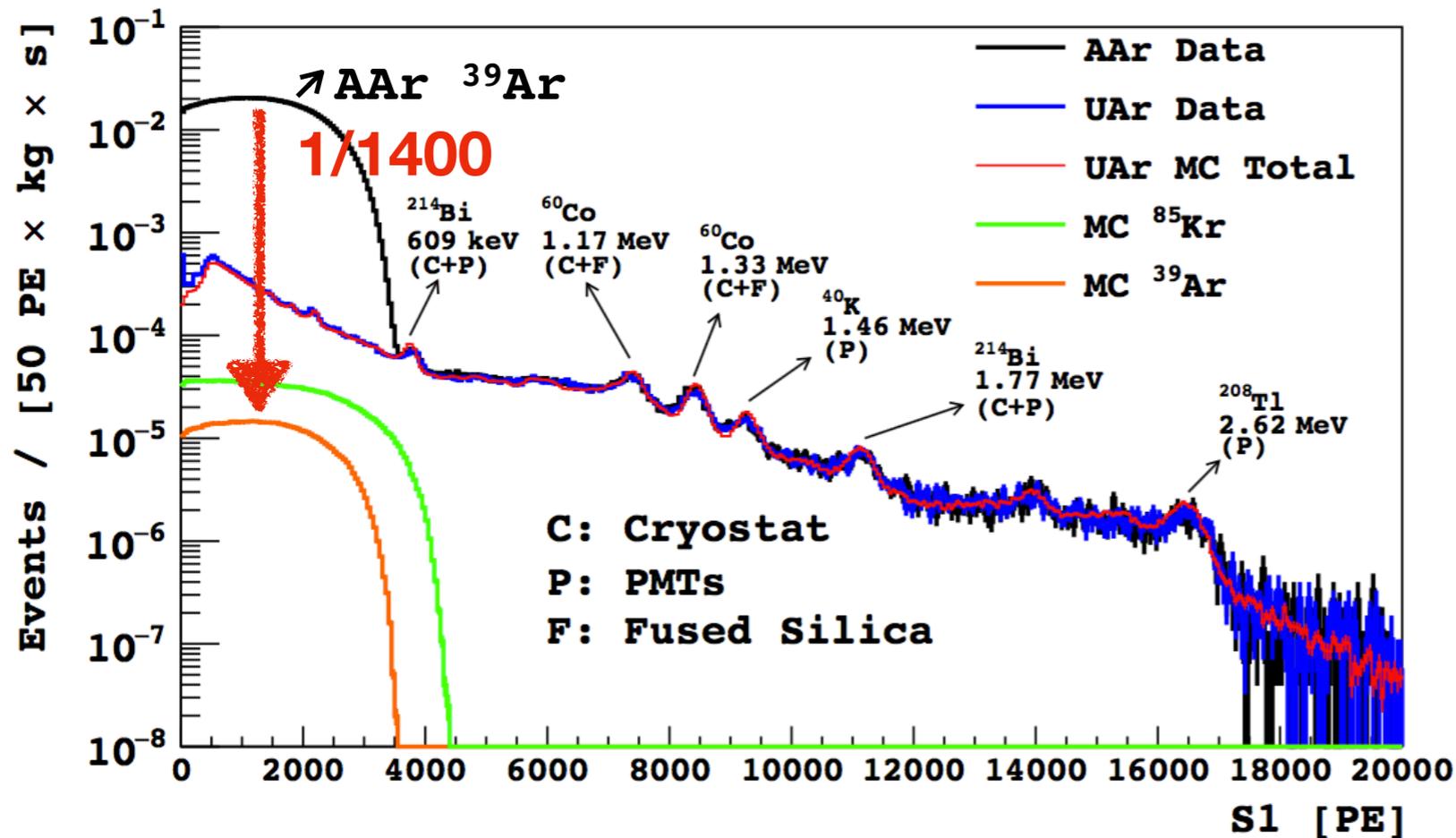
30-tonne Liquid Scintillator
Neutron and γ 's Veto

Inner detector **TPC**



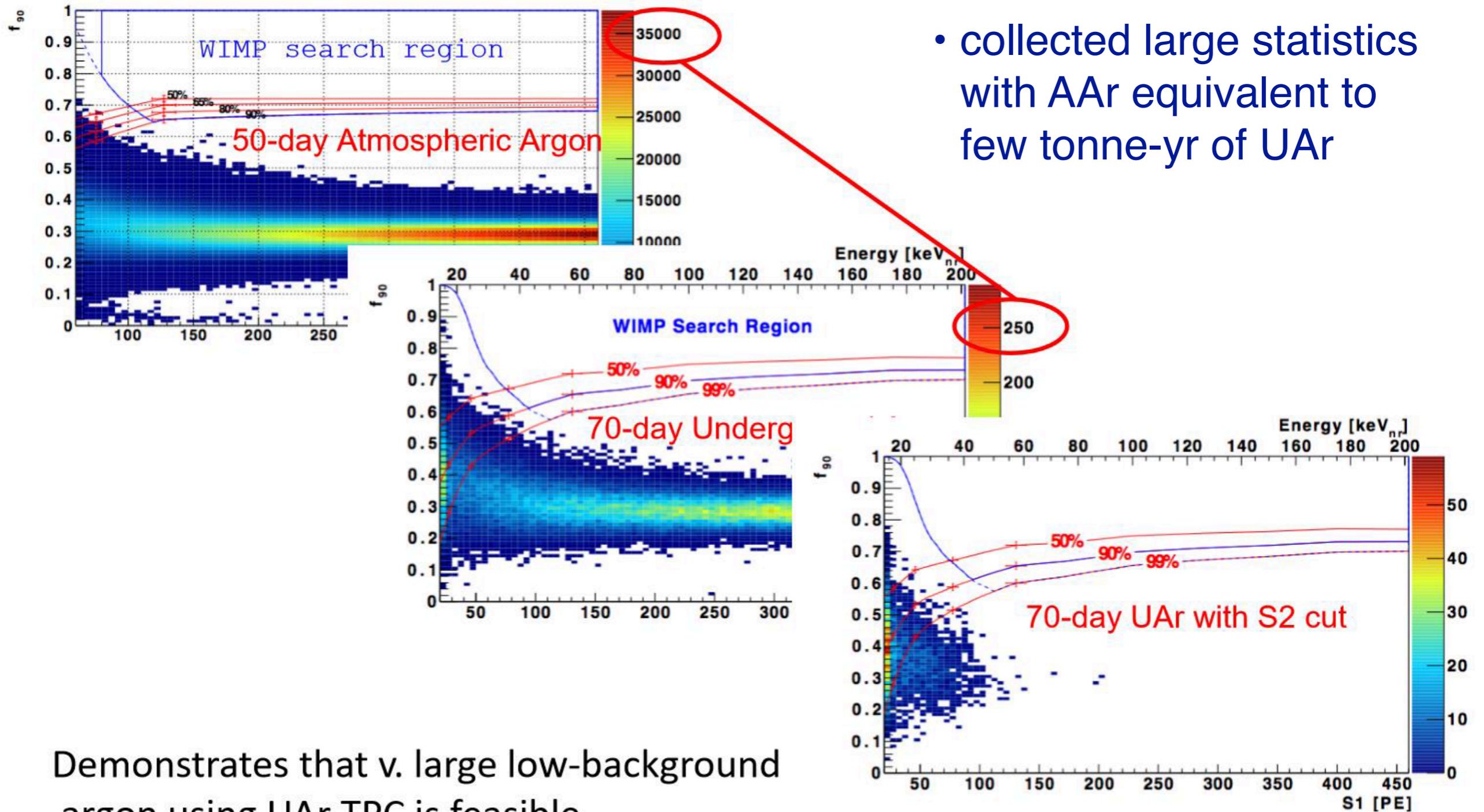


AAr vs UAr. Live-time-normalized S1 pulse integral spectra at **Zero** field. ^{39}Ar reduction factor of **~1400**



