



LDMA 2017

NEWS-G

Ultra-Light Dark Matter searches with a Spherical Gaseous Detector

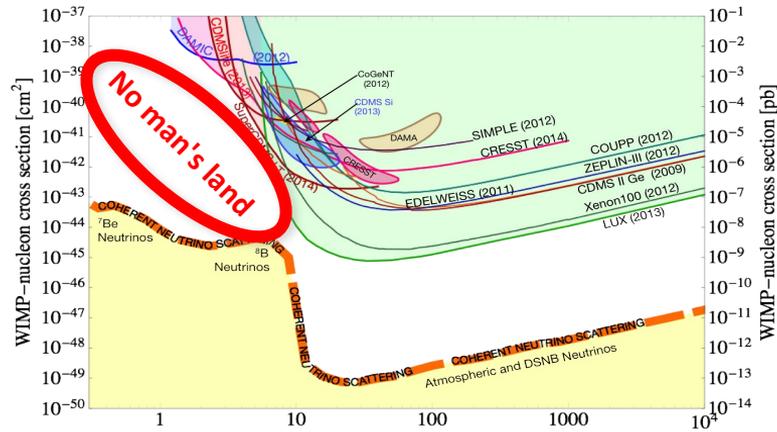
Ioannis KATSIIOULAS

CEA, Saclay

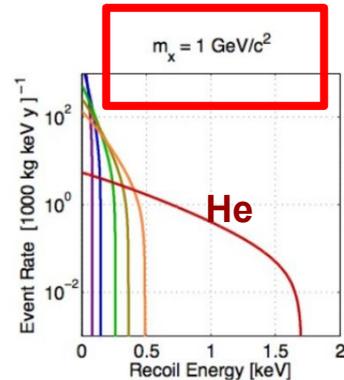
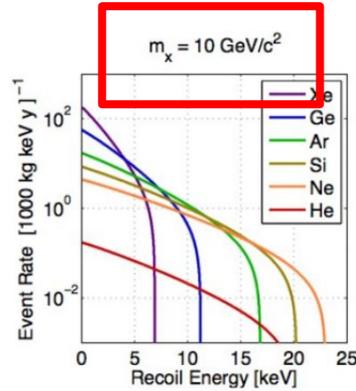
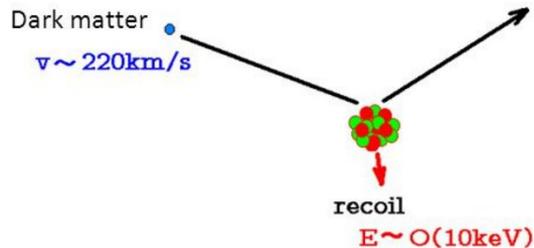
ioannis.katsioulas@cea.fr

New Experiments With Spheres

Search for WIMPs in the 100 MeV – 10 GeV mass range



Spin Independent couplings



Motivation:

Non findings at:

- Passive experiments
- LHC

Method:

Direct detection using:

- A Novel spherical gaseous proportional counter

- Light gases as target (H, He, Ne) for a better projectile - target kinematical match

Spherical Proportional Counter - Spherical TPC

Fun facts

Old LEP RF cavities



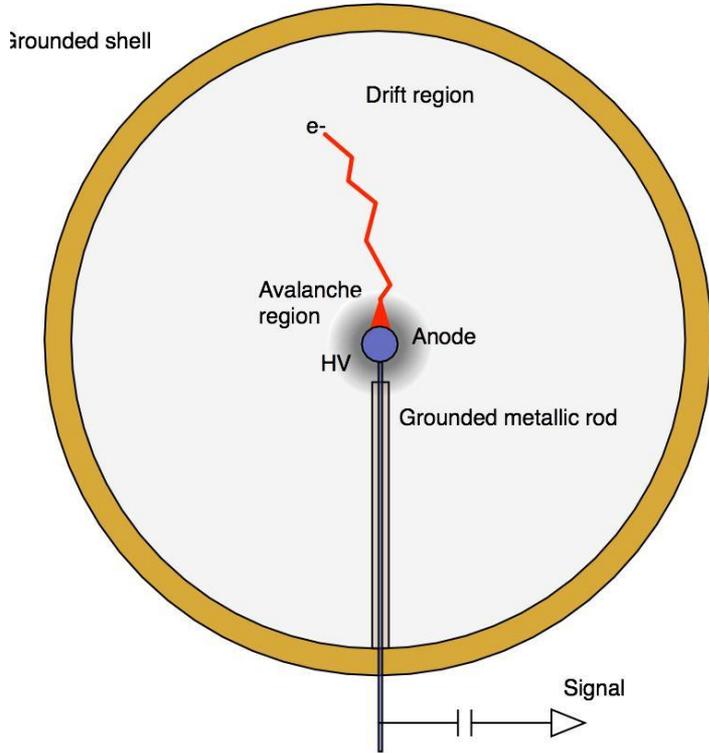
Spherical gaseous detectors



*In the picture:
I.Giomataris, G.Charpak*

Spherical Proportional Counter

Principal of operation



Greatly varying field
along the radius

$$E = \frac{V_0}{r^2} \frac{r_1 r_2}{r_2 - r_1} \approx \frac{V_0}{r^2} r_1$$

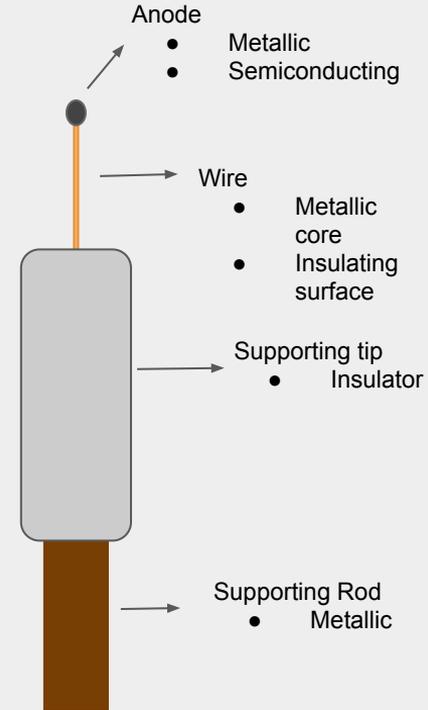
$r_1 = \text{anode radius}$

$r_2 = \text{cathode radius}$

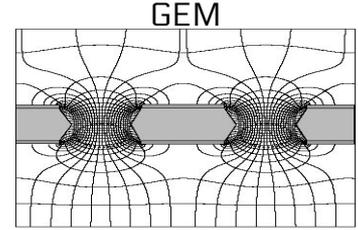
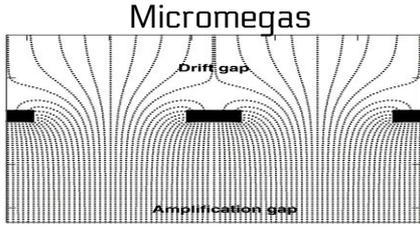
Natural division of the
volume in two

