

#### Summary talk of the collider session

Walter M. Bonivento

**INFN Cagliari** 

also on behalf of Eder Izaguirre who could not come to ther meeting but organised the session together with me!











17:45 - 19:00	IV Coll 17:45	iders B <b>DM Theory at Neutrino Facilities</b> 20' Speaker: Dr. Claudia Frugiuele (Weizmann institute) Material: <b>Slides</b>
	18:05	SHiP Dark Sector searches 20'Speaker:Laura Fabbri (BO)Material:Slides
	18:25	DM searches at Belle/Belle II 20' Speaker: Enrico Graziani (ROMA3) Material: Slides
	18:45	Discussion 15'





The dark sector, i.e. portals, and DM at accelerators, are relatively new research fields

Many experiments that were designed with other goals in mind (such as the LHC ones) are discovering now their physics potential for dark sector and DM particle searches

I do not personally believe (as sometimes is said by ourselves) that the dark sector would not be a strong enough physics case on its own

Indeed can we classify new physics ?? is a neutralino worth two axions or three? nobody knows where the truth is (especially now) and therefore we should keep an open mind and consider whatever theoretically allowed as viable for experiments (if we can afford them!) and leave no stones unturned





### A different way to search for NP!



## HS portals is yet another different way!





# Dark sector particles can potentially solve many outstanding issues in particle physics such as

Dark Matter: WIMP-like particles, dark photons, axions, sterile neutrinos

**Baryogenesis: Sterile neutrinos** 

Neutrinos masses: sterile neutrinos, see-saw





In this WG, we discussed dark sector particles searched with:

- pp collisions at LHC
- Pb-Pb collisions at LHC
- e+e- collisions at Belle2
- pp collisions in fixed target at SHiP

Apart from the last one, clearly pp and e+e- experiments are there and designed for other purposes, but are also providing and will provide for a long time very relevant information about dark sector particles (LHCb and Babar slipped into other sessions but they belong to the group!)

SHiP is indeed the first experiment at CERN designed for dark sector particle studies





Light Dark Matter 2017

24-28 May 2017

La Biodola - Isola d'Elba

# SEARCHES FOR DARK MATTER AT ATLAS AND CMS

Rocío Vilar (for Atlas and CMS collaboration) Instituto de Física de Cantabria (IFCA)





### RIMENTAL SETUP

INFN



LDMA ELBA 28 MAY 2017



## INTRODUCTION

- Look everywhere, wide range and generic signatures
  - Dark Sector: dijets, dileptons, displaced, portals, etc (see <u>M. Vertucci´s talk</u>)
  - Dark Matter candidates: pp-> Met + X (mono-X search)







## SEARCHES AT LHC

	ATLAS				<u>CMS</u>		
		ref.	data	L(fb <sup>-1</sup> )	ref	data	L(fb <sup>-1</sup> )
Mono-γ		arxiv new 1704.03848	2015+2016	36.1	CMS-PAS- EXO-16-039	2016	12.9
Mono-V		Phys. Lett. B 763 (2016) 251	2015	3.2	<u>CMS-PA</u> <u>EXO-16-048</u>	2015+2016	35.9
Mono-jet		PRD 94 (2016) 032005	2015	3.2			
mono- Higgs	bb	ATLAS new CONF-2017-028	2015+2016	36.1	arXiv 1703.0	2015	2.3
	γγ	ATLAS new CONF-2017-024	2015+2016	36.1	<u>CMS-PAS-</u> EXO-17-054	2015+2016	35.9
	ZZ	ATLAS- CONF-2015-059	2015	3.2			
mono-Z		ATLAS- CONF-2016-056	2015+2016	13.3	CMS-PAS new EXO-16-052	2016	35.9
mono-top					<u>CMS-PAS-</u> EXO-16-040	2015+2016	12.9
mono- tt	had	ATLAS- CONF-2016-077	2015+2016	I 3.3 CMS-PAS-		2015	2.2
	semi-l	ATLAS- CONF-2016-050	2015+2016 13.2		EXO-16-005		
	lep	<u>ATLAS-</u> CONF-2016-076	2015+2016	13.3	CMS-PAS- EXO-16-028	2015	2.2
mono-bb		ATLAS- CONF-2016-086	2015+2016	13.3	<u>CMS-PAS-</u> <u>B2G-15-007</u>	2015	2.17