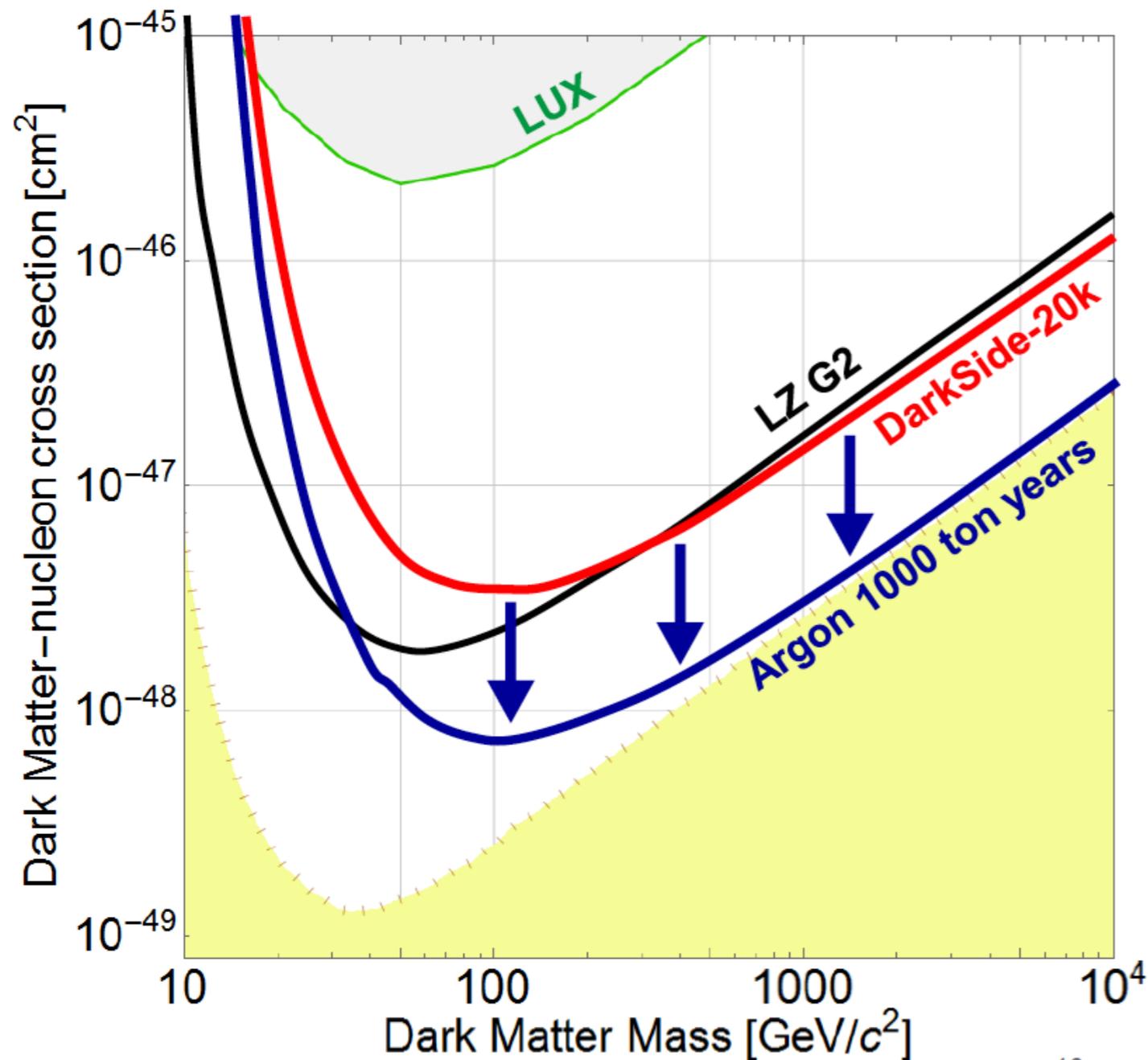
- pioneered the use of low radioactivity argon from underground sources
- introduced the use of an active neutron veto design
- low bg operation of a LAr TPC
- high light yield ~ 8 p.e./keV_{ee}

Low level of ^{39}Ar allows extension of DarkSide program to **ton-scale** detector.



Demonstrates that v. large low-background argon using UAr TPC is feasible

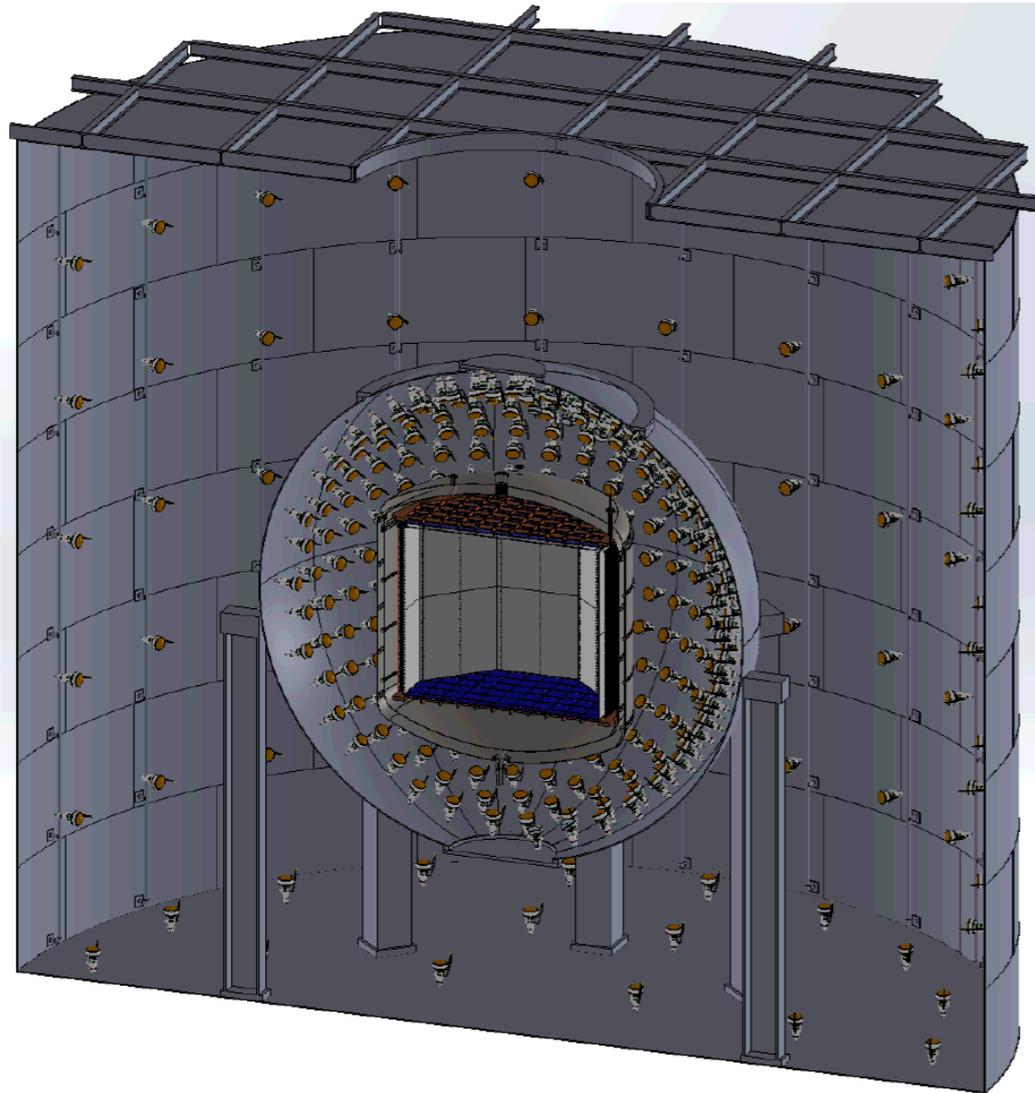
DS-20k and beyond — zoom in on heavy WIMPs



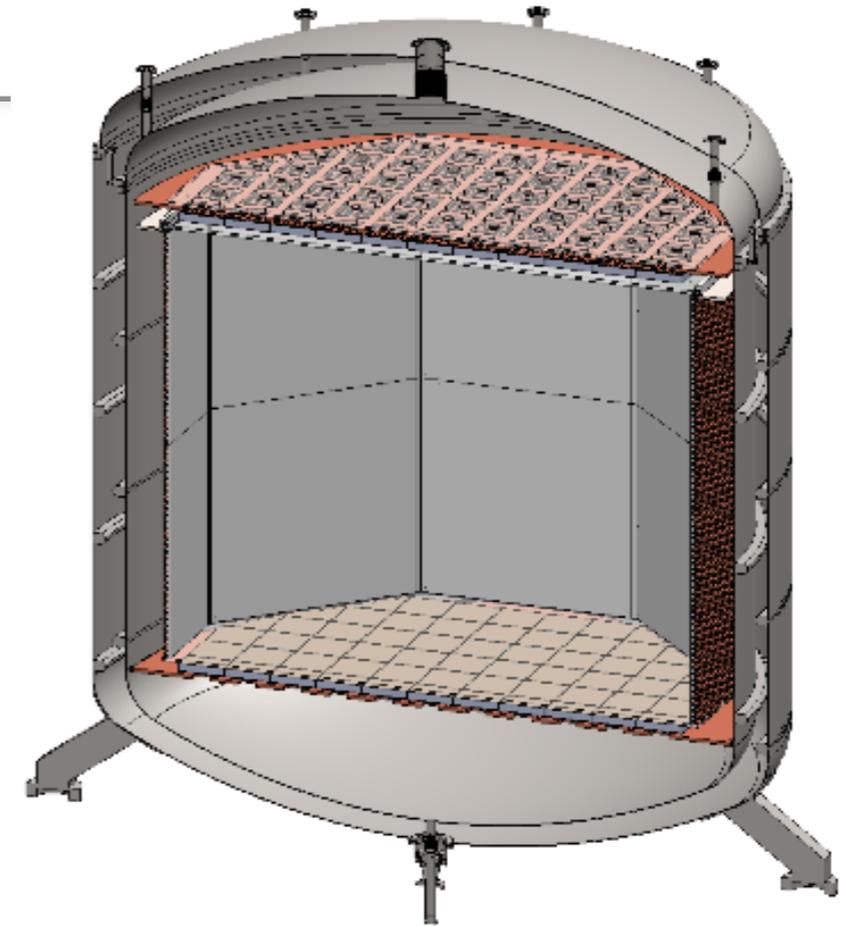
16

- Argon has good sensitivity in high-mass region
- DS-20K (20 tonnes UAr) competitive with LZ — start operation 2021
- 1000-tonne years (future detector) reaches down to neutrino floor
- Complimentary to xenon — only other target allowing such large exposure
- β/γ discrimination: solar pp neutrino ES background not a concern — in X1T, LZ expected dominant bkg at ~ 0.5 event per tonne-year after recoil discrimination