- Drift volume
- Multiplication volume

The Sensor

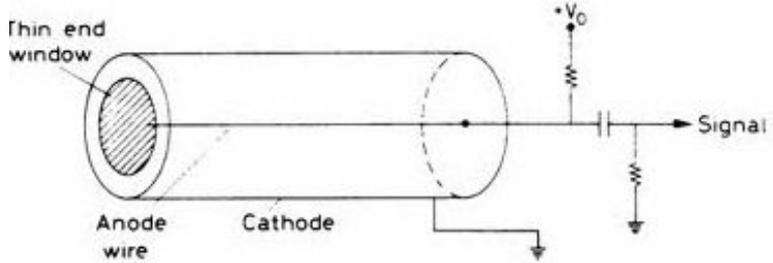


Capacitance dependence on size



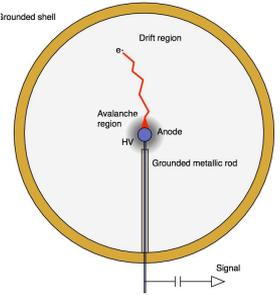
Parallel Plate Detector

$$C \approx S > 1nF$$



Cylindrical Proportional Counter

$$C = 2pL / \ln(b/a) \gg 10 \text{ pF}$$

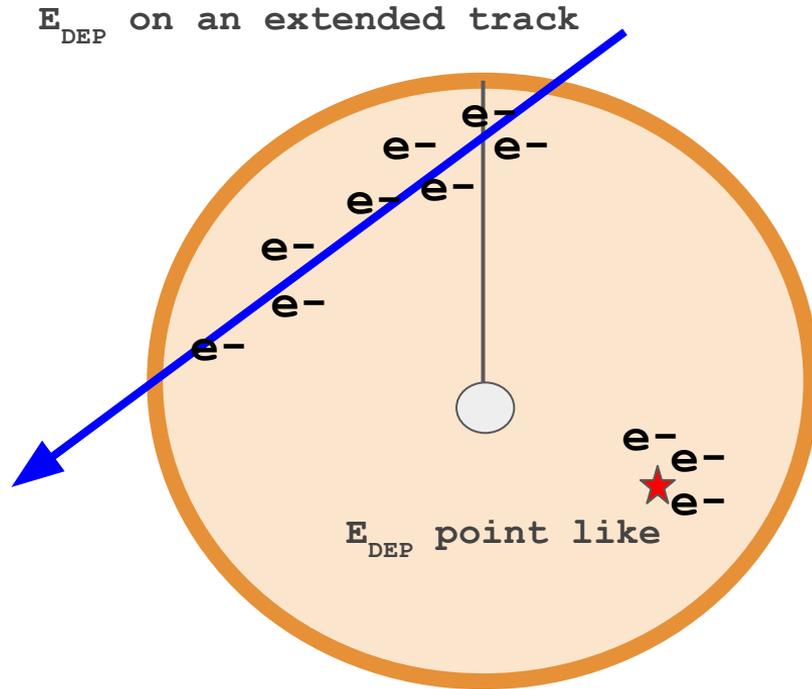


Spherical Proportional Counter

$$C \approx r_1 < 1pF$$

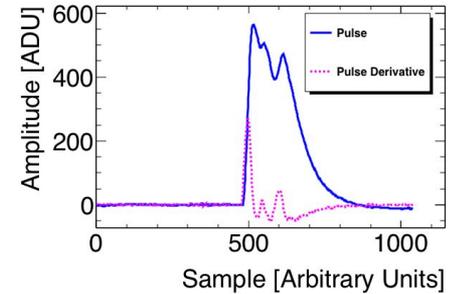
Large Size Detector
+
Robust construction