### SUMMARY OF DM +X SEARCH

95% CL exclusion region in the mediator- DM mass plane for different MET+X searches for spin-1/0 mediators with couplings

$$g_q = 0.25, g_l = 0, g_{DM} = 1$$
 or  $g_q = 0.1, g_l = 0.1, g_{DM} = 1$ 

The <mark>mono-jet</mark> has better sensitivity with full 2016+2015 lumi. Recast of MET+X analysis to set DM-nucleon limits:

- SD the strongest
- SI the strongest below 5.5 GeV
- this is only valid for the coupling and models chosen





the comparison with direct searches is model dependent —> important lesson: we need to pursue direct searches for low mass DM anyway!







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# Low Mass Dark Sector Searches at ATLAS and at CMS





Monica Verducci

INFN and University of Roma Tre





# Dark Photons in Lepton Jets (LJ)

#### Low-mass dark photons can be produced via cascade decays of heavier states

Benchmark models predicting Higgs boson or heavy scalar particle decays to neutral long-lived particles (dark photons  $\gamma_d \gamma_d$  of 0.4 GeV Mass each), which in turn produce:

 collimated jets of light leptons and mesons, so-called "lepton-jets"

The Low-mass of the Dark Photon produces:

- -> large boost
- -> collimated decay products

# Dark photon lifetime depends on the size of kinetic mixing $\varepsilon$ : small $\varepsilon \rightarrow$ displaced decays

Leptonic decays prominent over wide (low) mass range. The dark photons and two LJs are expected to be produced back-to-back in the azimuthal plane.





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# does it mean exclusion region? NB: here we add the additional assumptions on H decays



# Light Bosons Decaying into Muon Pairs

- This analysis searches for new light bosons with a mass in the range 0.25-8.5  $GeV/c^2$  decaying into muon pairs.
- The results are interpreted in the context of two benchmark models, namely, the nextto-minimal supersymmetric standard model, and dark SUSY models including those predicting a non-negligible light boson lifetime.

#### **Event Identification:**

main background is bb events. Background is modelled as a two dimensional (2D) template bb (m1, m2) in the plane of the invariant masses of the two dimuons in the selected events, where (1) always refers to the dimuon containing a muon with pT > 17 GeV/c and  $|\eta| < 0.9$ .



 $\mu^+$ 

H ->aa(ss)-> 4SM (or 2SM+X)

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CMS-PAS-HIG-16-035

 $n_1$ 

 $\gamma_D$ 

 $n_D$ 



#### does it mean exclusion region? model dependent but shows that it would be in a region that is not already excluded









# Heavy-ion as a $\gamma$ source

- ♦ Old idea A. Balantekin et.al. 1985, M. Greiner et.al. 1993...
- Enhanced production at Heavy-ion collisions (~GeV)
- ♦ QED is strongly coupled!
  ♦ Z<sup>4</sup> ~ 50 × 10<sup>6</sup> enhancement



♦ Can observe LBL scattering!





# Signal & Background

#### Signal

- Intact Pb-Pb ions, no tracks, very little calorimeter activity
- ♦ Veto on tracks
- Two ~GeV back to back
  photons (cut on Δφ<sub>γγ</sub>)
  - Otherwise ions will likely breakup
- ♦ Intact ions may be tagged
- Prominent mass peak!