DS-20k at LNGS



- 20 tonnes low radioactivity Ar
- Scale-up from DS-50
- Global LAr community
- Advanced design (Yellow Book)



- background-free, 100 tonnes-yr exposure
- SiPMs replace cryogenic PMTs (single-p.e. 5x5 cm² tile demonstrated in March 2017)
- complementary to LHC searches

Urania (Underground Argon):

- Expansion of the argon extraction plant in Cortez, CO, to reach capacity of ~**150 kg/day** of Underground Argon

Aria (UAr Purification):

- Very tall column in the Seruci mine in Sardinia, Italy, for high-volume chemical and isotopic purification of Underground Argon

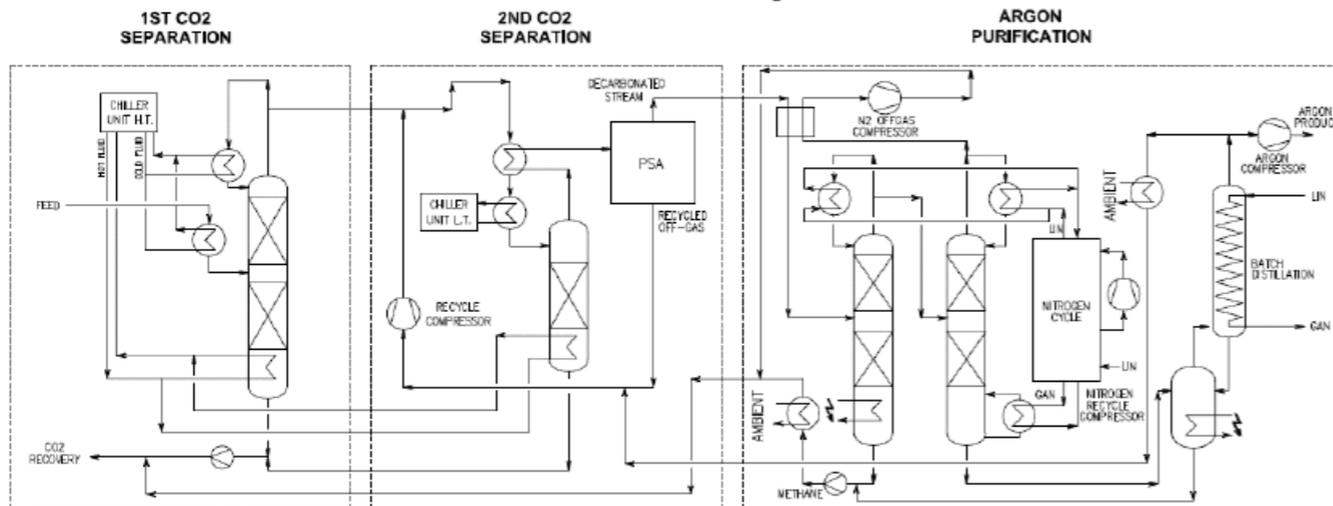
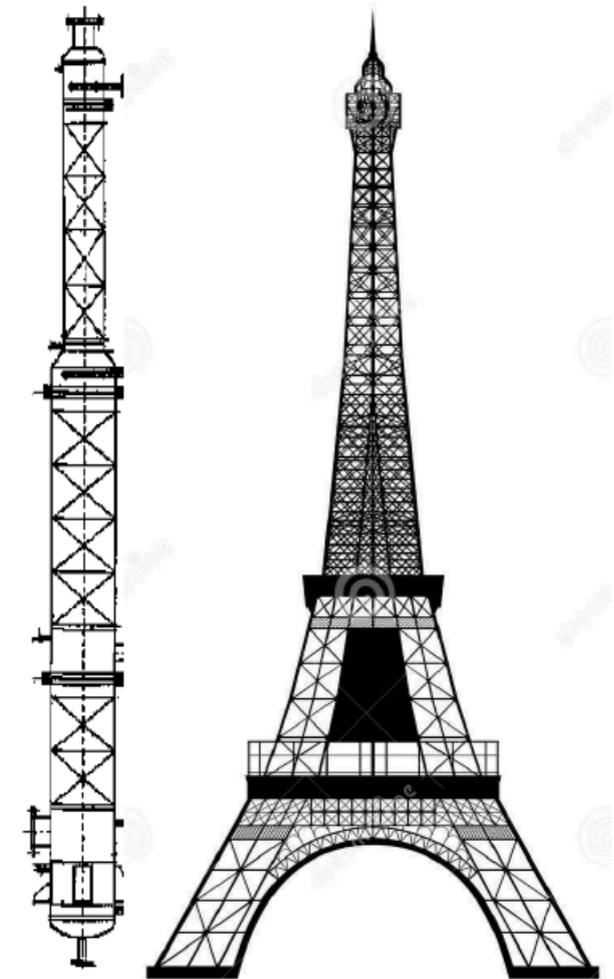
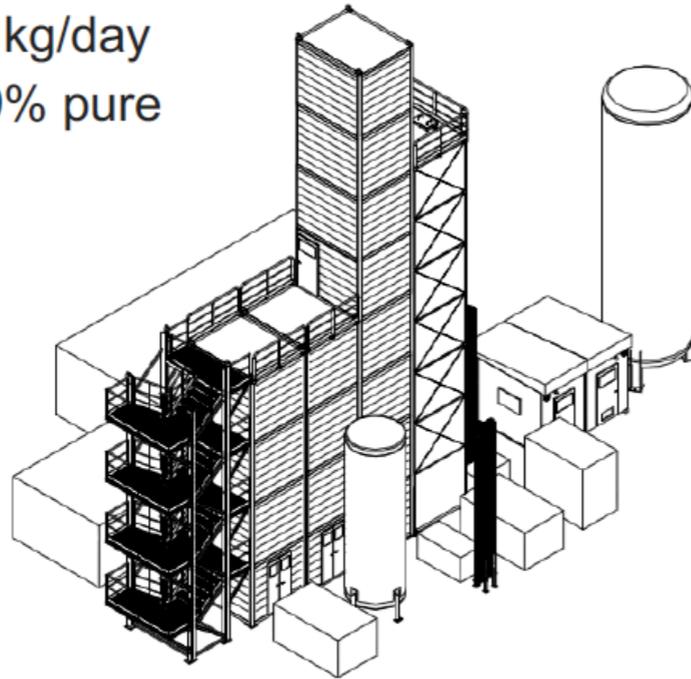
Urania and ARIA

► Urania (planned)

- Extracts argon from CO₂
- Same source as DarkSide-50 target
- Production:
 - 100 kg/day
 - 99.9% pure

► Aria (under construction)

- Final argon purification
- Capable of isotope separation (³⁹Ar depletion)
- Located in the Seruci mine in Sardinia, Italy



Beyond DS-20k — few hundred tonne LAr detector

- 1000 tonne-yr, background-free exposure reaching the neutrino floor
- Sensitivity to WIMPs $>$ few tens of GeV/c^2
- mid 2020's

- Large collaboration will pursue an integrated program
- Single/dual-phase design under consideration

- Physics program largely complementary with the LHC and to the expanding effort in searching low-mass WIMPs
- Solar neutrino program possible