Extended versus point like energy deposition

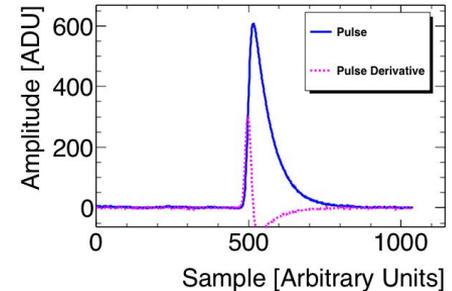


$$\sigma(r) \propto (r/r_{\text{sphere}})^3, \text{ e- drift time dispersion}$$

Muon induced pulse



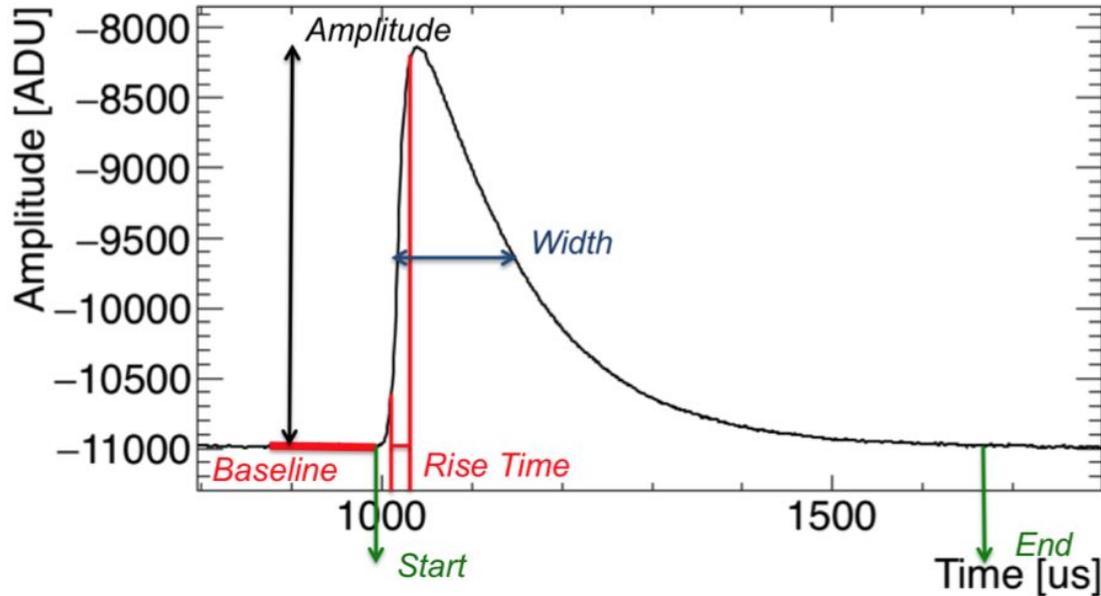
Ion induced pulse



Induced Pulses

Pulse Shape Analysis (PSA) parameters

Long Tail Pulse



Parameters

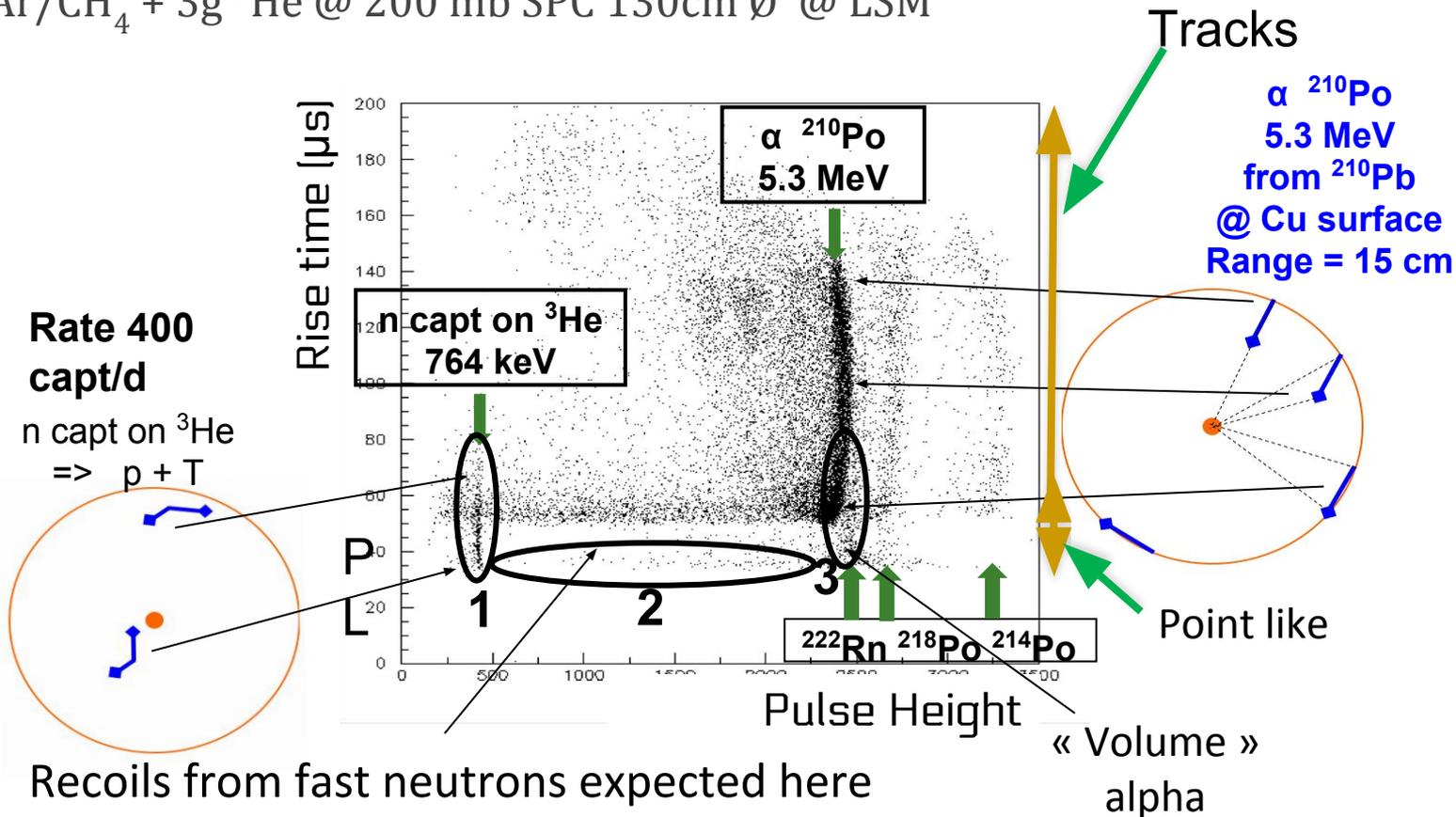
- Baseline
- Noise
- Peak height
- Rise time
- Width
- Integral
- Number of peaks

A lot of information hiding in the pulse shape

Pulse Rise time & Width ∞ Drift time dispersion

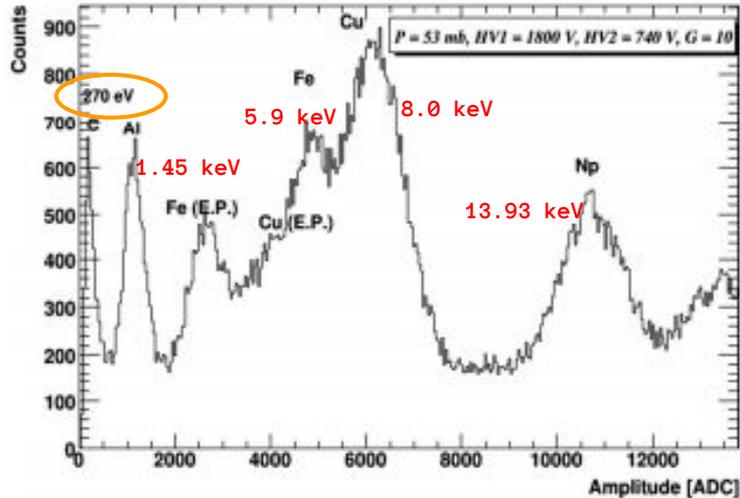
Illustration of particle identification – Background rejection

Run with Ar/CH₄ + 3g ³He @ 200 mb SPC 130cm Ø @ LSM



Low energy capabilities

E. Bougamont et al, Journal of Modern Physics, Vol. 3 No. 1, 2012, pp. 57-63.



SPC Φ 130 cm

Gas: Ar+2%CH₄

Detection of fluorescence

X-rays

$^{241}\text{Am} \rightarrow ^{237}\text{Np} + ^4\text{He} + 5.6 \text{ MeV}$

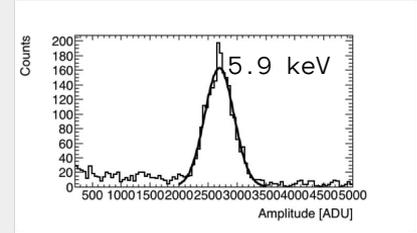
Lines

Al \rightarrow 1.45 keV

Cu \rightarrow 13.93 keV

$^{237}\text{Np} \rightarrow$ 13.93 keV(L _{α})

17.60 keV(L _{β})



SPC Φ 30 cm

**Irradiation by an ^{55}Fe
source (5.9 keV)**

Resolution (σ) < 9%

- **Single electron detection**
- **Energy threshold < 50 eV**

Structure of the experiment



- R&D
- Electronics
- Gas studies
- Simulations
- Data analysis methods
- ...