#### Background

- ♦ LBL scattering
  - ♦ Irreducible
  - Has been measured! Can be calculated reliably
- Electrons fakes/brem or CEP (photons from QCD)
  - ♦ Photons not back to back
- \* All background smoothly falling in  $m_{\gamma\gamma}!$

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# ATLAS recast

- $\sqrt{s_{NN}} = 5.02 \text{ TeV}$
- $\& \int Ldt = 480 \ \mu b^{-1}$
- Current best limit on ALPs with EM coupling!
- First time heavy-ion yields best limit on BSM physics













#### Light Dark Matter 2017 24-28 May 2017 La Biodola - Isola d'Elba

# SHiP Dark Sector searches

Laura Fabbri on behalf of the SHiP Collaboration







this experiment is designed to detect dark sector particles (it is not the only case as we learned also in this conference)



#### 0,<sup>101</sup> 0,<sup>103</sup> B (visible) CHARM B (invisible) CHARM 10



Primary Physics goals

- independent measurement of  $\sigma_{v_r}$  and  $\sigma_{v_r}$ 

1. Exploring hidden portals and extensions of the SM:

Laura Fabbri - Università di Bologna - INFN





- searches for Dark Matter, Sterile Neutrinos and Dark Photons => long-lived

and very weakly interacting particles through their decays to SM particles



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# Sensitivity to Hidden Sector

**BENCHMARK SCENARIO** 

TO BARYOGENESIS SHiP sensitivity to HNLs SHiP sensitivity to HNLs SHiP sensitivity to HNLs 10 10 10 CHARM. NuTeV 💛 CHARM PS191 BAU PS191 BAU °5<sup>10°</sup> ° 10 MS at 10 PS191 BAU 10 HNL coupling to SM U<sup>2</sup> NuTeV WS pt 10 BBN BBN HNL coupling 1 Bujdnoo BBN Ŧ 10 10 BAU / Seesaw BAU / Seesaw BAU / Seesaw 10"1 10-11 10 10 10 HNL mass (GeV) HNL mass (GeV) HNL mass (GeV)  $U_e^2: U_\mu^2: U_\tau^2 = 52:1:1$  $U_{\epsilon}^{2}: U_{\mu}^{2}: U_{\tau}^{2} = 1:16:3.8$  $U_{\epsilon}^{2}: U_{\mu}^{2}: U_{\tau}^{2} = 48:1:1$ 10-2 previcus experimental limits 101 10 e coupling to SM 101 10<sup>-6</sup> Sensitivity to dark photon as a function of the 10 mixing with the SM photon and the dark 10  $p \rightarrow p+\gamma'$ photon mass 101 mesons  $\rightarrow \gamma' X$ >−10<sup>-10</sup> <sup>4</sup>He BBN excluded areas 10-11 10<sup>-12</sup> D/H .i/H 10"2 2×10"2 10<sup>-1</sup> 2×10<sup>-1</sup> Technical Proposal CERN-SPSC-2015-016 2 3 4 5 3×10 1 γ' mass (GeV) Laura Fabbri - Università di Bologna - INFN



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INVERTED HIERARCHY SCENARIO

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SCENARIO SENSITIVE

# Dark Matter Search at Belle/Belle2

## Enrico Graziani

INFN – Roma 3

on behalf of the Belle II Collaboration



#### **OUTLINE OF THE TALK**

- Belle II and SuperKEKB
- Search of the dark photon invisible decay in Phase 2
- Alternative LDMA searches
- Search of ALP
- Summary

# Light Dark Matter 2017

La Biodola – Isola d'Elba, 24-28 May 2017







## **Breaking news**

## Belle II succesfully rolled in on April 11th

#### perfectly in time with plans (and sakura)



E. Graziani – Dark Matter at Belle II – LDMA 2017



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## Belle II + SuperKEKB schedule



I N F N

# Single photon trigger(s)

Two level 1 single photon triggers, both excluding the innermost ECL crystal towers

•  $E_{\gamma 1}^{CMS} > 1 \text{ GeV}$ , with the second cluster energy  $E_{\gamma 2}^{CMS} < 0.2 \text{ GeV} \longrightarrow \sim 4 \text{ nb}$ 







## **Belle II projection (simulation)**







### very exciting times ahead of us!