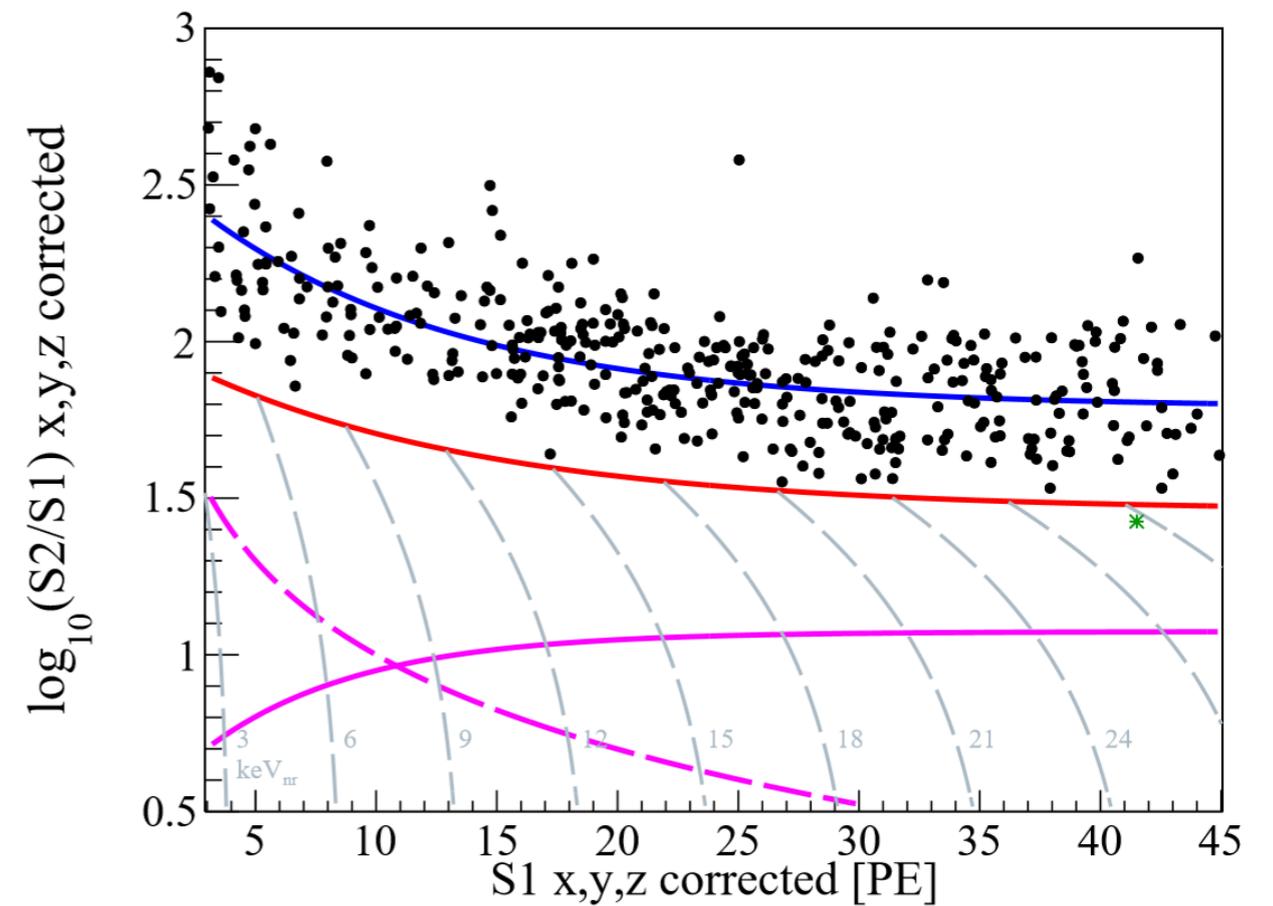
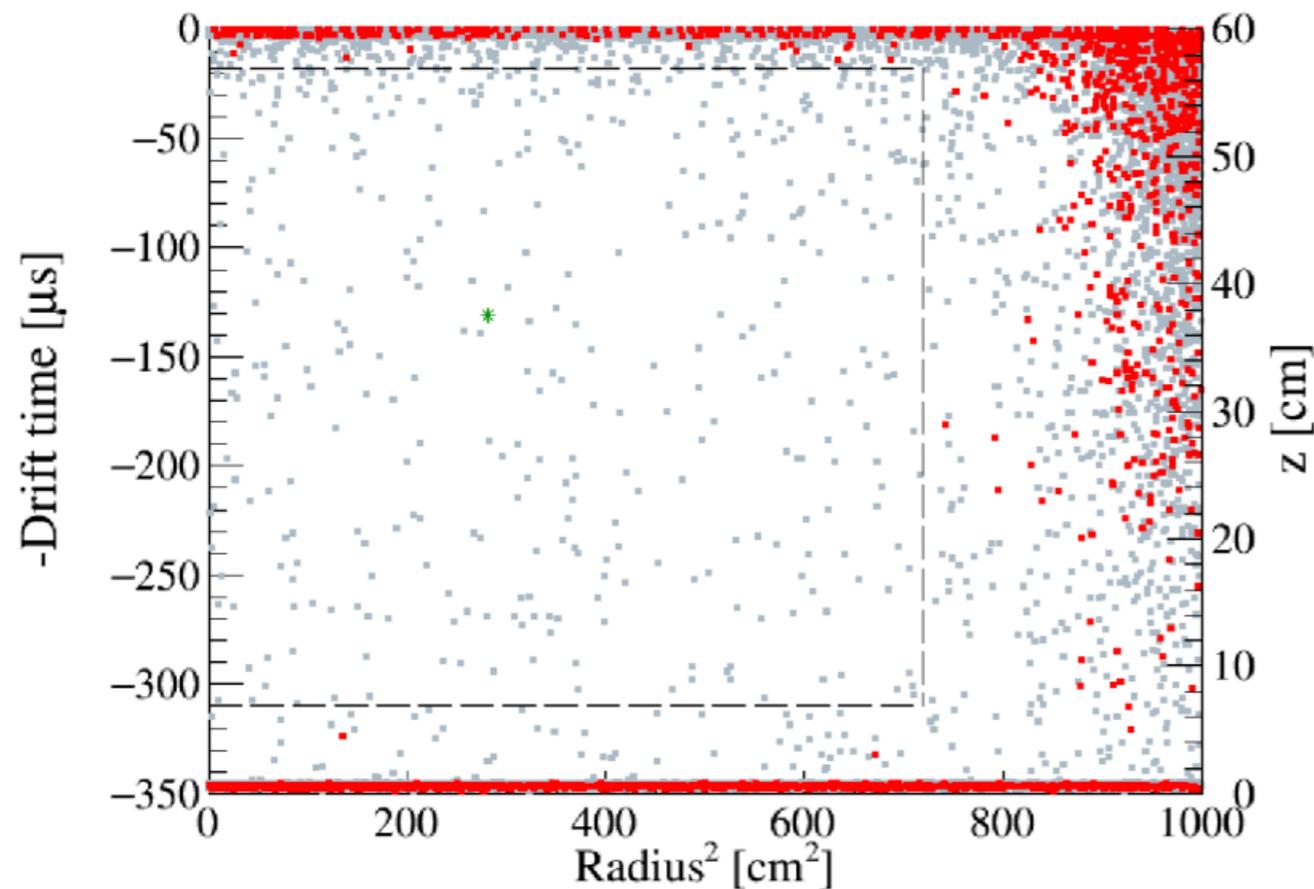


- LXe and LAr detectors are leading players in the search for WIMP dark matter heavier than a few GeV
- The tonne-scale dark matter search is under way with XENON1T and DEAP-3600 — more to follow in the next few years
- LXe 2-phase TPCs allow for single electron sensitivity for low-mass WIMPs ($\sim\text{GeV}/c^2$ scale)
- Future detectors with target mass of tens \rightarrow hundreds of tonnes are planned to search for WIMPs down to the neutrino floor
- Compelling, complementary physics program to the LHC (and searches for light dark matter)