- SEDINE detector in operation
- Competitive - will also act as the testbench for SNO detector

- The NEWS-G leading actor

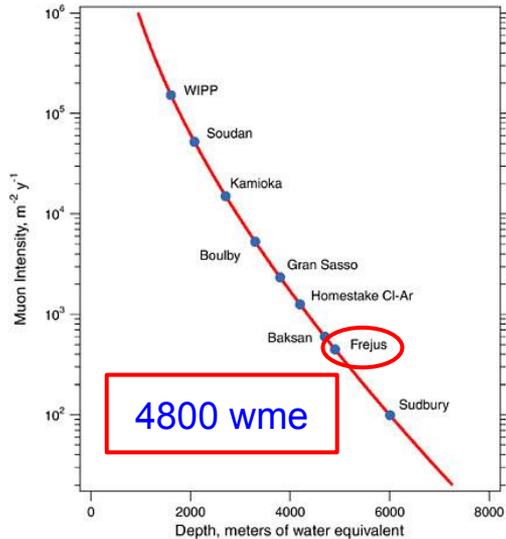
**LPSC
Grenoble**

Ionization Quenching Factor measurements using a low energy ion accelerator

Prototype low background SPC at LSM, Modane (SEDINE)

A competitive detector and a testing ground for NEWS-G / SNO

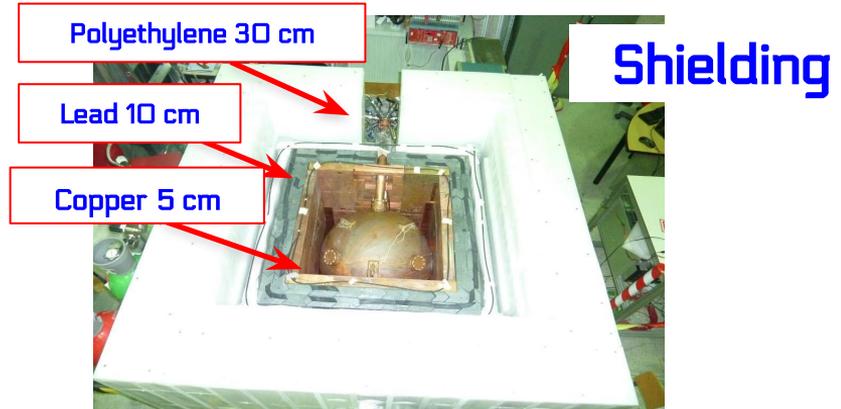
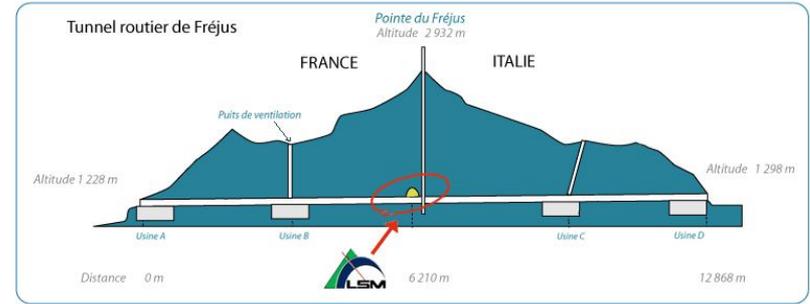
- Copper vessel (\varnothing 60 cm)
- Equipped with a 6.3mm \varnothing sensor
- Chemically cleaned several times for Radon deposit removal



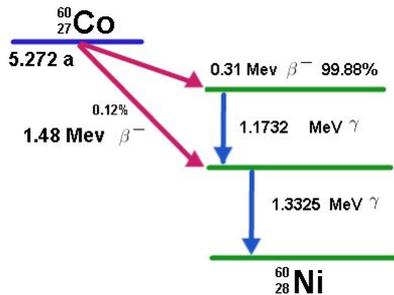
The SEDINE sensor



Laboratoire Souterrain de Modane



Main background sources for LSM detector



^{60}Co Contamination of 1 mBq/kg
BG Rate = 0.3-0.5 evnt/keV/kg/day

Solution: Limit time exposure on ground for pure copper.

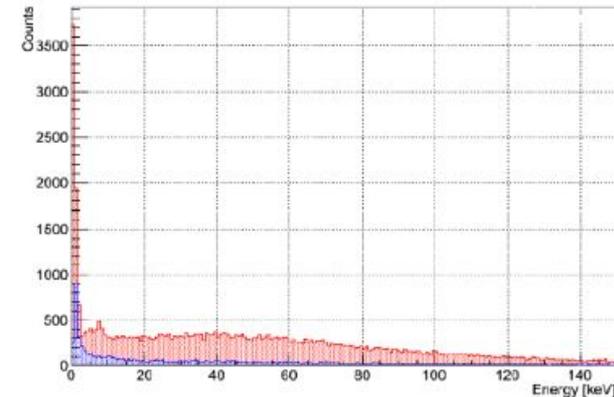
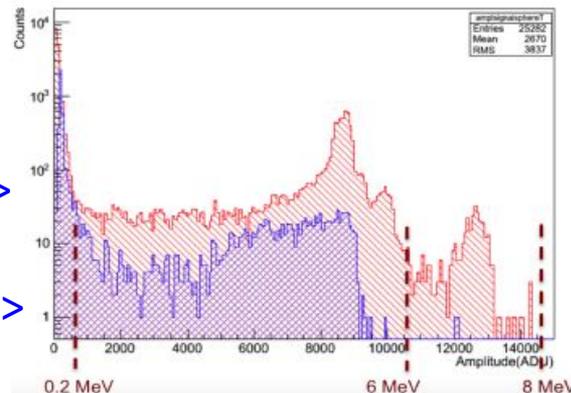
^{210}Pb , ^{210}Bi Contamination of 1 nBq/kg
BG Rate = 0.1 evnt/keV/kg/day

Competitive BG levels

Solution: Chemical cleaning

Effect of cleaning:

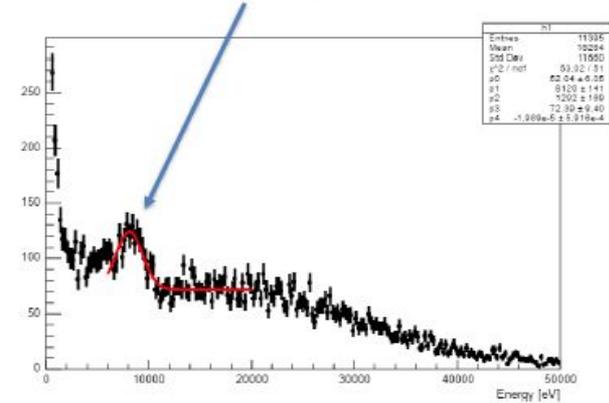
- High energy events 180 mHz => ~2 mHz
- Low energy events 400 mHz => ~20 mHz



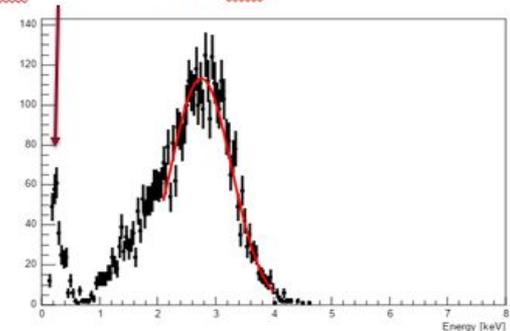
Operation and data taking conditions

- Data taking continuously during 42 days
- 99.3% Neon + 0.7 % CH₄ at 3.1 bar (310 gr)
- High voltage on anode set to 2520 V, no sparks
- Absolute Gain around 3000.
- Seal mode, no recirculation.
- Loss of gain 4% along 42 days
- Canberra charge sensitive preamplifier 2006 (RC=50 μs)
- Calibration with an ³⁷Ar gaseous source (from (n, α) reaction of ⁴⁰Ca) and the 8 keV fluorescence line of Copper

8 keV peak from Cu fluorescence



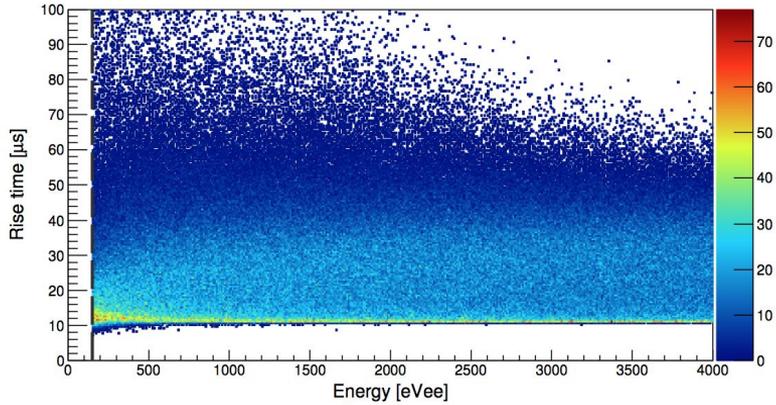
L capture, Auger e / X 0.27 keV K capture, Auger e / X 2.82 keV



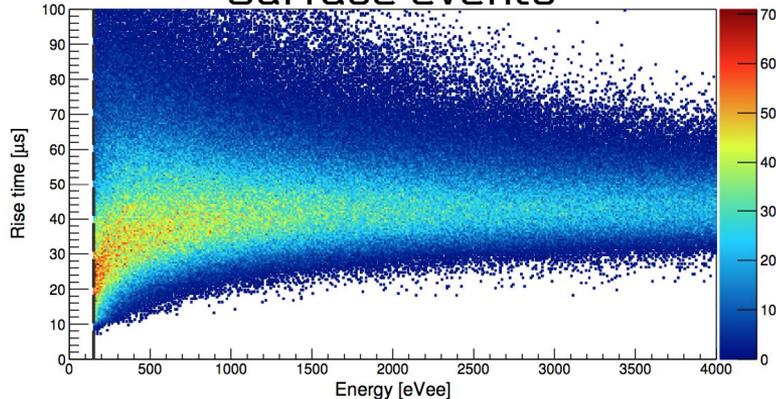
³⁷Ar X rays calibration

Simulation of volume and surface events

Volume events



Surface events



Anticipated main backgrounds:

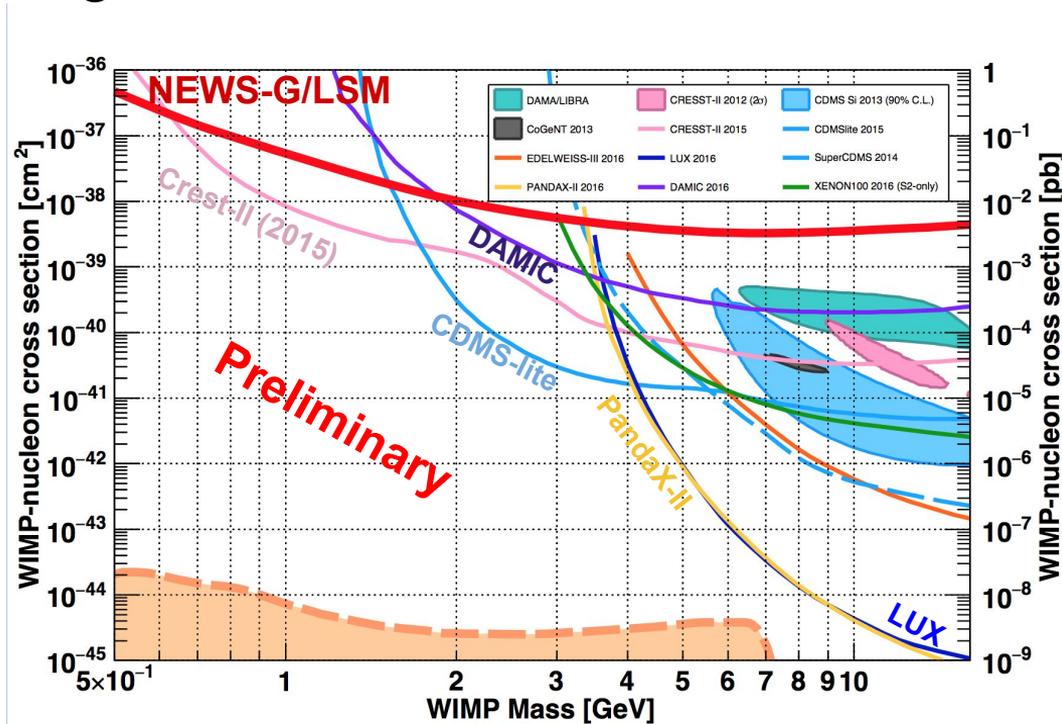
- Compton electrons (volume)
- ^{210}Pb decay products (surface)

Pulse simulations include:

- Electric field (FEM)
- Diffusion (Magboltz)
- Avalanche process
- Signal induction
- Preamplifier delta response

The Simulation results were used at input to a **Boosted Decision Tree (BDT)** algorithm to determine a fine tuned **ROI optimized for Signal/Background discrimination**

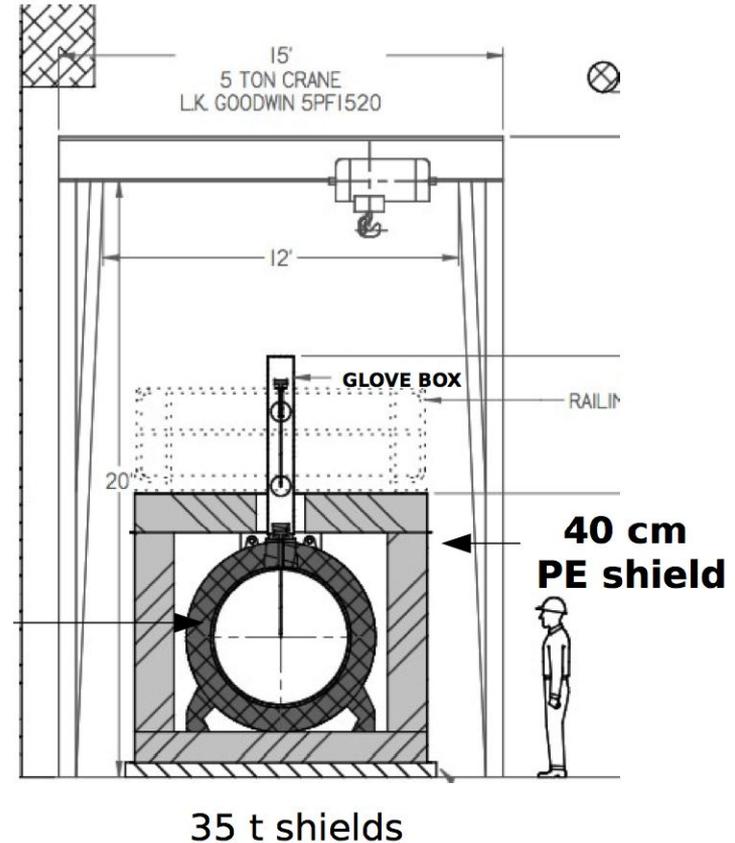
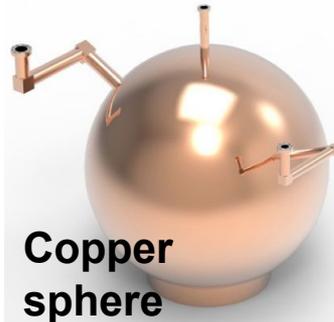
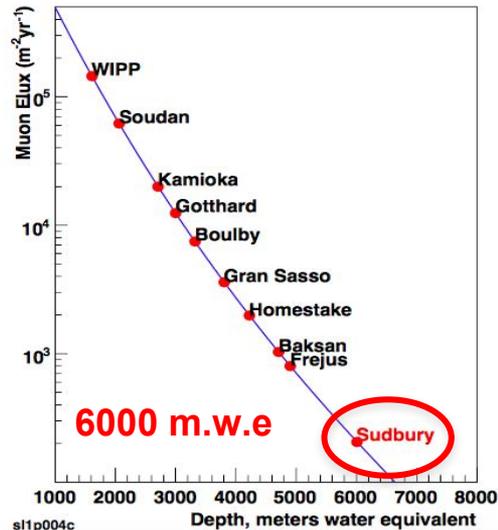
Exclusion limit of NEWS-G/ LSM for Spin Independent Wimp couplings



Limit set on spin independent WIMP coupling with standard assumptions on WIMP velocities, escape velocity and with quenching factor of Neon nuclear recoils in Neon calculated from SRIM

Implementation at SNOLAB by fall 2018

- 140 cm \varnothing detector, 10 bars, Ne, He, CH₄
- Low BG Copper Shell (12 cm thick)
- Shield: 40cm PE + Boron sheet + 25cm Lead + 3cm Ancient Lead



Estimated background

Simulation done with 12mm thick 140cm diam copper sphere full with 99% Ne 1%CH4, 11.43 kg of gas

Source Position	Mass (kg) or Surface (cm ²)	Source	evts/kg/day/[(μBq/kg) or (nBq/cm ²)]	contamination units	evts/kg/day < 1ke
CopperSphere	627.83 kg	Co60	0.0018	30 μBq/kg	0.054
CopperSphere	627.83 kg	U238	0.0036	3 μBq/kg	0.011
CopperSphere	627.83 kg	Th232	0.0049	12.9 μBq/kg	0.063
InnerSurface	57255 cm ²	Pb210	0.012	0.16 nBq/cm ²	0.002
ArchLead	2108.95 kg	U238	0.001	61.8 μBq/kg	0.062
ArchLead	2108.95 kg	Th232	0.0011	9.13 μBq/kg	0.010
Rod	0.0931721 kg	Co60	2.95E-007	30 μBq/kg	0.000
Rod	0.0931721 kg	U238	1.81E-006	3 μBq/kg	0.000
Rod	0.0931721 kg	Th232	2.11E-006	12.9 μBq/kg	0.000
Wire	2.66005e-05 kg	Co60	1.48E-010	31000 μBq/kg	0.000
Wire	2.66005e-05 kg	U238	2.12E-009	300000 μBq/kg	0.001
Wire	2.66005e-05 kg	Th232	1.42E-009	50000 μBq/kg	0.000
Wire	2.66005e-05 kg	K40	5.41E-010	1660000 μBq/kg	0.001
LabArea		T1208/K40			0.076