Gray: all

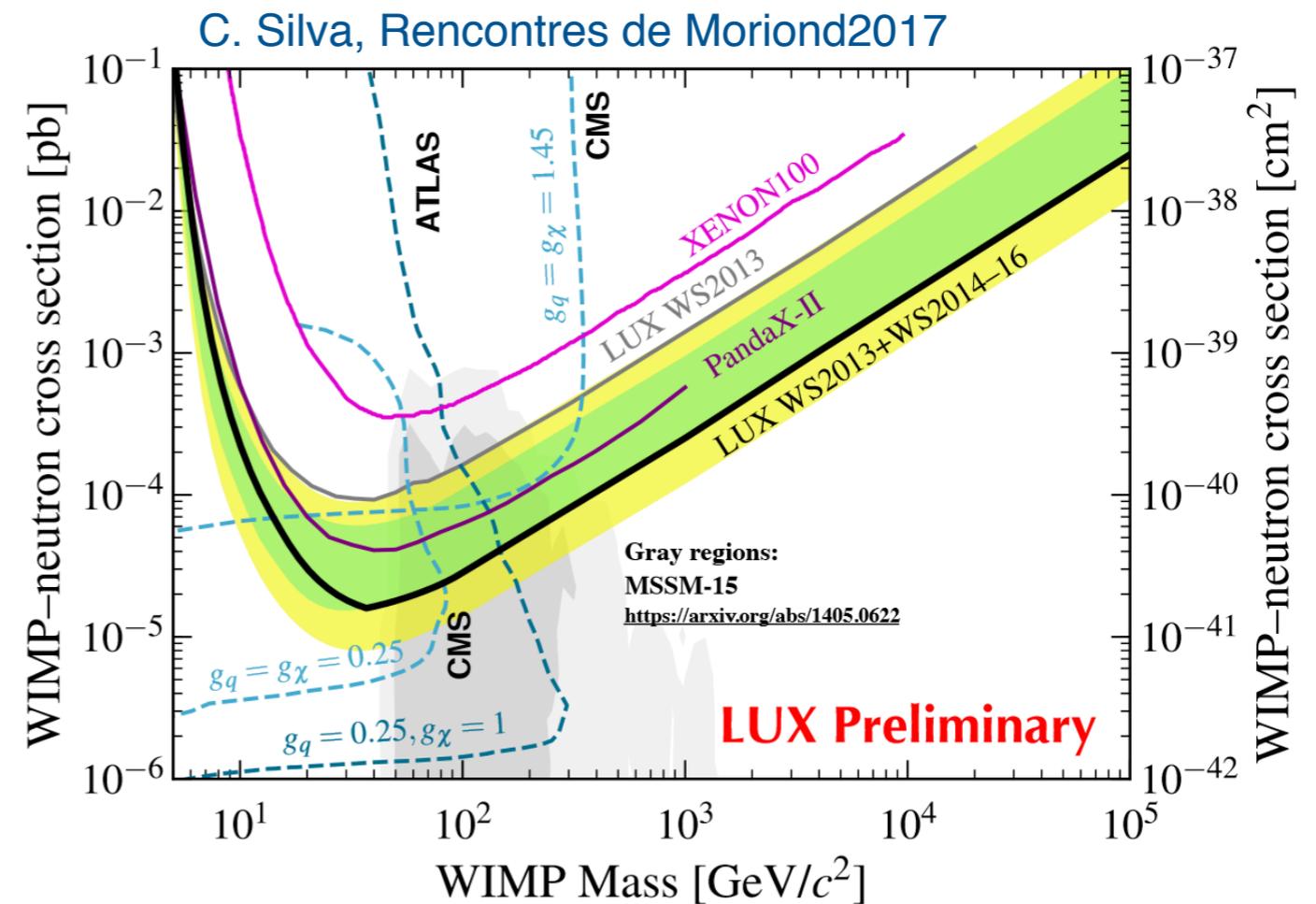
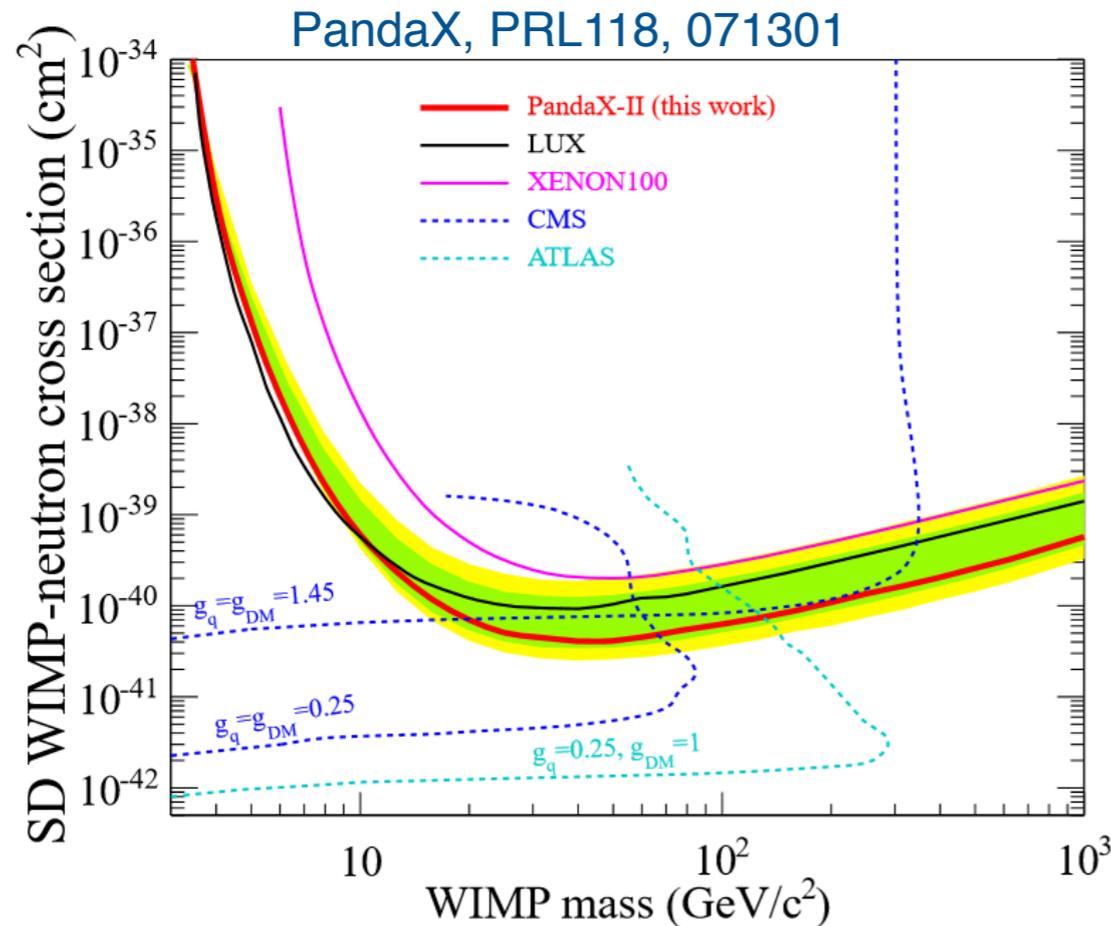
Red: below NR median

Green: below NR median and in FV



Spin-Dependent WIMP interactions

adapted from Jianglai Liu, Pheno 2017



- Spin-dependent WIMP-neutron cross-section on ^{129}Xe and ^{131}Xe
- complementary to collider searches at lower masses

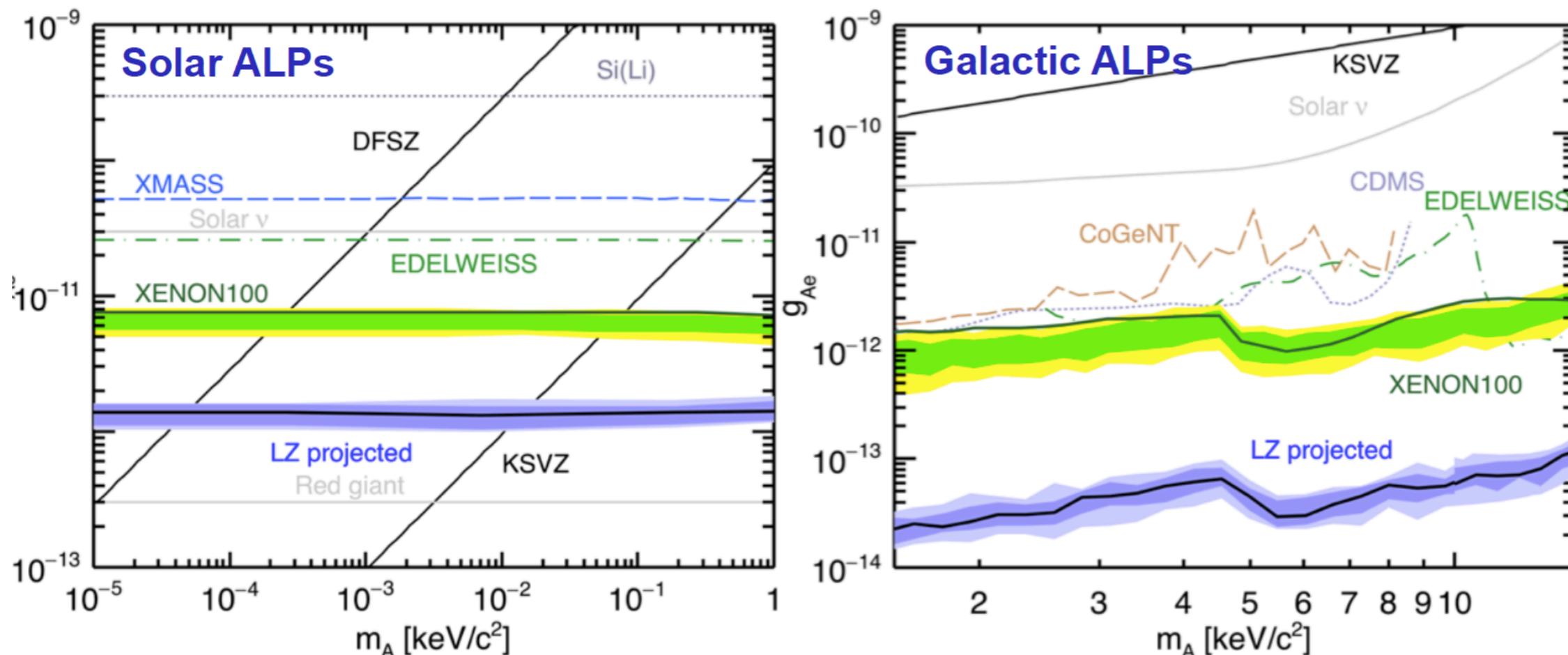
LXe TPCs: sensitivity to axion-like particles (ALPs)