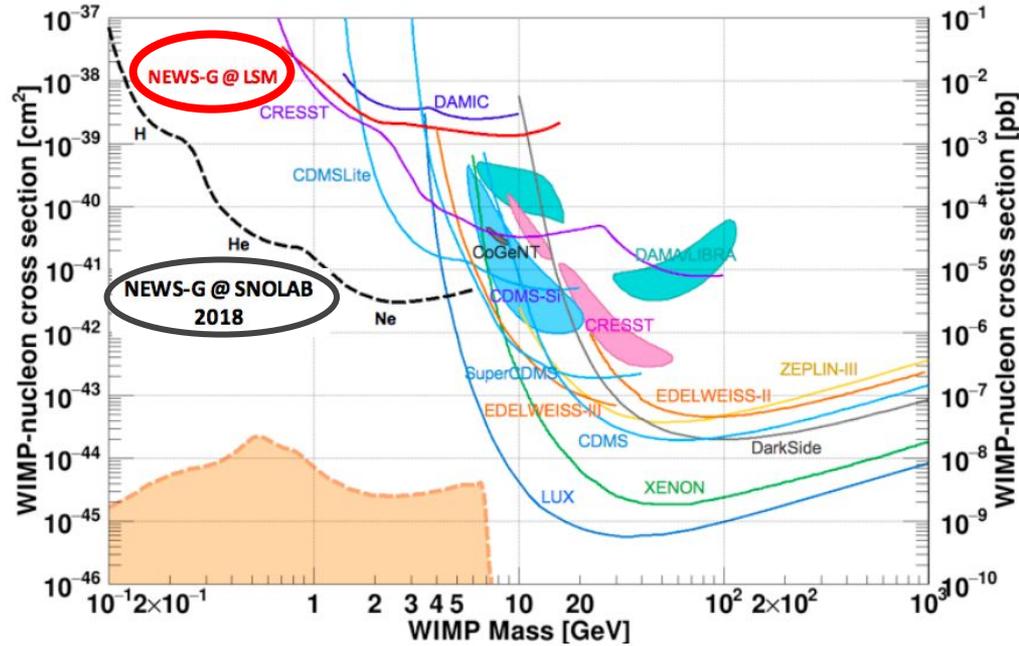
Total 0.279

Copper

Internal surface
Lead shield

External BG with SNO
Flux

Predicted exclusion limits for NEWS-G / SNO



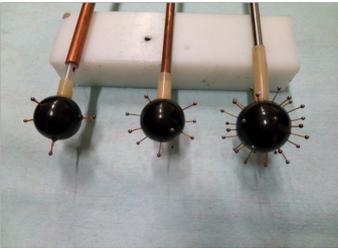
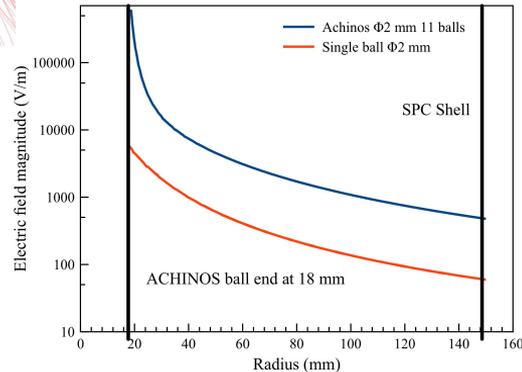
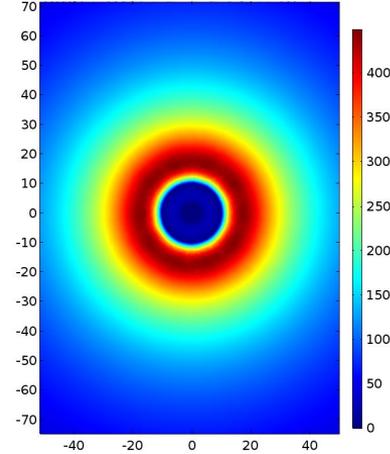
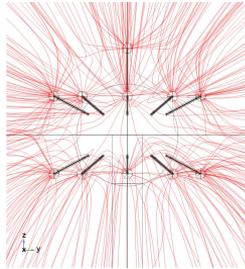
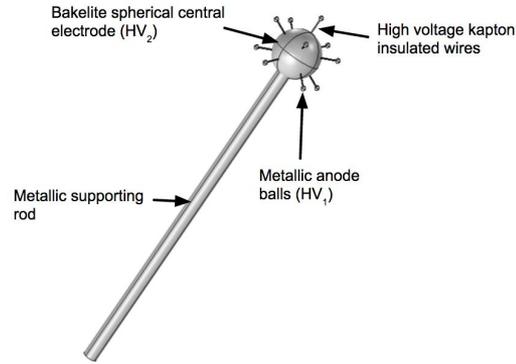
Hypothesis for NEWS-SNO expected sensitivity:

100 kg.days exposure @ 10 bar, Threshold 1 electron (~ 40 eV), 200eV window

A sensor upgrade

Sensor development - ACHINOS

If instead of one ball we use a number of them placed at equal distance on a sphere you can have the same gain but increased field at the outer region of the detector



AIM:

1. Operation in high pressure
2. Build larger volume detectors

Conundrum:

Both Gain and Drift time are a function of E/P

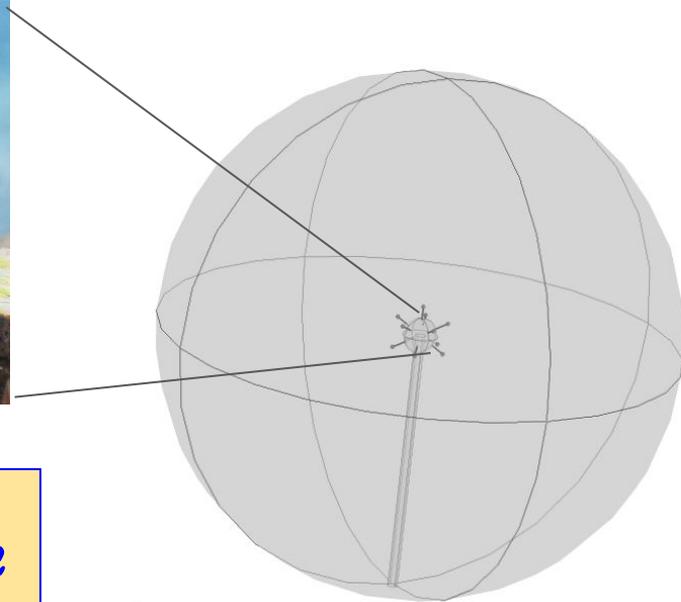
$$\ln M = \int_{E(r^1)}^{E(r^2)} a(E/P) \frac{dE}{E}$$

$$v_{drift} = \mu \frac{E}{P}$$

The elegant solution - ACHINOS

- Decoupling Gain - Drift
- Tunes Volume electric field
- Anodes can be read out individually

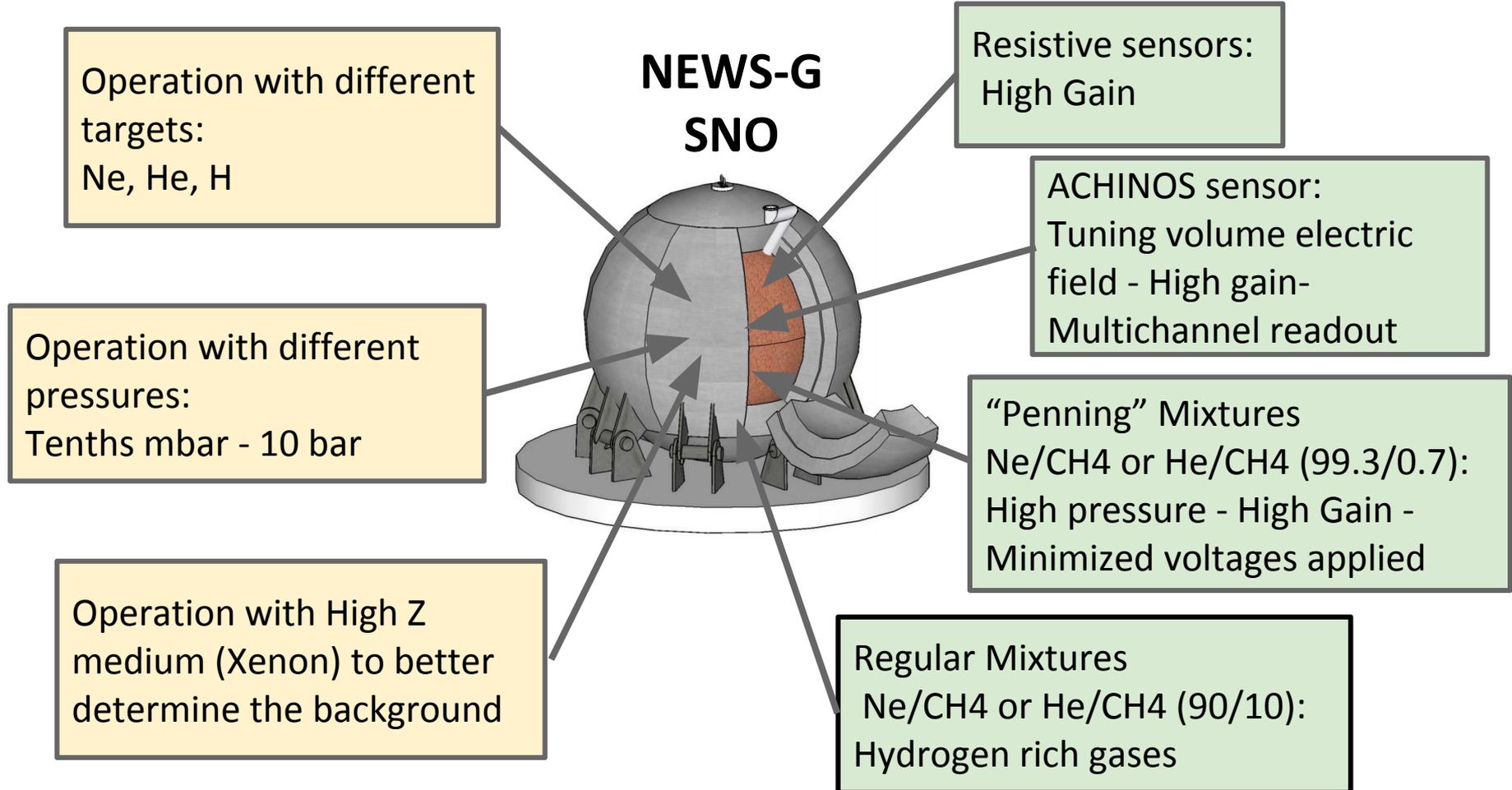
And just in case you wondered about the name ...



ACHINOS = Sea urchin = Riccio di mare

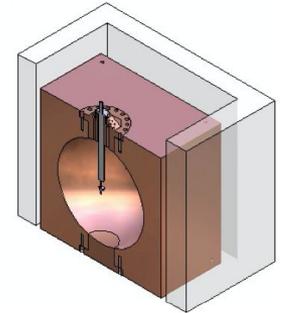


Detector versatility



Conclusions and Outlook

- The SPC is a simple, robust detector, providing background rejection capabilities
- The first very competitive gaseous detector in DM searches
- First runs with He/CH₄ mixtures have already started in LSM, Modane
- The Sedine detector is an excellent testing ground for NEWS-G/SNO
- The NEWS-G/SNO will have better Shield-Materials-Installation Procedure
- Construction is expected to start by fall 2017 and the installation by 2018
- R&D for the detector development
 - Underground electroformed sphere (PNNL)
 - Cubic sphere
- Some extra investigations with our detector
 - KK solar axions through 2 photon decay
 - Coherent neutrino nucleus elastic scattering (Reactor/Supernova neutrinos)





- **Queen's University Kingston** – G Gerbier, P di Stefano, R Martin, T Noble, D Dunford
A Brossard, A Kamaha, P Vasquez dS, Q Arnaud, K Dering, J Mc Donald, M Clark, M Chapellier
 - Copper vessel and gas set-up specifications, calibration, project management
 - Gas characterization, laser calibration, on smaller scale prototype
 - Simulations/Data analysis
- **IRFU (Institut de Recherches sur les Lois fondamentales de l'Univers)/CEA Saclay** -I Giomataris, M Gros, C Nones, I Katsioulas, T Papaevangelou, JP Bard, JP Mols, XF Navick,
 - Sensor/rod (low activity, optimization with 2 electrodes)
 - Electronics (low noise preamps, digitization, stream mode)
 - DAQ/soft
- **LSM (Laboratoire Souterrain de Modane), IN2P3, U of Chambéry** - F Piquemal, M Zampaolo, A Dastgheibi Fard
 - Low activity archeological lead
 - Coordination for lead/PE shielding and copper sphere
- **Thessaloniki University** – I Savvidis, A Leisos, S Tzamaras, C Eleftheriadis, L Anastasios
 - Simulations, neutron calibration
 - Studies on sensor
- **LPSC (Laboratoire de Physique Subatomique et Cosmologie) Grenoble** - D Santos, JF Muraz, O Guillaudin
- **Technical University Munich** – A Ulrich, T Dandl
 - Gas properties, ionization and scintillation process in gas
- **Pacific National Northwest Lab** – E Hoppe, D Asner
 - Low activity measurements, Copper electroforming
- **RMCC (Royal Military College Canada) Kingston** – D Kelly, E Corcoran
 - ^{37}Ar source production, sample analysis
- **SNOLAB –Sudbury** – P Gorel
 - Calibration system/slow control
- **University of Birmingham** – Kostas Nicolopoulos
 - Simulations, analysis, R&D
- **Associated lab : TRIUMF** - F Retiere
 - Future R&D on light detection, sensor





Thanks for your attention !