Axion-Like Particles (ALP)



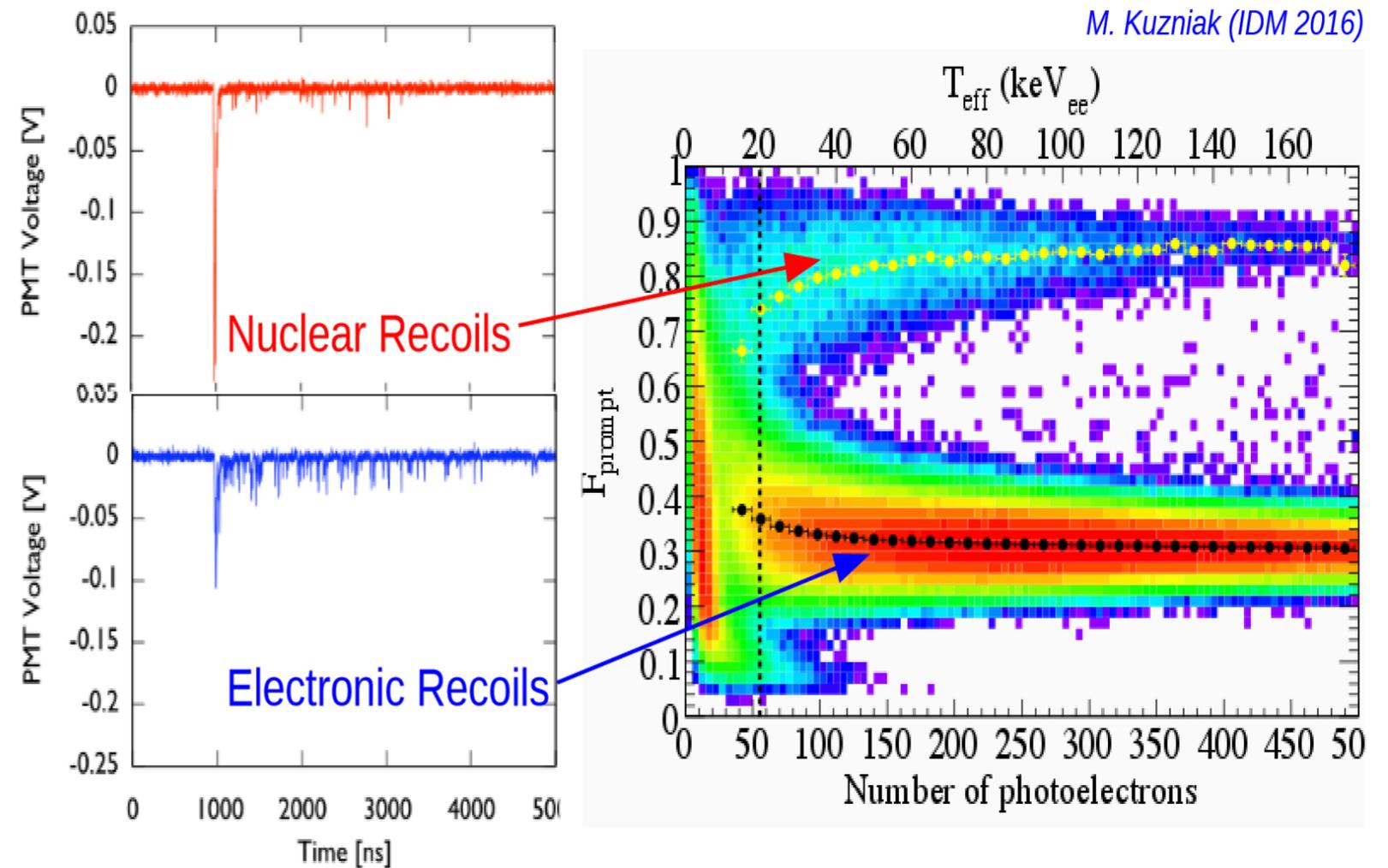
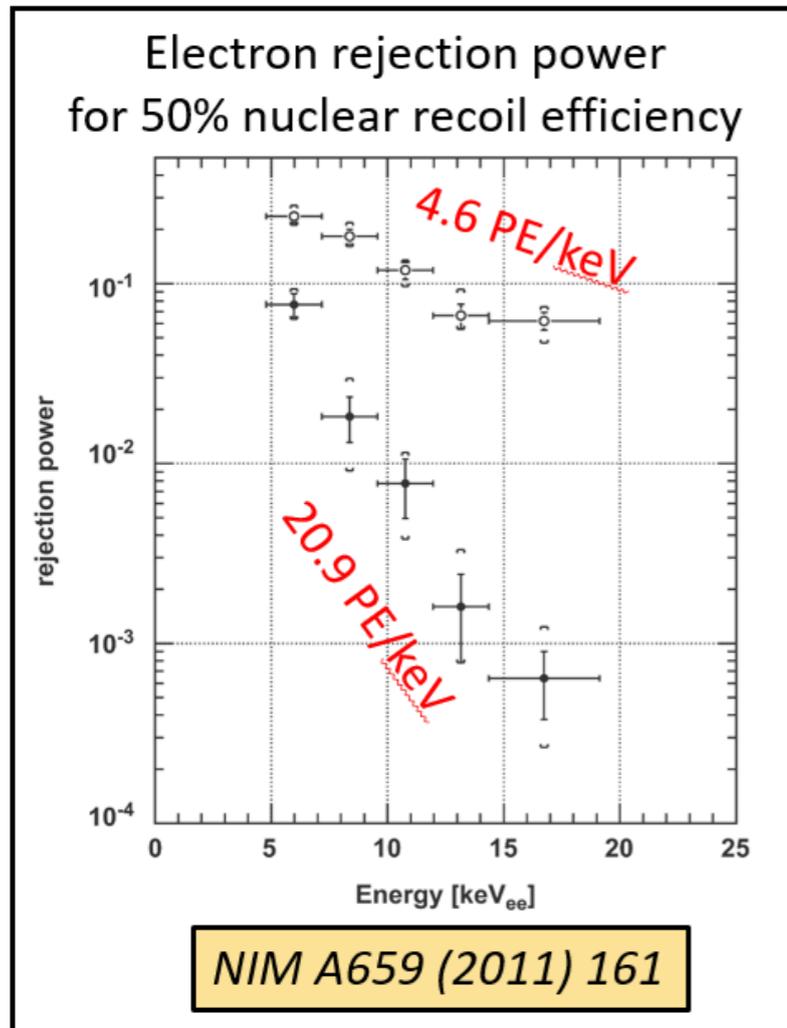
- **Sensitivity to axions and ALPs via the axioelectric effect:**
 - Nonrelativistic galactic ALPs (DM candidates)
 - ALPs emitted by brems/Compton in the Sun
- **Technique pioneered in Xenon100 (see arXiv:1404.1455)**

adapted from
Maria Elena Monzani



LZ TDR, in preparation

PSD study with a small setup



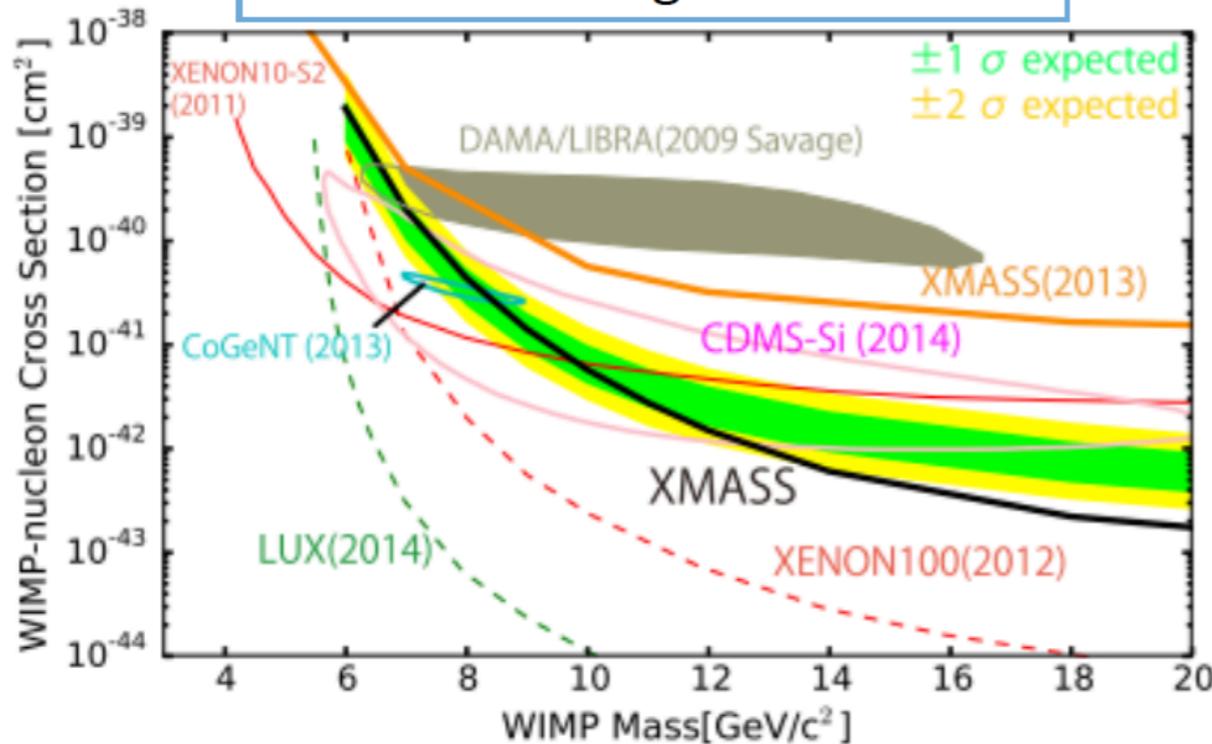
XMASS: annual modulation results

adapted from Shigetaka Moriyama

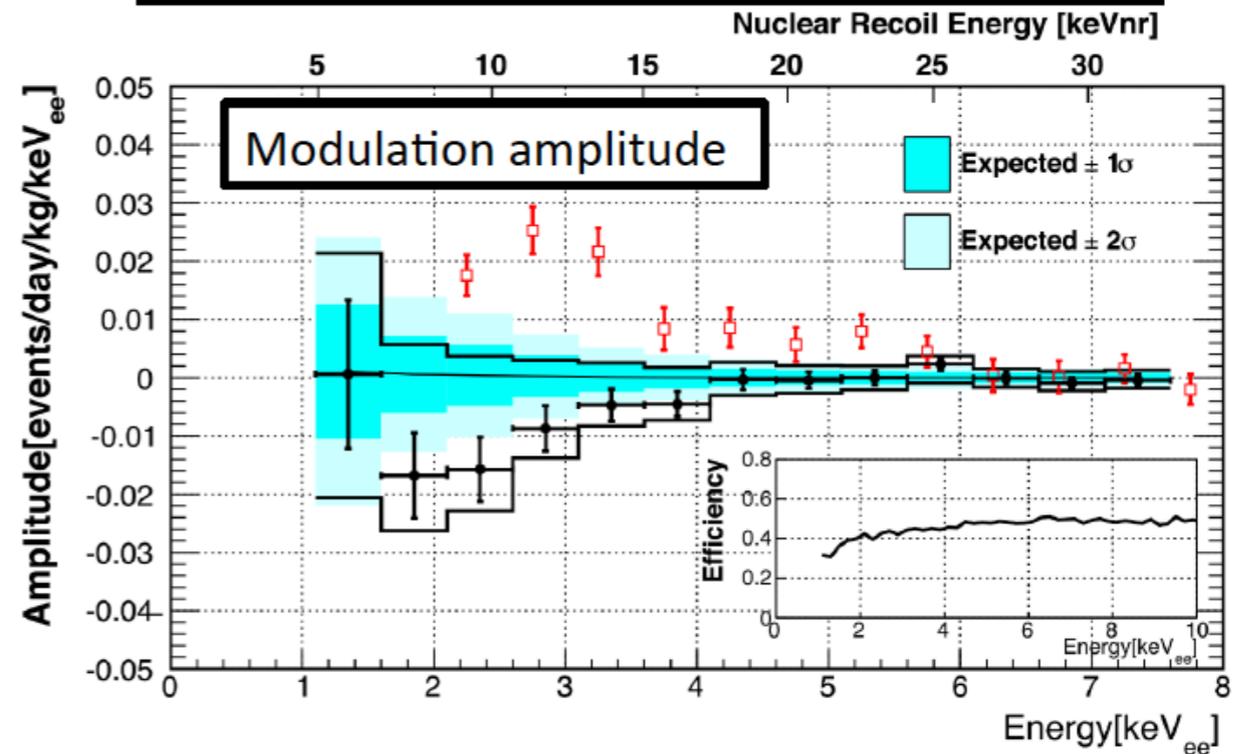
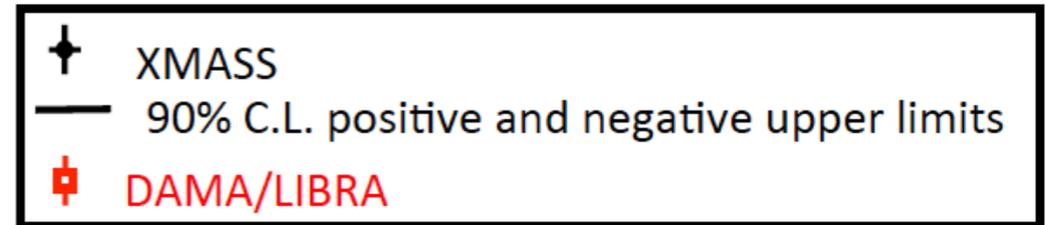
- DAMA/LIBRA (14 cycles)
~250 kg NaI(Tl): **1.33 ton×yr, no PID**
 - XMASS-I (1.5 cycle)
 - ~800kg LXe: **0.82 ton×yr, no PID**
- Phys. Lett. B 759 (2016) 272

First extensive annual modulation search proving this region with an exposure comparable to DAMA/LIBRA.

Result assuming WIMP model

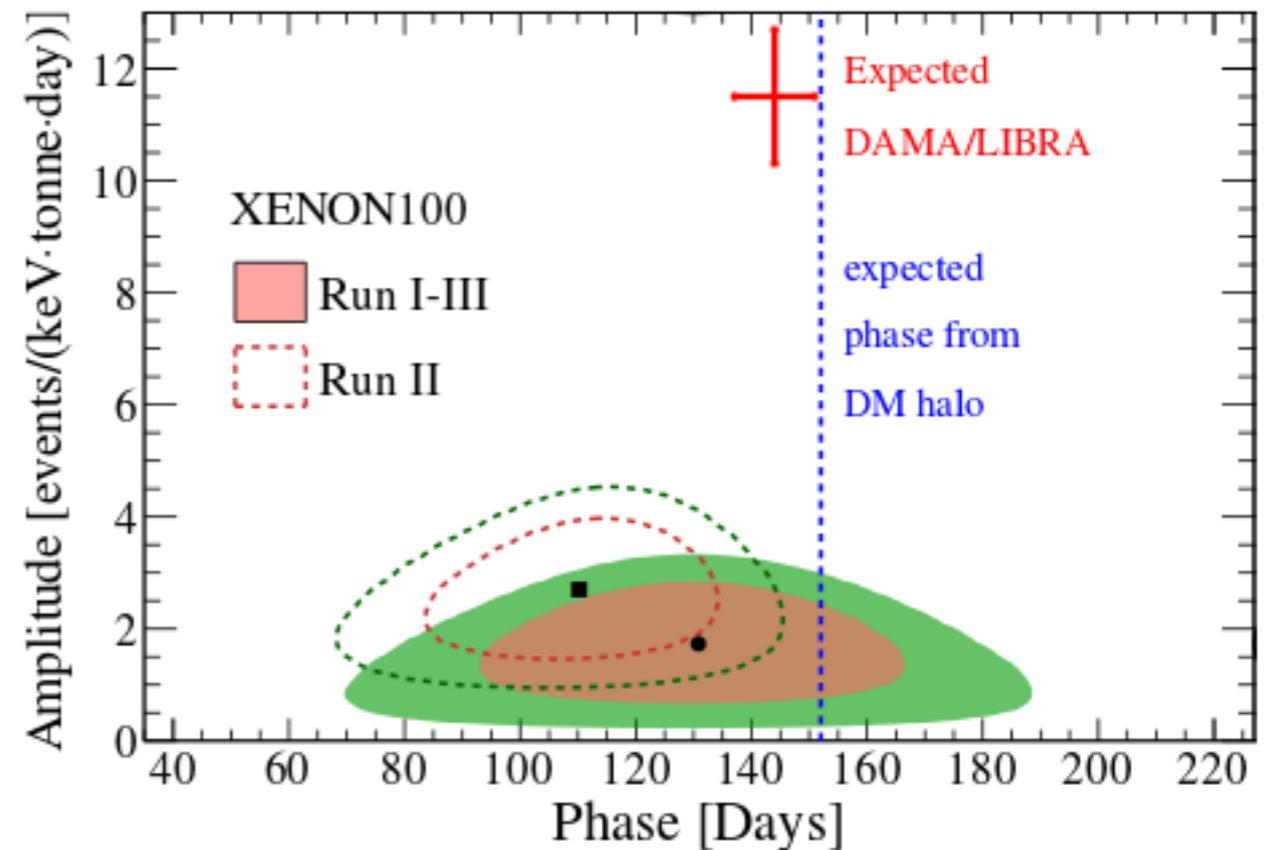
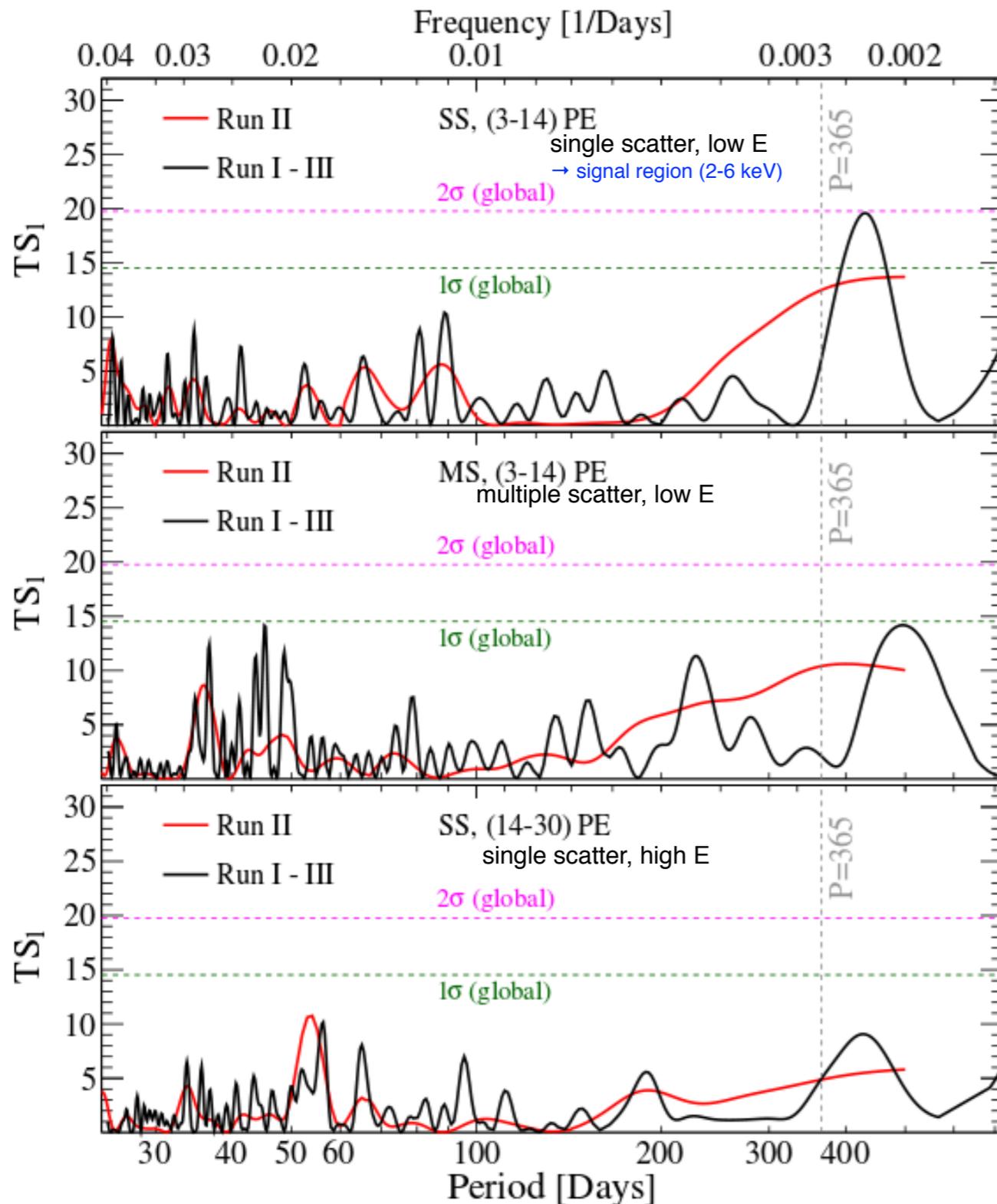


Results on individual energy bin



XENON100 — 4-year annual modulation

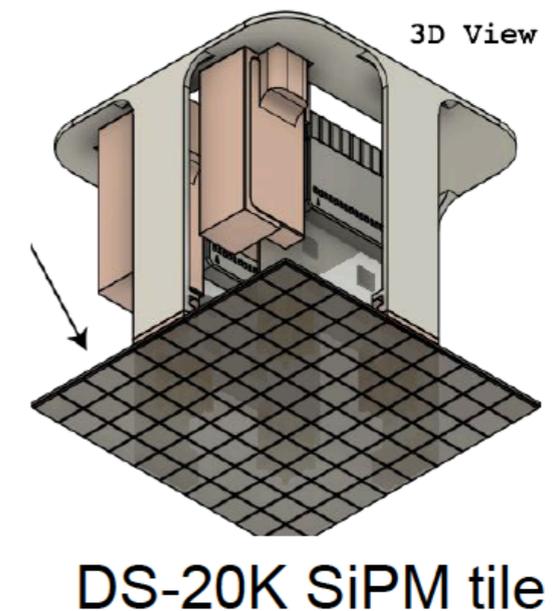
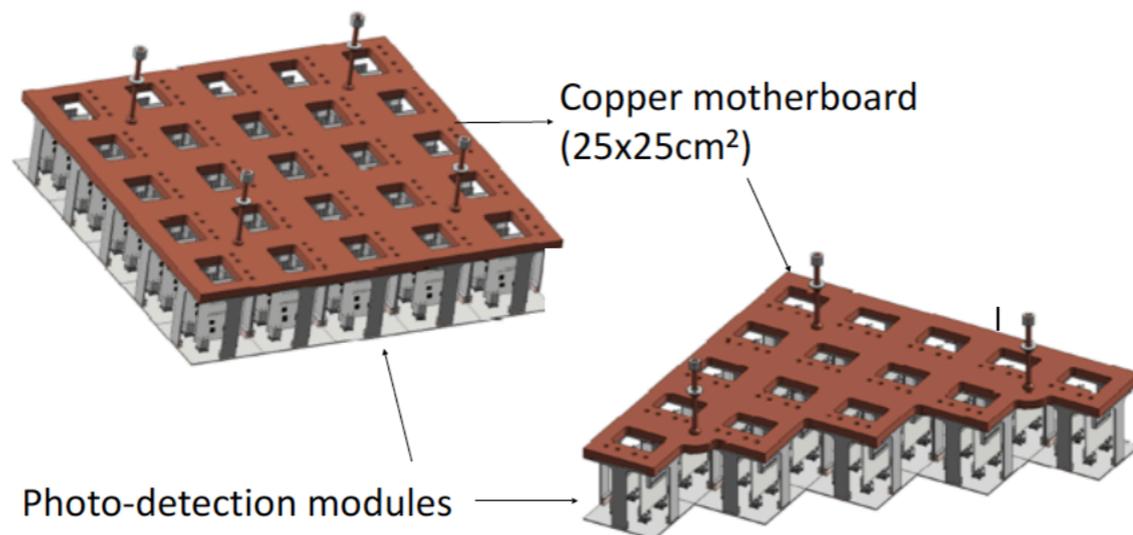
adapted from Marc Schumann



- improves upon previous analysis PRL 115, 091302 (2015)
- no significant modulation observed
- Dark matter explanation of DAMA/LIBRA
- signal excluded @ 5.7σ

DS-20k SiPMs

- Photon Detection Efficiency (PDE): 45% requirement met and surpassed
- Dark Count Rate (DCR): 0.1 Hz/mm² requirement met and surpassed
- Challenge in tiling due to 50 pf/mm² capacity. Signal-to-Noise Ratio (SNR) rapidly decreases with increasing surface. The steps:
 - 2×2 cm² tile: fully demonstrated September 2016
 - **5×5 cm² tile: fully demonstrated March 2017**



ARIA: isotopic purification via cryogenic distillation

Prototype column allows 10 kg/day purification

Full column: 100 kg/day

x10 reduction per pass

